

Virtual DPG Spring Meeting 2022

of the Matter and Cosmos Section (SMuK)

with its Divisions

Extraterrestrial Physics; Gravitation and Relativity; Particle Physics; Radiation and Medical Physics; Theoretical and Mathematical Physics;

and the further Divisions History of Physics; Physics Education;

as well as the Working Groups

Physics, Modern IT and Artificial Intelligence; Philosophy of Physics



21 – 25 March 2022

heidelberg22.dpg-tagungen.de

Verhandlungen der Deutschen Physikalischen Gesellschaft (ISSN 2751-0522 [Online]) Reihe VI, Band 57 (2022) Zitiertitel: Verhandl. DPG (VI) 57, 2/2022 Erscheinungsweise: Jährlich 3 - 6 Online-Hefte, je nach Bedarf

Verantwortlich für den Inhalt:

Dr. Bernhard Nunner, DPG e. V., Hauptstraße 5, 53604 Bad Honnef Telefon: +49 (0)2224 9232-0, Telefax: +49 (0)2224 9232-50 © Deutsche Physikalische Gesellschaft e. V., 53604 Bad Honnef

Content

Greeting	
Organisation	
Organiser	
Local Organiser	
Scientific Organisation	
Head of the Programme Committee	4
Chairs of the Participating Divisions of SMuK	4
Chairs of further Participating Divisions	5
Chairs of the Participating Working Groups	5
Symposia	
Information for Participants	
Conference Location	б
Conference Time Zone	6
Conference Website	6
Conference Office	б
Technical Requirements	6
How to use the Conference Platform MeetAnyway	7
Notice Board	7
Wilhelm and Else Heraeus Communication Programme	7
Information for Speakers	7
Information for Poster Presentations	7
Social Events	
Opening	
Virtual visits of Facilities of Particle and Astro-Particle Physics	
Public Evening Lecture	
Award Presentation of the SMuK Dissertation Prize	
Annual Meeting of Young Scientists in High Energy Physics (yHEP) in 2022	
Annual General Meetings of the DPG Divisions and Working Groups	
Synopsis of the Daily Programme	
Programme	
Plenary and Evening Talks	
Symposia	
SMuK Dissertationspreis 2022 (SYMD)	25
Awards Symposium (SYAW)	
The Nature of Science (SYNS)	29
Evisions Extraterrestrial Physics (FP)	21
Gravitation and Relativity (GR)	
Theoretical and Mathematical Physics (MP)	57
Radiation and Medical Physics (ST)	66
Particle Physics (T)	75
Physics Education (DD)	192
History of Physics (GP)	
Working Groups	<i>-</i>
Devoice Modern IT and Articial Intelligence (AKDIK)	000
Philosophy of Dhysios (ACDbil)	
FTIIIUSUPHY UFFHYSICS (AUFTIII)	
Authors	

Dear conference guests,

On behalf of the German Physical Society and also personally, I would like to welcome you to the virtual DPG-Frühjahrstagung (DPG Spring Meeting) of the Matter and Cosmos Section (SMuK) in Heidelberg.

I am very pleased that this Spring Meeting continues to take place despite the pandemic. Maintaining scientific exchange cannot be valued highly enough. Just as important in these times are the DPG conferences once again as outstanding symbols of the importance of scientific thinking in our society: Natural science produces hypotheses that have to be verified experimentally – that is the core of basic research.

I agree with Niels Bohr, who is said to have said: "Forecasts are difficult, especially when they concern the future." Nevertheless, I see physics and the DPG in particular as having a special responsibility to enter into a dialogue with politics on the basis of the findings from basic research in order to meet the major challenges facing society – and thus also to enable future generations to live well on this planet. For this dialogue, the solidarity of the scientific community with the colleagues who dare to go public and stead-fastly represent their results is particularly crucial.

I would like to express my sincere thanks to all those involved for the success of this Spring Meeting. First of all, I would like to thank the local conference management, Prof. Ulrich Uwer, Universität Heidelberg, Physikalisches Institut, for its support as well as the programme committee – consisting of the chairpersons of the divisions and working groups involved – for the outstanding programme of this conference. I would also like to thank the staff of the DPG Head Office for their support and supervision of all meetings.

I would also like to express my sincere thanks to the Wilhelm and Else Heraeus-Stiftung for again providing generous financial support to our young members.

I wish you all an exciting conference and many new insights.

Dr. Lutz Schröter President of the Deutsche Physikalische Gesellschaft e.V.

Organisation

Organiser

Deutsche Physikalische Gesellschaft e. V. Hauptstraße 5, 53604 Bad Honnef Phone +49 (0) 2224 9232-0 Email dpg@dpg-physik.de Homepage www.dpg-physik.de

Local Organiser

Prof. Dr. Ulrich Uwer Universität Heidelberg Physikalisches Institut Im Neuenheimer Feld 226, 69120 Heidelberg Email uwer@physi.uni-heidelberg.de

Scientific Organisation

Head of the Programme Committee

Prof. Dr. Karl-Henning Rehren Institut für Theoretische Physik Universität Göttingen Friedrich-Hund-Platz 1, 37077 Göttingen Email rehren@theorie.physik.uni-goettingen.de

Chairs of the participating Division of the Matter and Cosmos Section (SMuK)

Extraterrestrial Physics Division (EP)

Prof. Dr. Miriam Sinnhuber Karlsruher Institut für Technologie Hermann-von-Helmholtz Platz 1, 76344 Eggenstein-Leopoldshafen Email miriam.sinnhuber@kit.edu

Gravitation and Relativity Division (GR)

Prof. Dr. Bernd Brügmann Theoretisch-Physikalisches Institut Universität Jena Max-Wien-Platz 1, 07743 Jena Email bernd.bruegmann@uni-jena.de

Theoretical and Mathematical Physics Division (MP)

Prof. Dr. Johanna Erdmenger Lehrstuhl für Theoretische Physik III Am Hubland, 97074 Würzburg Email: erdmenger@physik.uni-wuerzburg.de

Radiation and Medical Physics Division (ST)

Anna Bakenecker Fraunhofer IMTE Mönkhofer Weg 239a, 23562 Lübeck Email bakenecker@dpg-mail.de

Particle Physics Division (T)

Prof. Dr. Kerstin Borras Deutsches Elektronen-Synchrotron DESY Notkestraße 85, 22607 Hamburg Email kerstin.borras@desy.de

Chairs of further participating Divisions

Physics Education (DD)

Prof. Dr. Susanne Heinicke Institut für Didaktik der Physik Westfälische Wilhelms-Universität Wilhelm-Klemm-Str. 10, 48149 Münster Email: susanne.heinicke@uni-muenster.de

History of Physics (GP)

PD Dr. Christian Forstner FSU Jena – Ernst-Haeckel-Haus AG Wissenschaftsgeschichte Kahlaische Str. 1, 07745 Jena Email: christian.forstner@uni-jena.de

Chairs of the participating Working Groups

Philosophy on Physics (AGPhil)

Prof. Dr. Dennis Lehmkuhl Universität Bonn Institut für Philosophie Am Hof 1, 53113 Bonn Email dennis.lehmkuhl@uni-bonn.de

Physics, Modern IT and Artificial Intelligence (AKPIK)

Dr. rer. nat. Tim Ruhe TU Dortmund Experimentelle Physik 5 Otto-Hahn-Straße 4 a, 44227 Dortmund Email: tim.ruhe@tu-dortmund.de

Symposia

SMuK Dissertation Prize 2022 (SYMD)

Organisation: Claus Lämmerzahl, ZARM Bremen

Awards Symposium (SYAW)

Organisation: Karl-Henning Rehren, Universität Göttingen

The Nature of Science (SYNS)

Organisation: Susanne Heinicke, Universität Münster

Programme

The scientific programme consists of 1.356 contributions:

9	Plenary Talks	1	Workshop
1	Evening Talk	24	Group Reports
3	Prize Talks	1150	Talks
56	Invited Talks	88	Posters

24 Topical Talks

Information for Participants

The virtuell conference will be held in the period 21–25 March, 2022.

Conference Location

Web-based Conference – Login information will be provided a few days before the event starts.

Conference Time Zone

All times mentioned on the website and in the programme are in Central European Time (CET), UTC+1.

Conference Website

https://heidelberg22.dpg-tagungen.de/

Conference Office

The virtual conference office is situated on the conference platform and will be open daily from 08:30 - 16:00 (Friday 08:30 - 12:00) for questions round the conference. You will find it on the conference platform under the *"Welcome"* tab immediately after signed up.

Technical Requirements

The MeetAnyway platform will be used for the conference. In order to participate in the conference, you need a MeetAnyway account with which you can register on the conference platform.

If you do not have a MeetAnyway account yet, please create one on meetanyway.com in good time before the start of the conference – using the email address you used for participant registration. To do this, click on "Create New Account" and follow the instructions on the screen. During registration, you will receive a 6-digit code by email from MeetAnyway to activate your account.

After creating the account, please add a profile picture and your affiliation to your personal profile.

To use all features of the conference platform, you need an up-to-date browser. Chrome is currently the most stable and reliable browser for using the conference platform. Firefox and Safari are browsers that should work but are often less performant. MeetAnyway support staff is highly trained in resolving Chrome issues. If you are using a different browser (e.g. Firefox, Safari or Edge), the support staff cannot provide in-depth troubleshooting support for you.

In case you have not received the activation code for your MeetAnyway account or you have technical difficulties on the meeting platform please contact the MeetAnyway support staff

- by email: support@meetanyway.com
- · via the participant helpdesk: https://help.meetanyway.com or
- directly on the conference platform via the (?) symbol at the right top.

All lectures will be held and broadcast via Zoom video conferencing service. For the best experience we recommend that you download or update to the latest version of the Zoom client for meetings before the start of the conference. A Zoom account is not required to use the application. Alternatively, joining the Zoom Meetings is also possible via all common browsers (Chrome, Firefox, Safari and Edge).

A video chat room is linked to each poster, where you can discuss in small groups during the poster sessions.

In addition, numerous video chat rooms are offered for exchange and networking.

For video chats, permission to access your microphone and camera is required. Please note that firewalls of company or institute networks can limit the functionality.

How to use the Conference Platform MeetAnyway

For using the platform, you will find detailed step-by-step instructions at *heidelberg22.dpg-tagungen.de/ tagungsplattform*.

Notice Board

All changes regarding the schedule of the conference will be updated currently. The information is identical to the programme updates of the scientific programme and available at the scientific programme in other formats as well (ordered by publication date, filterable by conference part and as an rss-feed). Please use the form at *https://heidelberg22.dpg-tagungen.de/programm/notice-board-form* to submit amendments, cancellations, etc.

Wilhelm and Else Heraeus Communication Programme

Within this programme, the active participation by young DPG members – from Germany and abroad – at the virtual DPG Meetings is financially supported.

For the virtual DPG-Meetings, the conference fee (exclusively the "early bird rate") is subsidised at 100% (*submission of an application was open until 17 February 2022. Subsequent applications are not possible*). After the conference, your participation in the conference will be checked on the basis of the login data and the funding will be finally confirmed or rejected if no participation took place. Payment will be made – after prior notification by email – by the end of April 2022 at the latest by bank transfer to the account you specified in your application.

The Deutsche Physikalische Gesellschaft thanks the Wilhelm and Else Heraeus-Stiftung for the generous financial support of young academic talents. We hope that young physicists will continue to seize the offered opportunity for active scientific communication at scientific conferences. A total of about 37,800 young academics were supported by this programme so far.

Information for Speakers

All speakers are invited to use our offer for a test session one week before the conference starts. The necessary information for the test session about day, time and login information will be sent out by email to the speakers. We would like to ask you to consider the following points for your presentation:

- Please use the same equipment with which you successfully completed your technical check to avoid technical problems during your presentation.
- Please be in the Zoom session of the virtual room where you will give your presentation at least 10 minutes before the session starts. Access for session chairs, speakers and participants to the individual sessions is via the virtual rooms on the conference platform (via the "Join" option). No separate login information will be sent to presenters.
- Please sign in at Zoom with your full name so that the technical support and the conference organisers can identify you as a speaker and give you the rights to share your screen, microphone and camera in Zoom.
- Please make sure that you respect your presentation time!

Information for Poster Presentations

A poster presentation can consist of up to six files – directly visible to the poster visitor – with the following requirements:

- Poster file as PDF without format restriction up to a size of 10 MB
- Image file as PNG, JPG or GIF in 4:3 format (min.1600x1200) up to a size of 10 MB
- animated GIF file in 4:3 format (min.1600x1200) up to a size of 10 MB

The criteria are based on the technical requirements of the conference platform. Therefore, different file formats are not possible.

We recommend creating an image file as a preview image and creating the poster as a PDF file in classic portrait format (DIN A0). If only a poster file is created, this also functions as the preview image.

In addition, up to six further files (in all common formats) of 100 MB each can be attached as downloads

to the presentation. Upon receipt of the login data – a few days before the start of the conference – the upload of the created file(s) is possible for the authors.

Once the uploads have been approved by the conference organisers, the posters will be available to all registered conference participants throughout the conference via the password-protected conference platform.

Presenting authors are requested to be available to answer questions and discuss via group video chat during the entire poster session at their poster.

Social Events

Opening

by the Local Organiser Prof. Dr. Ulrich Uwer, Universität Heidelberg Monday, 21 March, 11:30, Audimax. All participants are kindly invited.

Virtual visits of Facilities of Particle and Astro-Particle Physics

Monday - Thursday, 18:30 - 19:30

Live visits of the LHC-experiments ATLAS, CMS und LHCb, of the Belle II detector, the KATRIN and XENON experiments, the SNOLAB underground facility as well as of the H.E.S.S. and the of IceCube observatories are planned.

The virtual visits will be broadcasted via Zoom and the detailed programme can be found at heidelberg22. dpg-tagungen.de/veranstaltungen. The conference platform will take you directly to the corresponding Zoom rooms.

Particle Slam – The Science Slam for Particle Physics

Tuesday, 22 March 2022, 20:00 - 21:30, Raum PEL

Presentations on scientific topics are boring? Lay people cannot understand complex science anyway and they are also not interested in it? This may be a wide spread opinion – but that this is not always true is shown by Science Slams!

With the Particle Slam, we are inviting five early career researchers from particle physics and other neighbouring disciplines to the stage. Each of them has ten minutes to present their research topic to a general audience – easy to understand, accessible as well as in an entertaining way. Whoever does this best is decided upon by the audience who selects a winner of the presentation competition.

And last but not least: the slam also is an open stage: those who are for example PhD students in (particle) physics and are interested in sharing their research topic with the public can join by contacting the moderator via email: philipp@schroegel.de

Public Evening Lecture (in German language)

Wednesday, 23 March 2022, 20:00 - 21:00

Prof. Dr. Stephanie Hansmann-Menzemer, University of Heidelberg, is speaking about "Kleine Teilchen – große Chancen: Mit Quantenkorrekturen auf der Suche nach neuen Physikphänomen"

The Public Evening Lecture is open to all conference participants and interested public.

Award Presentation of the SMuK Dissertation Prize

Four selected finalists will give their presentations at the SMuK Dissertation Prize 2022 symposium (SYMD). The Award Presentation will take place on Wednesday, 23 March at 14:00 in the Audimax.

Annual Meeting of Young Scientists in High Energy Physics (yHEP) in 2022

On Wednesday, 23 March 2022, 19:30 – 20:30, the annual 2022 meeting of young scientists in high energy physics (yHEP) will take place online. All doctoral candidates, post-docs and scientists on temporary contracts are cordially invited.

We will present our activities from the last year and would like to discuss plans for the coming year with you and hear your ideas and thoughts. Topics are current and future developments in high and low energy physics, i.e. particle, astroparticle, hadron and nuclear physics, as well as accelerator physics, including topics of the situation of early-career researchers, environmental sustainability, networking and shaping the future of our fields.

Please register to our mailing list which can be found from https://yhep.desy.de to receive details on the meeting.

Annual General Meetings of the DPG Divisions and Working Groups

Division	n / Working Group	Date	Time	Location
(DD)	Physics Education	Tuesday, 22 March	18:00 - 19:30	DD-MV
(EP)	Extraterrestrial Physics	Thursday, 24 March	12:45 - 13:45	EP-MV
(GP)	History of Physics	Tuesday, 22 March	17:45 – 19:45	GP-MV
(GR) (MP)	Gravitational and Relativity Theoretical and Mathematical	Thursday, 24 March	19:00 - 20:30	GR-MV
	Physics	Thursday, 24 March	19:30 - 20:30	MP-MV
(ST)	Radiation and Medical Physics	Tuesday, 22 March	16:15 - 17:30	ST-MV
(T)	Particle Physics	Thursday, 24 March	19:30 - 21:00	T-MV
(AKPIK)	Physics, Modern IT and			
(AGPhil)	Artificial Intelligence Philosophy of Physics	Wednesday, 23 March Wednesday, 23 March	19:00 – 21:00 18:30 – 19:30	AKPIK-MV AGPhil-MV

Synopsis of the Daily Programme

Monday, March 21, 2022

11:30	Audimax		Opening	
11:40	Audimax	PVI	Plenary Talk Making uncertainty certain?: Proton therapy, range and in-vivo verificatio •Tony Lomax	ın.
			SY	'MD
			Invited Talks	
14:00	Audimax	SYMD 1.1	Timeless Quantum Mechanics and the Early Universe	
14:25	Audimax	SYMD 1.2	First tritium β -decay spectrum recorded with Cyclotron Radiation Emissi Spectroscopy (CRES)	on
14:50	Audimax	SYMD 1.3	Watching the top quark mass run – for the first time!	
15:15	Audimax	SYMD 1.4	•Matteo M. Defranchis, Katerina Lipka, Sven-Olaf Moch Towards beam-quality-preserving plasma accelerators: On the precision tuning of the wakefield •Sarah Schröder	
14:00	Audimax	SYMD 1	Session Symposium SMuK Dissertation Prize 2022	
				EP
			Invited Talks	
14:00	EP-H1	EP 1.1	Asteroseismology of red-giant stars •Saskia Hekker	
16:45	EP-H1	EP 2.3	Cloudy with a hint of magnetic fields •Ludmila Carone	
			Sessions	
14:00 16:15	EP-H1 EP-H1	EP 1 EP 2	Astrophysics Astrophysics / Exoplanets and Astrobiology	
				GR
			Sessions	
09:00	GR-H2	GR 1	Black Holes	
16:00 16:15	GR-H2 GR-H3	GR 2 GR 3	Classical Theory Cosmology	
				MP
00.30	MP-H5	MP 1 1	Invited Talks	
09.50	IVII 115		•Pieter Naaijkens, Yoshiko Ogata	
10:00	MP-H5	MP 1.2	Symmetry-resolved quantum information measures for AdS gravity and beyond	
10:30	MP-H5	MP 1.3	 René Meyer, Suting Zhao, Christian Northe, Konstantin Weisenberger Rapid thermalization of spin chain commuting Hamiltonians Angela Capel 	

Monday, March 21, 2022

			MP
16:15	MP-H5	MP 2.1	The Page curve from quantum error correction
16:45	MP-H5	MP 2.2	•Daniel Harlow Classical black hole scattering from a worldline quantum field theory •Jan Plefka
09:30	MP-H5	MP 1	Sessions Entanglement: Thermalization
16:15	MP-H5	MP 2	Gravity, Amplitudes, AdS/CFT
			ST
			Sessions
14:00	ST-H4	ST 1	Radiation therapy I
16:15	Р	ST 2	Poster Session
			Т
			Invited Talks
09:30	T-H15	T 1.1	From scattering amplitudes to precision predictions for the LHC •Claude Duhr
10:00	T-H15	T 1.2	Tackling new physics at the fringe of precision: Standard Model physics at
			•Simone Amoroso
10:30	T-H15	T 1.3	Hunt for New Physics at the LHC
00.30	Т-Н15	Т 1	Sessions Invited Talks 1
16:15	T-H15	T 2	OCD (Theorie) 1
16:15	T-H16	T 3	Flavour Physics 1
16:15	T-H17	Τ4	Flavour Physics 2
16:15	T-H18	Т 5	Electroweak Interactions (Exp.) 1
16:15	T-H19	T 6	Top Quarks: Production (Exp.) 1
16:15	I-H20		Top Quarks: Properties 1
16:15	1-H21 T L 22		Higgs Boson: Decay in Fermions 1
16.15	1-H22	T 10	Search for New Particles -1
16.15	T-H24	T 10	Gaseous Detectors
16:15	T-H25	T 12	Pixel Detectors
16:15	T-H26	T 13	Semiconductor Detectors:Radiation Hardness, new Materials and Concepts
16:15	T-H27	T 14	DAQ and Trigger 1
16:15	T-H28	T 15	GRID Computing
16:15	T-H29	T 16	Experimental Methods (general) 1
16:15	I-H30	11/ T10	Gamma Astronomy 1
10:15	I-H3 I エロ22	1 18 T 10	Neutrino Astronomy I
16.15	T-H32	T 20	Neutrino Physics without Accelerators 1
16:15	T-H34	T 21	Neutrino Physics without Accelerators 2
16:15	T-H35	T 22	Search for Dark Matter 1
16:15	T-H36	T 23	Experimental Techniques in Astroparticle Physics 1
16:15	T-H37	T 24	Outreach Methods
16:15	T-H38	T 25	Data Analysis, Information Technology and Artificial Intelligence
16:15	T-H39	T 26	Data Analysis, Information Technology and Artificial Intelligence

Monday, March 21, 2022

			DD
09:00	DD-H8	DD 1.1	Invited Talk Von Naturphänomenen und solchen jenseits unserer Wahrnehmung •Michael Vollmer
09:00 10:15 10:15 10:15 10:15 11:30 11:30 11:30 11:30 11:30 11:30 11:30 11:30 11:30 11:30 11:30 15:30 15:30 15:30	DD-H8 DD-H9 DD-H10 DD-H11 DD-H12 P P P P P P P P P P P P P DD-H7 DD-H8 DD-H9 DD-H10 DD-H11 DD-H12	DD 1 DD 2 DD 3 DD 4 DD 5 DD 6 DD 7 DD 8 DD 9 DD 10 DD 11 DD 12 DD 13 DD 14 DD 15 DD 16 DD 17 DD 18 DD 17 DD 18 DD 19 DD 19 DD 19 DD 20	Sessions Hauptvortrag 1 Physikdidaktik und Inklusion Neue / digitale Medien - Konzeption Lehr-Lernforschung - Schülervorstellungen Science Hochschuldidaktik - Studieneingangsphase Fachwissen Quantenphysik - Experimente Postersession 1: Anregungen aus dem Unterricht für den Unterricht Postersession 1: Bildung für nachhaltige Entwicklung Postersession 1: Geschichte der Physik / Nature of Science Postersession 1: Lehr- und Lernforschung Postersession 1: Lehreraus- und -weiterbildung Postersession 1: Neue Konzepte Postersession 1: Neue Konzepte Postersession 1: Sprache und Physikunterricht Communicating Physics and its History Neue / digitale Medien - Schule Neue / digitale Medien - AR Lehr-Lernforschung - Schülervorstellungen fachbezogen neue Konzepte - Physikunterricht außerschulisches Lernen - Konzepte
			GP
13:30 14:00 14:30	GP-H7 GP-H7 GP-H7	GP 1.1 GP 1.2 GP 1.3	Invited Talks The development of The Lorentz Lab: bringing the scientific history of Tey- lers Museum to life with working replicas •Trienke van der Spek Physics in Information Comics •Heike Elisabeth Juengst Jenseits gewohnter Pfade: Ausstellungen neu denken •Christian Sichau
13:30 15:30 17:00 18:00	GP-H7 GP-H7 GP-H7 GP-MV	GP 1 GP 2 GP 3 GP 4	Sessions Communicating Physics and its History History and Teaching Physics and Media Meeting of Early Career Scholars
			ΑΚΡΙΚ
16:15	AKPIK-H13	AKPIK 1	Session Data Integration & Processing
			AGPhil
16:15	AGPhil-H14	AGPhil 2.1	Invited Talk To G or not to G: J. H. Poynting and the gravitational constant in the 19th century •Isobel Falconer
11:00 16:15	AGPhil-H14 AGPhil-H14	AGPhil 1 AGPhil 2	Sessions Symmetry and Geometry History and Philosophy of Gravity

			Plenary Talks
09:00	Audimax	PV II	The LHC legacy and prospects
09:45	Audimax	PV III	•Markus Klute Particle theory in a data-driven era •Michael Krämer
			SYNS
14:00	Audimax	SYNS 1.1	The Role of Nature of Science Education for Science Media Literacy •Dietmar Höttecke
14:30	Audimax	SYNS 1.2	What kinds of identities are deemed in/our of place in physics?
15:00	Audimax	SYNS 1.3	Some thoughts on the status of theoretical physics •Daniel Harlow
			Session
14:00	Audimax	SYNS 1	The Nature of Science
			EP
			Invited Talks
11:00	EP-H1	EP 3.1	A new era of Venus exploration - seen Venus in a new light •Jörn Helbert, Melinda Darby Dyar, Giulia Alemanno, Alessandro Maturilli, Nils Müller, Doris Breuer, VEM on VERTIAS team, VenSpec on Envision team
14:00	EP-H1	EP 4.1	Exploration of the Jupiter system with a small submillimetre wave telescope onboard the JUICE satellite
16:15	EP-H1	EP 5.1	•Paul Hartogn Investigating Earth's atmosphere and ionosphere from space: How GNSS radio occultation measurements contribute to monitor the atmosphere in a high spatial resolution
17:30	EP-H1	EP 5.5	•Christina Arras, Ankur Kepkar, Jens Wickert Solar extreme events and their impact on the middle atmosphere •Thomas Reddmann, Monali Borthakur, Miriam Sinnhuber, Ilya Usoskin, Jan- Maik Wissing
			Sessions
11:00	EP-H1	EP 3	Planets and Small bodies
14.00 16:15	EP-H1 EP-H1	EP 4 EP 5	Near Earth Space
			GR
			Invited Talks
11:00	GR-H2	GR 4.1	Pseudospectrum and black hole quasi-normal mode (in)stability
11:45	GR-H2	GR 4.2	 Roungo Panosso Macedo Observable Signatures of Quantized Gravity in Quantum Optical Experiments Dennis Rätzel
			Sessions
11:00	GR-H2	GR 4	General Relativity
16:15 16:15	GR-H2 GR-H3	GR 6	Cosmology
-	-		

			MP
			Invited Talks
11:00	MP-H5	MP 3.1	Dualities and categorical symmetries in quantum spin chains •Frank Verstraete
11:30	MP-H5	MP 3.2	Functional Integral and Stochastic Representations for Bosonic Ensembles •Manfred Salmhofer
12:00	MP-H5	MP 3.3	Color-Flavor Transformation Revisited •Martin Zirnbauer
			Sessions
11:00 16:15	MP-H5 MP-H5	MP 3 MP 4	Quantum dynamics Quantum field theory, AdS/CFT, non-equilibrium and quantum dynamics
			ST
			Invited Talk
14:00	ST-H4	ST 4.1	Present status and future challenges of magnetic resonance-guided radio- therapy •Christian P. Karger
			Sessions
11:00	ST-H4	ST 3	Artificial Intelligence in Medicine
14:00 16:15	ST-H4 ST-MV	ST 4 ST 5	Radiation monitoring and dosimetry I Annual General Meeting
			Τ
11.00	T1146	T 07 1	Invited Talks, Invited Topical Talks
11:00	I-H I 5	127.1	•Thomas Kuhr
11:30	T-H15	T 27.2	Flavour Anomalies
			•Christoph Langenbruch
12:00	I-H15	127.3	l he top quark is still going strong (and electroweak) •Andrea Knue
14:00	T-H15	T 28.1	Hadronic Jets: Accuracy and Precision of their Reconstruction and Calibrati-
			on in ATLAS
14.25	T-H15	T 28 2	•Christopher Young Direct searches testing BSM explanations of the flavor anomalies
14.20	11110	1 20.2	•Arne Christoph Reimers
14:50	T-H15	T 28.3	ATLAS probes QCD measuring photons
15:15	T-H15	T 28.4	•Heperth Torres The upgrade of the ATLAS trigger to augment the physics reach of Run-3
	-	-	•Daniele Zanzi
14:00	T-H16	T 29.1	Testing the Standard Model through Gauge-boson Self-interactions
14:25	T-H16	T 29.2	Axions and similar particles – how to cover 10 ¹⁷ orders of magnitude in
			mass
14.50	Т-Н16	Т 20 3	•Kristof Schmieden From GERDA to LEGEND – Hunting no neutrinos
14.50	11110	129.5	•Christoph Wiesinger
15:15	T-H16	T 29.4	Mapping Highly-Energetic Messengers throughout the Universe •Sara Buson
			Sessions
11:00	T-H15	T 27	Invited Talks 2
14:00	T-H15	T 28	Invited Topical Talks 1
14.00 16:15	T-H15	T 30	Flavour Physics 3

				Т
16:15	T-H16	T 31	Beyond the Standard Model (Theory) 1	
16:15	T-H17	T 32	QCD (Exp.) 1	
16:15	I-H18	1 33	Top Quarks: Production (Exp.) 2	
16:15	1-H19	134 T25	Top Quarks: Properties -2	
10.15	I-HZU T_U21	130	Higgs Boson: Associated Production 1 Higgs Boson: Extended Medels 1	
16.15	T-H27	T 30 T 37	Search for New Particles 2	
16:15	T-H23	T 38	Search for New Particles 3	
16:15	T-H24	T 39	Gaseous Detectors 2	
16:15	T-H25	T 40	Pixel Detectors 2	
16:15	T-H26	T 41	Calorimeters 1	
16:15	T-H27	T 42	Detector Systems 1	
16:15	T-H28	T 43	DAQ and Trigger 2	
16:15	T-H29	T 44	Experimental Methods (general) 2	
16:15	T-H30	T 45	Gamma Astronomy 2	
16:15	I-H31	I 46	Neutrino Astronomy 2	
10:15	I-H3Z	14/ T40	Cosmic Ray 2 Neutrine Dhysice without Appelarators 2	
10.15	1-1133 T-1137	1 48 T 40	Neutrino Physics without Accelerators 3	
16.15	T-H35	T 50	Search for Dark Matter 2	
16:15	T-H36	T 51	Experimental Techniques in Astroparticle Physics 2	
16:15	T-H37	T 52	Outreach Methods 2	
16:15	T-H38	T 53	Data Analysis, Information Technology and Artificial Intelligence 3	
			F	
				עי
			Invited Talk	
11:00	DD-H8	DD 21.1	Warum Lehrbuchdarstellungen der Physikgeschichte so schlecht sind – ur	۱d
			was wir daraus lernen können	
			•Oliver Passon	
			Sessions	
11:00	DD-H8	DD 21	Hauptvortrag 2	
12:00	DD-H8	DD 22	Lehreraus- und Lehrerfortbildung – neue Ansätze	
12:00	DD-H9	DD 23	Neue / digitale Medien -VR	
12:00	DD-H10	DD 24	Physikdidaktik und Inklusion – Experimentieren	
12:00		DD 25	Hochschuldidaktik – Studieneingangsphase Studierendenperspektive	
12:00		DD 26	Quantenphysik – Konzepte	
16.15	00-00 00-00	DD 27 DD 28	Geschichte der Physik und NoS	
16.15	DD-H10	DD 20 DD 29	Lehr-Lernforschung – Lernermerkmale	
16:15	DD-H11	DD 30	Lehreraus- und Lehrerfortbildung – neue Ansätze	
16:15	DD-H12	DD 31	Praktika und neue Praktikumsversuche	
17:00	Р	DD 32	Postersession 2: Astronomie	
17:00	Р	DD 33	Postersession 2: Außerschulisches Lernen	
17:00	Р	DD 34	Postersession 2: Hochschuldidaktik	
17:00	Р	DD 35	Postersession 2: Neue / digitale Medien	
17:00	Р	DD 36	Postersession 2: Praktika und neue Praktikumsversuche	
17:00	Р	DD 37	Postersession 2: Präsentation von Experimenten	
17:00	Р	DD 38	Postersession 2: Quantenphysik	
18:00	P DD-MV	DD 39 DD 40	Mitgliederversammlung des Fachverbands Didaktik der Physik	
				ìΡ

Sessions

Physics and the Museum

16:15	GP-H7	GP 6	GP Physicists as Popularisators
17:45	GP-MV	GP 7	Annual General Meeting of the History of Physics Division
			AGPhil
			Invited Talks
14:00	AGPhil-H14	AGPhil 4.1	Hypothetical Waveforms and Unmodeled or Pipeline Searches in Gravitatio- nal Wave Astronomy •Lvdia Patton
16:15	AGPhil-H14	AGPhil 5.1	Portrait of a Black Hole: Objectivity and the Imaging of M87* by the Event Horizon Telescope •Peter Galison
17:00	AGPhil-H14	AGPhil 5.2	When is a black hole spacetime "as large as it can be?" •Juliusz Doboszewski
			Sessions
11:30 14:00 16:15	AGPhil-H14 AGPhil-H14 AGPhil-H14	AGPhil 3 AGPhil 4 AGPhil 5	Black Holes I Gravitational and Electromagnetic Waves Black Holes II

Wednesday, March 23, 2022

09:00	Audimax	PV IV	Plenary Talks Teaching with Objects and Teaching with Video: The Challenges of Informa Education in Physics	al
09:45	Audimax	PV V	Quantum gravity, chaos, statistical physics and wormholes •Jan de Boer	
			SYA	W
14:00	Audimax		Award Presentation of the Dissertation Prize 2022 of the SMuK Section	
14:10	Audimax	SYAW 1.1	Prize Talks Wie überprüft man die Ziele der Lehramtsausbildung Physik? •Horst Schecker (Laureate of the Georg-Kerschensteiner-Prize 2022)	
14:40	Audimax	SYAW 1.2	Astronomy at Highest Angular Resolution – Adaptive Optics, Interferometr and Black Holes	y
15:10	Audimax	SYAW 1.3	•Alexander M. Polyakov (Laureate of the Max Planck Medal 2022)	
14:00	Audimax	SYAW 1	Session Awards Symposium	
			E	ΞP
11:00	EP-H1	EP 6.1	Invited Talk Three-dimensional topology-driven magnetic reconnection •Raquel Mäusle, Jean-Mathieu Teissier, Wolf-Christian Müller	
11:00 16:15	EP-H1 EP-H1	EP 6 EP 7	Sessions Astrophysics Astrophysics	
				R
11:00	GR-H2	GR 7.1	Invited Talk Numerical Relativity and Gravitational Wave Observations •Harald Pfeiffer	
11:00 11:45 16:15 16:15	GR-H2 GR-H2 GR-H2 GR-H3	GR 7 GR 8 GR 9 GR 10	Sessions General Relativity Gravitational Waves Gravitational Waves Foundations and Alternatives	
			N	IP
11:00 16:15	MP-H5 MP-H5	MP 5.1 MP 6.1	Invited Talks Falling through masses in superposition: quantum reference frames for inc finite metrics Anne-Catherine de la Hamette, Viktoria Kabel, Esteban Castro-Ruiz, •Caslav Brukne State sum models with defects	de- er
			•Catherine Meusburger	
11:00 16:15	MP-H5 MP-H5	MP 5 MP 6	Sessions Quantum information Mathematical physics	

Wednesday, March 23, 2022

			MP
16:45	MP-H5	MP 7	Quantum field theory
17:25	MP-H6	MP 8	Alternative approaches
			ST
			Sessions
11:00	ST-H4	ST 6	Detectors and Applications I
16:15	ST-H4	ST 7	Radiation monitoring and dosimetry II
			т
			Invited Topical Talks
11:00	T-H15	T 54.1	Hunting XYZ Beasts at Belle and Belle II •Elisabetta Prencine
11:25	T-H15	T 54.2	Precision tests of the Standard Model using CP violation in B meson decays
11:50	T-H15	T 54.3	Back to the top: charting the bounds of the standard model
12:15	T-H15	T 54.4	Dark matter from spin-2 mediators
11:00	T-H16	T 55.1	Machine Learning for LHC Theory
11:25	T-H16	T 55.2	Towards high-precision deep learning for astroparticle physics
11:50	T-H16	T 55.3	The quest for the mechanism behind the matter-antimatter asymmetry
12:15	T-H16	T 55.4	Towards the lightest dark matter in direct searches •Belina von Krosigk
			Sessions
11.00	T-H15	T 54	Invited Topical Talks 3
11:00	T-H16	T 55	Invited Topical Talks 4
16:15	T-H15	T 56	Flavour Physics 4
16:15	T-H16	T 57	Flavour Physics 5
16:15	T-H17	T 58	QCD (Exp.) 2
16:15	T-H18	T 59	Neutrino Physics with Accelerators 1
16:15	T-H19	T 60	Top Quarks: Decay and CP Violation and Mixing Angles
16:15	T-H20	T 61	Higgs Boson: Decay in Bosons
16:15	T-H21	T 62	Higgs Boson: Extended Models 2
16:15	T-H22	T 63	Search for New Particles 4
16:15	T-H23	T 64	Search for New Particles 5
16:15	T-H24	T 65	Silicon Strip Detectors
16:15	I-H25	1 66 T 67	Semiconductor Detectors: Radiation Hardness, new Materials and Concepts 2
16:15	I-H26	16/	Myon Detectors
10.15			Detector Systems 2
10.15		T 09 T 70	DAQ dilu Myyel 5 Experimental Methode (general) 2
16.15	T-H30	T 71	Neutrino Astronomy 3
16.15	T-H31	T 72	Cosmic Ray 3
16:15	T-H32	T 73	Cosmic Ray 4
16:15	T-H33	T 74	Neutrino Physics without Accelerators 5
16:15	T-H34	Т 75	Neutrino Physics without Accelerators 6
16:15	T-H35	T 76	Search for Dark Matter 3
16:15	T-H36	Т 77	Search for Dark Matter 4
16:15	T-H37	T 78	Experimental Techniques in Astroparticle Physics 3
16:15	T-H38	T 79	Data Analysis, Information Technology and Artificial Intelligence 4

Wednesday, March 23, 2022

			Sessions
10:45	DD-H8	DD 41	Neue / digitale Medien – Experimente
10:45	DD-H9	DD 42	Lehr- und Lernforschung – Repräsentatonsformen
10:45	DD-H10	DD 43	BNE – Lernendenperspektive
10:45	DD-H11	DD 44	Hochschuldidakitk – neue Konzepte
10:45	DD-H12	DD 45	Astronomie
12:00	DD-H8	DD 46	BNE – Konzepte
12:00	DD-H9	DD 47	Lehreraus- und -weiterbildung – Lehrkonzepte
12:00	DD-H10	DD 48	Lehr- und Lernforschung – Methodik
12:00	DD-H11	DD 49	Lehreraus- und -weiterbildung – digitale Medien
12:00	DD-H12	DD 50	Praktika und neue Praktikumsversuche
16:00	DD-H8	DD 51	Workshop: Konsequenzen aus drei Jahren Studienreformforschung
			GP
			Sessions
10.20			Develop and Instrumente
10.30		GP 0	Physics and instruments
10.40		GP 9 CD 10	History of Physics Develop and Culture
18:00	GP-H/	GP IU	Physics and Culture
			AKPIK
			Sessions
16:15	AKPIK-H13	AKPIK 2	Data Analytics & Machine Learning
19:00	AKPIK-MV	AKPIK 3	Annual General Meeting of the Working Group on Physics, Modern IT and Artificial Intelligence
			AGPhil
			Invited Talks
14.00		ACPhil 6 1	Spacetime Conventionalism Revised: Tidal Forces and Weyl Curvature
14.00	AGEIII-III4	AGFIII 0.1	•Karim Théhault Hfuk Tasdan
16.15		AGPhil 7 1	On an inferential role of spacetime in particle physics
10.10	Adrimitit	Adrini 7.1	•Tushar Menon
			Sessions
14:00	AGPhil-H14	AGPhil 6	Foundations of Gravity
16:15	AGPhil-H14	AGPhil 7	Symmetries and Principles
18:30	AGPhil-MV	AGPhil 8	Annual General Meeting of the Working Group on Philosophy of Physics
			Public Evening Lecture (free entrance)
20:00	PEL	PV VI	Kleine Teilchen – große Chancen: Mit Quantenkorrekturen auf der Suche
			nach neuen Physikphänomen
			Stephanie Hansmann-Menzemer

Thursday, March 24, 2022

			Plenary Talks
09:00	Audimax	PV VII	New perspectives onto the Universe in the era of multi-messenger astrophysics •Samaya Nissanke
09:45	Audimax	PV VIII	The Sun as a source of high-energy particles •Rami Vainio
			EP
			Sessions
11:00	T-H15	EP 8	Astroparticles: Invited talks
12:45	EP-MV	EP 9	Annual General Meeting of the Extraterrestrial Physics Division
14:00	T-H15	EP 10	Astroparticles: Invited topical talks
16:15	EP-H1	EP 11	Astroparticles: From the source to the detector
			GR
			Sessions
11:00	GR-H2	GR 11	Relativistic Astrophysics
11:00	MP-H5	GR 12	Quantum gravity
14:00	GR-H2	GR 13 CP 14	Numerical Relativity Cravitational Wave Detectors
10.15	GR-MV	GR 14 GR 15	Annual General Meeting of the History of Physics Division
	GRIM		
			MP
			Invited Talks
11:00	MP-H5	MP 9.1	Reduced phase space quantisation in Loop quantum gravity and loop quan- tum cosmology •Kristina Giesel, Bao-Fei Li, Parampreet Singh
11:30	MP-H5	MP 9.2	Quantum fields propagating on curved backgrounds and their influence on spacetime curvature •Nicola Pinamonti
			Sessions
11:00	MP-H5	MP 9	Quantum gravity
16:15	MP-H5	MP 10	Classical and quantum gravity
16:35	MP-H6	MP 11	Thermodynamics and fundamental aspects of field theory
17:15 19:30	MP-H5 MP-MV	MP 12 MP 13	Annual General Meeting of the Theoretical and Mathematical Physics Division
			ST
			Invited Talk
16:15	ST-H4	ST 10.1	Artificial intelligence in PET image reconstruction and quantitative analysis •Zhaoheng Xie
11.00	07.114	o T 0	Sessions
11:00	SI-H4 ST-H4	518 979	Radiation therapy II
14.00	ST-H4	ST 10	Total-Body PFT
16:55	ST-H4	ST 11	Prize Ceremony and Closing Session
			Т
			Invited Talks, Invited Topical Talks
11:00	T-H15	T 80.1	Borexino looks in the direction of solar neutrinos •Livia Ludhova

Thursday, March 24, 2022

			-
11:30	T-H15	T 80.2	Gravitational waves – a new probe of the early Universe
12:00	T-H15	T 80.3	 Valerie Domcke Gravitational wave detectors – current and future challenges
14.00	T-H15	T 81 1	•Michèle Heurs
14.00	11110	101.1	the Moon
			•Sonke Burmeister, Shenyi Zhang, Jia Yu, Zigong Xu, Stephan Bottcher, Ro- bert Wimmer-Schweingruber
14:25	T-H15	T 81.2	Energetic Pulsar Environments and the Origins of Galactic Cosmic Rays •Alison Mitchell
14:50	T-H15	T 81.3	Looking forward to exciting physics with FASER
15:15	T-H15	T 81.4	Astroparticle physics at the LHC: Exploring the forward region with cross-section measurements
14:00	T-H16	T 82.1	Searches for new scalar particles at the LHC
14:25	T-H16	T 82.2	Novel approaches to search for new physics in rare charm decays
14:50	T-H16	T 82.3	Constraining the Higgs-charm Yukawa coupling with the CMS experiment
15:15	T-H16	T 82.4	Characterization of \$H\$ boson events in the \$\tau\tau\$ decay channel with
			•Sebastian Wozniewski
			Sessions
11:00	T-H15	T 80	Invited Talks 3
14:00	T-H15	T 81	Invited Topical Talks 5
14.00	T-H16	T 82	Invited Topical Talks 6
14.00		T 92	Astroteilchen: Von der Quelle zum Detektor (contributed talks)
16.15		T 03	Elevent Diverse
10.15		1 04	Flavour Physics
16:15	I-H16	185	Beyond the Standard Model (Theory) 2 and QFT and Lattice Gauge Theory T
16:15	I-H1/	186	Neutrino Physics (Theory) 1 and Theoretical Astroparticle Physics and Cos-
			mology 1
16:15	T-H18	T 87	Electroweak Interactions (Exp.) 2
16:15	T-H19	T 88	Top Quarks: Production (Exp.) 3
16:15	T-H20	T 89	Higgs Boson: Decay in Fermions 2
16:15	T-H21	T 90	Higgs Boson: Associated Production 2
16:15	T-H22	T 91	Higgs Boson: Rare Decays
16:15	T-H23	T 92	Higgs Boson: Extended Models 3
16.15	T-H24	T 93	Search for New Particles 6
16.15	T-H25	T 94	Silicon Strip Detectors 2
16.15	T-H26	T 95	Pixel Detectors 3
16.15	TH20	T 06	Detectors S
10.15		T 90	Electronice 1
10.15		T 97	Electionics i Even evine entre Martha da (non eval) 4
10:15	I-H29	198	Experimental Methods (general) 4
16:15	I-H30	199	Neutrino Astronomy 4
16:15	I-H32	1 100	Cosmic Ray 5
16:15	I-H33	I 101	Cosmic Ray 6
16:15	T-H34	T 102	Neutrino Physics without Accelerators 7
16:15	T-H35	T 103	Neutrino Physics without Accelerators 8
16:15	T-H36	T 104	Search for Dark Matter 5
16:15	T-H37	T 105	Search for Dark Matter 6
16:15	T-H38	T 106	Experimental Techniques in Astroparticle Physics 4
16:15	T-H39	T 107	Data Analysis, Information Technology and Artificial Intelligence 5
19:30	T-MV	T 108	Annual General Meeting of the Particle Physics Division (for DPG members)

Thursday, March 24, 2022

			AKPIK
		Session	
16:15	AKPIK-H13 AKPIK 4	Deep Learning	
			AGPhil
		Sessions	
11:00	AGPhil-H14 AGPhil 9	Quantum Mechanics I	
14:00	AGPhil-H14 AGPhil 10	Quantum Mechanics II	
16:15	AGPhil-H14 AGPhil 11	Time and Temperature	

Friday, March 25, 2022

09:00	Audimax	PV IX	Plenary Talks Shedding light on the axion: one particle solving two problems
09:45	Audimax	PV X	Quantum Computing: a future path for High Energy Physics •Karl Jansen
			EP
			Invited Talks
11:00	EP-H1	EP 12.1	Linking Solar Eruptions and Energetic Particles through Observations and Modeling
12.00	FD-H1	FD 12 /	•Frederic Ettenberger Solar Orbiter: two years of operations and first results
12.00		LI 12.4	•Frederic Schuller, Alexander Warmuth, Gottfried Mann
			Session
11:00	EP-H1	EP 12	Sun and Heliosphere
			Т
			Invited Talks
11:00	T-H15	T 109.1	Ten years of Higgs boson measurements: what we know and what we don't know
11:30	T-H15	T 109.2	Future of Silicon Tracking Detectors: LHC Upgrades and Beyond Georg Steinbrück
12:00	T-H15	T 109.3	The dawn of high energy neutrino astronomy •Elisa Resconi
			Session
11:00	T-H15	T 109	Invited Talks 4
			AGPhil

11:00 AGPhil-H14 AGPhil 12

Session

Plenary and Evening Talks

Plenary Talk

PV I Mon 11:40 Audimax Making uncertainty certain?: Proton therapy, range and in-vivo verification. — •TONY LOMAX — WPTA/140 Forschungsstrasse 111 5232 Villigen PSI Schweiz

Proton therapy (RT) is an increasingly important weapon against cancer. By controlling proton energy, and thus proton range, the dose can be concentrated in tumour more effectively than with conventional radiotherapy with photons. However, although proton ranges in water can be determined with sub-millimeter precision, the story in the patient can be quite different due to uncertainties in CT data of the patient, positioning inaccuracies and anatomical changes of the patient during the treatment course. As such, in vivo range uncertainty is one of the major challenges for proton therapy. Consequently, methods such as proton CT/probes, PET activation and prompt gamma imaging are being proposed and developed for direct and in vivo imaging of proton range. This presentation will review and compare each of these, both from their technical practicality and potential clinical impact.

Plenary Talk

PV II Tue 9:00 Audimax The LHC legacy and prospects — •MARKUS KLUTE — KIT

The Large Hadron Collider (LHC) at CERN has had two successful and highly productive runs (2009-2013 and 2015-2018), colliding protons and heavy ions with center-of-mass energies of up to 13 TeV and collecting an unprecedented amount of data. Its highlight, the Higgs Boson discovery in 2012, completed the Standard Model of fundamental particle interactions. While the impact of the collected data has been tremendous, many open questions in the world of elementary particle physics remain. At the dawn of the LHC Run 3 (2022-2025), the experiments have an extensive program exploring the uncharted territory at the energy frontier. I will review the main conclusions from the LHC to date and present the prospects of the LHC program in Run 3 and beyond.

Plenary Talk

PV III Tue 9:45 Audimax Particle theory in a data-driven era — • MICHAEL KRÄMER — RWTH Aachen University

Theory has played a crucial role in developing the Standard Model (SM) and in formulating the questions that point to physics beyond the SM, such as the origin of electroweak symmetry breaking or the particle nature of dark matter. In the first phase of the LHC many searches for physics beyond the SM have been driven by specific new physics models that would address some of the shortcomings of the SM. However, since no conclusive evidence for new physics has been established so far, more model-independent approaches such as effective field theories, or the search for anomalies by machine learning methods, are being pursued. I will try to clarify the role of theory for the discovery of physics beyond the SM in this data-driven era of LHC physics, and highlight the challenges and opportunities for the next decade.

Plenary Talk

PV IV Wed 9:00 Audimax Teaching with Objects and Teaching with Video: The Challenges of Informal Education in Physics — • ALLISON MARSH — University of South Carolina

Books and lectures are popular ways to impart information, but what about objects and videos? In this talk I will focus on these two approaches to teaching.

The first will focus on material culture - the objects and artifacts that reflect human progress. I will draw on my experience over the last four years writing a monthly column for the magazine Spectrum, the flagship publication of the IEEE. Each article focus on a museum object from the history of computer science or electrical engineering, broadly defined.

The second part of the talk will draw from my experience as the consultant for Crash Course History of Science, an extremely popular series of 46 short (10-12 minute), fast paced videos with goofy animation. I will talk about the collaborative process of deciding what information makes the cut and how that affects the overall content.

I will end with a reflection on audience and how these formats can draw in new people to the subjects, but also the drawbacks of the formats.

PV V Wed 9:45 Audimax **Plenary Talk** Quantum gravity, chaos, statistical physics and wormholes - •JAN DE BOER

 University of Amsterdam The information paradox for black holes has been an important source of inspiration for research in quantum gravity, as the paradox implies that low-energy effective field theories that include gravity must break down in an interesting and unexpected way. Recent work has managed to clarify the physics behind the paradox considerably and uncovered fascinating relations between gravity and wormholes on the one hand and chaos, quantum information and statistical physics on the other hand. One lesson seems to be that low-energy observers in gravitational theories do in fact have some access to high-energy informa-

tion, but only to statistical and not to microscopic information. In this picture wormhole solutions are a direct manifestation of this statistical high-energy information.

In this talk I will try to give an informal summary of some of these developments, trying to emphasize the general lessons rather than the technical details.

PV VI Wed 20:00 PEL **Evening Talk** Kleine Teilchen - große Chancen: Mit Quantenkorrekturen auf der Suche nach neuen Physikphänomen — • STEPHANIE HANSMANN-MENZEMER — Universität Heidelberg

Die Frage was die Welt im Innersten zusammenhält treibt Teilchenphysiker*innen weltweit an. Was sind die kleinsten Bausteine des Universums? Wie wechselwirken sie miteinander? Woraus besteht dunkle Materie und warum gibt es in unserem Universum mehr Materie als Antimaterie?

Am Large Hadronen Collider an der internationalen Forschungseinrichtung CERN in der Nähe von Genf werden Protonen bei höchsten Energien 40.000.000 mal pro Sekunde aufeinandergeschossen. Die Zerfallsprodukte der Kollisionen werden in riesigen Teilchendetektoren vermessen, die größer sind als ein Mehrfamilienhaus. Wie Detektive analysieren die Wissenschaftler*innen die Zerfallsprodukte um so zu rekonstruieren was genau bei dem Zusammenstoß der Protonen passierte. Im Rahmen der Heisenbergschen Unschärferelation kann für eine kurze Zeit die Energieerhaltung in Quantenkorrekturen verletzt werden und so können auch sehr schwere Teilchen kurzfristig produziert werden. Deshalb ist die Studie von Quantenkorrekturen induzierten Prozessen ein vielversprechendes Werkzeug für die Suche nach neuen Teilchen und neuen Physikphänomenen.

Der Vortrag führt in die Teilchenphysik am Large Hadron Collider ein und stellt ausgewählte aktuelle Ergebnisse zur Suche nach neuen Physikphänomenen in Quantenkorrekturen vor.

PV VII Thu 9:00 Audimax **Plenary Talk** New perspectives onto the Universe in the era of multi-messenger astrophysics - •SAMAYA NISSANKE - GRAPPA, University of Amsterdam, The Netherlands — Nikhef, The Netherlands

Since the revolutionary discovery of gravitational wave (GW) emission from a binary black hole merger in 2015, the exquisite GW detectors LIGO, Virgo and KAGRA have detected more than 90 compact object mergers. Most notably, one of these mergers corresponds to the first binary neutron star merger, dubbed GW170817. This event has been transformative because it was observed in both gravitational and electromagnetic radiation, thus opening up a new era in multimessenger astrophysics. The multi-messenger characterisation of such an event has enabled major advances into diverse fields of modern physics from gravity, high-energy and extragalactic astrophysics, nuclear physics, to cosmology. In this talk, I will discuss work in strong-field gravity astrophysics and how combining observations, theory and experiment is key to make progress in this field. I will present the opportunities and challenges that have emerged in multimessenger astrophysics, and what the future holds in this new era.

Plenary Talk PV VIII Thu 9:45 Audimax The Sun as a source of high-energy particles — \bullet RAMI VAINIO — Department of Physics and Astronomy, University of Turku, Turku, Finland

Particle acceleration at the Sun occurs in Solar Energetic Particle (SEP) events related to solar eruptions: solar flares and coronal mass ejections. Composition, timing and duration as well as other observational properties of SEP events can be used to assess the mechanisms of particle acceleration at play in these events. However, one particular complication in these analyses has been the relatively large distance of our observers to the source, both in radial and angular distance from the core of the solar eruption. Transport effects can distort the observed particle event substantially and signatures of the acceleration process may be washed out. In the last years, however, the heliospheric research community has obtained new tools to assess a large number of problems with the launch of three new space missions to the inner heliosphere in orbits around the Sun: NASA's Parker Solar Probe and ESA's Solar Orbiter and BepiColombo. In combination with older probes at Earth's distance from the Sun, they offer an unprecedented possibility to disentangle transport effects from the imprints of particle acceleration in the data.

I will provide an overview of observations and models of SEP events and discuss what has been learnt of particle acceleration at the Sun based on the comparison of models and observations. I will conclude with some of the open questions that we expect to solve over the next solar maximum with the new heliospheric fleet.

Plenary Talk

PV IX Fri 9:00 Audimax

Shedding light on the axion: one particle solving two problems — •FRIEDERIKE JANUSCHEK — Deutsches Elektronen-Synchrotron DESY, Notkestr. 85, 22607 Hamburg, Germany

Very lightweight and feebly interacting new bosons like the axion, originally introduced to solve the strong CP problem of QCD, or axion-like particles (ALPs) predicted by theories beyond the standard model, have been found to also be viable cold dark matter candidates. Thus, the existence of axions and axion-like particles has a particular appeal theoretically –and is additionally supported by experimental hints.

Worldwide, experiments have been and continue to employ creative methods to look for these –previously considered invisible– particles and promise to shed further light onto them. At DESY in Hamburg, the ALPS II experiment, a lightshining-through-walls experiment in the laboratory, is close to data-taking. The planned (Baby)IAXO and MADMAX experiments are complementary experiments that will search for axions and ALPs from the sun and dark matter halo, respectively.

In this talk, I will give an introduction to axions and discuss strategies for detecting them, focusing on activities at DESY.

Plenary Talk

PV X Fri 9:45 Audimax

Quantum Computing: a future path for High Energy Physics — •KARL JANSEN — DESY, Platanenallee 6, 15738 Zeuthen

Quantum computers offer the fascinating possibility to outperform classical computers or to even solve problems that cannot be addressed by classical computations. We will describe particular examples of quantum computing approaches such as the hybrid quantum/classical variational quantum eigensolver and quantum machine learning, including graph neural networks and quantum generative adversarial neural networks. We demonstrate successful applications of these methods at selected problems in theoretical and experimental high energy physics. In addition, we show, how by developing new methods for error mitigation and for analyzing the expressivity of quantum circuits the quantum noise of existing hardware platforms can be mildened.

Symposium SMuK Dissertationspreis 2022 (SYMD)

jointly organised by the divisions of the Matter and Cosmos Section (SMuK)

> Claus Lämmerzahl ZARM, Universität Bremen Am Fallturm 28359 Bremen claus.laemmerzahl@zarm.uni-bremen.de

The Matter and Cosmos Section, with the Extraterrestrial Physics division (EP), the Gravitation and Relativity Division (GR), the Hadronic and Nuclear Physics Division (HK), the Theoretical and Mathematical Physics Division (MP), the Plasma Physics Division (P), the Radiation and Medical Physics Division (ST), and the Particle Physics Division (T), awards a dissertation prize in recognition of outstanding research in the context of a doctoral thesis and its excellent communication. The award committee has selected four candidates from the nominations who will present their doctoral theses at this symposium. The dissertation prize will be awarded in front of the Awards Symposium (SYAW) on Wednesday at 2 pm.

Overview of Invited Talks and Sessions

(Lecture hall Audimax)

Invited Talks

SYMD 1.1	Mon	14:00-14:25	Audimax	Timeless Quantum Mechanics and the Early Universe — •LEONARDO CHATAIGNER
SYMD 1.2	Mon	14:25-14:50	Audimax	First tritium β -decay spectrum recorded with Cyclotron Radiation Emission
				Spectroscopy (CRES) — • CHRISTINE CLAESSENS
SYMD 1.3	Mon	14:50-15:15	Audimax	Watching the top quark mass run - for the first time! — • MATTEO M. DEFRANCHIS,
				Katerina Lipka, Sven-Olaf Moch
SYMD 1.4	Mon	15:15-15:40	Audimax	Towards beam-quality-preserving plasma accelerators: On the precision tuning
				of the wakefield — •SARAH SCHRÖDER

Sessions

SYMD 1.1–1.4 Mon 14:00–15:40 Audimax Symposium SMuK Dissertation Prize 2022

Sessions

Invited Talks -

SYMD 1: Symposium SMuK Dissertation Prize 2022

Time: Monday 14:00-15:40

Invited Talk SYMD 1.1 Mon 14:00 Audimax Timeless Quantum Mechanics and the Early Universe - •LEONARDO

CHATAIGNER — Department of Physics and Astronomy, University of Bologna - Institut für Theoretische Physik, Universität zu Köln

Candidate theories of quantum gravity must answer the questions: how can the dynamics of quantum states of matter and geometry be defined in a diffeomorphism-invariant way? How are the quantum states related to probabilities in the absence of a preferred time? To address these issues, we discuss the construction and interpretation of relational observables in quantum theories with worldline diffeomorphism invariance, which serve as toy models of quantum gravity. In this context, we present a method of construction of quantum relational observables which is analogous to the construction of gaugeinvariant extensions of noninvariant quantities in usual gauge (Yang-Mills) theories. Furthermore, we discuss how the notion of a physical propagator may be used to define a unitary evolution in the quantum theory, which is to be understood in terms of a generalized clock, as is the classical theory. We also discuss under which circumstances this formalism can be related to the use of conditional probabilities in a generalization of the Page-Wootters approach. Finally, we also examine how our formalism can be adapted to calculations of quantumgravitational effects in the early Universe.

Invited Talk SYMD 1.2 Mon 14:25 Audimax First tritium β -decay spectrum recorded with Cyclotron Radiation Emission Spectroscopy (CRES) — •CHRISTINE CLAESSENS — Center for Experimental Nuclear Physics and Astrophysics, University of Washington, WA, USA

The observation of neutrino flavor oscillation proved that neutrinos have mass, thereby requiring us to extend the Standard Model of particle physics. Until now, laboratory experiments could only set upper limits on the electron-weighted neutrino mass $m_{\beta} < 0.8 \,\mathrm{eV/c^2}$. The Project 8 collaboration aims to determine the absolute neutrino mass scale from the distortion of the tritium beta decay spectrum near the endpoint. To this end, the collaboration has successfully established CRES, a frequency-based approach to detect electrons and measure their kinetic energy. In this work, an event detection system consisting of realtime triggering and offline event reconstruction has been developed. Since the neutrino mass is determined from the shape distortion it induces in the tritium spectrum, it is essential to quantify any dependence of the electron detection efficiency on energy or, equivalently, frequency. This work demonstrates the importance of including the detection efficiency in the analysis of the first tritium spectrum recorded with CRES for an accurate endpoint measurement and the extraction of the electron-weighted antineutrino mass. In addition, a requirement for the precision of detection efficiency measurements for a future CRES experiment with a 40 meV/c^2 target sensitivity has been determined.

Invited Talk

Location: Audimax

SYMD 1.3 Mon 14:50 Audimax Watching the top quark mass run - for the first time! — •MATTEO M. DEFRANCHIS¹, KATERINA LIPKA², and SVEN-OLAF MOCH³ — ¹CERN, Geneva, Switzerland — ²DESY, Hamburg, Germany — ³UHH, Hamburg, Germany In the Standard Model of particle physics, the masses of elementary particles are understood as the fundamental couplings to the Higgs field. A special role is played by the mass of the top quark, the most massive elementary particle currently known, whose value affects conclusions about the stability of the vacuum state of our universe. The interaction between quarks and gluons is described by a sector of the Standard Model known as Quantum Chromodynamics, with the strength of the interaction depending on a quantity called strong coupling constant. According to Quantum Chromodynamics, the strong coupling constant rapidly decreases at higher energy scales. The effect is known as the "running of the coupling constant". The same is also true for the masses of the quarks, and the experimental verification of this effect is an essential test of the validity of the Standard Model. Furthermore, the presence of physics beyond the Standard Model can lead to modifications of the mass running by means of the effect of virtual particles. In this work, the running of the top quark mass is measured using high-energy proton-proton collision data collected by the CMS experiment at the CERN Large Hadron Collider. In this way, the fundamental quantum effect of the mass running is investigated for the first time for the most massive elementary particle known.

Invited Talk SYMD 1.4 Mon 15:15 Audimax Towards beam-quality-preserving plasma accelerators: On the precision tuning of the wakefield - •SARAH SCHRÖDER - Deutsches Elektronen-Synchrotron (DESY), Hamburg

Plasma wakefields enable record-setting GeV/m-level acceleration gradients, making them a promising avenue for reducing the size and associated costs of future particle accelerators - with potentially revolutionary implications for basic research and a wealth of industrial and medical applications. The control and optimisation of the acceleration process in the plasma is fundamentally linked to the ability to fine-tune the μ m-scale wakefield structure, calling for diagnostics with femtosecond resolution. In this contribution, a novel methodology that allows measurement [1] and tuning [2] of the plasma wakefields at the femtosecond level in a simple way is presented. These novel capabilities allowed the detailed structure of a GV/m-level wakefield acting on the electron bunch throughout the acceleration process to be directly measured for the first time [1] and ultimately fine-tuned such that the energy spread and charge of the bunch were preserved while achieving an energy-transfer efficiency of 42% [3]. These results mark a crucial step towards quality-preserving and efficient high-gradient acceleration - a necessary development to meet the continuous demand for everhigher beam energies in high-energy particle physics.

[1] S. Schröder, et al. Nat. Commun. 11, 5984 (2020)

[2] S. Schröder et al., J. Phys.: Conf. Ser. 1596 012002 (2020)

[3] C.A. Lindstrøm, J.M. Garland, S. Schröder et al., Phys. Rev. Lett. 126, 014801 (2021)

Awards Symposium jointly organised by

jointly organised by the Matter and Cosmos Section (SMuK)

Karl-Henning Rehren Institut für Theoretische Physik Georg-August-Universität Göttingen 37077 Göttingen krehren@gwdg.de

The laureates of the Georg Kerschensteiner Prize 2022, the Stern Gerlach Medal 2022, and the Max Planck Medal 2021 present their work to a broader audience.

Overview of Invited Talks and Sessions

(Lecture hall Audimax)

Prize Talks

SYAW 1.1	Wed	14:10-14:40	Audimax	Wie überprüft man die Ziele der Lehramtsausbildung Physik? — •HORST SCHECKER
SYAW 1.2	Wed	14:40-15:10	Audimax	Astronomy at Highest Angular Resolution - Adaptive Optics, Interferometry and
				Black Holes — •Frank Eisenhauer
SYAW 1.3	Wed	15:10-15:40	Audimax	Turbulence in one dimension — •Alexander M. Polyakov

Sessions

SYAW 1.1-1.3 Wed 14:00-15	:40 Audimax	Awards Symposium
---------------------------	-------------	------------------

Sessions

– Prize Talks –

SYAW 1: Awards Symposium

Time: Wednesday 14:00-15:40

Announcement of the Dissertation Prize 2022

Prize TalkSYAW 1.1Wed 14:10AudimaxWie überprüft man die Ziele der Lehramtsausbildung Physik? — •HORSTSCHECKER — Universität Bremen, FB 1Physik/Elektrotechnik, Didaktik derPhysik — Träger des Georg-Kerschensteiner-Preises 2022

Die fachbezogenen Ziele der Lehramtsausbildung werden von der Kultusministerkonferenz im Fachprofil Physik beschrieben. Die Schwerpunkte der universitären Phase liegen in den Bereichen Fachwissen, Fachdidaktik und fachliche Erkenntnis- und Arbeitsmethoden. Physikdidaktisch sollen die (zukünftigen) Lehrkräfte u.a. solide Kenntnisse der Grundkonzeptionen für die Gestaltung von Physikunterricht und der Forschungsergebnisse über typische physikalische Lernschwierigkeiten entwickeln.

Aber in welchem Maße werden die Ausbildungsziele auch erreicht? Die Physikdidaktik untersucht seit einigen Jahren empirisch die Wirkungen des Lehramtsstudiums. Dazu gehören die Dynamik des Erwerbs fachlichen Wissens, die Fähigkeit zum Erklären physikalischer Zusammenhänge, das Planen von Unterricht oder das Diagnostizieren von Lernschwierigkeiten. Es wurden differenzierte Kompetenzmodelle und vielfältige Methoden entwickelt. Neben Papierund-Bleistift-Tests stehen Verfahren, die Kompetenzen in realitätsnachempfundenen simulierten Handlungssituationen erheben. Im Vortrag werden insbesondere Arbeiten aus dem Forschungsverbund "Professionswissen in der Lehramtsausbildung Physik (ProfiLe-P)" vorgestellt.

Prize TalkSYAW 1.2Wed 14:40AudimaxAstronomy at Highest Angular Resolution - Adaptive Optics, Interferometryand Black Holes• FRANK EISENHAUERMax Planck Institute for extrater-restrial PhysicsLaureate of the Stern Gerlach Medal 2022

The last 25 years have seen wonderful times in high angular resolution, optical/infrared ground-based astronomy: active optics enabled ever larger telescopes, adaptive optics now routinely remove the image blur from the turbulent earth atmosphere, and integral field spectroscopy adds the third, spectral dimension to highest resolution imaging. And most recently, optical/infrared interferometry brought yet another revolution by synthesizing a telescope with an equivalent diameter of a hundred meter and more. This presentation will describe the development and advances in the field, and especially the pioneering SINFONI and GRAVITY instruments. The two instruments have been crucial to the discovery and characterization of the supermassive black hole at the center of our Galaxy, and have revolutionized the study of nearby exoplanets and star forming galaxies in the Early Universe. The presentation will end with an outlook on the next generation of interferometric instruments and extremely large telescopes, which are already in construction for the next revolution to come.

Prize TalkSYAW 1.3Wed 15:10AudimaxTurbulence in one dimension — •ALEXANDER M. POLYAKOV — Princeton University, Department of Physics, Jadwin Hall — Laureate of the Max Planck Medal2021

I will discuss general features of turbulence using a soluble model. The topics include spontaneous breaking of the Galilean symmetry, universality, dissipative anomalies, general covariance, and non-Gaussian probability distributions.

Location: Audimax

The Nature of Science

jointly organised by the Physics Education Division (DD), the History of Physics Division (GP), the Gravitation and Relativity Division (GR), and the Theoretical and Mathematical Physics Division (MP)

Christian Forstner Ernst-Haeckel-Haus - Wissenschaftsgeschichte Friedrich Schiller University Jena christian.forstner@uni-jena.de

For some decades, the term "Nature of Science" has given a headline to discussions in the history and philosophy of science about the ways in which science is developed by people within certain personal, historical, social and theoretical frameworks. Likewise, science education has considered what understanding of the nature of science learners hold and how such an understanding is developed and might be guided and supported.

In this symposium we take up new and challenging perspectives on and developments in science and society to enhance the discussion. The perspectives presented in the contribution to the symposium relate to three different aspects:

Lucy Avraamidou will share empirical results about the role of women and their representation in science. Developments in theoretical physics as well as their growing discrepancies to ongoing structural practices will be discussed by Daniel Harlow. As a completion, the change in medial communication and its effect on the discussion of scientific results and knowledge with a special emphasis on the role of social media will be presented by Dietmar Hoettecke.

Overview of Invited Talks and Sessions

(Lecture hall Audimax)

Invited Talks

SYNS 1.1	Tue	14:00-14:30	Audimax	The Role of Nature of Science Education for Science Media Literacy – •DIETMAR
				Ноттеске
SYNS 1.2	Tue	14:30-15:00	Audimax	What kinds of identities are deemed in/our of place in physics? — •LUCY AVRAAMI-
				DOU
SYNS 1.3	Tue	15:00-15:30	Audimax	Some thoughts on the status of theoretical physics — •DANIEL HARLOW

Sessions

SYNS 1.1–1.3 Tue 14:00–15:40 Audimax The Nature of Science

Sessions

– Invited Talks –

SYNS 1: The Nature of Science

Time: Tuesday 14:00-15:40

Invited Talk SYNS 1.1 Tue 14:00 Audimax The Role of Nature of Science Education for Science Media Literacy —

•DIETMAR HÖTTECKE — Faculty of Education for Science Melanburg •DIETMAR HÖTTECKE — Faculty of Education, University of Hamburg Science is inevitably mediated to the public sphere and both professional journalistism and social media networks play important roles. For well-informed decision-making, it is essential for citizens to know how scientists communicate with each other, as well as with the public. Until recently, the conventional mass media (e.g. newspapers) typically functioned as gatekeepers, helping to assess the reliability and trustworthyness of scientific claims. In today's culture, media and their gatekeeping roles are rapidly vanishing. In social media information flows along existing networks, sometimes heedless of scientific expertise and quality of information. As a result, we need an expanded conception of nature of science (NOS): First, students need to learn about the epistemics of communicative practices, within science and in society, science as a system of distributed knowledge and expertise characterized by division of labor as well

distributed knowledge and expertise, characterized by division of labor as well as a social system of checks and balances, trust and credibility. Second, students have to learn about the epistemic structure of science communcation and the role of "gatekeepers". Here, the role of social media and its correlated phenomena have to be considered like aggregated news, filter bubbles, echo chambers, spirals of silence, fake news, and purposeful disinformation. Third, the "consumer" of science has to learn about him- and herself, including the role of confirmation bias, motivated reasoning, and the social context of trust. These three perspectives finally lead to a the idea of Science Medica Literacy as an expansion of more traditionally NOS perspecitve.

Invited Talk

SYNS 1.2 Tue 14:30 Audimax

Location: Audimax

What kinds of identities are deemed in/our of place in physics? — •LUCY AVRAAMIDOU — University of Groningen, Groningen, Netherlands

By adopting an intersectional approach, I will present the findings of an empirical study that aimed to examine the ways in which physics identity intersects with other identities (i.e., racial identity, gender identity, Islamic religious identity, social-class identity, single-motherhood, and ethnic identity) and influences women's recognition. To do that I draw upon a life history, multiple case-study of three women in physics: a native to Western Europe, late-career white woman; two immigrant women to Western Europe, one is an undergraduate student of color, and the other, an early-career Muslim woman. With evidence gathered from this empirical study I will argue that a conceptualization of physics identity that does not value people for who they are in their entirety, made up of multiple and intersectional identities, but only values how people produce or consume scientific knowledge is exclusionary and only serves to create suffering. For an exploration of women's participation in physics, any attempt to examine gender in isolation instead of in intersection with other multiple identities is an ill-equipped way of examining the complexities and dynamics of contemporary identity politics embedded in recognition.

Invited TalkSYNS 1.3Tue 15:00AudimaxSome thoughts on the status of theoretical physics• DANIEL HARLOWMIT Department of Physics, Cambridge, MA

Abstract: Theoretical physics has changed substantially from what it was in the post-war period, but these changes are often not reflected in funding and hiring structures. I will present a contemporary assessment of what the goals of and methods of theoretical physics are, and how it relates to society more broadly.

General Discussion

Extraterrestrial Physics Division Fachverband Extraterrestrische Physik (EP)

Miriam Sinnhuber Karlsruher Institut für Technologie Hermann-von-Helmholtz-Platz 1 76344 Eggenstein-Leopoldshafen miriam.sinnhuber@kit.edu

Overview of Invited Talks and Sessions

(Lecture hall EP-H1)

Plenary and Prize Talks related to EP

SYAW 1.2	Wed	14:40-15:10	Audimax	Astronomy at Highest Angular Resolution - Adaptive Optics, Interferometry and
				Black Holes — •Frank Eisenhauer
PV VIII	Thu	9:45-10:30	Audimax	The Sun as a source of high-energy particles — •RAMI VAINIO

Special topic day T/EP: Astroparticles - From the source to the detector

EP 8.1-8.3	Thu	11:00-12:30	T-H15	Astroparticles: Invited talks (joint session T/EP)
EP 10.1-10.4	Thu	14:00-15:40	T-H15	Astroparticles: Invited topical talks (joint session T/EP)
EP 11.1–11.9	Thu	16:15-18:30	EP-H1	Astroparticles: From the source to the detector (joint session EP/T)

Invited Talks

ED 1 1	Man	14.00 14.20	ED III	Astonossiemala av af nod giant atoma Stavy A Haveyan
EP 1.1	MOII	14:00-14:50	EP-III	Asteroseismology of red-giant stars — •SASKIA FIERKER
EP 2.3	Mon	16:45-17:15	EP-H1	Cloudy with a hint of magnetic fields — •LUDMILA CARONE
EP 3.1	Tue	11:00-11:30	EP-H1	A new era of Venus exploration - seen Venus in a new light — •JÖRN HELBERT,
				Melinda Darby Dyar, Giulia Alemanno, Alessandro Maturilli, Nils Müller,
				Doris Breuer, VEM on VERTIAS team, VenSpec on Envision team
EP 4.1	Tue	14:00-14:30	EP-H1	Exploration of the Jupiter system with a small submillimetre wave telescope onboard
				the JUICE satellite — •PAUL HARTOGH
EP 5.1	Tue	16:15-16:45	EP-H1	Investigating Earth's atmosphere and ionosphere from space: How GNSS radio occul-
				tation measurements contribute to monitor the atmosphere in a high spatial resolution
				— •Christina Arras, Ankur Kepkar, Jens Wickert
EP 5.5	Tue	17:30-18:00	EP-H1	Solar extreme events and their impact on the middle atmosphere — •THOMAS RED-
				dmann, Monali Borthakur, Miriam Sinnhuber, Ilya Usoskin, Jan-Maik Wissing
EP 6.1	Wed	11:00-11:30	EP-H1	Three-dimensional topology-driven magnetic reconnection — •RAQUEL MÄUSLE, JEAN-
				Mathieu Teissier, Wolf-Christian Müller
EP 12.1	Fri	11:00-11:30	EP-H1	Linking Solar Eruptions and Energetic Particles through Observations and Modeling
				— •Frederic Effenberger
EP 12.4	Fri	12:00-12:30	EP-H1	Solar Orbiter: two years of operations and first results – •Frederic Schuller,
				Alexander Warmuth, Gottfried Mann

Invited Talks of the joint symposium SMuK Dissertation Prize 2022 2022 (SYMD)

See SYMD for the full program of the symposium.

SYMD 1.1	Mon	14:00-14:25	Audimax	Timeless Quantum Mechanics and the Early Universe — •LEONARDO CHATAIGNER
SYMD 1.2	Mon	14:25-14:50	Audimax	First tritium β -decay spectrum recorded with Cyclotron Radiation Emission
				Spectroscopy (CRES) — • CHRISTINE CLAESSENS
SYMD 1.3	Mon	14:50-15:15	Audimax	Watching the top quark mass run - for the first time! — • MATTEO M. DEFRANCHIS,
				Katerina Lipka, Sven-Olaf Moch
SYMD 1.4	Mon	15:15-15:40	Audimax	Towards beam-quality-preserving plasma accelerators: On the precision tuning
				of the wakefield — •SARAH SCHRÖDER

Invited Talks of the joint Awards Symposium (SYAW) See SYAW for the full program of the symposium.

SYAW 1.1	Wed	14:10-14:40	Audimax	Wie überprüft man die Ziele der Lehramtsausbildung Physik? — •Horst Schecker
SYAW 1.2	Wed	14:40-15:10	Audimax	Astronomy at Highest Angular Resolution - Adaptive Optics, Interferometry and Black Holes — •FRANK EISENHAUER
SYAW 1.3	Wed	15:10-15:40	Audimax	Turbulence in one dimension — •Alexander M. Polyakov

Sessions

EP 1.1–1.6	Mon	14:00-15:45	EP-H1	Astrophysics
EP 2.1–2.9	Mon	16:15-18:45	EP-H1	Astrophysics / Exoplanets and Astrobiology
EP 3.1-3.7	Tue	11:00-13:00	EP-H1	Planets and Small bodies
EP 4.1-4.6	Tue	14:00-15:35	EP-H1	Planets and Small bodies
EP 5.1-5.8	Tue	16:15-18:45	EP-H1	Near Earth Space
EP 6.1-6.8	Wed	11:00-13:15	EP-H1	Astrophysics
EP 7.1–7.11	Wed	16:15-18:50	EP-H1	Astrophysics
EP 8.1-8.3	Thu	11:00-12:30	T-H15	Astroparticles: Invited talks (joint session T/EP)
EP 9	Thu	12:45-13:45	EP-MV	Mitgliederversammlung Extraterrestrische Physik
EP 10.1-10.4	Thu	14:00-15:40	T-H15	Astroparticles: Invited topical talks (joint session T/EP)
EP 11.1–11.9	Thu	16:15-18:30	EP-H1	Astroparticles: From the source to the detector (joint session EP/T)
EP 12.1-12.11	Fri	11:00-14:00	EP-H1	Sun and Heliosphere

Annual General Meeting of the Extraterrestial Physics Division

Thursday 12:45-13:45 EP-MV

Sessions

- Invited, Invited Topical, and Contributed Talks -

EP 1: Astrophysics

Time: Monday 14:00-15:45

Invited TalkEP 1.1Mon 14:00EP-H1Asteroseismology of red-giant stars — •SASKIA HEKKER — Centre for Astronomy (ZAH/LSW), Heidelberg University — Heidelberg Institute for TheoreticalStudies (HITS) — Stellar Astrophysics Centre, Aarhus, Denmark

Over the past decades we experienced a revolution in asteroseismology of redgiant stars. In this talk, I will discuss this revolution and first insights gained from asteroseismology into the stellar structure of red-giant stars.

EP 1.2 Mon 14:30 EP-H1

Finite time singularities and their relations to eruptive stellar mass loss events — •DIETER NICKELER and MICHAELA KRAUS — Astronomical Institute, Czech Academy of Sciences, Ondrejov, Czech Republic

In recent years rings, spiral and arc structures have been detected around massive stars with strong stellar winds. These structures are proposed to be a result of stellar mass ejections. Taking a (fast) stationary stellar wind as initial condition, the time-dependent amplitude modulation is analyzed with respect to eventually occurring finite-time instabilities. These so called blow-up solutions can be interpreted as non-linear instabilities in the stellar wind, leading to eruptive mass loss of the star. We analyze under which circumstances such blow-up solutions can exist.

EP 1.3 Mon 14:45 EP-H1

Molecular environment of the yellow hypergiant HD 269953 — •MICHAELA KRAUS¹, MARIA LAURA ARIAS², and LYDIA CIDALE² — ¹Astronomical Institute, Czech Academy of Sciences, Ondrejov, Czech Republic — ²Instituto de Astrofísica de La Plata, CONICET-UNLP, Argentina

Yellow hypergiants are massive stars, most likely in post-red supergiant evolutionary state. Stars in this phase can undergo multiple outbursts, and the ejected material might enshroud the stars in gaseous and dusty shells or envelopes. The object HD 269953 has been suggested to be a candidate yellow hypergiant. Although no historic outburst has been reported for that object, its environment hosts a substantial amount of warm CO gas. To unveil the dynamics within the molecular gas we obtained a high-resolution (R ~ 45 000) K-band spectrum of HD 269953 with IGRINS at GEMINI-South. We find that the spectrum is rich in emission features. In particular, we detect emission from the ¹²CO and ¹³CO molecular bands. The latter are strongly enriched, in agreement with the hypothesis that the environment contains processed matter that has been released from an evolved object. Moreover, we identified emission of hot water vapor, which is, to our knowledge, the first detection of water in the vicinity of an evolved massive star. We will present first results from our analysis of the circumstellar molecular gas and discuss scenario for its origin.

EP 1.4 Mon 15:00 EP-H1

Kilonovae, gamma-ray bursts, and heavy elements from neutron star mergers — •Oliver Just¹, Andreas Bauswein¹, Thomas Janka², Stephane Goriely³, Ina Kullmann³, Shigehiro Nagataki⁴, Hirotaka Ito⁴, and Christine Collins¹ — ¹GSI, Darmstadt, Germany — ²MPA, Garching, Germany — ³ULB, Brussels, Belgium — ⁴RIKEN, Tokyo, Japan

The collision of two neutron stars, which was first observed in 2017, is one of the most luminous astrophysical explosions, in which not only electromagnetic radiation from radio to gamma-ray frequencies is emitted, but also huge amounts of energy in the form of neutrinos and gravitational waves. In order to decipher these multi-messenger events, e.g. for inferring the mass and composition of the material thrown out in the course of such a merger, one needs to build detailed theoretical models of the processes that lead to matter ejection and the emission of radiation. In this talk I will present our recent efforts to model the merger and its remnant using multi-dimensional hydrodynamics simulations including the transport of neutrinos. Moreover, I will outline how these simulations were used to predict the optical/near infrared signal, called kilonova, the flash of gamma-radiation, called short gamma-ray burst, as well as the amount of heavy elements, such as Gold, produced in the ejecta by means of the rapid neutron-capture (or r-) nucleosynthesis process.

EP 1.5 Mon 15:15 EP-H1

Location: EP-H1

Core-Collapse Simulations of Very Massive Star: Gravitational Collapse, Black-hole Formation, and Beyond — •NINOY RAHMAN — GSI Helmholtz Centre for Heavy Ion Research

We investigate the final gravitational collapse of rotating and non-rotating pulsational pair-instability supernova progenitors with zero-age-main-sequence masses of 60, 80, and 115 M_{\odot} and iron cores between 2.37 M_{\odot} and 2.72 M_{\odot} by 2D (axi-symmetric) hydrodynamics simulations. Using the general relativistic NADA-FLD code with neutrino transport allows us to follow the evolution beyond the moment when the transiently forming neutron star (NS) collapses to a black hole (BH), which happens in all cases. Because of high neutrino luminosities and mean energies, neutrino heating leads to shock revival before BH formation in all cases except in the rapidly rotating 60 M_{\odot} model, where centrifugal effects support a higher NS mass but reduce the neutrino-heating rate by roughly a factor of two compared to the non-rotating counterpart. After BH formation the neutrino luminosities drop steeply but continue on a 1-2 orders of magnitude lower level for several 100 ms because of aspherical accretion of neutrino and shock-heated matter, before the ultimately spherical collapse of the outer progenitor shells suppresses the neutrino emission to negligible values. In all shock-reviving models BH accretion swallows the entire neutrino-heated matter and the diagnostic explosion energies decrease to zero within a few seconds latest. Nevertheless, the shock or a sonic pulse move outward and may trigger mass loss, which we estimate by long-time simulations with the PROMETHEUS code.

EP 1.6 Mon 15:30 EP-H1

Location: EP-H1

Modeling disk fragmentation and multiplicity in massive star formation — •ROLF KUIPER¹ and ANDRÉ OLIVA² — ¹Universität Heidelberg — ²Universität Tübingen

There is growing evidence that massive stars grow by disk accretion in a similar way to their low-mass counterparts. Early in evolution, these disks can achieve masses that are comparable to the current stellar mass, and therefore the forming disks are highly susceptible to gravitational fragmentation. We investigate the formation and early evolution of an accretion disk around a forming massive protostar, focussing on its fragmentation physics.

We used a grid-based, self-gravity radiation hydrodynamics code including a sub-grid module for stellar evolution and dust evolution. We purposely do not use a sub-grid module for fragmentation such as sink particles to allow for all paths of fragment formation and destruction, but instead we keep the spatial grid resolution high enough to properly resolve the physical length scales of the problem, namely the pressure scale height and Jeans length of the disk.

The cloud collapses and a massive (proto)star is formed in its center surrounded by a fragmenting Keplerian-like accretion disk with spiral arms. The fragments have masses of \sim 1 Msol, and their continuous interactions with the disk, spiral arms, and other fragments result in eccentric orbits. Fragments form hydrostatic cores surrounded by secondary disks with spiral arms that also produce new fragments. Based on this, we study the multiplicity from spectroscopic multiples to companions at distances at 1000 au.

EP 2: Astrophysics / Exoplanets and Astrobiology

Time: Monday 16:15-18:45

EP 2.1 Mon 16:15 EP-H1 Cross-matching Low Frequency Array (LOFAR) Sources — •Lukas Böhme

— Fakultät für Physik, Universität Bielefeld, Deutschland Multi-frequency studies of the radio sky provide insight into the nature of the observed objects. To do so, I cross-match resolved and unresolved radio sources at different frequencies and angular resolutions. The combination of several radio bands from the Low Frequency Array (LO-FAR) and other radio telescopes with multi-frequency optical and infrared data allows for photometric redshift estimates, not only for LOFAR Two-metre Sky Survey (LoTSS) sources, but also for lower resolution LOFAR LBA Sky Survey (LoLSS) and NRAO VLA Sky Survey (NVSS) sources.

I present a new cross-matching algorithm incorporating the radio source ex-

I study the number of components of LoLSS radio sources and their spectral properties. I find an average spectral index of $\alpha = -0.77 \pm 0.17$ for sources matched in all surveys. This spectral index is flux density independent above $S_{1400} = 30$ mJy and appears to be redshift independent.

EP 2.2 Mon 16:30 EP-H1

2-Kanal - Kohonenkarten zur Klassifizierung von Radioquellen und Identifizierung von optischen Host-Galaxien — **•**SCHWARZ KILIAN¹, STEVENS SIMON¹, MAH ZHEE KEIN JEREMY¹, EL-BEIT SHAWISH SARA¹, FARJOUD MA-SOULEH NEGAR¹, IMHOF DENNIS¹, RAPP STEFAN¹ und HOEFT MATTHIAS² — ¹Hochschule Darmstadt, Haardtring 100, 64295 Darmstadt — ²Thüringer Landessternwarte, Sternwarte 5, 07778 Tautenburg

Aufgrund der großen Datenmengen durch aktuelle Himmelsdurchmusterungen gewinnt die Klassifikation mit Hilfe von Methoden des maschinellen Lernens an Bedeutung. Für die ausgedehnten Quellen in der ersten Veröffentlichung von Daten des LOFAR Two-Metre Sky Survey (LoTSS) hat sich die automatische Gruppierung der Quellmorphologien mit Hilfe der Self-Organising-Maps (SOMs) als sehr leistungsvoll erwiesen [Mostert et. al 2020, Astronomy & Astrophysics]. In diesem Beitrag wird gezeigt, wie diese Methode weiterentwickelt werden kann, um die morphologische Klassifikation von Radioquellen zu verbessern und die zugehörigen Wirtsgalaxien in optischen Karten [Chambers et. al 2019, arXiv] zu identifizieren. Hierzu werden 2-Kanal-Kohonenkarten mit Hilfe der PINK-Software [Polsterer et. al 2016, 24th European Symposium on Artificial Neural Networks] trainiert und anschließend analysiert. Durch das Setzen eines Begrenzungsrahmens wird die Anzahl der möglichen zugehörigen Wirtsgalaxien erheblich eingeschränkt. In weiteren Schritten werden andere Eigenschaften der Quellen, z.B. Rotverschiebung, zusätzlich in dem Lernverfahren berücksichtigt. Die aktuellen Ergebnisse werden präsentiert.

 Invited Talk
 EP 2.3
 Mon 16:45
 EP-H1

 Cloudy with a hint of magnetic fields
 - LUDMILA
 CARONE
 St Andrews

 University, St Andrews, UK — Max Planck Institute for Astronomy, Heidelberg, D
 D
 D
 D

In this review talk, I will present the state of the art of numerical models of exoplanet atmospheres. I will show why it is important to consider that exoplanets are three dimensional objects that can change their local observed atmosperic gas phase composition (C/O ratio) via cloud formation and disequilibrium chemistry. Since C/O ratios are used as proxies for exoplanet formation it is important that 3D processes are incorporated in the interpretation of upcoming detailed observations with next generation space telescopes: i.e. the James Webb Space Telescope, ARIEL and also PLATO. Last but not least, tackling magnetic fields interaction and how they impact observable properties in exoplanet atmospheres is an ongoing numerical challenge.

EP 2.4 Mon 17:15 EP-H1 Gap in Solar System's Proto Planetary Disk Likely Confirm Features in the Distribution of Exoplanet Semi-Major Axes — •STUART F. TAYLOR — SETI Affiliate, Mountain View, CA USA — Participation Worldscope, Hong Kong

The recent discovery of a gap dividing the solar system's protoplanetary disk (SS's PPD) may be a confirmation of the peak-gap-peak (PGP) feature in the distribution of semi-major axes of exoplanets hosted by stars most like the sun. This PGP feature was published before the SS PPD gap was presented. It has long been thought that PPDs form with separate inner and outer disks separated by a gap, but seeing this structure in the distribution of exoplanet semi-major axes is unexpected due to how it is thought that primordial features are erased by subsequent planet migration. The solar system gap is reported to be closer than 3 AU, while the PGP feature's gap extends from 1.5 to 1.9 AU for planets of solar mass stars. The two results taken together suggest that planets of stars that are similar to the sun or with higher metallicity may generally start their evolution with a gap in this range, likely associated with a snow line. We are now finding that the semi-major axis of this feature appears to scale with the square root of stellar mass. We also propose that the study of planetary system architectures and demographics be organized in the form of a new additional section of an exoplanet catalog, which would include results on exoplanet occurrence distributions and findings of features and correlations among exoplanet parameters.

EP 2.5 Mon 17:30 EP-H1

Composition of super-Earths, super-Mercuries, and their host stars — •VARDAN ADIBEKYAN — nstituto de Astrofísica e Ciências do Espaço (IA) Because of their common origin, it is expected (or assumed) that the composition of planet building blocks should (to a first order) correlate with stellar atmospheric composition, especially for refractory elements. In fact, information on the relative abundance of refractory and major rock-forming elements such as Fe, Mg, Si are commonly used to improve interior estimates for terrestrial planets (e.g. Dorn et al. 2015; Unterborn et al. 2016) and has even been used to estimate planet composition in different galactic populations (Santos et al. 2017). However, there is no direct observational evidence for the aforementioned expectation/assumption and was even recently questioned by Plotnykov & Valencia (2020). By using the largest possible sample of precisely characterized low-mass planets and their host stars, we show that the composition of the planet building blocks indeed correlates with the properties of the rocky planets. We also find that on average the iron-mass fraction of planets is higher than that of the primordial values, owing to the disk-chemistry and planet formation processes. This result can bring important implications for the future modeling of exoplanet composition.

EP 2.6 Mon 17:45 EP-H1

Planetary interiors via Love-number determined from radial velocities — •LIA MARTA BERNABÒ and SZILÁRD CSIZMADIA — DLR Berlin, Institut für Planetenforschung (Deutsches Zentrum für Luft-und Raumfahrt) - Rutherfordstr. 2, 12489 Berlin

We study the inner structure of planets by determining the Love numbers kn and hn (Love, 1911), which describes the susceptibility of their shape to change in response to a tidal and rotational potential. The second degree Love number k2 is highly sensitive to the thickness of the interior layers and rheology of the planet and it is proportional to the concentration of mass towards the centre of the planet, therefore it is used to infer the internal structure of the body. We will present the method how to analyze the radial velocity curve in presence of apsidal motion caused by tidal interaction between the star and the planet, and by general relativity. The former one can be linked to the Love-number k2 (Kopal, 1959) which constrains the planetary interior as a third measurable parameter beyond the mass and radius (Baumeister et al., 2020). We also study the effect of the rotationally and tidally distorted stellar shape on the radial velocity curves. This causes a distortion on the RV-shape as well, and leads to the presence of an apparent eccentric orbit in the RV-curve. We show that the correct estimate of this effect must be taken into account and some other studies overestimated its amplitude and significance. Finally, we show our first results when we applied our method to real systems.

EP 2.7 Mon 18:00 EP-H1 Stellar Flares and Habitable(?) Worlds from the TESS Primary Mission — •MAXIMILIAN N. GÜNTHER — European Space Agency (ESA), European Space Research and Technology Centre (ESTEC), Keplerlaan 1, 2201 AZ Noordwijk, Netherlands

On our search for habitable worlds, we have to account for explosive stellar flaring and coronal mass ejections (CMEs) impacting exoplanets. These stellar outbursts are a double-edged sword. On the one hand, flares and CMEs are capable of stripping off atmospheres and extinguishing existing biology. On the other hand, flares might be the (only) means to deliver the trigger energy for prebiotic chemistry and initiate life. This talk will highlight our study of all stellar flares from the TESS primary mission, driven by a convolutional neural network. I will discuss our new insights on flaring as a function of stellar type, age, rotation, spot coverage, and other factors. Most importantly, I will link our findings to prebiotic chemistry and ozone sterilisation, identifying which worlds might lie in the sweet spot between too much and too little flaring. With future extended missions and increased coverage, flare studies and new exoplanet discoveries will ultimately aid in defining criteria for habitability.

EP 2.8 Mon 18:15 EP-H1

INCREASE - An updated model suite to study the INfluence of Cosmic Rays on Exoplanetary AtmoSpherEs — •KONSTANTIN HERBST¹, J. LEE GRENFELL², MIRIAM SINNHUBER³, and FABIAN WUNDERLICH² — ¹Institut für Experimentelle and Angewandte Physik, Christian-Albrechts-Universität zu Kiel, 24118 Kiel — ²Institut für Planetenforschung, Deutsches Zentrum für Luftund Raumfahrt, 12489 Berlin — ³Institut für Meteorologie und Klimaforschung, Karlsruher Institut für Technologie, 77344 Eggenstein-Leopoldshafen

Much remains to be explored when it comes to the diversity of exoplanetary atmospheres, much remains to be explored. For a few selected objects such as Proxima Centauri b, first observations of the atmosphere have already been achieved. At the same time, much more information is expected soon with the help of JWST launched in December 2021. However, to interpret existing and upcoming observations, model studies of planetary atmospheres that account for various processes - e.g., atmospheric escape, outgassing, climate, photochemistry, and the physics of air showers - are necessary. Here, we present our unique model suite INCREASE investigating the impact of cosmic rays on diverse exoplanetary atmospheres and their habitability.

EP 2.9 Mon 18:30 EP-H1 Viruses first? — •KARIN MOELLING — Inst. med. Mikrobiologie, Universität Zürich, Gloriastr 30, 8006 Zürich, Schweiz — Max Planck Institut Mol.Gen. Ihnestr 73 14195 berlin Deutschland

The discovery of exoplanets within putative habitable zones and spaceships to Mars and Eraly Earth raise interest in the origin of life. Could viruses play a role at the beginning of life on Earth and elsewhere? Viruses are the most successful species on Earth in every habitat. Contemporary viruses reflect evolution ranging from the early RNA world to the present DNA-protein world. Earliest replicating and evolving entities are enzymatic non-coding RNA molecules,

which developed indivdual survival metabolisms depending on extreme living conditonas as extrempohils. Autonomy of early life forms may have been given up by energy-saving parasitic life forms today.(Moelling and Breocker,2019, doi: 10.3389/fmicb.2019.0052310.3389/fmicb.2019.00523)

EP 3: Planets and Small bodies

Time: Tuesday 11:00-13:00

Invited Talk

EP 3.1 Tue 11:00 EP-H1 A new era of Venus exploration - seen Venus in a new light — •JÖRN HELBERT¹, Melinda Darby Dyar², Giulia Alemanno¹, Alessandro Maturilli¹, Nils Müller¹, Doris Breuer¹, VEM on VERTIAS TEAM¹, and VenSpec on En-VISION TEAM $^1-^1$ Institute for Planetary Research, DLR, Berlin, Germany – ²Mount Holyoke College, USA

Venus is our next-door planet. It is almost identical in size to Earth and yet we know so little about it. Recently three new mission have been selected to study Venus - the ESA EnVision and the NASA VERITAS and DAVINCI missions. The new interest in Venus has partly come through discussions with the exoplanet community about why their models always lean towards Earth-like planets. They have asked for fundamental parameters for Venus to improve their models, like the surface composition of Venus, and we just do not have the answers currently. This led to the realisation that we need to find out more about this planet that has evolved in such a different way from the Earth in order to understand how habitable planets evolve in general.

All three recently selected Venus missions include in their payload instruments focused on the 1 micron region. The NASA VERITAS and ESA EnVision missions use the DLR build Venus Emissivity Mapper (VEM) as a multi-spectral imaging system. The DAVINCI mission has a 1 micron descent imager. These new instruments have been made possible in part by a dedicated effort to set up a new Venus high temperature spectroscopy laboratory at DLR to routinely obtain VNIR emissivity spectra at relevant Venus surface temperatures.

EP 3.2 Tue 11:30 EP-H1

The impact of large solar particle events on the chemical composition of the **Martian atmosphere** — •MIRIAM SINNHUBER¹, JOHN LEE GRENFELL², KON-STANTIN HERBST³, and FABIAN WUNDERLICH² — ¹Karlsruher Institute für Tech-nologie, Karlsruhe, Germany — ²DLR Institut für Planetenforschung, Berlin, Germany — ³Universität Kiel, Kiel, Germany

Large solar coronal mass ejections are known to have a large impact on the chemical composition of the high-latitude atmosphere of Earth. Collision of the incident protons and resulting secondary electrons with the most abundant atmospheric constituents leads to dissociation, ionization, and dissociative ionization of these substances; in the Earth's atmosphere, these are N_2 and $\mathrm{O}_2,$ and the main products are nitrogen radicals and nitrogen oxides (NOx: N, NO, NO₂) as well as hydrogen oxides (HOx: OH, HO₂) from the uptake of water vapor into large cluster ions. Both NOx and HOx species contribute to catalytic ozone loss, and very rapid loss of ozone in the terrestrial polar stratosphere and mesosphere is well-documented. However, not much is known about the impact of these events on other planets. Here, we present results from model experiments for the atmosphere of Mars, considering three different solar particle events: a groundlevel event (1956), the Carrington white light flare (1859), and one of the largest ground-level events found in the paleo-record so far, the AD774/775 event. The analysis focuses on the different responses of the thin Martian atmosphere with its low amounts of nitrogen and high CO_2 mixing ratio.

EP 3.3 Tue 11:45 EP-H1

The interaction of the Martian with the solar wind as observed by MAVEN -•LUKAS MAES and MARKUS FRAENZ — Max Planck Institute for Solar System Research, Göttingen, Germany

The Mars Atmosphere and Volatile EvolutioN (MAVEN) mission was launched in 2013 to study the atmosphere and ionosphere of Mars, its interaction with the solar wind, and the consequences for the erosion of the Martian atmosphere. With a comprehensive and complementary set of plasma and neutral gas instruments, it has offered higher resolution data than ever before, with a dataset of over 6 years now. In this talk we will look at a some results about the plasma physical processes in and around the Martian ionosphere observed by the MAVEN satellite and discuss them in the context of Mars' atmospheric evolution, the effect of Mars' crustal magnetic fields, and what we can learn from it about other planets.

EP 3.4 Tue 12:00 EP-H1

Dynamo models reproducing the offset dipole of Mercury's magnetic field -•PATRICK KOLHEY¹, DANIEL HEYNER¹, JOHANNES WICHT², THOMAS GASTINE³, and FERDINAND PLASCHKE¹ — ¹Institut für Geophysik und extraterrestrische Physik, Technische Universität Braunschweig, Braunschweig, Germany -

Location: EP-H1

²Max-Planck-Institut für Sonnensystemforschung, Göttingen, Germany -³Institut de Physique du Globe de Paris, Université de Paris, Paris, France

Since the discovery of Mercury's peculiar magnetic field it has raised questions about the dynamo process in its fluid core. The global magnetic field at the surface is rather weak compared to other planetary magnetic fields, strongly aligned to the planet's rotation axis and its magnetic equator is shifted towards north. Especially the latter characteristic is difficult to explain using common dynamo model setups. In this study we present new direct numerical simulations of the magnetohydrodynamical dynamo problem which include a stably stratified layer on top of the outer core, which can also reproduce the shift of the magnetic equator towards north. We revisit a model configuration for Mercury's dynamo action, which successfully reproduced the magnetic field features, in which core convection is driven by thermal buoyancy as well as compositional buoyancy (double-diffusive convection). While we find that this model configuration produces Mercury-like magnetic field only in a limited parameter range (Rayleigh and Ekman number), we show that also a simple codensity model is sufficient over a wide parameter range to produce Mercury-like magnetic fields.

EP 3.5 Tue 12:15 EP-H1

Revised Modular Model of Mercury's Magnetospheric Magnetic Field -•KRISTIN PUMP, DANIEL HEYNER, and FERDINAND PLASCHKE — Institut für Geophysik und extraterrestrische Physik, TU Braunschweig

Mercury is the smallest an innermost planet of our solar system and has a dipoledominated internal magnetic field that is relatively weak, very axisymmetric and significantly offset towards north. Through the interaction with the solar wind, this field leads to a magnetosphere. Compared to the magnetosphere of Earth, Mercury's magnetosphere is smaller and more dynamic. To understand the magnetospheric structures and processes we use in-situ MESSENGER data to develop a semi-empiric model, which can explain the observations and help to improve the mission planning for the BepiColombo mission en-route to Mercury.

We will present this semi-empiric KTH-model, a modular model to calculate the magnetic field inside the Hermean magnetosphere. Korth et al. (2015 and 2017) published a model, which is the basis for the KTH-Model. In this new version, the calculation of the magnetic field for the neutral current sheet is restructured based on observations rather than ad-hoc assumptions so that the description is more realistic. Furthermore, a new model is added to depict the partial ring current. An analysis of the residuals shows a better visibility of the field-aligned currents. In addition, this model offers the possibility to improve the main field determination.

EP 3.6 Tue 12:30 EP-H1

BepiColombo at Mercury: First close-in magnetic field measurements from the southern hemisphere — •DANIEL HEYNER — TU Braunschweig

The internal magnetic field of Mercury is best described by a northward offset dipole with almost zero obliquity. Its offset, weakness, axisymmetry and lack of secular variation still poses a challenge to dynamo theory. After NASA's Mariner 10 flybys in the 1970's and MESSENGER's orbital mission in 2011-2015, Bepi-Colombo performed a flyby at Mercury in October 2021. For the first time, magnetic field measurements are obtained from the southern hemisphere close to the planet by the fluxgate magnetometer MPO-MAG. We will present an overview of the flyby data highlighting different plasma regions and compare the new insitu data to magnetospheric models obtained from the previous missions to the innermost terrestrial planet. Does the flyby data reveal any secular variation? Has the dipole offset changed? These are some of the questions we will discuss with this unprecedented magnetometer data. We will close with a discussion on what is to be expected from the orbital phase of BepiColombo.

 $EP~3.7\quad Tue~12:45\quad EP-H1\\ \textbf{Binary main-belt comet}~\textbf{288P}-\textbf{\bullet}Jessica~Agarwal^{1,2} and Yoonyoung~Kim^1$ - ¹Institut für Geophysik und Extraterrestrische Physik, TU Braunschweig — ²Max-Planck-Institut für Sonnensystemforschung, Göttingen

Main-belt comets are asteroids that activate during subsequent perihelion passages, emitting dust like comets. The object 288P is currently the only known comet-like object that is also a gravitationally bound binary system. While binaries are common in the asteroid population, 288P is special even among these because of its combination of two similarly sized components and wide separation of about 100 times the object radius.

Possible formation scenarios include the likely involvement of rotational splitting, radiation-induced torques and outgassing-induced torques, but the detailed evolutionary history of the system remains to be understood. A key question in this context is whether one or both components are active.

We present data obtained with the Hubble Space Telescope in autumn 2021 that show the onset of activity in 288P as it approached perihelion.

EP 4: Planets and Small bodies

Time: Tuesday 14:00-15:35

Invited Talk

EP 4.1 Tue 14:00 EP-H1 Exploration of the Jupiter system with a small submillimetre wave telescope onboard the JUICE satellite — •PAUL HARTOGH — Justus-von-Liebig-Weg 3,

37077 Göttingen, Germany JUICE - JUpiter ICy moons Explorer it the first large class mission of ESA's Cosmic Vision 2015 - 2025 program. The JUICE satellite is planned to be launched in 2023 and will arrive in 2032. The primary mission in the Jupiter system will take about three years. The focus of the mission is Jupiter itself and the Galilean satellites, their internal oceans and potential habitability. Recent ground-based observations of Europa and Ganymede showed water vapor plumes, probably related to geysers on their surfaces. JUICE intends to identify the geysers, monitor their potential activity and molecular and isotopic composition in order to constrain satellite formation models and development processes (of chemical, physical and potentially biologic nature) in the interior of their oceans. Jupiter itself is seen as an archetype of a gas giant. A better understanding of its atmospheric processes will be a baseline for a better understanding of gas giants outside our solar system. JUICE will characterize the general circulation of Jupiter's atmosphere, its meteorology, chemistry and structure between the upper cloud deck and the ionosphere and magnetosphere. The Submillimetre Wave Instrument (SWI) is part of the JUICE science payload. SWI covers two spectral bands between 530 and 1275 GHz. The SWI functionalities and specifications as well as required technology developments during the last decades and how the unique capabilities of SWI will help to answer JUICE key science questions will be presented.

EP 4.2 Tue 14:30 EP-H1

A PE-based radio propagation simulation for glaciers and ice moons with depth-dependent permittivity profiles — •GIANLUCA BOCCARELLA, ALEX KYRIACOU, and PIA FRIEND — Bergische Universität Wuppertal, Gaußstraße 20, 42119 Wuppertal

Enceladus Explorer (EnEx) is an initiative from the DLR to develop a melting probe that can reach a near surface water pocket on the ice moon Enceladus to search for microbial life. To find the water pocket a radar imaging system is needed, but the unknown permittivity of the surface ice will affect the accuracy of the measurements. We present an ice layering model of Enceladus including dielectric properties for the different types of ice layers, derived from Cassini data. We introduce a simulation which utilises Parabolic Equations (PE) to model radio propagation through inhomogeneous dielectric media such as ice environments on terrestrial glaciers or ice moons. Target objects embedded in the ice can be identified together with further reflections from the interface layers in the time-domain spectrum. By extracting the time of flight of respective reflected signals one can calculate the distance from the antennae to the target. Radar images, which would be obtained if a transmitter and receiver were placed on the melting probe moving vertically downwards through the ice, can be simulated and compared to each other for different permittivity profiles and targets. *This project is funded by the Enceladus Explorer Initiative of the DLR Space Administration

EP 4.3 Tue 14:45 EP-H1

On the Subsurface Exploration of Ocean Worlds in the Outer Solar System with the TRIPLE project — •MIA GIANG DO¹, JAN AUDEHM¹, DIRK HEINEN¹, JOHANNA HERMANNSGABNER¹, SHARIF EL MENTAWI¹, ANDREAS Nöll¹, Shreyans Sakhare¹, Christopher Wiebusch¹, Yuting Ye¹, Simon ZIERKE¹, CLEMENS ESPE², MARCO FELDMANN², and GERO FRANCKE² for the TRIPLE-nanoAUV-Collaboration -¹RWTH Aachen University - Physics Institute III B, Aachen, Germany -²GSI - Gesellschaft für Systementwicklung und Instrumentierung mbH, Aachen, Germany

In search of extraterrestrial life, the icy moons in the solar system are prime targets as they are suspected to harbor a subsurface ocean of liquid water. Future space missions to explore these subsurface ocean worlds are of great interest. A mission scenario includes landing on the surface, penetrating through the massive ice shell with a probe, and diving into the ocean with an underwater vehicle that is collecting samples for identifying potential habitats. Within TRIPLE (Technologies for Rapid Ice Penetration and subglacial Lake Exploration), a system containing three components: a melting probe, a miniaturized autonomous underwater vehicle and an in-situ astrobiological laboratory, is in development. Aiming for a mission to Jupiter's moon Europa, the system will be tested in an analogue environment in Antarctica. The talk will give an overview of the TRIPLE project with a focus on the technological challenges of the melting probe Location: EP-H1

and the latest status of the system. TRIPLE is a project line that has been initiated by the German Space Agency at DLR.

EP 4.4 Tue 15:00 EP-H1

TRIPLE-IceCraft - A Retrievable Melting Probe to Transport Scientific Payloads into Subglacial Lakes or Oceans — •DIRK HEINEN¹, SIMON ZIERKE¹, JAN AUDEHM¹, MIA GIANG DO¹, YUTING YE¹, CHRISTOPHER WIEBUSCH¹, MARCO FELDMANN², GERO FRANCKE², and CLEMENS $Espe^2 - {}^1RWTH$ Aachen University, III. Physikalisches Institut B, Otto-Blumenthal-Str., 52074 Aachen – ²GSI - Gesellschaft für Systementwicklung und Instrumentierung mbH, Liebigstraße 26, 52070 Aachen

Within TRIPLE, initiated by the German Space Agency at DLR, Technologies for Rapid Ice Penetration and subglacial Lake Exploration are being researched. TRIPLE aims to explore the subglacial ocean of the Jovian moon Europa. Prior to this flight mission a technology demonstration is planned in Antarctica. For accessing the subglacial water reservoir, a drill or melting probe needs to penetrate the ice first.

The TRIPLE-IceCraft melting probe is currently in development. It is a modular bus system for transporting standardized payloads through ice. The design will be suitable for transporting a scientific payload through several hundred meters of ice, penetrating into a subglacial ocean or lake, and returning later to the surface. The TRIPLE-IceCraft will be demonstrated in an analog scenario at the Ekström Ice Shelf in Antarctica in 2023. In this talk we present the design and first results on subsystem tests of the TRIPLE-IceCraft.

EP 4.5 Tue 15:15 EP-H1

TRIPLE-IceCraft: A retrievable melting probe for transporting scientific payloads through glacial ice — •SIMON ZIERKE¹, DIRK HEINEN¹, JAN AUDEHM¹ MIA GIANG DO¹, YUTING YE¹, CHRISTOPHER WIEBUSCH¹, MARCO FELDMANN², GERO FRANCKE², and CLEMENS $ESPE^2 - {}^1RWTH$ Aachen University - Physics Institute III B, Aachen, Germany $- {}^2GSI$ - Gesellschaft für Systementwicklung und Instrumentierung mbH, Aachen, Germany

The exploration of subglacial worlds is one of the greatest technological challenges for both space science and terrestrial glaciology. Within TRIPLE, initiated by the German Space Agency at DLR, technologies are being developed to address these challenges. One of the projects within TRIPLE is TRIPLE-IceCraft. Its goal is the development of an electrothermal drill as a terrestrial demonstrator for transporting payloads through an ice sheet of several hundred meters at a drilling velocity of several meters per hour. The same-named drill, TRIPLE-IceCraft, is a modular system including a cable which can be coiled and uncoiled into a dedicated compartment. The cable bears the weight of the drill and is used for communication and power. This allows the refreezing of the melt hole including the cable, so it can be operated even in cold glacial ice. The demonstration of the TRIPLE-IceCraft is planned close to the Antarctic research station Neumayer III located on the Ekström Ice Shelf in 2023. In this poster, we focus on the technical design of the TRIPLE-IceCraft.

EP 4.6 Tue 15:25 EP-H1

The In-Ice Sonar System for the TRIPLE Forefield Reconnaissance System -•JAN AUDEHM, BEN BURGMANN, MIA GIANG DO, SHARIF EL MENTAWI, DIRK Heinen, Johanna Hermannsgabner, Andreas Nöll, Shreyans Sakhare, CHRISTOPHER WIEBUSCH, YUTING YE, and SIMON ZIERKE - RWTH Aachen University - Physics Institute III B, Aachen, Germany

The icy moons Europa and Enceladus belong to the most interesting sites for the search of extra-terrestrial life in the solar system. It is assumed that in the oceans beneath their thick ice crusts preconditions for the emergence of life are fulfilled.

In the TRIPLE-project Technologies for Rapid Ice Penetration and subglacial Lake Exploration are developed to enable future exploration missions to the subsurface oceans of icy moons. These technologies will be demonstrated in a terrestrial analog scenario in Antarctica. Here, a melting probe will transport a small autonomous underwater vehicle (TRIPLE-nanoAUV) through the ice to an underlying water reservoir.

For this mission it is of great importance that the probe avoids obstacles on the path through the ice and can detect the transition between ice and water to anchor there. This task is addressed by TRIPLE-FRS being a Forefield Reconnaissance System using a hybrid system consisting of a radar, a sonar, and a permittivity sensor. Combining the complementary techniques assures a high performance of the FRS in ice as well as in water. This poster presents our concept for the sonar system and simulations of the acoustic transducer characteristics.
EP 5: Near Earth Space

Time: Tuesday 16:15-18:45

Invited Talk

EP 5.1 Tue 16:15 EP-H1

Investigating Earth's atmosphere and ionosphere from space: How GNSS radio occultation measurements contribute to monitor the atmosphere in a high spatial resolution — •CHRISTINA ARRAS¹, ANKUR KEPKAR^{1,2}, and JENS WICKERT^{1,2} — ¹German Research Centre for Geosciences GFZ, Potsdam, Germany — ²Technische Universität Berlin, Germany

The GNSS radio occultation (RO) technique has been established successfully during the previous two decades. It evolved into a valuable observation tool for precise atmospheric and ionospheric vertical profiling. Radio occultation measurements provide globally distributed precise profiles of the refractivity of the Earth's atmosphere that can be converted into profiles of temperature, pressure, and water vapor in the lower neutral atmosphere and into electron density values in the ionosphere. Until today, there are about 14 million RO recordings available.

GNSS RO signals are very sensitive to vertical electron density gradients in the Earth's ionosphere. This becomes visible as strong fluctuations in, e.g., signal-tonoise ratio recordings, which allow detecting ionospheric disturbances like sporadic E layers in the lower ionospheric E region and equatorial plasma bubbles in the F-layer.

In this presentation, we will give an overview on RO data availability. We will review the data analysis to derive information on ionospheric disturbances in the E and F layer. Further, we discuss the sporadic E and plasma bubble formations that result from complex coupling processes in the thermosphere-ionospheremagnetosphere system.

EP 5.2 Tue 16:45 EP-H1

Measurements of cosmic rays by a mini neutron monitor aboard the German research vessel Polarstern — •M. ZOSKA¹, S. BANJAC¹, S. BURMEISTER¹, H. GIESE¹, B. HEBER¹, K. HERBST¹, L. ROMANEESEN¹, C. SCHWERDT², D. STRAUSS³, C. WALLMANN¹, and M. WALTER¹ — ¹Christian-Albrechts-Universität Kiel, Germany — ²Deutsches Elektronen-Synchrotron Zeuthen, Germany — ³North West University, Potchefstroom, South Africa

Galactic cosmic rays (GCRs) are a direct sample of material from far beyond the solar system. Measurements by various particle detectors have shown that the intensity varies on different timescales, caused by the Sun's activity. Many studies on GCR intensity decreases are based on the analysis of ground-based neutron monitors. Their measurements depend on the geomagnetic position, and the processes in the Earth's atmosphere. In order to get a better understanding of the geomagnetic filter over the solar cycle, the above groups agreed on a continous monitoring of the GCR flux as a function of latitude, by installing a portable device aboard the German research vessel Polarstern in 2012. The vessel is ideally suited for this research campaign because it covers extensive geomagnetic latitudes (i.e. goes from the Arctic to the Antarctic) at least once per year. Here we present themeasurements for different latitude surveys including the solar maximum in 2014 and solar minimum in 2019.

The Kiel team received funding from the European Union Horizon 2020 programme under grant agreement No 870405. We thank the crew of the Polarstern and the AWI for supporting our research campaign.

EP 5.3 Tue 17:00 EP-H1

Utilizing Cosmic Ray data as input for neutron-based soil moisture measurement — •HANNA GIESE¹, BERND HEBER¹, KONSTANTIN HERBST¹, and MARTIN SCHRÖN² — ¹Institut für Experimentelle und Angewandte Physik, Christian-Albrechts-Universität zu Kiel, Kiel, Germany — ²UFZ Helmholtz Center for Environmental Research, Leipzig, Germany

Neutrons on Earth interact with the soil and are substantially moderated by hydrogen atoms. Since the reflected neutron flux is a function of the soil water content, cosmic-ray neutron measurements above the ground can be used to estimate the average field soil moisture. Thus, if the local incoming neutron flux and the abundance of nearby hydrogen pools are known, the reflected neutron flux could be modeled and compared to observed detector count rates. However, the incoming neutrons are secondaries produced by interacting energetic Galactic Cosmic Rays (GCRs) in the atmosphere. The total neutron flux on the ground depends on the solar modulation-dependent GCR flux, the geomagnetic position, and the altitude within the atmosphere. So far, measurements of either the Jungfraujoch neutron monitor (NM) or a NM of similar cutoff rigidity have been used and altered to estimate the neutron flux at the position of each neutron detector. In this contribution we present a new method based on the Dorman function to directly compute the local neutron flux using remote neutron monitor data.

We received funding from the European Union's Horizon 2020 research and innovation programme under grant agreement No. 870405

Location: EP-H1

EP 5.4 Tue 17:15 EP-H1

Empirical modelling of SSUSI-derived auroral ionization rates — •STEFAN BENDER^{1,2}, PATRICK ESPY^{1,2}, and LARRY PAXTON³ — ¹Norwegian University of Science and Technology, Trondheim, Norway — ²Birkeland Centre for Space Science, Bergen, Norway — ³APL, Johns Hopkins University, Laurel, Maryland, USA

Solar, auroral, and radiation belt electrons enter the atmosphere at polar regions leading to ionization and affecting its chemistry. Climate models usually parametrize this ionization and the related changes in chemistry based on satellite particle measurements. Precise measurements of the particle and energy influx into the upper atmosphere are difficult because they vary substantially in location and time. Widely used particle data are derived from the POES and GOES satellite measurements which provide in-situ electron and proton spectra.

Here we use the electron energy and flux data products from the Special Sensor Ultraviolet Spectrographic Imager (SSUSI) instruments on board the Defense Meteorological Satellite Program (DMSP) satellites. The currently three operating satellites directly observe the auroral emissions in the UV from which electron energies and fluxes are inferred in the range from 2 keV to 20 keV. We use these observed electron energies and fluxes to calculate auroral ionization rates in the lower thermosphere (90–150 km). We present an empirical model of these ionization rates according to magnetic local time and geomagnetic latitude. The model is based on geomagnetic and solar flux indices and will be particularly targeted for use in climate models that include the upper atmosphere, such as WACCM-X or EDITh.

Invited Talk

EP 5.5 Tue 17:30 EP-H1 niddle atmosphere — •THOMAS

Solar extreme events and their impact on the middle atmosphere — •Thomas Reddmann¹, Monali Borthakur¹, Miriam Sinnhuber¹, Ilya Usoskin², and Jan-Maik Wissing³ — ¹Karlsruhe Institute of Technology, Germany — ²University of Oulu, Finland — ³University of Rostock, Germany

The Sun occasionally produces strong eruptions on its surface and in the corona. Within these strong eruptions, energetic particles are accelerated to high energies which hit the Earth within hours after the event. In addition, ejected plasma clouds can be accelerated towards the Earth causing severe geomagnetic disturbances which also cause particle precipitation. Here we study the impact of a one per millenium solar event on the atmosphere.

We use historical records and analyzed distributions of energy spectra to derive ionization rates for a combined extreme solar proton event and a geomagnetic storm which typically impact different parts of the atmosphere. The ionization rates for the extreme event are then used in simulations in the KASIMA and EMAC model which both include energetic particle induced chemistry. Their simplified production efficiency of NOx and HOx is compared to an ion chemistry model. We select specific dynamical situations for the event which represent a different vertical coupling in the atmosphere. Finally, we estimate the impact of the extreme event under those different dynamical situations on the chemical state in the atmosphere on the seasonal time scale in terms of ozone change and the global NOx budget together with the additional UV dose.

EP 5.6 Tue 18:00 EP-H1

Impact of chlorine ion chemistry on the ozone loss during very large solar proton events — •MONALI BORTHAKUR¹, THOMAS REDDMANN¹, MIRIAM SINNHUBER¹, GABRIELLE STILLER¹, THOMAS VON CLARMANN¹, and ILYA USOSKIN² — ¹Karlsruhe Institute of Technology, Karlsruhe, Germany — ²University of Oulu, Oulu, Finland

Strong eruptions on the Sun can accelerate charged particles, mostly protons, to high energies, causing solar proton events (SPEs). Such energetic particles can precipitate upon the Earth's atmosphere, mostly in polar regions because of the geomagnetic shielding. SPE induced chlorine activation and its impact on stratospheric ozone in the polar northern atmosphere has been investigated using a 1D stacked-box model, of atmospheric ion and neutral composition, EXOTIC. Two SPEs were used as test fields: the Halloween SPE late October 2003 and an extreme event of 775 AD, which was derived from historical records of cosmogenic nuclides. We used observations of chlorine species and ozone from the Michelson Interferometer for Passive Atmospheric Sounding (MIPAS) on ENVISAT to evaluate model results for the Halloween event. Model experiments were carried out with full ion chemistry as well as with a common parameterisation scheme considering the formation of HOx and NOx. Additional ozone destruction due to the chlorine catalytic cycles was observed as well in the upper stratosphere and lower mesosphere. For the parameterised model runs, ozone depletion was observed to be the largest for full parameterisation.

EP 5.7 Tue 18:15 EP-H1

Ring current electron precipitation during storm events -• Alina GRISHINA^{1,2}, YURI SHPRITS^{1,2,3}, MICHAEL WUTZIG¹, HAYLEY ALLISON¹, NIKITA Aseev^{1,2}, Dedong Wang¹, and Matyas Szabo-Roberts^{1,2} - ¹GFZ Potsdam, Potsdam, Germany — ²University of Potsdam, Potsdam, Germany — ³University of California, Los Angeles, Los Angeles, CA, USA

The particle flux in the near-Earth environment can increase by orders of magnitude during geomagnetically active periods. This leads to intensification of particle precipitation into Earth's atmosphere. The process potentially further affects atmospheric chemistry and temperature.

In this research, we concentrate on ring current electrons and investigate precipitation mechanisms on a short time scale using a numerical model based on the Fokker-Planck equation. We focus on understanding which kind of geomagnetic storm leads to stronger electron precipitation. For that, we considered two storms, corotating interaction region (CIR) and coronal mass ejection (CME) driven, and quantified impact on ring current. We validated results using observations made by POES satellite mission, low Earth orbiting meteorological

EP 6: Astrophysics

Time: Wednesday 11:00-13:15

Invited Talk

EP 6.1 Wed 11:00 EP-H1 Three-dimensional topology-driven magnetic reconnection - •RAQUEL Mäusle¹, Jean-Mathieu Teissier¹, and Wolf-Christian Müller^{1,2} – ¹Technische Universität Berlin, Berlin, Germany – ²Max-Planck/Princeton Center for Plasma Physics, Princeton, NJ, USA

Magnetic reconnection is a dissipative process, by which the magnetic field structure within a plasma is changed. It is important in many astrophysical systems, such as the Sun's corona, where it is believed to be of importance for the generation of solar flares.

We study a three-dimensional model of magnetic reconnection driven by magnetic field topology. Magnetic field lines in three-dimensional systems have the tendency to become highly entangled, making them exponentially sensitive to very small non-ideal effects. Therefore, entanglement could be the dominant mechanism for fast reconnection in low-resistivity plasmas, requiring far smaller current densities than otherwise needed. We investigate this model numerically using a fourth-order finite-volume scheme to solve the magnetohydrodynamic (MHD) equations. We start from a system with an initially constant magnetic field and line-tied boundaries, and drive it into a chaotic state in which reconnection is occuring. We study the dynamics of this system, the correlation between the field line entanglement and the occurrence of reconnection events, as well as the dependence of the reconnection rate on the resistivity.

In this talk I will introduce the model and numerical method employed and present our preliminary results.

EP 6.2 Wed 11:30 EP-H1

Very high-order subsonic magnetohydrodynamics solvers — •JEAN-MATHIEU TEISSIER¹ and WOLF-CHRISTIAN MÜLLER^{1,2} – ¹Technische Universität Berlin, Berlin, Germany — ²Max-Planck/Princeton Center for Plasma Physics, Princeton, NJ, USA

Magnetohydrodynamics (MHD) solvers are very important tools to analyze the large-scale, long time behaviour of astrophysical plasmas. Direct numerical simulations are however linked with high computational costs, so that a trade-off between accuracy of the results and the available resources has to be made. Higherorder solvers, i.e. solvers with a discretization order strictly higher than two, can typically achieve higher accuracy at a lower resolution than second-order ones, leading to an overall gain in performance. However, with increasing discretization order, solvers based on 3D-reconstructions may become prohibitively expensive. We present a finite-volume dimension-by-dimension method which allows to solve for the MHD equations at arbitrarily high discretization orders (we show results up to order ten), while maintaining affordable numerical costs for 3D problems. The magnetic field solenoidality is preserved up to machine precision with the constrained-transport approach. The study is limited to subsonic systems since shocks are not handled properly by higher-order methods. For a discretization order of six and above, the numerical dissipation is too low to prevent a pile-up of energy at small scales in turbulent systems, so that explicit diffusive terms need to be added. We present a formulation to do so, respecting the finite-volume and the constrained-transport frameworks.

EP 6.3 Wed 11:45 EP-H1

ComPol - A Compton polarimeter in a Nanosat — •MATTHIAS MEIER for the ComPol-Collaboration - Excellence Cluster ORIGINS, Garching, Germany -Technical University of Munich (TUM), Munich, Germany - Max-Planck Institute for Physics (MPP), Munich, Germany

The possibilities to investigate astrophysical compact objects are strongly limited. Due to the small size of these objects it is hardly possible to resolve their

satellites, and Van Allen Probes, and produced a dataset of precipitated fluxes that covers energy range from 1 keV to 1 MeV.

EP 5.8 Tue 18:30 EP-H1

Solar energetic particle event on 28 October 2021 as seen by the neutron monitor network — •CHRISTIAN STEIGIES¹ and ROLF BÜTIKOFER² — ¹Christian-Albrechts-Universität zu Kiel, Germany — ²Universität Bern, Switzerland

The first solar energetic particle event during the current solar cycle 25 observed on ground, a so-called ground level enhancement (GLE), was detected on 28 October 2021 by a few neutron monitors of the worldwide network. At 16:09 UT the Athens GLE Alert system issued successfully an automatic alert. After the initial rapid increase in cosmic ray (CR) intensity, it slowly declined to background level within several hours. During times of raised CR intensities, the ionization rates in the atmosphere are increased and thereby radiation dose rates at aircraft altitudes may be enhanced depending on the location and the altitude. We present a first GLE analysis and assess the additional radiation doses that may be acquired on selected flight routes.

Location: EP-H1

geometry. The CubeSat mission ComPol will investigate the black hole binary system Cygnus X-1. The goal is to improve its physical model by measuring the polarization of the hard X-ray spectrum. The information about the polarization can be extracted from the kinematics of the Compton scattering. A Silicon drift detector (SDD) is used as a scatterer. The SDD is stacked onto a CeBr3 calorimeter to be able to measure the full kinematics.

The talk will give an overview of the scientific motivation, the underlying physics and the detector setup.

This research was supported by the Excellence Cluster ORIGINS which is funded by the Deutsche Forschungsgemeinschaft (DFG, German Research Foundation) under Germany's Excellence Strategy - EXC-2094-390783311

EP 6.4 Wed 12:00 EP-H1

A background study via simulations for the ComPol CubeSat experiment - •CYNTHIA GLAS for the ComPol-Collaboration - Max-Planck-Institut für Physik, München, Deutschland — Technische Universität München, Deutschland

The objective of the CubeSat mission ComPol is to investigate the black hole binary Cygnus X-1 and to improve its physical model by measuring spectrum and polarization in the hard X-ray range. The information about the polarization can be extracted from the kinematics of Compton scattering.

To ensure a correct measurement, all physical processes occurring in the detectors, as well as common background sources need to be understood in advance. Therefore, the response of the setup to cosmic radiation in low earth orbit, as well as the south Atlantic anomaly is simulated with the Monte Carlo toolkit Geant4. Special focus was set on the cosmogenic activation of the CubeSat material. In order to determine which orbits are suitable for the ComPol mission, the activation of each material and the influence of radioactive decays originating from this activation on the individual detectors are studied in detail. To optimize the shielding strategy, simulations with different shielding configurations were performed. The talk will give an overview of the performed simulations, present the conclusions drawn from them, as well as discuss different shielding configurations.

This research was supported by the Excellence Cluster ORIGINS, which is funded by the Deutsche Forschungsgemeinschaft (DFG, German Research Foundation) under Germany*s Excellence Strategy.

EP 6.5 Wed 12:15 EP-H1

ComPol's way to space - preparing a CubeSat Compton Telescope for its in orbit verification (IOV) at the ISS — $\bullet {\rm Katrin}$ Geigenberger for the ComPol-Collaboration - Max-Planck Institute for Physics (MPP), Munich, Germany -Technical University of Munich (TUM), Munich, Germany

ComPol will be a 3U-CubeSat with a Compton polarimeter consisting of a Silicon Drift Detector and a CeBr3 calorimeter. It will perform a long-term measurement of Cygnus X-1 in the hard X-ray range. The goal is to determine the polarization between 20 and 300 keV - a region where the polarization of Cygnus X-1 was hardly observed before. This talk is about the steps towards the first big milestone: a scaled-down version of ComPol's future instrumentation will be operated along with essential parts of the CubeSat bus on an external platform aboard the International Space Station (ISS). This IOV-mission will be a first real life demonstration of the detector system in space environment (background radiation, thermal loads and vacuum). The goal is to receive and process coincident signals from both detectors and to determine the influence of solar, x-ray, and cosmic radiation backgrounds. An overview of the current status of both mechanical and electrical hardware for the very first ComPol prototype will be given with focus on the space-mission specific challenges: e.g. shortage of room, protection of the sensitive detectors from solar radiation, constraints due to vacuum, and the investigation of the behavior at launch loads. This research was supported by the Excellence Cluster ORIGINS which is funded by the Deutsche Forschungsgemeinschaft (DFG).

EP 6.6 Wed 12:30 EP-H1

Stochastic Fluctuations of Cosmic Rays: From Voyager Data to Ionization Rate in Molecular Clouds — •Vo Hong Minh Phan¹, FLORIAN SCHULZE¹, PHILIPP MERTSCH¹, SARAH RECCHIA², and STEFANO GABICI³ — ¹Institute for Theoretical Particle Physics and Cosmology (TTK), RWTH Aachen University, Aachen, Germany — ²Dipartimento di Fisica, Universitá di Torino, Torino, Italy — ³University of Paris, CNRS, Astroparticule et Cosmologie, Paris, France

Data from the Voyager probes have provided us with the first measurement of cosmic ray intensities at MeV energies, an energy range that had previously not been explored. Simple extrapolations of models that fit data at GeV energies, e.g., from AMS-02, however, fail to reproduce the Voyager data in that the predicted intensities are too high. Oftentimes, this discrepancy is addressed by adding a break to the source spectrum or the diffusion coefficient in an ad hoc fashion, with a convincing physical explanation yet to be provided. In this talk, we will show that the discrete nature of cosmic-ray sources, which is usually ignored, is instead a more likely explanation. We model the distribution of intensities expected from a statistical model of discrete sources and show that is expectation value is not representative but has a spectral shape different from that for a typical configuration of sources. The Voyager proton and electron data are however compatible with the median of the intensity distribution. We will also discuss some preliminary results concerning the ionization rate induced by low-energy cosmic rays in molecular clouds.

EP 6.7 Wed 12:45 EP-H1

Efficient numerical simulations of dynamical cosmic-ray transport -

•STEFAN KIS¹, PRANAB DEKA², RALF KISSMANN¹, and LUKAS EINKEMMER² — ¹Institut für Astro- und Teilchenphysik, Leopold-Franzens-Universität Innsbruck, A-6020 Innsbruck, Austria — ²Institut für Mathematik, Leopold-Franzens-Universität Innsbruck, A-6020 Innsbruck, Austria

Understanding the transport of charged particles in the Galaxy requires numerically solving a parabolic linear partial differential equation known as the cosmicray transport equation. Various codes aim at doing so, however most rely on a single approach in solving the time dependent problem and some do not provide substantial testing grounds for their solvers. We present a study case for various numerical time integrator schemes employed in solving the dynamical cosmicray transport equation. We asses the stability of the time integrator schemes and compare the numerical output against an analytical solution to determine their accuracy as well as judge their efficiency in the context of cosmic-ray transport.

EP 6.8 Wed 13:00 EP-H1

Relativistic regularized kappa distributions — •LINH HAN THANH — Ruhr-Universität Bochum

Within the framework of relativistic kinetic theory, a special relativistic generalization of isotropic kappa distributions is proposed, based on the requirement that in the non-relativistic limit the original distributions are recovered. By studying the moments, it is found that the relativistic description of the standard kappa distribution leads to an even greater restriction on allowed kappa values than in the non-relativistic case, whereas the relativistic regularized kappa distribution is able to remove all divergences.

EP 7: Astrophysics

Time: Wednesday 16:15-18:50

EP 7.1 Wed 16:15 EP-H1

Comparison of HAWC's Eye data and Monte Carlo simulations — •MARK MEYERS for the HAWC's Eye-Collaboration — Physics Institute III A, RWTH Aachen, Germany

HAWC's Eye is an array of Imaging Air-Cherenkov Telescopes which runs in coincidence with the High altitude water cherenkov observatory (HAWC), forming the first hybrid detector for cherenkov light, observing high energetic gamma rays and cosmic rays in the range of 100 GeV to 100 TeV. The goal is to enhance the performance, by improving the energy calibration and arrival direction reconstruction. In a Monte Carlo simulation, extensive air showers of protons and gammas are observed by an array of 55 telescopes. Properties of the produced detector images in the single telescopes are compared to those that have been produced in measurements, to verify the simulation. These measurements were obtained in stereo observations with two telescopes in three campaigns during the years 2019 and 2020. Potential ways to improve the simulation according to the obtained results will be investigated.

EP 7.2 Wed 16:30 EP-H1

Indication for a local source in the Northern Hemisphere from the joined energy spectrum and mass composition fit of the Pierre Auger Observatory and Telescope Array experiment — \bullet PAVLO PLOTKO¹, ARJEN VAN VLIET¹, XAVIER RODRIGUES¹, DOMENIK EHLERT², and WALTER WINTER¹ — ¹DESY, Platane-nallee 6, 15738 Zeuthen, Germany — ²Institutt for fysikk, NTNU, Trondheim, Norway

The Pierre Auger Observatory (PAO) and Telescope Array (TA) energy spectrum working groups report significant differences for the energy spectra above $10^{19.5}$ eV for the full fields of view of both experiments. In this work, we present a joined fit to the energy spectrum and mass composition measured by TA and PAO. Our fitting procedure includes the systematic uncertainty for both experiments. We carry out a detailed 3D model of UHECR sources that depends on the cosmological source evolution and a universal spectral power-law index and maximal rigidity. UHECR propagation is simulated within the "Propagation including Nuclear Cascade equations" (PriNCe) framework. Our fit results suggest that a local source in the Northern Hemisphere, that is only visible by the TA experiment, can reconcile PAO and TA measurements. For this proposed local source we discuss two possibilities: First, an object located at a few hundred Mpc with a heavy composition (corresponding to our best fit) and second a source at an even closer distance with a lighter composition (within the 1σ interval of our fit). Finally, we discuss possible source candidates taking into account the latest results of the TA hotspots.

Location: EP-H1

EP 7.3 Wed 16:45 EP-H1

Measurement of the mass composition of ultra-high energy cosmic rays at the Pierre Auger Observatory using a novel model based on air-shower universality — MARKUS ROTH, DAVID SCHMIDT, •MAXIMILIAN STADELMAIER, and DARKO VEBERIC — KIT, Karlsurhe, Germany

We present a model of extensive air showers that is based on the implications of air-shower universality. The model comprises the spatial and temporal densities of particles expected from extensive air-showers induced by cosmic rays. Therewith, the depth of the shower maximum and the relative muon content of an air shower can be reconstructed solely from data collected by the surface detector of the Pierre Auger Observatory and the upgraded AugerPrime. Using these two observables, the logarithmic atomic mass of cosmic rays can be reconstructed on an event-by-event basis with sufficient accuracy to discriminate light from heavy primary particles. The method is calibrated using hybrid measurements from the surface detector and the fluorescence telescopes of Auger. Furthermore, we present results on the mass composition of ultra high-energy cosmic rays as estimated using the universality-based reconstruction.

EP 7.4 Wed 17:00 EP-H1

Analytic examination of AGN variability in a two-zone model — •VITO ABERHAM and FELIX SPANIER — Institut für Theoretische Astrophysik, Albert-Ueberle-Straße 2, 69120 Heidelberg

The variability of active galactic nuclei is examined analytically using a two-zone model that injects particles in a monoenergetic fashion. The elemental PDEs of both zones are solved neglecting terms behaving as second-order Fermi acceleration. The time-dependent behavior of the obtained particle distributions is inspected by employing a variable source function. As time approaches infinity the steady state distribution functions are recovered displaying the expected spectral indices. Being typical for BL Lac objects parameter estimates from a model fitting the SED of the blazar TXS 0506+056 are incorporated. Corresponding photon and neutrino fluxes in both local jet and earth frame are determined and compared to observations. The general time-dependence of photon light curves in various bands and the neutrino flux is evaluated by implementing a timedependent initial source function. The variability timescales of relevant quantities are ultimately inferred also considering the effect of different particle species. It is demonstrated how the analytical results can be utilized to cross-check results from numerical simulation as well as to interpret observational data collected at different times.

EP 7.5 Wed 17:15 EP-H1

ExHaLe-jet: An extended hadro-leptonic jet model for blazars — •MICHAEL ZACHARIAS^{1,2}, ANITA REIMER³, CATHERINE BOISSON¹, and ANDREAS ZECH¹ — ¹LUTH, Observatoire de Paris, Meudon, France — ²Centre for Space Research, North-West University, Potchefstroom, South Africa — ³Institut für Astro- und Teilchenphysik, Universität Innsbruck, Austria

Blazars – active galaxies with the jet pointing at Earth – emit across all electromagnetic wavelengths. The so-called tne-zone model has described well both quiescent and flaring states, however it cannot explain the radio emission. In order to self-consistently describe the entire electromagnetic spectrum, extended jet models are necessary. Notably, kinetic descriptions of extended jets can provide the temporal and spatial evolution of the particle species and the full electromagnetic output. Here, we present the initial results of a recently developed lepto-hadronic, extended-jet code. As protons take much longer than electrons to lose their energy, they can transport energy over much larger distances than electrons and are therefore essential for the energy transport in the jet. Furthermore, protons can inject additional leptons through pion production and decay, as well as Bethe-Heitler pair production, which can explain a dominant leptonic radiation signal while still producing neutrinos. We will present a detailed parameter study and provide insights into the different blazar sub-classes.

EP 7.6 Wed 17:30 EP-H1 Modeling the leptonic origin of the low-frequency emission from blazar PKS

1502+106 ••FREDERIKE APEL — Ruhr-Universität Bochum In this work, we model the multi-wavelength emission from the blazar PKS 1502+106. This object is a candidate source of a high-energy neutrino observed by the IceCube experiment in 2019. We show that the emission from the source in the range from radio to infrared can be well explained by radiative (synchrotron) cooling of electrons accelerated in the relativistic jet. First, we show how simple analytical considerations can be used to obtain a rough estimate of the physical parameters of the source. Then, using time-dependent numerical simulations, we obtain a more accurate result that reproduces the observed spectral energy distribution. This result sheds light on the physical properties of the source, and can be used in the future to better constrain its nature as a hadronic accelerator.

EP 7.7 Wed 17:45 EP-H1

Active galactic nucleis have been discussed as possible accelerators of highenergy cosmic rays for quite some time. When the Blazar TXS 0506+056 was identified as the source of the high-energy Muon neutrino (IceCube-170922A) detected by the IceCube telescope in 2017 as a result of a large-scale multimessenger campaign, it was a first indication of possible correlations of the increased spectral activity of such sources and their neutrino production. Studies of this correlation by simulating the acceleration processes taking place in the jet and their photon and neutrino emission therefore allow conclusions to be drawn about the composition of the jet plasma by comparing them with the observed fluxes. In this talk, the two-zone model UNICORN-0d is used to perform a selfconsistent modeling of the eruption of TXS 0506+056 in 2017 based on the existing multimessenger data. A special focus will be put upon the data in the very high energy regime. The importance of these data being taken simultaneously is discussed by using time-resolved simulations of an emission flare. In contrast to previous models this approach addresses the possibility that the data from 2017 might represent different emission-states of the Blazar.

EP 7.8 Wed 18:00 EP-H1

The candidate tidal disruption event AT2019fdr coincident with a highenergy neutrino — •SIMEON REUSCH — Deutsches Elektronen Synchrotron DESY, Platanenallee 6, D-15738 Zeuthen, Germany — Institut für Physik, Humboldt-Universität zu Berlin, D-12489 Berlin, Germany

The origins of the high-energy cosmic neutrino flux remain largely unknown.

Last year, a high-energy neutrino was associated with the tidal disruption event (TDE) AT2019dsg.

In this talk we present AT2019fdr, an exceptionally luminous TDE candidate, coincident with another high-energy neutrino detected by IceCube. We will present observations that further support a TDE origin of this flare. These include a bright dust echo and soft late-time X-ray emission. The probability of finding two such bright events in neutrino follow-up by chance is just 0.034%.

Furthermore, we have evaluated several models for neutrino production and show that AT2019fdr is capable of producing the observed high-energy neutrino. This reinforces the case for TDEs as neutrino sources.

EP 7.9 Wed 18:15 EP-H1 Search for short time-scale transients from the Sculptor galaxy — •ANNANAY JAITLY — Humboldt-Universität zu Berlin, Berlin, Germany

Astrophysical sources show variability in their emissions over a range of timescales. For short time-scale transients like fast radio bursts, no very-high-energy gamma-ray counterparts have been detected so far and there is a general lack of tools suited to search for such phenomena. We developed and tested a plugin for the gammapy python package. It is is a tool capable of searching gamma-ray telescope data for transient phenomenon over arbitrary timescales. It scans the given field of view for clusters of events within user-defined time and angular separation intervals. To test the performance of said tool, we studied the Sculptor galaxy (NGC 253); it was chosen because Fermi-LAT previously reported a magnetar giant flare near this source. In this contribution we present the main features of the developed software, and our results from searching the Sculptor galaxy for short-timescale candidates using it.

EP 7.10 Wed 18:30 EP-H1

Understanding the multi-wavelength variability of TeV blazar VER J0521+211 with high-energy particle interactions — •ANASTASIIA OMELIUKH — Ruhr University Bochum, Faculty of Physics and Astronomy, Astronomical Institute, Universitätsstr. 150, 44801 Bochum, Germany

In spring 2020, the MAGIC collaboration detected photons with energies above 200 GeV from the source VER J0521+211 which provides evidence of efficient particle acceleration in a relativistic jet. The monitoring of the source's short-term variability in very-high-energy gamma rays is complemented by simultaneous data from other observatories in radio, optical, X-ray, and GeV gamma rays. We perform multi-wavelength modeling of this source with a fully self-consistent one-zone leptohadronic model to explain four different multi-wavelength data sets. While the radio and optical fluxes seem to originate from electron synchrotron emission, in this model the gamma-ray fluxes are well explained by electromagnetic cascades induced by proton interactions.

EP 7.11 Wed 18:40 EP-H1

Location: T-H15

Gravitation explained as a physical interaction instead of a geometric spacetime model. — •OSVALDO DOMANN — Stephanstr. 42, 85077 Manching

GR is the theory of gravitation of the SM. It is a mathematical approach from 1915 based on geometric reflections, arriving to the wondrous concept of spacetime curvature. GR resists all intents of integration into a unified field theory and is not compatible with quantum mechanics. An approach is presented for a gravitation theory that is based on the representation of a subatomic particle (SP) as a focal point of rays of Fundamental Particles (FPs) that move from infinite to infinite. The energy of a subatomic particle is stored at its FPs as rotation defining angular momenta. With this representation all SPs interact permanently through the angular momenta of their FPs. The approach explains gravitation as the result of the reintegration of migrated electrons and positrons to their nuclei. Gravitation is composed of a Newton and an Ampere component, with the Newton component dominant at sub galactic distances and the Ampere component at galactic distances. A positive Ampere component explains the speed flattening of galaxies and a negative Ampere component the expansion. Neither dark matter nor dark energy is required and the model is compatible with quantum mechanics. More at: www.odomann.com

EP 8: Astroparticles: Invited talks (joint session T/EP)

Time: Thursday 11:00-12:30

Invited Talk EP 8.1 Thu 11:00 T-H15 Borexino looks in the direction of solar neutrinos — •LIVIA LUDHOVA for

the Borexino-Collaboration — Forschungszentrum Jülich, Jülich, Germany — RWTH Aachen University, Aachen, Germany

Borexino is a 280-ton liquid scintillator detector located at the LNGS in Italy. Characterized by an unprecedented radio-purity, it has succeeded in providing several milestone measurements of MeV-scale neutrinos, with the main focus on solar neutrinos. The latter are the only direct probe of the Hydrogen-to-Helium fusion powering our Sun. The European Physical Society awarded the 2021 Giuseppe and Vanna Cocconi Prize to the Borexino Collaboration for the ground-breaking observation of solar neutrinos from the pp chain and CNO cycle that provided unique and comprehensive tests of the Sun as a nuclear fusion engine. Borexino has developed a new method, Correlated and Integrated Directionality (CID), to exploit the sub-dominant directional Cherenkov light in a liquid scintillator detector. This technique can disentangle the solar neutrino signal, correlated with the known position of the Sun, from the isotropic background. In the region of interest dominated by the signal from 0.862 MeV Be-7 solar neutrinos, the no-solar neutrino hypothesis has been excluded with

 $>5\sigma$ C.L. This novel method is readily applicable to next generation experiments. The talk will focus on the recent Borexino solar neutrino results, including the motivation, analysis details, as well as their interpretation.

EP 8.2 Thu 11:30 T-H15 Invited Talk Gravitational waves - a new probe of the early Universe — • VALERIE DOMCKE - CERN, Geneva, Switzerland

Due to their extremely weak interactions with the matter content of the Universe, gravitational waves generated right after the Big Bang can traverse the Universe basically unperturbed, carrying information about their production processes and the expansion history of our Universe. This makes them a unique probe of BSM physcis at very high energies. I will talk about possible next steps in this field, including the search for the stochastic gravitational wave background and new ideas for searching for gravitational waves at ultra-high frequencies.

EP 9: Mitgliederversammlung Extraterrestrische Physik

Invited Talk

HEURS — Leibniz Universität Hannover

Time: Thursday 12:45-13:45

Mitgliederversammlung EP / General assembly EP

EP 10: Astroparticles: Invited topical talks (joint session T/EP)

Time: Thursday 14:00-15:40

Invited Topical Talk

EP 10.1 Thu 14:00 T-H15 LND - A ("Made in Germany") Radiation Monitor Operating at the far Side of the Moon — •Sönke Burmeister¹, Shenyi Zhang², Jia Yu¹, Zigong Xu¹, Stephan Böttcher¹, and Robert Wimmer-Schweingruber¹ — ¹Institut für Experimentelle und Angewandte Physik, Uni Kiel – ²NSSC, Chinese Academy of Science

Space Radiation is one of the major concerns in human space flight. Of course, this also applies to human exploration of the Moon. On the lunar surface, this consists of chronic exposure to galactic cosmic rays and sporadic solar particle events. The interaction of this radiation field with the lunar soil leads to a third component that consists of neutral particles, i.e., neutrons and gamma radiation. Chang'E 4 is the Chinese mission that landed on the far side of the Moon on January 3rd, 2019. It consists of a lander, a rover, and a relay spacecraft. The LND (Lunar Lander Neutrons and Dosimetry) instrument that was built by CAU is located inside the lander under an opening lid. It consists of a stack of ten segmented Si solid-state detectors (SSDs), which form a particle telescope to measure charged particles (electrons from 0.5 MeV to several MeV, protons 8-35 MeV, and heavier nuclei 17-75 MeV/nuc). A special geometrical arrangement allows observations of fast neutrons (and gamma-rays) that are also important for dosimetry purposes. Thermal neutrons are measured by using a very thin Gd conversion foil sandwiched between two SSDs. The Lunar Lander Neutrons and Dosimetry experiment aboard China's Chang'E 4 lander has made the first ever measurements of the radiation exposure to both charged and neutral particles on the lunar surface.

Invited Topical Talk EP 10.2 Thu 14:25 T-H15 Energetic Pulsar Environments and the Origins of Galactic Cosmic Rays -•ALISON MITCHELL — Erlangen Centre for Astroparticle Physics, FAU, Erlangen, Germany

Cosmic Rays - and their origins - have fascinated Physicists for over a hundred years. Within our Milky Way Galaxy, particles are known to reach energies beyond the so-called Cosmic Ray *knee*, a spectral break at ~3 PeV in the all particle cosmic ray spectrum. However, evidence for the particle accelerators reach PeV energies - PeVatrons - has proven elusive. Only within the last five years have astrophysical sources of gamma-rays above 100 TeV been identified; gamma-rays produced through interactions of particles with PeV energies. Many of these sources are associated with known energetic pulsars.

In this talk, I will review the current census of PeVatrons and discuss implications for our understanding of pulsar environments. There are several open

questions to grapple with: Which particle species are being accelerated - leptonic or hadronic? How are the particles transported through the surrounding medium? What is the maximum energy limit for particle acceleration in pulsar environments? In the near future, data from current and forthcoming facilities will help us to address these questions.

EP 10.3 Thu 14:50 T-H15 **Invited Topical Talk**

Looking forward to exciting physics with FASER — •FELIX KLING — DESY Physics searches and measurements at high-energy collider experiments traditionally focus on the high-pT region. However, if particles are light and weaklycoupled, this focus may be completely misguided: light particles are typically highly collimated around the beam line, allowing sensitive searches with small detectors, and even extremely weakly-coupled particles may be produced in large numbers there. The FASER experiment will use the opportunity and extend the LHC's physic potential by searching for long-lived particles and studying neutrino interactions at TeV energies. In this talk, I will present the physics potential of FASER for new physics searches, neutrino physics and QCD and astro-particle physics.

Invited Topical Talk EP 10.4 Thu 15:15 T-H15 Astroparticle physics at the LHC: Exploring the forward region with crosssection measurements — •HANS DEMBINSKI — Fakultät Physik, Technische Universität Dortmund, Dortmund, Germany

Astroparticle physics is the study of the non-thermal universe with gamma rays, neutrinos, and cosmic rays. Cosmic rays are abundantly produced in cosmic accelerators, like supernova remnants. Some gamma rays and neutrinos are produced indirectly in interactions of cosmic rays with matter in the source, and cosmic rays interact with Earth's atmosphere to produce air showers, which are observed by ground-based cosmic ray observatories and contribute the main background to gamma ray and neutrino observatories. QCD cross-sections for the forward production of hadrons with light and heavy flavor are therefore needed to interpret astroparticle measurements. The experiments at the Large Hadron Collider (LHC) have powerful instruments to measure forward production, but data are more sparse compared to central production. I will summarize the stateof-the-art of forward cross-section measurements at the LHC from the point of view of astroparticle physics and give an outlook into the opportunities in near future with the upcoming run of the LHC and the planned pilot run with oxygen beams.

EP 11: Astroparticles: From the source to the detector (joint session EP/T)

Time: Thursday 16:15-18:30

EP 11.1 Thu 16:15 EP-H1

Multi-messenger studies with gravitational waves and neutrinos - •Тіята MUKHERJEE for the IceCube-Collaboration — Institute for Astroparticle Physics (IAP), Karlsruhe Institute of Technology (KIT), Karlsruhe, Germany IceCube is a neutrino observatory located in Antarctica. Since its discovery of a high-energy neutrino (IC170922A) from the blazar TXS0506+056 in 2017, neutrino astronomy has been established as a viable option to probe the high-energy Universe. Neutrinos can carry undistorted information about their respective astrophysical sources, thus can serve as a cosmic 'messenger' to us. There are other potential messengers as well, e.g. gravitational waves (GW) and cosmic rays other than the traditional photons of various wavelengths. Combining interesting signals of such messengers available from different observatories leads

Since the first direct detection in 2015, gravitational wave signals have been enriching the field of multi-messenger astronomy with insights into formerly "invisible" regimes of the universe. Despite their mind-boggling sensitivities, the current (second) generation of ground-based gravitational wave detectors are limited by various noise sources in their detection band, in particular quantum noise, thermal noise, and seismic noise. Next-generation detectors (e.g. Einstein

Gravitational wave detectors - current and future challenges - •MICHÈLE

Telescope, Cosmic Explorer) aim for sensitivities one or two orders of magnitude better even, making innovative techniques for noise reduction or mitigation a requirement. This talk will present challenges and technical developments on the road to ever-higher gravitational wave event detection rates.

EP 8.3 Thu 12:00 T-H15

Location: EP-MV

Location: T-H15

us towards multi-messenger searches, allowing us to address many of the so far unanswered questions about the fundamentals of this Universe, such as the origin of ultra-high-energy cosmic ray sources. So far, we have the knowledge of detecting electromagnetic signal in multiple wavelengths, spatially and temporally correlated with GW and high-energy neutrinos, as two separate events. However, there is still a missing link as we have not been able to correlate GW with neutrino signals. The aim of my work is to contribute in this aspect, searching for Virgo detected GW counterparts of neutrino events detected by IceCube, including low-energy neutrinos as well as sub-threshold GW events in our analysis. The work plan and initial results will be discussed in this talk.

EP 11.2 Thu 16:30 EP-H1

Multi-messenger characterization of Mrk501 during historically low Xray and gamma-ray activity — •Lea HECKMANN¹, DAVID PANEQUE¹, SAR-GIS GASPARYAN², MATTEO CERRUTI³, NAREK SAHAKYAN², and AXEL ARBET-ENGELS¹ for the Multi-wavelength collaborators and the MAGIC and Fermi-LAT-Collaboration — ¹Max-Planck-Institut für Physik, D-80805 München, Germany — ²ICRANet-Armenia, Marshall Baghramian Avenue 24a, Yerevan 0019, Armenia — ³Institut de Ciències del Cosmos (ICCUB), Universitat de Barcelona (IEEC-UB), Martí i Franquès 1, 8 E08028 Barcelona, Spain

Blazars are the most numerous very-high-energy (>0.2 TeV, VHE) gamma-ray emitters, due to their continuous and very luminous emission; but they are far from being understood.

In this contribution we describe the multi-wavelength behavior of Mrk501, one of our closest and therefore brightest blazars, from 2017 to 2020. Alongside the dense monitoring campaign over this four years, three long observations with NuSTAR were conducted displaying various low-activity flux levels for the hard X-ray emission. This very comprehensive data set reveals a historically low X-ray and VHE gamma-ray emission period lasting two years. Using the low-activity broadband spectral energy distribution (SED) and data published by IceCube, we investigate the nature of the low state. Additionally, we try to explain the evolution of the broadband SED data during the low state to evaluate its potential of being the baseline emission of Mrk501 that is usually outshone by more dominant and variable components.

EP 11.3 Thu 16:45 EP-H1

Hadronic models of active galaxies to constrain cosmic-ray acceleration — •XAVIER RODRIGUES — DESY Zeuthen

In a new era of multi-messenger observatories, numerical models can help shed light on what are the sources of the astrophysical neutrinos and the ultra-highenergy cosmic rays. In this talk I discuss recent results on active galactic nuclei (AGN) as multi-messenger sources, based on numerical simulations of photonuclear cosmic-ray interactions. Assuming AGN jets can reaccelerate cosmic rays up to the EeV regime, I will show that an AGN population may in fact dominate the observed flux and chemical composition of ultra-high-energy cosmic rays. Under certain conditions, the accompanying neutrino flux may be observable by future EeV neutrino telescopes, while respecting the current IceCube limits at PeV energies.

EP 11.4 Thu 17:00 EP-H1

Neutrino Emission during Supermassive and stellar mass Binary Black Hole Mergers — •ILJA JAROSCHEWSKI¹, JULIA BECKER TJUS¹, and PETER L. BIERMANN^{2,3} — ¹Theoretische Physik IV, Ruhr-Universität Bochum — ²MPI for Radioastr., Bonn — ³Dept. of Phys. & Astron., Univ. Alabama, Tuscaloosa, AL, USA

Ever since the discovery of a diffuse astrophysical neutrino flux by IceCube, the question arose which sources contribute most. With several neutrino-blazar associations since the first high-probability association of such a neutrino to the blazar TXS 0506+056 in 2017, there is an indication that at least a non-negligible part of this diffuse neutrino flux emerges from blazars.

As over ninety stellar mass binary black hole mergers were already detected via gravitational waves (GWs), with more to come, there are strong indications that supermassive black holes (SMBHs) in galaxy centers, and thus blazars, also merge and have had at least one merger in their lifetime. Such a merger is almost always accompanied by a change of observable jet direction, leading to interactions of a preceding jet with surrounding molecular clouds and neutrino productions.

By creating a connection between neutrinos and GWs, we set limits on how much energy can be emitted in form of neutrinos in each merger of binary SMBHs and stellar mass black holes and estimate their contributions to the diffuse neutrino flux that is measured by IceCube. As neutrino production is directly connected to high energy cosmic ray interactions, the contribution of these sources to the cosmic ray injection rate is established.

EP 11.5 Thu 17:15 EP-H1

Search for high-energy neutrinos from blazars with IceCube — •CRISTINA LAGUNAS GUALDA for the IceCube-Collaboration — DESY Zeuthen

The IceCube Neutrino Observatory is the world's largest neutrino telescope, instrumenting one cubic kilometre of Antarctic ice. IceCube started its operation with full configuration in 2011 and a diffuse flux of neutrinos was discovered in In 2018 IceCube reported the first clearly identified observation of an astrophysical high-energy neutrino, IC170922A, in spatial and temporal coincidence with blazar TXS 0506+056. Other examples of coincidences that have been observed with lower significance are, but not limited to, IC190730A with blazar PKS 1506+012 and IC141209A with blazar GB6 J1040+0617. What these have in common is that they involve a blazar and a high-energy neutrino with a high probability of being astrophysical in origin (neutrino alert). These coincidences can be combined to calculate a global p-value by performing a stacking analysis. Here we present the results obtained with the Fourth Catalog of Active Galactic Nuclei detected by Fermi-LAT (4LAC-DR2, for Data Release 2) and neutrinos detected by IceCube between 2011 and 2020 that would have passed the neutrino alert criteria.

EP 11.6 Thu 17:30 EP-H1

Comparison of Models for Predicting Periodic Gamma-Ray & Neutrino Emissions From Blazars — •ARMIN GHORBANIETEMAD, ILJA JAROSCHEWSKI, and JULIA BECKER TJUS — Theoretische Physik IV, Ruhr-Unviersität Bochum There are several indications that electromagnetic emissions from blazars have quasi-periodic variability, ranging from minutes to years. The long-term periodicity in the span of years is particularly evident in gamma-ray observations with the Fermi LAT instrument. Two separate high probability associations of neutrinos, detected by IceCube in 2014/15 and 2017, to the blazar TXS0506+056 further indicate that blazars are neutrino emitters. These two flares can be interpreted as a possible periodicity. It is the aim of this work to develop a general set of models that can explain the periodic gamma-ray and neutrino emissions from blazars.

In this talk, we present models with single supermassive black holes as well as supermassive binary black hole mergers at the centers of blazers. Our focus lies on supermassive binary black hole mergers, due to them radiating gravitational waves which could be detectable by the Laser-interferometer Space Antenna (LISA). The binary systems are characterized by the change of jet direction accompanied by jet precession close to an imminent merger. This allows predictions of possible neutrino and gravitational wave emissions from blazars with quasi-periodic behavior.

EP 11.7 Thu 17:45 EP-H1

First science results from the X-ray telescope STIX on Solar Orbiter — •Alexander Warmuth, Frederic Schuller, and Gottfried Mann — Leibniz -Institut für Astrophysik Potsdam (AIP)

The ESA mission Solar Orbiter was successfully launched in 2020, with the main goal of improving our understanding of how the Sun creates and controls the heliosphere. The Spectrometer/Telescope for Imaging X-rays (STIX) is one of six remote-sensing instruments on board and provides imaging spectroscopy of solar flares in the energy range of 4 to 150 keV. Thus, STIX is able to measure quantitatively both the parameters of the hot flare plasma and the characteristics of the accelerated electrons. Together with the other instruments on Solar Orbiter as well as with other space-borne and ground based observational assets, STIX studies energy release and particle acceleration in solar flares. This talk will be focused on the first science results obtained during the cruise phase of Solar Orbiter (2020 and 2021). This includes observations of microflares, constraints on flare energetics, collaborative studies of gamma-ray flares together with Fermi, and the investigation of flare-associated solar energetic particle events.

EP 11.8 Thu 18:00 EP-H1

Unfolding the muon neutrino energy spectrum from 10 years of IceCube data with DSEA+ — •LEONORA KARDUM, KAROLIN HYMON, JOHANNES WERTHE-BACH, PASCAL GUTJAHR, TIM RUHE, and JEAN-MARCO ALAMEDDINE for the IceCube-Collaboration — Astroparticle Physics WG Rhode, TU Dortmund University, Germany

Neutrinos, the most elusive particles in the Standard Model, can travel tremendous distances unaffected by magnetic fields or encountered particles from distant sources in the Universe. As this makes them perfect information carriers, many attempts at uncovering their properties are made. The IceCube Neutrino Observatory, a cubic kilometer detector embedded in the South Pole ice, is capable of detecting neutrinos from several GeV up to PeV energies enabling precise reconstruction of the neutrino spectrum. Determining the accurate spectrum is of great importance to neutrino physics, especially in differentiating the three predicted components - prompt, conventional, and astrophysical, of which only the latter two have been detected so far. The Dortmund Spectrum Estimation Algorithm (DSEA+) is a novel approach to unfolding the energy spectrum from measured experimental quantities that effectively translates ill-posed problems to multinomial classification solvable using readily available machine learning tools. The current status of applying DSEA+ on 10 years of IceCube data will be presented.

EP 11.9 Thu 18:15 EP-H1

Recent solar and geoneutrino results from Borexino — •SINDHUJHA KU-MARAN for the Borexino-Collaboration — Forschungszentrum Jülich - Institute for Nuclear Physics, IKP-2, Jülich, Germany — III. Physikalisches Institut B, RWTH Aachen University, Aachen, Germany

Borexino is a 280-ton liquid scintillator experiment which ran from May 2007 until October 2021 at the Laboratori Nazionali del Gran Sasso (LNGS), Italy. The main goals of Borexino include the measurement of solar neutrinos and geoneutrinos. The extreme radiopurity and thermal stability of the detector have proven to be valuable assets in achieving these goals. Borexino has not only performed a complete spectroscopy of the dominant pp-chain solar neutrinos but has also

EP 12: Sun and Heliosphere

Invited Talk

Time: Friday 11:00-14:00

Invited Talk EP 12.1 Fri 11:00 EP-H1 Linking Solar Eruptions and Energetic Particles through Observations and Modeling — •FREDERIC EFFENBERGER — Ruhr-Universität Bochum, Bochum, Germany

Cosmic rays and energetic particles constitute one of the fundamental components of space plasmas and our Heliospheric environment. However, the relation between different energetic particle populations accelerated in the solar atmosphere and detected in interplanetary space is not well established. Observational studies during the last years demonstrated the still poorly understood existence of a connection between solar flare signatures of accelerated particles at the Sun and the corresponding solar energetic particles (SEPs) detected at 1 AU. It is thus important to make progress towards answering the question: Under which circumstances do these two observations point to the same population of accelerated particles? Here, we will discuss recent progress concerned with this issue. We illustrate the potential for observations and simultaneous modeling of the escaping and precipitating electron populations to constrain the plasma properties of the flaring region and interplanetary medium. In particular, with the recently launched Parker Solar Probe and Solar Orbiter missions, which explore the Sun from a close distance and with unprecedented detail, new insights into these questions can be expected. We emphasize the importance of such studies for the fundamental understanding of physical processes in space plasmas and for our space weather forecasting capabilities.

EP 12.2 Fri 11:30 EP-H1

Determining Pitch-Angle Diffusion Coefficients for Electrons in Whistler Turbulence — •FELIX SPANIER¹, CEDRIC SCHREINER^{2,3}, and REINHARD SCHLICKEISER^{4,5} — ¹Institut für Theoretische Astrophysik, Universität Heidelberg, Albert-Ueberle-Strasse 2, 69120 Heidelberg — ²Centre for Space Research, Northwest-University, Potchefstroom 2520, South Africa — ³Max-Planck-Institute for Solar System Research, Justus-von-Liebig-Weg 3, 37077 Göttingen, — ⁴Institut für Theoretische Physik IV, Ruhr-Universität Bochum, Universitätsstrasse 150, 44801 Bochum, Germany — ⁵Institut für Theoretische Physik und Astrophysik, Christian-Albrechts-Universität zu Kiel, Leibnizstr. 15, 24118 Kiel

Transport of energetic electrons in the heliosphere is governed by resonant interaction with plasma waves, for electrons with sub-GeV kinetic energies specifically with dispersive modes in the whistler regime.

We have performed Particle-in-Cell simulations of kinetic turbulence using parameters similar to those found in the heliosphere. Test-particle electrons are injected into the simulation. The pitch-angle diffusion coefficients of these test particles were analyzed using a novel method.

An analytical model for electron transport in left- and right-handed is derived and the numerical results are compared to this model.

EP 12.3 Fri 11:45 EP-H1

On Compressible Turbulence in the Inner Heliosheath — •HORST FICHTNER¹, JENS KLEIMANN¹, PETER YOON², KLAUS SCHERER¹, SEAN OUGHTON³, and EU-GENE ENGELBRECHT⁴ — ¹Institut für Theoretische Physik IV, Ruhr-Universität Bochum, 44780 Bochum, Germany — ²Institute for Physical Science and Technology, University of Maryland, College Park, USA — ³Department of Mathematics and Statistics, University of Waikato, Hamilton 3240, New Zealand — ⁴Centre for Space Research, North-West University, Potchefstroom, 2522, South Africa

Measurements made with the Voyager 1 spacecraft indicate that significant levels of compressive fluctuations exist in the so-called inner heliosheath, i.e. the region between the solar wind termination shock and the heliopause. Here we extend previous studies of the mirror-mode instability to the whole inner heliosheath. Employing quasilinear theory combined with results from a global magnetohydrodynamic model of the heliosphere allows for a computation of the time evolution of the temperature anisotropy and the energy density of the magnetic fluctuations related to the mirror mode. We demonstrate the likely provided the first direct experimental evidence of the rare CNO-cycle neutrinos. These measurements have several implications for solar and stellar Physics and further improvements are envisioned using the full dataset. In addition, it has recently presented the first directional measurement of sub-MeV solar neutrinos using the sub-dominant Cherenkov light, through a novel technique called Correlated and Integrated Directionality (CID). This method can be further combined with a typical spectral fit and can also prove valuable for next-generation liquid scintillator detectors. The latest geoneutrino measurement from Borexino included a substantial improvement in the precision as well as the rejection of the no-mantle signal with a high significance. This group report will summarize all these recent solar and geoneutrino results of Borexino.

Location: EP-H1

presence of the latter in the inner heliosheath. Furthermore, we compute the associated, locally generated density fluctuations. The results can serve as inputs for future models of the transport of compressible fluctuations in this outermost region of the heliosphere.

EP 12.4 Fri 12:00 EP-H1

Solar Orbiter: two years of operations and first results — •FREDERIC SCHULLER, ALEXANDER WARMUTH, and GOTTFRIED MANN — Leibniz Institute for Astrophysics (AIP) Potsdam, Germany

The Solar Orbiter spacecraft was launched in February 2020 and will remain the most significant observatory for solar physics research in the next decade. During the cruise phase, which ended in November 2021, several crucial activities took place, starting from the commissioning of the various instruments until the testing and validation of all possible observing modes. We will provide an overview of the mission and describe the regular operations, focussing on the day-to-day work of the STIX instrument team. Then, we will highlight some technical improvements that were achieved during the early phase of the mission. Finally, we will briefly present initial scientific results obtained so far, whereas the nominal phase has only just started.

EP 12.5 Fri 12:30 EP-H1

MHD avalanches in truly curved coronal arcades: proliferation and heating — •JACK REID¹, JAMES THRELFALL², and ALAN W. HOOD¹ — ¹University of St Andrews, St Andrews, Fife, United Kingdom — ²Abertay University, Dundee, United Kingdom

MHD avalanches involve small, narrowly localized instabilities spreading across neighbouring areas in a magnetic field. Cumulatively, many small events release vast amounts of stored energy. Straight cylindrical flux tubes are easily modelled, between two parallel planes, and can support such an avalanche: one unstable flux tube causes instability to proliferate, via magnetic reconnection, and an ongoing chain of like events. True coronal loops, however, visibly curve between footpoints on the same solar surface. With 3D MHD simulations, we verify the viability of MHD avalanches in the realistic, curved geometry of an arcade. MHD avalanches thus amplify instability in strong astrophysical magnetic fields and disturb wide regions of plasma. Contrasting with the behaviour of straight cylindrical models, a modified ideal MHD kink mode occurs, more readily and preferentially upwards. Instability spreads over a region far wider than the original flux tubes and their footpoints. Sustained heating is produced in a series of 'nanoflares', collectively contributing substantially to coronal heating. Overwhelmingly, viscous heating dominates, generated in shocks and jets produced by individual small events. Reconnection is not the greatest contributor to heating, but rather facilitates those processes that are. Localized and impulsive, heating shows no strong spatial preference, except a modest bias away from footpoints, towards the apex.

EP 12.6 Fri 12:45 EP-H1

Quasi-discontinuous solar wind models – •LUKAS WESTRICH – Institut für theoretische Physik IV, Ruhr-Universität Bochum, Deutschland

Recently Shergelashvili et al. (2020) developed a new class of discontinuous solar wind solutions. They considered a case of quasi-adiabatic radial expansion with a jump in the flow velocity, density, and temperature but a continuous Mach number at the critical point and derived analytical solutions. Therefore, they proposed a localized external heating source without actual modeling. After a brief discussion of this concept, I will present continuous numerical solutions, more similar to the classical Parker solar wind model, but with quasi-adiabatic radial expansion with an explicitly formulated localized heating source. This kind of solutions can reproduce the analytically derived solutions without discontinuous jumps in the physical properties.

EP 12.7 Fri 13:00 EP-H1

Solar Surface Stereoscopy with Solar Orbiter's Polarimetric Helioseismic Imager — •Amanda Romero Avila, Bernd Inhester, Johann Hirzberger, and Sami Solanki — Max Planck Institute for Solar System Research

A compound method for a stereoscopic analysis of the height variations in the solar photosphere is presented. This method allows to estimate relevant quantities (i.e. the Wilson depression) and to study structures in the solar photosphere and within sunspots. We will demonstrate the feasibility of the method using simulated Stokes I continuum observations derived from a radiative transfer model using the plasma properties of a MHD simulation of the solar surface. The large scale variations in our method are estimated by shifting and correlating two signals of the same region as observed from two different view directions. This result is then introduced as an initial height estimate in a least squares optimization algorithm in order to reproduce smaller scale structures. This method has been developed to be applied to the high resolution images of the PHI instrument on board Solar Orbiter or similar instruments on other Sun-observing spacecraft. It will allow to perform direct stereoscopic studies of solar surface observations in different wavelengths of the solar spectrum. Preliminary results, advantages and limitations, applications and particular considerations for PHI data will be discussed.

EP 12.8 Fri 13:15 EP-H1

A new global nonlinear force-free coronal magnetic-field extrapolation code implemented on a Yin Yang grid — •ARGIRIS KOUMTZIS and THOMAS WIEGEL-MANN — Max Planck Institute for Solar System Research

The solar magnetic field dominates and structures the coronal plasma and detailed insights are important to understand almost all physical processes. While direct routine measurements of the coronal magnetic field are not available, we have to extrapolate the photospheric vector field measurements into the corona. To do so, we developed a new code that performs state-of-the-art nonlinear force-free magnetic field extrapolations in spherical geometry. Our new implementation is based on an optimization principle and is able to reconstruct the magnetic field in the entire corona, including the polar regions. Because of the nature of the finite-difference numerical scheme used in the past, extrapolation close to polar regions was computationally inefficient. In the current code, the so-called Yin Yang grid is used. Both the speed and accuracy of the code is improved compared to previous implementations. We tested our new code with a well known semi-analytical model (Low and Lou solution). This new Yin and Yang implementation is timely because the Solar Orbiter mission is expected to provide reliable vector magnetograms also in the polar regions within the following years. Thus, this code can be used in the future when these synoptic magnetograms are available to model the magnetic field of the solar corona for the entire Sun including the polar regions.

EP 12.9 Fri 13:30 EP-H1

Automatic computation of magneto-hydro-static equilibria from magnetograms and EUV-images — •THOMAS WIEGELMANN and MARIA MADJARSKA — MPS, Göttingen

We present a newly developed tool that models the magnetic field in the solar atmosphere and matches individual field lines with observed structures with enhanced emission in EUV images. Presently, for quiet Sun regions, we can only measure the photospheric line-of-sight magnetic field, as accurate horizontal field measurements are not available. The photospheric magnetic-field measurements are extrapolated into the upper photosphere, chromosphere and corona with a magneto-hydro-static model. Free model parameters are then optimized with a downhill simplex method by comparing magnetic field lines quantitatively with the enhanced emission of various structures recorded in EUV images. The tool can be employed to obtain the magnetic and plasma properties of these structures above the photosphere. This could help to achieve a better understanding of the solar atmosphere and will help the constrain of the modelling of atmospheric structures.

EP 12.10 Fri 13:40 EP-H1

Coronal Magnetic Field Extrapolation Using a Specific Family of Analytical 3D Magnetohydrostatic Equilibria - Practical Aspects — •LILLI NADOL and THOMAS NEUKIRCH — University of St Andrews, Scotland, UK

With current observational methods it is not possible to determine the magnetic field in the solar corona accurately. Hence, coronal magnetic field models have to rely on extrapolation methods using photospheric magnetograms as boundary conditions. In recent years, due to the increased resolution of observations and the need to resolve non-forcefree lower regions of the solar atmosphere, there have been efforts to use magnetohydrostatic (MHS) field models instead of force-free extrapolation methods. Although numerical methods to calculate MHS solutions can deal with non-linear problems more accurately, analytical 3D MHS equilibria can also be used as a numerically "cheaper" method.

We discuss a family of analytical MHS equilibria that allows for a transition from a non-force-free to a force-free region. The solution involves hypergeometric functions. While routines for the calculation of these are available, this can affect the speed and accuracy of the calculations. We look into the asymptotic behaviour of this solution in order to approximate it through exponential functions to improve the numerical efficiency. We present an illustrative example by comparing field line profiles, density and pressure differences between the exact solutions and the asymptotic solution.

EP 12.11 Fri 13:50 EP-H1

Confined and Subsequent Full Flux Rope Eruption as a Model for Homolous Solar Events — Alshaimaa Hassanin¹, •Bernhard Kliem², Norbert Seehafer², and Tibor Török³ — ¹Department of Astronomy, Space Science & Meteorology, Faculty of Science, University of Cairo, Egypt — ²Institute of Physics and Astronomy, University of Potsdam, Germany — ³Predictive Science Inc., San Diego, CA 92121, USA

We present the first numerical model of a sequence of a confined and a full eruption (i.e., a coronal mass ejection, CME). The first eruption results from the helical kink instability of a sufficiently twisted magnetic flux rope; it remains confined because the flux rope does not reach the critical height for onset of the torus instability. A two-step reconnection process reforms a flux rope with subcritical twist near the position of the original flux rope. The full eruption develops as a result of converging motions imposed at the photospheric boundary, which drive flux cancellation. In this process, a part of the positive and negative sunspot flux converge toward the polarity inversion line, reconnect, and cancel each other. Flux of the same amount as the canceled flux transfers to the flux rope, increasing its free magnetic energy. With sustained flux cancellation and the associated progressive weakening of the magnetic tension of the overlying flux, we find that a flux reduction of $\approx 9\%$ leads to the ejective eruption. These results demonstrate that homologous eruptions, eventually leading to a coronal mass ejection, can be driven by flux cancellation.

Gravitation and Relativity Division Fachverband Gravitation und Relativitätstheorie (GR)

Bernd Brügmann Theoretisch-Physikalisches Institut Friedrich-Schiller-Universität Jena 07743 Jena bernd.bruegmann@uni-jena.de

Welcome to the annual meeting of the Gravitation and Relativity Division of the DPG in "virtual" Heidelberg. Despite the adverse circumstances, we again have a rich scientific program, including several genuine highlights. Let me also draw your attention to our General Assembly of Members on Thursday evening, starting at 19:00. I wish all of us a pleasant and informative week!

Overview of Invited Talks and Sessions

(Online lecture halls, in particular GR-H2, GR-H3)

Plenary Talk of GR

PV VII	Thu	9:00- 9:45	Audimax	New perspectives onto the Universe in the era of multi-messenger astrophysics — \bullet SAMAYA <code>NISSANKE</code>

Invited Talks

GR 4.1	Tue	11:00-11:45	GR-H2	Pseudospectrum and black hole quasi-normal mode (in)stability — •Rodrigo Panosso Macedo
GR 4.2	Tue	11:45-12:30	GR-H2	Observable Signatures of Quantized Gravity in Quantum Optical Experiments – •DENNIS Rätzel
GR 7.1	Wed	11:00-11:45	GR-H2	Numerical Relativity and Gravitational Wave Observations — •HARALD PFEIFFER

Invited Talks of the joint symposium SMuK Dissertation Prize 2022 (SYMD)

See SYMD for the full program of the symposium.

SYMD 1.1	Mon	14:00-14:25	Audimax	Timeless Quantum Mechanics and the Early Universe — •LEONARDO CHATAIGNER
SYMD 1.2	Mon	14:25-14:50	Audimax	First tritium β -decay spectrum recorded with Cyclotron Radiation Emission
				Spectroscopy (CRES) — • CHRISTINE CLAESSENS
SYMD 1.3	Mon	14:50-15:15	Audimax	Watching the top quark mass run - for the first time! — • MATTEO M. DEFRANCHIS,
				Katerina Lipka, Sven-Olaf Moch
SYMD 1.4	Mon	15:15-15:40	Audimax	Towards beam-quality-preserving plasma accelerators: On the precision tuning
				of the wakefield — •SARAH SCHRÖDER

Invited Talks of the joint symposium The Nature of Science (SYNS)

See SYNS for the full program of the symposium.

SYNS 1.1	Tue	14:00-14:30	Audimax	The Role of Nature of Science Education for Science Media Literacy $- \bullet$ Dietmar
				Ноттеске
SYNS 1.2	Tue	14:30-15:00	Audimax	What kinds of identities are deemed in/our of place in physics? — •LUCY AVRAAMI-
				DOU
SYNS 1.3	Tue	15:00-15:30	Audimax	Some thoughts on the status of theoretical physics — •DANIEL HARLOW

Prize Talks of the joint Awards Symposium (SYAW)

See SYAW for the full program of the symposium.

SYAW 1.1	Wed	14:10-14:40	Audimax	Wie überprüft man die Ziele der Lehramtsausbildung Physik? - •HORST
				Schecker
SYAW 1.2	Wed	14:40-15:10	Audimax	Astronomy at Highest Angular Resolution - Adaptive Optics, Interferometry and
				Black Holes — •Frank Eisenhauer
SYAW 1.3	Wed	15:10-15:40	Audimax	Turbulence in one dimension — •Alexander M. Polyakov

Sessions

GR 1.1–1.6	Mon	9:00-11:00	GR-H2	Black Holes
GR 2.1–2.8	Mon	16:00-18:40	GR-H2	Classical Theory
GR 3.1–3.6	Mon	16:15-18:15	GR-H3	Cosmology
GR 4.1-4.2	Tue	11:00-12:30	GR-H2	General Relativity
GR 5.1–5.7	Tue	16:15-18:35	GR-H2	Gravitational Waves
GR 6.1–6.6	Tue	16:15-18:15	GR-H3	Cosmology
GR 7.1–7.1	Wed	11:00-11:45	GR-H2	General Relativity
GR 8.1–8.2	Wed	11:45-12:25	GR-H2	Gravitational Waves
GR 9.1–9.7	Wed	16:15-18:35	GR-H2	Gravitational Waves
GR 10.1–10.7	Wed	16:15-18:35	GR-H3	Foundations and Alternatives
GR 11.1–11.5	Thu	11:00-12:40	GR-H2	Relativistic Astrophysics
GR 12.1–12.4	Thu	11:00-12:40	MP-H5	Quantum gravity (joint session MP/GR)
GR 13.1–13.5	Thu	14:00-15:40	GR-H2	Numerical Relativity
GR 14.1–14.6	Thu	16:15-18:15	GR-H2	Gravitational Wave Detectors
GR 15	Thu	19:00-20:30	GR-MV	Member Assembly

Annual General Meeting of the Gravitation and Relativity Division

Thu 19:00-20:30 GR-MV

- Bericht des Vorsitzenden
- Aktuelles und Planung des kommenden Jahres
- Verschiedenes

Location: GR-H2

Sessions

- Invited and Contributed Talks -

GR 1: Black Holes

Time: Monday 9:00-11:00

GR 1.1 Mon 9:00 GR-H2

Photon spheres and shadows of time-dependent black holes – JAY SOLANKI¹ and •VOLKER PERLICK² – ¹Sardar Vallabhbhai National Institute of Technology, Surat, 395007 Gujarat, India – ²ZARM, University of Bremen, 28359 Bremen, Germany

The Vaidya spacetimes are a class of time-dependent and spherically symmetric solutions to Einstein's field equation with a null dust as the source. They describe the gravitational field around a spherically symmetric body that either gains mass by absorbing the null dust or loses mass by emitting it. Here we consider a subclass of the Vaidya spacetimes where the source is a black hole and the metric is conformally static. We demonstrate that there is an unstable photon sphere and we calculate the shadow of such black holes, first as seen by conformally static observers and then as seen by moving observers. The considered spacetimes may be viewed as simple models of black holes that are either accreting or (Hawking) radiating.

GR 1.2 Mon 9:20 GR-H2

Gravitational Lensing in the NUT Spacetime – •TORBEN C. FROST – ZARM, University of Bremen, 28359 Bremen, Germany – Institute for Theoretical Physics, Leibniz University Hannover, 30167 Hannover, Germany

The existence of a gravitomagnetic charge, the gravitational analogon to a hypothetical magnetic charge in electrodynamics, is a long-standing open question in physics. In my talk I will discuss how we can use gravitational lensing to identify if a black hole in nature carries a gravitomagnetic charge. For this purpose I will assume that they are described by the NUT metric. We will solve the geodesic equations using Legendre's elliptic integrals and Jacobi's elliptic functions. Then we will rederive the angular radius of the shadow, formulate a lens equation, derive redshift and travel time.

GR 1.3 Mon 9:40 GR-H2

A global view on Kerr spacetime – First person visualization of general relativity — •Тномая Reiber — Universität Hildesheim, Germany

The maximal analytic extension of slow Kerr spacetime contains an infinity of asymptotically flat "exterior" regions connected by a strongly curved "interior" region. An observer may stay in one of the exterior regions or – crossing event horizons – pass through the strongly curved region to reach one of the other asymptotically flat regions. We calculate videos of what an observer would see on different journeys through Kerr spacetime by using general relativistic ray tracing. For that purpose we use a covering of Kerr spacetime by an atlas consisting of Kerr-Schild and Kruskal-like coordinate patches.

GR 1.4 Mon 10:00 GR-H2

Non-teleology and motion of a tidally perturbed Schwarzschild black hole — •ZEYD SAM — School of Mathematical Sciences and STAG Research Centre, University of Southampton, United Kingdom — Institute for Physics and Astronomy, University of Potsdam, Germany

The prospect of gravitational wave astronomy with EMRIs has motivated increasingly accurate perturbative studies of binary black hole dynamics. Studying the apparent and event horizon of a perturbed Schwarzschild black hole, we find that the two horizons are identical at linear order regardless of the source of perturbation. This implies that the seemingly teleological behaviour of the linearly perturbed event horizon, previously observed in the literature, cannot be truly teleological in origin. The two horizons do generically differ at second order in some ways, but their Hawking masses remain identical. In the context of tidal distortion by a small companion, we also show how the perturbed event horizon in a small-mass-ratio binary is effectively localized in time, and we numerically visualize unexpected behaviour in the black hole's motion around the binary's center of mass.

GR 1.5 Mon 10:20 GR-H2

Black hole temperature in Horndeski gravity — •KAMAL HAJIAN¹, STEFANO LIBERATI², MOHAMMAD MEHDI SHEIKH-JABBARI³, and MOHAMMAD HASAN VAHIDINIA⁴ — ¹Carl von Ossietzky University of Oldenburg Department of Physics D-26111 Oldenburg — ²3SISSA, Via Bonomea 265, 34136 Trieste, Italy and INFN, Sezione di Trieste — ³5School of Physics, Institute for Research in Fundamental Sciences (IPM), P.O.Box 19395-5531, Tehran, Iran — ⁴Department of Physics, Institute for Advanced Studies in Basic Sciences (IASBS), P.O. Box 45137-66731, Zanjan, Iran

In Horndeski gravities, which are the most generic scalar-tensor theories without ghosts, the speed of graviton can be different w.r.t other massless particles/waves such as photons. We will show that this leads to a black hole temperature which is different from the standard Hawking temperature by an overall factor. The factor depends on black hole properties as well as the Lagrangian. Using this modified temperature, the first law of thermodynamics for black holes in Horndeski gravities is recovered.

GR 1.6 Mon 10:40 GR-H2

Black hole shadows enlightening quantum gravity — •MICHAEL FLORIAN WONDRAK, KOLJA KUIJPERS, JESSE DAAS, FRANK SAUERESSIG, and HEINO FAL-CKE — Institute for Mathematics, Astrophysics and Particle Physics (IMAPP), Radboud University, Nijmegen, The Netherlands

With the advent of the Event Horizon Telescope, the shadow of supermassive black holes could be resolved for the first time. Tests of the general theory of relativity at this strong-field regime come into reach. Probing quantum gravity is particularly interesting.

In this talk, we focus on extending the Einstein–Hilbert action by higher curvature terms which necessarily arise as counterterms upon quantization. We numerically solve the equations, find solutions with and without horizons, and extract the corresponding shadow radii. The results are discussed regarding observability.

GR 2: Classical Theory

Time: Monday 16:00-18:40

GR 2.1 Mon 16:00 GR-H2

A covariant formulation of violations of the Equivalence Principle — •CKAUS LÄMMERZAHL — ZARM, University of Bremen, Germany

In the Newtonian framework a violation of the Equivalence Principle can be described by introducing inertial and graviational masses. Within General Relativity it is not obvious how to couple different particles differently to the same space-time geometry. Here we propose a scheme where in a given geometrical background a violation of the universality of free fall as well as of the frame dragging can be formulated.

GR 2.2 Mon 16:20 GR-H2

A Collapsing Mass Shell with High Angular Momentum — ANDREAS KING¹, •MARKUS KING², and JÖRG FRAUENDIENER³ — ¹Preysingstraße 40, 81667 München, Germany — ²Fakultät Engineering, Hochschule Albstadt-Sigmaringen, 72458 Albstadt, Germany — ³Department of Mathematics and Statistics, University of Otago, Dunedin 9054, New Zealand

We calculate the free-fall collapse of a dust shell of mass M and radius R(s), with s being proper time along the worldlines of its collisionless particles, and rotating fast in the initial angular velocity ω_0 around an axis through its center in an asymptotically flat spacetime. The Einstein equations are solved to second order perturbation theory in ω_0 . We show existence of such a system with flat interior, free of gravitational waves produced by the quadrupolar deformation of the shell due to centrifugal effects. Stationary solutions of the underlying master equation of gravitational perturbations are given.

GR 2.3 Mon 16:40 GR-H2

Energy Conditions in Reverse-Engineered Metrics — •SEBASTIAN SCHUSTER¹, JESSICA SANTIAGO², and MATT VISSER³ — ¹Charles University, Prague, Czech Republic — ²Aristotle University of Thessaloniki, Thessaloniki, Greece — ³Victoria University of Wellington, Wellington, New Zealand

Many familiar metrics of general relativity have been achieved by integrating the Einstein equations for a particular source stress-energy tensor. In other cases,

Location: GR-H2

physical considerations lead to a motivation for studying metrics of a particular form; here, then, the question is to derive the stress-energy tensor from a given metric. While mathematically simpler—as differentiation is simpler than integration—this opens a long list of questions which stress-energy tensors should be considered physical. An early heuristic for such an evaluation were so-called energy conditions. In this talk, we will describe a proof for why "warp drives" will always violate pointwise energy conditions, and how even rather benign, reverse-engineered metrics ("tractor beams") will do so.

GR 2.4 Mon 17:00 GR-H2 Relativistic Geodesy: What do we know? — •DENNIS PHILIPP — ZARM, Uni-

versity of Bremen Geodesy is the science of the properties of our Earth, in particular its gravity field. Conventional geodesy builds on (the concepts of) Newtonian gravity. Thus, at the level of a relativistic theory of gravity, the underlying framework needs to be renewed and basic notions need to be generalized. This opens an entirely new perspective on the matter - chronometric geodesy - which investigates gravity by, e.g., the use of clocks and clock networks. In this talk, I will review the status of the theoretical aspects of relativistic geodesy and address concepts such as the potential, multipole moments, geoid, reference ellipsoid, and height notions in the conventional and in the relativistic framework. Moreover, observables and measurement prescriptions are discussed and an outlook on future developments is given.

GR 2.5 Mon 17:20 GR-H2

On free fall of quantum matter — •VIACHESLAV EMELYANOV — Institute of Theoretical Physics, Karlsruhe Institute of Technology, Wolfgang-Gaede-Straße 1, 76131 Karlsruhe, Germany

According to Newton's gravitational law, any object to have a non-zero gravitational mass is a source of gravity. It is a result of numerous experiments that gravitational mass is equal with good accuracy to inertial mass of a macroscopic object. Thus, all objects fall down equally fast, assuming same initial position and velocity. This circumstance is promoted to the weak equivalence principle in General Relativity, that is related to the concept of affine connection, giving in its turn the concept of geodesic corresponding to particles' trajectories in curved spacetime. However, there is an expectation that the free-fall universality may not hold for quantum matter.

In this talk, we intend to introduce our approach to quantum particle physics in curved spacetime. It is based on quantum field theory and the general principle of relativity, which are used to build a model for quantum particles in gravity. We then obtain by its means a deviation from a classical geodesic in the Earth's gravitational field. This shows that free fall depends on quantum-matter properties. Specifically, we find that the free-fall universality and the wave-packet spreading are mutually exclusive phenomena. Assuming that the latter is more fundamental, we present the first-ever estimate of the Eötvös parameter for a pair of atoms used nowadays in quantum tests of the universality of free fall in atom interferometers and compare that with recent experimental results.

GR 2.6 Mon 17:40 GR-H2

Gravitational field recovery via inter-satellite redshift measurements — •JAN HACKSTEIN, EVA HACKMANN, CLAUS LÄMMERZAHL, and DENNIS PHILIPP — Center of Applied Space Technology and Microgravity, Bremen, Germany Satellite gravimetry is a promising technique to monitor global changes in the Earth system. High-precision atomic clocks are already being compared to mea-

Time: Monday 16:15–18:15

GR 3.1 Mon 16:15 GR-H3

Through the Big Bang – • PAULA REICHERT – Mathematisches Institut, LMU München, Germany

This talk presents latest results regarding the evolution of the universe through the Big Bang singularity on shape space. Relationalism in the form of modern shape dynamics suggests that the Big Bang is only a turning or Janus point within an overall time-symmetric, eternal evolution - a common past in a universe with one past (i.e. the Big Bang) and two futures in both directions away from it. This idea is supported by the 2016 result that, for the quiescent Bianchi IX model of GR, the shape (i.e. angular) degrees of freedom can be evolved uniquely through the (otherwise singular) point of zero spatial volume. Studies of the total collision singularity on non-relativistic shape space further suggest that, at the point of total collision (i.e. the Big Bang of the Newtonian N-body universe), the system is exceptionally homogenous, forming a state of minimal shape complexity and minimal entropy. At the same time, both complexity and entropy increase as the system expands and galaxies form in both directions away from the Janus point, thereby marking two gravitational and entropic arrows of time.

sure physical heights in terrestrial gravimetry. In relativistic gravity, a clock comparison is sensitive to the clocks' positions and relative velocity in the gravity field. Thus, clocks are an ideal tool to investigate the Earth's gravity field. To cover the whole Earth, orbiting satellites can be equipped with clocks and observed by terrestrial ground stations. One important obstacle for Earth-satellite gravimetry, however, is the low measurement accuracy of a satellite's velocity, which enters into the redshift via the Doppler effect. Here we follow an alternative approach without absolute velocity measurements based on the framework of general relavity. We consider an idealised satellite setup in the Schwarzschild spacetime where the monopole moment is recovered from pairwise redshift measurements between multiple satellites equipped with clocks. We investigate whether or not the redshift between two satellites can be retrieved as a function depending only on relative observables between the satellites. This method promises a higher accuracy for gravity field recovery by bypassing the Doppler effect. We compare the results and error estimates of these inter-satellite measurements with conventional Earth-satellite measurements and conclude with future applications of this theoretical setup.

GR 2.7 Mon 18:00 GR-H2

Probing gravitational parity violation with compact binaries — •**H**ECTOR O. SILVA and JAN STEINHOFF — Max Planck Institute for Gravitational Physics (Albert Einstein Institute), Am Mühlenberg 1, 14476 Potsdam, Germany

The detection of gravitational-waves signals produced by the inspiral and coalescence of compact binaries have opened a new vista into the nonlinear, highly dynamical regime of gravity. These observations have allowed us to perform new tests of general relativity and also to probe (or constrain) modifications to Einstein's theory. In this work, we report on an ongoing effort to study the effects of higher-curvature parity violating modifications to general relativity on the inspiral of compact binaries with the goal of possibly placing (or forecasting) constraints on parity violation in gravity with present or future gravitational wave observatories.

GR 2.8 Mon 18:20 GR-H2

Konzept eines satellitengestützten Tests des gravitomagnetischen Uhreneffektes — •JAN SCHEUMANN, DENNIS PHILIPP, EVA HACKMANN, SVEN HERR-MANN, BENNY RIEVERS und CLAUS LÄMMERZAHL — ZARM, Universität Bremen, 28359 Bremen, Deutschland

Die Allgemeine Relativitätstheorie besagt, dass die Rotation eines Körpers einen nicht-newtonschen Einfluss auf Objekte in seinem Orbit ausübt. Ein Beispiel für einen solchen Effekt ist die als Lense-Thirring-Effekt bekannte Drehung der Knotenlinie eines Satelliten. Ein weiterer Effekt ist der zuerst von Cohen und Mashhoon beschriebene sogenannte gravitomagnetische Uhreneffekt, der den Unterschied in der Eigenzeit zweier Uhren in gegenläufigem Orbit um den Zentralkörper beschrieben. Dieser Effekt ist bisher für verschiedene idealisierte Orbits theoretisch beschrieben worden, wurde jedoch noch nicht experimentell bestätigt.

Nachdem zwei der Galileo-Satelliten auf nicht für GNSS-Zwecke geeigneten Orbits bereits dafür genutzt werden konnten, die Unsicherheit im Nachweis der gravitativen Rotverschiebung mittels der an Bord befindlichen passiven Wasserstoffmaser zu verringern, wird die Nutzung dieser Satelliten für weitere Tests relativistischer Effekte untersucht.

In diesem Vortrag wird ein Konzept vorgestellt, mit dem der gravitomagnetische Uhreneffekt untersucht werden könnte und die hierfür notwendigen Voraussetzungen mit dem aktuellen Stand der Technik verglichen.

GR 3: Cosmology

Location: GR-H3

GR 3.2 Mon 16:35 GR-H3

Graviton corrections to the Newtonian potential using invariant observables — •MARKUS B. FRÖB, CONSTANTIN REIN, and RAINER VERCH — Institut für Theoretische Physik, Universität Leipzig, Brüderstraße 16, 04103 Leipzig, Germany

We consider the effective theory of perturbative quantum gravity coupled to a point particle, quantizing fluctuations of both the gravitational field and the particle's position around flat space. Using a recent relational approach to construct gauge-invariant observables, we compute one-loop graviton corrections to the invariant metric perturbation, whose time-time component gives the Newtonian gravitational potential. The resulting quantum correction consists of two parts: the first stems from graviton loops and agrees with the correction deviced by other methods, while the second one is sourced by the quantum fluctuations of the particle's position and energy-momentum, and may be viewed as an analog of a "Zitterbewegung". As a check on the computation, we also recover classical corrections which agree with the perturbative expansion of the Schwarzschild metric.

GR 3.3 Mon 16:55 GR-H3

Spatial Geometry of the Large-Scale Universe: The Role of Quantum Gravity, Dark Energy and Other Unknowns — •MARC HOLMAN — Utrecht University, Utrecht, Netherlands

Most key features of contemporary concordance cosmology can be directly linked to observational facts, such as Hubble's law, the existence and properties of the Cosmic Microwave Background (CMB) - in particular its extreme uniformity - light element abundances and large-scale flatness. In some cases, these features first appeared in the form of further model constraints in the light of new observational data - e.g., the discoveries of distance-proportional galactic redshifts and the CMB, which were taken as irreconcilable with static and steady-state cosmological models, respectively. In other cases, they first appeared in the form of additional ingredients in the light of largely existent, but seemingly unaccounted for, observational data - e.g., the near-flatness of the Universe's large-scale spatial geometry and the existence of mass discrepancies, which were argued irreconcilable with standard Big Bang cosmology assuming only "normal matter" to be present. As recent work has emphasized however, the observed near-flatness of the large-scale Universe as a partial, but key motivation for assuming the existence of an ultra-short, inflationary expansion of the very early Universe, has a long and troubled history. In this respect, the present work strengthens earlier results regarding the absence of a cosmological flatness problem of the sort that could potentially be resolved by inflation.

GR 3.4 Mon 17:15 GR-H3

Compact objects from effective quantum gravity — Piero Nicolini^{1,2,3} and •Salvatore Samuele Sirletti^{2,4} — ¹New York University Abu Dhabi, Abu Dhabi, UAE — ²Frankfurt Institute for Advanced Studies, Frankfurt am Main, Germany — ³Johann Wolfgang Goethe-Universität, Frankfurt am Main, Germany — ⁴Università degli Studi di Napoli Federico II, Naples, Italy

It has been shown that the UV finiteness of Superstring Theory can led to the derivation of a family of regular black hole solutions in the gravity-matter decoupling limit. The latter is a regime governed by stringy effects like noncommutativity and T-duality. The most natural realization of a non-local structure inheriting noncommutative geometry effects is the Gaussian profile for the energy density in the relativistic stress tensor.

In this talk, we present two interesting regular black hole/compact object alternatives that stem from postulating a smooth transition between a quantum gravity dominated region at the origin, and a corona of degenerate nuclear matter around it. The derivation of the resulting metric allows for the description of a regular horizonless Planckian object and a neutron star with a quantum vacuum at its center.

GR 3.5 Mon 17:35 GR-H3

Generalized Uncertainty Relations and the Problem of Dark Energy -•Маттнеw J. Lake — Frankfurt Institute for Adavanced Studies, Frankfurt am Main, Germany

We outline a new model in which generalised uncertainty relations are obtained without modified commutation relations. While existing models introduce modified phase space volumes for the canonical degrees of freedom, we introduce new degrees of freedom for the background geometry. The background is treated as a genuinely quantum object, with an associated state vector, and the model naturally gives rise to the extended generalised uncertainty principle (EGUP). Importantly, this approach solves (or rather, evades) well known problems associated with modified commutators, including violation of the equivalence principle, the soccer ball problem for multi-particle states, and the velocity dependence of the minimum length. However, it implies two radical conclusions. The first is that space must be quantised on a different scale to matter and the second is that the fundamental quanta of geometry are fermions. We explain how, in the context of the model, this gives rise to an effective dark energy density, without contradicting established results including the no go theorems for multiple quantisation constants, which still hold for species of material particles, and the spin-2 nature of gravitons.

GR 3.6 Mon 17:55 GR-H3

Power spectrum for perturbations in an inflationary model for a closed universe — •TATEVIK VARDANYAN and CLAUS KIEFER — Institute for Theoretical Physics, University of Cologne, Köln, Germany

We derive the power spectrum of primordial quantum fluctuations in an inflationary universe for curvature parameter $\mathcal{K} = 1$. This is achieved through a Born-Oppenheimer type of approximation scheme from the Wheeler-DeWitt equation of canonical quantum gravity using gauge-invariant variables. Compared to the flat model, the closed model exhibits a deficit of power at large scales.

Reference: arXiv:2111.07835

GR 4: General Relativity

Time: Tuesday 11:00-12:30

GR 4.1 Tue 11:00 GR-H2 Invited Talk Pseudospectrum and black hole quasi-normal mode (in)stability — •Rodrigo Panosso Macedo — University of Southampton

Black hole spectroscopy is as a powerful approach to extract spacetime information from gravitational wave observed signals. However, quasinormal mode (QNM) spectral instability under high wave-number perturbations has been recently shown to be a common classical general relativistic phenomenon. I will discuss these recent results on the stability of QNM in asymptotically flat black hole spacetimes by means of a pseudospectrum analysis.

Invited Talk GR 4.2 Tue 11:45 GR-H2 Observable Signatures of Quantized Gravity in Quantum Optical Experiments — •DENNIS RÄTZEL — Institut für Physik, Humboldt Universität zu Berlin, Newtonstraße 15, 12489 Berlin, Germany

For nearly a century there has been an apparent tension between known laws of physics. The classical theory of General Relativity describes all macroscopic gravitational phenomena, while Quantum Theory is the basis for the description of matter at the microscopic scale. Yet, so far there has been no consensus on how, or even if, they can fit together. A final and conclusive answer to the question of whether gravity ought to be quantized must be based on empirical evidence.

In this talk, I will discuss the search for observable signatures of quantized gravity in quantum optics by means of two examples of gravitationally interacting quantum systems: photons in a polarization-entangled state and small masses in spatial superposition states. I will present experimental proposals and discuss their feasibility and what can be learned from them in principle.

GR 5: Gravitational Waves

Time: Tuesday 16:15-18:35

GR 5.1 Tue 16:15 GR-H2

The scientific potential of gravitational waves from Extreme Mass Ratio Inspirals — •LORENZO SPERI — Max Planck Institute for Gravitational Physics (Albert Einstein Institute) Am Muehlenberg 1, 14476 Potsdam, Germany One of the primary sources for the future space-based gravitational wave detector, the Laser Interferometer Space Antenna, are the inspirals of small compact objects into massive black holes in the centers of galaxies. The gravitational wave observations of such Extreme Mass Ratio Inspiral (EMRI) systems have a huge scientific potential. The compact object typically completes hundred thousand cycles in band, during which time it is orbiting in the strong field region close to the central rotating black hole. Because of this, EMRI signals encode a detailed map of the background space-time and offer a unique opportunity to measure the properties and environment of Massive Black Holes, and to test for deviations from General Relativity (GR). Properly modeling EMRIs is of paramount imporLocation: GR-H2

Location: GR-H2

tance to unlock such potential. In this talk I will review how EMRI waveform models are constructed, and show how environmental and beyond GR effects can be included. I will discuss the ability of EMRI signals to constrain accretion disk and beyond GR parameters. I will conclude by highlighting the future challenges for EMRI gravitational wave modeling.

GR 5.2 Tue 16:35 GR-H2

Detecting long-duration gravitational wave signals — \bullet LIUDMILA FESIK and MARIA ALESSANDRA PAPA — Max Planck Institute for Gravitational Physics (Albert Einstein Institute) and Leibniz University, Callinstraße 38, 30167 Hannover, Germany

Spinning neutron stars are sources of long-duration continuous waves (CWs) that may be detected by interferometric detectors. We focus on glitching pulsars with abrupt spin-ups and long term spin-down, which imprint in CWs as

transient signals from weeks to months. Standart method for identifying transient signals is the match-filtering, which combines a coherent detection statistics over time intervals of different duration. We propose a new method, where the most information from an initial search is considered in order to set up the post-following transient searches. We characterize the method by determining the false alarm and false dismissal probabilities for different signal strengths, and appropriate choices of the relative detection thresholds. We compare the sensitivity of this method with the standart match-filtering.

GR 5.3 Tue 16:55 GR-H2

Interference of strongly lensed Gravitational Waves — •STEFANO SAVASTANO¹, FILIPPO VERNIZZI², and MIGUEL ZUMALACARREGUI¹ — ¹Max Planck Institute for Gravitational Physics (Albert Einstein Institute), Am Mühlenberg 1, D-14476 Potsdam-Golm, Germany — ²Institut de Physique Theorique, Universite Paris Saclay CEA, CNRS, 91191 Gif-sur-Yvette, France

Gravitational waves (GW) can be lensed by inhomogeneities on their journey from source to observer, just as electromagnetic radiation. Lens parameters can be extracted from the phase evolution of lensed signals. Continuous GWs (CGW) have a negligible frequency evolution, so the full degeneracy between source and lens parameters limits the extraction. I will discuss how studying the interference pattern produced by strongly lensed CGWs images can enhance the inference of lens parameters. In particular, I will show that the relative motion of the lens to the source or a small frequency evolution can break this degeneracy for some systems. Finally, I will discuss detection perspectives of CGWs lensing for Earth and space-based detectors and elaborate on some possible applications of this tool.

GR 5.4 Tue 17:15 GR-H2

Scattering gravitational waves off effective spinning black holes — • VENKATA SAI SAKETH MUDDU and JUSTIN VINES — Max Planck institute for gravitational physics, Potsdam Science Park Am Mühlenberg 1 D-14476 Potsdam, Germany The scattering of gravitational waves off a black hole contains valuable information characterizing the response of a black hole to an external gravitational field. We present a new method for computing scattering amplitudes for this process, using an effective worldline theory where the black hole is treated as a point particle equipped with multipole moments. These moments can be intrinsic or induced. The effective action couples the black hole's moments to an external gravitational field via interaction terms, which contain unknown coefficients. We consider a plane wave impinging on the black hole and consider its response to the incoming wave. In the effective theory, this leads to a scattered wave produced by the black hole. Comparing the scattered wave to the incident wave gives us the amplitude. We solve for the amplitude iteratively in powers of spin (the Kerr ring radius). Comparing the amplitude from the effective theory to that obtained by solving the Teukolsky equation in a Kerr background can fix the unknown coefficients in the effective action. Using our amplitude as an effective "gravitational Compton amplitude" in a triangle unitarity cut, we can also compute contributions to the relativistic scattering amplitude and subsequently the radial action for black-hole-black-hole scattering.

GR 5.5 Tue 17:35 GR-H2

Black-hole ringdown as a probe of higher-curvature gravity theories — •ABHIRUP GHOSH¹, HECTOR O. SILVA¹, and ALESSANDRA BUONANNO^{1,2} — ¹Max Planck Institute for Gravitational Physics (Albert Einstein Institute), Am Mühlenberg 1, Potsdam 14476, Germany — ²Department of Physics, University of Maryland, College Park, MD 20742, USA

GR 6: Cosmology

Time: Tuesday 16:15-18:15

GR 6.1 Tue 16:15 GR-H3

Psi in the sky - the cosmology window on wavelike dark matter - •CORA UH-LEMANN — Newcastle University, Newcastle-upon-Tyne, UK

Despite the astonishing success of cosmological probes in constraining the LCDM model, the dark matter mass remains one of the least constrained physical parameters. Wavelike dark matter is an intriguing alternative to standard cold dark matter with key particle physics motivations (like the QCD axion or ultralight axion-like particles) and distinct astrophysical signatures. With a simple dynamical model for the evolution of the dark matter wavefunction (Psi), I will demonstrate how to predict the formation of topological defects and granules arising from destructive and constructive wave interference. Together with the wave interference imprint on substructures that leads to exciting and varied probing mechanisms bridging cosmology, astrophysics and particle physics.

Observations of gravitational waves from the mergers of compact objects like black holes have allowed us to test, for the first time, the nature of strong-field gravity. Albert Einstein's theory of General relativity (GR) remains our best description of gravitational interaction, and most of the strong-field tests of gravity demonstrated on LIGO-Virgo signals have been null tests of GR, i.e, they check

for consistency between observations and predictions of GR and place bounds on possible deviations from these predictions. Issues such as the quantization of GR and the cosmological constant problem suggest that Einstein's theory might not be a complete description of gravity and might require modifications. In this work, we use observations of the black hole ringdown from the latest LIGO-Virgo catalog of gravitational wave signals (GWTC-3) in an attempt to constrain possible deviations due to well-motivated higher-curvature theories of gravity.

GR 5.6 Tue 17:55 GR-H2

Probing new physics on the horizon of black holes with gravitational waves — •ELISA MAGGIO — Max Planck Institute for Gravitational Physics (Albert Einstein Institute), Am Mühlenberg 1, 14476 Potsdam, Germany

Black holes are the most compact objects in the universe. According to general relativity, black holes have a horizon that hides a singularity where Einstein's theory breaks down. Recently, gravitational waves opened the possibility to probe the existence of horizons and investigate the nature of compact objects. This is of particular interest given some quantum-gravity models which predict the presence of horizonless and singularity-free compact objects. Such exotic compact objects can emit a different gravitational-wave signal relative to the black hole case. In this talk, I overview the gravitational-wave phenomenology of exotic compact objects. I infer how extreme mass-ratio inspirals observable by future gravitational-wave detectors will allow for model-independent tests of the black hole paradigm.

GR 5.7 Tue 18:15 GR-H2

General Relativity from Worldline Quantum Field Theory — •GUSTAV UHRE JAKOBSEN — Humboldt-Universität zu Berlin, Berlin, Germany — Max-Planck-Institut für Gravitationsphysik, Potsdam, Germany

The worldline quantum field theory (WQFT) has recently been developed in order to describe classical gravitational interactions. While the goal is to analyse bound dynamics of binaries and their gravitational waves, this framework most naturally describes scattering (unbound) events. This in line with other quantum field theoretic approaches. However, several insights from this field has shown that the two regimes of unbound vs bound motion are intrinsically related.

I will present our current supersymmetric WQFT which describes spinning black holes or stars. Here, the supersymmetry encodes the symmetries of the spin degrees of freedom. I will consider several of the scattering observables that we have derived from this WQFT and how they can be related to bound ones. Examples are the eikonal, the spin kick and the total deflection.

Finally, I will consider future perspectives and challenges of the WQFT. This includes extending the WQFT in order to describe new phenomenological aspects of e.g. neutron stars and improving the integration techniques in order to increase precision of observables.

GR 6.2 Tue 16:35 GR-H3

The Schrödinger-Poisson Equation in One Spatial Dimension — •NICO SCHWERSENZ¹, TIM ZIMMERMANN², VICTOR LOAIZA³, JAVIER MADROÑERO³, MASSIMO PIETRONI^{4,5}, LUCA AMENDOLA¹, and SANDRO WIMBERGER^{4,5} — ¹Institut für Theoretische Physik, Universität Heidelberg, Germany — ²Institutt for Teoretisk Astrofysikk, Universitetet i Oslo, Norway — ³Departamento de Física, Universidad del Valle, Cali, Colombia — ⁴Dipartimento di Scienze Matematiche, Fisiche e Informatiche, Università di Parma, Italy — ⁵INFN, Sezione di Milano Bicocca, Gruppo Collegato di Parma, Italy

Despite the success of Wave-like Dark Matter [1] in explaining cosmological processes, its major issue is the high demand in computational resources. Not only does the non-linear, non-local nature of the underlying Schrödinger-Poisson equation pose a problem, but also the range of scales that have to be resolved. We construct two distinct one-dimensional toy models [2] that are less expensive from a numerical viewpoint, but still provide analogues to the phenomena observed in three dimensions. Our high-precision numerical technique is tested by an independent method. Some exemplary results will be shown for two different

GR 6.5 Tue 17:35 GR-H3

ways of treating the transverse dimensions, assuming uniform matter distribution in the first and strong confinement - effectively renormalizing the mass - in the second case.

J. C. Niemeyer, Progress in Particle and Nuclear Physics 113, 103787 (2020)
 T. Zimmermann, N. Schwersenz, M. Pietroni, S. Wimberger, Physical Review D 103, 083018 (2021)

GR 6.3 Tue 16:55 GR-H3

The Hawking energy of a cosmic observer in linearly perturbed FLRW — •DENNIS STOCK and RUTH DURRER — Université de Genève, Département de Physique Théorique and Center for Astroparticle Physics, 24 quai Ernest-Ansermet, CH-1211 Genève 4, Switzerland

Addressing cosmological questions exclusively based on observations requires a formulation on the past lightcone of the cosmic observer. In this talk, the question of how to define gravitational energy associated with the past lightcone of a cosmic observer is studied by introducing Hawking's quasi-local energy as a tentative energy measure of the observable Universe. The Hawking energy phenomenologically quantifies energy in terms of light bending. This talk will mainly focus on the relation of the Hawking energy to cosmological observables within linear perturbation theory on an FLRW background.

GR 6.4 Tue 17:15 GR-H3

Influence of cosmological expansion in local experiments — •FELIX SPENGLER¹, ALESSIO BELENCHIA^{1,2}, DENNIS RÄTZEL³, and DANIEL BRAUN¹ — ¹Intitut für theoretische Physik, Universität Tübingen, Tübingen, Germany — ²Centre for Theoretical Atomic, Molecular, and Optical Physics, School of Mathematics and Physics, Queens University, Belfast, United Kingdom — ³Humboldt Universität zu Berlin, Institut für Physik, Berlin, Germany

Whether the cosmological expansion can influence the local dynamics, below the galaxy clusters scale,has been the subject of intense investigations. We consider McVittie and Kottler spacetimes, embedding a spherical object in an expanding Friedmann-Lemaître-Robertson-Walker spacetime as a rough approximation of a local environment immersed in a globally expanding universe. We then calculate the influence of the cosmological expansion on the frequency shift of a resonator moving along different trajectories and estimate its effect on the exchange of light signals between local observers. Our results show the impact of the global expansion on these local experiments and give and upper estimate on the effects we can expect in more realistic conditions below the galaxy clusters scale.

Reaching precision cosmology faster with velocities — •MIGUEL QUARTIN — Institute of Theoretical Physics, Heidelberg University, Philosophenweg 16, Heidelberg — Instituto de Física, Universidade Federal do Rio de Janeiro, 21941-972, Rio de Janeiro, RJ, Brazil

We will show how standard candles such as type Ia supernovae and standard sirens can be used as tracers of both density and velocity fields, thus serving a new purpose in cosmology beyond mere distance indicators. We discuss how these new tracers can be combined with galaxy surveys, combining galaxy and supernova position and redshift data with supernova peculiar velocities, obtained through their magnitude scatter. The full method relies on a 6x2pt analysis which includes six power spectra. We proceed then to forecast the performance of future surveys like LSST and 4MOST with a Fisher Matrix analysis, adopting both a model-dependent and a model-independent approach. We compare the performance of the 6x2pt approach to the traditional one using only galaxy clustering and some recently proposed combinations of galaxy and supernovae data and quantify the possible gains by optimally extracting the linear information. We show that the 6x2pt method shrinks the uncertainty in growth of structure parameters significantly. The combined clustering and velocity data on the growth of structures has uncertainties at similar levels to those of the CMB but exhibit orthogonal degeneracies, and the combined constraints yield very large improvements in parameters both at the background and perturbation-level.

GR 6.6 Tue 17:55 GR-H3

First constraints on the intrinsic CMB dipole and our velocity with Doppler and aberration — •PEDRO DA SILVEIRA FERREIRA — Observatório do Valongo, Universidade Federal do Rio de Janeiro, Rio de Janeiro - RJ, Brazil

We test the usual hypothesis that the Cosmic Microwave Background (CMB) dipole, its largest anisotropy, is due to our peculiar velocity with respect to the Hubble flow by measuring independently the Doppler and aberration effects on the CMB using Planck 2018 data. We remove the spurious contributions from the conversion of intensity into temperature and arrive at measurements which are independent from the CMB dipole itself for both temperature and polarization maps and both SMICA and NILC component-separation methods. Combining these new measurements with the dipole one we get the first constraints on the intrinsic CMB dipole. Assuming a standard dipolar lensing contribution we can put an upper limit on the intrinsic amplitude: 3.7 mK (95% CI). We estimate the peculiar velocity of the solar system without assuming a negligible intrinsic dipole contribution: $v = (300^{+111}_{-93}) \text{ km/s with } (l, b) = (276 \pm 33, 51 \pm 19)^{\circ}$ [SMICA], and $v = (296^{+111}_{-88}) \text{ km/s}$ with $(l, b) = (280 \pm 33, 50 \pm 20)^{\circ}$ [NILC] with negligible systematic contributions. These values are consistent with the peculiar velocity hypothesis of the dipole.

GR 7: General Relativity

Time: Wednesday 11:00–11:45

Invited TalkGR 7.1Wed 11:00GR-H2Numerical Relativity and Gravitational Wave Observations• HARALDPFEIFFER — Max-Planck-Institute for Gravitational Physics

Gravitational wave detectors have observed nearly 100 mergers of compact object binaries since the ground-breaking first observation in 2015. Direct supercomputer calculations are the only means to access the highly dynamic and non-

GR 8: Gravitational Waves

Time: Wednesday 11:45-12:25

GR 8.1 Wed 11:45 GR-H2

On the accuracy of gravitational-wave observations — •FRANK OHME — Max-Planck-Institut für Gravitationsphysik (AEI), Hannover, Deutschland — Leibniz-Universität Hannover, Deutschland

The catalogue of gravitational-wave observations continues to grow with every LIGO-Virgo observing run. Nearly 100 signals from compact binary mergers have been identified in the first three runs. Each signal's source properties are determined by comparing theoretical models with the data. This process is subject to two categories of uncertainties (or errors): statistical errors account for the presence of detector noise; systematic errors arise from inaccuracies of the signal models. I this talk, I will highlight ways to measure and compare the two sources of errors and assess how they impact current and future gravitational-wave observations.

linear merger phase of such binaries. The simulations elucidate the behavior of space-time and matter in the extreme conditions near merger and play a crucial part in detection and analysis of gravitational wave observations. This talk introduces the techniques of numerical relativity and surveys some recent results. The talk then highlights the importance of numerical relativity simulations for gravitational wave astronomy.

Location: GR-H2

Location: GR-H2

GR 8.2 Wed 12:05 GR-H2

Dynamical tides and gravitational scattering – •JAN STEINHOFF – Max Planck Institute for Gravitational Physics (Albert Einstein Institute), Potsdam, Germany

Gravitational wave astronomy offers a promising tool to infer the nuclear physics of neutron stars from observations of inspiraling and merging binaries. This talk addresses the necessity to accurately model dynamical tidal effects during the inspiral, for instance the resonant excitation of neutron star oscillation modes, for the era of third generation detectors. For this purpose, a fully relativistic effective-field-theory model for tidal effects is put forward. Gravitational scattering is discussed as a promising Gedanken experiment to determine the tidal parameters of the model, and for modeling binary systems in general.

GR 9: Gravitational Waves

Time: Wednesday 16:15-18:35

GR 9.1 Wed 16:15 GR-H2

Machine Learning Gravitational-Wave Search Mock Data Challenge — •MARLIN BENEDIKT SCHÄFER — Albert-Einstein-Institut, D-30167 Hannover, Germany — Leibniz Universität Hannover, D-30167 Hannover, Germany

Gravitational wave astronomy is a rapidly growing field and the number of detections is rising faster with each observational period. With this come new challenges when extracting the signals from noise. A new approach to handle large quantities of data and possibly search regions of parameter space that are computationally prohibitive to search with state-of-the-art classical algorithms is the utilization of machine learning techniques. This projects aims to clarify the capabilities of current deep learning algorithms and how they compare to traditional methods. The challenge provides mock data of gradually increasing realism to aid the adoption of machine learning based algorithms in detection pipelines and wants to help establishing the wide adoption of astrophysically motivated evaluation metrics.

GR 9.2 Wed 16:35 GR-H2

New generation effective-one-body waveforms for binary black holes with non-precessing spins — SERGUEI OSSOKINE¹, DEYAN MIHAYLOV¹, •LORENZO POMPILI¹, ALESSANDRA BUONANNO^{1,2}, MICHAEL PÜRRER¹, and MOHAMMED KHALIL^{1,2} — ¹Max Planck Institute for Gravitational Physics (Albert Einstein Institute), Am Mühlenberg 1, Potsdam 14476, Germany — ²Department of Physics, University of Maryland, College Park, Maryland 20742, USA

We present an improved inspiral-merger-ringdown gravitational waveform model for quasi-circular, spinning, non-precessing binary black holes within the effective-one-body (EOB) formalism. Compared to its predecessor SEOBNRv4HM the waveform model i) incorporates recent high-order post-Newtonian results in the inspiral, with improved factorizations ii) includes the gravitational modes $(\ell, |m|) = (3, 2), (4, 3)$, in addition to the $(\ell, |m|) =$ (2, 2), (3, 3), (2, 1)(4, 4), (5, 5) modes already implemented in SEOBNRv4HM iii) has been recalibrated to larger mass ratios and spins using a catalog of 441 numerical-relativity (NR) simulations, and to 13 additional waveforms from black-hole perturbation theory. The accuracy of the waveform model is quantified by computing the unfaithfulness against the NR catalog used for its construction. The waveform model has been implemented in a new Python framework, that makes it easily extensible to include spin-precession and eccentricity effects, thus making it the starting point for a new generation of EOB waveform models (SEOBNRv5) to be employed for upcoming observing runs of the LIGO-Virgo-KAGRA detectors.

$GR~9.3 \quad Wed~16:55 \quad GR-H2 \\ \textbf{TEOBResumS: an advanced waveform model for O4-} \bullet ROSSELLA GAMBA-}$

Friedrich-Schiller-Universität Jena, Jena, Germany

The detection of Gravitational Waves by LIGO and Virgo opened a new, exceptional avenue for studying the physics of binary systems of compact objects, such as black holes and neutron stars. Source properties can be extracted from the LIGO/Virgo data via matched filtering techniques that employ waveform templates, i.e. theoretical models of the gravitational waves (GWs) emitted by the two coalescing bodies. To be able to obtain the largest amount of information from the data, such models must incorporate a large amount of physics while retaining high faithfulness to waveforms from numerical relativity simulations. In this talk I will present TEOBResumS, an efficient state-of-the-art waveform model for GWs from generic binary systems. I will detail the physics included, highlight its computational efficiency and faithfulness and show applications to real and simulated data in view of the fourth observing run O4, planned for late 2022.

GR 9.4 Wed 17:15 GR-H2

TEOBResumS for black-hole-neutron star merger waveforms — •ALEJANDRA GONZALEZ¹, ROSSELLA GAMBA¹, MATTEO BRESCHI¹, FRANCESCO ZAPPA¹, SEBASTIANO BERNUZZI¹, and ALESSANDRO NAGAR² — ¹Theoretisch-Physikalisches Institut, Friedrich-Schiller-Universität Jena, 07743, Jena, Germany — ²INFN Sezione di Torino, Via P. Giuria 1, 10125 Torino, Italy

We present a new effective-one-body (EOB) model for black-hole-neutron-star merger waveforms based on a numerical-relativity (NR) informed model of the remnant BH state and a ringdown that deforms the EOB ringdown for binary black holes. The new model reproduces the (2,2) mode waveform of NR simulations with typical phase agreement within 0.5 rad to merger and within 1 rad including ringdown. Comparing to other available BHNS waveform models, the NR phasing is captured with a comparable accuracy. The model also Location: GR-H2

includes higher modes (2,1), (2,2), (3,2), (3,3), (4,4) and (5,5). We present a full Bayesian analysis of the gravitational-wave events GW190814, GW200105, and GW200115, and find consistent results with previous studies.

GR 9.5 Wed 17:35 GR-H2

Realistic observing scenarios for the next decade of early warning detection of binary neutron stars — RYAN MAGEE¹ and •SSOHRAB BORHANIAN² — ¹California Institute of Technology, Pasadena, USA — ²Friedrich-Schiller-Universität Jena, Jena, Germany

We describe realistic observing scenarios for early warning detection of binary neutron star mergers with the current generation of ground-based gravitational-wave detectors as these approach design sensitivity. Using Fisher analysis, we compute both the number of detections and the sky localizations to expect from future detector network configurations. We estimate that the Advanced LIGO and Advanced Virgo facilities will detect two signals before merger in their fourth observing run, while the addition of the Kagra and LIGO-India detectors, at design sensitivites, should increase these numbers to the order of 10 early warning detections per year in the fifth observing run. More than 70% of these events will be localized to less than 100 deg², with one achieving a localization of $\sim 20 \deg^2$. Given uncertainties in sensitivities, participating detectors, and tuty cycles, we include a data release that allows for full generalizability of future detector networks so electromagnetic observers can tailor preparations towards their preferred models.

GR 9.6 Wed 17:55 GR-H2 Quantifying modelling uncertainties when combining multiple gravitational-wave detections from binary neutron star sources — \bullet NINA KUNERT¹, PETER TSUN HO PANG^{2,3}, INGO TEWS⁴, MICHAEL WILLIAM COUGHLIN⁵, and TIM DIETRICH^{1,6} — ¹Institute for Physics and Astronomy, University of Potsdam, D-14476 Potsdam, Germany — ²Nikhef, Science Park 105, 1098 XG Amsterdam, The Netherlands — ³Institute for Gravitational and Subatomic Physics (GRASP), Utrecht University, Princetonplein 1, 3584 CC Utrecht, The Netherlands — ⁴Theoretical Division, Los Alamos National Laboratory, Los Alamos, NM 87545, USA — ⁵School of Physics and Astronomy, University of Minnesota, Minneapolis, Minnesota 55455, USA — ⁶Max Planck Institute for Gravitational Physics (Albert Einstein Institute), Am Muehlenberg 1, Potsdam 14476, Germany

The combined analysis of multiple gravitational-wave signals from binary neutron star sources enables to constrain the neutron-star radius with unprecedented accuracy in the near future. However, it is crucial to ensure that uncertainties inherent in the gravitational-wave models will not lead to systematic biases when multiple detections are combined. To quantify waveform systematics, we perform an extensive simulation campaign of binary neutron-star sources and analyse them with a set of four different waveform models. We find that statistical uncertainties in the neutron-star radius decrease to $\pm 250m$ (2% at 90% credible interval), while systematic differences among currently employed waveform models can be twice as large emphasizing the need for waveform models with increased accuracy.

GR 9.7 Wed 18:15 GR-H2

Neutrino and viscosity effects on binary neutron star dynamics, gravitational waves and emitted material — •FRANCESCO ZAPPA — Friedrich-Schiller-Universität Jena Theoretisch-Physikalisches Institut

We present a multi-resolution and multi-physics comparison of simulations performed with the WhiskyTHC code of a single binary*neutron star merger with component neutron star mass of 1.3 M_{\odot} using the finite-temperature equation of state SLy4. Our simulations set consists of pure General Relativistic Hydrodynamic evolution; simulations with a leakage scheme for neutrino production; runs which include M0 treatment for the propagation of free-streaming neutrinos; simulations which include an effective treatment for magnetic-driven turbulent viscosity; simulations which make use of the THC_M1 scheme recently implemented in the code for neutrino treatment. We find that the effect of resolution is dominant with respect to the different physics simulated regarding the gravitational wave production and the merger dynamics, as well as on the onset of black hole collapse. The post-merger disk mass, the disk composition and its thermodynamic properties are instead affected by the production of neutrinos or the presence of turbulent viscosity. The ejected mass and its proton fraction has a very strong dependence on the neutrino treatment employed. Our higher resolution runs confirm that the proton-richest matter is produced when employing M1, at small angles from the orbital plane.

GR 10: Foundations and Alternatives

Time: Wednesday 16:15-18:35

GR 10.1 Wed 16:15 GR-H3

Lorentzianische Relativität — •ALBRECHT GIESE — Taxusweg 15, 22605 Hamburg

Warum lorentzianisch? Die Relativität von Einstein beruht auf der Annahme, dass die gemessene Konstanz der Lichtgeschwindigkeit eine physikalische Realität ist, kein bloßes Messergebnis. Sie führt jedoch zu Komplikationen. Sie erforderte den Ansatz einer verwobenen Raum-Zeit, welche die viel einfachere euklidische Geometrie ersetzten musste. Sie (Einsteins Ansatz) führt dabei zu logischen Konflikten, sobald es nicht um lineare Bewegung, sondern um Drehung geht. Einstein hat diese Konflikte sogar gegenüber seinem Kollegen Lorentz eingeräumt. Doch hat er nie eine echte Lösung dafür angeboten.

Folgt man dem Ansatz von Lorentz, werden sowohl die mathematische Behandlung als auch die Vorstellbarkeit grandios einfacher. Die logischen Konflikte bei Einstein werden vermieden. Offene Probleme der heutigen RT wie vor allem die dunkle Materie und dunkle Energie entfallen gänzlich. Dabei sind die Ergebnisse die gleichen wie bei Einstein, sowohl für die spezielle als auch die allgemeine Relativität.

Further info: www.ag-physics.org/relat

GR 10.2 Wed 16:35 GR-H3

Relativitätstheorie 2.0 — •RALF R. LENKE — 73466 Lauchheim Das Relativitätsprinzip ist eine fundamentale Eigenschaft unseres Universums. Albert Einstein leitete daraus auf 'mathematischem' Weg seine Relativitätstheorie (RT) ab. Bis heute ungeklärt ist allerdings, warum unser Universum dem Relativitätsprinzip folgt. Bis heute gibt es dafür kein in sich geschlossenes, 'physikalisch anschauliches' Modell.

Das grundlegende Experiment für die RT war das Michelson-Morley-Experiment (MME). Aus ihm leitet sich u.a. die 'Lichtuhr' ab, als anschauliche Erklärung für die relativistische Zeitdilatation. Aber darüber hinaus gibt es keine weiteren, wirklich anschaulichen Erklärungen; weder für die Gleichzeitigkeit, und schon gar nicht für die relativistische Längenkontraktion. Diese Asymmetrie in der Anschauung sollte eigentlich verwundern, angesichts der ansonsten herrschenden Symmetrie in Raum und Zeit, der sogenannten 4-dimensionalen Raumzeit.

Tatsächlich steckt im MME mehr: Eine rigorose Betrachtung des MME mit den Mitteln der klassischen Optik führt zum 'seitlich driftenden' Lichtstrahl, welcher Zeitdilatation, Längenkontraktion und Gleichzeitigkeit auf einen Blick offensichtlich werden lässt. Daraus folgt unmittelbar, dass die spezielle RT eine direkte Folge des Wellencharakters der Natur ist und die Lorentz-Transformationen nichts weiter sind als der Doppler-Effekt.

Diese Arbeit geht weit über einen didaktischen Aspekt hinaus. Relativitätstheorie und Quantenmechanik erscheinen unter einem ganz neuen Blickwinkel. Weitere Informationen unter: https://die-neue-relativitaetstheorie.de/

GR 10.3 Wed 16:55 GR-H3

Noether's Theorem and Gravitation in Multi-Particle Systems — •WALTER SMILGA — Geretsried, Deutschland

According to Noether's theorem, the total momentum of an isolated multiparticle system (of massive particles) is conserved as consequence of translational invariance. In agreement with the inverse of Noether's theorem, the individual particle momenta are in general not conserved but time-dependent. The conservation of the total momentum constrains the temporal variability of the particle momenta. Constraints on the particle momenta imply constraints on the particle trajectories. The constrained particle trajectories define a pseudo-Riemannian spacetime described by the field equations of Conformal Gravity. This spacetime is 'quantised' without problems, leading to a consistent quantum gravity.

GR 10.4 Wed 17:15 GR-H3

Theses for a Closed, Self Sustaining and Timeless Universe — •THOMAS Wäscher — IBW Engineering, 69231 Rauenberg

To overcome the persistent non-detectability of presumed Dark Matter and Dark Energy theses are given which implicate the need for a strong paradigm shift in the standard model of cosmology.

1. The universe is curved and closed by its own gravity, it is in the mean homogenious, isotropic and adiabat (Einstein universe) with a constant all over visible CMB-event horizon short before R = c/H [m].

2. The universe is dominated by a stabilizing internal process cycle of the stationary power emission of radiation $P = c^5/2G$ [W] balanced by the equivalent mass flow of $\dot{m} = c^3/2G$ [kg/s]

3. Normalised by the lookback time t = r/c the Hubble eqn. v = Hr results in $|a| = Hc \text{ [m/s}^2$], a scalar background field creating gravity deviations e.g. rotation anomalies.

4. Crossing the universe the radiation is subject to redshift z by gathering the grav. potential $z = \Phi/\Phi o = ar/c^2$ (linear up to $z \approx 0.1$, relativistic $z = ((1 + ar/c^2))(1 - ar/c^2))^{0.5} - 1)$

Location: GR-H3

5. The understanding of the Hubble parameter *H* changes from a velocity per distance (\approx 70km/sMpc) to a decay constant $H \approx 2.27 \times 10^{-18}$ [s⁻¹], which might be the real origin of gravity.

6. Theses 1-5 demand recycling of elements Z>1 to generate fresh hydrogen. Most probably this happens in the core of millions of neutron stars (i.e. pulsars) where nuclei can be unlocked down to the scale of quarks by the extreme gravity. As observed for long, huge clouds of H are outflowing bidirectional from the disc level of galaxies.

GR 10.5 Wed 17:35 GR-H3 Zur Sommerfeld Feinstrukturkonstante aus der Sicht Einsteins — •GEILHAUPT MANFRED — Mönchengladbach, Webschulstr. 31

Einstein (1923) zu seinem Kollegen Dirac: "Eine Theorie die Ladung und Masse des Elektrons a priori setzt ist unvollständig." Das gilt nun seit 100 Jahren sowohl für die ST also auch für die ART. Dirac (1965 in Scientific American) mit seiner Hypothese, Elektron-Masse (m) und Ladung (e) und Lichtgeschwindigkeit (c) sind DIE fundamentalen Naturkonstanten, die Gravitationskonstante (G) und die Planck-Konstante (h) aber nicht, hatte mit diesen Annahmen keine Chance, eine Lösung zu finden. Wenn man die ART mit den Prinzipien der TD verbindet, liefert die Bewegungsgleichung der ART für ein RUHENDES Elektron, sowohl Masse, Ladung und überraschenderweise auch die Sommerfeld FSK (siehe DPG 2019). Es zeigt sich, dass diese FSK auch mit der Ruhemasse in Beziehung steht und nicht nur mit der Elementarladung. Darüberhinaus hängt die FSK von der Metrik des Einstein-Raumes ab. Wenn Parker (1/137.035999048(27)) und Morel (1/137.035999206(11)) beide im Januar und im Juli ihre Messungen wiederholen, wird der Unterschied in der Metrik aufgrund des Sonnenstandes um +-(46) variieren. (The last word has the experiment.)

GR 10.6 Wed 17:55 GR-H3

Explanation of Quantum Physics by Gravity and Relativity — •HANS-OTTO CARMESIN — Gymnasium Athenaeum, Harsefelder Straße 40, 21680 Stade — Universität Bremen, Fachbereich 1, Postfach 330440, 28334 Bremen — Studienseminar Stade, Bahnhofstraße 5, 21682 Stade

Since Planck discovered quantization in 1900, the nature of quanta was a mystery. That problem has now been resolved [1]. For it, I derived the postulates of quantum physics from the equivalence principles, gravity and relativity, whereby I analyzed the vacuum.

Using that derivation, I explain many quantities and properties of quantum physics in a precise manner. Examples are the nature of nonlocality, the physical quantity corresponding to the wave function ψ , the mathematical transformation describing the particle wave duality and the origin of the dynamics inherent to the Schrödinger equation. Moreover, I propose and derive the generalized Schrödinger equation. Especially, I solve the EPR paradox. Furthermore, I identify the physical basis of the Planck constant *h*.

Altogether, quantum physics has now been derived, explained and extended in a direct and transparent manner on the basis of space, time and gravity. As an additional and global test of that explanation, I derive the density parameter Ω_{Λ} of the vacuum or of dark energy in the universe by using the wave function ψ . The result is in precise accordance with observation, whereby I do not apply any fit.

[1] Carmesin, H.-O. (February 2022): Explanation of Quantum Physics by Gravity and Relativity. Berlin: Verlag Dr. Köster.

GR 10.7 Wed 18:15 GR-H3

5th edition of "Special and general theory of relativity for ..." – \bullet Jürgen Brandes – Karlsbad, Germany

Exactly and comprehensibly are discussed in [1]: The experimental proofs of relativity theory, the solutions of the paradoxes and the EINSTEIN- and LORENTZ-interpretation of special and general relativity. Included are the twin paradox and the paradoxes of BELL, EHRENFEST and SAGNAC.

THORNE, Nobel Prize 2017, calls these interpretations the *curved spacetime paradigm* and the *flat spacetime paradigm* and states: "It is extremely useful, in relativity research, to have both paradigms at one's fingertips." Consenting, both interpretations are discussed in equal rights.

An important point in [1] concerns energy conservation. Within NEWTON's theory there is a negative gravitational potential, on account of the famous relation $E = mc^2$ this means negative masses. Negative masses don't exist. Neither NEWTON's nor EINSTEIN's theory can explain the meaning of the negative energy of particles resting in the gravitational field. Additionally, in certain limiting cases there exist contradictious formulas of total energy. In both cases LORENTZ-interpretation gives a clear, experimentally verifiable answer.

[1] J. Brandes, J. Czerniawski, L. Neidhart: Spezielle und Allgemeine Relativitätstheorie für Physiker und Philosophen - Einstein- und Lorentz-Interpretation, Paradoxien, Raum und Zeit, Experimente, 5th edition, VRI: 2022. [2] homepage www.grt-li.de

GR 11: Relativistic Astrophysics

Location: GR-H2

Time: Thursday 11:00-12:40

GR 11.1 Thu 11:00 GR-H2

Constraining supranuclear-dense matter with nuclear physics and multimessenger astrophysics — •TIM DIETRICH — Institute of Physics and Astronomy, University of Potsdam, Potsdam, Germany — Max Planck Institute for Gravitational Physics (Albert Einstein Institute), Potsdam, Germany

Our knowledge about dense matter explored in the cores of neutron stars remains limited. Fortunately, the detection of gravitational waves emitted from the merger of neutron stars and the corresponding electromagnetic signals allows us to place constraints on the properties of matter at supranuclear densities. Furthermore, such densities are not only probed in astrophysical observations, but also in terrestrial experiments. In this work, we use Bayesian inference to combine a large set of data from astrophysical multi-messenger observations of neutron stars and from heavy-ion collisions of gold nuclei at relativistic energies with microscopic nuclear theory calculations. Our findings show that constraints from heavy-ion collision experiments show a remarkable consistency with multi-messenger observations and provide complementary information on nuclear matter at intermediate densities. This work combines nuclear theory, nuclear experiments, and astrophysical observations from multiple messengers. Such joint analyses will be the key to shed light on the properties of neutron-rich supranuclear matter over the density range probed in neutron stars.

GR 11.2 Thu 11:20 GR-H2

GR 12.1 Thu 11:00 MP-H5

Tidally-driven crustal failure in coalescing binaries of neutron stars as triggers for precursor flares of short gamma-ray burst — •HAO-JUI KUAN^{1,2}, ARTHUR SUVOROV³, and KOSTAS KOKKOTAS¹ — ¹Theoretical Astrophysics, IAAT, University of Tübingen, Tübingen, D-72076, Germany — ²National Tsing Hua University, Hsinchu 300, Taiwan — ³Manly Astrophysics, 15/41-42 East Esplanade, Manly, NSW 2095, Australia

In some short gamma-ray bursts, precursor flares occurring ~ seconds prior to the main episode have been observed. These flares may then be associated with the last few cycles of the inspiral when the orbital frequency is a few hundred Hz. During these final cycles, tidal forces can resonantly excite quasi-normal modes in the inspiralling stars, leading to a rapid increase in their amplitude. It has been shown that these modes can exert sufficiently strong strains onto the neutron star crust to instigate yieldings. Among other possible modes, the typical frequencies of *g*-modes being ~ 100 Hz warrant further investigation since their resonances with the orbital frequency match the precursor timings. Adopting realistic equations of state (EOS) and solving the general-relativistic pulsation equations, we study *g*-mode resonances in coalescing quasi-circular binaries, where we consider various stellar rotation rates, degrees of stratification, and magnetic field structures. We show that for some combination of stellar parameters, the resonantly excited g_1 - and g_2 -modes may lead to crustal failure and trigger precursor flares, and also some EOS are more likely to cause the crustal cracks.

GR 11.3 Thu 11:40 GR-H2 Non-thermal electromagnetic counterparts to binary neutron star mergers and supernovae — •VSEVOLOD NEDORA — Max-Planck-Institut für Gravitationsphysik, Potsdam, Germany — Universität Potsdam, Potsdam, Germany In 2017 the merger of two neutron stars was observed in gravitational waves and across the electromagnetic spectrum. The event was associated with the short gamma-ray burst (GRB) GRB170817A, the afterglow from which peaked 160 days after the burst, before starting to decay in agreement with short GRB models. Much later, 1234 days after the burst, a change in the GRB170718A light curve behavior was observed, inconsistent with most GRB models.

We design a new numerical tool to model the synchrotron radiation arising from interactions between the ejected matter and the interstellar medium, employing the recent theoretical advancements in our understanding of relativistic jetted and oblate mildly relativistic outflows.

Assuming that the change of the afterglow is due to an emerging new component (e.g., kilonova ejecta), we investigate the properties of this component by modeling kilonova and GRB afterglows simultaneously. This allows us to take advantage of all the gathered data and obtain new constraints on the dynamics of the ejected matter and NS EOS.

GR 11.4 Thu 12:00 GR-H2

Fast Rotating Relativistic Stars: Spectra and Stability without Approximation — •CHRISTIAN KRUEGER and KOSTAS KOKKOTAS — Universitaet Tuebingen, Auf der Morgenstelle 10, 72076 Tuebingen

The oscillations and instabilities of relativistic stars are studied by taking into account, for the first time, the contribution of a dynamic space-time. The study is based on the linearised version of Einstein's equations and via this approach the oscillation frequencies, the damping and growth times as well as the critical values for the onset of the secular (CFS) instability are presented. The ultimate universal relations for asteroseismology are derived which can lead to relations involving the moment of inertia and Love numbers in an effort to uniquely constrain the equation of state via all possible observables. The results are important for all stages of neutron star's life but especially to nascent or post-merger cases.

GR 11.5 Thu 12:20 GR-H2

Exploring black holes as particle accelerators: hoop-radius and escaping conditions — STEFANO LIBERATI², •CHRISTIAN PFEIFER¹, and JAVIER RELANCIO³ — ¹University of Bremen, Bremen, Germany — ²SISSA and INFN, Trieste, Italy — ³Università di Napoli Federico II, Napoli, Italy

The possibility that rotating black holes could be natural particle accelerators has been subject of intense debate. While it appears that for extremal Kerr black holes arbitrarily high center of mass energies could be achieved, several works pointed out that both theoretical as well as astrophysical arguments would severely dampen the attainable energies. In this talk I study particle collisions near Kerr black holes, in particular collision between an infalling particle from infinity and a target particle that is already in the near vicinity of the black hole horizon. Most importantly, I will show how to implement the hoop conjecture and discuss the astrophysical relevance of these target particle collisions can be ultra high. Moreover, I discuss that for nearly extremal black holes, the energy of the particle secaping the black hole region can be similarly large. Thus, these target particle collisional Penrose processes could contribute to the observed spectrum of ultra high-energy cosmic rays, even if the hoop conjecture is taken into account.

GR 12: Quantum gravity (joint session MP/GR)

Time: Thursday 11:00-12:40

Invited Talk

Reduced phase space quantisation in Loop quantum gravity and loop quantum cosmology — •KRISTINA GIESEL¹, BAO-FEI LI², and PARAMPREET SINGH² — ¹Institute for Quantum Gravity, Department of Physics, FAU Erlangen-Nürnberg, Staudtstr. 7, 91058 Erlangen, Germany — ²Department of Physics and Astronomy, Louisiana State University, Baton Rouge, LA 70803, USA

In this talk an overview over results that apply a reduced phase space quantisation to formulate the dynamics of loop quantum gravity and loop quantum cosmology will be presented. It will be briefly discussed how a reduced phase space for GR can be derived by coupling additional reference matter. The reduced phase space for GR is taking as a starting point for a loop quantisation either in full LQG or in the symmetry reduced case of LQC. Different choices of reference matter yield in general different quantum models and several existing models will be compared. In the framework of LQC it will be analysed how different choices of reference matter can lead to different physical properties of the models and it will be discussed what kind of conditions appropriate reference matter should satisfy in this context. Invited Talk

GR 12.2 Thu 11:30 MP-H5

Location: MP-H5

Quantum fields propagating on curved backgrounds and their influence on spacetime curvature – •NICOLA PINAMONTI – Department of Mathematics, University of Genova, Italy

We shall review the theory of quantum fields propagating on curved backgrounds in the semiclassical approximation. Within this approximation matter is described by quantum fields propagating on a classical curved spacetime and their influence on spacetime curvature is taken into account by means of semiclassical Einstein equations. Typical known effects in this approximation are particle creation on cosmological spacetime, Hawking radiation in the case of black hole backgrounds and their evaporation. The semiclassical analysis we would like to present requires a careful study of the form of the correlation functions of the state which describes matter. The thorough analysis of their ultraviolet divergences and their renormalization is necessary to obtain meaningful expressions for the expectation values of the matter stress energy tensor. The resulting stress energy tensor will have nontrivial effects on the curvature on cosmological spacetimes too. In the latter case, semiclassical Einstein equations give origin to a well posed dynamical system provided the quantum state for matter is chosen

54

in an appropriate way. The question of existence of exact solutions of such system will be discussed and some implication for cosmology will be presented. In the case of black hole physics, exploiting the stress tensor properties, we will give a local model of black hole evaporation.

GR 12.3 Thu 12:00 MP-H5

A master equation for gravitationally induced decoherence of a scalar field — •MAX JOSEPH FAHN, KRISTINA GIESEL, and MICHAEL KOBLER — Institute for Quantum Gravity, Department of Physics, FAU Erlangen-Nürnberg, Staudtstr. 7, 91058 Erlangen, Germany

In the talk, we present the derivation of a decoherence model containing a scalar field coupled to a gravitational environment. With such a model, one can predict quantum gravity effects in the scalar field's dynamics that arise due to the presence of gravity. Starting with full general relativity in Hamiltonian form expressed in Ashtekar's connection formulation, we focus on weak gravitational interactions in an asymptotically flat universe, modelled by gravitational waves propagating on a fixed background, and the scalar field as matter component. Firstly, we construct appropriate observables (gauge invariant quantities) for the model which become elementary variables in the reduced phase space of the model. Afterwards a reduced phase space quantisation using Fock quantisation is performed. With the help of the projection operator technique, we derive a second order time-convolutionless master equation, that is an effective evolution equation which predicts the temporal evolution of the scalar field, including the gravitational interaction effectively in terms of certain operators, whose form is a result of the model under consideration and several physical assumptions. These additional terms lead to physical effects like dissipation or decoherence of the matter field induced by gravity. Finally, we briefly discuss possible applications of the model's master equation.

GR 12.4 Thu 12:20 MP-H5 Quantum (supersymmetric) black holes in loop quantum gravity — •KONSTANTIN EDER — FAU Erlangen-Nürnberg

Black holes are immediate and unavoidable consequences of Einstein's theory of gravity whose existence nowadays has been confirmed with remarkable accuracy. Albeit leading to a huge success of the theory, there are also various open questions that point out its incompleteness. One of them is about the huge amount of microstates needed to explain the entropy of black holes as predicted by Bekenstein and Hawking: What are these micostates? Answering it poses a challenge for any formulation for quantum theory of gravity. In loop quantum gravity a very intriguing picture has been developed that suggests an answer in terms of topological Chern-Simon degrees of freedom induced by quantized geometry on the horizon. In this talk, we first give a general review on the description of black hole horizons in LQG. Then, we will give an outlook on recents results towards their generalization to the supersymmetric context. There, it turns out, using tools from super Cartan geometry, that the unique boundary theory in chiral supergravity corresponds to that of a super Chern-Simons theory. This enables one to transfer ideas from the bosonic to the supersymmetric setting.

GR 13: Numerical Relativity

Time: Thursday 14:00-15:40

GR 13.1 Thu 14:00 GR-H2 GRHD simulations with GR-Athena++ — •WILLIAM COOK — Theoretisch-Physikalisches Institut, Friedrich-Schiller Universitat Jena, 07743, Jena, Germany

For the first time we demonstrate the performance of the code GR-Athena++ in evolving general relativistic hydrodynamics (GRHD) in a dynamically evolving spacetime. GR-Athena++ utilises the task-based parallelism and block based adaptive mesh refinement of the Athena++ code, as well as its approach to solving GRHD problems in stationary spacetimes; combined with new functionality to solve the Einstein equations in the Z4c formulation. We demonstrate the performance of this new code by simulating the evolution of Neutron Stars in GR-Athena++, removing the Cowling approximation assumed in previous work, presenting a fully dynamical spacetime evolution.

GR 13.2 Thu 14:20 GR-H2

Evolution of mixed binaries initial data produced with Elliptica — •FRANCESCO MARIA FABBRI¹, ALIREZA RASHTI², BERND BRÜGMANN¹, SWAMI VIVEKANANDJI CHAURASIA³, TIM DIETRICH^{4,5}, MAXIMILIANO UJEVIC⁶, and WOLFGANG TICHY² — ¹Theoretical Physics Institute, University of Jena, 07743 Jena, Germany — ²Department of Physics, Florida Atlantic University, Boca Raton, FL 33431, USA — ³The Oskar Klein Centre, Department of Astronomy, Stockholm University, AlbaNova, SE-10691 Stockholm, Sweden — ⁴Institut für Physik und Astronomie, Universitä Potsdam, Haus 28, Karl-Liebknecht-Strasse 24/25, 14476, Potsdam, Germany — ⁵Max Planck Institute for Gravitational Physics (Albert Einstein Institute), Am Mühlenberg 1, Potsdam 14476, Germany — ⁶Centro de Ciências Naturais e Humanas, Universidade Federal do ABC, 09210-170, Santo André, São Paulo, Brazilx

In this work, we present the evolution of black hole - neutron star initial data produced with the new pseudo spectral code Elliptica. The code makes use of Schur complement domain decomposition method with a direct solver using cubed sphere coordinate maps and the fields are expanded using Chebyshev polynomials of the first kind. To stress the capability of the code we designed different configurations of mass ratios and spin orientations and we made use of the BAM code to evolve such initial data. Arbitrary spin magnitudes and orientations are evolved, with a maximum adimensional spin of 0.8 for the black hole and an arbitrary high spin for the neutron star.

GR 13.3 Thu 14:40 GR-H2

Comparison of eccentric numerical relativity simulations to small mass-ratio perturbation theory — •ANTONI RAMOS-BUADES, MAARTEN VAN DE MEENT, and HARALD PFEIFFER — Max Planck Institute for Gravitational Physics, Potsdam, Germany

During the third observing run of the LIGO and Virgo detectors a few gravitational wave (GW) signals from binary black hole (BBH) mergers with unequal Location: GR-H2

masses have been detected. As detectors' sensitivity continues to increase, more systems with more asymmetric masses are expected to be detected, and therefore modelling of BBHs at all mass ratios is of preeminent relevance. In this work we compare two approaches to modeling binary black holes (BBHs): 1) small mass-ratio (SMR) perturbation theory, and 2) numerical relativity (NR). We extend recent work on combining information from quasicircular nonspinning NR simulations of BBHs with results from SMR perturbation theory to nonspinning eccentric BBHs. We produce a dataset of long and accurate eccentric nonspinning NR simulations with the Spectral Einstein Code (SpEC) from mass ratios 1 to 10, and eccentricities up to 0.7. We analyze these NR simulations, compute gauge invariant quantities from the gravitational radiation, and develop tools to map points in parameter space between eccentric NR and SMR waveforms. Finally, we discuss discrepancies between SMR and NR predictions for the energy and angular momentum fluxes due to eccentricity, and limitations of such comparisons due to the limited parameter space in mass ratio covered by the NR simulations.

GR 13.4 Thu 15:00 GR-H2 **Critical Collapse with bamps.** — DANIELA CORS AGULLÓ¹, •SARAH RENKHOFF¹, ISABEL SUÁREZ FERNÁNDEZ², HANNES RÜTTER³, DAVID HILDITCH², and BERND BRÜGMANN¹ — ¹Theoretisch-Physikalisch Institut, Friedrich-Schiller-Universität, Jena, Germany — ²CENTRA, Instituto Superior Tecnico, Lisbon, Portugal — ³Max-Planck-Institut für Gravitationphysik, Potsdam, Germany

Our pseudo-spectral code bamps, with its new adaptive mesh refinement, allows us to tune closer to the critical point between gravitational collapse and dispersed fields. We study, on the one hand critical phenomena in spherical symmetry by evolving scalar fields using generalised harmonic gauge and compare among several gauge source functions. On the other hand, we can assess critical phenomena in an axisymmetric setting by evolving gravitational waves in vacuum. We evolve six one-parameter families of Brill wave initial data: three prolate and three oblate, including two centred and four off-centred. Time permitting, we will discuss the relevance of our results in the context of critical collapse beyond spherical symmetry.

GR 13.5 Thu 15:20 GR-H2

A new approach to helical Killing vectors — •HANNES RÜTER — Potsdam, Germany

Helical Killing Vectors are an important ingredient for the description of binary systems in quasi-equilibrium, which makes them one of the central quantities in the construction of numerical-relativity initial data of compact binaries. Current approaches to helical Killing vectors and their generalisations to eccentric orbits are defined in a foliation-dependent manner. In this presentation I will a discuss a new approach that attempts to break this foliation dependence.

GR 14: Gravitational Wave Detectors

Time: Thursday 16:15-18:15

GR 14.1 Thu 16:15 GR-H2

Options for ultra-high-frequency gravitational wave detection at DESY -Axel Lindner, Krisztian Peters, •Christoph Reinhardt, Andreas Ring-WALD, and AARON SPECTOR — DESY, Hamburg, Germany

Fueled by recent achievements of km-scale gravitational-wave observatories, efforts towards the development of "down to lab-scale" detectors are currently gaining momentum. By virtue of their smaller dimensions, these detectors target gravitational-waves above the established observation window (i.e., 10 to 10^4 Hz) at "ultra-high-frequencies" (i.e., 10^4 to 10^{19} Hz). The sources proposed for generating potentially observable signals in this frequency range are of cosmological origin or require new physics beyond the Standard Model.

Current work at DESY towards three on-site experiments for axion searches (ALPS II, BabyIAXO, and MADMAX) provides the unique opportunity to exploit synergies with regard to the development of ultra-high-frequency gravitational wave detectors: On the one hand, these axion experiments are also sensitive to gravitational waves via "graviton-to-photon conversion" in the presence of a magnetic field. On the other hand, the underlying infrastructure of the axion experiments, comprising, e.g., a distribution system for liquid helium, can be a prerequisite for implementing dedicated detectors. This talk will present options for ultra-high-frequency gravitational wave searches at DESY, Hamburg.

GR 14.2 Thu 16:35 GR-H2

The implementation of stray light simulation tool for the Einstein Telescope - •Наппа Макоzava and Тномая Bretz — RWTH Aachen University

The Einstein Telescope will be the first gravitational wave detector of the third generation. Stray light is a severe problem for modern interferometers with high sensitivity, as another noise source contributing to the interferometer output. Simulation is requied to optimize detector settings to avoid undesirable light paths. Stray light simulation can be done with special optical analysis software, but existing tools are not ideally suited for general application, or require expensive licences. Therefore ET optical simulation task is still urgent.

The goal of this work is development of dedicated free tool for stray light simulation, easy to use and including higher precision and different aspects of stray light. Simulation involves modelling the system design, calculating the intersection points of the rays with the surfaces and simulating of stray light paths, and algorithms for some of this steps are already found.

This talk will present the progress of stray light simulation tool development, including the analysing of existing optical software, implementation of simulation and first results.

GR 14.3 Thu 16:55 GR-H2

Development and testing of composite vacuum tubes for Einstein Telescope — • Purnalingam Revathi¹, Ralf Schleichert², Achim Stahl¹, Tim KUHLBUSCH¹, ROBERT JOPPE¹, and OLIVER POOTH¹ — ¹III. Physikalisches Institut B, RWTH Aachen — ²Institut für Kernphysik, Forschungszentrum Jülich The Einstein Telescope, a proposed third generation gravitational wave detector, requires 120 km long vacuum tubes in total with a diameter of 1 m. Due to the vacuum requirements and mechanical integrity, stainless steel tubes are the standard for ultra high vacuum applications. Composite vacuum tubes are a promising alternative that could reduce material cost and opens possibilities for easier on site production. This talk presents the details of development and testing of prototypes made of Glass Fiber Reinforced Plastic wound around a stainless

steel liner. Pressure stability and aging tests are performed and the possibility of integrating sensors to measure temperature and pressure will be discussed.

GR 14.4 Thu 17:15 GR-H2

Test setup for cryogenic sensors and actuators working towards the Einstein **Telescope** — •Robert Joppe¹, Tim Kuhlbusch², Thomas Hebbeker¹, Vivek Pimpalshende², Oliver Pooth², Achim Stahl², and Franz-Peter Zantis¹ ¹III. Physikalisches Institut A, RWTH Aachen — ²III. Physikalisches Institut B, RWTH Aachen

The Einstein Telescope will be the first gravitational wave detector of the third generation. The sensitivity goal, especially in the low frequency region, will be achieved among other improvements by cooling the main parts of the interferometer. The required electronic components, sensors and actuators needed for mirror alignment and active dampening of suspension resonances have to perform at cryogenic temperatures.

The talk presents the progress on the development of electronics and mechanics within the E-TEST project. The performance of our cryogenic ultra high vacuum test setup will be furthermore explicated.

GR 14.5 Thu 17:35 GR-H2 A cryogenic displacement sensor and actuator for the E-TEST project – •Kuhlbusch Tim^1 , Thomas Hebbeker², Robert Joppe², Vivek Pimpalshende¹, Oliver Pooth¹, Achim Stahl¹, and Franz-Peter Zantis² ¹III. Physikalisches Institut B, RWTH Aachen — ²III. Physikalisches Institut A, RWTH Aachen

The Einstein Telescope will be the first third generation gravitational wave detector. In achieving a sensitivity increase of more than one order of magnitude at low frequencies compared to current detectors, mitigating thermal noises is essential. Cooling the mirrors to cryogenic temperatures is thus required. This also creates the need for parts of the vibration isolations systems of the mirrors to be working at low temperatures.

This talk will present the developement of an actuator with an integrated absolute displacement sensor operating below 20 K as part of the E-TEST project. The sensitivity of the sensor, the forces of the actuator and the thermal design will be discussed.

GR 14.6 Thu 17:55 GR-H2

Commissioning and characterization of a shadow sensor - •Vivek Pimpalshende¹, Thomas Hebbeker², Robert Joppe^{1,2}, Tim Kuhlbusch¹, $\mathsf{Oliver}\,\mathsf{Pooth}^1,\mathsf{Achim}\,\mathsf{Stahl}^1,\mathsf{and}\,\mathsf{Franz-Peter}\,\mathsf{Zantis}^2-{}^1\mathsf{III}.\,\mathsf{Physikalis}$ ches Institut B, RWTH Aachen — ²III. Physikalisches Institut A, RWTH Aachen The Einstein Telescope will be the first gravitational wave detector of the third generation. It will extend the detection sensitivity in the low frequency domain by more than an order of magnitude compared to current second generation detectors. An essential factor in achieving the sensitivity goals is to mechanically decouple the optic system from the ground. For this, a seismic isolation system is in place in the interferometric gravitational wave detectors. This talk will discuss one of the essential position sensors in this system: a shadow sensor.

Displacement sensors are a part of a control loop that monitors the relative motion of the components of the suspension system and controls the absolute motion of these components. This talk will present the commissioning of a shadow sensor along with its characteristics at room temperature.

GR 15: Member Assembly

Time: Thursday 19:00-20:30 **MV-GR Member Assembly**

Location: GR-H2

Theoretical and Mathematical Physics Division Fachverband Theoretische und Mathematische Grundlagen der Physik (MP)

Johanna Erdmenger Julius-Maximilians-Universität Würzburg Am Hubland, 97074 Würzburg

Overview of Invited Talks and Sessions

(Lecture halls MP-H4 and MP-H5)

Plenary Talk of MP

PV V	Wed	9:45-10:30	Audimax	Quantum gravity, chaos, statistical physics and wormholes — •JAN DE BOER
------	-----	------------	---------	--

Invited Talks

MP 1.1	Mon	9:30-10:00	MP-H5	Long-range entanglement and the split property $- \bullet$ Pieter Naaijkens, Yoshiko
MD 1 2	Mon	10.00 10.30	MD 115	UGATA Symmetry resolved quantum information measures for AdS gravity and beyond
WIF 1.2	WIOII	10.00-10.50	IVIT-115	•René Meyer, Suting Zhao, Christian Northe, Konstantin Weisenberger
MP 1.3	Mon	10:30-11:00	MP-H5	Rapid thermalization of spin chain commuting Hamiltonians — •ANGELA CAPEL
MP 2.1	Mon	16:15-16:45	MP-H5	The Page curve from quantum error correction — • DANIEL HARLOW
MP 2.2	Mon	16:45-17:15	MP-H5	Classical black hole scattering from a worldline quantum field theory — •JAN PLEFKA
MP 3.1	Tue	11:00-11:30	MP-H5	Dualities and categorical symmetries in quantum spin chains — •FRANK VERSTRAETE
MP 3.2	Tue	11:30-12:00	MP-H5	Functional Integral and Stochastic Representations for Bosonic Ensembles -
				•Manfred Salmhofer
MP 3.3	Tue	12:00-12:30	MP-H5	Color-Flavor Transformation Revisited — •MARTIN ZIRNBAUER
MP 5.1	Wed	11:00-11:30	MP-H5	Falling through masses in superposition: quantum reference frames for indefinite met-
				rics — Anne-Catherine de la Hamette, Viktoria Kabel, Esteban Castro-Ruiz,
				•Caslav Brukner
MP 6.1	Wed	16:15-16:45	MP-H5	State sum models with defects — • CATHERINE MEUSBURGER
MP 9.1	Thu	11:00-11:30	MP-H5	Reduced phase space quantisation in Loop quantum gravity and loop quantum cos-
				mology — •Kristina Giesel, Bao-Fei Li, Parampreet Singh
MP 9.2	Thu	11:30-12:00	MP-H5	Quantum fields propagating on curved backgrounds and their influence on spacetime
				curvature — •Nicola Pinamonti

Invited Talks of the joint symposium SMuK Dissertation Prize 2022 (SYMD)

See SYMD for the full program of the symposium.

SYMD 1.1	Mon	14:00-14:25	Audimax	Timeless Quantum Mechanics and the Early Universe — •LEONARDO CHATAIGNER
SYMD 1.2	Mon	14:25-14:50	Audimax	First tritium β -decay spectrum recorded with Cyclotron Radiation Emission
				Spectroscopy (CRES) — • CHRISTINE CLAESSENS
SYMD 1.3	Mon	14:50-15:15	Audimax	Watching the top quark mass run - for the first time! — • MATTEO M. DEFRANCHIS,
				Katerina Lipka, Sven-Olaf Moch
SYMD 1.4	Mon	15:15-15:40	Audimax	Towards beam-quality-preserving plasma accelerators: On the precision tuning
				of the wakefield — •SARAH SCHRÖDER

Invited Talks of the joint symposium The Nature of Science (SYNS)

See SYNS for the full program of the symposium.

SYNS 1.1	Tue	14:00-14:30	Audimax	The Role of Nature of Science Education for Science Media Literacy $- \cdot$ DIETMAR HÖTTECKE
SYNS 1.2	Tue	14:30-15:00	Audimax	What kinds of identities are deemed in/our of place in physics? — •LUCY AVRAAMI-
SYNS 1.3	Tue	15:00-15:30	Audimax	DOU Some thoughts on the status of theoretical physics — •DANIEL HARLOW

Prize Talks of the joint Awards Symposium (SYAW) See SYAW for the full program of the symposium.

SYAW 1.1	Wed	14:10-14:40	Audimax	Wie überprüft man die Ziele der Lehramtsausbildung Physik? – •HORST
				Schecker
SYAW 1.2	Wed	14:40-15:10	Audimax	Astronomy at Highest Angular Resolution - Adaptive Optics, Interferometry and
				Black Holes — •Frank Eisenhauer
SYAW 1.3	Wed	15:10-15:40	Audimax	Turbulence in one dimension — •Alexander M. Polyakov

Sessions

MP 1.1-1.3	Mon	9:30-11:00	MP-H5	Entanglement; Thermalization
MP 2.1–2.6	Mon	16:15-18:35	MP-H5	Gravity, Amplitudes, AdS/CFT
MP 3.1–3.3	Tue	11:00-12:30	MP-H5	Quantum dynamics
MP 4.1-4.7	Tue	16:15-18:35	MP-H5	Quantum field theory, AdS/CFT, non-equilibrium and quantum dynamics
MP 5.1-5.4	Wed	11:00-12:30	MP-H5	Quantum information
MP 6.1-6.1	Wed	16:15-16:45	MP-H5	Mathematical physics
MP 7.1–7.6	Wed	16:45-18:45	MP-H5	Quantum field theory
MP 8.1-8.3	Wed	17:25-18:25	MP-H6	Alternative approaches
MP 9.1–9.4	Thu	11:00-12:40	MP-H5	Quantum gravity (joint session MP/GR)
MP 10.1-10.3	Thu	16:15-17:15	MP-H5	Classical and quantum gravity
MP 11.1-11.5	Thu	16:35-18:15	MP-H6	Thermodynamics and fundamental aspects of field theory
MP 12.1-12.3	Thu	17:15-18:15	MP-H5	Quantum field theory for particle physics and plasmas
MP 13	Thu	19:30-20:30	MP-MV	Annual General Meeting

Annual General Meeting of the Theoretical and Mathematical Physics Division

Thursday 19:30-20:30 MP-MV

Sessions

Invited and Contributed Talks –

MP 1: Entanglement; Thermalization

Time: Monday 9:30-11:00

Invited Talk MP 1.1 Mon 9:30 MP-H5 Long-range entanglement and the split property $- \bullet$ PIETER NAAIJKENS¹ and

Yosніко Ogata² — ¹School of Mathematics, Cardiff University, United Kingdom — ²Graduate School of Mathematical Sciences, The University of Tokyo, Japan

Long-range entanglement has been crucial in understanding topologically ordered phases of matter. In particular, it is believed to be a necessary condition for the existence of non-trivial anyons in 2D. In this talk I will explain how we can prove this statement in the context of superselection sector theory. In particular, we show that if a ground state is not long-range entangled, in a way that we make precise, then we only have the trivial superselection sector. I will also indicate how this is related to the split property of certain von Neumann algebras generated by the observables of the system.

Invited Talk MP 1.2 Mon 10:00 MP-H5 Symmetry-resolved quantum information measures for AdS gravity and beyond — •René Meyer, Suting Zhao, Christian Northe, and Konstantin WEISENBERGER - Institut für Theoretische Physik und Astrophysik, Julius-Maximilians-Universität Würzburg, Am Hubland, 97074 Würzburg

Quantum entanglement is the key resource employed in modern quantum computation. Different entanglement measures such as the entanglement entropy and Renyi entropies also provide useful information about the entanglement structure of quantum field theories, in particular at critical points. I will discuss the symmetry resolved entanglement and Renyi entropies, a fine-grained version of the usual entanglement and Renyi entropies, both in the context of quantum field theory and AdS/CFT. In the presence of global conserved charges,

they quantify the entanglement content of the reduced density matrix in a fixed charge sector. These entanglement measures can in particular be calculated in two-dimensional conformal field theories with U(1) Kac-Moody structure at level k, and are found to not depend on the value of the charge. This charge independence is called equipartition of entanglement, and implies that no charge sector is distinguished in terms of its entanglement content. Finally, I will discuss the symmetry resolved entanglement in the AdS3/CFT2 dual of the U(1) Kac-Moody CFT. Agreement with CFT results provides a further test of the AdS3/CFT2 correspondence. I finish with some results about the violation of the equipartition property in CFTs with W3 symmetry. This talk is based on hep-th/2012.11274 and hep-th/2108.09210.

Invited Talk

MP 1.3 Mon 10:30 MP-H5 Rapid thermalization of spin chain commuting Hamiltonians - •ANGELA CAPEL — Universität Tübingen

In this talk, we will show that spin chains weakly coupled to a large heat bath thermalize rapidly at any temperature for finite-range, translation-invariant commuting Hamiltonians, reaching equilibrium in a time which scales logarithmically with the system size. From a physical point of view, this result establishes the absence of dissipative phase transition for Davies evolutions over translationinvariant spin chains. The result also applies in the case of Symmetry Protected Topological phases where the evolution is respecting the symmetry of the phase. We will comment on the possible extensions of this result to higher dimensions, as well as on some applications to the study of many-body in and outof-equilibrium quantum systems.

MP 2: Gravity, Amplitudes, AdS/CFT

Time: Monday 16:15-18:35

Invited Talk

MP 2.1 Mon 16:15 MP-H5 The Page curve from quantum error correction — •DANIEL HARLOW — Massachusetts Institute of Technology, Cambridge, Massachusetts, USA

The "Page curve", meaning the von Neumann entropy of the radiation of a black hole plotted as a function of time, is one of the main tools for diagnosing the black hole information problem. Recently semiclassical calculations of this curve were given, but they have several mysterious features. In particular, it is not so clear how they can be interpreted from a Hilbert space point view, as one would like in a real theory of quantum gravity. In this talk I'll explain how these calculations can be given such an interpretation through a novel generalization of quantum error correction to "non-isometric" codes.

Invited Talk

MP 2.2 Mon 16:45 MP-H5 Classical black hole scattering from a worldline quantum field theory — •JAN PLEFKA — Humboldt Universität zu Berlin, Germany

When two massive, gravitationally interacting bodies (black holes, neutron stars or stars) fly past each other they are deflected and emit gravitational Bremsstrahlung. I shall discuss how this classical two-body problem in general relativity can be efficiently described using worldline quantum field theory methods. Including spin (i.e. Kerr Black Holes or spinning neutron stars) leads to an effective N=2 supersymmetric description valid up to quadratic order in spins. The emitted gravitational waveform, scattering angle and spin-kick may be efficiently computed in a weak gravitational field (post Minkowskian) expansion and represent state of the art results. The worldline quantum field theory approach innovates over traditional approaches and imports modern technology from perturbative quantum field theory to classical perturbative GR.

MP 2.3 Mon 17:15 MP-H5

Sustained convergence of hydrodynamics in rapidly spinning quark-gluon plasma — •Matthias Kaminski¹, Casey Cartwright¹, Markus Garbiso Amano¹, Jorge Noronha², and Enrico Speranza² — ¹Department of Physics and Astronomy, University of Alabama,514 University Boulevard, Tuscaloosa, AL 35487, USA — ²Illinois Center for Advanced Studies of the Universe, Department of Physics, University of Illinois at Urbana-Champaign, Urbana, IL 61801, USA

We compute the radius of convergence of the linearized relativistic hydrodynamic gradient expansion around a non-trivially rotating strongly coupled \mathcal{N} = 4 Super-Yang-Mills plasma. Our results show that the validity of hydrodynamics is sustained and can even get enhanced in a highly vortical quark-gluon plasma, such as the one produced in intermediate-energy heavy-ion collisions. The hydrodynamic dispersion relations are computed using a rotating background that is an analytical solution of the ideal hydrodynamic equations of motion with non-vanishing angular momentum and large vorticity gradients, giving rise to a particular boost symmetry. Analytic equations for the transport coefficients of the rotating plasma as a function of their values in a plasma at rest are given.

MP 2.4 Mon 17:35 MP-H5

Location: MP-H5

Scalar flat holography — •LORENZO IACOBACCI 1,2 and MASSIMO TARONNA 1,2,3 ⁻¹Dipartimento di Fisica Ettore Pancini, Università degli Studi di Napoli Federico II Via Cintia, 80126 Napoli, Italy- 2 Istituto Nazionale di Fisica Nucleare, Sezione di Napoli Via Cintia, 80126 Napoli, Italy – ³Scuola Superiore Meridionale, Università degli Studi di Napoli Federico II, Largo San Marcellino 10, 80138 Napoli, Italy

Recently, flat space holography has played a central role in the study of scattering amplitudes. In particular, the (d+2)-dimensional holographic flat space-time is closed in the far past and in the far future by a d-dimensional boundary, called celestial sphere. On this boundary, a d-dimensional conformal field theory lives; therefore, researchers have been studying the imprint of the d-dimensional conformal symmetry in (d+2)-dimensional scattering amplitudes.

Further, it is well-known that the (d+2)-dimensional Minkowski space-time can be sliced in (d+1)-dimensional AdS spaces inside the light-cone and by (d+1)-dimensional dS spaces outside. This suggests us that a connection between flat, dS and AdS holography should exist.

In our talk, we shall see the strict connection between the holography realized on AdS, dS and Minkowski. By analogy with the AdS case, we will arrive to write down a dictionary to pass between AdS, dS and flat holography. Moreover, these stringent connections will allow us to exploit the results obtained in AdS holography to compute scattering amplitudes both in flat and dS holography.

MP 2.5 Mon 17:55 MP-H5 graphy — •GIUSEPPE DI GIULIO — Julius-

Hyperbolic tilings and discrete holography — •GIUSEPPE DI GIULIO — Julius-Maximilians-Universität Würzburg

The AdS/CFT correspondence is one of the most important breakthroughs of the last decades in theoretical physics. A recently proposed way to get insights on various features of this duality is achieved by discretising the Anti-deSitter spacetime. Within this program, we consider the Poincaré disk (fixed time slice of three-dimensional Anti-deSitter spacetime) discretised by introducing a regular hyperbolic tiling on it. This breaks the isometry group of the Poincaré disk down to a discrete subgroup. In a possible discrete version of the AdS/CFT correspondence, such a discrete group is expected to characterise also the quantum theory living on the boundary of the hyperbolic tiling. In this talk, we discuss the properties that such a boundary theory must have. The bulk tessellation we study induces an aperiodic spatial modulation on any model defined on the boundary. This provides an interesting bridge with the aperiodic quantum chains, wellknown systems in the literature of condensed matter theory. Furthermore, in view of examining possibilities for establishing a duality in this setup, we investigate the entanglement properties of the tiled bulk (obtained through a discretised version of the Ryu-Takayanagi formula for the holographic entanglement entropy) and of some possible boundary theories.

MP 2.6 Mon 18:15 MP-H5

 $\label{eq:Classical and quantum gravitational scattering with Generalized Wilson Lines — Domenico Bonocore, Anna Kulesza, and •Johannes Pirsch — Institut für Theoretische Physik, WWU Münster, Münster, Germany$

The all-order structure of scattering amplitudes is greatly simplified using Wilson line operators, describing eikonal emissions from straight lines extending to infinity. A generalization at subleading powers in the eikonal expansion, known as Generalized Wilson Line (GWL), has been proposed some time ago, and has been applied both in QCD phenomenology and in the high energy limits of gravitational amplitudes. In this talk I discuss the construction of the gravitational GWL starting from first principles in the worldline formalism. This includes identifying the correct Hamiltonian, which leads to a simple correspondence between the soft expansion and the weak field expansion. The resulting path integral representation of the GWL makes it possible to isolate the relevant contributions to the classical limit.

MP 3: Quantum dynamics

Time: Tuesday 11:00-12:30

Invited Talk MP 3.1 Tue 11:00 MP-H5 Dualities and categorical symmetries in quantum spin chains — • FRANK VER-STRAETE — Ghent University

Categorical symmetries play a central role in the characterization of the entanglement features of quantum spin systems exhibiting topological order in 2+1D or criticality in 1+1D. In this talk, we will discuss those symmetries from the point of view of tensor networks, and demonstrate that matrix product operators realize representations of the corresponding module categories. This on its turn allows for a constructive approach to building dualities for quantum spin chains.

Invited Talk MP 3.2 Tue 11:30 MP-H5

Functional Integral and Stochastic Representations for Bosonic Ensembles — •MANFRED SALMHOFER — Institut für theoretische Physik, Universität Heidelberg, Philosophenweg 19, 69120 Heidelberg

I discuss rigorously defined coherent-state functional integral representations for the partition function and correlation functions of many-boson systems, both for the canonical and the grand-canonical ensemble, and the relation of these

essential differences between the canonical and grand-canonical ensemble, outline a simplified proof of equivalence of different functional integral representations in the time-continuum limit, and discuss their use in obtaining convergent

Invited TalkMP 3.3Tue 12:00MP-H5Color-Flavor Transformation Revisited•MARTIN ZIRNBAUERInstitutefor Theoretical Physics, Uni Köln, Germany•MARTIN ZIRNBAUER•MARTIN ZIRNBAUER

expansions for the correlation functions.

representations to ensembles of interacting random walks. I will highlight a few

The "color-flavor transformation", conceived as a kind of generalized Hubbard-Stratonovich transformation, is a variant of the Wegner-Efetov supersymmetry method for disordered electron systems. Tailored to quantum systems with disorder distributed according to the Haar measure of any compact Lie group of classical type (A, B, C, or D), it has been applied to Dyson's Circular Ensembles, random-link network models, quantum chaotic graphs, disordered Floquet dynamics, and more. We review the method and, in particular, explore its limits of validity and some implications for the theory of Anderson localizationdelocalization transitions.

MP 4: Quantum field theory, AdS/CFT, non-equilibrium and quantum dynamics

Time: Tuesday 16:15-18:35

MP 4.1 Tue 16:15 MP-H5

Wormholes from Berry phases in AdS_3/CFT_2 — SOUVIK BANERJEE, MORITZ DORBAND, JOHANNA ERDMENGER, RENÉ MEYER, and •ANNA-LENA WEIGEL — Institute for Theoretical Physics and Astrophysics and Würzburg-Dresden Cluster of Excellence ct.qmat, Julius-Maximilians-Universität Würzburg, Am Hubland, 97074 Würzburg, Germany

The AdS/CFT correspondence states that certain CFTs admit a description in terms of a gravitational theory in asymptotically AdS geometries of one dimension more. A central question in understanding the mechanism behind the duality is how the geometry in the bulk spacetime is encoded in the dual CFT state. Berry phases present a useful tool for understanding this. In their most general form, Berry phases are geometric phases acquired by states due to the presence of holonomies when parallel transported around a closed loop in parameter space. The AdS/CFT correspondence admits the description of bulk geometries with semi-classical spacetime wormholes in terms of two entangled CFTs. Wormholes are a topological feature of the bulk spacetime that presents as a holonomy and thus can be probed with Berry phases. The entanglement induced by the wormhole in the bulk geometry implies the dual CFTs no longer factorize. We show that non-factorization in the dual entangled CFTs is evident in Berry phases for such systems. We briefly discuss further applications of Berry phases for probing spacetime holonomies in geometries without wormholes and their CFT interpretation.

MP 4.2 Tue 16:35 MP-H5

Quantum Complexity as Hydrodynamics — •PABLO BASTEIRO^{1,2}, JO-HANNA ERDMENGER^{1,2}, PASCAL FRIES¹, FLORIAN GOTH^{1,2}, IOANNIS MATTHAIAKAKIS^{1,2}, and RENÉ MEYER^{1,2} — ¹Institut für Theoretische Physik und Astrophysik, Julius-Maximilians-Universität Würzburg, 97074 Würzburg, Germany — ²Würzburg-Dresden Cluster of Excellence ct.qmat Location: MP-H5

Location: MP-H5

In recent years, many concepts of quantum information theory have been introduced to the AdS/CFT correspondence and have lead to several meaningful insights. In particular, quantum computational complexity has been suggested as a candidate to describe the late-time behavior of black hole interiors. We contribute to its understanding by considering Nielsen's geometric approach to operator complexity for the SU(N) group. We develop a tractable large N limit which leads to regular geometries on the manifold of unitaries. To achieve this, we introduce a particular basis for the $\mathfrak{su}(N)$ algebra and define a maximally anisotropic metric with polynomial penalty factors. We implement the Euler-Arnold approach to identify incompressible inviscid hydrodynamics on the twotorus as a novel effective theory for the evaluation of operator complexity of large qudits. We discuss the resulting complexity geometry in view of essential properties of holographic complexity measures, such as ergodicity and conjugate points.

MP 4.3 Tue 16:55 MP-H5

The free energy of the two-dimensional dilute Bose gas — •ANDREAS DEUCHERT¹, SIMON MAYER², and ROBERT SEIRINGER² — ¹Institute of Mathematics, University of Zurich — ²Institute of Science and Technology Austria (IST Austria)

We prove bounds for the specific free energy of the two-dimensional Bose gas in the thermodynamic limit. We show that the free energy at density ϱ and inverse temperature β differs from the one of the non-interacting system by the correction term $4\pi \varrho^2 |\ln(a^2\varrho)|^{-1}(2-[1-\beta_c/\beta]^2_+)$. Here *a* is the scattering length of the interaction potential, $[x]_+ = \max(0, x)$ and β_c is the inverse Berezinskii– Kosterlitz–Thouless critical temperature for superfluidity. The result is valid in the dilute limit $a^2\varrho \ll 1$ and if $\beta \varrho \ge 1$.

MP 4.4 Tue 17:15 MP-H5

Analytical and numerical methods for nonlinear diffusive transport and shock acceleration — •Dominik Walter¹, Horst Fichtner¹, Frederic ЕFFENBERGER¹, and YURI LITVINENKO² — ¹Ruhr-Universität-Bochum; Institut theoretische Physik IV — ²University of Waikato, Department of Mathematics, New Zealand

We explore a nonlinear diffusive type of particle/cosmic ray transport. A special focus will be put on particles/cosmic rays, escaping from a shock or other localized acceleration sites and their acceleration process. Instead of solving coupled differential equations, as is the more common method of describing the interaction of diffusing particles with the backgrond medium, we analyse a single nonlinear advection-diffusion equation. In a first step we analyse the effect of the nonlinear model on particle transport, we apply an analytical expansion technique to cartesian and spherical symetrical geometries, to derive approximate solutions to the resulting equations and establish numerical models to compare and expand on this results. As a foundation for the numerical models we use the grid based Code VLUGR3, to provide numerical solutions, when there is no analytical way of solving distinct models. As a second step we construct a model for shock acceleration, to investigate the impact of nonlinear diffusion on shock acceleration, again using VLUGR3. We recreate a linear cartesian case of reference and use it as a groundwork to investigate the effects of nonlinearity.

MP 4.5 Tue 17:35 MP-H5

Electrons and their interactions: A deduction from Quantum Field Theory — •NADINE CETIN and NILS SCHOPOHL — Institut für Theoretische Physik, Universität Tübingen, Auf der Morgenstelle 14, 72076 Tübingen

We suggest a new unitary transformation of the fundamental electron-positron field theory Hamiltonian of Quantum Electrodynamics (QED), that brings with help of Wegner's flow equation method this Hamiltonian to a block-diagonal form, each block labeled by an eigenvalue of the operator $\hat{N} = \hat{N}^{(e)} + \hat{N}^{(p)}$, with $\hat{N}^{(e)}$ counting the number of electrons and $\hat{N}^{(p)}$ counting the number of positrons, while keeping all interactions with with long wavelength photons $\lambda_{ph} \gg \frac{h}{m_e c}$ intact. A representation of that block-diagonal Hamiltonian as a perturbation series with regard to the fine structure constant $\alpha_{FS} \approx \frac{1}{137}$ establishes to order α_{FS}^2 the well known gauge invariant and particle number conserving Hamiltonian for nonrelativistic electrons and positrons in its second quantized guise. Our derivation of the effective two-particle interactions between particles carrying mass, charge and spin while moving at nonrelativistic speed arises

directly from QED throughout treating electron fields and positron fields in an entire symmetrical fashion. Similarities and salient contrast with related early work of Cohen-Tannoudji et al., Takashi Itoh and I. Bialynicki-Birula are discussed in detail.

MP 4.6 Tue 17:55 MP-H5

Location: MP-H5

Recent Results on Quantum Spin Glasses — •CHOKRI MANAI¹, SIMONE WARZEL¹, HAJO LESCHKE², and RAINER RUDERER² – ¹Department of Mathematics, TU München, Garching bei München, Germany – ²Institut für Theoretische Physik, FAU Erlangen, Erlangen, Germayn

In the past few decades, the theory of spin glasses has become a major field of interest in condensed matter physics, mathematical physics and proabbility theory. While in the classical spin glass theory many problems remain unsolved, at least a rough understanding of the underlying physics has been established. The milestone so far has been the derivation of a formula for the free energy in the classical Sherrington-Kirkpatrick model by this year's Nobel price winner Giorgio Parisi based on his replica method and its rigorous proof by Guerra and Talagrand.

The situation is vastly different for quantum spin glasses, where quantum effects are for example incorporated via a transversal field. In this case no closed formula has been found for the Quantum SK-model; and most physical results are based on numerical simulations. In this talk, I will give an introduction to the topic of quantum spin glasses. I will discuss recent results on hierarchical Quantum spin glasses, where one can rigorously prove a formula for the free energy. Moreover, we will have a quick look on the Quantum SK model, where one can at least prove the existing of replica symmetry breaking.

MP 4.7 Tue 18:15 MP-H5 Local composite operators in the Sine-Gordon model — DANIELA CADAMURO and •MARKUS B. FRÖB — Institut für Theoretische Physik, Universität Leipzig The Sine-Gordon model is a widely studied two-dimensional quantum field theory, which depending on the value of the coupling β is finite (for $\beta^2 < 4\pi$), superrenormalizable $(4\pi \le \beta^2 < 8\pi)$ or just renormalizable $(\beta^2 = 8\pi)$. However, local composite operators have not been studied in the theory, apart from a few simple examples. We show that even in the finite range $\beta^2 < 4\pi$ composite operators need additional renormalization beyond the free-field normal-ordering at each order in perturbation theory, and prove convergence of the renormalized perturbative series.

MP 5: Quantum information

Time: Wednesday 11:00-12:30

Invited Talk

MP 5.1 Wed 11:00 MP-H5 Falling through masses in superposition: quantum reference frames for indefinite metrics — Anne-Catherine de la Hamette^{1,2}, Viktoria Kabel^{1,2}, Esteban Castro-Ruiz³, and •Caslav Brukner^{1,2} — ¹Institute for Quantum Optics and Quantum Information, Vienna, Austria $-\,^2$ Faculty of Physics, University of Vienna, Vienna, Austria — ³Institute for Theoretical Physics, ETH Zürich, Zürich, Switzerland

The current theories of quantum physics and general relativity on their own do not allow us to study situations in which the gravitational source is quantum. In my talk, I will propose a strategy to determine the dynamics of objects in the presence of mass configurations in superposition, and hence an indefinite spacetime metric, using quantum reference frame (QRF) transformations. Specifically, I will show that as long as the mass configurations in the different branches are related via isometries, one can use an extension of the current framework of QRFs to "quantum isometries" to change to a frame in which the mass configuration becomes definite. Assuming covariance of dynamical laws under quantum coordinate transformations, this allows to use known physics to determine the dynamics. I will apply this procedure to find the motion of a probe particle and the behavior of clocks near the mass configuration, and thus find the time dilation caused by a gravitating object in superposition.

MP 5.2 Wed 11:30 MP-H5

Exploiting Graph Symmetries for Quantum Dynamics — Armin J. RÖMER^{1,2}, EMANUEL MALVETTI^{1,2}, ROBERT ZEIER³, and •THOMAS SCHULTE-HERBRÜGGEN^{1,2} — ¹Technical University of Munich (TUM) — ²Munich Centre for Quantum Science and Technology (MCQST) and Munich Quantum Valley (MQV) — ³Forschungszentrum Jülich GmbH, Peter Grünberg Institute, Quantum Control (PGI-8)

Systems of coupled spins can easily be represented by coloured graphs, where the vertices relate to the local spins while the edges stand for pairwise couplings of different type (colour). Potential graph symmetries then naturally simplify quantum dynamics in terms of generators.

We present the background for an efficient algorithmic way to exploit the graph symmetry for arriving (automatically) at a symmetry-adapted basis. It avoids explicit calculation of the entire underlying graph automorphism groups (usually taking the form of wreath products of permutation groups). It connects the well-known Weisfeiler-Leman algorithm (occurring in the context of graph isomorphism problems) with cutting-edge versions of calculating central and orthogonal idempotents.

Worked examples illustrate principles and practice as well as the advantageous connections to graph theory in a widely applicable manner.

MP 5.3 Wed 11:50 MP-H5 Markovian Quantum Systems with Full and Fast Hamiltonian Control - •EMANUEL MALVETTI^{1,2}, FREDERIK VOM ENDE^{1,2}, THOMAS SCHULTE-HERBRÜGGEN^{1,2}, and GUNTHER DIRR³ – ¹Dept. Chem., TU-München (TUM) - 2 Munich Centre for Quantum Science and Technology (MCQST) and Munich Quantum Valley (MQV) - 3 Institute of Mathematics, Universität Würzburg

Markovian quantum systems with full and fast Hamiltonian control can be reduced to an equivalent control system on the eigenvalues of the density matrix describing the state. First we consider the case of a single qubit, presenting explicit solutions of the optimal control problem for a large family of Lindblad operators. For the cases where analytic solutions seem out of reach, we can still efficiently compute numerical solutions. Second we consider quantum systems of arbitrary finite dimension. While analytic solutions to optimal control problems do not exist in the general case, the reduced control system on the eigenvalues is still a powerful tool. As an example, we derive necessary and sufficient conditions for a Markovian quantum system to be coolable.

MP 5.4 Wed 12:10 MP-H5 On the Alberti-Uhlmann Condition for Unital Channels - SAGNIK Chakraborty¹, Dariusz Chruściński¹, Gniewomir Sarbicki¹, and •FREDERIK VOM $ENDE^{2,3} - {}^{1}$ Institute of Physics, Faculty of Physics, Astronomy and Informatics, Nicolaus Copernicus University, Grudziadzka 5/7, 87-100 Toruń, Poland — ²Department of Chemistry, Technische Universität München, 85747 Garching, Germany — ³Munich Centre for Quantum Science and Technology & Munich Quantum Valley, Schellingstr. 4, 80799 München, Germany

We address the problem of existence of quantum channels between two sets of density matrices. We refine the result of Alberti and Uhlmann and derive a necessary and sufficient condition for the existence of a unital channel between two pairs of qubit states which ultimately boils down to three simple inequalities.

MP 6: Mathematical physics

Time: Wednesday 16:15-16:45

MP 6.1 Wed 16:15 MP-H5 Invited Talk State sum models with defects — •CATHERINE MEUSBURGER — Department Mathematik, FAU, Cauerstr. 11, 91058 Erlangen

State sum models arise in topological quantum field theories, quantum topology, 3d quantum geometry, in the quantisation of Chern-Simons gauge theories Location: MP-H5

Wednesday

and are also closely related to topological models in condensed matter physics such as Kitaev's quantum double model and Levin-Wen models. The inclusion of defects into such models is of interest in several of these research fields.

In the talk I introduce a model that includes defects in Turaev-Viro-Barrett-Westbury state sum models. It involves defect surfaces, defect lines and defect points decorated with higher categorical data and yields topological invariants.

MP 7: Quantum field theory

Time: Wednesday 16:45-18:45

MP 7.1 Wed 16:45 MP-H5

Fermionic integrable models and graded Borchers triples - •HENNING BOSTELMANN¹ and DANIELA CADAMURO² - ¹University of York, Department of Mathematics, York YO10 5DD, United Kingdom — ²Universität Leipzig, Institut für Theoretische Physik, Brüderstraße 16, 04103 Leipzig, Germany

The operator-algebraic construction of 1+1-dimensional integrable quantum field theories has received substantial attention over the past decade. These models are characterized by their asymptotic particle spectrum and their two-particle S-matrix; so far, those particles have been bosonic. By contrast, we consider the case of asymptotic fermions. Abstractly, they arise from a grading of the underlying operator algebraic structures (Borchers triples). Many of the technical methods required can be carried over from the bosonic case, mutatis mutandis; most importantly, existing results on the technically hard part of the construction (i.e., establishing the modular nuclearity condition) do not require modification.

Thus we are lead to a new family of rigorously constructed quantum field theories which are physically distinct from the bosonic case (with a different net of local algebras and a different scattering theory). Their local operators fulfill modified form factor axioms, consistent with the physics literature.

MP 7.2 Wed 17:05 MP-H5

Interacting massless infra particles in 1+1 dimensions — •WOJCIECH DYBALSKI¹ and JENS MUND² — ¹AMU Poznań — ²Universidade Federal de Juiz de Fora - UFJF

The Buchholz' scattering theory of waves in two dimensional massless models suggests a natural definition of a scattering amplitude. We computed such a scattering amplitude for charged infraparticles that live in the GNS representation of the 2d massless scalar free field and obtained a non-trivial result. It turns out that these excitations exchange phases, depending on their charges, when they collide. Perspectives for obtaining a similar effect in higher dimensions will also be discussed.

MP 7.3 Wed 17:25 MP-H5

Quantum energy inequalities in integrable QFT models — •JAN MANDRYSCH - Institut für theoretische Physik, Leipzig

Many results in general relativity rely crucially on classical energy conditions imposed on the stress-energy tensor. Quantum matter, however, violates these conditions since its energy density can become arbitrarily negative at a point. Nonetheless quantum matter should have some reminiscent notion of stability, which can be captured by the so-called quantum (weak) energy inequalities (QEIs), lower bounds of the smeared quantum-stress-energy tensor. QEIs could be proven in many free quantum field theory (QFT) models on both flat and curved spacetimes. In interacting theories only few results exist. We are here presenting numerical and analytical results on QEIs in interacting integrable QFT models in 1+1 dimension, in particular the O(N)-nonlinear-sigma model at 1particle level.

MP 7.4 Wed 17:45 MP-H5 **Dyson-Schwinger Equations in Tensorial** Φ^4 **Theory** — •JOHANNES THÜRIGEN - Mathematisches Institut der Westfälischen Wilhelms-Universität, Münster,

Deutschland In quantum field theory, the Connes-Kreimer Hopf algebra captures not only the structure of perturbative renormalization but allows also to describe the nonperturbative regime in terms of "combinatorial" Dyson-Schwinger equations. This algebra generalizes from usual Feynman diagrams to "strand graphs", the combinatorial objects underlying a broad class of non-local theories, in particular tensorial field theory. Here we show how this can be used to derive Dyson-Schwinger equations in the case of Φ^4 theory with tensorial interactions.

MP 7.5 Wed 18:05 MP-H5

Dyson series approach for understanding quadratic interactions - •AYUSH PALIWAL¹ and KARL HENNING REHREN² - ¹Institut für Theoretische Physik, Universität Göttingen — ²Institut für Theoretische Physik, Universität Göttingen

We analyze for general quadratic interactions whether, and in which sense, the two point function of the perturbed field, written in terms of the unperturbed field as a Dyson expansion, taken in the vacuum state of the unperturbed field, converges to the vacuum state of the perturbed field when the adiabatic limit is taken. The answer is affirmative in a number of simple cases, where the perturbed field is explicitly known. The result is not in conflict with Haag's theorem. It suggests to use the same method to compute the two-point function in some cases of interest where the perturbed field is not known.

MP 7.6 Wed 18:25 MP-H5

A C*-algebraic approach to the classical limit of quantum systems -•CHRISTIAAN VAN DE VEN¹ and VALTER MORETTI² — ¹Max Planck Institute for Mathematics in the Sciences, Leipzig, Germany — ²Università di Trento, Trento, Italy

Quantization refers to the passage from a classical to a corresponding quantum theory. The converse problem, called the classical limit of quantum theories, is considered as a much more difficult issue and poses an important question for various areas of modern mathematical physics.

In this talk I will present this notion, first by introducing a natural framework based on the theory of deformation quantization. Subsequently, I will show that this setting is perfectly suitable to model the classical limit of quantum theories. More precisely, the corresponding quantization maps allow us to take the limit for $\hbar \, \rightarrow \, 0$ of a suitable sequence of algebraic states induced by $\hbar\text{-dependent}$ eigenvectors of several quantum models, in which the sequence converges to a probability measure on the pertinent phase space.

In addition, since this C*-algebraic approach allows for both quantum and classical theories, it provides a convenient way to study the theoretical concept of spontaneous symmetry breaking (SSB) as an emergent phenomenon when passing from the quantum realm to the classical world by switching off the semi-classical parameter \hbar . This is illustrated for several physical models, e.g. Schrödinger operators and mean-field quantum spin systems.

MP 8: Alternative approaches

Time: Wednesday 17:25-18:25

Treppendorf 38

MP 8.1 Wed 17:25 MP-H6 Speicherung elektrischer Energie — • ROLAND RENZ — 96138 Burgebrach

Unter Vernachlässigung der elektischen Verschiebungsstromdichte ist das Maxwellsche Gleichungssystem analytisch gelöst worden. Die neue Lösung ist zugeschnitten auf einen großen Bereich von elektromagnetischen Randbedingungen, inclusive der von DC- und AC-Betrieb. Man kann diese analytische Lösung auf das älteste Problem der Energietechik anwenden, der Speicherung von elektrischer Energie. Dazu nutzen wir die in elektrischen Maschinen rotatorisch erzeugte Spannung, die jetzt analytisch berechnet werden kann. Diese rotatorisch erzeugte Spannung beruht auf der räumlich verteilten elektrischen Feldstärke, die aus einem Vektorprodukt entsteht, gebildet aus der Umlaufgeschwindigkeit des koaxialen Kondensators und der ruhenden axialen magnetischen Induktion einer DC-Spule. Die so erzeugte räumliche E-Feldstärke ist radial gerichtet und überlagert das elektrische Feld des geladenen koaxialen Kondensators. Die resultierende E-Feldstärke bestimmt die Spannung am Kondensator. Mit Hilfe des analytischen Durchgriffs optimieren wir die Maschine bezüglich aller Parameter. Wir berechnen die elektrische Ladung des Kondensators durch eine zeitliche Integration des Ladestroms, der bestimmt ist durch die Spannungsdifferenz von aktueller Kondensatorspannung und aktueller DC-Busbar-Spannung. Regelungstechnisch gesehen hat der so gebaute elektrische Energiespeicher integrale Eigenschaften, die notwendig sind, um ein HGÜ-Netz sicher und sehr stabil machen zu können.

MP 8.2 Wed 17:45 MP-H6 Physik und Sehgewohnheiten — •HELMUT HILLE — Fritz-Haber-Straße 34, 74081 Heilbronn

Time: Thursday 11:00-12:40

Invited Talk

MP 9.1 Thu 11:00 MP-H5 Reduced phase space quantisation in Loop quantum gravity and loop quan-

tum cosmology — •Kristina Giesel¹, Bao-Fei Li², and Parampreet Singh² ¹Institute for Quantum Gravity, Department of Physics, FAU Erlangen-Nürnberg, Staudtstr. 7, 91058 Erlangen, Germany – ²Department of Physics and Astronomy, Louisiana State University, Baton Rouge, LA 70803, USA In this talk an overview over results that apply a reduced phase space quantisation to formulate the dynamics of loop quantum gravity and loop quantum cosmology will be presented. It will be briefly discussed how a reduced phase space for GR can be derived by coupling additional reference matter. The reduced phase space for GR is taking as a starting point for a loop quantisation

either in full LQG or in the symmetry reduced case of LQC. Different choices of reference matter yield in general different quantum models and several existing models will be compared. In the framework of LQC it will be analysed how different choices of reference matter can lead to different physical properties of the models and it will be discussed what kind of conditions appropriate reference matter should satisfy in this context.

Invited Talk

MP 9.2 Thu 11:30 MP-H5

Quantum fields propagating on curved backgrounds and their influence on spacetime curvature — •NICOLA PINAMONTI — Department of Mathematics, University of Genova, Italy

We shall review the theory of quantum fields propagating on curved backgrounds in the semiclassical approximation. Within this approximation matter is described by quantum fields propagating on a classical curved spacetime and their influence on spacetime curvature is taken into account by means of semiclassical Einstein equations. Typical known effects in this approximation are particle creation on cosmological spacetime, Hawking radiation in the case of black hole backgrounds and their evaporation. The semiclassical analysis we would like to present requires a careful study of the form of the correlation functions of the state which describes matter. The thorough analysis of their ultraviolet divergences and their renormalization is necessary to obtain meaningful expressions for the expectation values of the matter stress energy tensor. The resulting stress energy tensor will have nontrivial effects on the curvature on cosmological spacetimes too. In the latter case, semiclassical Einstein equations give origin to a well posed dynamical system provided the quantum state for matter is chosen in an appropriate way. The question of existence of exact solutions of such system will be discussed and some implication for cosmology will be presented. In the case of black hole physics, exploiting the stress tensor properties, we will give a local model of black hole evaporation.

Es ist eine Sehgewohnheit, getrennt Gesehenes als getrennt Existierendes einzustufen, Das kann für das Überleben nützlich sein, können wir doch dadurch Einzelnes besser fokussieren. Aber was ist z.B. mit Erde, Sonne und Mond? Reagieren sie nicht als ein Ganzes? Das Wasser der Meere türmt sich in Richtung Mond. Und wenn Mond und Sonne in einer Linie stehen gibt es die gefürchtete Springflut. Und so ist es doch mit allen kosmischen Objekte, die ewig umeinander herumtanzen und um ein geheimes Zentrum kreisen. Keines täte es von sich aus. Einstein sagte: *Es gibt keine Fernwirkungen* Richtig! Die Dinge sind sich nicht fern, wir sehen sie nur so! Alles wurde beim Urknall miteinander verschränkt und reagiert dadurch von Anfang an als ein Ganzes. Wir nennen die Kraft die dies bewirkt Schwerkraft oder Gravitation, Das Problem der Fernwirkungen erledigt sich so ohne jede Hypothese, allein durch Beachtung unserer Sehgewohnheit. Ebenso ist es mit dem Problem der Bewegung bei Unbelebten, was näher ausgeführt wird. So wie wir Biologie und Physik unterscheiden müssen, so müssen wir die *Bewegung* von Belebten und Unbelebten unterscheiden, wollen wir physikalische Objekte angemessen verstehen.

MP 8.3 Wed 18:05 MP-H6

Die Vereinigung der vier Grundkräfte der Natur - • ADOLF BABLITZKA - Alpenblick 6, 88682 Salem, Baden

Der Vortrag beschreibt die Vereinigung der vier Grundkräfte der Natur durch Einführung der von mir sog. Planckkraft

Kou = h * c / Oo aus Oo = G * h / (c * c * c),

sowie der Einführung der Spannung Uo mit Ersatz von ec durch eo.

Es werden Terme vorgeschlagen, die alle genannten Naturkonstanten mit beliebiger Genauigkeit ermitteln lassen.

MP 9: Quantum gravity (joint session MP/GR)

Location: MP-H5 MP 9.3 Thu 12:00 MP-H5

A master equation for gravitationally induced decoherence of a scalar field - •Max Joseph Fahn, Kristina Giesel, and Michael Kobler — Institute for Quantum Gravity, Department of Physics, FAU Erlangen-Nürnberg, Staudtstr. 7, 91058 Erlangen, Germany

In the talk, we present the derivation of a decoherence model containing a scalar field coupled to a gravitational environment. With such a model, one can predict quantum gravity effects in the scalar field's dynamics that arise due to the presence of gravity. Starting with full general relativity in Hamiltonian form expressed in Ashtekar's connection formulation, we focus on weak gravitational interactions in an asymptotically flat universe, modelled by gravitational waves propagating on a fixed background, and the scalar field as matter component. Firstly, we construct appropriate observables (gauge invariant quantities) for the model which become elementary variables in the reduced phase space of the model. Afterwards a reduced phase space quantisation using Fock quantisation is performed. With the help of the projection operator technique, we derive a second order time-convolutionless master equation, that is an effective evolution equation which predicts the temporal evolution of the scalar field, including the gravitational interaction effectively in terms of certain operators, whose form is a result of the model under consideration and several physical assumptions. These additional terms lead to physical effects like dissipation or decoherence of the matter field induced by gravity. Finally, we briefly discuss possible applications of the model's master equation.

MP 9.4 Thu 12:20 MP-H5 Quantum (supersymmetric) black holes in loop quantum gravity -•KONSTANTIN EDER — FAU Erlangen-Nürnberg

Black holes are immediate and unavoidable consequences of Einstein's theory of gravity whose existence nowadays has been confirmed with remarkable accuracy. Albeit leading to a huge success of the theory, there are also various open questions that point out its incompleteness. One of them is about the huge amount of microstates needed to explain the entropy of black holes as predicted by Bekenstein and Hawking: What are these micostates? Answering it poses a challenge for any formulation for quantum theory of gravity. In loop quantum gravity a very intriguing picture has been developed that suggests an answer in terms of topological Chern-Simon degrees of freedom induced by quantized geometry on the horizon. In this talk, we first give a general review on the description of black hole horizons in LQG. Then, we will give an outlook on recents results towards their generalization to the supersymmetric context. There, it turns out, using tools from super Cartan geometry, that the unique boundary theory in chiral supergravity corresponds to that of a super Chern-Simons theory. This enables one to transfer ideas from the bosonic to the supersymmetric setting.

63

MP 10: Classical and quantum gravity

Time: Thursday 16:15-17:15

MP 10.1 Thu 16:15 MP-H5

Space - Time - Matter: Finite Projective Geometry as a Quantum World with Elementary Particles — •KLAUS MECKE — Institut für Theoretische Physik, Universität Erlangen-Nürnberg

A unified theory for space-time and matter might be based on finite projective geometries instead of differentiable manifolds and fields. In contrast to general relativity the metric is given over a finite Galois field which defines neighbors in the finite set of points. Due to the projective equivalence of all quadratic forms in finite projective geometries this world exhibits necessarily a 4-dimensional, locally Lorentz-covariant space-time with a gauge symmetry G(3)xG(2)xG(1) for points at infinity which represent elementary particle degrees of freedom. Thus, matter appears as a geometric distortion of an inhomogeneous field of quadrics and physical properties of the standard model such as spins and charges seem to follow from its finite geometric structure in a continuum limit. The finiteness inevitably induces a fermionic quantization of all matter fields and a bosonic for gauge fields. The main difference to Einstein's general theory of relativity is the use of finite fields instead of real numbers to parametrize points of events.

K. Mecke, Biquadrics configure finite projective geometry into a quantum spacetime, EPL 120, 10007 (2017).

MP 10.2 Thu 16:35 MP-H5

On consistent gauge-fixing conditions in polymerized gravitational systems - KRISTINA GIESEL¹, BAO-FEI LI², PARAMPREET SINGH², and •STEFAN AN-- $^{1} {\rm Institute}$ for Quantum GravityDepartment of Physics, Theo-DREAS WEIGL retical Physics III, FAU Erlangen-Nürnberg, Staudtstr. 7, 91058 Erlangen, Germany — ²Department of Physics and Astronomy, Louisiana State University, Baton Rouge, LA 70803, USA

Gauge fixing is a standard method for deriving the physical sector of a gauge the-

MP 11: Thermodynamics and fundamental aspects of field theory

Time: Thursday 16:35-18:15

MP 11.1 Thu 16:35 MP-H6

Relation between the Cartesian multipole expansion and the spherical harmonic expansion — •NILS WALTER SCHWEEN and BRIAN REVILLE — Max-Planck-Institut für Kernphysik, Saupfercheckweg 1, Heidelberg 69117, Germany The multipole expansion, which is, for example, used to approximate an electrostatic potential (or a gravitational potential), has two equivalent forms. First, it is a Taylor expansion, i.e.

$$4\pi\epsilon_0\phi(\mathbf{r}) = \int \frac{\rho(\mathbf{r}')}{|\mathbf{r}-\mathbf{r}'|} dr'^3 = \sum_{l=0}^{\infty} \frac{1}{l!} \frac{r^{i_1}\cdots r^{i_l} \mathbf{Q}_{i_1\cdots i_l}}{r^{2l+1}}$$

Note the Einstein summation convention. Secondly, it is a spherical harmonic expansion, i.e.

$$4\pi\epsilon_0\phi(\mathbf{r}) = \sum_{l=0}^{\infty}\sum_{m=-l}^{l}\frac{4\pi}{2l+1}q_l^m\frac{r^lY_l^m(\theta,\varphi)}{r^{2l+1}}.$$

We show that the relation between these two expansions can be formalised as a series of basis transformations of spaces of homogeneous polynomials of increasing degree *l*. These basis transformations allow us to derive an algorithm to express the components of the multipole tensors, i.e. $\mathbf{Q}_{i_1 \cdots i_l}$, as linear combinations of the spherical multipole moments q_l^m for an arbitrary degree l. Since the spherical multipole moments are

$$q_l^m := \int Y_l^{m*}(\theta, \varphi) r'^l \rho(\mathbf{r}) \, \mathrm{d} r'^3 \,,$$

this opens the opportunity to compute $\mathbf{Q}_{i_1 \cdots i_l}$ solving the above integral instead of performing the derivatives and integrations needed to compute the multipole tensor components directly.

On the unification of mechanics and thermodynamics — •GRIT KALIES — HTW University of Applied Sciences, Dresden, Germany

Mechanics and thermodynamics (TD) are usually considered to be unified. This view is based on the work of Clausius and Boltzmann [1, 2], who attributed the behavior of complex systems to idealized moved material points. Nevertheless, mechanics and TD have been incompletely unified so far. They are taught separately, use different types of energy, namely the external and internal energy, and while the equations of mechanics, relativistic mechanics, and quantum mechanics are invariant under time reversal, those of TD are not.

Based on recent work [3-6], a unified formulation of mechanical and thermodynamic process equations is presented: The internal and external energies of an object are described in a unified manner. The equation of the ideal gas is derived in a simple way using process equations. Force, entropy, chemical potential, Gibbs free energy and other state variables are vividly interpreted at the fundamental level. The findings, consistent with experiment, enable a realistic and causal quantum physics and cosmology to be developed.

[1] R. Clausius, Pogg. Ann. 142 (1871) 433-461; [2] L. Boltzmann, Sitzungsber. kaiserl. Akad. Wiss. Wien 66 (1872) 275-370; [3] G. Kalies, Z. Phys. Chem. 234 (2020) 1567-1602; [4] G. Kalies, Z. Phys. Chem. 235 (2021) 849-874; [5] G. Kalies: Back to the roots: The concepts of force and energy, Z. Phys. Chem. (2021) 1-53; DOI: 10.1515/zpch-2021-3122;[6] G. Kalies: On the unification of mechanics and thermodynamics, submitted (2021).

MP 11.3 Thu 17:15 MP-H6 The fundamental role of the proper time parameter in general relativity and in quantum mechanics — • RENÉ FRIEDRICH — Strasbourg

Special relativity provides time with a precise physical concept, splitting the absolute Newtonian time into two different time concepts, proper time and coordinate time: In a first step, time is generated by rest energy, in the form of the proper time parameter of a worldline of a particle, and in a second step, the worldline is observed, by measuring the corresponding coordinate time parameter of the observer. Conversely, from the observed coordinate time we may retrieve by calculation the underlying proper time of the worldline. Proper time is time before time dilation, and coordinate time is time after time dilation. This is why for fundamental, unresolved questions about time we should not refer to the coordinate time parameter of spacetime but to the more fundamental parameter of proper time.

MP 11.4 Thu 17:35 MP-H6 Electrodynamics and the Isoclinic Decomposition of SO(4) — •ALEXANDER UNZICKER — Pestalozzi-Gymnasium München

According to William Rowan Hamilton, quaternions are a 'fundamental building block of the physical universe'. By visualizing unit quaternions with the stereographic projection, it is shown that unit quaternions display many characteristic features of electrodynamics. The question as to the origin of this interaction thus might be approached with an alternative description of spacetime by means of quaternions.

MP 11.5 Thu 17:55 MP-H6 On Negative Mass, Dark Energy and Dark Matter — •ANTOINE A.J. VAN DE VEN — Alumnus of Utrecht University, P.O. Box 19027, 3501 DA, Utrecht, The Netherlands

Back in 1957 Bondi published about negative mass in general relativity. In 2003

64

Thursday

ory. In the context of symmetry reduced models of loop quantum gravity a polymerisation has been applied to gauge fixed models to obtain so called effective theories that mimic the underlying quantum theory to some extend. Motivated from the question whether gauge fixing and polymerization commute, in this talk we will discuss the subtleties of implementing dynamical consistent gauge fixings in the effective theory and present a procedure to determine in a given model the effective lapse and shift. Although we can proof for a range of models that gauge fixing and polymerization does indeed commute and discuss consequences, for most models in the literature this is not the case. We further discuss how for a given choice of effective lapse and/or shift one can obtain a corresponding gauge fixing condition and show that in general this requires non-standard polymerisations or gauge fixing conditions with different classical limits. Based on these results we will then conclude with a discussion of some models from the literature.

MP 10.3 Thu 16:55 MP-H5

Local normal forms for Riemannian metrics with infinitesimal symmetries of their pre-geodesics. — •ANDREAS VOLLMER — Corso Duca degli Abruzzi, 24, 10129 Torino TO Italy

Projective vector fields are infinitesimal symmetries that preserve pre-geodesics, i.e. geodesics up to reparametrisation. A classical problem formulated by Sophus Lie is to describe the local metrics that admit such symmetries.

The list of local normal forms in dimension two has only recently been established, by Aminova (1990, 2003), Bryant-Manno-Matveev (2009), Matveev (2012) and Manno-V (2020). In joint work with G. Manno, we have recently derived such a list for dimension three, too (arXiv:2110.06785). The talk will provide an overview of the normal forms and their projective symmetry algebras, and will explain the techniques used to obtain them.

Theoretical and Mathematical Physics Division (MP)

the author published a new formalism and redefinition of mass and energy in which antimatter has negative mass while its energy is still always positive and wrote 'It is predicted that antimatter has antigravitational properties and that antimatter is the dark energy in the universe'. In later presentations it was also further elaborated that this can also explain away dark matter without further assumptions. Antimatter with repulsive gravitational fields surrounding a galaxy will generate an extra inward gravitational force without needing dark matter. This was also presented at 'The Dark Universe Conference' in Heidelberg in 2011 and other places and has been used and cited by others. See the references in [1].

MP 12: Quantum field theory for particle physics and plasmas

Time: Thursday 17:15-18:15

MP 12.1 Thu 17:15 MP-H5 Quantum Backreaction in Laser-driven Plasma — Gudrid Moortgat-Pick, ANANTA EFFENDIE, and •VICTOR ROGNER — Universität Hamburg

High-intensity laser-matter interactions are of interest for both experimental and theoretical physicists alike. New and more intense lasers could introduce inconsistencies into the classical approach on which the theories of the interactions have been calculated.

In light of new approaches to the quantisation of laser-driven Plasma that do not rely on particle-in-cell codes that are commonly used to model events on the quantum-scale, field theories are constructed through means including various attributes (like multi-particle effects) from the out-set.

In this Talk we will try to enable a better understanding of quantum-fieldtheory under the constraints given by a uniform plasma. Central to our work, we follow a different Ansatz namely the path integral quantisation of a Bi-Scalar field, which will be constructed through the successive introduction of quantum properties to the classical theory of laser driven plasma.

Starting from the classical equations governing the behaviour of laser-driven Plasma, we will extrapolate equations from the path integral quantisation and in doing so, find linearised field equations for our scalar fields that dictate the dynamics of a monochormatic Laser propagating through a uniform plasma.

We will then compare with results from literature.

MP 12.2 Thu 17:35 MP-H5 Quantum Backreaction in Laser-Driven Plasma — GUDRID MOORTGAT-PICK,

VICTOR ROGNER, and •ANANTA EFFENDIE — University of Hamburg Similarly to the statistical mechanics as they pertain to gases, the sets of mechanics for a plasma will need to be modeled separately from that of other states of matter. Location: MP-H5

In this presentation an update will be given including an alternative formulation and interpretation of the Dirac equation with quaternions in which negative energies are avoided and in which it is derived that antimatter has negative mass in this formalism.

Reference:

[1] S. Thakur and S. Mukerji, "A Brief Study of the Universe," International Journal of Astronomy and Astrophysics, Vol. 3 No. 2, 2013, pp. 137-152. doi:10.4236/ijaa.2013.32016

2. Quantum neid theory for particle physics and plasmas

In particular, the electrons behave as if free inside a plasma, the intense laserpulse interactions of which are of interest in the understanding of a laser-driven plasma. In contrast to the particle-in-cell approach already explored thoroughly and canonized in EPOCH, an approach considering multi-particle effects will yield separate mathematical insights.

The proposed method of integral quantisation utilizes the action in respect to the Lagrangian and allows quantum considerations from quantum-field-theory and quantum-electrodynamics to be introduced, the effects of which are largely ignored through the particle-in-cell approach.

This yields field equations describing a laser-driven plasma, including backreactions of quantum fluctuations, which can be linearised to describe a uniform, monochromatic laser beam propagating through a uniform plasma. The resulting quantum perturbations appear non-trivial.

MP 12.3 Thu 17:55 MP-H5

Higgs Mechanism — •ABHISHEK GOSWAMI — Adam Mickiewicz University, Poznan, Poland

In the Standard Model of particle physics, the interaction of a particle with the Higgs boson is responsible for its mass generation. This principle is known as the Higgs mechanism. In this talk, I will discuss a rigorous, non-perturbative proof of the Higgs mechanism. I will start with a weakly coupled U(1) Higgs theory on a unit lattice and show the exponential decay of correlations of the observable electromagnetic field strength tensor. This is the mass gap. I will also discuss the application of a new power series cluster expansion to this problem and explain how it provides a clean and simple alternative to decoupling cluster expansions in Constructive QFT.

Reference- Goswami, A. Mass Gap in Weakly Coupled Abelian Higgs on a Unit Lattice. Ann. Henri Poincaré 20, 3955-3996 (2019).

MP 13: Annual General Meeting

Location: MP-MV

Time: Thursday 19:30–20:30 Annual general meeting

Radiation and Medical Physics Division Fachverband Strahlen- und Medizinphysik (ST)

Anna C. Bakenecker Institute for Bioengineering of Catalonia (IBEC) Baldiri Reixac, 10-12 08028 Barcelona, Spain bakenecker@dpg-mail.de Ronja Hetzel RWTH Aachen University Templergraben 55 52056 Aachen ronja.hetzel@physik.rwth-aachen.de Reimund Bayerlein EXPLORER Molecular Imaging Center Department of Radiology University of California Davis rbayerlein@ucdavis.edu

Overview of Invited Talks and Sessions

(Lecture hall ST-H4; Poster P)

Plenary Talk of ST

See plenary section for details.

PV I	Mon	11:40-12:25	Audimax	Making uncertainty certain?: Proton therapy, range and in-vivo verification. — •Tot LOMAX	NY
------	-----	-------------	---------	--	----

Invited Talks

ST 4.1	Tue	14:00-14:40	ST-H4	Present status and future challenges of magnetic resonance-guided radiotherapy —
ST 10.1	Thu	16:15-16:55	ST-H4	•CHRISTIAN P. KARGER Artificial intelligence in PET image reconstruction and quantitative analysis —
				•Zhaoheng Xie

Invited Talks of the joint symposium SMuK Dissertation Prize 2022 (SYMD)

See SYMD for the full program of the symposium.

SYMD 1.1	Mon	14:00-14:25	Audimax	Timeless Quantum Mechanics and the Early Universe — • LEONARDO CHATAIGNER
SYMD 1.2	Mon	14:25-14:50	Audimax	First tritium β -decay spectrum recorded with Cyclotron Radiation Emission
				Spectroscopy (CRES) — • CHRISTINE CLAESSENS
SYMD 1.3	Mon	14:50-15:15	Audimax	Watching the top quark mass run - for the first time! — •MATTEO M. DEFRANCHIS,
				Katerina Lipka, Sven-Olaf Moch
SYMD 1.4	Mon	15:15-15:40	Audimax	Towards beam-quality-preserving plasma accelerators: On the precision tuning
				of the wakefield — •SARAH SCHRÖDER

Prize Talks of the joint Awards Symposium (SYAW)

See SYAW for the full program of the symposium.

SYAW 1.1	Wed	14:10-14:40	Audimax	Wie überprüft man die Ziele der Lehramtsausbildung Physik? — •HORST
				Schecker
SYAW 1.2	Wed	14:40-15:10	Audimax	Astronomy at Highest Angular Resolution - Adaptive Optics, Interferometry and
				Black Holes — •Frank Eisenhauer
SYAW 1.3	Wed	15:10-15:40	Audimax	Turbulence in one dimension — •Alexander M. Polyakov

Sessions

ST 1.1–1.6	Mon	14:00-15:30	ST-H4	Radiation therapy I
ST 2.1–2.4	Mon	16:15-17:30	Р	Poster Session
ST 3.1–3.5	Tue	11:00-12:15	ST-H4	Artificial Intelligence in Medicine
ST 4.1-4.5	Tue	14:00-15:40	ST-H4	Radiation monitoring and dosimetry I
ST 5	Tue	16:15-17:30	ST-MV	Annual General Meeting

ST 6.1–6.5	Wed	11:00-12:15	ST-H4	Detectors and Applications I
ST 7.1–7.5	Wed	16:15-17:30	ST-H4	Radiation monitoring and dosimetry II
ST 8.1–8.4	Thu	11:00-12:00	ST-H4	Radiation therapy II
ST 9.1–9.6	Thu	14:00-15:30	ST-H4	Detectors and Applications II
ST 10.1-10.1	Thu	16:15-16:55	ST-H4	Total-Body PET
ST 11	Thu	16:55-17:10	ST-H4	Prize Ceremony and Closing Session

Annual General Meeting of the Radiation and Medical Physics Division

Tuesday 16:15-17:30 ST-H4

Networking and Coffee Breaks

Everybody is kindly invited to join the meeting area during the coffee breaks. Continue the discussion with the speakers of the previous session, meet your colleagues and get to know new people. Please feel free to bring a cup of coffee and enjoy the networking!

Sessions

- Invited Talks, Contributed Talks, and Posters -

ST 1: Radiation therapy I

Time: Monday 14:00-15:30

ST 1.1 Mon 14:00 ST-H4

Usage of the track reconstruction framework Corryvreckan in proton therapy — •CHRISTOPHER KRAUSE, VALERIE HOHM, KEVIN KRÖNINGER, JENS WEIN-GARTEN, and FLORIAN MENTZEL — TU Dortmund, Dortmund, Germany

The Inner Tracker of the ATLAS experiment requires the optimal performance of its pixel sensors. To test their efficiency, a reliable track reconstruction and analysis for testbeam data is necessary to ensure the precise detection of particles. In the last years, track reconstruction was mostly done with the EUTelescope software, a generic and versatile framework.

In 2017, the new track reconstruction software Corryvreckan was published with the intention to reduce external dependencies without reducing the quality and versatility of track reconstruction in complex environments.

Efforts are made in TU Dortmund to use pixel sensors and track reconstruction software for proton computed tomography. The usage of Corryvreckan for low energy high density proton beams is investigated.

This talk presents the performance tests of Corryvreckan with simulated data. The simulated data is generated with Allpix^2 and serves to test the usability of Corryvreckan with beam properties used in proton therapy. Results are further improved through the use of machine learning algorithms to separate true and false reconstructed tracks.

ST 1.2 Mon 14:15 ST-H4

Simulationen zur Ortsauflösung in der Protonenradiographie unter Verwendung eines ATLAS Pixeldetektors — •JACQUELINE SCHLINGMANN, MARIUS HÖTTING, KEVIN KRÖNINGER und JENS WEINGARTEN — TU Dortmund, Lehrstuhl für Experimentelle Physik IV

Bei der Protonentherapie muss eine Sicherheitsspanne um den zu bestrahlenden Bereich eingehalten werden um sicherzustellen, dass das gesamte zu bestrahlende Gewebe behandelt wird. Unsicherheiten durch Lagerungsveränderungen und durch die Bildgebung mit Elektronen sind ausschlaggebend für die Sicherheitsspanne.

Um die Protonentherapie zu optimieren, soll zur Verifizierung der Position des zu bestrahlenden Bereiches eine Bildgebung mit Protonen am Ort der Bestrahlung stattfinden. Desweiteren kann der Patient vor und während der Bestrahlung überwacht werden. Für die Bildgebung wird zunächst die Auflösung über ein edge Phantom mit Hilfe der Modulation transfer function (MTF) untersucht. Außerdem werden Methoden zur Bildverarbeitung und Objektextraktion getestet.

In diesem Vortrag werden die Ergebnisse und die Methode zur Auflösungsbestimmung präsentiert. Desweiteren werden Bildverarbeitungsmethoden mit den verwendeten Filtern in Hinblick auf die Möglichkeit zur Positionsfindung von Objekten im Phantom diskutiert.

ST 1.3 Mon 14:30 ST-H4

Study of prompt gamma emission in c^{12} (p, p*y) c^{12} nuclear reactions close to a bragg peak — •MARIAM ABULADZE^{1,2}, RONJA HETZEL³, JONAS KASPER³, RE-VAZ SHANIDZE^{1,2}, and ACHIM STAHL³ — ¹Tbilisi State Univerity, Tbilisi, Georgia — ²Kutaisi International University, Kutaisi, Georgia — ³RWTH Aachen University, Aachen, Germany

Proton therapy is a high-quality radiation therapy which uses a proton beam to irradiate human tissue. The advantage of this type of treatment is a highly conformal dose deposition due to the presence of the Bragg peak. Though it is often required to irradiate the tumor volume with a precision better than 1 - 2 mm, which means that proton therapy needs not only precise treatment planning but also monitoring and proton range verification during the treatment. One of the ways to monitor the proton range is Prompt Gamma Imaging (PGI) which means to detect gamma rays produced by the excitation of the target nuclei by incident protons.

In this work, a results of the Geant4 simulation (version 10.6.3.) of interactions of 17.56 MeV protons with a carbon target are shown. This includes the study of 4.4 MeV and 9.6 MeV line properties, as multiple differences were observed between simulation and experiment, one of which is a double peak for the $C^{12}(p, p^*y)C^{12}$ spectral line. The shortcomings of the current physical models in Geant4 in describing the shape and intensity of the 4.4 MeV and 9.6 MeV gamma lines will be discussed.

ST 1.4 Mon 14:45 ST-H4 Granularity and Photomultiplier studies for Prompt Gamma Spectra in Proton Therapy — •OLGA NOVGORODOVA and ARNO STRAESSNER — IKTP TU Dresden, Dresden, Germany Prompt gammas (PG) in proton therapy are one of the promising techniques for non-invasive measurements of in-vivo proton range and it is already in implementation in clinical research. It can be based on time or spectral measurements. We concentrate on the spectral properties by developing systems measuring PG in the range 2-8 MeV, which already were shown to provide a time resolution of about 100 ns. A big problem in the recording prompt gammas during the irradiation of patients is the data load due to the size of the crystal. By decreasing the size of the crystals and forming them into the matrices the load to each channel can be reduced and more PGs can be detected. We are investigating different sizes and types of crystals to find an optimum. Availability on the market and the choice of photomultipliers also plays an important role for the granularity of crystals. Three different types of photomultipliers were used: a regular PMT, MAPMT and several SiPMs, which have high photon detection efficiency, good time resolution, low bias voltage and can operate in magnetic fields. For the SiPMs, PETsys readout electronics was used. In the presentation we present effects due to crystal types and sizes, and compare different photomultipliers and read out systems.

ST 1.5 Mon 15:00 ST-H4

Use of PET Readout Electronics for a Scintillating Fiber-Based Compton Camera — •SARA MÜLLER^{1,3}, RONJA HETZEL^{1,3}, MAREIKE PROFE¹, ACHIM STAHL¹, ALEKSANDRA WRONSKA², MING LIANG WONG², MAGDALENA KOŁODZIEJ², KATARZYNA RUSIECKA², DAVID SCHUG³, BJÖRN WEISSLER³, and VOLKMAR SCHULZ^{1,3} — ¹RWTH Aachen University - Physics Institute III B, Aachen, Germany — ²M. Smoluchowski Institute of Physics, Jagiellonian University Kraków, Poland — ³RWTH Aachen University, Physics of Molecular Imaging Systems, Aachen, Germany

A powerful tool in cancer treatment is hadron therapy. Its precision can be optimised by online-monitoring the Bragg peak position by using prompt gamma radiation emitted in the process. A Compton camera is a promising setup for this task as it provides the possibility to reconstruct the shape and the location of the deposited dose.

The SiFi-CC (SiPM and scintillating Fiber-based Compton Camera) is developed by researchers of the RWTH Aachen University, the Jagiellonian University in Kraków and the University of Lübeck. The two parts of the SiFi-CC, the scatterer and the absorber, both consist of closely packed LYSO:Ce fibres and are read out by digital SiPMs. The SiPMs are arranged in a sensor tile which was originally developed for PET systems and which consists of 36 Digital Photon Counters containing four readout channels each. Thus, the light produced in the scintillating fibers is spread over several SiPMs. A small-scale prototype of the camera has been assembled and was tested under laboratory conditions. First results of these measurements will be presented.

ST 1.6 Mon 15:15 ST-H4

Proton field verification by implant activation — •CLAUS MAXIMILIAN BÄCKER^{1,2,3}, CHRISTIAN BÄUMER^{1,2,3}, WALTER JENTZEN⁵, SANDRA LAURA KAZEK⁵, KEVIN KRÖNINGER¹, FLEUR SPIECKER¹, NICO VERBEEK^{2,3}, JENS WEINGARTEN¹, JÖRG WULFF^{2,3}, and BEATE TIMMERMANN^{2,3,4,6} — ¹TU Dortmund University, Department of Physics, D-44221 Dortmund — ²West German Proton Therapy Centre Essen, D-45122 Essen — ³University Hospital Essen, West German Cancer Center, D-45122 Essen — ⁴University Hospital Essen, Clinic for Nuclear Medicine, D-45122 Essen — ⁶University Hospital Essen, Faculty of Medicine, D-45147 Essen

The delivered dose in particle therapy is sensitive to the correctly predicted range. Uncertainties arise e.g. from estimation of tissue stopping powers. In order to identify range uncertainties of the treatment fields, multiple techniques have been developed in the past. However, these techniques come with several limitations for the clinical applicability. In this study, the potential use of implant activation for field verification is investigated. Therefore, a method validation is performed focusing on the activation of titanium implants during proton therapy treatments and subsequent PET imaging. The parameters delivered dose and variations in the beam range are investigated to estimate the detectability of these changes during fractional treatment. While first pre-clinical studies demonstrated the general potential of implant activation for range verification, the current limitations and necessary developments in PET imaging will be presented in this talk.

Location: ST-H4

ST 2: Poster Session

Time: Monday 16:15-17:30

ST 2.1 Mon 16:15 P

Characterization of a multi-arm Compton-camera setup - Henning GEESMANN¹, GIULLIO LOVATTI¹, TIM FITZPATRICK^{1,2}, JENNIFER ZHOU¹, MOhammad Safari¹, Florian Schneider², Vassia Anagnostatou¹, Katia PARODI¹, and PETER G. THIROLF¹ - ¹Ludwig-Maximilians-Universität, Munich, Germany — ²KETEK GmbH, Munich, Germany

PET, *y*-PET and prompt-*y* imaging are all *y*-ray based imaging techniques that can be used for beam range verification in particle therapy. They can all be realised by a multi-arm Compton camera (CC) setup.

We characterized a system where a single CC arm of the setup comprises a 16×16 pixelated GAGG crystal array as scatterer (26×26×6 mm³, read out by one 25µm microcell SiPM array) and a monolithic absorber (LaBr₃:Ce or CeBr₃ crystal, $51 \times 51 \times 30 \text{ mm}^3$), read out by four $50 \mu \text{m}$ microcell SiPM arrays). The characterization of the setup in terms of spatial resolution and efficiency is based on the use of different y-ray point sources (511, 662 and 1274 keV) and several spatial arrangements of up to four CC arms. Further, the angular dependence of the Compton scattering kinematics can be exploited to identify spatial arrangements (scatterer/absorber) in a single CC arm that allow to select a certain regime of energies being deposited in the scatterer and, thus, allow for enhancing the performance of the setup. A summary of the latest status of the ongoing studies will be given.

This work was supported by the Bayerische Forschungsstiftung, by the ERC grant No. 725538 (SIRMIO) and by Sklodowska-Curie Individual Fellowship project HIPPOCRATE.

ST 2.2 Mon 16:15 P

Stabilization of magnetic fields in a low-field magnetic resonance tomograph for student learning - •Velda Azalia Abraham, Thomas Heinzel, and MIHAI CERCHEZ — Condensed Matter Physics Laboratory (IPKM), Heinrich-Heine University Dusseldorf

EFNMR (Earth's Field Nuclear Magnetic Resonance) is a low-field NMR (Nuclear Magnetic Resonance) technique. Employing the Earth's magnetic field, this is an excellent tool for Medical Physics students learning as it is easy to operate and low-cost. However, EFNMR also has some downsides related to signal-tonoise ratio and magnetic field stability. Either natural fluctuations in the Earth's magnetic field or room-related changes may influence the long-time 3D imaging through changes in the Larmor frequency. Therefore a stabilization system, that would keep the magnetic field stable at all times is needed. We present here a two-directional Helmholtz correction coil system that compensates for changes

Location: P

of the homogeneous magnetic field in the room, leading to increased imaging quality.

ST 2.3 Mon 16:15 P

Image registration of CT and MRI scans using deep learning - •ALEXANDER RATKE, HANNAH TIMM, JOHANNES WINTZ, GÖKSU ÜNLÜ, and BERNHARD SPAAN — TU Dortmund University, Dortmund, Germany

In radiation therapy, precise localisation of tumour and risk structures is important for therapy planning. Medical imaging methods, such as computed tomography (CT) and magnetic resonance imaging (MRI), offer different advantages due to the respective physical process, which can be combined by image fusion.

In this project, CT scans and T_1 - and T_2 -weighted MRI scans of different body regions are used. For their registration, two processing steps are performed. The preparatory part includes equalising the formats of the images, which is required for a neural-network-based registration. Deep learning is then used to filter structures of an image and to match them to a second image. The results of the registration of three-dimensional CT and MRI scans of the skull and the thorax will be presented as well as studies of quality and uncertainty, performed with a Shepp-Logan phantom.

ST 2.4 Mon 16:15 P

Design and validation of a Digital Twin for tumor stage prediction of prostate cancer patients — •Anna-Katharina Nitschke, Carlos Andres Brandl, and MATTHIAS WEIDEMÜLLER - Physikalisches Institut, Ruprecht-Karls Universität Heidelberg, Im Neuenheimer Feld 226, 69120 Heidelberg, Germany

Digital Twins (DT) are virtual representations of physical assets. In healthcare, DT can help to personalize diagnosis and treatment for complex diseases like prostate cancer.

To be able to generate a DT which is supporting the complete clinical journey of a cancer patient, a conceptual development for combining several individual algorithms is needed. This approach also ensures best possible practices for patient data protection during collaborative work. As interpretability and uncertainty quantification are crucial in medical applications, on hast to find a solution that meets these requirements as well as clinical guidelines. Through using ensemble methods, we develop a concept of a DT, which is combining machine learning algorithms and expert systems with the medical doctor's opinion to improve decision-making. We exemplify the approach with tumor stage prediction (TNM) before treatment planning.

This work is part of the project CLINIC 5.1 and is supported by the BMWi.

ST 3: Artificial Intelligence in Medicine

Time: Tuesday 11:00-12:15

ST 3.1 Tue 11:00 ST-H4

Optimierung und Evaluierung der Registrierung von CT- und MRT-Bildern mithilfe maschinellem Lernens — •ELENA DARSHT, ALEXANDER RATKE und BERNHARD SPAAN — Experimentelle Physik 5, TU Dortmund

In der Strahlentherapie ist der Erfolg einer Bestrahlung abhängig von der Genauigkeit der Tumorlokalisierung. Diese erfolgt üblicherweise anhand der zur Bestrahlungsplanung verwendeten CT-Bildern. Im Vergleich dazu bieten MRT-Bilder eine bessere Darstellung von gesunden und kranken Weichteilstrukturen für eine präzisere Konturierung an. Durch die Bildregistrierung und -fusion von CT- und MRT-Bildern kann die Bestrahlungsplanung anhand von CT-Bildern erfolgen und zusätzlich können die Informationen der MRT-Bilder genutzt werden.

Eine Möglichkeit die Registrierung umzusetzen, ist die Verwendung eines neuronalen Faltungsnetzes, wodurch eine Rechenzeit von nur einigen Minuten pro Registrierung erreicht werden kann. Es wird der Aufbau und die Evaluierung des verwendeten neuronalen Netzes präsentiert. Dabei werden insbesondere die zur Optimierung für dreidimensionale Schädelaufnahmen genutzten Parameter vorgestellt. Zur Evaluierung der Registrierung wird der Dice-Koeffizient verwendet, bei dem die Überlappung von segmentierten Bildern ermittelt wird.

ST 3.2 Tue 11:15 ST-H4

Towards a Digital Twin for clinical decision support of a prostate cancer patient — •Carlos Andres Brandl, Anna-Katharina Nitschke, and MATTHIAS WEIDEMÜLLER — Physikalisches Institut, Ruprecht-Karls Universität Heidelberg, Im Neuenheimer Feld 226, 69120 Heidelberg, Germany

We present concepts to realise a Digital Twin for decision support in medicine using the example of prostate cancer. The aim for a medical Digital Twin is to support clinical decision making by providing an intuitively interpretable way for the doctors' decisions. Inspired by concepts from engineering, we devise a combination of data-driven models with evidence-based knowledge which comprise a wide range of parameters and data types available. As a higher integrated approach, we discuss the implementation of latent representations of the patient. Compatibility with the clinical guidelines and physicians' decision-making processes must be ensured by finding appropriate and interpretable representations. This work is part of the project CLINIC 5.1 and is supported by the BMWi.

ST 3.3 Tue 11:30 ST-H4

Location: ST-H4

Using generative adversarial networks to predict proton beam dose distributions in mice — •Lara Bussmann, Kevin Kröninger, Armin Lühr, Florian MENTZEL, JANINE SALEWSKI, and JENS WEINGARTEN — TU Dortmund

The clinically used generic relative biological effectiveness (RBE) of 1.1 for protons compared to photons does not consider variations along the beams axis. For a better estimation of the varying RBE and to assess potential adverse effects, mouse brains are irradiated and excised to visualize DNA double-strand breaks.

In order to deduct conclusions about the RBE, the observed irradiation damage in the tissue is compared to the expected damage from Monte Carlo simulations of the dose distribution.

Using Monte Carlo simulations for dose distribution predictions can be very time-consuming. Machine learning models can be trained to predict dose distributions based on the phantom geometry.

In this talk, a deep learning dose prediction model for proton mouse irradiations based on generative adversial networks (GANs) is presented. GANs can be trained to generate data samples following a learnt distribution, which are indistinguishable from a ground truth distribution. In this study, MC simulation samples are used to train the GAN, using geometrical information about the target phantom as conditional input.

ST 3.5 Tue 12:00 ST-H4

ST 3.4 Tue 11:45 ST-H4

A step towards treatment planning for microbeam radiation therapy: fast dose predictions with generative adversarial networks - •FLORIAN Mentzel¹, Micah Barnes², Kevin Kröninger¹, Michael Lerch², Olaf Nackenhorst¹, Jason Paino², Anatoly Rosenfeld², Ayu Saraswati³, Ah CHUNG TSOI³, JENS WEINGARTEN¹, MARKUS HAGENBUCHNER³, and SUSANNA GUATELLI² - ¹TU Dortmund University, Department of Physics - ²Centre for Medical Radiation Physics, University of Wollongong, Australia – ³School of Computing and Information Technology, University of Wollongong, Australia Microbeam radiation therapy is a novel and currently pre-clinical radiotherapy treatment based on planar arrays of high intensity sub-millimetre synchrotron gamma rays. Due to good healthy tissue sparing it is a promising candidate e. g. for the treatment of glioblastoma. The dose computations required to plan treatments are currently performed using time-consuming Monte Carlo (MC) simulations with Geant4. The dose computations are complex as steep dose gradients near the beams require very high spatial resolution while the need to take into account the effect of stray radiation over large distances makes small voxel sizes infeasible.

A two-fold approach to these problems is explored: first, a novel data taking method for MC simulations is presented. The method considers both high resolution and long-range effects of stray radiation. Second, a deep learning model based on 3D-UNet GANs is created to mimic dose simulations of Geant4, allowing for very short prediction times.

A neural network for the event identification of a Compton camera -•Mareike Profe¹, Jonas Kasper¹, Awal Awal², Aleksandra Wrońska³, and ACHIM STAHL¹ - ¹RWTH Aachen University - Physics Institute III B, Aachen, Germany — ²Forschungszentrum Jülich - Institute for Nuclear Physics, IKP-4, Jülich, Germany — ³Jagiellonian University - M. Smoluchowski Institute of Physics, Cracow, Poland

In proton therapy, a main approach to diminish the problem of range uncertainties is to verify the range of the proton beam on-line. By imaging prompt gamma photons originating from interactions of the proton beam within the body this challenge can be mastered. For the on-line range verification, the SiFi-CC project aims to develop a Compton camera based on scintillating fibers read out with silicon photomultipliers. The Compton events need to be discriminated from other event types: random coincidences and physical background. Subsequently, their features crucial for imaging need to be determined, i.e. the position of the Compton effect, the energy transferred to the electron, the position of the first interaction of the Compton-scattered gamma and its energy. Neural networks are an promising alternative to a classical selection algorithm to tackle this task. Here, a convolutional neural network is trained to predict the type and parameters of the events. The design of the neural network as well as the evaluation of the performance are presented.

ST 4: Radiation monitoring and dosimetry I

Time: Tuesday 14:00-15:40

Invited Talk

ST 4.1 Tue 14:00 ST-H4 Present status and future challenges of magnetic resonance-guided radiotherapy — • CHRISTIAN P. KARGER — German Cancer Research Center (DKFZ), Heidelberg, Germany

As a special realization of image-guided radiotherapy (IgRT), magnetic resonance (MR)-guided radiotherapy (MRgRT) has been implemented clinically by integrating linear accelerators (Linacs) into MR imaging devices. These so-called MR-Linacs allow for fast imaging with superior soft-tissue contrast without extra dose prior or even during radiation treatment. The acquired images may be used to adapt the treatment plan to the actual anatomical situation and to compensate for motion by beam gating or tracking. Besides technical challenges of integrating and shielding of the two devices, dose distributions and the detector responses are changed by the magnetic field requiring modified dose calculation techniques and dosimetry protocols. Clinically, the development and validation of new workflows is challenging as adaptive treatment planning while the patient is lying on the table (online adaptive radiotherapy) requires fast and robust planning techniques as well as new methods for validating the adapted treatment plan. In addition special end-to-end tests have to be developed to quantify the overall dosimetric and geometric accuracy. The presentation gives an overview on the present status and future developments in MR-guided radiotherapy.

ST 4.2 Tue 14:40 ST-H4

The performance of scintillating fibre based beam profile monitor for ion therapy in magnetic field — •QIAN YANG, LIQING QIN, and BLAKE LEVERING-TON — Physikalisches Institut Universitaet Heidelberg

In the Heidelberg Ion-beam Therapy Center (HIT), proton, helium, carbon and oxygen ions are used for cancer therapy. The existing scanning technique, it is called Raster scanning. The beam is not switched off between spots but adjusted by the 2 dimensions fast deflection magnet. A tracking system is used to monitor the beam and feedback to adjust the scanning magnet currents online. The commercial MWPC (multiwire proportional chamber) currently used has several drawbacks. An MR magnetic is currently installed at the HIT facility for studying treatment with prompt MR imaging. The BMBF funded ARTEMIS project is focused on the establishment of a unique MR-guided ion-beam therapy prototype for clinical application. A scintillating fibre based detector is now studied as a possible monitor replacement for this system. The detector performance in the environment of MR magnetic fields is also be studied, to complement the ARTEMIS project at HIT. Recently, the performance of the detector was tested under the condition inside the Helmoltz coil by changing the magnetic field (from 0.1 mT ~ 99.9 mT) and it will also be tested inside the prompt MR imaging in the following days.

ST 4.3 Tue 14:55 ST-H4

First study on energy resolution in proton radiography — •MARIUS HÖTTING, KEVIN KRÖNINGER, ISABELLE SCHILLING, JACQUELINE SCHLINGMANN, HEN-DRIK SPEISER, and JENS WEINGARTEN — TU Dortmund, department of physics Proton radiography offers the possibility of real-time patient positioning before or during the treatment. Typically, pixelated silicon sensors are used in the implementation, which makes it possible to contour various structures of phantoms or patients from the pure evaluation of the proton hit position.

With the additional information of the deposited charge of each particle per pixel, the initial proton energy, the residual range and finally the water equivalent thickness (WET) can be calculated. This allows conclusions about the material traversed without an additional external energy detector and gives the opportunity for cross-check with the treatment plan.

The challenge is to determine the proton energy with the limited resolution in deposited charge and a low number of detected protons per pixel. The measurement of the WET of a phantom is therefore done in two steps, one without and one with an additional homogeneous absorber of known width. The energy difference in both images allows minimization of the systematic uncertainties. Finally, the conversion to the WET is done with a simulated calibration curve.

In this talk, we describe the currently investigated method for WET calculation with pixel sensors and first simulations with Allpix² to determine depthdependent uncertainties.

ST 4.4 Tue 15:10 ST-H4

Location: ST-H4

New Optimisation Method for Proton Radiography Images - MARIUS HÖT-TING, KEVIN KRÖNINGER, FLORIAN MENTZEL, •HENDRIK SPEISER, and JENS WEINGARTEN — TU Dortmund, department of physics

For years, proton therapy for cancer treatment has been experiencing an increasing application, as it has known advantages such as the high dose conformity of protons. However, using this precision requires enhanced imaging techniques to ensure the accurate patient alignment. This results in a reduction of the safety margin around the target volume and in the protection of the surrounding healthy tissue.

As part of a master thesis, a new method is being developed to improve the quality of proton radiography images and thereby achieve greater therapeutic success as mentioned above. At present, the image quality and the resolution of structures in the patient are degraded by proton scattering. The goal of the project is to combine the projections of the widened proton trajectories onto the image plane with conventional proton hit maps to increase the resolution of object edges in radiography images.

This talk will include a brief description of the developed method and the evaluation of the image quality using Monte Carlo simulations. Subsequently, the application of the new approach will be compared with results of previous imaging methods.

ST 4.5 Tue 15:25 ST-H4

Monte Carlo feasibility study for neutron radiography in proton therapy -•HANNAH ROTGERI, MARIUS HÖTTING, KEVIN KRÖNINGER, ALINA LANDMANN, and JENS WEINGARTEN — TU Dortmund University, Department of Physics

Due to their depth dose dose distribution, proton beams used in cancer treatment can reduce damage to healthy tissue when compared to irradiation with photons. Neutrons, and other secondary particles, are produced during proton therapy, causing dose deposition outside the treatment field. Being unavoidable, the feasibility for using these neutrons for imaging during the treatment is studied in the context of a master thesis.

Radiation and Medical Physics Division (ST)

The study is performed in the Monte Carlo framework Geant4. It is investigated whether different energy ranges of neutrons are suitable for radiography, due to the different ways they can be detected. This talk will present the recent results of this study. This includes the investigation of different angles of the detector with respect to the proton beam. Finally it will be shown whether the different materials found in a human body can be distinguished.

ST 5: Annual General Meeting

Time: Tuesday 16:15-17:30

Annual General Meeting of the Radiation and Medical Physics Division

ST 6: Detectors and Applications I

Time: Wednesday 11:00-12:15

ST 6.1 Wed 11:00 ST-H4

Upgraded Proton Irradiation Site at Bonn University — •PASCAL WOLF¹, DENNIS SAUERLAND², JOCHEN DINGFELDER¹, DAVID-LEON POHL¹, and NORBERT WERMES¹ — ¹Physikalisches Institut, Universität Bonn — ²Helmholtz Institut für Strahlen- und Kernphysik, Universität Bonn

The Bonn Isochronous cyclotron delivers 14 MeV (≈ 12.5 MeV on-device) protons with typical beam currents of 1 μ A and beam diameters of a few millimeters to the irradiation site. Enhanced beam diagnostics as well as R/O electronics allow for online monitoring across several orders of magnitude of beam currents with a relative uncertainty of \approx 1%. Devices are irradiated by being scanned through the beam in a row-wise pattern while housed in a thermally-insulated cooling box kept at \approx -20 °C to minimize annealing. Online monitoring of the beam current at extraction facilitates a measurement of the fluence per scanned row with an accuracy of a few %, ensuring homogeneous irradiation. The setup allows one to power and read out DUTs during irradiation as well as pause irradiations for in-between measurements. Latest irradiations of thin CMOS pixel test structures yield a proton hardness factor with reduced uncertainty, compatible with previous measurements and simulation, facilitating irradiations up to 10¹⁶ neq / cm² within a few hours. The setup, its reworked components, the irradiation procedure as well as the latest proton hardness factor measurements are presented in this talk.

ST 6.2 Wed 11:15 ST-H4

Comparison of different methods measuring the beam energy in proton therapy using pixelated silicon detectors — •ISABELLE SCHILLING¹, CLAUS MAXIMILIAN BÄCKER^{1,2,3}, CHRISTIAN BÄUMER^{1,2,3}, CARINA BEHRENDS^{1,2,3}, MARIUS HÖTTING¹, KEVIN KRÖNINGER¹, BEATE TIMMERMANN^{2,3,4,5}, and JENS WEINGARTEN¹ — ¹TU Dortmund University, Department of Physics, 44221 Dortmund — ²West German Proton Therapy Centre Essen, 45122 Essen — ³West German Cancer Center, University Hospital Essen, 45122 Essen — ⁴University Hospital Essen, Clinic for Particle Therapy, 45122 Essen — ⁵Faculty of Medicine, University of Duisburg-Essen, 45147 Essen

The accurate measurement of beam range for quality assurance (QA) in proton therapy is important for optimal patient treatment. Conventionally used detectors mostly calculate the energy by detecting the depth dose distribution of the protons. In contrast to this, the ATLAS pixelated silicon detector measures the deposited energy in the sensor for single protons, allowing the determination of the stopping power. The restriction on the dynamic energy range of the measurement is given by the readout chip. Hence, there are different ways to use the detector whose applicability is being examined. For range consistency checks during the QA, an absorber with different thicknesses is used to investigate the variation of the charge production in the sensor. In comparison, this talk also presents results of energy calculations by measuring the stopping power in the silicon sensor directly, all performed at the West German Proton Therapy Centre Essen.

ST 6.3 Wed 11:30 ST-H4

Real-time analysis for a scintillating fiber-based beam profile monitor for charged particle beams — •LIQING QIN, QIAN YANG, and BLAKE LEVERING-TON — Physikalisches Institut, Heidelberg, Germany

A lighter, faster and more precise real-time beam profile monitor (BPM) is desired by the Heidelberg Ion-beam Therapy Center (HIT) to upgrade their original Multiwire chambers. A Scintillating Fibre based BPM will offer real-time information of the beam conditions, including position and width, with a readout rate of 10 kHz using photodiode arrays with a channel size of 0.8mm. Currently the data from the BPM is saved offline for processing and analysis, but the goal is to reconstruct the beam profile in real-time on board the device. A reconstruction algorithm has been designed and the goal is to implement this within the FPGA of the readout electrons. The calibration of the detector as well as the beam reconstruction steps will be presented, describing how radiation damage effects to the fibres will be managed.

ST 6.4 Wed 11:45 ST-H4

Evaluation of HV-CMOS Sensors in a Beam Monitoring System for Ion Therapy — •MARTIN PITTERMANN¹, ALEXANDER DIERLAMM¹, ULRICH HUSEMANN¹, STEFAN MAIER¹, HANS JÜRGEN SIMONIS¹, PIA STECK¹, MATTHIAS BALZER², FELIX EHRLER², IVAN PERIĆ², RUDOLF SCHIMMASEK², and ALENA WEBER² — ¹Institute of Experimental Particle Physics (ETP), Karlsruher Institute of Technology (KIT) — ²Institute for Data Processing and Electronics (IPE), KIT

Ion therapy is an advanced tool for the treatment of cancer by means of irradiation. The characteristic Bragg peak of ionizing radiation creates a highly localized energy deposit. Additionally, very narrow beams (pencil beam) and raster scan techniques are used. This allows the dose distribution to conform to the tumor while minimizing damage to surrounding tissue. Fast and precise feedback of the beam parameters is required for closed-loop control of the beam optics.

We investigate the feasibility of using HV-CMOS pixels sensors to monitor the position, size and shape of such a medical ion beam in real-time. The high intensities encountered in a primary particle beam prohibit the use of traditional single-hit-readout sensors used in high energy particle physics. Instead, a dedicated counting pixel sensor is being developed at the IPE. The radiation hardness and high-rate capabilities of this sensor are tested at the therapeutic ion beam line. Further development steps towards a beam monitoring system replacing the current wire chambers are also discussed.

ST 6.5 Wed 12:00 ST-H4

Coincident Detection of Cherenkov Photons from Electrons for Medical Applications — •Каvен Кооsнк, Ivor Fleck, and Daniel Berker — Center for Particle Physics Siegen, Experimentelle Teilchenphysik, Universität Siegen

The need for medical imaging devices capable of detecting high energy photons prompts research into new detection methods such as Compton camera in nuclear medicine. A new detection method for Compton electrons using Cherenkov radiation is proposed in this work as a proof of principle. Electrons from beta minus decay of Strontium 90/Yttrium 90 with energies up to 2.28 MeV are used. They are directed through a vacuum channel within which an EM field from an electromagnet allows only a specific energy to reach a collimator at the end of the path. After the collimator, the energy spread of the electrons is less than 6% around the nominal energy which can vary between 0.5 and 2.28 MeV. The electrons subsequently reach a radiator material (PMMA) and produce Cherenkov photons, which are detected via a 8x8 Silicon-Photomultiplier array with 64 readout channels. For each electron, the Cherenkov photons are collected within a time-window of 100 ns. The spatial distribution of the Cherenkov photons and their total number are recorded and will be investigated as a function of electron energy, and the results will be compared with theoretical data.

Location: ST-H4

Location: ST-MV

ST 7: Radiation monitoring and dosimetry II

Time: Wednesday 16:15-17:30

ST 7.1 Wed 16:15 ST-H4

Augenlinsendosimetrie für den klinischen Expositionsalltag - Untersuchungen zur Dosisverteilung sowie Risikobewertung hinsichtlich der Kataraktentwicklung — •JENNIFER SCHLÜSS¹, KEVIN KRÖNINGER¹, JÖRG WALBERSLOH² und JENS WEINGARTEN¹ — ¹Technische Universität Dortmund, Dortmund, Deutschland — ²Materialprüfungsamt NRW, Dortmund, Deutschland

Der Jahresgrenzwert für Augenlinsen bei beruflich strahlenexportierten Personen wurde auf Empfehlung der Internationalen Strahlenschutzkommission (ICRP) im Jahr 2017 auf 20 mSv/Jahr herabgesetzt. Dieser Grenzwert beruht nicht auf einer empirischen Absicherung hinsichtlich des immanenten schädigenden Potentials der Arbeitsumgebung. Entsprechend der Arbeitsschutzbestimmungen für das Klinikpersonal wird verhindert, dass sich die Personen im direkten Strahlengang aufhalten, stattdessen wirkt hier Streustrahlung. Zur Detektion der aufgenommenen Strahlendosis ist das Tragen eines Augenlinsendosimeter unabkömmlich. Vorgestellt wird eine Versuchsreihe zur Bestimmung einer optimalen Trageposition für ein Augenlinsendosimeter (TL-DOS-Dosimeter) vorgestellt. Zur Modellierung wird ein Wasserphantom, ein anthropomorphes Phantom und ein PhantomX verwendet, mithilfe derer zunächst Winkelmessungen an TL-DOS-Dosimetern mittels H_p(3)-Dosismessungen durchgeführt werden. Maßnahmen des persönlichen Strahlenschutzes am Auge wie das Tragen einer Brille oder eines Visiers werden hinsichtlich ihrer Wirkung analysiert, um daraus eine Trageempfehlung für das Klinikpersonal abzuleiten.

ST 7.2 Wed 16:30 ST-H4

Implementation and characterization of a virtual Cs-137 gamma irradiation facility — •Lena Olbrich¹, Evelin Derugin¹, Kevin Kröninger¹, Florian Mentzel¹, Jörg Walbersloh², and Jens Weingarten¹ — ¹TU Dortmund — ²Materialprüfungsamt NRW

A Cs-137 gamma irradiation facility is operated at the *Materialprüfungsamt NRW* for research purposes in the field of thermoluminescence dosimetry and calibration of dosimeters, mainly for official personal dose monitoring. The facility is sufficiently characterized for routine dose monitoring. For the development of new dosimeters, a complete simulation of the irradiation facility is useful to perform virtual prototype tests before anything is built. To understand all the processes involved and their influence on the radiation field at different points, e.g. in the reference point for calibration, all relevant aspects of the entire irradiation facility are implemented and investigated successively using Geant4 simulations.

In this talk, the results of the implementation and characterization of a virtual Cs-137 irradiation facility will be presented.

ST 7.3 Wed 16:45 ST-H4

Automatisierte Vermessung des Feldes der Cs-Bestrahlungsanlage am MPA NRW — •POLINA STECHER¹, EVELIN DERUGIN¹, KEVIN KRÖNINGER¹, JENS WEINGARTEN¹ und Jörg WALBERSLOH² — ¹TU Dortmund, Lehrstuhl für Experimentelle Physik IV — ²Materialprüfungsamt NRW

Die Cäsium-137 Bestrahlungsanlage am Materialprüfungsamt NRW wird für die Kalibrierung der thermolumineszenz Personendosimeter und für Forschungsbestrahlungen genutzt. Mit einer bekannten Dosisrate, die vorher mit Hilfe einer Ionisationskammer an einem definierten Referenzpunkt bestimmt wurde, werden die Dosimeter an der gleichen Position bestrahlt und anschließend ausgelesen. Für die Kalibrierbestrahlung ist die Kenntnis der räumlichen Dosisverteilung der Bestrahlungsanlage sehr wichtig. Zum einen lässt sich überprüfen, wie groß und homogen das Bestrahlungsfeld tatsächlich ist und zum anderen feststellen, welche potenzielle Streuquellen im Bestrahlungsraum die Homogenität des Feldes beeinträchtigen. Informationen der räumlichen Dosisverteilung der Cs-Anlage lassen sich aktuell nur aus Simulationen gewinnen. Eine manuelle räumliche Dosismessung konnte erste valide Messergebnisse liefern. Wir präsentieren den Fortschritt einer automatisierten Feldvermessung. Mit Hilfe von Schrittmotoren wird eine Ionisationskammer an verschiedene Positionen innerhalb einer Ebene des Bestrahlungsfeldes platziert. Mit dem Messaufbau soll das gesamte Bestrahlungsfeld der Cäsium-Anlage abgetastet und die zugehörige Dosisverteilung aufgenommen werden.

 $ST 7.4 \ \ Wed \ 17:00 \ \ ST-H4$ Thermoluminescence glow curve generation using generative adversarial networks (GANs) — •Evelin Derugin¹, Florian Mentzel¹, Kevin Kröninger¹, Jens Weingarten¹, and Jörg Walbersloh² — ¹Department of Physics, TU Dortmund University — ²Materialprüfungsamt NRW

Personal dose monitoring is essential for a successful radiation protection program for occupationally exposed persons. The Materialprüfungsamt NRW produces thermoluminescence detectors based on LiF:Mg,Ti. First studies on artificial neural network (ANN) analysis techniques have been successfully performed based on both measured data and artificial glow curves from empirical simulations using parameter interpolation. ANNs require large data sets to be trained before they can be used to predict parameters of new measurements. Therefore the Department of Physics at TU Dortmund is developing multivariate methods for the simulation of thermoluminescence glow curves using GANs in cooperation with the MPA NRW. These glow curves can be used to predict additional information about the irradiation scenario such as the irradiation date or the number of irradiation fractions. In this study, GANs are trained to simulate a glow curve model using a measured data set of 4200 glow curves. Due to a large number of different irradiation times in the data set, the stochastic properties can be used to improve the simulation process. In this talk, we present the comparison of the simulated glow curves with the measured ones and provide information about the performance and the optimization of the neural network.

ST 7.5 Wed 17:15 ST-H4

Location: ST-H4

Neutron Detection using B4C-coated Silicon Detectors — KEVIN ALEXANDER KRÖNINGER, • ALINA JOHANNA LANDMANN, and JENS WEINGARTEN — TU Dortmund, Fakultät Physik, Otto Hahn Str. 4, 44227 Dortmund

He(3) is a frequently used element in neutron detection. However, the world is suffering from an extreme He(3)-shortage increasing the need for alternative detection methods. Semiconductors, commonly known from high energy physics, coated with B(10) as neutron converters, represent a promising alternative. The coating process of the first prototype on-site of the TU Dortmund was performed. Geant4 simulations are used to investigate how specific parameter changes within the coating process can increase the detection efficiency. Results indicate that the detection efficiency is suitable for high neutron flux particle fields which can be found, for example, within the thermal column of a research remains difficult. Promising candidates for that purpose are scintillation detectors enriched with neutron converters.

In this talk, we will present the results obtained from the Geant4 simulations regarding the increase in detection efficiency of coated silicon detectors; and discuss alternative neutron detector candidates, suitable for lower neutron flux.

ST 8: Radiation therapy II

Time: Thursday 11:00-12:00

ST 8.1 Thu 11:00 ST-H4

Development of an end-to-end verification method for Gamma Knife treatments — IRENÄUS ADAMIETZ^{2,3}, JAN BOSTRÖM^{2,3}, MORITZ BUDDE³, •FELINE HEINZELMANN¹, and KEVIN KRÖNINGER¹ — ¹TU Dortmund University, Department of Physics, Dortmund, Germany — ²University Hospital at Ruhr-Universität Bochum, Gamma Knife Zentrum, Bochum, Germany — ³Marien Hospital Herne, University Hospital at Ruhr-Universität Bochum, Clinic for Radiotherapy and Radiation Oncology, Herne, Germany

Gamma Knife is a stereotactic radiosurgery (SRS) instrument using Cobalt-60 radiation sources to treat malign and benign brain tumours and other skull base disorders. With the introduction of inverse radiation planning, which has recently been adopted in Gamma Knife treatment, dose verification is recommended for Gamma Knife radiosurgery. Such plan verification is already standard practice for other teletherapies. For plan verification in Gamma Knife therapy, a new SRS phantom based on film dosimetry and an end-to-end verification

protocol has been developed and implemented. The gamma analysis is used to evaluate the conformance between the measured dose distribution and the dose distribution calculated by the radiation planning system. The focus of this lecture is on the adaptations of the standardized plan verification method to stereotaxy and the associated challenges such as the small target volumes and the high applied dose, which demand high accuracy.

ST 8.2 Thu 11:15 ST-H4 Simulations of silicone-tungsten shieldings for ruthenium eye applicators — •JUSTINE GEMMECKE¹, HENNING MANKE¹, MICHELLE STROTH¹, SASKIA MÜLLER¹, DIRK FLÜHS², and BERNHARD SPAAN¹ — ¹Experimental Physics 5, TU Dortmund — ²Department of Radiotherapy, University hospital Essen Brachytherapy with Ruthenium-106 eye plaques is a standard treatment modality for ocular tumours. A silver calotte with an integrated Ruthenium-106 layer is attached to the patient's eye for a calculated duration, applying the prescribed

Location: ST-H4
dose in order to destroy the tumour tissue. Surrounding healthy tissue and highrisk organs should be spared as much as possible.

To optimize the protection of healthy tissue, precision fit shieldings made of a silicone-tungsten mixture, fixed to the eye plaque surface, are currently being developed. Their dosimetric properties are investigated by means of both simulations in Geant4 and measurements in a water phantom. Simulations based on real patient data also allow retrospective estimation of the dose distribution to the tissue, measuring dose-volume histograms of the irradiated area and an evaluation of the effect of the shielding under clinical conditions.

When developing patient individual shieldings, two main aspects have to be considered. First, the highest possible tungsten fraction in the shieldings has to be determined in order to achieve the maximal protection of the healthy tissue. Second, an easy manufacturing process has to be developed to apply this technique to the clinical routine.

ST 8.3 Thu 11:30 ST-H4

A combination of Brachy- and X-ray-therapy as a novel concept for intraocular tumors — •Henning Manke¹, Dirk Flühs², Justine Gemmecke¹, Saskia Müller¹, Michelle Stroth¹, and Bernhard Spaan¹ — ¹Experimental Physics 5, TU Dortmund — ²Department of Radiotherapy, Essen University Hospital

Eye plaque brachytherapy with the beta emitter Ruthenium-106 is applied to a large fraction of ocular tumors. Due to the emitted particles' steep dose gradient, there is a limitation to the clinically treatable tumor height. Depending on the indications, eyes affected by a tumor larger than the maximum height have to be enucleated in many cases, especially if located close to the posterior pole. As this massively impairs the patients' quality of life, a novel concept to treat such tumors is currently investigated.

External photon irradiation through insensitive parts of the eye can be used

ST 9: Detectors and Applications II

Time: Thursday 14:00-15:30

ST 9.1 Thu 14:00 ST-H4

Polyethylene naphthalate as alpha and heavy nucleus detection material — •KIM TABEA GIEBENHAIN, HANS-GEORG ZAUNICK, ROMAN BERGERT, and KAI-THOMAS BRINKMANN — Justus-Liebig Universität, Gießen, Germany

Polyethylene naphthalate (PEN) is a material with intrinsic scintillation capabilities. Using a thin film of PEN read out by SiPM photo sensors has shown to be an excellent detector for alpha particles with potential applications for alpha spectroscopy. Measurements with alpha sources were conducted to determine the achievable energy resolution. Furthermore, optimization of the detector by means of optical surface coating was studied and results will be discussed.

ST 9.2 Thu 14:15 ST-H4

Technical aspects of simulation-based scatter correction in total-body Positron Emission Tomography (PET) imaging using the uEXPLORER PET scanner — •REIMUND BAYERLEIN, EDWIN K. LEUNG, ZHAOHENG XIE, ERIC BERG, BENJAMIN A. SPENCER, NEGAR OMIDVARI, QIAN WANG, LORENZO NARDO, SIMON R. CHERRY, and RAMSEY D. BADAWI — University of California Davis

Positron Emission Tomography (PET) is a powerful tool for molecular imaging and has brought enhancements to biological research with widespread oncological and clinical adoption. The limited sensitivity of conventional PET scanners with a short axial field of view (AFOV) has been overcome by the uEXPLORER total-body PET scanner with a total AFOV of 194 cm. With a 15-68-fold increase in sensitivity and a spatial resolution of 3.0 mm the uEXPLORER can provide improved image quality, or reduced scan duration, or reduced radioactivity in the subject, or late time point imaging, or some combination of these. This stateof-the-art PET scanner constitutes a paradigm shift in nuclear medicine with the ability to address open questions in medicine and biology. However, the large number of detectors and the widened acceptance angle dramatically increase the data sizes, setting higher demands on the image reconstruction algorithms. Specifically, quantitative techniques for the correction of scattered events become more complex and computationally expensive. Due to the high number of lines of response (92×10^9) scatter correction by direct computation using the Klein-Nishina formula is challenging in total-body PET and Monte-Carlo (MC) methods are preferred. In the spirit of a technical note, this contribution will describe the procedure of scatter correction in total-body PET using MC simulations embedded in a framework using a list-mode ordered subset expectation maximization image reconstruction. The method was developed and validated using phantom studies conducted at the EXPLORER Molecular Imaging Center at UC Davis. In the presentation, mathematical, physical, and computational aspects will be highlighted.

to enhance the applied dose at the tumor apex. A confocal irradiation concept results in a low exposure of healthy tissue while the tumor control dose can be reached in all parts of the target volume. The X-ray therapy can be performed while the Ruthenium-plaque is positioned on the eye. Therefore the plaque itself, which is mainly made of silver, can be used as a beam stopper. This leads to an additional significant reduction of dose in the healthy tissue behind the tumor.

This contribution presents the general concept of the combined therapy, our methods to examine the physical and radiobiological properties and first results.

ST 8.4 Thu 11:45 ST-H4

Monte Carlo simulations of a combination of brachytherapy and external X-ray irradiation for the treatment of intraocular tumors — •Michelle Stroth¹, Henning Manke¹, Justine Gemmecke¹, Saskia Müller¹, Dirk Flühs², and Bernhard Spaan¹ — ¹Experimental Physics 5, TU Dortmund — ²Department of Radiotherapy, University hospital Essen

Brachytherapy with ruthenium-106 eye plaques is a standard modality for the treatment of intraocular tumors. In order to ensure a sufficient tumor apex dose even in larger tumors without exceeding dose limits at neighboring organs at risk, the concept of combining eye plaque brachytherapy with an external beam from a precisely positioned X-ray tube is currently being investigated.

For this purpose, a generic eye model is created using the CAD software Fusion 360 and adapted to the data of real patient cases. The area of the exit window of the X-ray tube is created for various positions depending on the tumor apex, so that different directions of photon irradiation are simulated. A decomposition of the model into many sub-volumes and their implementation in a Monte Carlo simulation in Geant4 allows the analysis of local dose profiles and dose-volume-histograms in organs at risk and the tumor.

This presentation shows the setup of the simulation and first results of combining brachytherapy with external X-ray irradiation.

Location: ST-H4

ST 9.3 Thu 14:30 ST-H4

Two Photon Absorption - TCT: Characterisation of LGAD and other silicon sensors with a newly developed table-top TPA-TCT system — •SEBASTIAN PAPE^{1,5}, ESTEBAN CURRÁS¹, MARCOS FERNÁNDEZ^{1,2}, MICHAEL MOLL¹, RAÚL MONTERO³, F. ROGELIO PALOMO⁴, CHRISTIAN QUINTANA², IVAN VILA², and MORITZ WIEHE^{1,6} — ¹CERN — ²Instituto de Física de Cantabria — ³Universidad del Pais Vasco (UPV-EHU) — ⁴Universidad de Sevilla — ⁵TU Dortmund University — ⁶Universität Freiburg

The Two Photon Absorption – Transient Current Technique (TPA-TCT) uses fspulsed infrared lasers, with photon energies below the silicon band gap to only generate excess charge carriers in a small volume (approximately $1\mu m \times 1\mu m \times 20\mu m$) around the focal point of the laser beam. Therefore, a resolution in all three spatial directions is achieved to characterise silicon devices. Following the initial success of the method, a compact TPA-TCT setup was developed at CERN and first measurements were performed. The setup, measurements on non-irradiated and irradiated PIN diodes, and measurements on LGAD sensors with focus on the gain suppression mechanism will be presented.

ST 9.4 Thu 14:45 ST-H4

Studies towards a Time-of-Flight system equipped with LGADs — •Valerie Hohm¹, Kevin Kröninger¹, Sebastian Pape^{1,2}, and Jens Weingarten¹ — ¹TU Dortmund, Department of Physics — ²CERN

Treatment planning is a crucial part in particle therapy of cancer in order to prevent healthy tissue from being irradiated falsely. One approach to improve the precision in proton therapy is the usage of proton computed tomography (pCT) for imaging. In such a system the paths of the protons as well as their energy loss in a phantom are measured to create an image of the phantom.

To measure the energy loss of protons, a Time-of-Flight (ToF) system can be used. The time an ionising particle needs to traverse two detectors in a given distance depends on its energy. For a precise energy measurement, the time resolution of the used detectors needs to be as small as possible.

In high-energy physics so called Low Gain Avalanche Detectors (LGADs) were developed for the ATLAS and CMS experiment upgrades. These n-in-p silicon sensors with an additional gain layer are designed for a typical charge amplification in the range of 10 to 30 for unirradiated LGADs which results in a time resolution up to 30 ps.

This talk focuses on the usage of LGADs in a ToF system. The feasibility of the system for the energy measurement in a proton tomography system will be presented as well as first measurements with a prototype.

ST 9.5 Thu 15:00 ST-H4

Single Photon Avalanche Diodes with an on-chip integrated preamplifier to improve single photon time resolution — •JONATHAN PREITNACHER¹, WOLF-GANG SCHMAILZL¹, SERGEI AGEEV², and WALTER HANSCH¹ — ¹Bundeswehr University Munich, Neubiberg, Germany — ²The Moscow Engineering Physics Institute- Kashira Hwy, 31, Moscow, Russland, 115409

Silicon photomultiplier (SiPM) are solid-state detectors used for applications requiring high timing resolution and single photon sensitivity and play an important role in various measurement methods in high energy physics or in fields of medical imaging. To further improve the measurements in such applications, an enhanced single photon time resolution (SPTR) on the SiPM is required. For this we designed and implemented a small array of 4x4 single photon avalanche diodes (SPAD) in CMOS 350 nm technology and combined it with an on-chip integrated amplifier. The amplifier is a modified Regulated Common Gate (RCG) circuit and consists of an n-MOS based current follower with additional amplification stages that provides a stable signal and a fast slew rate, which are necessary conditions for a good SPTR. The standard characterization of the SiPM using parameters such as IV-curves, breakdown voltage, dark count rate, crosstalk, gain and afterpulsing shows acceptable results. Furthermore, first measurements of SPTR with a femtosecond laser and an oscilloscope with 10 GHz bandwidth show that it is possible to measure SPTR down to 42 ps FWHM. The results and different contributions to the SPTR are also discussed.

ST 9.6 Thu 15:15 ST-H4

Space applications - Measuring the effect of total ionization dose on field-effect transistors — •ERIK JOZSEF¹, ANDREAS REEH², HANS-GEORG ZAUNICK², KAI-THOMAS BRINKMANN², and UWE PROBST¹ — ¹Technische Hochschule Mittelhessen, Gießen, Germany — ²Justus-Liebig University, Gießen, Germany

Compared to terrestrial applications the utilization of electronics in space environment meets several additional requirements to ensure functional reliability. One of the key requirements is the radiation hardness of the electronic components. Field-effect transistors are vital for modern electronics and are commonly used in power electronics. This presentation shows a method how radiation hardness of switching transistors can be investigated qualitatively. Parameters relevant to operation, such as threshold voltage, parasitic capacitances and leakage currents are to be measured. Measuring methods, process and equipment are presented. EU regional development funding via the EFRE scheme of the State of Hesse is gratefully acknowledged.

ST 10: Total-Body PET

Time: Thursday 16:15-16:55

 Invited Talk
 ST 10.1
 Thu 16:15
 ST-H4

 Artificial intelligence in PET image reconstruction and quantitative analysis
 - •ZHAOHENG XIE — University of California Davis

Positron emission tomography (PET) in vivo visualizes the molecular pathway and is the most sensitive molecular imaging modality routinely applied in clinic. Recent developments in PET technology dramatically increased the effective sensitivity by increasing the geometric coverage or improved time of flight (TOF) resolution. In this talk, I will discuss a few examples of deep learning-based solutions to address the specific challenges in ultra-low-dose or ultra-fast scanning, which enables more convenient and safer clinical practice, medical research and drug screening. The talk will start with an overview of the applications of AI in PET imaging. Then I will provide some frontline applications, which will cover the latest works we published including learning-based PET image reconstruction, scatter correction, motion correction, and kinetic modeling. A special emphasis is on deep-learning-based methods. We will discuss their potential benefits and limitations. The talk will conclude with a few challenging opportunities in various research and clinical applications.

ST 11: Prize Ceremony and Closing Session

Time: Thursday 16:55-17:10

Location: ST-H4

Location: ST-H4

In this last session we would like to take the opportunity to thank all participants for their attendance and contributions. We will announce the winner of this years award for the *Best contribution in the radiation and medical physics devision at the DPG Spring Meeting 2022*. We welcome everyone to celebrate a successful conference with us, to provide some final feedback and to take the chance to meet other participants one last time at this conference.

Particle Physics Division Fachverband Teilchenphysik (T)

Kerstin Borras Deutsches Elektronen-Synchrotron DESY and RWTH Aachen Notkestraße 85 22607 Hamburg kerstin.borras@desy.de

Overview of Invited Talks and Sessions

(Lecture halls T-H15 to T-H39)

Invited Talks

T 1.1	Mon	9:30-10:00	T-H15	From scattering amplitudes to precision predictions for the LHC – •CLAUDE DUHR
T 1.2	Mon	10:00-10:30	T-H15	Tackling new physics at the fringe of precision: Standard Model physics at the LHC -
				•Simone Amoroso
T 1.3	Mon	10:30-11:00	T-H15	Hunt for New Physics at the LHC — •SWAGATA MUKHERJEE
T 27.1	Tue	11:00-11:30	T-H15	First Results From the Next Generation B-Factory Experiment Belle II - • THOMAS
				Kuhr
T 27.2	Tue	11:30-12:00	T-H15	Flavour Anomalies — • Christoph Langenbruch
T 27.3	Tue	12:00-12:30	T-H15	The top quark is still going strong (and electroweak) — • ANDREA KNUE
T 80.1	Thu	11:00-11:30	T-H15	Borexino looks in the direction of solar neutrinos — •LIVIA LUDHOVA
T 80.2	Thu	11:30-12:00	T-H15	Gravitational waves - a new probe of the early Universe — • VALERIE DOMCKE
T 80.3	Thu	12:00-12:30	T-H15	Gravitational wave detectors - current and future challenges — •MICHÈLE HEURS
T 109.1	Fri	11:00-11:30	T-H15	Ten years of Higgs boson measurements: what we know and what we don't know $-$
				•Christian Grefe
T 109.2	Fri	11:30-12:00	T-H15	Future of Silicon Tracking Detectors: LHC Upgrades and Beyond — •GEORG STEIN-
				BRÜCK
T 109.3	Fri	12:00-12:30	T-H15	The dawn of high energy neutrino astronomy — •ELISA RESCONI

Invited Topical Talks

-			
Tue	14:00-14:25	T-H15	Hadronic Jets: Accuracy and Precision of their Reconstruction and Calibration in AT- LAS – •CHRISTOPHER YOUNG
Tue	14.25-14.50	T-H15	Direct searches testing BSM explanations of the flavor anomalies $-$ •ARNE CHRISTOPH
140	11.20 11.00	1 1110	Reimers
Tue	14:50-15:15	T-H15	ATLAS probes QCD measuring photons — •HeBerth Torres
Tue	15:15-15:40	T-H15	The upgrade of the ATLAS trigger to augment the physics reach of Run-3 $-$ •DANIELE
			ZANZI
Tue	14:00-14:25	T-H16	Testing the Standard Model through Gauge-boson Self-interactions — • PHILIP SOMMER
Tue	14:25-14:50	T-H16	Axions and similar particles - how to cover 10^{17} orders of magnitude in mass — •KRISTOF
			Schmieden
Tue	14:50-15:15	T-H16	From GERDA to LEGEND - Hunting no neutrinos — •Christoph Wiesinger
Tue	15:15-15:40	T-H16	Mapping Highly-Energetic Messengers throughout the Universe — •SARA BUSON
Wed	11:00-11:25	T-H15	Hunting XYZ Beasts at Belle and Belle II — •ELISABETTA PRENCIPE
Wed	11:25-11:50	T-H15	Precision tests of the Standard Model using CP violation in B meson decays - • THIBAUD
			Humair
Wed	11:50-12:15	T-H15	Back to the top: charting the bounds of the standard model — •AFIQ ANUAR
Wed	12:15-12:40	T-H15	Dark matter from spin-2 mediators — • STEFAN VOGL
Wed	11:00-11:25	T-H16	Machine Learning for LHC Theory — • ANJA BUTTER
Wed	11:25-11:50	T-H16	Towards high-precision deep learning for astroparticle physics — •CHRISTOPH WENIGER
Wed	11:50-12:15	T-H16	The quest for the mechanism behind the matter-antimatter asymmetry — •JULIA HARZ
Wed	12:15-12:40	T-H16	Towards the lightest dark matter in direct searches — •Belina von Krosigk
	Tue Tue Tue Tue Tue Tue Wed Wed Wed Wed Wed Wed Wed Wed	Tue 14:00-14:25 Tue 14:25-14:50 Tue 14:50-15:15 Tue 15:15-15:40 Tue 14:00-14:25 Tue 14:50-15:15 Tue 14:50-15:15 Tue 15:15-15:40 Wed 11:00-11:25 Wed 11:50-12:15 Wed 11:00-11:25 Wed 11:00-11:25 Wed 11:00-11:25 Wed 11:00-11:25 Wed 11:00-11:25 Wed 11:25-12:40 Wed 11:25-12:40 Wed 11:25-12:40 Wed 11:25-12:40 Wed 11:25-12:40	Tue 14:00-14:25 T-H15 Tue 14:25-14:50 T-H15 Tue 14:50-15:15 T-H15 Tue 14:00-14:25 T-H16 Tue 14:00-14:25 T-H16 Tue 14:50-15:15 T-H16 Tue 14:50-15:15 T-H16 Tue 14:50-15:15 T-H16 Wed 11:00-11:25 T-H15 Wed 11:50-12:15 T-H15 Wed 11:00-11:25 T-H16 Wed 11:25-11:50 T-H16 Wed 11:25-12:40 T-H16

T 81.1	Thu	14:00-14:25	T-H15	LND - A ("Made in Germany") Radiation Monitor Operating at the far Side of the Moon
				— •Sönke Burmeister, Shenyi Zhang, Jia Yu, Zigong Xu, Stephan Böttcher, Robert
				Wimmer-Schweingruber
T 81.2	Thu	14:25-14:50	T-H15	Energetic Pulsar Environments and the Origins of Galactic Cosmic Rays - •ALISON
				Mitchell
T 81.3	Thu	14:50-15:15	T-H15	Looking forward to exciting physics with FASER — •FELIX KLING
T 81.4	Thu	15:15-15:40	T-H15	Astroparticle physics at the LHC: Exploring the forward region with cross-section mea-
				surements — •Hans Dembinski
T 82.1	Thu	14:00-14:25	T-H16	Searches for new scalar particles at the LHC — • DOMINIK DUDA
T 82.2	Thu	14:25-14:50	T-H16	Novel approaches to search for new physics in rare charm decays — •DOMINIK STEFAN
				Mitzel
T 82.3	Thu	14:50-15:15	T-H16	Constraining the Higgs-charm Yukawa coupling with the CMS experiment - •LUCA
				Mastrolorenzo
T 82.4	Thu	15:15-15:40	T-H16	Characterization of H boson events in the $\tau\tau$ decay channel with the full CMS Run-2
				data set — •Sebastian Wozniewski

Invited Talks of the joint symposium SMuK Dissertation Prize 2022 2022 (SYMD) See SYMD for the full program of the symposium.

SYMD 1.1	Mon	14:00-14:25	Audimax	Timeless Quantum Mechanics and the Early Universe — • LEONARDO CHATAIGNER
SYMD 1.2	Mon	14:25-14:50	Audimax	First tritium β -decay spectrum recorded with Cyclotron Radiation Emission
				Spectroscopy (CRES) — • CHRISTINE CLAESSENS
SYMD 1.3	Mon	14:50-15:15	Audimax	Watching the top quark mass run - for the first time! — • MATTEO M. DEFRANCHIS,
				Katerina Lipka, Sven-Olaf Moch
SYMD 1.4	Mon	15:15-15:40	Audimax	Towards beam-quality-preserving plasma accelerators: On the precision tuning
				of the wakefield — •SARAH SCHRÖDER

Prize Talks of the joint Awards Symposium (SYAW) See SYAW for the full program of the symposium.

SYAW 1.1	Wed	14:10-14:40	Audimax	Wie überprüft man die Ziele der Lehramtsausbildung Physik? — •Horst Schecker
SYAW 1.2	Wed	14:40-15:10	Audimax	Astronomy at Highest Angular Resolution - Adaptive Optics, Interferometry and
				Black Holes — •Frank Eisenhauer
SYAW 1.3	Wed	15:10-15:40	Audimax	Turbulence in one dimension — •Alexander M. Polyakov

Sessions

Т 1.1–1.3	Mon	9:30-11:00	T-H15	Invited Talks 1
Т 2.1–2.9	Mon	16:15-18:30	T-H15	QCD (Theorie) 1
Т 3.1–3.9	Mon	16:15-18:35	T-H16	Flavour Physics 1
Т 4.1–4.7	Mon	16:15-18:00	T-H17	Flavour Physics 2
Т 5.1–5.8	Mon	16:15-18:20	T-H18	Electroweak Interactions (Exp.) 1
Т 6.1–6.9	Mon	16:15-18:30	T-H19	Top Quarks: Production (Exp.) 1
Т 7.1–7.8	Mon	16:15-18:15	T-H20	Top Quarks: Properties 1
Т 8.1–8.9	Mon	16:15-18:30	T-H21	Higgs Boson: Decay in Fermions 1
Т 9.1–9.9	Mon	16:15-18:30	T-H22	Search for Supersymmetry
T 10.1–10.9	Mon	16:15-18:30	T-H23	Search for New Particles -1
Т 11.1–11.8	Mon	16:15-18:15	T-H24	Gaseous Detectors
Т 12.1–12.9	Mon	16:15-18:30	T-H25	Pixel Detectors
Т 13.1–13.8	Mon	16:15-18:15	T-H26	Semiconductor Detectors:Radiation Hardness, new Materials and Concepts
Т 14.1–14.9	Mon	16:15-18:30	T-H27	DAQ and Trigger 1
T 15.1–15.7	Mon	16:15-18:00	T-H28	GRID Computing
Т 16.1–16.9	Mon	16:15-18:30	T-H29	Experimental Methods (general) 1
Т 17.1–17.8	Mon	16:15-18:20	T-H30	Gamma Astronomy 1
T 18.1–18.9	Mon	16:15-18:30	T-H31	Neutrino Astronomy 1
Т 19.1–19.8	Mon	16:15-18:20	T-H32	Cosmic Ray 1
Т 20.1–20.9	Mon	16:15-18:40	T-H33	Neutrino Physics without Accelerators 1
T 21.1–21.9	Mon	16:15-18:35	T-H34	Neutrino Physics without Accelerators 2

Т 22.1–22.8	Mon	16:15-18:20	T-H35	Search for Dark Matter 1
T 23.1–23.6	Mon	16:15-17:45	T-H36	Experimental Techniques in Astroparticle Physics 1
T 24.1–24.5	Mon	16:15-17:35	T-H37	Outreach Methods
T 25.1–25.8	Mon	16:15-18:25	T-H38	Data Analysis, Information Technology and Artificial Intelligence
T 26.1–26.9	Mon	16:15-18:30	T-H39	Data Analysis, Information Technology and Artificial Intelligence
Т 27.1–27.3	Tue	11:00-12:30	T-H15	Invited Talks 2
T 28.1–28.4	Tue	14:00-15:40	T-H15	Invited Topical Talks 1
T 29.1–29.4	Tue	14:00-15:40	T-H16	Invited Topical Talks 2
T 30.1-30.9	Tue	16:15-18:30	T-H15	Flavour Physics 3
Т 31.1–31.8	Tue	16:15-18:15	T-H16	Beyond the Standard Model (Theory) 1
Т 32.1–32.8	Tue	16:15-18:15	T-H17	QCD (Exp.) 1
Т 33.1–33.9	Tue	16:15-18:30	T-H18	Top Quarks: Production (Exp.) 2
Т 34.1–34.7	Tue	16:15-18:00	T-H19	Top Quarks: Properties -2
T 35.1–35.8	Tue	16:15-18:15	T-H20	Higgs Boson: Associated Production 1
T 36.1–36.9	Tue	16:15-18:30	T-H21	Higgs Boson: Extended Models 1
Т 37.1–37.9	Tue	16:15-18:30	T-H22	Search for New Particles 2
T 38.1–38.7	Tue	16:15-18:00	T-H23	Search for New Particles 3
T 39.1–39.8	Tue	16:15-18:15	T-H24	Gaseous Detectors 2
T 40.1–40.9	Tue	16:15-18:30	T-H25	Pixel Detectors 2
T 41.1–41.9	Tue	16:15-18:30	T-H26	Calorimeters 1
T 42.1–42.7	Tue	16:15-18:00	T-H27	Detector Systems 1
T 43.1–43.8	Tue	16:15-18:15	T-H28	DAQ and Trigger 2
T 44.1–44.6	Tue	16:15-17:45	T-H29	Experimental Methods (general) 2
T 45.1–45.8	Tue	16:15-18:15	T-H30	Gamma Astronomy 2
T 46.1–46.9	Tue	16:15-18:30	T-H31	Neutrino Astronomy 2
T 47.1–47.8	Tue	16:15-18:15	T-H32	Cosmic Ray 2
T 48.1–48.10	Tue	16:15-18:45	T-H33	Neutrino Physics without Accelerators 3
T 49.1–49.9	Tue	16:15-18:40	T-H34	Neutrino Physics without Accelerators 4
T 50.1–50.7	Tue	16:15-18:00	T-H35	Search for Dark Matter 2
T 51.1–51.8	Tue	16:15-18:20	T-H36	Experimental Techniques in Astroparticle Physics 2
T 52.1–52.6	Tue	16:15-17:45	T-H37	Outreach Methods 2
T 53.1–53.9	Tue	16:15-18:30	T-H38	Data Analysis, Information Technology and Artificial Intelligence 3
T 54.1–54.4	Wed	11:00-12:40	T-H15	Invited Topical Talks 3
T 55.1–55.4	Wed	11:00-12:40	T-H16	Invited Topical Talks 4
T 56.1–56.9	Wed	16:15-18:30	T-H15	Flavour Physics 4
T 57.1–57.8	Wed	16:15-18:15	T-H16	Flavour Physics 5
T 58.1–58.7	Wed	16:15-18:00	T-H17	QCD (Exp.) 2
T 59.1–59.6	Wed	16:15-17:50	T-H18	Neutrino Physics with Accelerators I
T 60.1–60.9	Wed	16:15-18:30	T-H19	Top Quarks: Decay and CP Violation and Mixing Angles
1 61.1-61.8	Wed	16:15-18:15	1-H20	Higgs Boson: Decay in Bosons
T 62.1-62.9	Wed	16:15-18:30	1-H21 T 1122	Higgs Boson: Extended Models 2
1 63.1–63.9	Wed	16:15-18:30	1-H22	Search for New Particles 4
1 64.1–64.9	Wed	16:15-18:30	1-H23	Search for New Particles 5
1 65.1-65.8	wea	16:15-18:15	1-H24 T 1125	Silicon Strip Detectors
1 66.1-66.9	wed	16:15-18:30	1-H25 Т 1126	Semiconductor Detectors: Radiation Hardness, new Materials and Concepts 2
1 6/.1-6/.9	wed	16:15-18:40	1-H20 т ц27	Myon Detectors
T 60.1-00.0	Wed	10:15-16:15	1-П2/ Т U20	Detector Systems 2
T 70 1 70 0	Wed	16:15 18:30	1-п20 Т Ц20	DAQ and Trigger 5 Experimental Methods (general) 3
T 71.1 71.0	Wed	16.15 18.30	1-1129 Т Ц20	Experimental Methods (general) 5
T 72 1 72 9	Wed	16.15 18.30	Т-1150 Т H31	Cosmic Pay 3
T 73 1_73 8	Wed	16.15 - 18.15	T-H32	Cosmic Ray 5
T 74 1_74 8	Wed	16.15 - 18.20	T-H33	Neutrino Physics without Accelerators 5
T 75.1–75.9	Wed	16:15-18:35	T-H34	Neutrino Physics without Accelerators 6
T 76.1–76.9	Wed	16:15-18:30	T-H35	Search for Dark Matter 3
T 77.1–77.8	Wed	16:15-18:15	T-H36	Search for Dark Matter 4
T 78.1–78.7	Wed	16:15-18:00	T-H37	Experimental Techniques in Astronarticle Physics 3
T 79.1–79.9	Wed	16:15-18:30	T-H38	Data Analysis, Information Technology and Artificial Intelligence 4
T 80.1-80.3	Thu	11:00-12:30	T-H15	Invited Talks 3 (ioint session T/EP)
T 81.1-81.4	Thu	14:00-15:40	T-H15	Invited Topical Talks 5 (joint session T/EP)
Т 82.1-82.4	Thu	14:00-15:40	T-H16	Invited Topical Talks 6
T 83.1-83.9	Thu	16:15-18:30	EP-H1	Astroteilchen: Von der Quelle zum Detektor (contributed talks) (joint session
				EP/T)

T 84.1–84.9	Thu	16:15-18:30	T-H15	Flavour Physics
Т 85.1–85.9	Thu	16:15-18:30	T-H16	Beyond the Standard Model (Theory) 2 and QFT and Lattice Gauge Theory 1
Т 86.1–86.7	Thu	16:15-18:00	T-H17	Neutrino Physics (Theory) 1 and Theoretical Astroparticle Physics and Cosmol-
				ogy 1
Т 87.1–87.8	Thu	16:15-18:20	T-H18	Electroweak Interactions (Exp.) 2
Т 88.1–88.9	Thu	16:15-18:30	T-H19	Top Quarks: Production (Exp.) 3
Т 89.1–89.7	Thu	16:15-18:00	T-H20	Higgs Boson: Decay in Fermions 2
Т 90.1–90.8	Thu	16:15-18:15	T-H21	Higgs Boson: Associated Production 2
Т 91.1–91.6	Thu	16:15-17:45	T-H22	Higgs Boson: Rare Decays
Т 92.1–92.7	Thu	16:15-18:00	T-H23	Higgs Boson: Extended Models 3
Т 93.1–93.10	Thu	16:15-18:45	T-H24	Search for New Particles 6
Т 94.1–94.7	Thu	16:15-18:00	T-H25	Silicon Strip Detectors 2
Т 95.1–95.8	Thu	16:15-18:15	T-H26	Pixel Detectors 3
T 96.1–96.8	Thu	16:15-18:15	T-H27	Detector Systems 3
Т 97.1–97.7	Thu	16:15-18:00	T-H28	Electronics 1
T 98.1–98.7	Thu	16:15-18:05	T-H29	Experimental Methods (general) 4
T 99.1–99.9	Thu	16:15-18:35	T-H30	Neutrino Astronomy 4
T 100.1–100.9	Thu	16:15-18:30	T-H32	Cosmic Ray 5
T 101.1–101.9	Thu	16:15-18:30	T-H33	Cosmic Ray 6
T 102.1–102.9	Thu	16:15-18:35	T-H34	Neutrino Physics without Accelerators 7
T 103.1–103.9	Thu	16:15-18:30	T-H35	Neutrino Physics without Accelerators 8
T 104.1–104.7	Thu	16:15-18:05	T-H36	Search for Dark Matter 5
T 105.1–105.10	Thu	16:15-18:45	T-H37	Search for Dark Matter 6
T 106.1–106.9	Thu	16:15-18:30	T-H38	Experimental Techniques in Astroparticle Physics 4
Т 107.1–107.8	Thu	16:15-18:15	T-H39	Data Analysis, Information Technology and Artificial Intelligence 5
T 108	Thu	19:30-21:00	T-MV	General assembly - Particle Physics Division (for DPG members)
T 109.1–109.3	Fri	11:00-12:30	T-H15	Invited Talks 4

Overview

Annual General Meeting of the Particle Physics Division

Thursday 19:30-21:00 T-MV

Sessions

- Invited Talks, Invited Topical Talks, Group Reports, and Contributed Talks -

T 1: Invited Talks 1

Time: Monday 9:30-11:00

Invited Talk T 1.1 Mon 9:30 T-H15 From scattering amplitudes to precision predictions for the LHC - •CLAUDE DUHR — Bethe Center for Theoretical Physics, Bonn University

Scattering amplitudes are the main theory tool to compute precise predictions for collider experiments like the Large Hadron Collider (LHC) at CERN. Over the last decade, we have reached a new level of understanding of the mathematics describing scattering amplitudes. This has resulted in the development of novel powerful computational techniques that are often inspired by cutting-edge results in pure mathematics. We give a review of these recent developments and techniques, and we illustrate their use for precision predictions on several recent milestone computations for LHC observables.

Invited Talk

T 1.2 Mon 10:00 T-H15 Tackling new physics at the fringe of precision: Standard Model physics at the LHC — •SIMONE AMOROSO — DESY, Hamburg

Despite the lack for direct evidence for physics beyond the Standard Model, the potential of the Large Hadron Collider is far from exhausted. The large datasets accumulated, combined with advancements in detector calibrations, data analysis, and theory calculations, allows for measurements of Standard Model paLocation: T-H15

rameters with precision which were unthinkable even a decade ago, and for the observation of new and rarer processes. These high energy and high precision measurements can be used to probe the behavior of the Standard Model at scales well beyond the direct reach of the Large Hadron Collider, providing a promising avenue for the investigation of New Physics. In this talk recent Standard Model results from the LHC Collaborations will be reviewed, and their impact in constraining the Standard Model and its extensions will be illustrated.

Invited Talk T 1.3 Mon 10:30 T-H15 Hunt for New Physics at the LHC — •SWAGATA MUKHERJEE — RWTH Aachen University

The search for new physics is a major goal of the LHC physics program. The excellent quality of the Run-2 data set collected by the LHC experiments provides a promising avenue to search for signatures of physics beyond the Standard Model. In this talk I will review some of the searches from Run-2. These searches have covered a wide range of new physics scenarios including supersymmetry, new hidden sectors, dark matter, and long-lived particles. In addition to reviewing some of the innovative techniques that made the analyses possible, I will summarise what we have learned from the results and briefly discuss prospects for Run-3 which is starting this year.

T 2: QCD (Theorie) 1

Time: Monday 16:15-18:30

T 2.1 Mon 16:15 T-H15

Soft photon bremsstrahlung at Next-to-Leading Power — • DOMENICO BONO-CORE and ANNA KULESZA - Institut für Theoretische Physik, Westfälische Wilhelms-Universität Münster

A long-standing discrepancy in the soft photon bremsstrahlung has attracted a renewed attention in view of the proposed measurements with a future upgrade of the ALICE detector in the upcoming runs of the LHC. In this talk I will discuss the possibility to implement techniques that have been recently developed for soft gluon resummation at Next-to-Leading-Power (NLP) to the soft photon spectrum.

T 2.2 Mon 16:30 T-H15

Leading-Color Two-Loop Amplitudes for Four Partons and a W-Boson in QCD — SAMUEL ABREU 1,2,3 , FERNANDO FEBRES CORDERO⁴, HARALD ITA⁵, •MAXIMILIAN KLINKERT⁵, BENJAMIN PAGE¹, and VASILY SOTNIKOV⁶ - ¹Theoretical Physics Department, CERN, Geneva - ²Higgs Centre for Theoretical Physics, School of Physics and Astronomy, The University of Edinburgh - ³Mani L. Bhaumik Institute for Theoretical Physics, Department of Physics and Astronomy, UCLA, Los Angeles — ⁴Physics Department, Florida State University, Tallahassee — ⁵Physikalisches Institut, Albert-Ludwigs-Universität Freiburg — ⁶Max Planck Institute for Physics (Werner Heisenberg Institute), München

Leading-Color Two-Loop Amplitudes for Four Partons and a W-Boson in QCD I will present the leading-color two-loop QCD corrections for the scattering of four partons and a W boson, including its leptonic decay. The amplitudes are assembled from the planar two-loop helicity amplitudes for four partons and a vector boson decaying to a lepton pair. The analytic expressions are obtained by setting up a dedicated Ansatz and constraining the free parameters from numerical samples obtained within the framework of numerical unitarity. Our results are expressed in a basis of one-mass pentagon functions, which opens the possibility of their efficient numerical evaluation.

T 2.3 Mon 16:45 T-H15

Fast simulations with NNLO QCD accuracy — •LUCAS KUNZ — Karlsruhe Institute of Technology, Karlsruhe, Germany

The calculation of theoretical predictions for hadron colliders at higher orders in perturbation theory involves computing time expensive iterative procedures. The same is true for the extraction of parton distribution functions (PDFs) from measured data. Hence, to produce results in reasonable time, a very efficient and flexible setup is needed. The APPLfast project fulfills these requirements by linking the parton-level Monte Carlo program NNLOjet with both the APPLgrid and fastNLO grid libraries, thereby allowing for an a posteriori choice of a set of Location: T-H15

PDFs or value of the strong coupling constant. This talk will give an overview of the project, focusing on an explanation of the general logic and on possible applications rather than technical details.

T 2.4 Mon 17:00 T-H15

MiNNLO_{PS} for dibosons: matching NNLO QCD with parton showers -•Daniele Lombardi¹, Marius Wiesemann¹, Giulia Zanderighi¹, Gabriël $KOOLE^1$, LUCA BUONOCORE², and LUCA ROTTOL² – ¹Max-Planck-Institut für Physik, Munich, Germany – ²University of Zurich, Zurich, Switzerland

The comparison of theory calculations against experimental measurements for many particle collider processes is nowadays one of the main roads towards the discovery of new physics. The tremendous amount of data collected and the great effort of the experimental collaborations have allowed us to reach an unprecedented accuracy in the measurements of many processes. To avoid that the theory uncertainties become the limiting factor of this comparison, a similar effort from the theory community is demanded. In this talk, I will first motivate why matching our best available fixed-order results with parton showers is a necessary step in this direction. Then I will focus on MiNNLO_{PS}, which is one of the latest methods that have been proposed to embed a next-to-next-to-leading order QCD calculation into a full-fledged Monte Carlo event generator. Originally formulated for single boson production, this method has been recently extended to general color-singlet final states and successfully applied to different diboson processes, such as $Z\gamma$, W^+W^- and ZZ. After a brief review of the underlying idea of MiNNLO_{PS} and a presentation of the main results, I will conclude by highlighting the potential of the method for future applications.

T 2.5 Mon 17:15 T-H15

Automating the calculation of jet functions and beam functions in SCET -GUIDO BELL, •KEVIN BRUNE, GOUTAM DAS, and MARCEL WALD - Center for Particle Physics Siegen, Theoretische Physik 1, Universität Siegen

In perturbative QCD large logarithms can arise in the computation of collider observables. These logarithms can be resummed via factorization theorems within Soft-Collinear Effective Theory(SCET). These factorization theorems include beam functions accounting for the initial-state collinear interactions and jet functions for the final-state collinear interactions. While these functions have been calculated case by case for different observables until now, we are investigating an automated approach for a general class of observables. For this, we study a general phase-space parameterization that factorizes the universal singularities of the functions. We have implemented this framework for different observables, by using the public code "pySecDec" to compute the next-to-nextto-leading order beam and jet function.

T 2.6 Mon 17:30 T-H15

Invertible Networks for the Matrix Element Method — ANJA BUTTER¹, •THEO Heimel¹, Till Martini², Sascha Peitzsch², and Tilman Plehn¹ — ¹Institut für Theoretische Physik, Universität Heidelberg, Germany – ²Institut für Physik, Humboldt-Universität zu Berlin, Germany

For many years, the matrix element method has been considered the perfect approach to LHC inference. We show how conditional neural networks can be used to unfold detector effects and initial-state QCD radiation, to provide the hard-scattering information for this method. We illustrate our approach for the CP-violating phase of the top Yukawa coupling in associated Higgs and singletop production.

Targeting Multi-Loop Integrals with Neural Networks - • RAMON WINTERHALDER^{1,2,3}, VITALY MAGERYA⁴, EMILIO VILLA⁴, STEPHEN P. JONES⁵, MATTHIAS KERNER^{4,6,7}, ANJA BUTTER^{1,2}, GUDRUN HEINRICH^{2,4}, and TILMAN Plehn^{1,2} — ¹Institut für Theoretische Physik, Universität Heidelberg, Germany — 2 HEiKA - Heidelberg Karlsruhe Strategic Partnership, Heidelberg University, Karlsruhe Institute of Technology (KIT), Germany – ³Centre for Cosmology, Particle Physics and Phenomenology (CP3), Université catholique de Louvain, Belgium — ⁴Institut für Theoretische Physik, Karlsruher Institut für Technologie, Germany — ⁵Institute for Particle Physics Phenomenology, Durham University, UK — ⁶Institut für Astroteilchenphysik, Karlsruher Institut für Technologie, Germany — ⁷Physik-Institut, Universität Zürich, Switzerland Numerical evaluations of Feynman integrals often proceed via a deformation of the integration contour into the complex plane. While valid contours are easy to construct, the numerical precision reached for a multi-loop integral can depend critically on the chosen contour. We present methods to optimize this contour using a combination of machine-learned complex shifts and a normalizing flow. This can, potentially, lead to a very significant gain in precision.

T 2.8 Mon 18:00 T-H15

Generative Networks for Precision Enthusiasts - • ANJA BUTTER, THEO HEIMEL, SANDER HUMMERICH, TOBIAS KREBS, TILMAN PLEHN, ARMAND ROUSselot, and Sophia Vent — U. Heidelberg, ITP

Generative networks are opening new avenues in fast event generation for the LHC. We show how generative flow networks can reach percent-level precision for kinematic distributions, how they can be trained jointly with a discriminator, and how this discriminator improves the generation. Our joint training relies on a novel coupling of the two networks which does not require a Nash equilibrium. We then estimate the generation uncertainties through a Bayesian network setup and through conditional data augmentation, while the discriminator ensures that there are no systematic inconsistencies compared to the training data.

T 2.9 Mon 18:15 T-H15

Development of transverse flow for small and large systems in conformal kinetic theory — •CLEMENS WERTHMANN¹, SÖREN SCHLICHTING¹, and VICTOR EUGEN AMBRUS^{2,3} — ¹Universität Bielefeld, Germany — ²Goethe-Universität Frankfurt, Germany — ³West University of Timisoara, Romania

We employ an effective kinetic description to study the space-time dynamics and development of transverse flow of small and large collision systems. By combining analytical insights in the few interactions limit with numerical simulations at higher opacity, we are able to describe the development of transverse flow from very small to very large opacities, realised in small and large collision systems. Surpisingly, we find that deviations between kinetic theory and hydrodynamics persist even in the limit of very large interaction rates, which can be attributed to the presence of the early pre-equilibrium phase. We discuss implications for the phenomenological description of heavy-ion collisions and the applicability of viscous hydrodynamics to describe small and large collision systems.

[1] V.Ambrus, S.Schlichting, C.Werthmann arXiv:2109.03290

T 3: Flavour Physics 1

Time: Monday 16:15-18:35

T 3.1 Mon 16:15 T-H16

Measurement of the ratios $\mathscr{R}(\mathbf{D}^{(*)})$ with leptonic τ and hadronic tag at Belle - •FELIX METZNER¹, FLORIAN BERNLOCHNER², MICHAEL FEINDT¹, and PABLO GOLDENZWEIG¹ for the Belle-Collaboration — ¹Karlsruhe Institute of Technology — ²University of Bonn

The discrepancy observed for the ratios $\mathscr{R}(D^{(*)})$ of the decays $B \to D^{(*)} \tau v_{\tau}$ relative to the normalisation modes $B \to D^{(*)} \ell \nu_{\ell}$ ($\ell = e, \mu$) between the experimental results and the Standard Model (SM) prediction is one of few longstanding tensions of the SM. The new Belle II software framework and the therein included conversion tool B2BII allows to reevaluate the Belle data set of 772 million BBpairs recorded from 1999 until 2010 using the improved algorithms of the modern framework. With this approach a new measurement of the ratios $\mathscr{R}(D^{(*)})$ with an improved hadronic tagging algorithm - the Full Event Interpretation (FEI) — is carried out. Profiting from a higher reconstruction efficiency, due to the new tagging algorithm, this analysis aims to provide new insights into these semileptonic B-decays. In this talk, the procedure and the current status of the analysis will be presented.

T 3.2 Mon 16:30 T-H16

Inclusive B-meson tagging for an $R(D^{(*)})$ measurement at Belle II – THOMAS KUHR, THOMAS LÜCK, and •SOFIA PALACIOS SCHWEITZER — Ludwig-Maximilians-Universität, München

The world average of previous measurements of $R(D^{(*)})$, defined as $R(D^{(*)}) =$

 $\frac{\mathscr{B}(B \to D^{(*)} \tau v_{\tau})}{\mathscr{B}(B \to D^{(*)} l v_{l})}, \text{ shows a 3.4}\sigma \text{ deviation from Standard Model predictions, which}$ could indicate some new physics phenomena, such as the existence of Leptoquarks. This analysis uses data and simulations from the Belle II experiment at the SuperKEKB electron-positron collider to measure $R(D^{(*)})$. To account for the challenge of multiple neutrinos as final state particles, an approach is considered, where besides the signal B-meson decay kinematical and topological properties of the other B-meson are reconstructed fully inclusively. In contrast to exclusive tagging used for previous measurements of $R(D^{(*)})$ by the B-factories, this inclusive tagging approach suffers from a larger background, but also offers a higher reconstruction efficiency. The current status of the analysis will be presented.

T 3.3 Mon 16:45 T-H16 Probing the $R(D^{(*)})$ discrepancy with inclusive $B \rightarrow X \tau v$ decays at Belle II — JOCHEN DINGFELDER, FLORIAN BERNLOCHNER, •HENRIK JUNKERKALE-FELD, and PETER LEWIS - Physikalisches Institut der Rheinischen Friedrich-Wilhelms-Universität Bonn

Location: T-H16 Excesses in the R(D) and $R(D^*)$ ratios measured by multiple experiments have

caused large interest in recent years and align with other measurements in the flavor sector that may hint at non-universality of lepton flavor. Exclusive decays of B mesons into τ leptons $(B \to D^{(*)} \tau \nu)$ seem to appear more frequently compared to the respective decays into light leptons than expected by theory.

The Belle II experiment in Japan enables a complementary test of these measurements. Due to the precise knowledge of the initial state of the collision and the controlled production of $B\bar{B}$ pairs, an inclusive measurement of $B \to X\tau v$ becomes possible. Here, the hadronic system X is not explicitly reconstructed, i.e. all possible hadrons contribute. This approach offers a better statistical precision than exclusive measurements at the expense of larger backgrounds.

In this talk, the current status of the inclusive R(X) measurement is presented. The event selection, the signal extraction strategy and the most important systematic uncertainties are discussed.

T 3.4 Mon 17:00 T-H16

Measuring $R(D^*)$ in hadronic one-prong τ decays at Belle II — FLORIAN BERNLOCHNER, WILLIAM SUTCLIFFE, and •ILIAS TSAKLIDIS for the Belle II-Collaboration - Physikalisches Institut der Rheinischen Friedrich-Wilhelms-Universität Bonn

In this work we are measuring the $R(D^*)$ ratio with hadronically decaying τ leptons. The Belle II experiment is producing exact $B\bar{B}$ pairs and it greatly benefits from the clean experimental environment of e^+e^- collisions. In this study we tag one of the two B mesons using the Full Event Interpretation algorithm in fully hadronic modes, in order to kinematically constrain the second B meson. We reconstruct $B \rightarrow D^* \tau \nu$ decays with a single charged track originating from the τ decay and two missing neutrinos. This gives us a unique access to quantities, sensitive to New Physics, such as the τ lepton polarization besides the $R(D^*)$ ratio. In this talk the reconstruction strategy, the current status and future targets of the analysis will be presented.

T 3.5 Mon 17:15 T-H16

Measurement of R(D*) with semileptonic tagging at Belle II - FLORIAN BERNLOCHNER, JOCHEN DINGFELDER, PETER LEWIS, and •Alina Manthei for the Belle II-Collaboration - Physikalisches Institut der Rheinischen Friedrich-Wilhelms-Universität Bonn

The Belle II experiment at the SuperKEKB asymmetric-energy electron-positron collider is able to collect a large number of events with $B\overline{B}$ pairs. The analysis of semitauonic decays of these B mesons allows for tests of lepton flavour universality. Existing experimental results on the ratios of the branching fractions $\mathscr{R}(D^{(*)} = \mathscr{B}(\overline{B} \to D^{(*)}\tau^-\overline{\nu})/\mathscr{B}(\overline{B} \to D^{(*)}l^-\overline{\nu}),$ where *l* denotes an electron or muon, are in tension with the Standard Model (SM) predictions, which might

hint at physics beyond the SM, such as the presence of charged Higgs bosons or leptoquarks. A combined analysis of $\mathscr{R}(D)$ and $\mathscr{R}(D^*)$ with measurements from Belle, BaBar and LHCb yields a divergence from the SM prediction by approximately 3.8 σ , where σ indicates the standard deviation. Thus, further investigations of these decays with the recently collected data by Belle II are necessary. This talk will present the current status and plans for a measurement of $\mathscr{R}(D^{(*)})$ using this data, while reconstructing the respective other *B* meson in the event in semileptonic modes.

T 3.6 Mon 17:30 T-H16 $m_h(m_Z)$ revisited with Zedometry — •STEFAN KLUTH — MPI für Physik, Föhringer Ring 6, 80805 München

Precision measurements of ${\rm Z}^0$ boson properties can allow a determination of the mass of the b quark at the scale of the Z^0 boson mass $m_b(m_Z)$. The dependence of Standard Model predictions by the programs zfitter and Gfitter on the b quark mass are studied. The precision of the currently available measurements by the LEP experiments and SLD, together with measurements from the LHC experiments for the mass of the top quark and the Higgs boson, is not sufficient for a relevant measurement. The predicted precision of Z⁰ boson resonance measurements at future e^+e^- colliders will allow a competetive determination of $m_b(m_Z)$.

T 3.7 Mon 17:45 T-H16

From sWeights to COWs: News about the sWeight method - \cdot HANS DEMBINSKI¹, MATT KENZIE², CHRISTOPH LANGENBRUCH³, and MICHAEL SCHMELLING⁴ — ¹TU Dortmund, Germany — ²University of Warwick, United Kingdom — ³RWTH Aachen, Germany — ⁴MPIK Heidelberg

A common problem in experimental flavour physics is the separation of signal and background, when the background is difficult to parameterise. We revisit the foundation of the popular method known as sWeights (or sPlot), which allows one to calculate estimates from the signal density in a control variable (for example, the decay time of a particle) using a fit of a mixed signal and background model to a discriminating variable (typically the invariant mass of decay candidates). sWeights are a special case of a larger class of Custom Orthogonal Weight functions (COWs), which can be applied to a more general class of problems in which the discriminating and control variables are not necessarily independent and still achieve close to optimal performance. We present new insights into the properties of parameters estimated from fits to sWeighted data, and provide closed formulas for the asymptotic covariance matrix of these parameters. To illustrate our findings, we show practical applications of these techniques.

T 3.8 Mon 18:00 T-H16

Normalization of the rare $K^+ \rightarrow \pi^+ \nu \bar{\nu}$ decay by a $K^+ \rightarrow \pi^+ \pi^0 / K^+ \rightarrow \mu \nu$ measurement with NA62 — • ATAKAN TUĞBERK AKMETE — Johannes Gutenberg Universität, Mainz

The ultra-rare $K^+ \rightarrow \pi^+ \nu \bar{\nu}$ decay has a precisely predicted Standard Model branching ratio of $8.4 \pm 1.0 \times 10^{-11}$ which is almost free from theoretical uncertainties. Therefore, this SM limit can be tested by a precision measurement. Further, the very high sensitivity of this decay also makes it one of the most suitable candidates to investigate indirect effects of new physics in the flavour sector.

The NA62 experiment at the CERN SPS was proposed and designed to measure this branching ratio by using a decay-in-flight technique. NA62 took data of the $K^+ \rightarrow \pi^+ v \bar{v}$ decay in 2016, 2017, 2018 and 2021. The previous analyses yielded the most precise branching-ratio measurement of the decay. In those analyses, the $K^+ \rightarrow \pi^+ \pi^0$ was used as normalization in order to extract the number of kaon decays for the $K^+ \rightarrow \pi^+ \nu \bar{\nu}$ branching ratio measurement.

However, it is also possible to use another normalization mode, $K^+ \rightarrow \mu \nu$, which can be useful for eliminating undesirable effects that were considered in the previous analyses such as photon and multiplicity rejections. In this talk, I will be discussing the results of this $K^+ \rightarrow \mu \nu$ normalization mode and representing a new measurement on $K^+ \to \pi^+ \pi^0 / K^+ \to \mu \nu$.

Group Report T 3.9 Mon 18:15 T-H16 Towards a Super Charm-Tau Factory in Russia — • MUSTAFA SCHMIDT, SI-MON BODENSCHATZ, LISA BRÜCK, MICHAEL DÜREN, JAN NICLAS HOFMANN, SOphie Kegel, Jhonatan Pereira de Lira, Marc Strickert, Chris Takatsch, LEONARD WELDE, and VINCENT WETTIG - II. Physikalisches Institut, Justus-Liebig-Universität Gießen

The Super Charm-Tau Factory (SCTF) is a future electron-positron collider that is planned to be built in Saroy, Russia. Its center-of-mass energy can be tuned between 2 and 6 GeV for studying a large variety of physics programs. The crabwaste method, a novel collision scheme for particle beams, makes it possible to reach an exceptionally high luminosity up to a value of $\mathcal{L} = 10^{35} \text{ cm}^{-2} \text{ s}^{-1}$ which is 3-4 orders larger than in previous accelerators. In combination with longitudinally polarized lepton beams in the interaction point and an excellent Particle Identification (PID), SCTF will offer the unique possibility to study open questions in the field of particle physics and to find new physics beyond the standard model (SM). Prominent examples are the investigation of CP-violation in the open-charm sector or Charged Lepton Flavor Violation (CLFV) in decays of τ -leptons. This talk will present some highlights of the physics program and the detector design. Our working group in Giessen focuses on the PID design using RICH and DIRC detectors.

T 4: Flavour Physics 2

Time: Monday 16:15-18:00

T 4.1 Mon 16:15 T-H17 Testing Lepton Flavour Universality with $B_s^0 \to \phi \ell^+ \ell^-$ decays using LHCb **data** — Christoph Langenbruch¹, Stefan Schael¹, •Sebastian Schmitt¹, and ELUNED SMITH² - ¹I. Phys. Inst. B - RWTH Aachen - ²University of Zürich

In the Standard Model of Particle Physics (SM), $b \rightarrow s\ell^+\ell^-$ transitions are forbidden at tree-level and may only occur at the loop-level. The branching fractions of these so-called Flavour Changing Neutral Currents (FCNCs) can thus be significantly affected by New Physics (NP) beyond the SM. While in the SM, the coupling of the electro-weak gauge-bosons is Lepton Flavour Universal (LFU), this universality can be broken in NP scenarios. Ratios of branching fractions of semileptonic rare decays with muons and electrons in the final state constitute clean SM tests. Recent measurements of LFU ratios have shown tensions of up to 3.1 σ.

The LHCb detector is located at the Large Hadron Collider (LHC) at CERN and is optimised to study rare b-hadron decays. For this purpose LHCb features high trigger efficiencies, excellent track reconstruction, and particle identification.

This talk gives an overview of the measurement of R_{ϕ} , which benefits from the experimentally clean $B_s^0 \rightarrow \phi \ell^+ \ell^-$ environment. The analysis uses the full Run 1 and Run 2 dataset collected by LHCb which corresponds to 9 fb⁻¹ of integrated luminosity.

T 4.2 Mon 16:30 T-H17 Search for the lepton flavour violating decays $B^0 \to K^*(892)^0 \mu^{\pm} e^{\mp}$ and $B_s^0 \rightarrow \phi(1020) \mu^{\pm} e^{\mp}$ with the LHCb experiment — •JAN-MARC BASELS, AN-DREAS GÜTH, and CHRISTOPH LANGENBRUCH — I. Physikalisches Institut B, RWTH Aachen University

The conservation of lepton flavour in interactions involving charged leptons is a central property of the Standard Model (SM). Thus, every discovery of lepton flavour violation (LFV) would simultaneously be a discovery of new physics.

Location: T-H17

Designed to study the decays of heavy flavour hadrons, the LHCb detector at the Large Hadron Collider (LHC) at CERN allows for the search for LFV in $b \to s\ell^+\ell'^-$ transitions of B-mesons with unprecedented sensitivity. An additional motivation for such searches arises by recent tests of lepton flavour universality in rare $b \rightarrow s\ell^+\ell^-$ decays, which have shown tensions with the SM prediction. Any discovery of lepton flavour non-universality would generally imply the existence of LFV decays.

This talk presents the status of a search for the LFV decays $B^0 \to K^*(892)^0 \mu^{\pm} e^{\mp}$ and $B_s^0 \to \phi(1020) \mu^{\pm} e^{\mp}$, based on a dataset taken with the LHCb detector during Run 1 and Run 2 of the LHC that corresponds to an integrated luminosity of $9\, {\rm fb}^{-1}.$ Particular focus is placed on the study and control of backgrounds and the determination of expected upper limits on the signal branching fraction.

T 4.3 Mon 16:45 T-H17

Search for the lepton flavour violating decays $B^+ \to K^+ e^{\pm} \mu^{\mp}$ with the full dataset of the LHCb experiment — JOHANNES ALBRECHT, •ALEXANDER BATтід, and Elena Dall'Occo — Technische Universität Dortmund

The conservation of lepton flavour in interactions of charged leptons is an important prediction of the Standard Model of particle physics, making searches for lepton flavour violating decays of B mesons an interesting probe for New Physics. In addition, hints of lepton non-universality in $b \rightarrow sll$ transitions (measurements of $R_{K^+},\,R_{K^{\pm 0}})$ imply the violation of lepton flavour conservation. Due to the abundance of produced B-mesons and ability to precisely study them, the LHCb experiment provides an ideal environment for searches for lepton flavour violating decays of B-mesons.

In this talk, the search for the lepton flavour violating decays $B^+ \to K^+ e^{\pm} \mu^{\mp}$ with the LHCb experiment is presented. The analysed data has been recorded during Run 1 and Run 2 of the LHC and corresponds to an integrated luminosity of 9.1 fb^{-1} .

T 4.6 Mon 17:30 T-H17

T 4.4 Mon 17:00 T-H17

Test of Lepton Flavour Universality (LFU) in $\Lambda_b \rightarrow \Lambda \ell \ell$ with the LHCb experiment — FLAVIO ARCHILLI¹, SIMONE BIFANI², VLAD DEDU³, LEX GREEVEN³, SIETSKE KEIJZER³, MICK MULDER⁴, NILADRI SAHOO², MARCO SANTIMARIA⁵, SILVIA SOLE³, PAUL SWALLOW², NIELS TUNING³, MAARTEN VEGHEL³, •CHISHUAI WANG¹, and NIGEL WATSON² — ¹Physikalisches Institut, Heidelberg, Germany — ²University of Birmingham, UK — ³Nikhef, Netherlands — ⁴University of Groningen, Netherlands — ⁵INFN, Frascati, Italy

The Flavour Changing Neutral Current (FCNC) transition $b \rightarrow s\ell\ell$ is highly suppressed in the Standard Model (SM), which makes it susceptible to the impact of possible New Physics (NP).

In the SM, the electroweak interaction does not distinguish between the three generations of leptons. However, several recent studies of LFU using $b \rightarrow s\ell\ell$ processes, e.g. $R_K = \frac{\mathscr{B}(B^1 \rightarrow K^+\mu\mu)}{\mathscr{B}(B^1 \rightarrow K^+ee)}$ and $R_{K^{*0}} = \frac{\mathscr{B}(B^0 \rightarrow K^{*0}\mu\mu)}{\mathscr{B}(B^0 \rightarrow K^{*0}ee)}$ measurements, have shown deviations from the SM expectations. These ratios have the advantage that theoretical uncertainties from the hadronization cancel, which makes them relatively clean observables. For the $\Lambda_b \rightarrow \Lambda\ell\ell$ system, a similar observable $R_\Lambda = \frac{\mathscr{B}(\Lambda_b \rightarrow \Lambda\mu\mu)}{\mathscr{B}(\Lambda_b \rightarrow \Lambda ee)}$ can be defined. The measurement of R_Λ will provide an independent test with respect to the tests of LFU performed with the B-mesons.

This talk will present the ongoing study of the R_{Λ} measurement at LHCb, using the data collected by the LHCb experiment during the years 2011-2012 and 2015-2018.

T 4.5 Mon 17:15 T-H17

Search for lepton flavour violation in four-body charm decays at LHCb — •DANIEL UNVERZAGT — Physikalisches Institut, Heidelberg, Germany

LHCb is playing a leading role in the study of rare and forbidden decays of charm hadrons, which might reveal effects beyond the Standard Model. This talk aims in particular to motivate the search for lepton flavour violation in four-body charm decays. In addition an overview about the current status of an analysis studying neutral charm hadrons decaying to two hadrons and an electron and muon, $D^0 \rightarrow hh\mu e$ with $h = \pi^{\pm}, K^{\pm}$, is provided. **Angular analysis of** $B^0 \to K^* e^+ e^-$ **decays** — MARTINO BORSATO, FABIAN GLASER, and •JIANGQIAO HU — Physikalisches Institut - Universität Heidelberg Over the last decade, the LHCb detector at the Large Hadron Collider (LHC) collected the world largest sample of beauty hadron decays. This dataset allowed to study rare transitions of a b quark into an s quark and a pair of charged leptons $(b \to s \ell \ell)$ with unprecedented precision and unearthed a series of anomalies that could be a sign of dynamics beyond the Standard Model. Namely, the angular distributions of $b \to s \mu \mu$ decays do not agree with theoretical predictions and their rate is lower than that observed in $b \to see$ decays, in contrast to the expectation dictated by the SM symmetry of lepton universality. A key measurement to solve this puzzle is the angular analysis of $B^0 \to K^* e^+ e^-$ decays. In this talk I will present my contribution to this analysis, focusing on the difficult kinematic region of high dielectron invariant mass. Advanced analysis techniques are used to classify signal candidates, characterise background contributions and measure the angular properties of the decay.

T 4.7 Mon 17:45 T-H17 **Angular analysis of the decay** $B^0 \rightarrow K^{*0}\mu^+\mu^-$ — LEON CARUS¹, CHRISTOPH LANGENBRUCH¹, •THOMAS OESER¹, and ELUNED SMITH² — ¹I. Physikalisches Institut B, RWTH Aachen University — ²Physik-Institut, University of Zurich Recently, several tensions between measurements and Standard Model predictions emerged in the area of $b \rightarrow s\ell\ell$ decays in measurements of branching fractions, angular observables, and tests of lepton universality. Additional experimental information, such as angular observables determined from the angular distribution of $b \rightarrow s\ell\ell$ decays, can provide deeper insight into the nature of potential New Physics contributions.

A previous measurement of the angular distribution of $B^0 \to K^{*0}(\to K^+\pi^-)\mu^+\mu^-$ decays, performed by the LHCb collaboration using data collected during Run 1 and 2016, found tensions with Standard Model predictions at the level of 3 standard deviations.

This talk will present an overview of the update of this analysis, including LHCb data collected in 2017 and 2018.

T 5: Electroweak Interactions (Exp.) 1

Time: Monday 16:15-18:20

In 2017, the ATLAS collaboration measured the W-boson mass using *pp*-collision data taken at $\sqrt{s} = 7$ TeV in 2011, resulting in the most precise single measurement with a precision of 19 MeV. We present a revised analysis of the same dataset, improving the fit methods and including a measurement of the width of the W-boson. A precise measurement of these quantities in the decay of the W-boson represent an excellent precision test of the Standard Model (SM).

A detailed comparison of the analysis design between the reanalysis and the 2017 analysis is carried out. Improvements are made in the estimation of the multijet background and the description of some systematic uncertainties. The new fitting method of a profile likelihood fit is studied carefully and cross-checked against the results of the revised analysis.

T 5.2 Mon 16:35 T-H18

Measurement of the mass and width of the *W*-boson at the ATLAS experiment — •PHILIP DAVID KENNEDY¹, JAKUB KREMER¹, PHILIPP KÖNIG², and MATTHIAS SCHOTT¹ — ¹Institute of Physics, Johannes Gutenberg Universität, D-55099 Mainz, Germany — ²Institute of Physics, Rheinische Friedrich-Wilhelms-Universität Bonn, D-53115 Bonn, Germany

We discuss the status of the W-boson mass and width measurements using data from the ATLAS experiment at the LHC. This work utilises a profile-likelihood fit to re-analyse the Run-1 dataset. This method provides an advantage over a χ^2 approach used in the original analysis as the likelihood is minimised over the whole parameter space, including all systematic uncertainties. It is then used to perform the first fit of Γ_W at the LHC. These results are of crucial importance for the EW-fit which bounds possible new physics scenarios. Comparison is then made between these results and those from other experiments. Particular emphasis will be placed on the performance of the profile-likelihood fit and its impact on the uncertainties of parton density functions.

T 5.3 Mon 16:50 T-H18

Measurement of the differential $W \rightarrow e + v$ cross-section at high transverse masses at $\sqrt{s} = 13$ TeV with the ATLAS detector — FRANK ELLINGHAUS and •FREDERIC SCHRÖDER — Bergische Universität Wuppertal

Location: T-H18

The charged-current Drell-Yan (DY) cross-section is measured for the leptonic decay of the W boson $W \rightarrow ev$. While the cross-section at the peak of the W boson mass is known very well, the measurement of the differential cross-section for transverse masses up to $\mathcal{O}(1 \text{ TeV})$ is measured for the first time. In addition, the double-differential cross-section will be measured as a function of the transverse mass of the W boson and the pseudorapidity of the lepton.

The charged-current DY can be used to constrain the density function that describes the partonic content of the proton and to measure fundamental parameters of the Standard Model. In particular, the high m_T^W region of the charged-current DY allows probing new physics by constraining effective field theory parameters, because these parameters are sensitive to small deviations in the cross-section with respect to the theory prediction.

An overview of the cross-section measurement focused on issues related to the reconstruction of the missing transverse momentum in the fake lepton background estimation will be presented. The data has been taken at the ATLAS experiment during Run-2 based on *pp*-collisions at a center-of-mass energy of $\sqrt{s} = 13$ TeV at the LHC.

T 5.4 Mon 17:05 T-H18

Measurement of the differential $W \rightarrow \mu + \nu$ cross section at high transverse masses at $\sqrt{s} = 13$ TeV with the ATLAS detector. — FRANK ELLINGHAUS, FREDERIC SCHRÖDER, and •JOHANNA WANDA KRAUS — Bergische Universität Wuppertal

The cross section of the charged-current Drell-Yan process in the decay $W \rightarrow \mu + \nu$ is measured with data taken with the ATLAS detector from pp-collisions at a center-of-mass energy of $\sqrt{s} = 13$ TeV. The full run 2 dataset of $\mathcal{L} = 139$ fb⁻¹ is analyzed.

While the inclusive cross-section is well-known, a differential measurement at very high transverse masses is done for the first time. The cross-section will also be measured double-differentially in the transverse mass of the W-boson m_T^W and the pseudorapidity of the lepton.

This measurement is important since it can be used to constrain the parton distribution function of the proton as well as electroweak parameters.

A quick overview of the complete analysis will be given while the main focus is on the unfolding strategy via Iterative Bayesian Unfolding.

T 5.5 Mon 17:20 T-H18

A data-driven multijet background estimation method for the measurement of the electroweak Wjj production with the ATLAS experiment — •LISA MARIE BALTES — Kirchhoff-Institute for Physics, University Heidelberg, Germany

The observation and measurement of self-interactions of weak gauge bosons provide an indirect search for physics beyond the Standard Model. The electroweak production of a W boson in association with two jets includes the triple gauge boson vertices WW γ and WWZ and is thus sensitive to the vector-boson-fusion (VBF) production of a W boson. In proton-proton collisions, the characteristic signature of a VBF includes two high-momentum jets at small angles with respect to the incoming beams and a centrally produced lepton-neutrino pair originating from the W boson decay. A significant background for this analysis is the multijet production via the strong interaction where a jet is misidentified as a lepton. It is difficult to model this background since it strongly depends on detector-related quantities such as lepton identification and isolation criteria. Therefore, data-driven techniques are used to estimate this background. In this talk, the results of the multijet background estimation using the matrix method are presented.

T 5.6 Mon 17:35 T-H18

Measurement of angular coefficients of the Z boson production at ATLAS — •JULIAN BLUMENTHAL and STEFAN TAPPROGGE — Institut für Physik, Johannes Gutenberg-Universität, Mainz

A better understanding of QCD production processes at hadron colliders is a key aspect for theoretical predictions of perturbative QCD at higher accuracy. It allows for more precise measurements of Standard Model parameters and background estimations for searches. This contribution focusses on the measurement of angular coefficients that are used to describe the differential cross section of the Z boson production and subsequent decay into leptons in the Collins-Soperframe. These angular coefficients can be used to probe QCD contributions in Z production processes in detail. Two of the coefficients in particular can also be used to make inferences about the effective weak mixing angle. For the measurement the full Run 2 ATLAS dataset with an integrated luminosity of L \approx 139 fb⁻¹ at \sqrt{s} =13 TeV is used, which increases the statistical accuracy significantly at a higher centre-of-mass energy than previous analyses. Major challenges of the measurement using centrally produced charged lepton pairs will be described and expected uncertainties discussed.

T 5.7 Mon 17:50 T-H18

Measurement of the anomalous magnetic moment of the tau lepton in heavy ion collisions with the ATLAS experiment - •Leonie Hermann, Valerie Lang, and Markus Schumacher — Albert-Ludwigs-Universität Freiburg The anomalous magnetic moment of leptons is an important property in the Standard Model of particle physics and is highly sensitive to new physics beyond the Standard Model. At the LHC the anomalous magnetic moment of the tau lepton (a_{τ}) can be measured in ultra-peripheral PbPb collisions exploiting the large photon flux via the partonic process $\gamma \gamma \rightarrow \tau \tau$. Anomalous values of a_{τ} change the total cross-section and differential cross-sections in various kinematic observables. The analysis is based on PbPb collision data with a center-of-mass energy of 5.02 TeV collected with the ATLAS experiment in 2018 with an integrated luminosity of 1.33 nb⁻¹. Events with one leptonically decaying tau lepton, i.e. an electron or muon in the final state, and the other decaying hadronically or leptonically are exploited in the analysis. The sensitivity of the measurement is determined by a maximum likelihood fit to the number of selected events, the shape of kinematic distributions and a combination of both in several signal regions corresponding to different final states. Studies of the expected sensitivity quantified by the length of the confidence interval for a_{τ} will be presented.

T 5.8 Mon 18:05 T-H18

Measurement of $Z\gamma\gamma$ and $ZZ\gamma$ final states with the ATLAS detector at the LHC — •ANKE ACKERMANN and PHILIPP OTT — Kirchhoff-Institute for Physics, Heidelberg University

The Standard Model of Particle Physics (SM) predicts the rare production of triboson states, in which three gauge bosons are produced simultaneously. Although suffering from small cross sections and hence a limited amount of signal events, such triboson states can be studied with the vast amount of data that is collected by the ATLAS detector in Run 2. In addition to validating the predictions of the SM for rare processes, sensitivity to New Physics is given via anomalous quartic couplings of e.g. four neutral gauge bosons. This talk will focus on the analysis of the simultaneous production of ZZy as well as Zyy. In order to determine the cross sections of those processes, it is crucial to separate signal events from events arising through background processes mimicking the signal topology. The most dominant background process contains fake photons, which are non-prompt photons within jets. Different data-driven methods are used to estimate the amount of fake photons in the signal region. After giving a general introduction about the triboson production of the processes $Z\gamma\gamma$ and $ZZ\gamma$, a short summary of the two analyses, including the event selection, the background estimation and a study for effects of New Physics, is presented.

T 6: Top Quarks: Production (Exp.) 1

Time: Monday 16:15-18:30

T 6.1 Mon 16:15 T-H19

Measurement of the Single-Top production cross section in the s-channel at \sqrt{s} =13 TeV with the ATLAS detector — •KREUL KEN — Humboldt-Universität zu Berlin

The production of single top-quarks in electroweak processes (Single-Top) is an important part for the study of the Standard Model and possible extensions. Single-Top production is possible in three channels: t-channel, s-channel and via associated production of a W-boson. In proton-proton collisions at the Large Hadron Collider (LHC), the s-channel has the lowest production cross section and is dominated by many background processes. During the LHC run at 8 TeV, the s-channel was already observed with a significance of 3.2σ using the Matrix Element Method. In this method, the matrix elements for the most important signal and background processes are integrated over the available phase space to compute process likelihoods, which can then be combined to a discriminant. The method is now applied to current ATLAS data at \sqrt{s} =13 TeV to improve the previous result using the higher luminosity of up to 139 fb^{-1} .

T 6.2 Mon 16:30 T-H19

Measurement of the t-channel single top-quark production cross-section in proton-proton collisions at a centre-of-mass energy of 13 TeV with the ATLAS detector — OLGA BESSIDSKAIA BYLUND¹, DOMINIC HIRSCHBÜHL¹, •JOSHUA REIDELSTÜRZ¹, MOHSEN REZAEI ESTABRAGH¹, WOLFGANG WAGNER¹, JOHANNES ERDMANN², and BENEDIKT GOCKE² — ¹Bergische Universität Wuppertal, Wuppertal, Deutschland — ²Technische Universität Dortmund, Dortmund, Deutschland

The measurement of the single top-quark t-channel production cross sections σ_{tq} and $\sigma_{\bar{t}q}$ and their fraction R_t as well as the total cross section $\sigma_{tq,\bar{t}q}$ is presented. These measurements provide a precise test of the standard model and are sensitive to new-physics phenomena by probing the properties of the *Wtb* vertex and placing limits on the CKM matrix element $|V_{tb}|$. Data taken with the ATLAS detector from 2015 to 2018 corresponding to an integrated luminosity of $\mathcal{L} = 139 \,\mathrm{fb}^{-1}$ at a center-of-mass energy of 13 TeV is analyzed using cor-

Location: T-H19

responding samples of simulated events. Requirements are applied to the data selecting events with the signature expected for the signal process. To further enhance the separation between signal and background events a neural network is trained using the Monte Carlo simulated data combining several kinematic variables. The neural network output distribution is then used in a binned profile maximum likelihood fit including all systematic uncertainties to determine the cross sections.

T 6.3 Mon 16:45 T-H19

Measurement of W-boson-associated single-top-quark production in boosted lepton-plus-jets final states with CMS — •CHRISTOPHER MATTHIES¹, JOHANNES HALLER¹, ROMAN KOGLER², and MATTHIAS SCHRÖDER¹ — ¹Institut für Experimentalphysik, Universität Hamburg — ²Deutsches Elektronen-Synchrotron DESY, Hamburg

We present a differential cross section measurement of the associated production of a single top quark and a W boson (tW) in boosted lepton-plus-jets final states in pp collisions at $\sqrt{s} = 13$ TeV with the CMS experiment. Boosted hadronic decays of both the W boson or the top quark are reconstructed as large-radius jets, using the Heavy Object Tagger with Variable R (HOTVR). Deep learning techniques are employed to discriminate tW from top quark pair production and other background processes. It is shown that a measurement up to a transverse momentum of several hundred GeV of the top quark or associated W boson is feasible, extending the phase space covered by previous measurements considerable.

T 6.4 Mon 17:00 T-H19 Measurements of differential cross sections and spin asymmetry in tZq — •DAVID WALTER and ABIDEH JAFARI — DESY, Hamburg, Germany

The associated production of a single top quark and a Z boson in pp collisions at the LHC includes the tZ coupling as well as the coupling of three vector bosons (WWZ) and is therefore a unique process to study the couplings of heavy particles in the SM. The top quark in this process is polarized due to its production

through the weak interaction. Since the top quark decays before it hadronizes, the spin information is conserved in the leptonic decay products and can be measured. In this talk the first differential measurement of the tZq cross section is presented where the full Run-2 data of 138 /fb is used. The tZq cross section is measured at parton and particle level as a function of various kinematic observables including leptons and jets. A maximum likelihood unfolding procedure is exploited to correct for detector and hadronization effects. Also presented is the first measurement of the spin asymmetry in tZq, which is proportional to the top quark polarization.

T 6.5 Mon 17:15 T-H19

Differential cross-section measurement of the tZq process with the ATLAS detector — •NILIMA AKOLKAR and IAN BROCK — Physikalisches Institut, Universität Bonn

The production of a single top quark in association with a Z boson (tZq) is a rare process that has been discovered by the CMS and ATLAS Collaborations. This process is of special interest, as it allows one to probe the couplings of the Z boson to the quark sector and to W boson simultaneously.

This talk will focus on the differential cross-section measurement of the tZq process, analyzed in the trilepton decay channel. The data used was collected with the ATLAS detector during Run 2 of the LHC, corresponding to an integrated luminosity of 139 fb⁻¹. The tZq differential cross-section is measured using different methods of unfolding and the preliminary results will be presented in the talk.

T 6.6 Mon 17:30 T-H19

Towards a WbWb differential cross-section measurement — \bullet Thomas McLachlan — DESY

Top quark pair production is a widely studied process at the Large Hadron Collider (LHC) and is a significant background in many searches beyond the Standard Model (BSM). The WWbb final states of this process interfere with the production of a single top quark in association with a W boson and a b-quark (tWb). Using data from Run-2, I will measure the WbWb production cross-section in a phase space that is maximally sensitive to the interference effects with the goal of improving the modelling of SM processes for BSM searches. An event selection using single lepton events is being developed. In this context, I will present a range of quantities and theoretical parameters that will be used in the differential cross-section measurement.

T 6.7 Mon 17:45 T-H19

Search for single production of top quarks in association with a photon with the ATLAS detector at $\sqrt{s} = 13$ TeV — •BJÖRN WENDLAND, JOHANNES ERD-MANN, and KEVIN KRÖNINGER — Technische Universität Dortmund, Fakultät Physik

Analyses of top quark production in association with a photon are important tests of the Standard Model. They probe top quark properties with respect to the electroweak interaction, such as the top quark charge or the structure of the top quark and photon vertex. Top quark pair production with a photon in leptonic final states was observed and investigated extensively by the ATLAS and CMS collaborations. No significant deviations from the Standard Model expectations were found by now. With the rich datasets collected by the ATLAS and CMS experiments during Run 2 of the LHC programme, it is feasible to observe single production of top quarks in association with a photon. The CMS collaboration reported evidence corresponding to $4.4\,\sigma$ for this process using a partial Run 2 dataset.

In this talk, studies of *t*-channel single production of top quarks with a photon using the full Run 2 dataset collected by the ATLAS detector are presented. As the leptonic decay channel of the top quark is used in this analysis, the final state consists of either an electron or a muon, a jet containing *B* hadrons, missing transverse energy, a photon and an additional jet produced in the forward direction.

T 6.8 Mon 18:00 T-H19

First studies on the Monte Carlo simulation of single-top quark production in association with a photon using effective field theory samples — •NILS JULIUS ABICHT, JOHANNES ERDMANN, and BJÖRN WENDLAND — Technische Universität Dortmund, Fakultät Physik

The observation of Standard Model (SM) single-top quark production in association with a photon with the ATLAS detector is expected to be possible with the full Run-2 dataset. A differential measurement of the process can be used to further test the SM. In particular, the interaction between the photon and the top quark is sensitive to modifications of the electroweak couplings of the top quark. An effective field theory (EFT) approach, with variations of the relevant parameters $c_{\rm tW}$ and $c_{\rm tB}$ is employed in order to further constrain current limits on the respective EFT operators. For a better understanding of the influence of these parameters on the simulation of the process in MADGRAPH, preliminary studies for future EFT interpretations are presented.

T 6.9 Mon 18:15 T-H19

Search for flavour-changing neutral-current interactions of a top quark and a gluon in pp collisions at $\sqrt{s} = 13$ TeV with the ATLAS detector — •WOLFGANG WAGNER, GUNNAR JÄKEL, and DOMINIC HIRSCHBÜHL — Bergische Universität Wuppertal

In the standard model, flavour-changing neutral currents (FCNCs) are strongly suppressed, in particular in the top-quark sector. Any observation of top-quarkrelated FCNCs will thus be a clear sign of physics beyond the standard model. A search is presented for the production of a single top quark via left-handed flavour-changing neutral-current (FCNC) interactions of a top quark, a gluon and an up or charm quark. Two production processes are considered: $u + g \rightarrow t$ and $c + g \rightarrow t$. The analysis is based on proton–proton collision data taken at a centre-of-mass energy of 13 TeV with the ATLAS detector at the LHC. The data set corresponds to an integrated luminosity of 139 fb⁻¹. Events with exactly one electron or one muon, exactly one b-tagged jet and missing transverse momentum are selected, resembling the decay products of a singly produced top quark. Neural networks based on kinematic variables differentiate between events from the two signal processes and events from background processes. The measured data are consistent with the background-only hypothesis, and limits are set on the production cross-sections of the signal processes. Based on the framework of an effective field theory, the cross-section limits are translated into limits on the strengths of the *tug* and *tcg* couplings occurring in the theory.

T 7: Top Quarks: Properties 1

Location: T-H20

11110: 1001100 10.110 10.110

T 7.1 Mon 16:15 T-H20

Measurement of the top quark pole mass using $t\bar{t}+1$ jet events with the CMS experiment — •Sebastian Wuchterl¹, Katerina Lipka¹, and Matteo Defranchis² — ¹Deutsches Elektronen Synchrotron (DESY) — ²CERN

The top quark is the most massive elementary particle known. Its mass, m_t , is a fundamental parameter of the Standard Model (SM), and its value needs to be determined experimentally. A precise measurement of m_t and the masses of the W and Higgs bosons play a crucial role in precision tests of the SM. Additionally, the value and the uncertainty of m_t are driving predictions for the energy dependence of the Higgs quartic coupling, which determines the stability of the electroweak vacuum. In proton-proton collisions at the LHC, top quark-antiquark (tt̃) production can be used to extract m_t in different renormalization schemes.

In this work, the pole mass of the top quark is measured using events in which the t \tilde{t} system is produced in association with one additional jet. This analysis is performed using proton-proton collision data collected by the CMS experiment in 2016-2018 with $\sqrt{s} = 13$ TeV, corresponding to a total integrated luminosity of 138 fb⁻¹. Events with two opposite-sign leptons in the final state are analyzed to measure the normalized differential cross section as a function of the inverse of the invariant mass of the t \tilde{t} +1 jet system. This observable has been chosen due to strongest sensitivity to m_t at the threshold of the t \tilde{t} pair production.

T 7.2 Mon 16:30 T-H20 Measurement of the jet mass distribution in hadronic decays of boosted top quarks and determination of the top quark mass with CMS – •ALEXANDER PAASCH¹, JOHANNES HALLER¹, ROMAN KOGLER², and DENNIS SCHWARZ³ – ¹Institut für Experimentalphysik, Universität Hamburg – ²DESY, Hamburg – ³Austrian Academy of Sciences, Vienna

The top quark is the heaviest known elementary particle. Due to its high mass it plays an important role in the electroweak sector of the Standard Model and the measurement of its properties is of special interest. In contrast to conventional top quark mass measurements, we provide an analysis in the boosted regime. At these high energies, the top quark decay products are collimated and are clustered into a single large-radius jet with a mass sensitive to the top quark mass.

In this talk, we present the measurement of the jet mass distribution and top quark mass in hadronic decays of boosted top quarks, using 137 fb^{-1} of data collected by the CMS experiment during the LHC Run-2. New techniques such as a refined calibration of the jet mass scale and improving the description of the final state radiation through a measurement of jet substructure variables substantially increase the precision compared to earlier analyses. The result represents a large step towards the precision observed in measurements at threshold production.

Time: Monday 16:15–18:15

T 7.6 Mon 17:30 T-H20

T 7.3 Mon 16:45 T-H20

Measurement of the top-quark mass in $t\bar{t}$ events using the template method in the lepton+jets channel with the ATLAS detector — •DIMBINIAINA RAFANOHARANA and ANDREA KNUE — Albert-Ludwigs-Universität Freiburg The top-quark mass is a fundamental parameter of the Standard Model (SM). Its precise determination is therefore crucial to test the consistency of the SM. A multitude of measurements was performed at the Tevatron and the LHC using different methods and final states.

The combination of the ATLAS measurements at $\sqrt{s} = 7$ TeV and at $\sqrt{s} = 8$ TeV has a relative overall uncertainty of 0.28% and a relative statistical uncertainty of 0.14%. The measurement precision is therefore limited by the understanding of systematic effects.

In this presentation, an investigation of the systematic effects in the top-quark mass measurement using the template method in $t\bar{t} \rightarrow$ lepton+jets channel at $\sqrt{s} = 13$ TeV is shown. The studies are performed using different observables sensitive to the top-quark mass as well as different event selections, aiming at reducing the overall uncertainty.

T 7.4 Mon 17:00 T-H20

Neural network based estimators to measure the top quark mass — CHRISTOPH GARBERS, JOHANNES LANGE, •NATHAN PROUVOST, PETER SCHLEPER, and HARTMUT STADIE — Universität Hamburg, Luruper Chaussee 149, 22761 Hamburg

The top quark is the heaviest known particle in the Standard Model. As such, the top quark mass is an important parameter for constraining and checking the validity of Standard Model predictions. In the semi-leptonic decay channel, a value of the top quark mass of 172.25 \pm 0.63 GeV has been measured with 35 fb⁻¹ of the 2016 data at the CMS experiment. The mass of the top quark from a kinematic fit and the reconstructed mass of the W boson are used as variables. It is expected that adding more variables will improve the measurement. This presentation focuses on the development of neural network based estimators of the top quark mass using nuisance parameters for the systematics and multiple observables.

T 7.5 Mon 17:15 T-H20

Measurement of the top quark width from Wb scattering — THORSTEN CHWALEK¹, MATTEO DEFRANCHIS², NILS FALTERMANN¹, JAN KIESELER², MATTHIAS KOMM², •MARCO LINK¹, MARTIJN MULDERS², THOMAS MÜLLER¹, MICHAEL PITT², and PEDRO SILVA² — ¹Institut für Experimentelle Teilchenphysik (ETP), Karlsruher Institut für Technologie (KIT) — ²CERN

The top quark is the heaviest elementary particle in the standard model (SM). As such it is especially interesting to search for hints of physics beyond the standard model (BSM). Current measurements of the top quark decay width make strong assumptions on the branching ratio of the top quark into a W boson and a bottom quark (Wb).

This talk explores the possibility of the model independent measurement of the top quark decay width in Wb \rightarrow Wb scattering via a top quark propagator at the CMS experiment. Generator level studies are used to develop reconstruction methods for the Wb system. In combination with a measurement of the single top cross section this measurement could be sensitive to effects from BSM physics. Given the limited size of the current data set and the small cross section of the signal process, the measurement will be interesting with the increase of data expected during HL-LHC data taking.

Measurement of top-quark pair spin correlation in the ℓ + jets channel using the ATLAS experiment — •OLEKSANDR BURLAYENKO, ANDREA KNUE, and ZUZANA RURIKOVA — Albert-Ludwigs Universität Freiburg, Experimentelle Teilchenphysik AG Herten

The top quark is the heaviest known fundamental particle and has a lifetime on the order of 10^{-25} s. This lifetime is shorter than the quantum chromodynamic (QCD) hadronization time scale $1/\Lambda_{QCD} \approx 10^{-24}$ s, and much shorter than the spin decorrelation time scale $m_t/\Lambda_{QCD}^2 \approx 10^{-21}$ s. This gives an opportunity to study the spin properties of a bare quark, as top-quark spin information is preserved in the angular distribution of its decay products.

The Standard Model predicts the $t\bar{t}$ pairs to have correlated spins. The degree of the this correlation is sensitive to the production mechanism of the top quark. The ATLAS collaboration measured it at 13 TeV in the dilepton channel. In this measurement, a discrepancy between the predicted and observed results was found.

This work presents ongoing studies of the $t\bar{t}$ spin correlation in the lepton + jet channel at $\sqrt{s} = 13$ TeV. While this channel provides a larger dataset to study, the analyzing power is reduced compared to the dilepton channel. In this talk, first studies will be presented including different event selections and different observables for this final state.

T 7.7 Mon 17:45 T-H20

Extracting top-Yukawa coupling from $t\bar{t}$ cross-section using ATLAS data — •SUPRIYA SINHA — DESY, Hamburg and Zeuthen, Germany

This work aims to extract the top-Yukawa coupling (Y_t) from $t\bar{t}$ cross section close to the threshold. In order to achieve this, one can use the kinematic distributions in $t\bar{t}$ production along with the virtual Higgs boson loop correction. This boson exchange modifies the differential distributions near $t\bar{t}$ production threshold energy. It becomes highly sensitive to Y_t , and hence, is used to extract its value.

This talk introduces the involved physics processes and gives an insight to the analysis strategy. The decay channel considered for the analysis is the lepton+jets final state. Full Run-II data with the integral luminosity of 139 fb⁻¹ taken from the ATLAS experiment at 13 TeV, is used.

T 7.8 Mon 18:00 T-H20

Sensitivity studies for the measurement of the top-Yukawa coupling using four-top final states — ARNULF QUADT, ELIZAVETA SHABALINA, and •SREELAKSHMI SINDHU — II. Physikalisches Institut, Georg-August-Universität Göttingen

The top quark is the heaviest particle in the Standard Model and hence a precise measurement of its properties is key to identifying evidence for physics beyond the Standard Model. One such property is the top-Yukawa coupling, which describes the strength of the interaction between the top quark and the Higgs boson. The production of four top quarks can be mediated by the Higgs boson, making this process highly sensitive to the top-Yukawa coupling. Various kinematic variables from the decay of the four top process in the trilepton and same-sign dilepton channels are studied to identify the observables that are most sensitive to the top-Yukawa coupling. To get better sensitivity, the top quarks are reconstructed to directly probe the properties of the top quark. The neural network is studied to further improve the sensitivity of the four top quark production to the top-Yukawa coupling. In this talk, a summary of these sensitivity studies will be presented.

T 8: Higgs Boson: Decay in Fermions 1

Location: T-H21

Time: Monday 16:15-18:30

T 8.1 Mon 16:15 T-H21

Measurement of Higgs boson production cross sections in the di- τ decay channel with the ATLAS detector and the combination with other decay channels — •FRANK SAUERBURGER, KARSTEN KÖNEKE, CHRISTOPHER YOUNG, and KARL JAKOBS — Albert-Ludwigs-Universität Freiburg, Deutschland

The coupling of the Higgs boson to τ -leptons is one of the most precisely measured couplings of the Higgs boson to fermions. A measurement of the production cross sections of the Higgs boson decaying into two τ -leptons is presented. The cross section is measured in the gluon-fusion, vector-boson-fusion, W/Z boson associated, and top-quark pair associated production channels. The study illustrates the application of machine-learning techniques. The analysis uses proton-proton collision data at a center-of-mass energy $\sqrt{s} = 13$ TeV corresponding to an integrated luminosity of 139 fb⁻¹ recorded during Run 2 with the ATLAS detector at the LHC.

In addition, the combination of this measurement with other Higgs boson production and decay channels is presented. The presentation focuses on the measurement of coupling modifiers in the κ framework.

T 8.2 Mon 16:30 T-H21 Optimization of di-tau mass reconstruction in the ATLAS experiment using a deep neural network — KLAUS DESCH, PHILIP BECHTLE, CHRISTIAN GREFE, LENA HERRMANN, and •RAMY HMAID — Rheinische-Friedrich-Wilhelms-Universität Bonn

A major challenge of identifying Higgs decays to tau leptons is the similarity of Higgs- and Z-Boson decays, due to their similar mass and at least least two unobservable neutrinos in the final state.

We will present a regression neural network that determines the invariant mass of the reconstructed di-tau system and compare its performance with existing solutions for this problem. In addition, the stability of the neural network response under different training conditions will be discussed.

T 8.3 Mon 16:45 T-H21

Improving the sensitivity of CP tests in VBF Higgs-boson production exploiting the $H \rightarrow \tau_e \tau_\mu$ decay with neural networks for the reconstruction of the Higgs-boson four-momentum at the ATLAS experiment — •ALEXANDRA SPITZER, MICHAEL BÖHLER, and MARKUS SCHUMACHER — Albert-Ludwigs-Universität Freiburg

85

Violation of CP invariance is one of the three Sakharov conditions needed to explain the observed baryon asymmetry in our universe. In the Standard Model CP violation is already introduced via the CKM matrix. However, its size is not sufficient to explain the amount of observed baryon asymmetry. The vector-boson fusion production of the Higgs particle allows to study its CP structure of the couplings to electroweak gauge bosons *HVV*.

CP invariance of *HVV* couplings is probed by investigating the mean value of the CP-odd Optimal Observable. Since the Higgs bosons's four-momentum is required for calculating the Optimal Observable the resolution of the Optimal Observable is limited by the reconstruction quality of the Higgs bosons' four-momentum. It is investigated how neural networks improve the reconstruction quality of the four-momentum in comparison to the Missing Mass Calculator. The hyperparameters of four different neural networks were optimized by using a hyperparameter optimization software framework OPTUNA. The impact on determing the strength of CP violation in the $H \rightarrow \tau_e \tau_\mu$ decay channel assuming $\mathcal{L} = 139 \text{ fb}^{-1}$ of pp collisions at $\sqrt{s} = 13$ TeV is presented.

T 8.4 Mon 17:00 T-H21 Signal selection and background estimation for testing CP invariance in vector boson fusion production of the Higgs boson in the $H \rightarrow \tau_e \tau_\mu$ decay channel using the ATLAS detector — •YE JOON KIM, VALERIE LANG, Ö. OĞUL ÖN-CEL, and MARKUS SCHUMACHER — Albert-Ludwigs-Universität Freiburg

The combination of charge conjugation symmetry (C) and parity symmetry (P), together denoted as CP invariance, is violated in the Standard Model (SM) of particle physics in the weak interaction. The amount of CP violation in the SM is not enough to explain the observed baryon asymmetry in the universe. Further sources of CP violation may exist in the Higgs sector. The CP nature of the Higgs boson coupling to vector bosons is investigated in vector-boson fusion production (VBF) of the Higgs boson exploiting the $H \rightarrow \tau_e \tau_\mu$ decay channel.

In this talk, contributions to the search for CP violation in VBF production in the $H \rightarrow \tau_e \tau_\mu$ channel based on *pp* collision data collected with the ATLAS detector at $\sqrt{s} = 13$ TeV corresponding to an integrated luminosity of 139 fb⁻¹ are discussed.

The estimation of background processes, in particular from events where jets are falsely identified as electrons or muons, with the data-driven matrix method is presented. Studies of the signal selection will also be shown.

T 8.5 Mon 17:15 T-H21

Signal selection and background estimation for testing CP invariance in vector boson production of the Higgs boson in the $H \rightarrow \tau_{lep} \tau_{had}$ decay channel using the ATLAS detector — •HELENA MOYANO, VALERIE LANG, Ö. OĞUL ÖNCEL, and MARKUS SCHUMACHER — Albert-Ludwigs-Universität Freiburg

The existence of charge conjugation and parity symmetry (CP) violating processes represents one of the three Sakharov conditions for baryogenesis, the mechanism that explains the different abundances of matter and antimatter in the universe. The Standard Model of particle physics includes a description of CP violating processes, however, their contribution is not sufficient to explain the observed baryon asymmetry in the universe. The discovery of the Higgs boson in 2012 opened the window to search for new sources of CP violation in the production or decay of the Higgs boson.

The analysis exploits Higgs-boson production via vector-boson fusion (VBF) and the decay $H \rightarrow \tau_{lep} \tau_{had}$ to search for CP-violating contributions to the HVV vertex based on data collected by the ATLAS detector at $\sqrt{s} = 13$ TeV corresponding to an integrated luminosity of $\mathscr{L} = 139$ fb⁻¹.

This presentation discusses the selection of $\tau_{lep}\tau_{had}$ events with neural networks as well as the estimations of the background contributions. Particularly, first studies will be shown regarding the estimation of the background contribution due to jets misidentified as hadronically decaying τ leptons using the Fake Factor method.

T 8.6 Mon 17:30 T-H21

Event reconstruction techniques in the context of a Higgs boson CP analysis in the di-tau lepton final state with the CMS experiment — WOLFGANG LOHMANN, ACHIM STAHL, and •ALEXANDER ZOTZ — RWTH Aachen University - Physics Institute III B, Aachen, Germany

In 2020 the first measurement of the effective CP mixing angle in Higgs boson decays into two tau leptons has been performed by the CMS experiment. It was determined to be $(4 \pm 17)^{\circ}$ using the Run 2 data set of pp collision of $137 f b^{-1}$ integrated luminosity. The mixing angle was extracted from a distribution of angles between the decay planes of the tau lepton decay products in the $H \rightarrow \tau \tau$ decay. In the case of hadronic tau lepton decays via the intermediate a_1 resonance the full tau lepton kinematics including its neutrino and furthermore its polarimetric vector can be reconstructed. Requiring both tau leptons to decay

via a_1 mesons allows for the reconstruction of a CP sensitive observable with higher sensitivity. However the a_1a_1 final state suffers from a small branching fraction and therefore these improvements have a neglible effect on the overall sensitivity once all final states are included.

In this talk, an extension of the polarimetric vector method via the inclusion of final states with an a_1 decay on one side and a single charged lepton or hadron on the other side of the $H \to \tau \tau$ decay is presented. To reconstruct the event a kinematic fit with external constraints is used and the potential improvement on the measurement of the CP mixing angle is discussed.

T 8.7 Mon 17:45 T-H21

Tau reconstruction exploiting machine learning techniques at CMS $- \cdot$ ZE CHEN - DESY, Hamburg, Germany

Reconstruction of hadronically decaying tau leptons (denoted as τ_h) in the CMS experiment at the Large Hadron Collider has been historically performed with the Hadron-plus-strip (HPS) algorithm. In the HPS algorithm, the τ_h final state signature is identified by combining information from charged hadrons, reconstructed by their associated tracks, and π_0 candidates, obtained by clustering photon and electron candidates in rectangular regions, called "strips". As of the LHC Run 2, deep-learning techniques have been implemented to improve the identification of genuine τ_h leptons and reduce contributions from backgrounds. This talk covers a study to improve the tau decay mode reconstruction using machine learning techniques. Its efficiency is shown and compared to the one of the HPS algorithm.

T 8.8 Mon 18:00 T-H21

Search for lepton-flavour violating decays of the Higgs boson using the symmetry method for background estimation with the ATLAS experiment at $\sqrt{s} = 13$ TeV — •KATHARINA SCHLEICHER, VALERIE LANG, and MARKUS SCHUMACHER — Albert-Ludwigs-Universität Freiburg

The discovery of the Higgs boson opened the window to a variety of interesting probes to physics beyond the standard model (SM), including searches for lepton-flavour violating (LFV) Higgs-boson decays. Such decays are predicted in several extension of the SM e.g. in models with two Higgs doublets. In nature, LFV was already observed in form of neutrino oscillations. In this analysis the decays of $H \rightarrow e\tau$ and $H \rightarrow \mu\tau$ with leptonic τ -decays leading to $e\mu + 2\nu$ final states are considered. A central part of the analysis is the precise estimation of the SM backgrounds. Therefore, a data-driven method is used - the so-called symmetry method. It exploits two principles: First, SM backgrounds with prompt leptons are symmetric w.r.t. the interchange of electrons and muons. And second, this symmetry is broken if the branching ratios of the two LFV decays are of different magnitude. The first principle implicates the challenge of restoring this symmetry since electrons and muons are experimentally different. The second principle is motivated by the upper limit on the $BR(\mu \rightarrow e\gamma)$. To obtain the best possible sensitivity, a dedicated statistical model was developed and neural networks for classification are deployed. In this talk, an overview of the analysis using the LHC Run-2 dataset recorded with the ATLAS detector in pp collisions at $\sqrt{s} = 13$ TeV is given.

T 8.9 Mon 18:15 T-H21

Sensitivity to lepton flavour violating Higgs boson decays at the HL-LHC using data-driven background estimation — •NAMAN KUMAR BHALLA, KATHA-RINA SCHLEICHER, VALERIE LANG, and MARKUS SCHUMACHER — Albert-Ludwigs-Universität Freiburg

One of the primary goals of the Large Hadron Collider (LHC) program is to look for processes beyond the Standard Model (SM) of particle physics. One such process predicted by many beyond-SM theories is lepton flavour violation (LFV) in the decays of the Higgs Boson. A search for LFV decays of the Higgs boson with $H \rightarrow e \tau_{\mu}$ and $H \rightarrow \mu \tau_{e}$ final states was performed using the full Run 2 data collected at $\sqrt{s} = 13$ TeV, corresponding to an integrated luminosity of 138 fb⁻¹. A part of this analysis used a data-driven background estimation, which takes advantage of the idempotency of SM backgrounds under the exchange of an electron and a muon. This symmetry is then broken only by the difference in the two LFV signals considered. Due to its data-driven nature, the sensitivity of this approach is limited by statistical uncertainties, which are expected to reduce with more data.

This talk describes the extrapolation of the Run 2 analysis' sensitivity to conditions at the high-luminosity LHC (HL-LHC), where a data set, collected in *pp* collisions at $\sqrt{s} = 14$ TeV corresponding to an integrated luminosity of 3000 fb⁻¹, is expected. The extrapolation also accounts for the expected improvements in systematic uncertainties from detector upgrades planned for the HL-LHC. The first expected sensitivities for LFV decays of the Higgs boson at the HL-LHC, based on the data-driven background estimation, are presented.

T 9: Search for Supersymmetry

Time: Monday 16:15-18:30

T 9.1 Mon 16:15 T-H22

Search for Higgsino production in SUSY scenarios with a compressed mass spectrum — •YUVAL NISSAN, SAM BEIN, PETER SCHLEPER, and GUDRID MOORTGAT-PICK — Universität Hamburg

A search for leptonic decays of Higgsino-like neutralinos in the case of a compressed mass spectrum using two soft lepton tracks and missing transverse momentum is presented. We consider the case of a second-lightest neutralino decaying into a dark matter candidate - lightest neutralino - and two leptons via an off-shell Z boson. In the case of a small mass differences between the neutralinos, the leptons produced are very soft, making them difficult to reconstruct at CMS. Signals of different mass splitting are probed and interpreted within a set of simplified models. Multivariate discriminants are employed in the event and object-level selection, and their performance is studied.

T 9.2 Mon 16:30 T-H22

Search for Higgsinos in final states with a low-momentum, displaced track at the CMS experiment — SAMUEL BEIN, YUVAL NISSAN, PETER SCHLEPER, ALEXANDRA TEWS, and •MORITZ WOLF — Universität Hamburg

Many supersymmetric extensions to the Standard Model predict the three lightest electroweakinos, χ_2^0 , χ_1^{\pm} , χ_1^0 , to be Higgsino-like with similar masses around the electroweak scale. The lightest chargino and the second-lightest neutralino can be pair-produced and decay to the lightest neutralino. To search for these particles, the best strategy depends on the differences between their masses. For $\Delta m(\chi_2^0, \chi_1^0) > \mathcal{O}(1 \text{ GeV})$ lepton pairs from the decay of the second-lightest neutralino leave an experimentally distinct signature, whereas $\Delta m(\chi_1^{\pm}, \chi_1^0) \leq 0.3 \text{ GeV}$ can lead to the chargino giving rise to a disappearing track. However, mass splittings in the range of $\Delta m(\chi_1^{\pm}, \chi_1^0) = 0.3 - 1.0 \text{ GeV}$ are still unexplored at the LHC.

In this analysis, a slightly displaced track with low momentum, corresponding to a pion originating from the chargino decay, is used to gain sensitivity to this intermediate range of mass splittings.

T 9.3 Mon 16:45 T-H22

Search for disappearing tracks with the CMS experiment at $\sqrt{s} = 13$ TeV – •VIKTOR KUTZNER¹, SAMUEL BEIN¹, SEH WOK LEE², SANG-IL PAK², PE-TER SCHLEPER¹, and SEZEN SEKMEN² – ¹Institute for Experimental Physics, Hamburg University, Luruper Chaussee 149, D-22761 Hamburg, Germany – ²Kyungpook National University, Daegu, South Korea

Long-lived heavy particles are often predicted in BSM theories with a small mass splitting between the two lightest new particles, for example a chargino and a neutralino in supersymmetry. Given a sufficiently small mass splitting in the range of $m_{\pi} \leq \Delta m \leq 200$ MeV, the chargino is expected to decay in the CMS tracker volume into soft non-reconstructed leptons or hadrons and a lightest supersymmetric particle, leaving a short track that then seems to disappear. This signature is characterized by missing hits in the outer layers of the tracker with little or no energy deposited in the calorimeter. In addition to events with one or more disappearing tracks, events with an additional lepton are considered as well to account for a second very long-lived chargino, which decays outside the tracker volume. For both topologies events with additional b-quark jets are investigated to account for gluino-/squark-associated chargino production. Datadriven methods are used to determine the dominant backgrounds arising from prompt leptons and fake tracks. Results are presented using proton-proton collision data with $\sqrt{s} = 13$ TeV collected with the CMS experiment during Run-2.

T 9.4 Mon 17:00 T-H22

Search for Elektroweak Production of Sleptons in Di-Lepton Final States with the ATLAS Detector — •MARIAN RENDEL, MICHAEL HOLZBOCK, HUBERT KROHA, and SANDRA KORTNER — Max-Planck-Institut für Physik (Werner-Heisenberg-Institut)

Supersymmetry (SUSY) is one of the best studied extensions of the Standard Model (SM) and as such a major part of the ATLAS physics program. SUSY scenarios with light scalar partners of the SM leptons (sleptons) and the neutralino as lightest SUSY particle (LSP) may address the observed muon g-2 anomaly, as well as provide a viable candidate for Dark Matter, and are thus of particular interest. The kinematic properties of final states in events with pair-produced sleptons primarily depend on the mass splitting (Δm) between the sleptons and the LSP. Two complementary ATLAS searches targeting large and small mass splittings, respectively, have been published. However, there remains a sensitivity gap in the intermediate region of moderate mass splittings between 20 GeV

In this talk, a new search for sleptons in events with an electron or muon pair produced with an initial-state radiation jet, is presented which particularly aims at closing the sensitivity gap.

The design of the signal region and the background estimation strategies are introduced and the expected sensitivity using 139 $\rm fb^{-1}$ LHC proton–proton col-

Monday

Location: T-H22

lision data collected by the ATLAS experiment during the years of 2015 and 2018 is presented.

T 9.5 Mon 17:15 T-H22

Search for supersymmetry in single lepton events with the full Run 2 data — •Frederic Engelke^{4,5}, Kerstin Borras^{4,5}, Kimmo Kallonen³, Henning Kirschenmann³, Pantelis Kontaxis¹, Dirk Krücker⁴, Isabell Melzer-Pellmann⁴, Ashraf Mohammed^{4,5}, Paris Sphicas^{1,2}, Costas Vellidis¹, and Lucas Wiens⁴ — ¹University of Athens — ²CERN — ³Helsinki Institute of Physics — ⁴DESY — ⁵RWTH Aachen IIIA

Results are presented from a search for supersymmetry in events with a single electron or muon, and multiple hadronic jets. The data corresponds to a sample of proton-proton collisions at $\sqrt{s} = 13$ TeV with an integrated luminosity of 138 fb⁻¹, recorded by the CMS experiment at the LHC.

We use the angular correlation between the lepton and the W boson's transverse momenta for a strong separation between the signal and the background region. The investigation of the two different signal models benefits from improved top and W tagging methods.

The search targets gluino pair production, where the gluinos decay into the lightest supersymmetric particle (LSP) and either a top quark-antiquark pair or a pair of light quarks in the final state.

T 9.6 Mon 17:30 T-H22

Reconstruction of the displaced τ for the long-lived τ slepton searches at CMS — •MYKYTA SHCHEDROLOSIEV¹, KONSTANTIN ANDROSOV^{2,3}, ANDREA CARDINI¹, DIRK KRÜCKER¹, and ISABELL MELZER-PELLMANN¹ — ¹DESY, Hamburg, Germany — ²École polytechnique fédérale de Lausanne — ³ETH Zurich Supersymmetric scenarios with long-lived tau sleptons are well motivated, e.g. within gauge mediated symmetry breaking scenarios. Direct searches of $\tilde{\tau} \rightarrow \tau \tilde{\chi}_0^1$ are limited by the reconstruction efficiency of displaced tau leptons at CMS that are produced up to 50 cm away from the IP. In our study, we explore the optimization of the displaced τ lepton reconstruction using deep neural networks for the corresponding stau searches.

T 9.7 Mon 17:45 T-H22

Constraints on Supersymmetry from Collider Searches and Other Experiments — SAMUEL BEIN, •MALTE MROWIETZ, and PETER SCHLEPER — Universität Hamburg, Institut für Experimentalphysik

Constraints from 13 TeV LHC searches for supersymmetry and other experiments on the minimal supersymmetric standard model (MSSM) are evaluated in the context of the 19-parameter phenomenological MSSM (pMSSM). Complementarity and possible tension between the LHC data, the recent g-2 result, and direct detection experiments are examined.

T 9.8 Mon 18:00 T-H22

Study of the chargino discovery at FCC-hh and inner tracker proposition — •CÉDRINE HÜGLI — DESY, Zeuthen, Germany

The Standard Model has been shown to be an incomplete theory. The most popular theory beyond the SM is Supersymmetry (SUSY). Up to now, no evidence for any of the SUSY particles was found, it was only possible to set exclusion regions and limits on their properties. The post LHC era Future Circular Collider (FCC) with a center of mass energy of 100 TeV could be a new possibility to find SUSY particles. In this work the discovery potential of the pure wino (higgsino) with mass 3 TeV (1 TeV) and average lifetime 0.2 ns (0.023 ns) at FCC-hh is studied in the context of the Minimal Supersymmetric Standard Model with Anomaly Mediated Supersymmetry breaking (tan(β) = 5, m_0 = 20 TeV and $sign(\mu) > 0$). Using simulations of minimum bias events and single chargino interactions in the reference design of the tracker, the fake rate is estimated. Additionally, the influence of the tracker performance on the discovery potential of the wino (higgsino) at FCC-hh. The possible mass reach for charginos at FCC-hh is found and the mass prediction possibility is studied.

T 9.9 Mon 18:15 T-H22

SUSY at future colliders - an overview — •MIKAEL BERGGREN — Deutsches Elektronen-Synchrotron DESY, Notkestr. 85, 22607Hamburg

A study of the prospects for discovering or excluding SUSY at various proposed future colliders is presented. The study is based on scanning the relevant parameter space of (weak-scale) SUSY parameters. In particular, I concentrate on the properties most relevant to evaluate the experimental prospects: mass differences, lifetimes and decay-modes. The observations are then confronted with estimated experimental capabilities, including - importantly - the detail of simulation these estimates are based upon. Conclusions on realistic prospects are presented.

T 10: Search for New Particles -1

Time: Monday 16:15-18:30

T 10.1 Mon 16:15 T-H23

 $\begin{array}{l} \textbf{Combination of diboson resonance searches with CMS Run2 data -- \bullet ENRICO} \\ \textbf{Wendrich}^1, \textbf{Anna Albrecht}^1, \textbf{Irene Zoi}^2, \textbf{Ankita Mehta}^1, and \textbf{Andreas} \\ \textbf{Hinzmann}^1 -- ^1 \textbf{Universität Hamburg, Institut für Experimental physik, Luruper} \\ \textbf{Chaussee 149, Hamburg, Germany -- }^2 Fermilab, Batavia IL, USA \end{array}$

In many extensions of the standard model, we have new physics resonances decaying to boson pairs. This talk focuses on new resonances decaying into a pair of W, Z and H bosons. Searches in two channels (fully hadronic VV and semileptonic VW) are statistically combined to receive a higher sensitivity for the upper cross section limits and the expected sensitivity is presented. The searches are based on CMS proton-proton collision data at 13 TeV, corresponding to an integrated luminosity of 138/fb.

T 10.2 Mon 16:30 T-H23

Search for high-mass dilepton resonances in association with b-jets with the ATLAS detector at \sqrt{s} =13TeV — VOLKER AUSTRUP, FRANK ELLINGHAUS, JENS ROGGEL, and •MAREN STRATMANN — Bergische Universität Wuppertal

Potential deviations from the Standard Model predictions observed in decay processes involving a $b \rightarrow s$ quark transition hint at Beyond Standard Model physics. One possible explanation for the deviations is the existence of a new heavy vector boson, the Z', which couples only to quarks of the second and third generation. In this talk, the status of a search for a lefthanded Z' produced in association with jets originating from b or s quarks is presented. The search is carried out in the dileptonic Z' decay channel. An overview of the analysis strategy, including two different approaches for defining a signal region, is given and a preliminary limit on the Z' production cross section is set.

T 10.3 Mon 16:45 T-H23

Search for heavy resonances in ATLAS experiment in events with four top final state — •ELIZAVETA SITNIKOVA, KRISZTIAN PETERS, ALICIA WONGEL, and PHILIPP GADOW — DESY, Hamburg, Germany

A search for heavy resonances in the four-top-quark final state is presented. This search provides a unique way to probe new physics, such as top-philic resonances (Z') produced in association with top quarks $(t\bar{t}Z')$ resulting in events containing four top-quarks. The study was performed using data of proton-proton collisions with center-of-mass energy of 13 TeV collected by the ATLAS experiment at the LHC in 2015-2018 with total luminosity of 139 fb⁻¹. Selected events contain a single lepton in association with multiple jets and are categorized into signal regions according to the multiplicity of jets and how likely these contain *b*-hadrons. In this presentation we will outline the analysis strategy and discuss the results which are interpreted in a model independent way, as well as in terms of a simplified model for top-philic resonances.

T 10.4 Mon 17:00 T-H23

Search for heavy resonances decaying to top quark pairs at CMS — •KSENIA DE LEO¹, JOHANNES HALLER¹, ROMAN KOGLER², ARTUR LOBANOV¹, and MATTHIAS SCHRÖDER¹ — ¹Institut für Experimentalphysik, Universität Hamburg — ²DESY, Hamburg

A search for new heavy resonances decaying to top quark pairs is presented. The analysis uses pp-collision data with a center-of-mass energy of 13 TeV, collected with the CMS experiment during Run-2 of the LHC. The data correspond to an integrated luminosity of 137 fb^{-1} .

The search is performed in final states with one lepton, missing transverse energy and jets, and exploits top-tagging techniques to identify the hadronic decay of top quarks. A multi-class neural network has been developed to categorise the events into the main backgrounds from known standard model processes. Exclusion limits on the production cross section of new particles are set for different benchmark models. While the main target of the analysis is the highest possible mass region, improvements have been implemented to increase the sensitivity for masses below 1 TeV, important for models where a scalar or pseudo-scalar particle decays to top quark pairs.

T 10.5 Mon 17:15 T-H23

A search for pair production of excited top quarks t* at CMS — •FINN LABE¹, JOHANNES HALLER¹, ROMAN KOGLER², ARTUR LOBANOV¹, and MATTHIAS SCHRÖDER¹ — ¹Institut für Experimentalphysik, Universität Hamburg — ²DESY

A search for pair production of excited top quarks t^* in the decay channel $t^*t^* \rightarrow tgtg$ is presented. The search uses proton-proton collision data at a center-of-mass energy of 13 TeV, collected with the CMS experiment during Run 2 of the LHC, corresponding to a total integrated luminosity of 137 fb⁻¹. The analysis is performed in the lepton+jets final state, where the fully hadronic decay of the highly boosted top quark is identified with top tagging techniques. For discrimination of signal from background, a deep neural network is used, which is decorrelated from the sum of all object momenta S_T using a "designing decorrelated taggers" (DDT) approach. Distributions of signal and background

Monday

events of S_T are used to evaluate the expected sensitivity of the search, which yields promising results over the full mass range analyzed.

T 10.6 Mon 17:30 T-H23

Categorizing final states and deriving limits on ALPs coupling to the SM Higgs boson in multiphoton events with ATLAS — PETER KRÄMER, KRISTOF SCMIEDEN, MATTHIAS SCHOTT, and •OLIVERA VUJINOVIĆ — Johannes Gutenberg University, Mainz, Germany

Some puzzling questions in particle physics, such as the strong CP problem or the discrepancy of the muon magnetic moment could be solved by introducing light scalar or pseudo-scalar axion-like particles (ALPs). Theoretical models allow a wide range of ALP-masses and couplings to SM particles such as photons and the Higgs boson. Therefore, parts of the ALP parameter space could be investigated with collider experiments like the ATLAS experiment at the LHC.

In the ongoing analysis, we search for the SM Higgs boson decaying into a pair of ALPs further decaying into two photons each. Depending on ALP properties such as mass and their coupling to photons, the signal is expected to form different final states, ranging from 2 to 4 photons. Each final state requires a dedicated approach to deriving the desired limits. In this talk it will be discussed which final states are expected, how they can be accessed and categorized.

T 10.7 Mon 17:45 T-H23

Multivariate photon classification for an analysis aiming to derive limits on ALPs coupling to the SM Higgs boson in multiphoton events — •PETER KRÄMER, KRISTOF SCHMIEDEN, MATTHIAS SCHOTT, and OLIVERA VUJINOVIC — Johannes Gutenberg Universität Mainz

Some puzzling questions in particle physics, such as the strong CP problem or the discrepancy of the muons magnetic moment could be solved by introducing light scalar or pseudo-scalar axion like particles (ALPs). Theoretic models allow a wide range of ALP-masses and couplings to SM particles such as the photon and the Higgs boson. Therefore, parts of the ALPs parameter space could be investigated with collider experiments like the ATLAS experiment at the LHC.

In the present analysis we search for SM Higgs bosons decaying to a pair of ALPs further decaying to two photons each. Depending on the mass of the ALP, the kinematic of the decay photons differ a lot. Especially for low ALP masses, the photons experience a Lorentz-boost which results in highly collinear photons. These collinear photons are not recognized by standart ATLAS photon reconstruction algorithms.

In this talk I want to discuss how events with highly collimated photons can be distinguished from $H \rightarrow \gamma \gamma$ events using multivariate techniques.

T 10.8 Mon 18:00 T-H23

Search for Axion-Like Particles in Non-Resonant tī Production Using Lepton+Jets Final States with the CMS Detector — •HENRIK JABUSCH¹, KSENIA DE LEO¹, JOHANNES HALLER¹, ROMAN KOGLER², ARTUR LOBANOV¹, and MATTHIAS SCHRÖDER¹ — ¹Institut für Experimentalphysik, Universität Hamburg — ²DESY, Hamburg

We search for signs of axion-like particles (ALPs) in top quark pair production at the LHC. In high-energy proton-proton collisions, ALPs could arise as off-shell mediators. Employing a model-independent effective field theory approach with ALP couplings to gluons and top quarks, ALPs lead to non-resonant signatures modifying the shape of the invariant mass distribution of the t \bar{t} system.

Our search targets these signatures using 137 fb^{-1} of pp collision data at $\sqrt{s} = 13 \text{ TeV}$ recorded with the CMS detector. Focusing on lepton+jets final states, we utilize a deep neural network for event classification and constrain the ALP-top quark coupling for the first time.

T 10.9 Mon 18:15 T-H23

Searching for ALPs in Light-by-light Scattering in pp Collisions Using AFP Proton Tagging with the ATLAS Detector — Peter Bussey¹, Tomas Chobola², Petr Dostal², •Hussain Kitagawa³, Patrick Odagiu⁴, Rafal Staszewski⁵, André Sopczak², Gen Tateno⁶, Junichi Tanaka⁶, Marek Tasevsky⁷, and Koji Terashi⁶ — ¹University of Glasgow — ²CTU in Prague — ³Okayama University — ⁴EPFL Lausanne — ⁵IFJ PAN Krakow — ⁶ICEPP Tokyo — ⁷CAS Prague

The search for an Axion-Like-Particle (ALP) is being performed using about 20 fb^{-1} data recorded with the ATLAS experiment and the ATLAS Forward Proton (AFP) detector in 2017. The two components of the AFP detector are positioned symmetrically at approximately 220 m on either side of the interaction point near the beam pipe and are used to measure the kinematics of surviving protons. The high-mass diphoton spectrum is studied for the search for an ALP mediated by light-by-light scattering. The investigated mass range is between 0.1 TeV and 2 TeV ALP with a typical coupling $g = 1 \text{ TeV}^{-1}$. A blinding strategy is established, and an optimization of the acoplanarity angle is performed. A Functional Decomposition study for the background modeling is applied.

T 11: Gaseous Detectors

Time: Monday 16:15-18:15

T 11.1 Mon 16:15 T-H24

Photon Detection by Structured Converter Layers in Micro-Pattern Gaseous Detectors — •Katrin Penski, Otmar Biebel, Stefanie Götz, Vitaliy HAVRYLENKO, RALF HERTENBERGER, CHRISTOPH JAGFELD, MAXIMILIAN RIN-NAGEL, CHRYSOSTOMOS VALDERANIS, and FABIAN VOGEL — LMU München Micro-Pattern Gaseous Detectors are high-rate capable with excellent spatial and temporal resolution. Developed for the detection of charged particles, the low density in the active gas volume of these detectors exhibit only a poor detection efficiency for electrically neutral particles. For photons the detection via the photoelectric effect can be increased using a solid converter cathode, which is made of high-Z materials. With our novel approach the detection efficiency can be optimized by incorporating several converter plates, which are mounted parallel to the electric drift field in the detector. With an optimized electric field, the created electrons are guided out of the conversion volume. First measurement results are presented and compared to corresponding simulations. This technique allows for higher photon detection efficiencies and more sensitive equipment, which might be applied to modern astrophysics, material research or medical physics.

T 11.2 Mon 16:30 T-H24

Characterisation of a neutron source using MICROMEGAS detectors — •Stefanie Götz, Otmar Biebel, Vitalii Havrylenko, Ralf Hertenberger, Christoph Jagfeld, Katrin Penski, Maximilian Rinnagel, Chrysostomos Valderanis, and Fabian Vogel — LMU München

MICRO MEsh GAseous Structure detectors (Micromegas) are high-rate capable micro-structured gaseous detectors with high spatial resolution due to small-scale readout strip pitch. This study uses the Micromegas detector technology to characterise the intensity profile and the interaction probability of the radiation composition of a 10 GBq Am-Be neutron source. The detector response from the neutron source is evaluated for different source positions relatively to the detector. Pieces of plastic and lead for radiation shielding allow differentiating between neutrons and gamma radiation when placed on the detector, thus disentangling the detector response of photons and neutrons. The interaction rate of the MeV neutrons and gammas is determined using random triggers. From the source characteristics, an interaction rate of 12 MHz is expected in a $40 \times 150 \text{ cm}^2$ Micromegas detector with 1000 strips which corresponds to 6 accidental hits per trigger. The goal is to resolve the beam profile of the neutron source in both horizontal and lateral direction.

T 11.3 Mon 16:45 T-H24

Particle position reconstruction using a segmented GEM foil in a microstructured gaseous detector — •Christoph Jagfeld, Otmar Biebel, Ste-FANIE GÖTZ, VITALII HAVRYLENKO, RALF HERTENBERGER, KATRIN PENSKI, MAXIMILIAN RINNAGEL, CHRYSOSTOMOS VALDERANIS, and FABIAN VOGEL — LMU, München

In Micromegas (Micro-MEsh GAseous Structures) detectors, a modern form of micro-pattern gaseous detectors, the signal is usually read out via readout strips on the anode. The signal created at the mesh is usually neglected for the particle position reconstruction. By replacing the mesh with a GEM (Gas Electron Multiplier) foil, which is segmented into 0.5 mm wide readout strips on one side, the particle position can be determined from the GEM foil signal as well. If the strips on the GEM foil are orientated perpendicular to the anode readout strips, a particle position can be reconstructed in two spatial coordinates without adding a second layer of readout strips on the anode. CERN test beam measurement results are presented. These show a good good particle reconstruction efficiency, spatial and angular resolution of this GEM-Micromegas concept.

T 11.4 Mon 17:00 T-H24

Research and Development of an Electret based Gasseous Detector — •VITALII HAVRYLENKO, OTMAR BIEBEL, STEFANIE GÖTZ, RALF HERTENBERGER, CHRISTOPH JAGFELD, KATRIN PENSKI, MAXIMILIAN RINNAGEL, CHRYSOSTOMOS VALDERANIS, and FABIAN VOGEL — LMU, München

Gaseous Electron Multiplier (GEM) detectors require a high electric field for the signal amplification. In order to simplify the operation of such a detector system, e.g. for use on X-ray astronomy satellites the GEM foil will be replaced by an electret GEM foil. An electret is a material that conserves an internal polarisation, thus creating an electrostatic potential on a long time scale. This detector does not need an external HV supply. Results of different techniques of electret production are presented. Different bipolar epoxies were hardened in strong electric fields yielding a static polarization which can be observed by the induced voltage. Furthermore, Teflon is charged via corona-discharges. The voltage dependence on time is measured for all electret samples and compared. Potential up to 1000 V stable over multiple days are reported.

Location: T-H24

T 11.5 Mon 17:15 T-H24

Prototype of a Cherenkov position sensitive Micromegas — •MAXIMILIAN RINNAGEL, OTMAR BIEBEL, STEFANIE GOETZ, CHRISTOPH JAGFELD, KATRIN PENSKI, CHRYSOSTOMOS VALDERANIS, FABIAN VOGEL, and RALF HERTENBERGER — LMU München

etectors utilizing the Cherenkov effect are well established for particle identification of charged particles in detector systems such as LHCb or HADES. In reverse it is possible to determine the momentum of a known particle by measuring the opening angle of the Cherenkov cone in Cherenkov media. Our goal with this 10x10 cm² prototype is a proof of principle using cosmic muons. A traversing muon creates around 1500 Cherenkov photons in our 19 mm thick ultra-violet transparent Lithium Fluoride (LiF) crystal (diameter 50 mm; UV optical refractive index 1.5). The conversion to electrons happens in transmission in a photosensitive CsI layer evaporated onto a 5 nm Cr layer, both applied to the bottom of the radiator. High voltage of -300 V, at the Cr layer, guides the ionization and photoelectrons into the drift region of a Micromegas gaseous micro pattern detector. First results utilizing muon tracks reconstructed from reference detectors will be shown. The typical signal shape of this detector as well as the spatial position reconstruction are compared to the Micromegas detector without the LiF Cherenkov radiator.

T 11.6 Mon 17:30 T-H24

Test of ATLAS Micromegas detectors with a ternary gas mixture at the CERN GIF++ facility — •Fabian Vogel, Otmar Biebel, Stefanie Götz, Vitaliy Havrylenko, Ralf Hertenberger, Christoph Jagfeld, Katrin Penski, MAXIMILIAN RINNAGEL, and CHRYSOSTOMOS VALDERANIS - LMU München The ATLAS collaboration at LHC has chosen the resistive Micromegas technology, along with the small-strip Thin Gap Chambers (sTGC), for the high luminosity upgrade of the first muon station in the high-rapidity region, the New Small Wheel (NSW) project. Four different sizes of Micromegas quadruplets have been constructed at four construction sites in Italy (SM1), Germany (SM2), France (LM1) and CERN/Greece/Russia (LM2). Achieving the requirements for these detectors revealed to be even more challenging than expected. One of the main features being studied is the HV stability of the detectors. Several approaches have been tested in order to enhance the stability, among them the use of different gas mixtures. A ternary $\mbox{Argon-CO}_2\mbox{-i}\mbox{C}_4\mbox{H}_{10}$ mixture has shown to be effective in dumping discharges and dark currents. It allows the operation of the Micromegas detectors at safe working points with high cosmic muon detection efficiency. The presence of Isobutane in the mixture required a set of aging studies, ongoing at the GIF++ radiation facility at CERN, where the expected HL LHC background rate is created by a ¹³⁷Cs 14 TBq source of 662 keV photons. Preliminary aging results and effectiveness of the ternary mixture will be shown.

T 11.7 Mon 17:45 T-H24

The Influence of Oxygen Defects at Measurements with a MicroMegas Detector Filled with an Ar-CO₂ Gas Mixture — •Burkhard Böhm¹, Deb Sankar Bhattacharya^{1,2}, Thorben Swirski¹, and Raimund Ströhmer¹ — ¹Universität Würzburg — ²INFN Trieste

In particle physics, Micro-Pattern Gaseous Detectors (MPGD) find high usage in different experiments like ATLAS, CMS or ALICE. In this study MicroMegas (MM) - a special type of MPGDs - are researched in terms of O_2 contamination. They are well known for their simple single-stage amplification, high and stable gain and excellent spatial and temporal resolutions. These detectors can be contaminated by O_2 from the air and due to the electronegativity of O_2 , electrons in the gaseous detector can be captured. Hence, even a low concentration of O_2 has an impact on the detector performance. By precisely controlled inflowing of O_2 inside a resistive MicroMegas chamber, the effect on the gas-gain,mainly due to attachment in the drift region, and the amplification of the number of primary electrons are studied. In parallel to the experimental study numerical investigations were done to estimate the amplification gap from the comparison of the simulated to the measured signal.

T 11.8 Mon 18:00 T-H24

A GridPix detector for IAXO — •TOBIAS SCHIFFER, KLAUS DESCH, JOCHEN KAMINSKI, SEBASTIAN SCHMIDT, and MARKUS GRUBER — Physikalisches Institut der Universität Bonn

In the scope of the search for axions with helioscopes, like the International Axion Observatory (IAXO) and its precursor BabyIAXO, detectors capable of measuring low energy X-rays down to the 200 eV range are necessary. For this purpose the GridPix detector, which was already successfully used at CAST, is an appropriate and constantly evolving solution.

The GridPix is a MicroMegas like readout consisting of a pixelized readout ASIC (Timepix/Timepix3) with a perfectly aligned gas amplification stage on top. Due to the very high granularity this detector is capable of detecting single electrons allowing the measurement of low energy X-rays. To convert these X-

rays into electrons a gas volume is built above the readout sealed with a vacuum tight X-ray entrance window.

For the goals of IAXO and BabyIAXO a very low detector background needs to be achieved, therefore only a few radiopure materials are contemplable. Also a good offline separation of signal and background events is to be achieved, here

T 12: Pixel Detectors

Time: Monday 16:15-18:30

T 12.1 Mon 16:15 T-H25 ATLAS ITk Pixel Detector Quad Module Building and Testing for the HL-

LHC Upgrade — JÖRN GROSSE-KNETTER, ARNULF QUADT, • YUSONG TIAN, and HUA YE — II. Physikalisches Institut, Georg-August-Universität Göttingen In the ATLAS detector upgrade for the High-Luminosity LHC (HL-LHC), the current Inner Detector will be upgraded to an all-silicon Inner Tracker (ITk), to operate under higher occupancy and radiation damage resulting from the higher luminosity. The pixel detector is the inner-most layer of the ITk, shaped like a cylinder with about 0.4m radius and 6m long. It is assembled with barrels and end-cap rings made of single and quad-chip modules. Each module is assembled with a bare module and a printed circuit board (PCB). A quad bare module consists of one sensor bump-bonded to 4 front-end readout chips. RD53A is the name for the front-end readout chip prototype, the name for the final version is ITkPix. The modules are referred to with the name of the front-ends. This talk shows the module assembly and testing results.

T 12.2 Mon 16:30 T-H25

ITk-pixel outer barrel demonstrator — JÖRN GROSSE-KNETTER¹, SUSANNE Kühn², •Silke Möbius¹, Arnulf Quadt¹, Benedikt Vormwald², and Hua $Ye^1 - {}^1II.$ Physikalisches Institut, Georg-August-Universität Göttingen -²CERN

For the upgrade of the LHC to the High-Luminosity-LHC, the ATLAS tracking detector will be replaced with a pure silicon detector, the Inner Tracker (ITk), as the higher luminosity requires radiation hard components that can deal with higher occupancies and radiation. Given the close proximity to the interaction point, the environment is especially challenging for the pixel detector which will comprise quad chip modules for outer barrel layers.

In order to characterize and test ITk-pixel prototype quad modules, up to 40 modules are built and tested for the outer barrel demonstrator. This so-called demonstrator is a larger structure which features some of the final mechanics to test the system functionalities. Systems like an interlock need to be implemented and tested and the modules behaviour before and after being powered in serial needs to be examined to allow for thorough tests.

After having done initial tests with a preliminary version of the demonstrator with older modules, the preparation of the set-up and commissioning of the new demonstrator is currently ongoing.

This talk will give an overview over the current state of the demonstrator, explain the set-up and summarize the potential tests to be performed on the demonstrator.

T 12.3 Mon 16:45 T-H25

Testing of loaded local supports of the ATLAS Inner Tracker - • WAEL ALKAкні, JOERN GROSSE-KNETTER, and ARNULF QUADT — II. Physikalisches Institut, Georg-August-Universität Göttingen

During the third long shut down (LS3) of the Large Hadron Collider (LHC), a high luminosity upgrade takes place in which the instantaneous luminosity reaches $7.5 \times 10^{34} \text{ cm}^{-2} \text{s}^{-1}$ in Run 4. The current inner detector of the ATLAS experiment is going to be replaced by inner tracker (ITk). ITk is composed of a pixel detector surrounded by a strip detector. The pixel detector consists of 5 layers where the first 2 layers represent the inner pixel detector and the last 3 layers represent the outer barrel and the endcap. The outer barrel consists of 4474 modules. These modules are mounted on loaded local supports of which there are two types : loaded longeron and loaded inclined half ring. Due to the large number of elements used during the assembly and production of the ITk, the ITk production database is used to store and track the information of the ITk components. Moreover a graphical user interface (GUI) is used to ease working with the ITk production database during the production of the ITk outer barrel.

This talk introduces the integration procedure of modules to local supports. Moreover it demonstrates my contribution to the implementation of the loaded local supports in the ITk production database and the GUI developments.

T 12.4 Mon 17:00 T-H25

ATLAS ITKPix Waferprobing — • MARK STANDKE, YANNICK DIETER, TOMASZ HEMPEREK, FLORIAN HINTERKEUSER, FABIAN HÜGGING, HANS KRÜGER, DAVID-LEON POHL, MARCO VOGT, and JOCHEN DINGFELDER - Forschungsund Technologie-Zentrum Detektorphysik - Kreuzbergweg 26, 53115 Bonn Wafer-probing is a process, in which each individual chip is tested for its key function parameters on wafer level. For this purpose, Bonn has developed a the insights from the CAST detector are used. Further, to get more signal the X-ray entrance window needs to be as transparent as possible for the low energy X-rays. This is achieved with an ultra thin (<200 nm) silicon nitride membrane. The challenges of the design process and some first results of the detector will

be presented.

fast and versatile testing and analysis environment, making large-scale testing

Location: T-H25

for ATLAS-ITkPix possible. ITkPix is the first full-scale 65 nm ATLAS - hybrid pixel-detector readout chip, developed by the RD53 collaboration. ITkPix consists of more than one billion transistors with high triplication ratio in order to cope with high particle and therefore radiation densities at the heart of ATLAS. The chips will be located as close as possible to the interaction point to optimize impact parameter resolution. ITkPix features a single low power, low noise analog front-end to ensure high readout speeds and low detection thresholds. A failure of such chips at the heart of ATLAS is assumed to be hard to correct. Therefore, thorough testing is necessary. This talk will give an overview over the testing environment, while summarizing the latest results and performance of ATLAS's future inner tracker performance driver, ITkPix.

T 12.5 Mon 17:15 T-H25

Bump bond stress tests with ITk-Pixel-style daisy-chain modules through thermal cycling - JÖRN GROSSE-KNETTER, •STEFFEN KORN, and ARNULF QUADT — II. Physikalisches Institut, Georg-August-Universität Göttingen

For the upgrade of the LHC to the HL-LHC, the Inner Detector will be replaced by the fully silicon-based Inner Tracker Detector (ITk). The pixel detector of the ITk uses hybrid modules where sensor and readout chips are connected by bump bonds. Early ITk module prototypes highlighted these bump bond connections as a possible point of failure in future ITk Pixel modules when exposed to thermally induced stress. In order to investigate this issue, daisy chain modules with realistic bump bond pitch were tested before and during exposure to thermal stress through cycling in a thermal shock chamber using a dedicated in-situ method in Goettingen. The results of these tests using different modules with different assembly options are presented in this talk.

T 12.6 Mon 17:30 T-H25

Pixel front-end masking and DQ monitoring - MARCELLO BINDI, JÖRN GROSSE-KNETTER, •ANDREAS KIRCHHOFF, and ARNULF QUADT - II. Physikalisches Institut, Georg-August-Universität Göttingen, Göttingen, Germanv

During Run 2 of the LHC, the ATLAS tracking software masked non-working pixel modules for offline reconstruction. The masking itself is applied if a module does not receive hits. To improve track reconstruction the number of holes should be reduced. Holes are defined as intersections of reconstructed tracks with sensitive detector elements that did not result in a hit. They are estimated by comparing the hits-on-track with the intersected modules. Inactive modules are excluded from the hole definition. As a consequence, a masked module is treated in the track reconstruction as if it received always a hit.

The ATLAS tracking group for Run 3 will increase the granularity of the masking (moving from module to single front-ends) in order to reduce the number of pixel holes and increase the tracking efficiency. Hit maps collected at the end of each run are chosen as input for the front-end masking. Due to the detector geometry the masking for the different front ends needs to be optimised individually. This talk will present the corresponding studies.

The second topic of the talk will be data quality monitoring. While taking data it is extremely crucial to check whether the detector performs as expected to avoid loss in data. Therefore different monitoring software exists. This talk will present the updated software GNAM2 and its current status.

T 12.7 Mon 17:45 T-H25

Data quality monitoring for the ATLAS pixel detector using large radius tracks — CARMEN DIEZ PARDOS, IVOR FLECK, •JAN JOACHIM HAHN, and ISKAN-DER IBRAGIMOV — Center for Particle Physics Siegen, Experimentelle Teilchenphysik, Universität Siegen

The ATLAS detector is a general-purpose detector at the Large Hadron Collider (LHC) at CERN. Its Inner Detector (ID), used for reconstruction of tracks of charged particles, consists of four barrel layers and two end-caps of three discs each of the silicon pixel detector, surrounded by the silicon strip detector and the transition radiation tracker. During operation of ATLAS, a small fraction of recorded collision events is reconstructed in real time, providing prompt information on the quality of data being taken. For the upcoming LHC Run 3, the reconstruction of tracks in the ID will include the reconstruction of large radius tracks (LRT). These tracks are of interest for the search of long-lived particles beyond the Standard Model. With respect to the standard tracks, they are reconstructed with relaxed constraints on their starting position. The information

T 12.9 Mon 18:15 T-H25

from LRTs can be used to create additional histograms for monitoring detector performance during operation of ATLAS. This talk will discuss how this can improve data quality monitoring of the pixel detector.

T 12.8 Mon 18:00 T-H25

Measurement of position resolution of small pixel sensors — •Amala Augusthy¹, Daniel Pitzl², Erika Garutti¹, and Jörn Schwandt¹ — ¹Institute für Experimentalphysik, Universität Hamburg, Luruper Chaussee 149, 22761 Hamburg, Deutschland — ²Deutsches Elektronen-Synchrotron, Notkestraße 85, 22607 Hamburg, Deutschland

In order to cope with the high radiation environment in the phase 3 upgrade of CMS and ATLAS, pixel detectors with spatial resolution below $1\mu m$ will become necessary. The resolution of pixel sensors can be improved by reducing their pitch. However, the pitch of pixel sensors is limited by the read-out electronics.

To study the intrinsic resolution capabilities, non-irradiated pixels sensors with sizes of $17 \times 150 \mu m^2$, $25 \times 100 \mu m^2$ and $50 \times 50 \mu m^2$ were tested with 5.2 GeV electron beam at the DESY test beam facility. The $25 \times 100 \mu m^2$ and $50 \times 50 \mu m^2$ sensors have already been characterized and are being considered for the phase 2 upgrade of the CMS experiment. The aim of this study is to reduce the pitch even further to $17 \mu m$ and investigate its performance.

These sensors have a thickness of $285\mu m$ and are bump bonded to a low noise read-out chip ROC4SENS. The spatial resolution as a function of angle of incidence of the particle is extracted from these measurements. In this talk, the results of these measurements will be presented.

Characterisation of planar pixel sensor for the CMS phase-2 upgrade – •MOHAMMADTAGHI HAJHEIDARI¹, MASSIMILIANO ANTONELLO¹, FINN FEINDT², ERIKA GARUTTI¹, DANIEL PITZL², JÖRN SCHWANDT¹, GEORG STEINBRÜCK¹, ANNIKA VAUTH¹, and IRENE ZOI³ – ¹Universität Hamburg, Luruper Chaussee 149, 22761 Hamburg, Deutschland – ²Deutsches Elektronen-Synchrotron, Notkestraße 85, 22607 Hamburg, Deutschland – ³Fermilab, Batavia, IL, 60510, USA

For the High Luminosity phase of the Large Hadron Collider (HL-LHC), the CMS pixel detector is expected to collect up to 1-MeV neutron equivalent fluence, Φ_{eq} , of 2.3 × 10¹⁶ cm⁻² for integrated luminosity of 3000 fb⁻¹. The pixel detector will be upgraded to withstand this range of fluence.

Planar n-in-p pixel sensors with an active thickness of 150 μ m and a pixel size of 25 × 100 μ m² have been produced by Hamamatsu Photonics (HPK) and FBK. The sensors were bumped bonded to the RD53A readout chip prototype. The sensor-chip modules were irradiated with 23 MeV protons to the equivalent fluence of up to 2.4 × 10¹⁶ cm⁻².

The modules were investigated in the DESY II beam test facility after irradiation. The hit efficiency and spatial resolution as a function of the incidence angle of pixel sensors were determined from these measurements. For non-irradiated modules, a single hit resolution of 2 μ m was achieved at the optimal angle. For irradiated modules with highest fluence, the single hit efficiency still reached 98% at bias voltages below 800 V.

T 13: Semiconductor Detectors:Radiation Hardness, new Materials and Concepts

Time: Monday 16:15-18:15

T 13.1 Mon 16:15 T-H26 Voltage scans on germanium detectors — •Felix Надемани for the GeDet-Collaboration — Max-Planck-Institut für Physik, München

Germanium detectors are used in fundamental research to search for neutrinoless double-beta decay or dark matter. In many of these experiments, a perfect understanding of the working principle and characteristics of these detectors is essential.

The electric field inside a germanium detector consists of two main components: one resulting from the potentials applied to the contacts to bias the detector, and one from the ionized impurities in the germanium crystal. While the overall number of ionized impurities can be estimated from Hall measurements on the surface of the detector, the spatial distribution of the impurities inside the germanium crystal is not well known. This results in large uncertainties on the resulting electric field.

One way to probe the impurity density profile is through voltage scans. In voltage scans, the pulses resulting from the same volumes of the detectors are recorded with different bias voltages applied to the detector. This way, the contribution from the bias voltage is varied and the constant contribution from the impurities can be separated and determined. This also requires reliable pulse shape simulation.

In my talk, I will present first data from voltage scans on a p-type segmented point-contact germanium detector and compare them to simulation results obtained from the open-source Julia software package *SolidStateDetectors.jl* to give an estimate of the impurity density profile.

T 13.2 Mon 16:30 T-H26

Eine neue Software zur 3D Simulation von Halbleiterdetektoren - SolidState-Detectors.jl — •MARTIN SCHUSTER für die GeDet-Kollaboration — Max-Planck Insitut für Physik, München, Deutschland

Halbleiterdetektoren, insbesondere aus Silizium und Germanium, haben schon lange einen festen Platz in zahlreichen Experimenten und Industriefeldern. In der GeDet (Germanium Detektor Entwicklung) Gruppe am Max-Planck Institut für Physik werden Germaniumdetektoren genau untersucht. Eine entscheidende Rolle spielt dabei der Vergleich von in Testständen aufgenommenen und simulierten Daten. In der Gruppe wurde eine neue "Open Source" Software in der jungen Programmiersprache Julia geschrieben, mit der das Verhalten aller auf Dioden basierenden Halbleiterdetektoren simuliert werden kann. Das Paket ermöglicht die schnelle Berechnung der elektrischen Potentiale und Felder und bietet die Möglichkeit der Pulsformsimulation basierend auf der Drift der Ladungsträger. Das Einlesen von GEANT4-generierten Ereignissen ist möglich. In diesem Vortrag werden die Funktionsweise und Möglichkeiten der Software erläutert. Als Beispiel dient ein vierfach segmentierter n-Typ Punktkontakt -Germaniumdetektor.

T 13.3 Mon 16:45 T-H26

Influence of radiation damage on the absorption of near-infrared light in silicon – •ANNIKA VAUTH, ROBERT KLANNER, and JÖRN SCHWANDT – Institut für Experimentalphysik, Universität Hamburg, Luruper Chaussee 149, Hamburg Radiation damage of silicon sensors is an important area of investitation in high energy physics today. Frequently, red and near-infrared light is used to generate electron-hole pairs to study the charge collection efficiency of radiationdamaged silicon sensors. In order to determine the absolute number of produced charge carriers, the light absorption coefficient, α , has to be known.

To study the change of α due to radiation-induced defects, we have measured the transmission of light with wavelengths between 1-2 μ m through silicon samples irradiated to 1 MeV-neutron-equivalent fluences between 0 and 1×10^{17} cm⁻².

In this contribution, the results of these measurements will be presented: the contribution of the irradiation to α was found to scale with fluence for the entire fluence range investigated. In the wavelength region around 1.8 μ m, evidence for the production of the radiation-induced divacancy defect V_{2i}^0 with a density approximately proportional to the fluence was found. A decrease of the effective band gap of silicon with irradiation fluence will be shown, up to a reduction of about 60 meV for a fluence of 1×10^{17} cm⁻².

T 13.4 Mon 17:00 T-H26

Location: T-H26

Studies of irradiated ATLASpix3.1 sensors for the LHCb MightyTracker — •JAN HAMMERICH for the LHCb MightyTracker group-Collaboration — University of Liverpool, Liverpool, United Kingdom

The Mighty Tracker is a proposed upgrade to the downstream tracking system of LHCb for operations at luminosities of up to 1.5×10^{34} cm $^{-2}$ s $^{-1}$ starting with the LHC Run 5 data taking period. It foresees the replacement of the most central area of the current scintillating fibre tracker with High Voltage CMOS (HV-CMOS) pixel sensors. HV-CMOS sensors have demonstrated a significant radiation tolerance and good performance making them an ideal choice of technology for the LHCb experiment.

Monolithic Active Pixel Sensors (MAPS) fabricated in HV-CMOS processes provide fast charge collection via drift and allow the implementation of the readout on the same die as the sensitive volume. Due to the use of commercial processes, these sensors can be fabricated at low cost as no hybridisation with bump bonds is required. Since they are not fully depleted, the inactive volume can be thinned away.

A dedicated sensor called the MightyPix is developed for this programme. It is based on the HV-MAPS families MuPix and ATLASpix and tailored to the requirements of LHCb. To demonstrate the feasibility of this technology for the LHCb environment, ATLASpix3.1 sensors have been irradiated. These sensors are studied in terms of time resolution and power dissipation in a temperature controlled environment.

T 13.5 Mon 17:15 T-H26

Full-size passive CMOS sensors for radiation tolerant hybrid pixel detectors — •Yannick Dieter¹, Michael Daas¹, Tomasz Hemperek¹, Fabian Hügging¹, Hans Krüger¹, David-Leon Pohl¹, Tianyang Wang², Norbert Wermes¹, Pascal Wolf¹, and Jochen Dingfelder¹ — ¹Physikalisches Institut der Universität Bonn — ²Zhangjiang Laboratory, China

CMOS process lines are an attractive option for the fabrication of hybrid pixel sensors for large-scale detectors like ATLAS and CMS. Besides the cost-

effectiveness and high throughput of commercial CMOS lines, multiple features like poly-silicon layers, metal-insulator-metal capacitors and several metal layers are available to enhance the sensor design.

After an extensive R&D programme with several prototype sensors in 150 nm LFoundry technology, passive CMOS pixel sensors have been manufactured for the first time as full-size sensors compatible with the RD53 readout chips designed for the ATLAS and CMS tracker upgrades.

This presentation will focus on IV-curve and hit-detection efficiency measurements and the characterization of the full-size sensors, before and after irradiation to fluences of $2\times10^{15}\,n_{eq}/cm^2$ and $5\times10^{15}\,n_{eq}/cm^2$, using a minimum ionising electron beam.

T 13.6 Mon 17:30 T-H26 Electrical characterisation of passive CMOS strip sensors — •Hannah Jansen¹, Jan-Hendrik Arling³, Marta Baselga¹, Leena Diehl², Ingrid Maria Gregor^{3,4}, Tomasz Hemperek⁴, Kevin Kröninger¹, Sven Mägdefessel², Ulrich Parzefall², Arturo Rodriguez², Surabhi Sharma³, Dennis Sperlich², and Jens Weingarten¹ — ¹TU Dortmund University — ²University of Freiburg — ³DESY Hamburg — ⁴University of Bonn

One of the major limitations of experiments in high energy physics are the production costs of the sensors used. Silicon sensors made with CMOS technology can solve this problem because they are very cost efficient and allow big production possibilities. Therefore, passive CMOS strip sensors with different implant layout types are investigated in this project to determine their suitability for high energy physics experiments as well as in medical applications. In particular, large passive strip sensors produced with CMOS technology are considered. In this talk, the current results of the electrical characterisation of the passive CMOS strip sensors will be presented.

T 13.7 Mon 17:45 T-H26 **Total Ionizing Dose effects on CMOS Image Sensor of the ULTRASAT space mission** — •VLAD DUMITRU BERLEA for the ULTRASAT-Collaboration — DESY Zeuthen Platanenallee 6, 15738 ULTRASAT (ULtraviolet TRansient Astronomy SATellite) is a wide-angle space telescope that will perform deep time-resolved surveys in the near ultraviolet spectrum. ULTRASAT is led by the Weizmann Institute of Science (WIS) in Israel and the Israel Space Agency (ISA) and is planned for launch in 2024. The telescope implements a backside-illuminated, stitched pixel detector. The pixel has 4T architecture with a pitch of 9.5 μ m and is produced in 180 nm process by Tower Semiconductor.

As part of the space qualification for the sensors, radiation tests are to be performed on both test sensors provided by Tower and the final flight design of the sensor. One of the main contributions to sensor degradation due to radiation for the ULTRASAT mission is Total Ionizing Dose (TID). TID measurements on the test sensors have been performed with Co-60 gamma source at Helmholz Zentrum Berlin (HZB) and CC-60 facilities at CERN, and preliminary results are presented in this talk.

T 13.8 Mon 18:00 T-H26 CVD diamonds for the use in radiation hard particle detectors — Holger Stevens, •Patrick Hölken, and Robert Konradi — Experimentelle Physik 5, TU Dortmund

Chemical-vapor deposition (CVD) diamond sensors have gained more influence in the area of particle detection in recent years. For example, they are used in the beam conditions monitor (BCM) of the LHCb experiment at CERN. The BCM monitors the quality of the particle beam and can abort the beam in flawed conditions, such as insufficient beam focus, to protect the experiment setup from possible radiation damage. The big advantages of CVD diamonds include their radiation hardness and fast response time. Furthermore, the CVD diamonds have a high tissue equivalence, which is why they are theoretically well suited for clinical dosimetry in hospitals.

In order to further investigate the radiation characteristics of the diamonds, measurements of the time behavior are carried out. In addition, a possible spatial resolution of the diamond sensors is of interest. This talk will give an overview about the results of these studies.

T 14: DAQ and Trigger 1

Time: Monday 16:15-18:30

T 14.1 Mon 16:15 T-H27

A readout system for the LHCb Beam Condition Monitor based on off-theshelf FPGA hardware — •MARTIN BIEKER, HOLGER STEVENS, and DIRK WIED-NER — Experimentelle Physik 5, TU Dortmund

The LHCb experiment is a single-arm forward spectrometer at the LHC and focuses on measurements in the *b* and *c* quark sector. Due to its unique geometry, featuring a sensitive tracking system located as close as 3.5 mm to the LHC beams, the detector is at risk of damage from adverse beam conditions. For this reason, the particle flux is monitored near the beam pipe by 8 diamond sensors in a circular arrangement at either side of and close to the interaction point.

This so-called Beam Condition Monitor (BCM) successfully protected the LHCb detector during Run I and Run II of the LHC. In preparation for the following Run III, the BCM is overhauled as part of a comprehensive upgrade of the LHCb detector. The development of a new readout system is an important part of the BCM related improvements. This system is responsible for monitoring the BCM data and initiating a beam abort in case adverse beam conditions are detected.

This talk will give an overview of the new readout system and the accompanying firmware developments: A flexible architecture based on commercially sourced FPGA boards is presented and the integration into the existing LHCb DAQ framework is outlined.

T 14.2 Mon 16:30 T-H27

Modular and scalable Timepix3 readout system — KLAUS DESCH, •MARKUS GRUBER, TOMASZ HEMPEREK, JOCHEN KAMINSKI, and TOBIAS SCHIFFER — Physikalisches Institut, Universität Bonn, Nußallee 12, 53115 Bonn

With the highly granular pixel ASIC Timepix3 several different detectors can be built by combining it either with a bump bonded sensor, with a photolithographically postprocessed MicroMegas gas amplification stage (InGrid) or with a micro-channel plate (MCP). With these combinations quite different applications like beam telescopes, X-ray detectors for axion search and polarimetry and neutron detectors can be realised.

Based on the basil framework we are developing a Timepix3 readout system which can efficiently adapt these applications that range from single chip to multi-chip designs and from low- to high-rate applications. Furthermore we are implementing a flexible monitoring system that is needed to support a range of data sources depending on the detector. The hardware consists of our own PCB designs and a support of multiple FPGA boards including the Scalable Readout System (SRS) which offers scalability in low to medium rate multi-chip applications.

In this talk I will present the readout and control system and how it scales for the applications. Furthermore, I will show how the modular approach enables several different detector designs and offers the needed functionality like calibration, equalisation, readout and monitoring.

T 14.3 Mon 16:45 T-H27

Location: T-H27

Upgrade of a cosmic muon teststand for the CMS DT-Project — •DMITRY ELISEEV, THOMAS HEBBEKER, MARKUS MERSCHMEYER, and MATEJ REPIK — Physics Institute III A, RWTH Aachen University

The Drift Tube (DT) system is one of the muon detectors of the Compact Muon Solenoid (CMS). The DT system is now upgraded to provide the increased luminosity for CMS Phase II. The upgrade pertains mainly the on-chamber digital readout and the subsequent data chain, which are completely replaced to enable higher acquisition rates and better trigger flexibility. The especially important component of the data chain is the electronics for digital readout of the muon hit data. This is the newly designed On-Board Drift Tube (OBDT) electronics, located directly at the DT chambers. This electronics enables higher data acquisition rates and is radiation tolerant. After production, the OBDT electronics will require numerous verification tests. One of the verification sites will be RWTH Aachen University with its cosmic muon teststand, originally developed during CMS construction. The core part of the muon teststand is a fully functional DT chamber, which is similar to the DT chambers at CMS. For the upcoming verification tests the teststand has undergone a comprehensive upgrade. The upgrade included the complete replacement of the data system as well as the redesign and upgrade of other systems: high- and low- voltage power-supply, gas mixing system. The teststand enables in-situ tests of the newly produced OBDT electronics as well as the verification of the data integrity of the muon-data along the readout chain.

T 14.4 Mon 17:00 T-H27

Calibration of the analogue signal path of the Level-1 Calorimeter Trigger after the ATLAS Phase-I upgrade — •THOMAS JUNKERMANN — Kirchhoff-Institut für Physik, Heidelberg

The Phase-I Upgrade of the ATLAS Level-1 Calorimeter Trigger targets a finer granularity of the spatial information of energy depositions when selecting events. To process the higher amounts of data a new digital trigger is installed. Therefore new electronic components get introduced in the on-detector elec-

tronics leading up to the trigger. These components effect the old analog trigger system and re-calibration of it is needed as it will be run in Run 3 parallel to the new system. Examples for the effects of the new components on the legacy system are the introduction of time delays to certain signal parts and amplitude losses. The effects are measured and corrected for by re-adjusting the calibration of various components.

T 14.5 Mon 17:15 T-H27

Test beam studies of the new ATLAS MDT front-end electronics in the GIF++ facility at CERN — •DAVIDE CIERI¹, GREGOR EBERWEIN¹, MARKUS FRAS¹, OLIVER KORTNER¹, HUBERT KROHA¹, CHRYSOSTOMOS VALDERANIS², ELENA VOEVODINA¹, and BASTIAN WESELY¹ — ¹Max-Planck-Institut für Physik, Munich, Germany — ²Ludwig-Maximilians-Universität, Munich, Germany

The front-end readout electronics of the ATLAS Monitored Drift Tube (MDT) detector will be replaced for High-Luminosity LHC operations in order to provide a triggerless read-out required by the upgraded first-level muon trigger. The read-out chain of an MDT chamber is composed of several Amplifier/Shaper/Discriminator (ASD) ASICs, capturing the arrival signal at the wires; Time-to-Digital Converter (TDC) ASICs, performing the required time measurements on 24 channels coming from three ASD ASICs, and a Chamber Service Module (CSM), which multiplexes data from up to 18 TDCs and sends the data via two optical fibres to the MDT trigger processor for the further trigger processing and transmission to the data acquisition (DAQ) system. The frontend electronics were designed to cope with a hit rate of 600 kHz/tube, which is twice the maximum expected rate.

Prototypes of the entire MDT front-end electronics were operated on MDT chambers in a muon beam at the GIF++ facility at CERN in 2021 using two small MDT chambers. The electronics were shown to work perfectly under different levels of the photon background irradiation, covering twice the range of the expected background fluxes at the HL-LHC.

T 14.6 Mon 17:30 T-H27 Scan Automated Testing for the ATLAS Pixel Detector — MARCELLO BINDI, ARNULF QUADT, and •CHRIS SCHEULEN — II. Physikalisches Institut, Georg-August Universität Göttingen

The ATLAS Pixel detector data acquisition system (DAQ) is distributed over several different physical components, such as front-end detector modules, read-out drivers, and PCs for operating and calibrating the detector. As a result, timeconsuming manual tests are currently required to ensure the correct operation of the entire system after software or firmware changes in any component.

To simplify software validation and free up manpower, a suite of automated tests of the is being developed for deployment in the DAQ software's continuous integration system on GitLab. Fully automated testing is only possible without involvement of the detector modules, whose operation requires some degree of manual supervision. Therefore, emulated detector responses are used for tests of readout-chain components under exclusion of the detector modules themselves.

This talk will give an overview over the first version of the automated calibration testing and validation framework currently deployed for code developments on GitLab. An outlook to further developments of the testing infrastructure will be presented as well.

T 14.7 Mon 17:45 T-H27

GUI framework and configuration database for ATLAS ITk Pixel system tests — GERHARD BRANDT, MARVIN GEYIK, •JONAS SCHMEING, and WOLFGANG WAGNER — Bergische Universität Wuppertal

For the LHC Phase-2 upgrade, a new ITk Pixel detector will be installed in the ATLAS experiment. It will allow for even higher data rates and will be thoroughly

tested in the ATLAS ITk Pixel system tests. To operate these tests, a GUI and configuration system is needed. A flexible and scalable GUI framework based on distributed microservices is introduced. Each microservice consists of a frontend GUI, a Python app served by a WSGI server, and a system-level backend. The frontend GUI is a single-page application built with the React JavaScript library. It uses PatternFly, which provides many UI elements as React components. The API for RESTful HTTP communication between the frontend and the Python app is defined via an OpenAPI specification. The Python app is the central part of each microservice. It connects to the microservices backend, such as a database or various DAQ applications that provide a Python binding. With this microservice framework, it is possible to serve specialized applications for different purposes: e.g., an API to access the data acquisition software, a service for configuration of hardware components, and a database to store these configurations. To enable users to access all services from a single web page, all frontend GUIs are compiled into one chassis. The REST and Python interfaces facilitate the maintainability and long-term upgradability of the system.

T 14.8 Mon 18:00 T-H27

The First Layer of the Mu3e Data Acquisition System — •MARTIN MÜLLER for the Mu3e-Collaboration — Institute for Nuclear Physics, JGU Mainz

The Mu3e experiment will search for the charged lepton flavor violating decay of a positive Muon into two positrons and one electron. The branching ratio of this decay in the Standard Model is predicted to be in the order of 10^{-54} and therefore any observation of such a decay would be a clear sign for new physics. Observing up to 10^8 muon decays per second, the phase I Mu3e detector will produce 100 GB/s of data from monolithic pixel chips, scintillating fibres and scintillating tiles.

The trigger-less data acquisition for the detector consists of multiple layers. Layer 1 is located inside of the Mu3e magnet and is directly connected to the detector readout ASICs with 1.25 Gbit data links. It is built from 112 Frontend-Boards which include two field programmable gate arrays (FPGAs). These FP-GAs are responsible for all communication with the different detector ASICs, including data readout, decoding, sorting, synchronisation and also the configuration of the detectors. The talk will discuss the firmware developed for the FPGAs in Layer 1 of the Mu3e DAQ and the interfaces to the other DAQ layers.

T 14.9 Mon 18:15 T-H27

Data Flow in the Mu3e Filter Farm — •MARIUS KÖPPEL for the Mu3e-Collaboration — Institute for Nuclear Physics, Johannes Gutenberg University, Mainz Germany

The Mu3e experiment at the Paul Scherrer Institute searches for the decay $\mu^+ \rightarrow e^+e^+e^-$. This decay violates charged lepton flavour conservation - any observation would be a clear indication for Physics Beyond the Standard Model. The Mu3e experiment aims for an ultimate sensitivity of one in $10^{16} \mu$ decays. The first phase of the experiment, currently under construction, will reach a branching ratio sensitivity of $2 \cdot 10^{-15}$ by observing $10^8 \mu$ decays per second over a year of data taking. The highly granular detector based on thin high-voltage monolithic active pixel sensors (HV-MAPS) and scintillating timing detectors will produce about 100 GB/s of data at these particle rates.

Since the corresponding data cannot be saved to disk, a trigger-less online readout system is required which is able to sort, align and analyze the data while running. A farm with PCs equipped with powerful graphics processing units (GPUs) will perform the data reduction. The talk presents the ongoing integration of the sub detectors into the Field Programmable Gate Array (FPGA) based readout system, in particular focusing on the data flow inside the FPGAs of the filter farm. It will also show insides of the DAQ system used in the Mu3e Integration Run performed in Spring 2021.

T 15: GRID Computing

Time: Monday 16:15-18:00

T 15.1 Mon 16:15 T-H28

Rucio Datamanagement: Expanding the capabilities for dealing with potentially problematic files — •CHRISTOPH AMES¹, GÜNTER DUCKECK¹, RODNEY WALKER¹, and CÉDRIC SERFON² — ¹Ludwig-Maximillians-Universität, Munich — ²Brookhaven National Laboratory

Rucio is a software framework that enables the management of large collections of data. Originally developed for the ATLAS experiment at the Large Hadron Collider (LHC), it has been expanded to encompass other scientific experiments, such as CMS and Belle II, whereby data is stored on and transferred between over 100 data centres across the world. Accessing files on these data centres doesn't always function without problems occurring, therefore dedicated services are used to track files that are being accessed or transferred. Files that cause problems are marked as suspicious by the tracking services and have to be handled by another service, the suspicious replica recoverer. The goal of this work is to expand the suspicious replica recoverer by adding more flexibility and nuance when decidLocation: T-H28

ing what should be done with a suspicious file. This is achieved by introducing a policy system based on file metadata, which allows each virtual organisation to individually set their own policies.

T 15.2 Mon 16:30 T-H28 **Probing a tiered storage for a WLCG Tier-3 center** — MICHAEL BÖHLER, •DIRK SAMMEL, and MARKUS SCHUMACHER — Albert-Ludwigs-Universität Freiburg

The storage of large amounts of data is an important topic in the HEP community, especially with the start of the High-Luminosity LHC in the near future. This data has to be available to local users at the institutes, which represent the Tier-3 level of the WLCG infrastucture. An interesting scenario would be a "tiered" Tier-3 system consisting of a combination of cold and warm storage. In such a setup, a parallel file system might serve as the warm storage and in addition, an object storage would be used as cold storage to avoid a too large occupancy of the

T 15.3 Mon 16:45 T-H28

Modelling Large Scale Distributed Computing Systems for Identification of Efficient Architectures — •MAXIMILIAN HORZELA, MANUEL GIFFELS, ARTUR GOTTMANN, GÜNTER QUAST, and ACHIM STREIT — Karlsruhe Institute of Technology

One approach to overcome the expected gap between available and required resources at future colliders like the High Luminosity Large Hadron Collider (HL-LHC) is to increase the efficiency of existing workflows, for example via caching of frequently used data. However, in complex environments with distributed computing systems and many users like the Worldwide LHC Computing Grid (WLCG), finding a solution is not trivial. Due to the complexity and size of such systems, it is not feasible to deploy several experimental test-beds at large scales.

Simulation of such systems has proved a scalable and versatile approach to identify efficient and practical computing architectures for High Energy Physics (HEP). A prominent example is the Monarc Simulation Framework, which lead to the present structure of the WLCG. However, the discontinuation of the old software and the demand for the ability to simulate recent scenarios including caching solutions requires a revision of suitable software.

In the context of this talk, a modern example for such applications based on the WRENCH and SimGrid simulation packages will be presented. Furthermore, first results obtained by the simulation will be shown.

T 15.4 Mon 17:00 T-H28

Dynamic and transparent provisioning of opportunistic compute resources for HEP — • RALF FLORIAN VON CUBE, RENÉ CASPART, MAX FISCHER, MANUEL GIFFELS, EILEEN KÜHN, GÜNTER QUAST, and MATTHIAS SCHNEPF — Karlsruhe Institute of Technology

The utilization of only temporarily available compute resources (opportunistic resources) not dedicated to HEP becomes more and more important for future HEP experiments. On the one hand, due to the unprecedented need and resulting scarcity of HEP compute resources. On the other hand, due to the desired integration of HPC, Cloud and locally available resources.

To meet the challenges posed by the resulting heterogeneous compute environment, the Karlsruhe Institute of Technology (KIT) developed the COBalD/TARDIS resource manager and an entire ecosystem around an HTCondor overlay batch system to allow for a dynamic, transparent and hassle-free integration of those resources into the World-wide LHC Computing Grid (WLCG) via a single point of entry.

In this contribution we will present the current status of COBalD/TARDIS developments as well as our experience with the integration of various opportunistic resources into the WLCG.

T 15.5 Mon 17:15 T-H28 AUDITOR: An accounting system for opportunistically used resources -MICHAEL BÖHLER, •STEFAN KROBOTH, and MARKUS SCHUMACHER — Albert-Ludwigs-Universität Freiburg

Computing clusters often experience varying workload over time, which may

lead to suboptimal utilization of available hardware. Sharing resources between multiple clusters can mitigate this inefficiency. The software COBalD/TARDIS enables the integration of resources in an opportunistic, dynamic, and transparent manner and is successfully operated on various sites for a wide range of scenarios. In such a setup, the question of how to account for computations conducted on shared resources arises.

In this work the prototype of AUDITOR, the AccoUnting Data handlIng Toolbox for Opportunistic Resources is presented. At its core it consists of a database and a server which can be interacted with using a REST API. This allows collectors to store data relevant for accounting in the database and plugins to publish the data on an external platform or perform other actions. The extensible nature of AUDITOR enables handling of various use cases by combining the appropriate collectors and plugins. An example use case is described, where the fairshare of one batchsystem is transfered to priorities of another based on data collected from COBalD/TARDIS.

T 15.6 Mon 17:30 T-H28

Optimization of performance for HEP ML applications on GPU Clusters •TIM VOIGTLÄNDER, RENÉ CASPART, MANUEL GIFFELS, GÜNTER QUAST, MATTHIAS SCHNEPF, and ROGER WOLF - Karlsruhe Institute of Technology, Karlsruhe, Germany

GPU clusters are gaining increased importance also in particle physics. To use GPUs most efficiently, concepts like multi-processing on a single GPU, multi-GPU usage for suitable applications or the balance between CPU and GPU resources must be considered. In particular, GPU support for applications in Machine Learning has become quite common, and they provide a wide variety of usage scenarios. The GPU performance in relation to CPUs depends on the complexity of the network topology, on the training strategy and other hyperparameters of the problem at hand. To illustrate the possible performance gains, a number of scenarios in neural network training on a shared GPU cluster attached to the TOpAS Tier3 at KIT are discussed.

T 15.7 Mon 17:45 T-H28 Belle II Grid Computing Developments at KIT - • MORITZ BAUER, R. FLO-RIAN VON CUBE, TORBEN FERBER, MANUEL GIFFELS, MAXIMILIAN HORZELA, GÜNTER QUAST, and MATTHIAS SCHNEPF — Karlsruhe Institute of Technology The Belle II experiment studies B-meson decays with high precision. Therefore, Belle II plans to record 50ab⁻¹. This will result in 50PB of raw data, which has to be reconstructed and analyzed. Furthermore, corresponding simulations have to be produced. To achieve this goal the Belle II collaboration uses several data centers around the world as a Grid, similar to the worldwide LHC Computing Grid.

While LHC experiments usually only analyze data from one data-taking period, Belle II analyses typically use the complete dataset over all periods. The high energy physics computing group at KIT works on several development projects to handle these challenges. In this presentation, we describe the challenges as well as the current development projects. These projects include the transparent integration of opportunistic resources to provide more computing resources and caching to provide additional copies of datasets to improve their accessibility automatically.

T 16: Experimental Methods (general) 1

Time: Monday 16:15-18:30

Measurement of the CMS offline tracking efficiency from the ratio of reconstructed D^* and D^0 mesons — •Yewon Yang and Achim Geiser — DESY, Hamburg, Germany

The efficiency for offline track reconstruction in CMS is measured directly from the data using a novel method. This method is based on taking the ratio of charm mesons reconstructed in the decay chains $D^{*\pm} \to K^{\mp} \pi^{\pm} \pi^{\pm}_s$ and $D^0 \to K^{\mp} \pi^{\pm}$, using the special kinematics of the so-called 'slow pion' π_s from D^* decay. It also requires the treatment of the a priori unknown mixture of prompt (from charm) and non-prompt (from beauty) contributions to the final states. Details of the method are explained and first results for the actual tracking efficiencies are presented.

T 16.2 Mon 16:30 T-H29 Clustering and tracking in dense environments with the ITk - •NICOLA DE

BIASE — DESY Hamburg Dense hadronic environments, encountered in particular in the core of high- p_T jets or hadronic τ decays, present specific challenges for the reconstruction of charged-particle trajectories (tracks) in the ATLAS silicon-pixel tracking detector, as the charge clusters left by different ionising particles in the silicon sensors can merge with a sizeable rate. Tracks competing for the same cluster are penalised for sharing it, leading to a loss in tracking efficiency.

In the current ATLAS Inner Detector, a machine learning algorithm is used for classifying and splitting merged clusters with minimal efficiency losses, leading to better performances of Clustering and Tracking in Dense Environments (CTIDE). The new Inner Tracker (ITk), which will replace the current Inner Detector as part of the ATLAS phase-2 upgrade, will benefit from an improved granularity thanks to its smaller pixel sensor size, which might render such a procedure unnecessary.

In this talk, the expected performance of the ITk in dense environments will be discussed, addressing the question of whether a cluster splitting procedure is necessary.

T 16.3 Mon 16:45 T-H29

Location: T-H29

Clustering and Tracking in Dense Environment studies of the ATLAS ITk Strip Detector for the HL - LHC Upgrade — KATHARINA BEHR, NICHOLAS STYLES, and •AKHILESH TAYADE — DESY, Hamburg, Germany

The new Inner Tracker (ITk) in the ATLAS detector is being built as a part of the HL-LHC upgrade. The High Luminosity upgrade will see an increase in track density, pile up and collisions per bunch crossing. Correspondingly, the current offline tracking reconstruction is being upgraded to handle this. Strongly boosted objects give highly collimated tracks. Reconstructing these tracks can be

crucial in discovering new physics. Highly collinear and dense tracks are likely to share pixel or strip clusters for which they are penalized in the track reconstruction. Machine learning techniques are employed for the current ATLAS Pixel Detector to resolve the ambiguities due to cluster merging. In this talk, we discuss whether similar techniques are needed for clusters in the ITk strip detector.

T 16.4 Mon 17:00 T-H29

Jet Vertex Tagger in release 22 – •ABDULLAH NAYAZ¹, TENG JIAN KHOO², and CIGDEM ISSEVER³ – ¹Humboldt University, Berlin, Germany – ²Humboldt University, Berlin, Germany – ³Humboldt University, Berlin, Germany

Pile-up mitigation is a crucial part of many important Particle Physics analysis e.g. HH->4b. The Jet Vertex Tagger (JVT) is a multivariate pile-up suppression variable developed for the ATLAS experiment that combines information from other track based pile-up variables and plays a major role in ATLAS analysis. In this study, as part of the preparation for Run 3 data-taking and analysis, the performance of JVT has been checked for the new release 22 Track to Vertex Association (TTVA) working points using Monte-Carlo simulated dijet data samples. First, the TTVAs that result in a good performance of the JVT have been identified. Furthermore, to increase the JVT performance, a Multilayer Perceptron Neural Network (NN) has been used to retrain the JVT for release 22. The training was done separately for offline and trigger level jets, varying the inputs to the NN to optimise the separation of hard scatter and pile-up jets. Some improvement on the JVT performance was observed after the training process which will be beneficial for Run 3 ATLAS analyses.

T 16.5 Mon 17:15 T-H29

Global χ^2 fitter for Acts — FLORIAN BERNLOCHNER, JOCHEN DINGFELDER, and •RALF FARKAS — Physikalisches Institut der Rheinischen Friedrich-Wilhelms-Universität Bonn

The reconstruction of trajectories of charged particles is a crucial task for most HEP experiments. The Acts (A Common Tracking Software) aims to be a generic, framework and experiment-independent toolkit for track reconstruction, initially started from the ATLAS tracking software. My talk summarizes the recent development of a global χ^2 fitter for Acts, which complements and validates the existing (Combinatorial) Kálmán Filter.

T 16.6 Mon 17:30 T-H29

Track Reconstruction of the FASER Experiment — FLORIAN BERNLOCHNER, •ТОВІАЅ ВӦСКН, JOCHEN DINGFELDER, and MARKUS PRIM — Physikalisches Institut der Rheinischen Friedrich-Wilhelms-Universität Bonn

FASER (ForwArd Search ExpeRiment) is a new, small and inexpensive experiment designed to search for light, weakly interacting particles during Run 3 of the LHC. Such particles may be produced in large quantities in proton-proton collisions, travel for hundreds of meters along the beam axis, and can decay in two charged Standard Model particles. To reach its physics goals, a good hit resolution, and track reconstruction to separate the two closely-spaced, oppositely charged tracks is essential. In this talk, I review the track reconstruction, which is based on the ACTS toolkit. ACTS aims to provide an experiment-independent toolkit for track reconstruction.

T 16.7 Mon 17:45 T-H29

detray - a GPU-friendly tracking geometry description — ANDREAS SALZBURGER¹, •JOANA NIERMANN^{1,2}, BEOMKI YEO^{3,4}, ATTILA KRASZNAHORKAY¹, and STAN LAI² — ¹CERN — ²II. Physikalisches Institut, Georg-August-Universität Göttingen — ³Department of Physics, University of California — ⁴Lawrence Berkeley National Laboratory

With the next generation high luminosity experiments, the computational demand of particle track reconstruction will increase strongly. A potential way to tackle this is by offloading highly parallelizable tasks to an accelerator device. Existing codebases need to be adapted to e.g. specific host and device memory management and the calling of dedicated compute kernels, while avoiding code duplication as much as possible. Designed to be integrated into ACTS (A Common Tracking Software), which provides efficient algorithms for common tracking tasks, detray is an ongoing R&D effort to formulate the tracking geometry description for heterogeneous hardware. In order to propagate track states through the geometry model, detray follows the navigation design established in ACTS, but presents the geometry in a GPU-friendly way. It makes use of flat container structures without runtime polymorphism, a dedicated memory management scheme provided by the vecmem library, as well as direct indexing to link the geometry data, which together allows to instantiate the geometry model in host and device code. This talk gives an overview of the detray tracking GPU-friendly approach.

T 16.8 Mon 18:00 T-H29 ATLAS Primary Vertexing with ACTS — •BASTIAN SCHLAG^{1,2}, AN-DREAS SALZBURGER¹, MARKUS ELSING¹, CHRISTIAN SCHMITT², and VOLKER BÜSCHER² — ¹CERN — ²Johannes Gutenberg-Universität Mainz

The reconstruction of particle trajectories and their associated vertices is a crucial task in the event reconstruction of most high energy physics experiments. In order to maintain or even improve upon the current performance of tracking and vertexing algorithms under the upcoming challenges of increasing energies and ever increasing luminosities in the future, major software upgrades are required.

Based on the well-tested ATLAS tracking and vertexing software, the ACTS (*A Common Tracking Software*) project aims to provide a standalone, modern and experiment-independent toolkit of track- and vertex reconstruction software, specifically designed for parallel code execution. The newly developed ACTS vertexing software suite provides thread-safe, highly performant and state-of-the-art vertex reconstruction tools that have been fully integrated and validated in the ATLAS software framework AthenaMT. Due to its superb physics and CPU performance, the ACTS vertexing software will be used as the default primary vertex reconstruction tool in ATLAS for LHC Run 3.

Additionally, an entirely new vertex seed finding algorithm with great physics performance and CPU speed-ups of up to a factor of 100 in high pile-up environments has been developed and implemented in ACTS.

This talk presents an overview of the ACTS vertexing software suite, its performance in ATLAS as well as latest developments.

T 16.9 Mon 18:15 T-H29

Electron reconstruction and identification with the ATLAS detector – •ASMA HADEF and LUCIA MASETTI – Johannes Gutenberg Universität, Mainz, Germany

Electrons are important objects both for the search for new physics and for precision measurements. In the ATLAS detector, electrons in the central detector region are triggered by and reconstructed from energy deposits in the electromagnetic (EM) calorimeter that are matched to a track in the inner detector. Electrons are distinguished from other particles using identification (ID) criteria with different levels of background rejection and signal efficiency. The electron ID used in ATLAS for Run 2 is based on a likelihood discrimination to separate isolated electron candidates from candidates originating from photon conversions, hadron misidentification and heavy flavor decays. The performance of the electron reconstruction and ID algorithms is evaluated by measuring efficiencies using tag-and-probe techniques with large statistics samples of isolated electrons from $Z \rightarrow ee$ and $J/\psi \rightarrow ee$ resonance decays. These measurements were performed with pp collisions data at $\sqrt{s} = 13$ TeV in 2015-2018 corresponding to an integrated luminosity of 139 fb^{-1} and studied as a function of the electron's transverse momentum, pseudorapidity and number of primary vertices. Furthermore, in order to achieve reliable physics results, the simulated samples need to be corrected to reproduce the measured data efficiencies as closely as possible. For this reason, the efficiencies are estimated both in data and in simulation. The scale factors (data to MC efficiency ratios) are then estimated and provided to all physics analyses involving electrons.

T 17: Gamma Astronomy 1

Time: Monday 16:15-18:20

 Group Report
 T 17.1
 Mon 16:15
 T-H30

 Status and First Results of the CTA Large-Sized Telescope
 •MARTIN WILL

 for the CTA Consortium
 Max-Planck-Institut für Physik, München

The prototype of the Large-Sized Telescope (LST), intended to become part of the Northern site of the Cherenkov Telescope Array (CTA) in the Canarian island of La Palma, is currently finishing its commissioning phase and started taking engineering data runs. With its reflective surface of 23 meter diameter, the LSTs are optimized to detect gamma-rays in the energy range between 20 and 200 GeV. In this presentation, some preliminary physics results, the performance of the prototype, and the plans to construct more LSTs as part of CTA will be shown.

Location: T-H30

T 17.2 Mon 16:35 T-H30

Muons as a tool for background rejection in Imaging Atmospheric Cherenkov Telescope arrays — •LAURA OLIVERA-NIETO¹, ALISON MITCHELL^{2,3}, KONRAD BERNLÖHR¹, and JIM HINTON¹ — ¹Max-Planck-Institut für Kernphysik, Heidelberg, Germany — ²Department of Physics, ETH Zurich, Zurich, Switzerland — ³Erlangen Centre for Astroparticle Physics, Erlangen, Germany

The presence of muons in air-showers initiated by cosmic ray protons and nuclei is well established as a powerful tool to separate such showers from those initiated by gamma rays. However, so far this approach has been fully exploited only for ground level particle detecting arrays. We explore the feasibility of using Cherenkov light from muons as a background rejection tool for imaging atmospheric Cherenkov telescope arrays at the highest energies. We adopt an analytical model of the Cherenkov light from individual muons to allow rapid simulation of a large number of showers in a hybrid mode. This allows us to explore the very high background rejection power regime at acceptable cost in terms of computing time. We show that for very large (≥ 20 m mirror diameter) telescopes, efficient identification of muon light can potentially lead to background rejection levels up to 10^{-5} whilst retaining high efficiency for gamma rays. While many challenges remain in the effective exploitation of the muon Cherenkov light in the data analysis for imaging Cherenkov telescope arrays, our study indicates that for arrays containing at least one large telescope, this is a very worthwhile endeavor.

T 17.3 Mon 16:50 T-H30

Adjusting Monte Carlo Simulations for the Cherenkov Telescope Array's Large-Sized Telescope Prototype — •LUKAS NICKEL and MAXIMILIAN NÖTHE FOR THE CTA LST PROJECT — Astroparticle Physics, WG Elsässer, TU Dortmund University, Germany

The lowest energy range of the Cherenkov Telescope Array, which is going to be the next-generation very-high energy ($\geq 20 \text{ GeV}$) gamma-ray observatory, will be covered by the Large-Sized Telescopes (LSTs). The prototype of the LST was inaugurated in October 2018 on the Canary Island of La Palma and has since performed observations of bright gamma-ray sources as part of the commissioning process.

One area that needs special care for any analysis regards potential differences between Monte Carlo simulations and observed data. In this talk current approaches to adjust existing simulations for different observational conditions will be presented.

T 17.4 Mon 17:05 T-H30

Data Volume Reduction for the Cherenkov Telescope Array's Large-Sized Telescope Prototype — •JONAS HACKFELD for the CTA-Collaboration — Institute for theoretical physics IV, Ruhr-University Bochum, Germany

The prototype of the Large-Sized Telescope (LST) of the Cherenkov Telescope Array (CTA), which is going to be the next-generation very-high energy (>20 GeV) gamma-ray observatory, was inaugurated in October 2018 and has already observed several bright gamma-ray sources during its commissioning phase. For the next years, in addition to 3 more LSTs, several Medium-Sized Telescopes (MST) are planned, which together will equip the northern site of the CTA Observatory. Due to the locally limited data transfer rates and the technical and economic effort to store data quantities of ~ 100 PByte/year permanently over a planned duration of ~ 30 years, a volume reduction for low level data is inevitable. In addition to lossless compression methods for volume reduction, there are lossy methods such as pixel selection. In this process, the pixels with signal are isolated from the night-sky background ones, so that the physics results are impacted as minimally as possible during subsequent event reconstruction. In this talk, pixel selection algorithms and their impact on higher level data analysis will be presented.

This project is funded via BMBF Verbundforschung, Project 05A20PC1

T 17.5 Mon 17:20 T-H30

Background Estimation: Towards an Analysis of MAGIC Data Using Gammapy — •SIMONE MENDER and LENA LINHOFF for the MAGIC-Collaboration — Astroparticle Physics WG Elsässer, TU Dortmund University, Germany

The Python package Gammapy is mainly developed for the high-level analysis of gamma-ray data of the future Cherenkov Telescope Array. As it can also be used to analyze data from existing imaging air Cherenkov telescopes, it is reasonable to perform the high-level analysis of MAGIC data with Gammapy and compare the results with existing analyses. Gammapy requires event-based data combined with the corresponding instrument response functions. In order to process these so-called DL3 data for MAGIC, the new database-based framework AutoMAGIC is developed. With AutoMAGIC it is possible to create DL3 data in an automated and reproducible way.

So far, the MAGIC DL3 data does not include background models, which are needed for the creation of skymaps. In this talk, the dependence of the background on parameters like the azimuth and zenith angle of the pointing position will be presented. Using the exclusion region method, background models are built from DL3 data. Using these background models, the first preliminary skymaps of MAGIC data created with Gammapy will be shown.

T 17.6 Mon 17:35 T-H30

Automatized Analysis of MAGIC Sum-Trigger-II Data — •JAN LUKAS SCHU-BERT and LENA LINHOFF for the MAGIC-Collaboration — Astroparticle Physics WG Elsässer, TU Dortmund University, Germany

The MAGIC telescopes are a stereoscopic system of Imaging Air Cherenkov Telescopes. They are used for the detection of gamma rays in the GeV to TeV range. With the Sum-Trigger-II, low-energy data with a threshold as low as 25 GeV can be recorded. This data requires a dedicated analysis adapted to the low energies. Since the analysis structure is complex, it is reasonable to automatize the analysis in order to save time for an analyzer and to deliver entirely reproducible results. The automatization of the analysis of Sum-Trigger-II data was implemented in the autoMAGIC project which aims to automatize the entire MAGIC analysis chain.

For testing the performance, currently an analysis of the Crab pulsar is performed based on the autoMAGIC output and compared to previous analyses.

In the future, the automatization of the analysis of Sum-Trigger-II data could be used for further optimizations of the low-energy analysis as well as for comparisons of low-energy data from MAGIC and the CTA-LST1.

T 17.7 Mon 17:50 T-H30

Characterization of the performance of the MAGIC LIDAR — •FELIX SCHMUCKERMAIER for the MAGIC-Collaboration — Max-Planck-Institut für Physik — TU München

The Major Atmospheric Gamma-ray Imaging Cherenkov (MAGIC) telescopes are a system of two Imaging Atmospheric Cherenkov Telescopes (IACTs). IACTs make calorimetric use of the Earth's atmosphere, which allows them to reach large effective areas, but also makes them strongly dependent on the quality of the atmosphere at the time of the observations. Suboptimal conditions can then lead to a wrong reconstruction of the gamma-ray data. In order to mitigate this problem, the MPP group built and has been operating a single wavelength elastic LIDAR (LIght Detection And Ranging) system to perform real time rangedresolved measurements of the atmospheric transmission. This information is then used to quantify the quality of the telescope data, as well as to correct the data taken under suboptimal atmospheric conditions. In this talk, I will present the first detailed characterization of the correction capabilities of the LIDAR system. The results describe the impact of the LIDAR corrections for a variety of atmospheric and observational conditions, and therefore contribute to a better understanding of the telescope's performance and related systematic uncertainties.

T 17.8 Mon 18:05 T-H30

Location: T-H31

Combined analysis pipeline for joint observations with MAGIC and CTA LST-1 — •GIORGIO PIROLA¹, YOSHIKI OHTANI², ALESSIO BERTI¹, DAVIDE DEPAOLI³, FEDERICO DI PIERRO³, DAVID GREEN¹, LEA HECKMANN¹, MORITZ HÜTTEN¹, RUBEN LÓPEZ-COTO⁴, ABELARDO MORALEJO⁵, DANIEL MORCUENDE⁶, MARCEL STRZYS², YUSUKE SUDA⁷, IEVGEN VOVK², and MARTIN WILL¹ for the CTA LST project and the MAGIC Collaboration — ¹Max Planck Institute for Physics, München, Germany — ²ICRR, the University of Tokyo, Japan — ³INFN Sezione di Torino, Italy — ⁴INFN Sezione di Padova, Italy — ⁵IFAE, Barcelona, Spain — ⁶Universidad Complutense de Madrid, Spain — ⁷Hiroshima University, Japan

The performance achievable with Imaging Atmospheric Cherenkov Technique is remarkably improved by using multiple telescopes. Currently, in La Palma (Canary Islands), there are three operative Cherenkov telescopes: the two MAGIC telescopes and the prototype Large-Sized Telescope (LST-1), intended for the future northern site of the Cherenkov Telescope Array (CTA), the next generation Very-High Energy gamma-ray observatory. The data acquired during nights of simultaneous observation of the same target have been used to develop and test an algorithm for the offline search of coincident events. Recently, this algorithm was implemented for the development of the combined analysis pipeline: an analysis chain meant to perform the 3-telescope event reconstruction. The talk aims to present the current status of the pipeline and to give an insight into the possible results achievable with the 3-telescope system.

T 18: Neutrino Astronomy 1

Time: Monday 16:15-18:30

T 18.1 Mon 16:15 T-H31

Follow-up of high-energy neutrino events in IceCube – •MARTINA KARL for the IceCube-Collaboration — Max Planck Institute for Physics, Munich, Germany — Technical University of Munich, Department of Physics, Garching, Germany We investigate the arrival direction of the most energetic track-like neutrino events in IceCube. These high-energy events allow for a good pointing back to their origin direction and have a high probability to be of astrophysical origin. Roughly 10 of these track-like high-energy neutrino events are detected per year. With these events acting as a source catalog, we present a search for cosmic neu-

96

trino emission on 11 years of IceCube neutrino-induced muon data. We explore the hypotheses of steady and transient neutrino emission, and present methods to find neutrino flares.

T 18.2 Mon 16:30 T-H31

High-Energy Neutrinos From Accretion Flares — \bullet Jannis Necker for the IceCube-Collaboration — DESY, Zeuthen, Deutschland

The past two decades have seen a revolution in astronomy as for the first time it became possible to gain information about astrophysical processes not only from (low energy) photons but also from other messengers such as gravitational waves and neutrinos. The IceCube observatory is a cubic kilometre neutrino detector array in the antarctic ice, looking for astrophysical, high-energy neutrinos. The collected data reveal a diffuse flux of these neutrinos over the whole sky, indicating an extragalactic origin. Recent observations suggest a contribution to this diffuse flux from accretion flares, radiation outbursts from Super-Massive Black Holes that accrete at an enhanced rate. In this contribution I will present results from a stacking analysis looking for IceCube neutrinos from these accretion flares.

T 18.3 Mon 16:45 T-H31

A study to detect neutrino signals from AGN using machine-learning methods for source classification — •SEBASTIAN SCHINDLER for the IceCube-Collaboration — Erlangen Centre for Astroparticle Physics (ECAP), University Erlangen-Nürnberg, Germany

The IceCube Neutrino Observatory is currently the world's largest high-energy neutrino detector. After the initial detection of a diffuse astrophysical neutrino flux in 2013, one of the main goals has been to associate parts of this flux with specific source classes. A few tentative "hot spots" at the three-sigma level have been found and associated with certain classes of Active Galactic Nuclei (AGN), among them blazars and Seyfert galaxies. The underlying physical mechanisms of neutrino production, however, remain poorly understood. One problem for neutrino source searches comes from the use of historically-driven class definitions of AGN, which are based on specific spectral properties that are not necessarily optimal for the selection of potential neutrino sources.

This talk will motivate a study that aims to address this problem in two stages. The first stage will use multi-wavelength data to define a source selection using modern machine-learning approaches in a way that emphasizes intrinsic physical properties and mostly disregards the general AGN classification. This will allow us to identify potential neutrino sources similar in physical properties to those associated with the currently detected "hot spots". The second part will perform a statistical analysis in the form of a correlation analysis, for example a stacking search, using these previously defined source selections.

T 18.4 Mon 17:00 T-H31

Blazar stacking using new and improved neutrino point-source analysis method — •TOMAS KONTRIMAS, MARTINA KARL, and CHIARA BELLENGHI for the IceCube-Collaboration — Physik-department, Technische Universität München, D-85748 Garching, Germany

The IceCube Neutrino Observatory has experienced remarkable success in the field of neutrino astronomy since its completion. One of the main goals is to identify sources of the diffuse astrophysical high-energy neutrino flux. In 2018, IceCube found evidence for a correlation between high-energy neutrinos and the blazar TXS 0506+056. Blazars, one of the most powerful objects in the universe, are among the most promising source candidates of high-energy neutrinos. We present a new method improving the accuracy of the likelihood function, including enhanced neutrino reconstruction and data calibration. Furthermore, we discuss a correlation study between IceCube high-energy neutrinos and different classes of blazars.

T 18.5 Mon 17:15 T-H31

Neutrino source searches with DNN based Cascade Dataset in IceCube — •MIRCO HÜNNEFELD for the IceCube-Collaboration — Astroparticle Physics WG Rhode, TU Dortmund University, Germany

IceCube has discovered a flux of astrophysical neutrinos and presented evidence for one neutrino source, a flaring blazar known as TXS 0506+056. However, the sources responsible for the majority of the astrophysical neutrino flux remain elusive. While charged-current muon-neutrino datasets (track events) are predominantly used for source searches due to their superior pointing resolution, cascade events (neutral-current interactions of all neutrino flavors and chargedcurrent interactions of electron- and tau-neutrinos) allow to lower the energy threshold in the southern sky for IceCube. In this contribution, searches for neutrino sources are presented utilizing an improved cascade dataset that builds upon recent advances in deep learning based reconstruction methods. The resulting dataset improves IceCube's sensitivity in the southern neutrino sky and is thus particularly promising for the identification of neutrino production from the galactic plane.

T 18.6 Mon 17:30 T-H31

Constraining populations of astrophysical neutrino sources with IceCube — •CHIARA BELLENGHI and KRISTIAN TCHIORNIY for the IceCube-Collaboration — Technische Universität München, Physik-Department, James-Franck-Str. 1, 85748 Garching

The discovery of a diffuse flux of high-energy astrophysical neutrinos in 2013 by the IceCube neutrino observatory has triggered a vast effort to identify the mostly unknown sources of this signal. We present an analysis optimized for identifying an excess of astrophysical neutrino clusterings produced by a population of sub-threshold point sources. We aim at testing the hypothesis of timeintegrated emission in the Northern Hemisphere using 9 years of IceCube data. We present here the methods and the potential of the analysis on constraining the neutrino flux contribution from populations of neutrino point sources.

T 18.7 Mon 17:45 T-H31

Targeting luminous optical transients in the search for high-energy neutrinos — •MASSIMILIANO LINCETTO for the IceCube-Collaboration — Astronomisches Institut, Ruhr-Universität Bochum, Bochum, Germany

Years after the discovery of astrophysical neutrinos by the IceCube Neutrino Observatory, the dominant sources of the measured flux are still to be determined. Despite existing evidence in favour of blazars, multi-messenger considerations suggest the need of sources that do not produce high-energy gamma rays. Recent observations, following the detection of a high-energy neutrino in coincidence with a tidal disruption event, point to accreting black holes as promising candidate sources. With the rise of wide-field optical surveys such as the Zwicky Transient Facility, an unprecedented amount of optical transients is being observed on a regular basis. Among these, superluminous supernovae (SLSNe) stand out as the most luminous class. The power source behind such extreme phenomena is still unclear: magnetars, black hole accretion or CSM interaction have been proposed to explain their increased luminosity. In this contribution, the prospects for targeting SLSNe in a search for high energy neutrinos with Ice-Cube data are presented, giving an overview of the candidate event catalogue and the proposed analysis methods.

T 18.8 Mon 18:00 T-H31

A Combined Analysis of IceCube's High Energy Muon Tracks and Cascades Neutrino Data — •ERIK GANSTER¹, MARKUS ACKERMANN², JAKOB BÖTTCHER¹, PHILIPP FÜRST¹, JONAS HELLRUNG¹, RICHARD NAAB², GEORG SCHWEFER¹, ROMAN SUVEYZDIS¹, and CHRISTOPHER WIEBUSCH¹ for the IceCube-Collaboration — ¹Physics Institute III B, RWTH Aachen University, Germany — ²DESY, Zeuthen, Germany

The IceCube Neutrino Observatory is observing a diffuse flux of high-energy astrophysical neutrinos in multiple detection channels. We combine two of these channels, through-going muon tracks and contained cascades, in a single analysis that employs a consistent treatment of signal and background as well as systematic uncertainties in a global fit. Then, the complementary information from the two channels reduces the overall uncertainties in signal and background. This improves our understanding of the astrophysical neutrino flux properties: measuring the energy spectrum and testing the flux composition. We will describe the method of this global fit and present first results from 10 years of Ice-Cube neutrino data.

T 18.9 Mon 18:15 T-H31

Sensitivity of IceCube-Gen2 for the identification of high-energy tau neutrinos and for the measurement of the flavour composition — •NEHA LAD¹, MAXIMILLIAN MEIER², and MARKUS ACKERMANN¹ for the IceCube-Collaboration — ¹DESY, Zeuthen, Germany — ²Dept. of Physics and Institute for Global Prominent Research, Chiba University, Japan

The IceCube neutrino observatory at the South Pole has disfavoured the absence of an astrophysical tau-neutrino flux at the 2.8σ level. IceCube-Gen2 is the planned extension of current IceCube detector, which will be about 8 times bigger than the current instrumented volume. In this work, we study the sensitivity of IceCube-Gen2 to the astrophysical flavour composition and investigate it's tau neutrino identification capabilities. We apply the IceCube analysis on a Gen2 dataset that mimics the High Energy Starting Event classification. Reconstructions are performed using sensors that have 3 times higher quantum efficiency and isotropic angular acceptance compared to the current IceCube optical modules. We present results of this sensitivity study.

T 19: Cosmic Ray 1

Time: Monday 16:15-18:20

Group Report

T 19.1 Mon 16:15 T-H32 The Pierre Auger Observatory - Status, Results, Prospects - • PHILIP RUEHL for the Pierre Auger-Collaboration - Center for Particle Physics Siegen, Exper-

imental Astroparticle Physics, University of Siegen With an instrumented area of about 3000 km², the Pierre Auger Observatory is the world's largest experiment for measuring cosmic particles. Its surface detector (SD) array is comprised of 1660 water Cherenkov detectors with a nextneighbor spacing of 1.5 km. The atmosphere above the SD is monitored by 27 fluorescence telescopes.

Precision measurements of the cosmic-ray energy spectrum have recently been extended down to 10¹⁷ eV using the SD-750 low-energy enhancement of the SD array, showing a gradual transition of the spectral index just above the "knee". A first measurement of the fluctuations in the muon content of air showers at ultra-high energies can be used to constrain hadronic interaction models trying to explain the observed muon deficit in air shower simulations. Recent advancements in the application of deep neural networks to Auger data enabled the successful reconstruction of X_{\max} from SD data and an extraction of the muon component of simulated SD signal traces. The construction of the AugerPrime upgrade together with the AMIGA enhancement is in progress and will further enhance the mass discrimination power of the Observatory with additional scintillation and radio detector units in the SD array.

The Pierre Auger project is supported by BMBF Verbundforschung Astroteilchenphysik.

T 19.2 Mon 16:35 T-H32

Analysis of laser shots of the Aeolus satellite in the Pierre Auger Observatory - •FELIX KNAPP for the Pierre Auger-Collaboration — Karlsruher Institut für Technologie

The Pierre Auger Observatory is a large-scale facility for the investigation of ultra-high-energy cosmic rays. It uses a combination of surface detectors and fluorescence telescopes to measure extensive air showers initiated by cosmic-ray particles. Aeolus is an ESA-operated satellite with the mission of conducting global wind profile measurements. To this end, a UV-lidar is employed which emits laser beams towards Earth. When passing over the Pierre Auger Observatory, light that scatters off the laser beam in the atmosphere is detected by the Fluorescence Detector of the Observatory. This allows for a reconstruction of the laser tracks from the Fluorescence Detector data for several overpasses each year. These laser tracks provide a unique opportunity for analyses of the atmosphere above the Observatory.

In this presentation, we will explain the process of reconstructing laser tracks from data taken by the fluorescence telescopes and give an overview of the possible application for aerosol measurements. Furthermore, some results of the laser reconstruction are shown, including the most recent overpasses under a special orbital configuration of the satellite.

T 19.3 Mon 16:50 T-H32

A method to determine baselines of time traces at the Pierre Auger Observatory — •Tobias Schulz, David Schmidt, Darko Veberič, and Markus Rотн for the Pierre Auger-Collaboration — Karlsruhe Institute of Technology, Institute for Astroparticle Physics, Karlsruhe, Germany

The calibration and identification of detector signals is crucial for minimizing systematic uncertainties in measurements. A constant offset, called the baseline, which is generated by the electronics, has to be determined to properly estimate the size of the signals. At the Pierre Auger Observatory, two detector types, namely Water Cherenkov and Surface Scintillation Detectors, are used to measure the lateral distribution of extensive air showers at the ground. To determine the signal produced by the particles that enter the detectors, photomultiplier tubes (PMTs) are used to collect the emitted Cherenkov or scintillation light. The PMTs have one low gain channel and one amplified high gain channel. The analog pulses are read out and sampled with a flash analog to digital converter in a FADC time trace.

After a signal, the output of the PMTs is reduced by an undershoot, resulting in a lowered baseline, that recovers after a certain time period. Accidental muons or late air shower components can result in additional signal in the traces, which complicate the estimation of the baseline. Here, we present a method to estimate the baseline, that is robust to signal contributions in the trace and accounts for undershoot of the PMT.

T 19.4 Mon 17:05 T-H32

Combined analysis of the ultrahigh-energy cosmic-ray mass composition and hadronic interaction cross-sections — •OLENA TKACHENKO for the Pierre Auger-Collaboration — Karlsruhe Institute of Technology, Karlsruhe, Germany Studies of the cosmic-ray mass composition and hadronic interaction properties can improve our understanding of the origin and nature of the ultrahighenergy cosmic rays. However, neither the mass composition nor the interaction cross sections are well measured at ultrahigh energies and normally the standard analyses require certain assumptions on either of these quantities to estimate the other one.

In this talk, we present a method for the independent and simultaneous estimation of the cosmic-ray composition fractions and proton-proton interaction cross-sections. We perform a standard mass composition fit using the measured distributions of the shower maximum of air showers (X_{max}) while varying at the same time the interaction cross-sections and thus getting the best-fit combination of the estimated quantities without making any underlying assumptions on either of them. For this purpose, we modify the proton-proton interactions and the corresponding nucleus-nucleus cross-sections are then rescaled via the Glauber theory. We test the performance of this method and its sensitivity for the different composition and cross-section scenarios and compare the outcomes to the standard approach. We also study the effects of the $X_{\rm max}$ range selection and $X_{\rm max}$ scale uncertainty on the fit. Finally, we apply the method to data collected at the Pierre Auger Observatory and discuss the results.

T 19.5 Mon 17:20 T-H32

Measurement of the Energy Spectrum of UHECRs with the Fluorescence Detector of the Pierre Auger Observatory - •KATHRIN BISMARK for the Pierre Auger-Collaboration - Karlsruher Institut für Technologie, Karlsruhe, Germany

The origin of ultrahigh-energy cosmic-rays (UHECRs) is one of the unsolved mysteries of modern-day astrophysics. The flux of UHECRs at Earth provides an important constraint on the luminosity density of their sources and the features in the UHECR energy spectrum shed light on the properties of astrophysical accelerators and on the propagation of cosmic rays through extragalactic photon fields.

Combining the measurements of the surface (SD) and fluorescence detector (FD) of the Pierre Auger Observatory allows us to determine a high resolution hybrid energy spectrum. Due to the partially redundant measurement of air showers with FD and SD, most event selection criteria and environmental influences on detection capabilities and reconstruction parameters can be investigated using measured data instead of simulations.

This presentation will focus in particular on the condition-independent visibility range of the FD, the so-called fiducial distance, given by the trigger efficiency of the FD. This trigger efficiency can be measured by determining the conditional probability to trigger a fluorescence telescope given an air shower detected by SD. The results of this study are compared to predictions from detector simulations and their impact on improvement of the precision of the measured spectrum will be discussed.

T 19.6 Mon 17:35 T-H32

Effects of magnetic fields on anisotropies in the arrival direction of ultrahigh-energy cosmic rays — •LUCA DEVAL, RALPH ENGEL, THOMAS FITOUSSI, and MICHAEL UNGER - Karlsruhe Institute of Technology, Institut für Kernphysik, Karlsruhe, Germany

The source of ultra-high-energy cosmic rays (UHECRs) is still an open question in astrophysics. The latest analysis of the dataset from the Pierre Auger Observatory revealed presence of anisotropy in the arrival direction of UHECRs which is an indication of the signal contribution of nearby sources. A maximum likelihood analysis found a statistical significance of 4σ for the correlation of the measured arrival directions with a sample of nearby starbust galaxies (SBG). Although, the dependence of the galactic magnetic field (GMF), which is expected to have a key role in the arrival direction of charged particles, has not been considered.

In this work we present a study of the effects of the GMF on the arrival directions of particles related to different source populations, namely SBG and active galactic nuclei. We assume an injected cosmic ray spectrum with a mixed composition and a maximum rigidity. The extragalactic propagation is simulated with CRPropa3 while the deflections of cosmic rays in the Galaxy are calculated assuming the GMF model of Jansson&Farrar (2012). The obtained results show that it is possible to recover scenarios which are compatible with the results obtained by the Pierre Auger Collaboration although the signal fraction related to the source contribution has to be increased. Moreover no contribution of the extragalactic magnetic field is necessary.

T 19.7 Mon 17:50 T-H32

Measuring the muon content of inclined air showers using the radio and particle detector of the Pierre Auger Observatory* — •Marvin Gottowik for the Pierre Auger-Collaboration — Bergische Universität Wuppertal, Gaußstraße 20, 42119 Wuppertal

A first measurement of the muon content of an air shower using hybrid radio and particle detection is presented. For inclined air showers with zenith angles

Location: T-H32

above 60°, the water Cherenkov detector (WCD) of the Pierre Auger Observatory performs an almost pure measurement of the muonic component, whereas the Auger Engineering Radio Array (AERA) reconstructs the electromagnetic energy independently using the radio emission of the air shower. The analysis of more than 6 years of AERA data shows a deficit of muons predicted by all current-generation hadronic interaction models for energies between 4 EeV and 20 EeV. This deficit, already observed with the Auger Fluorescence Detector, is now confirmed using for the first time radio data. The analysis is limited by low statistics due to the small area of AERA and the high energy threshold originating from the WCD reconstruction. With the AugerPrime Radio Detector currently being deployed, this analysis can be extended to the highest energies to allow for in-depth tests of hadronic interaction models with large statistics.

* Gefördert durch die BMBF Verbundforschung Astroteilchenphysik (Vorhaben 05A20PX1).

T 19.8 Mon 18:05 T-H32 Optimization of Radio Reconstruction for Inclined Air Showers with AERA at the Pierre Auger Observatory* — •JELENA PETEREIT for the Pierre Auger-Collaboration — Bergische Universität Wuppertal, Gaußstraße 20, Wuppertal, Germany The Pierre Auger Observatory is the world's largest cosmic ray observatory. Its Auger Engineering Radio Array (AERA) consists of more than 150 antenna stations that cover an area of about 17 km² and is used to detect radio signals emitted by extensive air showers. These measurements are used to reconstruct properties of the primary cosmic rays inducing the air showers.

This talk describes the improvements that have been made on the AERA analysis with the Auger reconstruction framework. Using CoREAS simulations for measured event geometries, noise extracted from data can be added to a simulated pure signal. Various parameters for identifying noise dominated stations for the rejection in the geometry reconstruction, such as the time difference between the pure signal and the signal with noise, are examined and modified in order to improve the event reconstruction.

*Gefördert durch die BMBF Verbundforschung Astroteilchenphysik (Vorhaben 05A20PX1)

T 20: Neutrino Physics without Accelerators 1

Time: Monday 16:15-18:40

Group Report

JUNO physics potential and status — •ALEXANDRE GÖTTEL for the JUNO-Collaboration — Forschungszentrum Jülich GmbH, Nuclear Physics Institute IKP-2, Jülich, Germany — III. Physikalisches Institut B, RWTH Aachen University, Germany

The Jiangmen Underground Neutrino Observatory (JUNO) is a next-generation liquid scintillator experiment being built in the Guangdong province in China. JUNO's target mass of 20 kton will be contained in a 35.4 m acrylic vessel, itself submerged in a water pool, under about 700 m of granite overburden. Surrounding the acrylic vessel are 17612 20" PMTs and 25600 3" PMTs. The main goal of JUNO, whose construction is scheduled for completion in 2022, is a 3- 4σ determination of the neutrino mass ordering (MO) using reactor neutrinos within six years, as well as a sub-percent measurement of the oscillation parameters θ_{12} , Δm_{21}^2 , and Δm_{31}^2 . JUNO's large target mass, low background, and dual calorimetry, leading to an excellent energy resolution and low threshold, allows for a rich physics program with many applications - including solar-, geo-, and atmospheric neutrino measurements. JUNO will also be able to measure neutrinos from galactic core-collapse supernovae, detecting about 10,000 events for a supernova at 10 kpc within 10 s, and achieve a 3σ discovery of the diffuse supernova neutrino background in ten years. JUNO is also suited for exotic searches and can be expected to give a lower limit of 8.34e33 years (90% C.L.) on the proton lifetime. This group talk covers the rich neutrino and astrophysics potential of the JUNO experiments and gives an update on the current experimental status.

T 20.2 Mon 16:35 T-H33

T 20.1 Mon 16:15 T-H33

Studies on the sensitivity for the Neutrino Mass Ordering Measurement of JUNO — •NIKHIL MOHAN^{1,3}, ALEXANDRE GOETTEL^{2,3}, PHILIPP KAMPMANN¹, RUNXUAN LIU^{2,3}, LIVIA LUDHOVA^{2,3}, LUCA PELICCI^{2,3}, MARIAM RIFAI^{2,3}, APEKSHA SINGHAL^{2,3}, and CORNELIUS VOLLBRECHT^{2,3} — ¹GSI Helmholtz Centre for Heavy Ion Research, Darmstadt — ²Forschungszentrum Jülich GmbH, Nuclear Physics Institute IKP-2, Jülich — ³III. Physikalisches Institut B, RWTH Aachen University, Aachen

JUNO is a multipurpose 20 kton liquid scintillator detector under construction in a 700 m underground laboratory in China, positioned 53 km from Yangjiang and Taishan nuclear power plants. The central detector is being built with high photocathode coverage of 78%, provided by 17,612 20-inch PMTs (LPMTs) and 25,600 3-inch PMTs (SPMTs). The unprecedented expected energy resolution at $3\%/\sqrt{E[MeV]}$ and the large fiducial volume anticipated for the JUNO detector offers exciting opportunities for addressing many important topics in neutrino and astroparticle physics.

This talk will focus on the primary physics goal of JUNO, which is the determination of Neutrino Mass Ordering (NMO) via the measurement of the vacuum oscillation pattern of the reactor antineutrinos. JUNO will detect the antineutrinos of electron flavor via the Inverse Beta Decay (IBD) interaction with a 1.8 MeV energy threshold. The estimated sensitivity to the NMO is a 3-4 σ significance with at least six years of data taking.

T 20.3 Mon 16:50 T-H33 Analysis of possible implications by the finestructure in the reactor neutrino spectrum on the JUNO NMO sensitivity — •TOBIAS HEINZ, LUKAS BIEGER, DAVID BLUM, MARC BREISCH, SRIJAN DELAMPADY, JESSICA ECK, GINA GRÜ- NAUER, BENEDICT KAISER, FRIEDER KOHLER, TOBIAS LACHENMAIER, AXEL MÜLLER, TOBIAS STERR, ALEXANDER TIETZSCH, and JAN ZÜFLE — Eberhard Karls Universität Tübingen, Physikalisches Institut

The Jiangmen Underground Neutrino Observatory (JUNO) is a 20 kt liquid scintillator detector currently constructed in southern China with the main goal to determine the neutrino mass ordering (NMO). Therefore, JUNO measures the reactor neutrino spectrum from two nuclear power plants located in a distance of around 53 km. The precise knowledge of the emitted reactor neutrino spectrum is one of the major aspects for the NMO determination. In recent years, new calculations of the spectrum predicted the existence of a spectral finestructure which could impede the measurement with the unprecedented energy resolution of the JUNO detector.

This talk will discuss possibilities to study the implications of the still unknown finestructure in the reactor neutrino spectrum for the sensitivity of the mass hierarchy determination with JUNO. Further, some preliminary results of these sensitivity studies will be presented.

This work is supported by the Deutsche Forschungsgemeinschaft.

Group ReportT 20.4Mon 17:05T-H33The Taishan Antineutrino Observatory- •HANSTHEODOR JOSEF STEIGER- Cluster of Excellence PRISMA+, Staudingweg 9, 55128MainzJohannes Gutenberg Universität Mainz, Staudingerweg 7, 55128MainzDepartment, Technische Universität München, James-Franck-Str.1, 85748Garching, Germany

The TAO (Taishan Antineutrino Observatory) detector is aiming for a measurement of the reactor neutrino spectrum at very low distances (<30m) to the core with a groundbreaking resolution better than 2 % at 1 MeV. The TAO experiment will realize the unprecedented neutrino detection rate of about 2000 per day, which is approximately 30 times the rate in the JUNO main detector. In order to achieve its goals, TAO is relying on yet to be developed, cutting-edge technology, both in photosensor and liquid scintillator (LS) development which is expected to have an impact on future neutrino and Dark Matter detectors. In this talk TAO's design, physics prospects as well as the status of its construction will be presented, together with a short excursion into its rich R&D program with a special focus on the German contribution to the development of the novel gadolinium-loaded liquid scintillator. This work is supported by the Cluster of Excellence PRISMA+ at the Johannes Gutenberg University in Mainz and the DFG research unit JUNO.

T 20.5 Mon 17:25 T-H33

Event reconstruction for the neutrino mass ordering measurement of JUNO — •MARIAM RIFAI^{1,2}, LIVIA LUDHOVA^{1,2}, PHILIPP KAMPMANN³, LUCA PELICCI^{1,2}, APEKSHA SINGHAL^{1,2}, ALEXANDRE GOETTEL^{1,2}, CORNIL-IUS VOLLBRECHT^{1,2}, RUNXUAN LIU^{1,2}, and NIKHIL MOHAN^{2,3} — ¹Forschungszentrum Jülich GmbH, Nuclear Physics Institute IKP-2, Jülich, Germany — ²III. Physikalisches Institut B, RWTH Aachen University, Aachen, Germany — ³GSI Helmholtz Centre for Heavy Ion Research, Darmstadt, Germany

The Jiangmen Underground Neutrino Observatory (JUNO) is a multipurpose liquid scintillator-based neutrino experiment with a target mass of 20 kt. The detector is currently under construction and plans to start the data-taking in 2023. Its main goal is the determination of the neutrino mass ordering (MO),

Location: T-H33

through a measurement of the oscillation pattern of reactor neutrinos over 53 km baseline. For a successful measurement of MO with at least 3\$\sigma\$ in 6 years, the energy resolution of JUNO must reach an unprecedented 3% at 1 MeV, which is challenging in terms of event reconstruction. Moreover, future JUNO results about neutrino MO could be further improved via a combined analysis with atmospheric neutrinos, which can be observed and reconstructed in JUNO. To achieve this target performance, a precise knowledge of the detector's energy scale has been studied and event reconstruction methods based on charge and time information of the PMTs will be presented in this talk.

T 20.6 Mon 17:40 T-H33 Machine learning based reconstruction of atmospheric neutrino events in

JUNO — • ROSMARIE WIRTH — Hamburg University, Hamburg, Germany The Jiangmen Underground Neutrino Observatory (JUNO) is a 20 kt liquid scintillation detector. By observing reactor anti-neutrinos with a 53 km baseline, JUNO aims to determine the mass hierarchy with 3 σ significance.

Due to JUNOs large volume, it could be suitable to measure atmospheric neutrino events with high precision. In that case, this channel could deliver further measurements on the mass ordering and the atmospheric oscillation parameters. To obtain this goal sufficient reconstruction methods are needed. This talk presents machine learning based reconstruction methods to analyze these atmospheric neutrino events at JUNO.

T 20.7 Mon 17:55 T-H33 Event Reconstruction in JUNO-TAO using Deep Learing - • VIDHYA THARA HARIHARAN — University of Hamburg

The primary goal of JUNO is to resolve the neutrino mass hierarchy using precision spectral measurements of reactor antineutrino oscillations. To achieve this goal a precise knowledge of the unoscillated reactor spectrum is required in order to constrain its fine structure. To account for this, Taishan Antineutrino Observatory (TAO), a ton-level, high energy resolution liquid scintillator detector with a baseline of about 30 m, is set up as a reference detector to JUNO. The 20% increase in the coverage of photosensors, the replacement of Photomultiplier Tubes (PMTs) with Silicon Photomultiplier (SiPM) tiles, the smaller dimension and the low temperature at -51°C, would enable TAO to achieve an yield of 4,500 p.e./MeV. Consequently TAO will achieve an energy resolution of 1.5%/E(MeV). The measurement of the reactor antineutrino spectrum with this energy resolution will provide a model-independent reference spectrum for IUNO.

The reconstruction can be performed using several approaches. However previous studies have proved Deep Learning yields competitive reconstruction results. Hence this work aims at demonstrating the general applicability of Graph neural networks (GNNs) to reconstruct vertex and energy and later at studying the directionality of TAO events.

T 20.8 Mon 18:10 T-H33 Search for the DSNB in JUNO: Development of new Methods for Background Event Identification — •MATTHIAS MAYER¹, LOTHAR OBERAUER¹, Raphael Stock¹, Hans Steiger², Konstantin Schweizer¹, Ulrike Fahrendholz¹, David Dörflinger¹, Sebastian Zwickel¹, Simon Appel¹, CARSTEN DITTRICH¹, VINCENT ROMPEL¹, LUCA SCHWEIZER¹, KORBINIAN STANGLER¹, and SIMON CSAKLI¹ for the JUNO-Collaboration -¹Technische Universität München, München, Germany – ²Institute of Physics and EC PRISMA+, Johannes Gutenberg Universität Mainz, Mainz, Germany

The diffuse supernova neutrino background (DSNB) is a constant, diffuse flux of relic neutrinos from past core-collapse supernovae over the entire visible universe. The upcoming Jiangmen Underground Neutrino Observatory (JUNO), a 20 kton liquid scintillator detector, expects to observe the DSNB through the inverse beta decay (IBD) detection channel. Besides IBD background from other electron anti-neutrino sources, there are also neutron-induced background events and NC interactions of atmospheric neutrinos of all flavours. This non-IBD background can be discriminated using pulse shape discrimination (PSD) methods. In this talk, I investigate the possibility to increase the fiducial volume available for the DSNB search using machine learning methods. Further, this talk discusses the effects of an electronics simulation and the flourescence parameter choice on the achievable PSD performance. This work is supported by the DFG research unit "JUNO", the DFG collaborative research centre 1258 "NDM", and the DFG Cluster of Excellence "Origins".

T 20.9 Mon 18:25 T-H33

Indirect dark matter search with neutrinos in JUNO and THEIA - LUKAS BIEGER, •DAVID BLUM, MARC BREISCH, SRIJAN DELAMPADY, JESSICA ECK, GINA GRÜNAUER, TOBIAS HEINZ, BENEDICT KAISER, FRIEDER KOHLER, TOBIAS LACHENMAIER, AXEL MÜLLER, TOBIAS STERR, ALEXANDER TIETZSCH, and JAN ZÜFLE - Eberhard Karls Universität, Physikalisches Institut, Tübingen, Germany

Neutrino detectors like the Jiangmen Underground Neutrino Observatory (JUNO) and the prospective neutrino detector THEIA are sensitive to a potential neutrino flux produced by dark matter self-annihilation in the Milky Way. The expected neutrino signals from dark matter self-annihilation and the relevant backgrounds in the energy range from 10 MeV to 100 MeV are investigated for both neutrino detectors. Further background suppression is realized by pulse shape discrimination analysis in JUNO and by studying the ratio between Cherenkov and scintillation light in THEIA. Results of a sensitivity study of JUNO and THEIA on the dark matter self-annihilation cross section are presented in this talk. This work is supported by the Deutsche Forschungsgemeinschaft.

T 21: Neutrino Physics without Accelerators 2

Time: Monday 16:15-18:35

Group Report

T 21.1 Mon 16:15 T-H34 Sterile neutrino search at the keV mass scale with KATRIN - •FRANK Edzards for the KATRIN collaboration¹, Marco Carminati^{2,3}, David Fink¹, Carlo Fiorini^{2,3}, Matteo Gugiatti^{2,3}, Pietro King^{2,3}, and Peter $\label{eq:Lechner} \begin{array}{l} {\rm Lechner}^4 - {}^1{\rm Max} \mbox{ Planck Institute for Physics, Munich, Germany} - {}^2{\rm DEIB}, \\ {\rm Politecnico \ di \ Milano, \ Milano, \ Italy} - {}^3{\rm INFN}, \mbox{ Sezione \ di \ Milano, \ Milano, \ Italy} \end{array}$ —⁴Halbleiterlabor der Max Planck Gesellschaft, Munich, Germany

Sterile neutrinos are a natural extension of the Standard Model of particle physics. If their mass is in the keV range, they are a viable dark matter candidate. One way to search for sterile neutrinos in a laboratory-based experiment is via tritium beta decay. A sterile neutrino with a mass up to 18.6\,keV would manifest itself in the decay spectrum as a kink-like distortion. The objective of the TRISTAN project is to extend the KATRIN experiment with a novel multi-pixel silicon drift detector and readout system to search for a keV-scale sterile neutrino signal. This talk will give an overview on the current status of the project. First characterization measurement results obtained with a 166 pixel system will be shown. This work is supported by BMBF (05A17PM3, 05A17PX3, 05A17VK2, 05A17WO3), KSETA, the Max Planck society, and the Helmholtz Association.

T 21.2 Mon 16:35 T-H34

Shifted Analyzing Plane: Optimizing spectrometer potentials and fields to reduce background in KATRIN — •BENEDIKT BIERINGER for the KATRIN-Collaboration — Institut für Kernphysik, Uni Münster, Germany

The KArlsruhe TRItium Neutrino (KATRIN) experiment aims at measuring the electron antineutrino mass to an estimated sensitivity of $0.2 \, \text{eV/c}^2$ at 90 % CL through spectroscopy of Tritium beta decay electrons using an electrostatic spectrometer of MAC-E filter type. For this level of precision, a low spectrometer background is required. The novel Shifted Analyzing Plane method achieves a significant reduction of this background through optimization of MAC-E filter electric potentials and magnetic fields. In this talk, computational and physical methods are presented that enabled the reduction of background of the KATRIN experiment by over a factor of two and fully eliminated measurable correlated background events from trapped high-energetic electrons. This includes a brief introduction to the inner workings of a MAC-E filter, a novel software collection to enable realtime field calculations based on Zonal Harmonic Field Expansion and background modelling for the largest ultra-high vacuum component in the KATRIN experiment.

This talk presents work funded via BMBF contract numbers 05A20VK3, 05A20PX3, 05A20PMA and 05A17WO3.

T 21.3 Mon 16:50 T-H34

Location: T-H34

Determination of Electromagnetic Fields in the Shifted Analyzing Plane of the KATRIN Main Spectrometer — •FABIAN BLOCK¹ and ALEXEY LOKHOV² for the KATRIN-Collaboration -¹Karlsruhe Institute of Technology -²WWU Münster

The KATRIN experiment aims to determine the effective electron antineutrino mass with a sensitivity of 0.2 eV (90 % C.L.) by high-resolution spectroscopy of the endpoint region of the tritium β decay spectrum. To reach the sensitivity goal, the experimental setup of KATRIN combines a windowless gaseous tritium source with a high-resolution MAC-E filter, called main spectrometer. The energy analysis of the β -decay electrons in the main spectrometer takes place via a complex interplay of electric and magnetic fields.

To improve the signal-to-background ratio during neutrino mass measurements, the electromagnetic field configuration in the main spectrometer is adapted to the so-called Shifted Analyzing Plane (SAP). The SAP electromagnetic fields need to be known with high precision in order for them to be taken accurately into account in the β -spectrum model applied in the fit of the data. We present in this talk the results of SAP characterization measurements employing conversion electrons of krypton-83m as sensitive probes for the electromagnetic fields. *This work is supported by the Helmholtz Association (HGF), the Ministry for Education and Research BMBF (05A17PM3, 05A17PX3, 05A17VK2, 05A20PMA, and 05A17WO3), the Helmholtz Alliance for Astroparticle Physics (HAP), and the Helmholtz Young Investigator Group (VH-NG-1055).*

T 21.4 Mon 17:05 T-H34

aTEF: Background reduction at KATRIN via an active transverse energy filter — •SONJA SCHNEIDEWIND¹, KEVIN GAUDA¹, VOLKER HANNEN¹, ALEXEY LOKHOV¹, HANS-WERNER ORTJOHANN¹, WOLFRAM PERNICE², RICHARD SALOMON¹, MAIK STAPPERS², and CHRISTIAN WEINHEIMER¹ for the KATRIN-Collaboration — ¹Institut für Kernphysik, Universität Münster, Wilhelm-Klemm-Str. 9, 48149 Münster, Germany — ²Physikalisches Institut, Universität Münster, Heisenbergstr. 11, 48149 Münster, Germany

The Karlsruhe Tritium Neutrino Experiment (KATRIN) aims at the direct measurement of the electron antineutrino mass with 0.2 eV/c² sensitivity from precision spectroscopy of the tritium beta decay. The analysis of its first two science runs yields a new upper limit of $m_{\nu} < 0.8$ eV (90% C.L.). Even in the shifted-analysis-plane (SAP) mode it is required to further lower the background rate to reach the target sensitivity. The background rate is dominated by electrons originating from ionisation of highly-excited (Rydberg) atoms produced by α -decays in the spectrometer walls. Thus, they cannot be distinguished from the signal electrons by energy but they possess much smaller angles w.r.t. the beam axis and, thus, much smaller cyclotron radii in the magnetic fields of KATRIN. The aTEF idea is to construct a detector by microstructuring that is mainly sensitive to the signal electrons because of their larger cyclotron radii. Investigations of first prototypes based on microstructured silicon PIN detectors are presented in this talk. The work of the authors for KATRIN is supported by BMBF under contract number 05A20PMA.

T 21.5 Mon 17:20 T-H34

Electron tracking simulations for the active transverse energy filter at KATRIN — •RICHARD SALOMON, KEVIN GAUDA, SONJA SCHNEIDEWIND, VOLKER HANNEN, ALEXEY LOKHOV, HANS-WERNER ORTJOHANN, and CHRISTIAN WEINHEIMER for the KATRIN-Collaboration — Institut für Kernphysik, Wilhelm-Klemm-Straße 9, 48149 Münster, Germany

The Karlsruhe Tritium Neutrino Experiment (KATRIN) aims at determining the electron neutrino mass with a sensitivity of 0.2 eV/ c^2 from a precision measurement of the tritium β -decay spectrum. To reach the desired sensitivity it is crucial to minimize experimental background events especially in the endpoint region of the electron spectrum.

One of the dominant backgrounds identified is the ionization of highly-excited (Rydberg) atoms inside the main spectrometer. To mitigate this background source, the concept of an active transverse energy filter (aTEF) is being investigated. As the electrons emitted by ionized Rydberg atoms, in contrast to most signal electrons from tritium beta decay, possess only a small amount of energy perpendicular to the guiding magnetic field, an angular-selective detector might be able to distinguish between the two. In order to test this novel detection technique, prototypes consisting of microstructured Si-PIN diodes are currently investigated in a test setup in Münster. This talk focuses on the accompanying particle tracking simulations which are essential for the analysis and interpretation of measurement data.

This project is supported by BMBF under contract number 05A20PMA.

T 21.6 Mon 17:35 T-H34

Combined analysis of the first five KATRIN measurement campaigns with KaFit — •STEPHANIE HICKFORD and LEONARD KÖLLENBERGER for the KATRIN-Collaboration — Institute for Astroparticle Physics, Karlsruhe Institute of Technology

The KATRIN collaboration aims to determine the neutrino mass with a sensitivity of $0.2 \text{ eV}/c^2$ (90 % CL). This will be achieved by measuring the endpoint region of the tritium β -electron spectrum. Combined analysis of the first two KATRIN measurement campaigns yielded a neutrino mass limit of $m_{\gamma} \leq 0.8 \text{ eV}$ (90 % CL).

Analyses of data from the first five measurements campaigns are currently underway. One of the combined analyses is performed using the KaFit/SSC model within the KASPER software framework. In this analysis systematic uncertainties are propagated as additional fit parameters with constraints (the "pull term" method). An overview of the collected data and the expected combined sensitivity on the neutrino mass from these five measurement campaigns will be presented in this talk.

This work is supported by the Helmholtz Association (HGF), the Ministry for Education and Research BMBF (05A17PM3, 05A17PX3, 05A17VK2, and 05A17WO3), the Helmholtz Alliance for Astroparticle Physics (HAP), and the Helmholtz Young Investigator Group (VH-NG-1055).

T 21.7 Mon 17:50 T-H34

Status of the KATRIN neutrino mass analysis using Monte Carlo propagation and a novel neural network approach — CHRISTIAN KARL^{1,2}, SUSANNE MERTENS^{1,2}, •ALESSANDRO SCHWEMMER^{1,2}, and CHRISTOPH WIESINGER^{1,2} for the KATRIN-Collaboration — ¹Max-Planck-Institut für Physik, München — ²Physik Department, Technische Universität München, Garching

The Karlsruhe Tritium Neutrino (KATRIN) experiment probes the effective electron anti-neutrino mass by a precision measurement of the tritium beta-decay spectrum near the endpoint. A world-leading upper limit of 0.8 eV c^{-2} (90 % CL) has been set with the first two measurement campaigns. New operational conditions for an improved signal-to-background ratio, the reduction of systematic uncertainties and a substantial increase in statistics allow to expand this reach.

The performance figures of three additional datasets, analysed with the Monte Carlo propagation method, and an outlook on their combination using a novel neural network technique will be presented in this talk.

T 21.8 Mon 18:05 T-H34

Measurement of the drift time in a silicon drift detector for the KATRIN experiment by laser pulsing — •KORBINIAN URBAN¹, MARCO CARMINATI^{2,3}, DAVID FINK¹, CARLO FIORINI^{2,3}, MATTEO GUGIATTI^{2,3}, PIETRO KING^{2,3}, and PETER LECHNER⁴ for the KATRIN-Collaboration — ¹Max Planck Institute for Physics, Munich, Germany — ²DEIB, Politecnico di Milano, Milano, Italy — ³INFN, Sezione di Milano, Milano, Italy — ⁴Halbleiterlabor der Max Planck Gesellschaft, Munich, Germany

The KATRIN experiment investigates the endpoint of the tritium beta-decay spectrum to search for the effective mass of the electron neutrino. Furthermore, the KATRIN experiment has the potential to also search for the signature of a sterile neutrino in the keV-mass regime by measuring the entire tritium beta-decay spectrum with an upgraded detector system. The new detector system, named TRISTAN, will be a multi-pixel silicon drift detector. This technology provides an improved energy resolution at high rates compared to PIN detector diodes. The radial drift of a charge cloud to the small anode of each pixel with 3 mm radius can be a significant contribution to the time resolution of the detector. This tak presents a measurement where a pulsed red laser is used to characterize the drift time in a 7-pixel TRISTAN detector device.

This work is supported by BMBF (05A17PM3, 05A17PX3, 05A17VK2, and 05A17WO3), KSETA, the Max Planck society, and the Helmholtz Association.

T 21.9 Mon 18:20 T-H34

SQL database caching for calculating the response function of the KATRIN experiment in HPC environments — •JAN BEHRENS and FABIAN BLOCK for the KATRIN-Collaboration — Institut für Astroteilchen-/Experimentelle Teilchen-physik, KIT Karlsruhe, Hermann-von-Helmholtz-Platz 1, 76344 Eggenstein-Leopoldshafen

The Karlsruhe Tritium Neutrino experiment aims to determine the mass of the electron antineutrino with a sensitivity of 0.2 eV/ c^2 (90% C.L.). The measurement of the shape of the tritium beta-spectrum enables a model-independent investigation of the absolute neutrino mass scale. The setup consists of a 70 m long beam line that magnetically guides electrons from a gaseous, windowless tritium source through an electrostatic spectrometer of the MAC-E filter type. The neutrino mass analysis involves a time-consuming calculation of the response function that depends on various experimental parameters, such as the magnetic fields along the beam line or the source column density. In order to facilitate a fit over hundreds of data runs with varying conditions, a caching mechanism is implemented which operates on a SQL database that can be shared between multiple users. Using a central database allows to distribute the calculations in a HPC cluster environment in order to improve the efficiency of existing parallelization techniques. This work is supported by the Helmholtz Association (HGF), the Ministry for Education and Research BMBF (05A17PM3, 05A17PX3, 05A17VK2, and 05A17WO3), the Helmholtz Alliance for Astroparticle Physics (HAP), and the Helmholtz Young Investigator Group (VH-NG-1055).

T 22: Search for Dark Matter 1

Time: Monday 16:15-18:20

 Group Report
 T 22.1
 Mon 16:15
 T-H35

 Current status of the XENONnT Dark Matter Search Experiment — •LUISA

 HÖTZSCH — Max-Planck-Institut für Kernphysik, Heidelberg — on behalf of the

XENON Collaboration The XENON experiments are among the most sensitive dark matter (DM) detectors, utilizing the concept of dual-phase xenon time projection chambers (TPCs) for the direct detection of weakly interacting massive particles (WIMPs). The XENON1T detector, which utilized a total of 3.2 tonnes of xenon target, was able to set the most stringent limits on the WIMP-nucleon spin-independent cross section for WIMP masses above $6 \text{ GeV}/c^2$, with a minimum of $4.1 \times 10^{-47} \text{ cm}^2$ at $30 \text{ GeV}/c^2$.

The latest iteration in the XENON experiment series is the XENONnT detector, which is currently running at the INFN Gran Sasso National Laboratory in Italy. With a total of 8.4 tonnes of xenon, it is projected to improve the sensitivity to WIMP dark matter by another order of magnitude. In addition, due to its further background reduction by a factor of approximately 6, XENONnT is expected to be able to clarify the nature of an electronic recoil event excess observed in XENON1T.

In this talk, I will give an overview of the XENONnT detector and its subsystems, and present the current status of the experiment.

T 22.2 Mon 16:35 T-H35

Light signal correction for the XENONnT experiment — •JOHANNA JAKOB for the XENON-Collaboration — Institut für Kernphysik - WWU, Münster, Germany

XENONnT, the latest stage of the XENON dark matter project, is currently taking science data with the science goals to detect WIMP-nucleus scattering and to search for other rare events. The detector is a dual-phase time projection chamber (TPC) filled with 8.5 tonnes of liquid xenon. The detector side walls reflect scintillation light caused by energy deposition in the detector, which is registered at the top and bottom by photomultiplier arrays. Free electrons, additionally created by the energy deposition in the detector, are drifted to the gaseous phase at the top of the detector where they create a second scintillation light pulse by electroluminescence. Both the light as well as the charge signal allow to perform a 3-dimensional position reconstruction of the recorded events.

This talk focuses on the light signal reconstruction, which requires a correction of the position dependent light collection efficiency and the understanding of the effects of the non-uniform electric drift field. Based on calibration data with several internal sources, light collection efficiency maps are derived and applied to the light signals.

This work is supported by BMBF under contract 05A20PM1 and by DFG within the Research Training Group GRK-2149.

T 22.3 Mon 16:50 T-H35

Energy calibration of the XENONNT Experiment — •HENNING SCHULZE EISSING for the XENON-Collaboration — Institut für Kernphysik - WWU, Münster, Deutschland

The XENON Dark Matter Project uses a dual phase time projection chamber filled with liquid xenon to search for Dark Matter in the form of weakly interacting massive particles (WIMPs). The current iteration, the XENONnT experiment with 8.5 t of xenon, is taking science data and will also allow the investigation of other science topics due to its extremely low background especially for low energies.

The energy deposition as well as the three-dimensional location of an event in the detector is reconstructed using fast scintillation light signal and a delayed charge signal. The latter is converted into a light signal by electroluminescence in the gaseous xenon phase above the liquid. The size of the primary scintillation light and of the charge signal are anticorrelated. This talk will outline the energy calibration of the XENONnT experiment using several mono-energetic gamma sources that can be found in the background data as well as in dedicated calibration data using external and internal sources.

This work is supported by BMBF under contract 05A20PM1 und by DFG within the Research Training Group GRK-2149.

T 22.4 Mon 17:05 T-H35

Calibration of XENONNT with tagged neutrons in its TPC and water Cherenkov neutron veto — •DANIEL WENZ — Institut für Physik & Exzellenzcluster PRISMA+, Johannes Gutenberg-Universität Mainz, 55099 Mainz, Germany

For more than a decade, liquid xenon (LXe) time projection chambers (TPC) have been playing a key role in the direct search for WIMP dark matter and other rare events. In 2018, XENON1T set the most stringent limits for WIMP-nucleon couplings for masses above 6 GeV/c2. Building on this success, an even larger LXe TPC called XENONnT was built, which features a ~3-times larger target mass of 5.9 t. The experiment reuses much of the existing infrastructure of

XENON1T, which has been augmented by additional sub-systems. One of these new systems is a water Cherenkov neutron veto encapsulating the TPC cryostat. Neutrons are capable of mimicking WIMP signals by undergoing a single-scatter nuclear recoil inside the TPC and escaping the TPC cryostat. The neutron veto system has the goal to reduce the intrinsic nuclear-recoil background by serving as an active veto for those neutrons.

In 2021, XENONnT successfully finished commissioning and started science data taking. In this talk, we will present the first results of XENONnT*s nuclear recoil calibration as well as the calibration of the neutron-veto tagging efficiency, using coincident gammas and neutrons from an Americium-Beryllium source.

T 22.5 Mon 17:20 T-H35

Towards an automated krypton assay in xenon at the ppq level — \bullet ROBERT HAMMANN, HARDY SIMGEN, and STEFFEN FORM — Max-Planck Institut für Kernphysik, Heidelberg, Germany

The beta-decaying isotope ⁸⁵Kr is one of the main internal background sources of liquid xenon (LXe) detectors. With purification techniques, it is possible to reduce the concentration of krypton in xenon below 100 ppq (parts per quadrillion). To model the background contribution of the remaining ⁸⁵Kr, it is crucial for low-background experiments such as XENONnT to precisely quantify the concentration of krypton in the detector material. The rare gas mass spectrometer (RGMS) at MPIK Heidelberg can meet this requirement by measuring the krypton concentration of an extracted xenon sample in two steps: First, krypton is separated from the bulk of xenon using a cryogenic gas chromatographic system. Then, the amount of krypton is analyzed using a mass spectrometer. A fully automatic rare gas mass spectrometer (AutoRGMS) is envisaged for the krypton assay of future low-background LXe detectors. This instrument will considerably facilitate the time-consuming measurement procedure, thus enabling a more frequent krypton monitoring. In addition, larger adsorption traps with a novel adsorbent will be employed, which makes the system sensitive to lower krypton concentrations. A proof of concept was demonstrated with a test setup of an automated gas chromatographic system in which stable conditions were maintained for more than 10 hours during the separation process. Moreover, the setup was used to test individual components and to find a working point for AutoRGMS.

T 22.6 Mon 17:35 T-H35

Radon removal system of the XENONnT experiment — •DENNY SCHULTE, HENNING SCHULZE EISSING, PHILIPP SCHULTE, CHRISTIAN HUHMANN, and CHRISTIAN WEINHEIMER — WWU MÜNSTER

A novel high flux radon removal system has been built for the dark matter experiment XENONnT reducing the dominant electron recoil background produced by Rn-222 and its progenies. Continuous emanation from detector components and its half-life of 3.8 days leads to a homogenous distribution of the Rn-222 within the detector system before it decays. Our radon removal method is based on the vapor pressure difference of xenon and radon. We built a cryogenic distillation column with a throughput of 200 slpm to exchange the liquid xenon mass of 8.5 tonnes within one mean lifetime of Rn-222 in order to decrease the radon concentration by a factor 2. An additional extraction flow of 25 slpm from the xenon gas phase at the top of the XENONnT detector, where some specific radon sources were identified, demonstrated to provide an additonal radon reduction factor of nearly 2. Both reduction methods aim for reaching for the first time a radon activity concentration of 1 μ Bq/kg in a xenon-based dark matter experiment. To provide the enormous cooling power of more than 3 kW at about -100°C we use a heat-pump concept with custom-built, radon-free xenon compressors and heat exchangers.

This talk will focus on the principle, construction and commissioning measurements of the radon removal system.

The project is funded by BMBF under contract 05A20PM1.

T 22.7 Mon 17:50 T-H35

Coating techniques for radon mitigation in liquid xenon detectors — •MONA PIOTTER, HARDY SIMGEN, and FLORIAN JÖRG — Max-Planck-Institut für Kernphysik, Heidelberg

Searching for rare events like dark matter interaction or neutrinoless double beta decay using liquid xenon detectors, requires a low radon background. Radon, which is part of the uranium and thorium decay chain, can continuously emanate from the detector materials. Current attempts to lower the radon induced background include the selection of radio-pure materials, techniques allowing to actively remove radon from xenon, as well as data selection criteria. However, next generation experiments will require even lower radon levels which likely can not be achieved by employing those methods alone. A new technique to stop radon emanation based on surface coatings has been investigated. This requires a tight and radium-free layer. We have developed a electro-deposited copper coating and present here the promising results. During the development,

Location: T-H35

we systematically investigated the coating parameters with the short-lived ²²⁰Rn emanating from tungsten rods or stainless steel plates. After this preliminary tests we applied the coating to a suitable ²²²Rn emanating stainless steel source, which has a longer half-life. It was produced at the CERN facility ISOLDE by implanting ²²⁶Ra in stainless steel plates. In the talk we present the results of the first coating of that sample.

T 22.8 Mon 18:05 T-H35 **S1-based position reconstruction in dual phase time projection chambers** — •JARON GRIGAT — Albert-Ludwigs-Universität, Freiburg, Deutschland Most particle interactions inside liquid xenon dual phase time projection cham-

T 23: Experimental Techniques in Astroparticle Physics 1

Time: Monday 16:15-17:45

T 23.1 Mon 16:15 T-H36

Monte Carlo simulation of background components in low level Germanium spectrometry — •NICOLA ACKERMANN¹, HANNES BONET¹, CHRISTIAN BUCK¹, JANINA HAKENMÜLLER¹, GERD HEUSSER¹, MATTHIAS LAUBENSTEIN², MANFRED LINDNER¹, WERNER MANESCHG¹, JOCHEN SCHREINER¹, and HERBERT STRECKER¹ — ¹Max-Planck-Institut für Kernphysik, Saupfercheckweg 1, 69117 Heidelberg, Germany — ²Laboratori Nazionali del Gran Sasso, Via G. Acitelli 22, 67100 Assergi L'Aquila, Italy

This talk presents Monte Carlo simulations of the background spectra of the 4 screening detectors GeMPI 1 - 4 at the Gran Sasso Underground Laboratory (LNGS) using the Geant4 based framework MaGe. The GeMPI detectors are low background Ge spectrometers located at a depth of 3500 m.w.e. and achieve extremely high sensitivites in material screening at a level of μ Bq/kg. They are used to test material samples on their suitability to use in rare event experiments. In the simulations, muons, neutrons and tiny radioactive contaminations of the detector and shielding materials are investigated as possible sources of background radiation and it was found that the ²¹⁰Pb contaminations in the detector shield and the neutrons coming from radioactive decays in the surrounding rock are the biggest contributors. A detailed understanding of the composition of the background spectra was achieved, allowing for the suggestion of two new possible shield designs for future GeMPI-like detectors

T 23.2 Mon 16:30 T-H36

Biases in the ⁷⁶Ge 0νββ tagging from calibrations — •TOMMASO COMELLATO¹, MATTEO AGOSTINI^{1,2}, and STEFAN SCHÖNERT¹ — ¹Technical University of Munich, Garching bei München, Germany — ²University College London, London, United Kingdom

The analysis of the time profile of electrical signals in germanium detectors provides a powerful tool for a high efficiency selection of neutrinoless double beta decay $(0\nu\beta\beta)$ of ⁷⁶Ge. The standard discrimination techniques are calibrated using samples of $0\nu\beta\beta$ -like events, which either occur at a different energy or contain a significant background contamination. With the help of a ⁵⁶Co source (which was custom produced by the Jagiellonian University in Krakow), we present a precision measurement of the biases of the standard event selection techniques in $0\nu\beta\beta$ experiments with ⁷⁶Ge, and propose an additional calibration method. This work has been supported in part by the ERC (Grant agr. No. 786430 - GemX) and by the SFB1258 funded by the DFG.

T 23.3 Mon 16:45 T-H36

Towards a low background SDD for IAXO — DAVID CASADO MORAN^{1,2}, FRANK EDZARDS^{1,2}, THIBAUT HOUDY³, SUSANNE MERTENS^{1,2}, JUAN PABLO UL-LOA BETETA^{1,2}, •CHRISTOPH WIESINGER^{1,2}, and MICHAEL WILLERS^{1,2} — ¹Max-Planck-Institut für Physik, München — ²Physik-Department, Technische Universität München, Garching — ³IJCLab, Université Paris-Saclay, Paris

The International Axion Observatory (IAXO) is aiming to detect solar axions, as they are converted into X-rays in a strong magnet pointing towards the sun. Excellent spectroscopic performance, high X-ray absorption efficiency at and below 10 keV and potential for ultra-low background operations are features of Silicon Drift Detectors (SSDs) that could facilitate this search. First measurements in the Munich shallow underground laboratory have shown promising background performance. Dedicated low-background designs, following a conventional passive shielding strategy and a novel all-semiconductor active shield approach, are under development. In this talk, we will report on the latest achievement towards a low background SDD for IAXO.

T 23.4 Mon 17:00 T-H36 The Pacific Ocean Neutrino Experiment: Results from three years of pathfinder data — Christopher Fink, Kilian Holzapfel, Stephan Meighen-Berger, Imma Rea, Li Ruohan, •Lisa Schumacher, Maria Sharshunova, and Laura Winter for the P-ONE-Collaboration — Experimental Physics with Cosmic Particles, TU Munich bers (LXe-TPCs) create two light signals. Besides the prompt scintillation light (S1), electrons from the interaction site are drifted in an electric field to the gas phase of the TPC. There, they create a delayed proportional scintillation signal (S2). Normally, the position in the x-y-plane is reconstructed from the hit pattern of the S2 signal on the top photosensor array. The depth of the interaction can be calculated from the time delay between S1 and S2. In this talk, we explore the possibility to reconstruct the 3D position by only looking at the S1 signal using machine learning techniques. We discuss possible applications of this additional information and show how this method can help us to characterize the electric field inside the XENONNT TPC.

Location: T-H36

The Pacific Ocean Neutrino Experiment (P-ONE) is a planned, cubic-kilometerscale neutrino telescope at 2660 m depth in Cascadia Basin located off the coast of Vancouver, Canada. The telescope is planned by a growing collaboration of Ocean Networks Canada (ONC), and Universities from Germany, Canada and the USA. Two pathfinder experiments have been deployed in Cascadia Basin using the already available infrastructure of ONC: STRAW (STRings for Absorption length in Water) in 2018 and STRAW-b in 2020. Both pathfinder experiments - and P-ONE in the future - are based on vertical, deep-sea cable lines equipped with multiple photosensors and calibration light sources. The main goal of the pathfinder lines is to characterize the optical properties of the site, which proved suitable to host P-ONE. Another purpose of the pathfinder lines is the monitoring of background light caused by K40 decay and bioluminescence, by now over more than three years. We will present an overview over recent results obtained with the STRAW and STRAW-b lines.

T 23.5 Mon 17:15 T-H36

Performance evaluation of the Wavelength-shifting Optical Module for the IceCube Upgrade — •YURIY POPOVYCH¹, SEBASTIAN BÖSER¹, ANNA POLLMANN², JOHN RACK-HELLEIS¹, and MARTIN RONGEN¹ for the IceCube-Collaboration — ¹JGU Mainz — ²Bergische Universität Wuppertal

In the upcoming IceCube Upgrade several new types of sensors will be deployed so to increase the sensitivity and explore possibilities for the envisioned IceCube Gen2 detector.

One of the modules to be deployed in the Upgrade is the Wavelength-shifting Optical Module (WOM). It consists of a quartz tube coated with Wavelength-Shifting (WLS-) paint with two Photomultiplier Tubes (PMTs) attached to its ends located inside a quartz pressure vessel filled with PFPE. The paint absorbs UV-photons and reemits them as visible light which is then captured through total internal reflection and propagate to a PMT on each side. This design results in a large photosensitive area, UV-sensitivity and a high signal-to-noise ratio. Through various improvements, like the choice of the filling material or coating techniques the efficiency of the modules can be optimized.

This talk will present the optical design of the WOM and explain the contributions of the single WOM components to efficiency of the module. Further, several simulation tools will be presented used to study and optimize the overall performance.

T 23.6 Mon 17:30 T-H36

Timing characteristics of the Wavelength-shifting Optical Module — •JOHN RACK-HELLEIS¹, SEBASTIAN BÖSER¹, MARTIN RONGEN¹, KLAUS HELBING², ANNA POLLMANN², NICK SCHMEISSER², YURIY POPOVYCH¹, and KYRA MOSSEL¹ for the IceCube-Collaboration — ¹Johannes Gutenberg Universität Mainz — ²Bergische Universität Wuppertal

The Wavelength-shifting Optical Module (WOM) uses the techniques of wavelength shifting and light guiding to achieve a large photosensitive area, UVsensitivity and improved signal-to-noise ratio. The centerpiece of the sen- sor is a hollow quartz cylinder coated with wavelength-shifting paint with a PMT (Photomultiplier Tube) optically coupled to each of its ends. Incident photons are absorbed, wavelength shifted and re-emitted into the tube walls. From there, they are guided towards one of the read out PMTs via total inter- nal reflection. While effective area and signal-to-noise ratio scale approximately linearly with the cylinder length, the average time it takes photons to reach one of the readout PMTs also increases. The timing of the WOM can be described by a convolution of three main components: The time response of the attached read out PMT, the photoluminescence characteristics of the WLS paint, and the path length distribution of photons inside the WLS tube. In this presentation we elaborate on the understanding of the timing of the WOM from a theoretical and experimental stand point. We present the intricacies of a device where everything seemingly runs in circles.

T 24: Outreach Methods

Time: Monday 16:15-17:35

Group Report

T 24.1 Mon 16:15 T-H37

Netzwerk Teilchenwelt als Plattform für Outreach in der Teilchenphysik, Astroteilchenphysik sowie Hadronen- und Kernphysik — •UTA BILOW und MICHAEL KOBEL für die Netzwerk Teilchenwelt-Kollaboration — Technische Universität Dresden und 28 weitere Standorte

Forschende sind heute verstärkt gefordert, Einblick in ihre Arbeit zu geben und den Dialog mit der fachfremden Öffentlichkeit zu führen. Für die Physik der kleinsten Teilchen existiert mit dem Netzwerk Teilchenwelt eine einzigartige Struktur, in der sich bundesweit Forschungsgruppen aus 29 Instituten zusammengeschlossen haben, um ihre wissenschaftliche Arbeit einem breiten Publikum zugänglich zu machen. Netzwerk Teilchenwelt stellt etablierte Programme und Strukturen bereit, mit denen Jugendliche bei Projekttagen die faszinierende Forschung an Beschleunigern kennenlernen oder eigene Messungen mit Detektoren durchführen. Gleichzeitig werden junge Forscherinnen und Forscher zur Wissenschaftskommunikation motiviert und befähigt. Ein mobiles Modul, das durch Deutschland tourt, spricht weniger wissenschaftsaffine Zielgruppen an. Die Aktivitäten werden durch das BMBF-Projekt KONTAKT2 gefördert und ausgebaut. Für Lehrkräfte als wichtige Multiplikator:innen führt Netzwerk Teilchenwelt regelmäßig Fortbildungen im Programm Forschung trifft Schule durch und bietet ein breites Spektrum an Unterrichtsmaterial an. Der Vortrag stellt die Angebote im Netzwerk Teilchenwelt sowie Beteiligungsmöglichkeiten für interessierte Forscherinnen und Forscher vor.

T 24.2 Mon 16:35 T-H37

T 25.1 Mon 16:15 T-H38

Urknall unterwegs: eine mobile Ausstellung zur Teilchenphysik — UTA BILOW¹, •CHRISTIAN KLEIN-BÖSING², MICHAEL KOBEL¹, PHILIPP LINDENAU¹ und JOSEPH PIERGROSSI¹ für die Netzwerk Teilchenwelt-Kollaboration — ¹TU Dresden, Institut für Kern- und Teilchenphysik — ²WWU Münster, Institut für Kernphysik

Seit Juli 2021 tourt die mobile Ausstellung Urknall unterwegs durch Deutschland und macht Station in Fußgängerzonen und auf öffentlichen Plätzen. Die Ausstellung besteht aus einem sechs Meter langen begehbaren Tunnel, in dem Besucherinnen und Besucher eine Reise durch die Geschichte des Universums erleben. Zwei Infosäulen präsentieren auf unterhaltsame Art Wissenswertes zu Elementarteilchen und Wechselwirkungen, erläutern Forschungsmethoden in der Teilchenphysik und stellen anhand von Spin-Offs dar, wie die Teilchenphysik und ihre Technologien unser tägliches Leben beeinflussen. Ein Pavillon mit Spielen komplettiert die Schau. Mithilfe des eigenen Smartphones können Besucherinnen und Besucher die Ausstellung interaktiv erkunden. Urknall unterwegs kann in kurzer Zeit aufgebaut werden, da Tunnel und Säulen aufblasbar sind.

Die Ausstellung hat zum Ziel, Menschen aller Altersgruppen zu erreichen, die eher weniger Berührung mit Wissenschaft haben. Sie entstand im BMBFgeförderten Vorhaben KONTAKT in einer Zusammenarbeit von Weltmaschine und Netzwerk Teilchenwelt. Der Vortrag stellt vor, welche Erfahrungen mit der Ausstellung gesammelt wurden und welche Möglichkeiten zur Nutzung bestehen.

Monday

Location: T-H37

T 24.3 Mon 16:50 T-H37

I am a Scientist als niedrigschwelliger Beitrag zur Wissenschaftskommunikation — UTA BILOW und •LYDIA DÖRING für die Netzwerk Teilchenwelt-Kollaboration — Technische Universität Dresden

Wissenschaft im Dialog gGmbH hat in Kooperation mit Netzwerk Teilchenwelt den Austausch zwischen Jugendlichen und Forschenden auf niedrigschwellige Art ermöglicht: Innerhalb von 30-minütigen Chats konnten Schulklassen ihre Fragen rund um die Welt der kleinsten Teilchen beantwortet bekommen. 59 Schulklassen und 25 Forschende hatten sich für die zweiwöchige Aktion im November 2021 registriert. Auch auf der Website I am a Scientist konnten die Schülerinnen und Schüler Fragen stellen. Diese und die entsprechenden Antworten, die neben physikalischen Fachfragen auch den beruflichen Alltag und den Werdegang der Wissenschaftlerinnen und Wissenschaftler betreffen, sind für die Öffentlichkeit auf der Website nachlesbar. Zum Abschluss der Aktion wurde als Lieblingswissenschaftler Dominik Koll gekürt. Er ist als Kernphysiker am Helmholtz-Zentrum Dresden-Rossendorf tätig und promoviert an der TU Dresden und in Australien. Er hat besonders engagiert die Fragen beantwortet und den Wert von Grundlagenforschung herausgestellt. Um fake news zu entgegnen, ist es wichtig zu wissen, wie wissenschaftliches Arbeiten funktioniert. Mit dem verliehenen Preisgeld möchte Koll Schulen besuchen, über wissenschaftliches Arbeiten aufklären und für die Forschung begeistern. Der Vortrag stellt die Aktion vor und begeistert für Wissenschaftskommunikation.

T 24.4 Mon 17:05 T-H37

Implementation of a Portal Dedicated to Higgs Bosons for Experts and the General Public — IVAN DEMCHENKO, MARTIN KUPKA, ANDRÉ SOPCZAK, ANTOINE VAUTERIN, and •PETER ZACIK — CTU in Prague

The implementation of a web portal dedicated to Higgs boson research is presented. A database is created with more than 1000 relevant articles using CERN Document Server API and web scraping methods. The database is automatically updated when new results on the Higgs boson become available. Using natural language processing, the articles are categorised according to properties of the Higgs boson and other criteria. The process of designing and implementing the Higgs Boson Portal (HBP) is described in detail. The components of the HBP are deployed to CERN Web Services using the OpenShift cloud platform. The web portal is operational and freely accessible on http://cern.ch/higgs.

T 24.5 Mon 17:20 T-H37

Location: T-H38

Management of conference presentations in CMS — •ARND MEYER — III. Physikalisches Institut A, RWTH Aachen, Germany

While presentations of the scientific output to the community in conferences and workshops constitute a major duty of any collaboration, large collaborations face the issue of ensuring the highest quality, a proper recognition of the work done by members, and an adequate representation of all the contributing bodies and institutions. In this talk, the management of conference presentations by the CMS collaboration as well as a statistical analysis over the past 14 years are summarized.

T 25: Data Analysis, Information Technology and Artificial Intelligence

Time: Monday 16:15-18:25

Group Report

CROWN - A Software framework for fast analysis ntuple production — •SEBASTIAN BROMMER¹, MARKUS KLUTE¹, NIKITA SHADSKIY¹, GUENTER QUAST¹, ROGER WOLF¹, and SEBASTIAN WOZNIEWSKI² — ¹Karlsruhe Institute of Technology — ²Universität Göttingen

With the ever-increasing data recorded by the LHC experiments, new software solutions are required to handle the rising demand in computational power and to ensure fast and efficient processing of the data. The CROWN framework is designed to provide such a fast and robust solution for the conversion of stuctured event data into flat ntuples for histogramming and further analysis. Within the framework, code generation is used to create compiled C++ executables based on ROOT data frames, ensuring fast data processing with minimal dependencies.

In this talk, the core concepts of the framework as well as performance comparisons to existing solutions are presented.

Group ReportT 25.2Mon 16:35T-H38Verringerung systematischer Unsicherheiten durch systematics-aware training — MARKUS KLUTE, GÜNTER QUAST, •LARS SOWA, ROGER WOLF und STEFAN WUNSCH — Karlsruhe Institute of Technology (KIT)

Eine Aufgabe für Analysen in der Hochenergiephysik besteht in der Trennung von Signal und Untergrundereignissen. Durch statistische Anpassung an Da-

ten werden mit Hilfe dieser Trennung Fitparameter und deren Unsicherheiten, die die Genauigkeit der Fitparameter quantifizieren, bestimmt. Um statistische Unsicherheiten dieser Fitparameter zu minimieren, nehmen moderne Teilchenbeschleuniger enorme Datenmengen auf. Infolgedessen treten systematische Unsicherheiten verstärkt in den Vordergrund und Methoden zu deren Unterdrückung gewinnen für Analysen zunehmend an Wichtigkeit.

Dieser Vortrag präsentiert Studien zur Verringerung systematischer Unsicherheiten. Dabei wird eine diagnostische, auf Taylorkoeffizienten basierende Methode verwendet, um den Einfluss systematischer Variationen der Eingangsparameter auf die Ausgabefunktion eines Neuronalen Netzes zu untersuchen. Darauf aufbauend werden erprobte Methoden für systematics-aware training erläutert und vielversprechende, zur Umsetzung geplante Methoden vorgestellt.

T 25.3 Mon 16:55 T-H38

Understanding Event-Generation Networks via Uncertainties — MARCO BELLAGENTE¹, MANUEL HAUSSMANN², •MICHEL LUCHMANN³, and MICHEL LUCHMANN⁴ — ¹Institut für Theoretische Physik, Universität Heidelberg, Germany — ²Heidelberg Collaboratory for Image Processing, Universität Heidelberg, Germany — ⁴Institut für Theoretische Physik, Universität Heidelberg, Germany — ⁶Institut für Theoretische Physik, Universität Heidelberg, Germany — Following the growing success of generative neural networks in LHC simula-

104

tions, the crucial question is how to control the networks and assign uncertainties to their event output. We show how Bayesian normalizing flows or invertible networks capture uncertainties from the training and turn them into an uncertainty on the event weight. Fundamentally, the interplay between density and uncertainty estimates indicates that these networks learn functions in analogy to parameter fits rather than binned event counts.

T 25.4 Mon 17:10 T-H38

Evaluating Uncertainties in Measurements of the Production of a Single Top-Quark in Association with a Photon with Bayesian Neural Networks — Jo-HANNES ERDMANN¹, BURIM RAMOSAJ², and •DANIEL WALL¹ — ¹TU Dortmund University, Department of Physics — ²TU Dortmund University, Department of Statistics

Multivariate approaches including neural networks constitute powerful and established methods in experimental particle physics. However, using these methods, it is difficult to account for uncertainties from statistical and systematic sources in a consistent and efficient way. By employing weight distributions instead of fixed weights and by utilising the process of Bayesian inference, Bayesian Neural Networks not only suffer significantly less from overfitting, but also allow to obtain an uncertainty estimate on the output.

These characteristics are of particular interest in measurements of processes suffering from limited statistics and challenging signal-to-background ratios. The analysis of top-quark production in association with a photon $(tq\gamma)$, probing the structure of the electroweak couplings of the top quark, is one of such processes, as the corresponding cross section is considerably lower than those of relevant background processes, most importantly top-quark pair production $(t\bar{t}\gamma)$.

In this talk, studies of Bayesian Neural Networks for their application in the classification of top-quark processes in association with a photon are presented.

T 25.5 Mon 17:25 T-H38 **Non-parametric background models for axion haloscopes** – •JOHANNES DIEHL¹, JAKOB KNOLLMÜLLER², and OLIVER SCHULZ¹ – ¹Max Planck Institute for Physics, Munich, Germany – ²Max Planck Institute for Astrophysics,

Munich, Germany Axions have been introduced to solve the strong CP problem of the standard model of particle physics and turned out to be an excellent candidate to explain cold dark matter. "Haloscopes" are searching world wide for axions from the galactic dark matter halo, mostly by axion conversion to photons at radio frequencies in a strong B-field. Finding an axion signal in haloscope data means finding a small peak in a vast non-uniform RF background. One crucial challenge is therefore to selectively suppress larger frequency scales while inducing as little attenuation and correlation as possible at smaller frequency scales. This has so far been tackled using filter theory, e.g. through Savitzky-Golay filters for the HAYSTAC experiment, but proof that this is the optimal filter to use is still lacking. Using simulated data from the MADMAX haloscope, I present a novel machine-learning based approach to separate scales and subtract the background without attenuating the signal which lends itself well to being incorporated into a final Bayesian analysis.

T 25.6 Mon 17:40 T-H38

Open Science in KM3NeT – •JUTTA SCHNABEL for the ANTARES-KM3NET-ERLANGEN-Collaboration — Erlangen Centre for Astroparticle Physics, FAU Erlangen-Nürnberg

The KM3NeT neutrino detectors are currently under construction at two locations in the Mediterranean Sea, with a first taking of data from high-energy neutrino interactions already under way. This scientific data is valuable both for the astrophysics and neutrino physics communities as well as for marine biologists. In order to facilitate FAIR data sharing of the research results, the KM3NeT collaboration is actively working towards an open science infrastructure to provide high-level scientific data, software, and analysis pipelines in an interoperable research environment suited both for research and education. This contribution introduces the open science program of KM3NeT and gives an overview of its current architecture and implementation.

 $T\ 25.7\ Mon\ 17:55\ T-H38$ Columnar data analysis with ATLAS analysis formats — •Nikolai Hart-Mann — Ludwig-Maximilians-Universität München

Future analysis of ATLAS data will involve new small-sized analysis formats to cope with the increased storage needs. The smallest of these, named DAOD_PHYSLITE, has calibrations already applied to allow fast downstream analysis and avoid the need for further analysis-specific intermediate formats. This allows for application of the "columnar analysis" paradigm where operations are applied on a per-array instead of a per-event basis. This presentation shows the latest developments of tools within the scientific python ecosystem and discusses a prototype analysis for testing both on single Machines as well as Analysis Facilities or similar scale-out systems.

T 25.8 Mon 18:10 T-H38 Information visualization platform for data quality monitoring of CMS tracker — • ABHIT PATIL — Ruprecht-Karls-Universität Heidelberg, Germany The tracker of the CMS detector consists of silicon sensors arranged in concentric cylinders and endcap disks to track muons, which requires continuous monitoring during operation and certification of the recorded data for physics analysis. The process relies on shifters who assess the data quality by comparing data distributions with references. This challenging task requires examining possible types of failures with expert-based rule systems and manual profiling of a large number of histograms. To assess the quality of data volumes with finer granularity and to improve the quality of the data certification, this work proposes to augment the monitoring process with information visualization based methods, which aims to pre-process large amount of multidimensional data during the data taking period and provide a visual abstraction of the data quality and provide hints for potential anomalies. The visualisation methods are deployed on a platform built using Python-Django framework and Postgres database.

T 26: Data Analysis, Information Technology and Artificial Intelligence

Time: Monday 16:15-18:30

T 26.1 Mon 16:15 T-H39

Investigation of robustness of b-Tagging algorithms for the CMS Experiment — Xavier Coubez^{1,2}, Nikolas Frediani¹, Spandan Mondal¹, Andrzej Novak¹, Alexander Schmidt¹, and •Annika Stein¹ — ¹RWTH Aachen University, Germany — ²Brown University, USA

Deep learning as one form of machine learning is utilized for various applications and shows its benefits also in the field of high-energy physics, or more specifically, for jet flavour tagging. However, subtle mismodelings in the simulation could be invisible to typical validation methods. Investigating the response to mismodeled input data is motivated by the later usage of the outputs in physics analyses, as the values for simulation and data are deviating. The vulnerability of b-tagging algorithms used at the CMS experiment is probed through application of adversarial attacks. In this talk, a corresponding defense strategy that improves the robustness, namely adversarial training, will be presented. Comparisons of the model performance and the susceptibility show that this method constitutes a promising candidate to reduce the vulnerability and that this could improve the capability to generalize to data.

T 26.2 Mon 16:30 T-H39

Performance Studies of the Conditional Attention Deep Sets *b***-Tagger for the ATLAS Experiment** — •ALEXANDER FROCH¹, MANUEL GUTH², and ANDREA KNUE¹ — ¹Albert-Ludwigs-Universität Freiburg — ²Université de Genève The identification of jets containing *b*-hadrons, called *b*-tagging, is crucial for most analyses performed at the ATLAS experiment. Several new multivariate

techniques have been developed for this purpose. One of these is the Deep-Impact-Parameter-Sets (DIPS) tagger.

The DIPS tagger is a deep neural network based on the Deep-Sets architecture. It uses track information of the particles inside the clustered jets for classification. It is part of a new generation of tagging algorithms currently developed in AT-LAS. DIPS itself can distinguish between different jet origins, like light, charm or bottom jets.

Although DIPS already outperforms the currently recommended RNNIP tagger, its high- p_T performance can still be improved further. While the number of fragmentation tracks increases rapidly with p_T , less heavy-flavour tracks are being reconstructed at high p_T . Therefore, it is more difficult for this kind of network to filter the most important tracks.

To further enhance the tagging capabilities of DIPS and fix the issues arising in the higher p_T region, DIPS will be extended with an attention mechanism conditioned on jet kinematics. This new version is the Conditional Attention Deep Sets (CADS) tagger.

The new CADS tagger will be discussed and its performance will be compared to the current best DIPS model.

T 26.3 Mon 16:45 T-H39

Location: T-H39

Charm tagger shape calibration for BDT-based signal-background discrimination — •Spandan Mondal¹, Xavier Coubez^{1,2}, Alena Dodonova¹, Ming-Yan Lee¹, Luca Mastrolorenzo¹, Andrzej Novak¹, Andrey Pozdnyakov¹, Alexander Schmidt¹, and Annika Stein¹ — ¹RWTH Aachen University — ²Brown University, USA

Identification of charm-quark-initiated jets at the LHC is especially challenging. Usage of deep learning based algorithms have enabled several CMS analyses to efficiently discriminate charm jets simultaneously from bottom and light jets. The charm probability scores yielded by such charm tagging algorithms can play a powerful role when used as inputs to a machine learning based algorithm for discrimination between signal and background. However, as jet identification algorithms are trained strictly on simulated jets, a direct usage of charm tagger output values requires calibrating the entire output probability distributions using real jets reconstructed from CMS data. This talk focuses on the calibration of the output discriminator values of charm-tagging algorithms using flavourenriched selections of jets in data. Additionally, the improvement resulting from a shape calibration approach, over the traditional approach of calibrating efficiencies at fixed c-tagger working points, is exemplified in the context of the resolved VHcc analysis.

T 26.4 Mon 17:00 T-H39 Reduction of the Irreducible Background in the $t\bar{t}H(b\bar{b})$ Analysis at ATLAS, Using a Deep-Sets-Based bb-Tagger — • JOSCHKA BIRK, ALEXANDER FROCH, and ANDREA KNUE — Albert-Ludwigs-Universität Freiburg

The search for the $t\bar{t}H(b\bar{b})$ signal suffers from the large irreducible $t\bar{t} + b\bar{b}$ background. This irreducible background contains the same final-state particles as the signal, including four b quarks. In the background process, a radiated gluon can split into a *b*-quark pair, which often leads to two *b*-jets that are very close together. With the currently used *b*-tagging algorithm, these $b\bar{b}$ -jets are often misidentified as a single *b*-jet.

In order to improve the rejection of these background events, the existing Deep-Impact-Parameter-Sets (DIPS) Tagger is extended with an additional output class dedicated to jets which contain two *b*-hadrons ($b\bar{b}$ -jets). DIPS is part of a new ATLAS b-tagging algorithm, based on the Deep Sets architecture, and has already shown promising performance compared to the RNNIP tagger, which is part of the DL1r tagger that is currently used in ATLAS analyses. Studies of this extended version of the DIPS tagger, including first results of a hyperparameter optimisation, are presented.

T 26.5 Mon 17:15 T-H39

High- $p_{\rm T}$ b-tagging using track classification in the ATLAS experiment - BEATRICE CERVATO, MARKUS CRISTINZIANI, GABRIEL GOMES, VADIM KOSTYUKHIN, •KATHARINA VOSS, and WOLFGANG WALKOWIAK — Center for Particle Physics Siegen, Experimentelle Teilchenphysik, Universität Siegen

In the ATLAS experiment the search for new physics beyond the Standard Model is of particular interest. In many theories, new physics processes result in the production of energetic *b*-quarks, fragmenting to high- $p_{\rm T}$ jets in the detector. Many successful *b*-jet identification algorithms suffer from the jet energy increase due to higher multiplicity of fragmentation tracks, reduced track reconstruction efficiency and, looking ahead to the High-Luminosity LHC, increased pile-up. As b-tagging algorithms typically use properties of the track-in-jet ensemble, the increased multiplicity of fragmentation tracks in energetic jets inevitably decreases the *b*-tagging efficiency at high $p_{\rm T}$. A *b*-tagging algorithm, based on a multivariate technique, particularly suited for very energetic events, is presented. The multiplicity problem is solved by considering only those tracks, which are most likely to stem from a B-hadron decay. These are identified through a multi-class multivariate track classification algorithm, considering heavy flavour, fragmentation and other tracks from material interactions, as well as pile-up tracks.

T 26.6 Mon 17:30 T-H39

Exploration of neural network architectures for Flavour Tagging algorithms at the LHCb experiment — • VUKAN JEVTIC¹, QUENTIN FÜHRING¹, CHRISTOPH HASSE³, NIKLAS NOLTE², and CLAIRE PROUVÉ¹ – ¹Experimentelle Physik 5, TU Dortmund $-{}^{2}MIT - {}^{3}CERN$

The LHCb detector at the LHC is specialised for measurements of B meson decays, which open a window into the nature of weak interactions through measurements of rare decays and charge parity (CP) violation. In the Standard Model, CP violation is enabled through a complex phase of the Cabibbo-Kobayashi-Maskawa quark-mixing matrix. B meson mixing refers to the property of neutral B mesons to oscillate between two states of matter, B_a^0 and $\overline{B_a^0}_a$, with different quark contents (i.e. different flavours).

The reconstruction of the flavour at the time of the *B* meson production is a difficult but indispensable component of measurements of time-dependent CP violation at LHCb. In this talk new approaches to Flavour Tagging via full-eventinterpretation techniques will be presented by the example of recurrent neural networks and deep sets.

T 26.7 Mon 17:45 T-H39

Regression of Missing Transverse Momentum (MET) with Graph Neural Net $works - \bullet Nikita Shadskiy^{1}, Matteo Cremonesi^{2}, Jost von den Driesch^{1},$ LINDSEY GRAY³, ULRICH HUSEMANN¹, YIHUI LAI⁴, and MICHAEL WASSMER¹ — 1 Institute of Experimental Particle Physics (ETP), Karlsruhe Institute of Technology (KIT) - 2 University of Notre Dame - 3 Fermilab - 4 University of Maryland

Neutral particles that are only interacting weakly, like neutrinos, which are known from the standard model, or other, still unknown, particles in theories beyond the standard model, can be measured indirectly using the missing transverse momentum (MET). Analyses which search for specific invisible particles or expect such particles in their final state need well reconstructed MET. The reconstruction of MET is sensitive to e.g. experimental resolutions, mismeasurements of reconstructed particles or pileup interactions and is therefore a challenging task.

This talk will give an overview about a new approach to reconstruct MET with graph neural networks using information from particle flow (PF) candidates. Particle flow is an algorithm used by the CMS collaboration to reconstruct particles by combining information from different detector parts. Using graphs is a more intuitive way to describe the topology of an event because it has the advantage to be permutation invariant. Thus, the order of the PF candidates is irrelevant. The graph neural network is optimized to predict a weight for each PF candidate. These predictions are then used to weight the contribution of each PF candidate in the calculation of MET.

T 26.8 Mon 18:00 T-H39 Studies of machine learning inspired clustering algorithms for jets — AMRITA $Bhattacherjee^{1}, Debarghya\,Ghoshdastidar^{1}, \bullet Siddha\,Hill^{2}, and\,Stefan$ KLUTH² — ¹TUM Informatik — ²MPI für Physik

We study several machine learning inspired hierarchical clustering algorithms algorithms to cluster the particles of hadronic final states in high energy e+e- collisions into jets. We compare their performance against well known algorithms such as JADE or Durham. Performance indicators are physically motivated such as angular distance or energy difference of matching jets at parton, hadron or detector level. We also study new performance indicators derived from computer science clustering theory.

T 26.9 Mon 18:15 T-H39

Location: T-H15

Improvement of the Jet-Parton Assignment in $t\bar{t}H(b\bar{b})$ Events using Symmetry-Preserving Attention Networks - •Daniel Bahner, Andrea KNUE, and GREGOR HERTEN — Albert-Ludwigs-University, Freiburg, Germany The associated production of a Higgs boson and a top-quark pair allows for a direct measurement of the top-Higgs Yukawa coupling, which can be sensitive to Beyond Standard Model physics. In the studies presented, the process of interest is the semileptonic decay of the $t\bar{t}$ -pair accompanied by a $b\bar{b}$ -pair resulting from the most prominent Higgs decay. In this topology, at least four *b*-jets and two light jets are expected. This Higgs decay channel suffers from irreducible background due to $t\bar{t} + b\bar{b}$ production. Furthermore, the full reconstruction of this final state proves difficult because of the ambiguities in assigning the jets to their original parton, which is called combinatorial background.

In the latest publication, a Boosted Decision Tree was used for the jet-parton assignment. In the studies presented in this talk, a novel Symmetry-Preserving Attention Network is exploited (suggested in arXiv:2106.03898). The training was performed and evaluated on two different samples: In the first sample the full detector simulation with GEANT4 and the new ATLAS b-tagging algorithm DL1r was used and in the second sample the DELPHES framework was used. The performances of the networks and possible future improvements will be presented.

T 27: Invited Talks 2

Time: Tuesday 11:00-12:30

Invited Talk

T 27.1 Tue 11:00 T-H15 First Results From the Next Generation B-Factory Experiment Belle II -•Тномая Кинк — Ludwig-Maximilians-Universität München

The first generation B-factory experiments BaBar and Belle had successfully confirmed the theory of CP violation in the Standard Model, but it is known to be insufficient to explain the observed matter anti-matter asymmetry in the universe. To address this and further open questions more precise measurements are needed to find evidence for a more general theory.

The Belle II experiment at the SuperKEKB accelerator in Tsukuba, Japan is designed to collect 50 times more data than its predecessor. The increase in luminosity and backgrounds is a challenge for the detector and the data processing and analysis. The physics data run started in 2019.

T 27.3 Tue 12:00 T-H15

Location: T-H15

The status of the experiment, with focus on the Pixel Vertex Detector, and first physics results, highlighting the complementarity to the flavor physics program at the LHC, will be presented.

Invited Talk T 27.2 Tue 11:30 T-H15 Flavour Anomalies — •CHRISTOPH LANGENBRUCH — RWTH Aachen, Germanv

Precision measurements of observables in heavy flavour decays constitute powerful tests of the Standard Model of particle physics. New heavy particles beyond the Standard Model can significantly affect flavour observables through (virtual) quantum corrections. Precision measurements of these observables can reveal potential deviations from the Standard Model predictions and probe energy scales far beyond the beam energies presently available at colliders.

The talk will present recent precision measurements of flavour observables by the B-factories BaBar and Belle (II), and the LHC experiments ATLAS, CMS, and LHCb. Particular focus will be on the flavour anomalies, recent tensions between measurements and SM predictions in the flavour sector, and prospects for their clarification in the near future.

Invited Talk

The top quark is still going strong (and electroweak) — •ANDREA KNUE — Albert-Ludwigs-University Freiburg

The unique properties of the top quark have fascinated particle physicists since several decades. With its large mass, the top quark is expected to be connected to the mechanism of electroweak symmetry breaking. Moreover, the top quark is key to most research areas at the LHC: it can be measured very precisely, it is involved in rare processes that are finally accessible and it plays a special role in beyond-standard-model theories that are constantly being tested.

In this talk, the latest top-quark results will be discussed, revealing a rich research landscape that is thriving on the abundance of collision data available at the LHC.

T 28: Invited Topical Talks 1

Time: Tuesday 14:00-15:40

Invited Topical Talk T 28.1 Tue 14:00 T-H15 Hadronic Jets: Accuracy and Precision of their Reconstruction and Calibration in ATLAS — • CHRISTOPHER YOUNG — University of Freiburg, Freiburg im Breisgau, Germany

Hadronic jets are prolifically produced in LHC collisions such that their reconstruction is essential for understanding many different physics processes. This talk will detail how such jets are precisely reconstructed in the ATLAS detector using a particle flow algorithm developed for the high pile-up environment of Run 2 of the LHC and beyond. Additionally the accuracy of the calibration of the energy scale of hadronic jets is a leading source of experimental systematic uncertainty in many searches and measurements. The derivation of the latest calibration using data, and the accuracy and understanding achieved will also he covered.

Invited Topical Talk

T 28.2 Tue 14:25 T-H15 Direct searches testing BSM explanations of the flavor anomalies - •ARNE Сняізторн Reimers — Universität Zürich, Switzerland

Anomalies measured in the decay of B mesons have revealed first indications for the possible existence of lepton flavor universality violation. If confirmed by future measurements, the presence of such processes in nature would imply physics beyond the standard model.

In this presentation, LHC results of direct searches for new particles that are commonly proposed as an explanation for the flavor anomalies are reviewed. Particular emphasis is given to final states with large transverse momenta.

Invited Topical Talk T 28.3 Tue 14:50 T-H15 ATLAS probes QCD measuring photons — •HEBERTH TORRES — TU Dresden,

Germany The production of prompt isolated photons in proton-proton collisions is an im-

portant test of perturbative QCD prediction. Despite its electromagnetic nature, photon production at the LHC is affected by surprisingly large strong-interaction effects. Thanks to the precise ATLAS measurements of photon processes, the collaboration is able to probe these effects and scrutinise state-of-the-art theoretical calculations. In this talk, we present the latest measurements of prompt photon production using proton-proton collision data collected by the ATLAS experiment at $\sqrt{s} = 13$ TeV.

T 28.4 Tue 15:15 T-H15 Invited Topical Talk The upgrade of the ATLAS trigger to augment the physics reach of Run-3 -•DANIELE ZANZI — Albert-Ludwigs-Universität Freiburg, Freiburg, Germany The ATLAS experiment uses a two-level trigger system to record about 1 kHz out of the 40 MHz of collisions delivered by the LHC. After a very successful operation in Run-2, the ATLAS trigger system has gone through a radical upgrade both in hardware and in software in preparation for Run-3. This upgrade is critical for maintaining the high data-taking efficiency achieved in Run-2, while opening up new trigger opportunities to augment the ATLAS physics programme. This talk will present an overview of the ATLAS trigger system in Run-3 together with studies on its expected performance.

T 29: Invited Topical Talks 2

Time: Tuesday 14:00-15:40

Invited Topical Talk T 29.1 Tue 14:00 T-H16 Testing the Standard Model through Gauge-boson Self-interactions •PHILIP SOMMER — The University of Sheffield, Sheffield, United Kingdom; CERN, Geneva, Switzerland

In the Standard Model, electroweak interactions of fermions proceed through the exchange of gauge bosons, the W and Z bosons, and the photon. In addition, self-interactions of the gauge bosons through trilinear and quartic couplings are predicted. At the LHC, these lead to the production of single- and multiboson final states. Measurements of such processes, thus, provide a sensitive probe of the gauge structure of the electroweak theory. The pp collision data from the second experimental phase of the LHC has allowed for precise measurements of processes proceeding through trilinear couplings and facilitated, for the first time, the observation of a number of processes that proceed through quartic electroweak couplings. Recent measurements of single- and multi-boson production by the ATLAS collaboration are presented. The agreement of the measurements with the theoretical predictions is quantified by constraining Effective Field Theory operators that modify the trilinear and quartic gauge-boson self-interactions in a general extension of the Standard Model.

T 29.2 Tue 14:25 T-H16 **Invited Topical Talk** Axions and similar particles - how to cover 10^{17} orders of magnitude in mass

— •Kristof Schmieden — Johannes Gutenberg-Universität Mainz Axions, the famous hypothetical particle that explains the absence of CP violation in QCD, was already though of in the 70ies. Yet only in the past decade the hunt for this and similar particles took up pace, which huge advancements in the recent years. The reason behind the growing interest is the understanding that axions and axion-like particles can contribute to the dark matter content of the universe. In fact, pseudo-scalar particles are a natural prediction of many extensions of the Standard model and even possibly explain the muon (g-2) anomaly. Considering the general case of pseudo-scalar particles a huge parameter space in mass and coupling to SM particles opens up, requiring a variety of experimental approaches to hunt for these particles. This talk will briefly introduce the phenomenology of axions and axion-like particles and discuss a selection of experiments and their latest results in more detail.

Invited Topical Talk

T 29.3 Tue 14:50 T-H16

Location: T-H16

From GERDA to LEGEND - Hunting no neutrinos — •CHRISTOPH WIESINGER – Max-Planck-Institut für Physik, München — Physik-Department, Technische Universität München, Garching

Hidden by their tiny mass, neutrinos may carry a profound secret with farreaching consequences for both particle physics and cosmology. Given zero electric charge and no color, they may be Majorana particles - fermions that are their own anti-particles. Double beta decay offers a unique probe for this hypothesis. Finding no neutrinos, but solely two electrons sharing the full available decay energy, would prove lepton number non-conservation and reveal the Majorana character of neutrinos. The superb spectroscopic performance of high-purity germanium detectors provides exceptional discovery potential for this monoenergetic peak. The Germanium Detector Array (GERDA) experiment has operated 40 kg of enriched germanium in an instrumented low-background liquid argon environment. In a total exposure of more than 100 kg yr, taken under record-

107

low background conditions, no signal was found. The corresponding half-life limit for neutrinoless double beta decay of ⁷⁶Ge is >1.8·10²⁶ yr at 90% C.L., and coincides with the median sensitivity for the null hypothesis. The Large Enriched Germanium Experiment for Neutrinoless double beta Decay (LEGEND) is about to expand this search towards 10²⁸ yr and beyond. Following LEGEND-200, the initial 200-kg phase currently under construction, LEGEND-1000 will probe the full parameter space spanned by the inverted ordering scenario.

Invited Topical TalkT 29.4Tue 15:15T-H16Mapping Highly-Energetic Messengers throughout the Universe — • SARA BU-
SON — Institut für Theoretische Physik und Astrophysik, Universität Würzburg

T 30: Flavour Physics 3

Time: Tuesday 16:15-18:30

T 30.1 Tue 16:15 T-H15

Time-Dependent Charge-Parity Violation at Belle II — •CASPAR SCHMITT — Max-Planck Institute for Physics, Munich, Germany

Overconstraining the unitarity triangle of the Cabibbo-Kobayashi-Maskawa mixing matrix by precision measurements is an essential test for the description of weak currents in the standard model (SM) of particle physics. We thereby test our present understanding of charge-parity (CP) violation in the quark sector and search for CP violation beyond the SM, necessary to explain the observed baryon-asymmetry in the universe.

We present time-dependent measurements of CP violation and flavor mixing in the B meson system using data from the Belle II Experiment. We introduce the measurement principles of time-dependent analyses at B factories and put emphasis on our latest analysis of flavor mixing in the neutral B meson system. In particular, we focus on the treatment and estimation of uncertainties resulting from residual background processes.

T 30.2 Tue 16:30 T-H15

Inclusive analysis of untagged $B \rightarrow Xl^{+}l^{-}$ decays at Belle II — •SVIATOSLAV BILOKIN and THOMAS KUHR — Geschwister-Scholl-Platz 1, Munchen, Germany The $b \rightarrow s(d)l^{+}l^{-}$ processes are sensitive to New Physics phenomena since these decays may only occur through loops in the Standard Model.

This contribution describes a study of $b \rightarrow s(d)l^+l^-$ decays at the Belle II experiment using a fully-inclusive approach, where we have no explicit restrictions on the quark hadron final states. So far, no results of fully inclusive $b \rightarrow s(d)l^+l^-$ studies have been published because of the small signal branching ratio in the Standard Model, limited efficiency of the established tagging methods, and high rate of background processes.

This analysis intends to use machine learning techniques to reject background processes and an unrestricted tag side to measure $R(X) = \mathscr{B}(B \rightarrow X\mu^+\mu^-)/\mathscr{B}(B \rightarrow Xe^+e^-)$ as a lepton universality test, similarly to R(K) and $R(K^*)$ measurements, which are known for their deviations from the Standard Model predictions.

T 30.3 Tue 16:45 T-H15

Belle II results on charmless hadronic B-decays and prospects — •JUSTIN SKORUPA, THIBAUD HUMAIR, HANS-GÜNTHER MOSER, MARKUS REIF, and BENEDIKT WACH — Max Planck Institute for Physics

The Belle II experiment at the SuperKEKB e+ e- accelerator in Tsukuba, Japan, aims to constrain the parameters of the Cabibbo-Kobayashi-Maskawa matrix by measuring the size of the sides and angles of the unitary triangle associated to B-meson decays. Possible non-closure of the triangle would provide a hint for physics beyond the Standard Model. Belle II will significantly improve the accuracy of the determination of the angle ϕ_2 of the unitary triangle due to the large yield of expected charmless hadronic B-meson decays. Measurements of branching ratios of several charmless hadronic B-meson decays using 190 fb⁻¹ of Belle II data are presented.

T 30.4 Tue 17:00 T-H15

Search for the lepton flavour violating decay $B^0 \rightarrow \tau^{\pm} \ell^{\mp} - \bullet$ NATHALIE EBER-LEIN, THOMAS LÜCK, and THOMAS KUHR — Ludwig- Maximilians-Universität, München

Lepton flavour is conserved in the Standard Model, but violated in many new physics models. An observation of the $B^0 \rightarrow \tau^{\pm} \ell^{\mp}$ decay, where $\ell = e/\mu$, would be a clear sign for new physics.

At B factories one can determine the kinematics of the signal B meson by fully reconstructing the accompanying B meson in $e^+e^- \rightarrow \Upsilon(4S) \rightarrow BB$ events. In the rest frame of the signal B meson the mono-energetic lepton provides a clean signature to identify the signal decays. This talk presents the current status of the search for $B^0 \rightarrow \tau^{\pm} \ell^{\mp}$ decays with Belle data using the Full Event Interpretation algorithm for the reconstruction of the accompanying B meson.

Cosmic rays prove that our Universe hosts elusive astrophysical "monsters" capable of continuously and efficiently accelerating particles at extreme energies. High-energy photons and neutrinos may be the key to ultimately decipher the mystery of cosmic rays. In 2017, the candidate detection of neutrino emission from the direction of the gamma-ray flaring blazar TXS 0506+056 has put forward gamma-ray blazars as promising neutrino point-sources, hence cosmic-ray accelerators. However, to date there is neither a consistent picture for the physical mechanism nor a theoretical framework capable of convincingly explain the full set of multi-messenger observations. I will present initial encouraging steps in this multimessenger (electromagnetic and neutrino) quest and finally discuss the latest status of the field.

Location: T-H15

T 30.5 Tue 17:15 T-H15

Search for the LFV Decay $\tau \rightarrow \mu \pi^0$ — •Marton Nemeth-Csoka, Felix Meggendorfter, and Christian Kiesling — Max-Planck-Institute for Physics Munich

During its runtime from 1999 to 2010 the Belle experiment was able to confirm the Kobayashi-Maskawa theory about the occurrence of \mathscr{CP} violation and by this played a decisive role in firmly establishing the Standard Model (SM). However, there is also convincing evidence for physics beyond the SM.

Belle's upgraded successor Belle II aims for a higher precision with a goal to collect 50 times more data than Belle, a total integrated luminosity of $L_{\text{int}} = 50ab^{-1}$.

This work focuses is on the lepton flavor violating decay $\tau \rightarrow \mu \pi^0$ with the goal to explore the prospects of finding New Physics in this particular channel.

In the analysis, the decay simulated by a Monte Carlo software including detectors and full reconstruction to get an understanding of the overall kinematics. When studying the background, the largest background is that of the pair production of muons, together with light quarks and τ that decay according to the predictions of the SM.

After applying kinematic cuts and requiring a moderate confidence threshold for the identification of the muon, only 5.3% of the signal is left, but the background is fully suppressed in a sample equaling $100 f b^{-1}$.

T 30.6 Tue 17:30 T-H15

Analysis of $B \rightarrow \mu \nu$ with inclusive tagging at Belle II — FLORIAN BERN-LOCHNER, JOCHEN DINGFELDER, •DANIEL JACOBI, PETER LEWIS, and MARKUS PRIM for the Belle II-Collaboration — Physikalisches Institut der Rheinischen Friedrich-Wilhelms-Universität Bonn

 $B\overline{B}$ meson pairs are the dominant decay products of the Y(4S) resonance, which is produced in large amounts in e^+e^- collisions at the SuperKEKB collider in Japan, and their decays are measured by the Belle II experiment. Leptonic *B* meson decays such as $B \to \mu \nu$ are highly CKM- and helicity-suppressed. In a two-body decay like $B \to \mu \nu$, the muon momentum is exactly known in the rest frame of the signal-side *B* meson. By boosting the signal-side muon into that frame, a better signal resolution and improved sensitivity can thus be achieved compared to the center-of-mass frame. This requires a high-precision for the boost vector, which can be determined from the rest of the event that contains the decay products of the second *B* meson. At the same time, this information can be used to reconstruct the kinematics of the signal-side *B* meson. Boosted decision trees are trained to suppress background and increase signal purity. This talk will discuss the current status of the analysis and present approaches to maximize the sensitivity of the measurement of $B \to \mu \nu$ at Belle II, and will additionally provide an outlook on the search for sterile neutrinos.

T 30.7 Tue 17:45 T-H15

First Results and Prospects for $\tau \rightarrow \ell + \alpha$ (invisible) at Belle II — •THOMAS KRAETZSCHMAR for the Belle II-Collaboration — Max-Planck-Insitut für Physik (Werner-Heisenberg-Insitut), München, Deutschland

The Belle II experiment at SuperKEKB, an asymmetric e^+e^- collider, aims at a total integrated luminosity of 50 ab⁻¹, to pursue a rich program of Standard Model and Beyond the Standard Model physics. Until the end of 2020 and the beginning of 2021, 62.8 fb⁻¹ were collected at the Y(4S) resonance. This data set results in a sizeable sample of τ pairs, enabling detailed studies of Standard and Beyond the Standard Model measurements, including searches for Lepton Flavor Violating (LFV) decays. One of the first channels where competitive limits are expected is the $\tau \rightarrow \ell + \alpha$ (invisible) process, where α is a Goldstone boson. Here, the currently best limit has been obtained by ARGUS with an integrated luminosity of 475 pb⁻¹. Belle II will improve on this result with the recorded data. This contribution will discuss the first results of this search.
Tau lifetime measurement at Belle II — •ANSELM BAUR and DANIEL PITZL — Deutsches Elektronen Synchrotron (DESY), Hamburg, Germany

The tau-lepton lifetime represents a fundamental parameter within the Standard Model framework, contributing to the test of lepton flavor universality. Exploiting the vertex detector resolution and the tiny beam spot size at the interaction point, Belle II is expected to improve the present tau-lifetime value. The event topology where one tau decays to three charged hadrons (3-prong) and the other tau goes to a charged pion or lepton, allows to have an higher event yield respect to 3-prong topology studied by Belle. Therefore, a measurement with a statistical uncertainty competitive with the world average could already be performed using Zech's Monte Carlo re-weighting method with an early Belle II dataset.

T 30.9 Tue 18:15 T-H15

Optimization of the K_L^0 **detection and rejection at Belle II** — FLORIAN BERN-LOCHNER, JOCHEN DINGFELDER, PETER LEWIS, and •LUCAS STÖTZER for the Belle II-Collaboration — Physikalisches Institut der Rheinischen Friedrich-Wilhelms-Universität Bonn

T 31: Beyond the Standard Model (Theory) 1

Time: Tuesday 16:15-18:15

T 31.1 Tue 16:15 T-H16

Constructing Effective Field Theories to Higher Mass Dimensions — ROBERT V. HARLANDER, TIM KEMPKENS, JAKOB W. LINDER, and •MAGNUS C. SCHAAF — Institute for Theoretical Particle Physics and Cosmology, RWTH Aachen University

The Standard Model Effective Field Theory (SMEFT) provides a framework to parametrise the effects of yet unseen heavy degrees of freedom in a model independent way. While in recent years the interest in higher-dimensional operators has increased, the construction of a complete and minimal set of operators is remarkably challenging. In this talk, I will report on the implementation of a recently proposed group-theoretical algorithm for the construction of an operator basis. It systematically takes into account the redundancies which arise due to equations of motion and integration-by-parts identities among the operators. The resulting program can be applied to phenomenologically relevant theories like the Standard Model or extensions of it, including new light particles and additional symmetry groups.

T 31.2 Tue 16:30 T-H16

Catching Heavy Vector Triplets with the SMEFT: from one-loop matching to phenomenology — •EMMA GEOFFRAY¹, ILARIA BRIVIO¹, SEBASTIAN BRUGGISSER¹, WOLFGANG KILIAN², MICHAEL KRÄMER³, MICHEL LUCHMANN¹, TILMAN PLEHN¹, and BENJAMIN SUMM^{3,4} — ¹Institute for Theoretical Physics, Heidelberg University, Germany — ²Department of Physics, University of Siegen, Germany — ³Institut für Theoretische Teilchenphysik und Kosmologie, RWTH Aachen University, Germany — ⁴Institut für Theoretische Physik und Astrophysik, Universität Würzburg, Germany

An important question for both phenomenologists and experimentalists is whether one can put limits on UV model parameters by matching the full theory onto the SMEFT. I will show that this is possible and explore the complementarity between SMEFT and model-specific approaches.

In particular, I will focus on an additional theory uncertainty arising from the matching at one-loop and discuss how this affects the limits set for the Heavy Vector Triplet extension of the Standard Model. I use the SFitter framework to derive limits, taking into account Higgs measurements and electroweak precision data previously implemented, as well as two new resonance searches for VH and VV. I will discuss the impact of those measurements on the fit and the complementarity of our results with direct searches.

T 31.3 Tue 16:45 T-H16 **Precision predictions for scalar leptoquark pair production at the LHC** – •CHRISTOPH BORSCHENSKY¹, BENJAMIN FUKS², ADIL JUEID³, ANNA KULESZA⁴, and DANIEL SCHWARTLÄNDER⁴ – ¹KIT, Karlsruhe, Germany – ²LPTHE/Sorbonne Université, Paris, France – ³KIAS, Seoul, South Korea – ⁴WWU Münster, Germany

Leptoquarks are particles that simultaneously carry lepton and baryon number, and appear in many extensions of the Standard Model. The appearance of socalled flavour anomalies has led to increased interest in leptoquark models which are known to mitigate the tensions between theoretical expectations and experimental measurements.

In my talk, I will present precision predictions for the production of scalar leptoquarks at the LHC, evaluated at next-to-leading order in QCD and improved by threshold resummation corrections at next-to-next-to-leading-logarithmic accuracy. Apart from QCD contributions, included are the lepton *t*-channel exchange diagrams relevant in the light of the recent *B*-flavour anomalies. The We investigate an optimization of the reconstruction of K_L^0 mesons in the Belle II experiment. K_L^0 mesons are challenging to detect and identify because they are chargeless, and due to their long lifetime they typically do not decay within the Belle II detector. The main way to detect them is via hadronic showers in the K-Long-Muon detector (KLM). However, the detection efficiency is low and background clusters are common, which limits the usability of the KLM clusters in physics analyses. Thus, we seek to improve the discrimination between KLM clusters produced by K_L^0 showers and background sources using a clean sample of K_L^0 mesons from the process $e^+e^- \rightarrow \gamma_{\rm ISR} \left[\phi \rightarrow K_L^0 K_S^0\right]$. The K_S^0 will mainly decay to two charged pions. Thus, by finding the high energy photon ($\gamma_{\rm ISR}$) and reconstructing the K_S^0 from the pions, the four-momentum of the K_L^0 can be inferred. This allows for a direct comparison between KLM clusters from K_L^0 and background sources. Further, we investigate whether K_L^0 showers deposited in the electromagnetic calorimeter can be used to improve the K_L^0 detection efficiency while maintaining high purity.

Location: T-H16

results exhibit an interesting interplay between the different contributions, affected considerably by the choice of parton distribution functions. Additionally, I discuss the impact of NLO-QCD corrections on the so-called off-diagonal production channels, i.e. the production of a pair of different components of a given leptoquark multiplet. These predictions consist of the most precise leptoquark cross section calculations available to date and are necessary for the best exploitation of leptoquark LHC searches.

T 31.4 Tue 17:00 T-H16 **Constraining BSM models using high precision observables** — •MARTEN BERGER¹, GUDRID MOORTGAT-PICK^{1,2}, GEORG WEIGLEIN^{1,2}, and SVEN HEINEMEYER³ — ¹II. Institute of Theoretical Physics, University of Hamburg, Germany — ²DESY, Hamburg, Germany — ³Instituto de Fisica Teorica UAM-CSIC, Madrid, Spain

The high experimental accuracy of the W boson mass, M_W , measurement provides a powerful tool to test the theory and differentiate between models. One of the best motivated extensions of the Standard Model (SM) is the Minimal Supersymmetric Standard Model (MSSM). The electroweak precision observables such as M_W are highly sensitive to loop contributions determined by the model parameters respectively. Therefore the precise experimental accuracy can be used to narrow down possible scenarios. In this talk a stand-alone mathematica code for predicting M_W in the MSSM is presented. It includes full one-loop as well as leading higher-order corrections of SUSY-type, which are combined with state-of-the-art SM-type contributions. The prediction for M_W is discussed in comparison with the current experimental result.

T 31.5 Tue 17:15 T-H16 Bachelor thesis: Vacuum stability constraints in the NMSSM — •FABIO CAMPELLO¹, GEORG WEIGLEIN², and THOMAS BIEKÖTTER² — ¹Universität Hamburg, Hamburg, Deutschland — ²DESY, Hamburg, Germany

In supersymmetric extensions of the Standard Model the electroweak (EW) vacuum is not generally the global minimum of the scalar potential, and tunneling to deeper minima is possible. Since the lifetime of the EW vacuum must be at least of the order of the age of the universe, constraints on the parameter space can be obtained from an analysis of vacuum stability. In this talk the vacuum structure of the next-to-minimal supersymmetric extension of the SM (NMSSM) is investigated, where the scalar potential receives contributions from an extended Higgs sector and from superpartners of the SM fermions. The results are discussed in comparison to the case of the minimal supersymmetric extension (MSSM).

T 31.6 Tue 17:30 T-H16

SU(6) Gauge Higgs Unification – •ANDREAS BALLY¹, ANDREI ANGELESCU¹, FLORIAN GOERTZ¹, and SIMONE BLASI² – ¹Max Planck Institute for Nuclear Physics, Heidelberg – ²Vrije Universiteit, Brussel, Belgium

We present a minimal viable Gauge-Higgs Grand Unification scenario in warped space based on a SU(6) bulk symmetry - unifying the gauge symmetries of the SM and their breaking sector. We show how the issue of light exotic new states is eliminated by appropriately breaking the gauge symmetry on the UV and IR boundaries by either brane scalars or gauge boundary conditions. The SM fermion spectrum is naturally reproduced including Dirac neutrinos and we compute the Higgs potential at one-loop, finding easily solutions with a realistic Higgs mass. The problem of proton decay is addressed by showing that baryon number is a hidden symmetry of the model. Among the phenomenological consequences, we highlight the presence of a scalar leptoquark and a scalar singlet.

The usual X,Y gauge bosons from SU(5) GUTs are found at collider accessible masses.

T 31.7 Tue 17:45 T-H16 Baryogenesis and Dark Matter in Extended Inert Doublet Model — •Sven FABIAN, FLORIAN GOERTZ, and MARÍA ISABEL DIAS ASTROS — Max-Planck-Institut für Kernphysik, Heidelberg, Deutschland

Despite the great success of the Standard Model (SM) of Particle Physics in explaining many experimental observations to an astonishing degree of accuracy, it cannot account for the long-standing conundrums of the nature of dark matter (DM) and of the obvious dominant abundance of matter compared to antimatter in our Universe. In this talk, we will discuss the Inert Doublet Model, augmented with a higher-dimensional operator tied to the SM gauge sector and – vital for baryogenesis – inducing CP violation. In addition to identifying the parameter space for the observed DM relic abundance, we investigate the potential of this operator for giving rise to the measured baryon asymmetry during the electroweak phase transition. We will find that the discussed extension of the IDM can, in principle, serve as an effective theory in which both DM and baryogenesis are accounted for. T 31.8 Tue 18:00 T-H16 Higher-Dimensional Operators in the Inert Doublet Model: Dark Matter and CP Violation — •María Isabel Dias Astros, Florian Goertz, and Sven FABIAN — Max-Planck-Institut für Kernphysik, Heidelberg, Deutschland

Facing the puzzles of dark matter (DM) with its properties yet to be deciphered, the Inert Doublet Model (IDM) has been widely studied as a possible theory explaining DM and in the context of the Electroweak Phase Transition, as a first step towards accommodating the three Sakharov conditions allowing for baryogenesis. Following this motivation, we will discuss in this talk the IDM as an effective field theory (EFT) with a higher-dimensional CP-violating operator added to the scalar potential. Working with the EFT approach, a comprehensive discussion of DM phenomenology with promising parameter space is given, while taking theoretical and the latest experimental constraints into account. In particular, for particles heavier than 500 GeV higher-dimensional derivative operators lead to an extended viable parameter space even for non-(quasi)degenerate scalar masses. We conclude that the discussed sources of CP violation do not spoil the predicted DM relic abundance.

T 32: QCD (Exp.) 1

Time: Tuesday 16:15-18:15

T 32.1 Tue 16:15 T-H17

Measurement and QCD analysis of inclusive jet production in deep inelastic scattering at ZEUS — •FLORIAN LORKOWSKI — DESY, Hamburg, Germany The cross sections of deep inelastic scattering processes at the electron-proton collider HERA are a well established tool to test perturbative QCD predictions. Additionally, they can be used to determine the non-perturbative parton distribution functions of the proton. Measurements of jet production cross sections are particularly well suited to also constrain the strong coupling constant.

In this talk, a measurement of inclusive jet cross sections in neutral current deep inelastic scattering using the ZEUS detector at the HERA collider is presented. The data was taken in the years 2003 to 2007 at a center of mass energy of 318 GeV and corresponds to an integrated luminosity of 344 pb⁻¹. Massless jets, reconstructed using the k_{\perp} -algorithm in the Breit reference frame, are measured as a function of the squared momentum transfer Q^2 and the transverse momentum of the jets in the Breit frame $p_{\perp,\text{Breit}}$.

The measured cross sections are compared to previous measurements as well as NNLO theory predictions. The consistency of the measurement is demonstrated by a simultaneous determination of parton distribution functions and the strong coupling constant.

T 32.2 Tue 16:30 T-H17 QCD and SMEFT analysis of CMS 13 TeV inclusive jet cross section data — •TONI MÄKELÄ and KATERINA LIPKA — DESY, Hamburg, Germany

The parton distributions of the proton, the strong coupling constant and the top quark mass are extracted simultaneously, using the cross sections of inclusive jet production and top quark-antiquark pair production at the LHC at a center of mass energy of 13 TeV. The standard model analysis is performed at NLO and NNLO. In an alternative analysis, the standard model cross section is extended with effective couplings for 4-quark contact interactions at NLO. In particular, left-handed vector-like or axial vector-like colour-singlet exchanges are considered. For the first time, the Wilson coefficients of contact interactions are extracted simultaneously with the standard model parameters using the LHC data.

T 32.3 Tue 16:45 T-H17

Triple-differential measurement of dijet production at $\sqrt{s} = 13$ TeV with the CMS detector — GÜNTER QUAST, KLAUS RABBERTZ, and •DANIEL SAVOIU — Institute of Experimental Particle Physics (ETP), Karlsruhe Institute of Technology (KIT)

Jet measurements at high precision are an essential probe of quantum chromodynamics (QCD) and constitute valuable experimental inputs to determinations of fundamental QCD parameters and of the parton distribution functions (PDFs) describing the structure of protons.

In this talk, we present a recent measurement of the dijet production cross section using proton-proton collision data collected at a center-of-mass energy of 13 TeV by the CMS detector at the CERN LHC, amounting to an integrated luminosity of 36.3 fb⁻¹. Jets are reconstructed using the anti- $k_{\rm T}$ algorithm for radius parameters of R = 0.4 and 0.8 and cross sections are measured triple-differentially as a function of the kinematic properties of the two jets with largest transverse momenta. After accounting for detector- and reconstruction-specific effects in a three-dimensional unfolding procedure, the data are compared to theoretical predictions derived at next-to-next-to-leading order in perturbative QCD and the impact of the data for determinations of the proton PDFs and the strong coupling constant $\alpha_{\rm s}$ is studied.

Location: T-H17

T 32.4 Tue 17:00 T-H17

Electroweak corrections to high pT jets — •MIKEL MENDIZABAL and HANNES JUNG — DESY, Hamburg, Germany

The production of electroweak (EW) bosons in association with jets has been extensively studied at particle colliders. The EW boson is considered the outcome of the hard process and the jets a product of parton evolution. These events are a great test of quantum chromodynamics and allow to study parton density functions and parton evolution equations. So far, light quarks and gluons are considered in the parton evolution. However, with increasing centre-of-mass energies the probability of radiating heavier particles increases.

In this analysis, the production of EW bosons in association with jets is studied specifically with the aim to investigate EW boson emitted in the parton shower. To this end, events with high transverse momentum jets are studied. Then, the contribution of EW boson emissions is measured.

Preliminary results are presented with data collected in 2016, corresponding to an integrated luminosituy of 36.3 fb^{-1} . The contribution of the Z boson is studied in the leptonic decay channel.

T 32.5 Tue 17:15 T-H17

Triple differential measurement of the inclusive Z+jet production — •CEDRIC VERSTEGE, KLAUS RABBERTZ, and GÜNTER QUAST — Institut für Experimentelle Teilchenphysik, Karlsruher Institut für Technologie (KIT)

The triple differential inclusive cross section for $Z (\rightarrow \mu \mu) + jet$ production is measured combining CMS Run II data from 2016-2018. The measurement uses the observables p_T^Z , the difference in rapidity between the Z-Boson and the jet y^* as well as the boost of the center of mass system y_b . Those variables allow for a suitable division of the phase space in order to obtain a better sensitivity to the partonic subprocesses.

Detector effects are corrected via a three-dimensional unfolding procedure. The resulting cross section is then compared to QCD theory predictions at nextto-next-to-leading order. The results can be used as constraints for fitting the PDFs.

T 32.6 Tue 17:30 T-H17

Measurement of jet mass distribution of hadronic W and Z bosons — •STEFFEN ALBRECHT¹, ANDREAS HINZMANN¹, DENNIS SCHWARZ², and ROMAN KOGLER³ — ¹Universität Hamburg — ²Austrian Academy of Sciences — ³DESY Hamburg

In this talk we introduce a new effort towards measuring the jet mass distribution of hadronically decaying W and Z bosons.

We study events in which the bosons have a large transverse momentum and thus produce strongly collimated decay products reconstructed as single fat jets. The substructure of such jets proves to be a useful handle in various procedures (e.g. jet calibration, jet tagging), but has room for improvement in its modelling. We aim to gain an in-depth understanding of the substructure by studying the unfolded jet mass distribution in dependence of the jet p_T and substructure tagger discriminants. While previous measurements of jet mass have been carried out for gluon, quark and top jets in dijet, Z(ll)+jet and $t\bar{t}$ samples, this is the first study of W and Z jet masses in the processes with W(qq)+jets, Z(qq)+jets as well as hadronic $t\bar{t}$ systems in the final states.

In addition the measurement of the difference $m_Z - m_W$ will be pursued, setting a first step towards a potential measurement of the W mass with jet substructure.

T 32.7 Tue 17:45 T-H17

Jet Energy Calibration for Ultra Legacy Data with Z+Jet Events at CMS -•ROBIN HOFSAESS, DANIEL SAVOIU, FLORIAN VON CUBE, and MAXIMILIAN HORZELA — KIT (ETP), Karlsruhe, Germany

High precision analyses in modern particle physics experiments rely on the measurement of jets coming from the particle interactions. Since jets comprise many different particles and the observation of such complex physics objects is affected by detector- and reconstruction-specific effects, sophisticated methods are necessary to get a reliable and accurate calibration of the jet energy.

At CMS, a factorized approach - collectively known as the jet energy calibration - is employed for correcting shifts in the jet energy. An important step in this process exploits events where a jet is balanced against a well-measured reference object such as the Z boson. By comparing the transverse momenta of the two objects, it is possible to determine the absolute jet energy scale, accounting for any residual differences between simulation and data.

In this talk, the methods for the determination of the jet energy scale will be described and the latest results for the legacy calibration of Run II will be presented.

T 33: Top Quarks: Production (Exp.) 2

Time: Tuesday 16:15-18:30

T 33.1 Tue 16:15 T-H18

An effective field theory approach using top quark polarisation and spin correlations in $t\bar{t}$ production at the LHC — •ANDRE ZIMERMMANE-SANTOS, AFIQ ANUAR, ALEXANDER GROHSJEAN, and CHRISTIAN SCHWANENBERGER — Deutsches Elektronen-Synchrotron, Hamburg, Germany

The Effective Field Theory (EFT) approach provides a systematic and modelindependent way to search for new physics. It assumes new heavier particles exist outside the energy reach of the LHC. Nonetheless, their effects can be parametrized by new effective interactions constructed with Standard Model (SM) fields. Hence, new couplings can be identified and measured via small deviations from SM predictions.

In this study, we aim to use particular sets of observables related to the top quark polarisation and spin correlation in top quark pair events with two leptons in the final state. Each of those sets are sensitive exclusively to a subset of EFT couplings. This provides a natural way of uncorrelating EFT effects, allowing limits on their strength to be drawn with unprecedented precision. We investigate various EFT scenarios using the *dim6top* and *SMEFTatNLO* models. Subsequently, we also verify necessary translations between different EFT formalism. Our findings pave the way for the EFT interpretation using full CMS Run 2 data.

T 33.2 Tue 16:30 T-H18 Measurement of the dileptonic tt differential cross section in a BSM phase

space at CMS — Valeria Botta, Lutz Feld, •Danilo Meuser, Philipp Nat-TLAND, and MARIUS TEROERDE — I. Physikalisches Insitut B, RWTH Aachen University

Measurements of the tt production cross section yield important precision tests of the Standard Model (SM), while also probing scenarios for physics beyond the SM (BSM).

This analysis aims to measure the $t\bar{t}$ cross section in a phase space where additional contributions from BSM scenarios could be present. It is based on the data set recorded by CMS in the years 2016 to 2018 at a center-of-mass energy of 13 TeV, corresponding to an integrated luminosity of 138 fb⁻¹. The BSM scenarios considered include supersymmetric and dark matter models, where, similarly to the dileptonic tt channel, two leptons, b jets and undetected particles are produced.

Unlike previous measurements, where the differential cross sections were mainly measured as a function of kinematic variables of the leptons or top quarks, this analysis focuses on observables related to the neutrinos, like the missing transverse momentum and the angular distance between the missing transverse momentum and the nearest lepton, to separate BSM from SM tt events. In order to increase the sensitivity of the analysis multivariant techniques are used which improve the resolution of the missing transverse momentum in SM tt events. In this talk the analysis strategy will be presented and preliminary results on the improved missing transverse momentum resolution and on systematic uncertainties will be shown.

T 33.3 Tue 16:45 T-H18

Studying prospects of a measurement of the cross section of top-quark pair production with additional charm quarks in the lepton+jets channel at AT-LAS at \sqrt{s} = 13 TeV — •Lukas Ehrke, Tobias Golling, Manuel Guth, JOHNNY RAINE, and KNUT ZOCH — Université de Genève, Geneva, Switzerland The goal of this analysis is to measure the inclusive cross section of $t\bar{t}$ production with additional charm quarks in the ATLAS collaboration. This talk focuses on

T 32.8 Tue 18:00 T-H17 Differential cross section for $Z\gamma$ +jets using the ATLAS detector — •VINCENT GOUMARRE — Deutsches Elektronen-Synchrotron (DESY), Hamburg, Germany Differential measurements of the production of Zy bosons in association with jet activity in pp collisions at $\sqrt{s} = 13$ TeV are presented in this work, using the full Run2 dataset of 139fb⁻¹ collected by the ATLAS detector at the LHC.

Jet activity is a crucial point to study since differential distributions can help constrain important Standard Model (SM) parameters and calibrate models used in inputs in other observables, such as PDFs functions. Moreover, due to the possibility to fully reconstruct the final state, and a large cross section with a small background, Zy is a good candle to test beyond the SM physic, with models such as ALPs or anomalous gauge couplings.

Distributions are measured in a fiducial space with transverse momentum of the photon greater than 30 GeV and considering only events where the Z-boson decays leptonically. The sum of the dilepton invariant mass and the dilepton plus photons invariant mass has to be greater than 182 GeV to suppress final state radiation.

Location: T-H18

the semileptonic $t\bar{t}$ decay channel for which it would be the first measurement of this cross section. The measurement will benefit several analyses where this process is a non negligible background, most notably in the search for a $t\bar{t}$ pair in association with a Higgs boson where the Higgs boson decays into two b-quarks.

To reduce background processes not containing a $t\bar{t}$ pair *b*-tagging is needed, whereas to identify the events with additional c-quarks c-tagging is needed. A further complication in the lepton+jets channel are *c*-quarks originating from the hadronically decaying W boson. Therefore, the existing flavour tagging methods are extended to allow for simultaneous b- and c-tagging. New working points are derived on a 2D plane, and based on the b- and c-multiplicity, multiple regions are defined with different contributions of the different $t\bar{t}$ +jets components. Initial studies show promising sensitivity to the cross section. The measurement in the lepton+jets channel benefits from larger statistics compared to the dilepton channel. However, the charm quarks from the hadronic W decays pose a greater modelling challenge.

T 33.4 Tue 17:00 T-H18

Measurement of the inclusive production cross sections of a top-quark pair in association with a Z boson at $\sqrt{s} = 13$ TeV in final states with three leptons using deep neural nets with the ATLAS detector - • STEFFEN KORN, AR-NULF QUADT, and ELIZAVETA SHABALINA — II. Physikalisches Institut, Georg-August-Universität Göttingen

Through the associated production of a top quark pair and a Z boson, the strength and structure of the neutral current weak coupling of the top quark and the Z boson can be measured. It provides sensitivity to the top quark's weak isospin in the framework of the Standard Model (SM). The measurement of this fundamental parameter of the SM also serves as a probe to new physics beyond the SM. The process was measured by ATLAS and CMS at $\sqrt{s} = 13$ TeV with the full run 2 data set. In a new, refined analysis multivariate techniques are used to improve the sensitivity of the measurement. The impact of the usage of multiclass deep neural network for event classification on the systematic uncertainties for a measurement of the inclusive cross section of $t\bar{t}Z$ final states with three charged leptons is presented.

T 33.5 Tue 17:15 T-H18

First simultaneous differential measurement of tZq and ttZ processes at the CMS Experiment — •FEDERICA COLOMBINA, ANDREAS MEYER, and ABIDEH JAFARI — Notkestraße 85, 22607 Hamburg, Germany

At the Large Hadron Collider (LHC) at CERN, about millions of top quark events have been produced. The data recorded during LHC Run-2, in the years 2016-2018, gave access to first differential measurements of top quark production in association with Z-bosons, and precisely probes the coupling between top quarks and Z bosons for the first time. The cross sections of top quark pair producton, tTZ, and single-top quark production, tZq, are similar, and both processes are mutual backgrounds to one another. Measurements of top-Z coupling and EFT analyses require measurements of both these processes and their correlation. In this analysis, tZq and $t\bar{t}Z$ are measured simultaneously for the first time, aiming to better understand the correlation between these two processes. Furthermore, the evaluation of their differential cross section can bring evidence of possible deviations from the standard model, providing information for EFT analyses and new physics scenarios.

T 33.6 Tue 17:30 T-H18

Background model in a $t\bar{t}W$ cross-section measurement — •Marcel Niemeyer, Arnulf Quadt, and Elizaveta Shabalina — Georg-August University Goettingen

The top-quark pair production in association with a W boson is an interesting process by itself and exhibit an important background to processes like $t\bar{t}H$ or 4-tops production. Due to higher order electroweak corrections, the process is difficult to model. In consequence, a mismodelling of $t\bar{t}W$ has been observed in previous analyses. Thus, it is of high importance to measure this process to improve our understanding of it. The analysis is performed in the multi-lepton channel requiring 2ℓ (same-sign) or 3ℓ . The resulting event sample has a significant contribution from fake backgrounds.

To estimate this background, an extended template fit is performed that uses a discriminant based on isolation and b-tagging variables, referred to as a prompt lepton veto. The fit, the calibration of the prompt lepton veto, and the related systematic uncertainties will be discussed in this talk.

T 33.7 Tue 17:45 T-H18

Measurement of differential cross-sections of the $t\bar{t}\gamma$ production in the dilepton channel in proton-proton collisions at $\sqrt{s} = 13$ TeV with ATLAS detector - •Buddhadeb Mondal, Ivor Fleck, and Carmen Diez Pardos - Center for Particle Physics Siegen, Experimentelle Teilchenphysik, Universität Siegen The top quark being the heaviest fundamental particle in the Standard Model (SM) plays a very important role in the study of fundamental interactions. It has a very short lifetime and it decays before it hadronizes, passing its properties to its decay products. Top quark pair production in association with a photon $(t\bar{t}\gamma)$ is a very important process for measuring the coupling between top quark and photon. A precise measurement of this coupling is necessary for testing the SM and also for probing any new physics effect at very high energy scale. Deviations from the SM coupling can be a limit of new physics phenomena that can be interpreted in the context of effective field theory approaches. In this talk, measurement of differential cross-section using 139 fb^{-1} of data collected by the ATLAS detector in proton-proton collisions at $\sqrt{s} = 13$ TeV will be presented. This measurement is done in the dileptonic decay channel of the $t\bar{t}$ pair.

T 33.8 Tue 18:00 T-H18 Measurement of $t\bar{t} + \gamma$ production with the full Run 2 ATLAS data set — •ANDREAS KIRCHHOFF, ARNULF QUADT, and ELIZAVETA SHABALINA — II. Physikalisches Institut, Georg-August-Universität Göttingen The optimal way to measure the top-photon coupling and later interpret it within an EFT-framework would be an e^+e^- collider with sufficient energy. As such a collider does not exist, another possibility to measure it is the production of $t\bar{t}$ pairs in association with a photon. Unfortunately, most of such photons will originate from the decay products of the top quarks and hence do not convey any information about the top-photon coupling. In contrast, photons produced in the production of the $t\bar{t}$ pair mostly originate from the top quark (beside a small contribution from ISR). The separation of photons originating from production and decay is tried for the first time in this ATLAS analysis. In this talk, the status of the currently ongoing full Run 2 analysis of the $t\bar{t} + y$ process in the *l*+jets channel will be presented. The talk will focus on showing how deep neural networks are used to measure the $t\bar{t} + y$ cross section, where the photon is emitted during production. First fit results will also be shown.

T 33.9 Tue 18:15 T-H18

Search for $t\bar{t}\gamma\gamma$ production in pp collisions at $\sqrt{s} = 13$ TeV with the ATLAS detector — •ARPAN GHOSAL, IVOR FLECK, CARMEN DIEZ PARDOS, and AMARTYA REJ — Center for Particle Physics Siegen, Experimentelle Teilchenphysik, Universität Siegen

The production of top quarks with photons gives access to measure the strength of the electroweak coupling of the top quark and the photon. While the production of top quark pairs $(t\bar{t})$ with one photon is being extensively studied, the $t\bar{t}$ production in association with two photons $(t\bar{t}\gamma\gamma)$ has not been observed yet. The $t\bar{t}\gamma\gamma$ is a rare process and not only is it a good candidate for probing the top EW coupling but is also relevant as an irreducible background process to $t\bar{t}$ production in association with a Higgs boson decaying to two photons $(H \rightarrow \gamma\gamma)$. Besides, new sources of CP-violation are expected from physics phenomena beyond the SM. These sources can appear as electric dipole moment terms in top-quark interactions, and their precise measurement is essential to determine the effects of new physics. Understanding the $t\bar{t}\gamma\gamma$ process can help put better bounds on top-quark dipole moments.

The cross-section of the process is expected to be of the order of 10 fb at $\sqrt{s} = 13$ TeV, much lower than the $t\bar{t}\gamma$ process. The presentation will discuss the ongoing efforts in the search for the process in semileptonic $t\bar{t}$ decay channel using the full Run 2 dataset collected by ATLAS detector at 13 TeV.

T 34: Top Quarks: Properties -2

Time: Tuesday 16:15-18:00

T 34.1 Tue 16:15 T-H19

Measurement and EFT interpretation of differential $t\bar{t}$ cross-sections in the boosted lepton+jets channel with the ATLAS detector at $\sqrt{s} = 13$ TeV — KEVIN KRÖNINGER, JOHANNES ERDMANN, and •KEVIN SEDLACZEK — TU Dortmund University, Department of Physics

Physics beyond the Standard Model (SM) can, in absence of resonances, be investigated in a model-independent way by using effective field theory (EFT) approaches. Without assumption of the underlying theory, effects of potential new high-mass particles at a low energy scale can be quantified by higher dimension expansions of the SM Lagrangian.

At the LHC, physics in the top sector is entering a phase of precision measurements combined with very accurate predictions. Meanwhile, many theories beyond the SM predict deviations in the top-quark couplings or new interactions of the top quark. These aspects make model-independent measurements in the top sector a very attractive way to test the SM for deviations arising from new physics at higher energy scales.

In this talk, a differential $t\bar{t}$ cross-section measurement is shown. The measurement is performed in the boosted lepton + jets channel on the full Run 2 dataset taken with the ATLAS detector at $\sqrt{s} = 13$ TeV. The differential measurements in different kinematic variables are unfolded to the particle level. One of the unfolded distributions is then used to derive bounds on the contributions of new physics within the EFT framework via two dimension-6 operators.

T 34.2 Tue 16:30 T-H19

Top-antitop energy asymmetry in jet-associated top-quark pair production at ATLAS — •Alexander Basan¹, Asma Hadef¹, Jessica Höfner¹, Lucia Masetti¹, Eftychia Tzovara¹, and Susanne Westhoff² — ¹Universität Mainz — ²Universität Heidelberg

The top quark is particularly well suited to probe the standard model (SM) and many extensions thereof at the electroweak symmetry-breaking scale and beyond.

At hadron colliders, the $t\bar{t}$ production is symmetric at leading order perturbation theory under the exchange of the top- and anti-top-quark, while interfer-

ences at higher orders create an asymmetry. This charge asymmetry can provide sensitive probes for many models beyond the standard model. Within the framework of standard model effective field theories (SMEFT), the charge asymmetry is especially sensitive to four-quark operators and one operator that modifies the top-gluon interaction.

In inclusive jet-associated top-quark pair production the asymmetry arises already at leading order in quark-gluon interactions. Furthermore, the $t\bar{t}j$ final states allow the definition of a new observable, the energy asymmetry, expressed in terms of the distribution of the energy difference $E_t - E_{\bar{t}}$.

This talk presents the measurement in lepton+jets events with a high p_T hadronically decaying top quark at ATLAS with a centre of mass energy of $\sqrt{s} = 13$ TeV as well as limits on the Wilson coefficients of four-quark operators within the SMEFT framework.

T 34.3 Tue 16:45 T-H19

Location: T-H19

Untersuchung neuer Physik über die Energieasymmetrie in der Top-Antitop-Jet Produktion am ATLAS — •JESSICA HÖFNER¹, ALEXANDER BASAN¹, ASMA HADEF¹, LUCIA MASETTI¹, EFTYCHIA TZOVARA¹ und SUSANNE WESTHOFF² — ¹Universität Mainz — ²Universität Heidelberg

Das Top-Quark ist das schwerste Teilchen im Standardmodell der Elementarteilchen und das einzige Quark das nicht direkt hadronisiert sondern zerfällt. Es eignet sich sehr gut dafür Physik außerhalb des Standardmodells zu suchen, denn es könnten noch unentdeckte schwerere Teilchen oder auch neue Wechselwirkungen mit dem Top-Quark interagieren.

Bei der Produktion eines Top-Antitop-Paares mit zusätzlichem Jet kann die Energieasymmetrie, eine neue Observable der Ladungsasymmetrie, bestimmt werden, die besonders sensitiv auf Physik jenseits des Standardmodells sein kann. Nach einer ersten veröffentlichten Messung der Energieasymmetrie mit dem ATLAS Experiment, werden Möglichkeiten untersucht sowohl den Phasenraum der Messung zu erweitern als auch die Ergebnisse mit denen aus Wirkungsquerschnittmessungen zu kombinieren. In diesem Vortrag wird die erwartete Sensitivität dieser Erweiterungen vorgestellt.

T 34.6 Tue 17:30 T-H19

T 34.4 Tue 17:00 T-H19

Measurements of observables sensitive to colour reconnection in $t\bar{t}$ events — DOMINIC HIRSCHBÜHL, WOLFGANG WAGNER, and •SHAYMA WAHDAN - Bergische Universität Wuppertal, Wuppertal, Germany

Colour reconnection (CR) is a mechanism that describes the interactions that can occur between colour fields during the hadronisation transition. In the context of precise top-quark mass measurements, it plays a crucial role. The modelling of CR has become one of the dominant sources of systematic uncertainty in these measurements. Ongoing top-quark mass analyses use PYTHIA 8 MC event generator for parton showering and hadronisation. PYTHIA 8 comes with several alternative CR models which should be explored to estimate the CR modelling uncertainty. At the same time, the models should be confronted with LHC data to test their validity. Only models which are in agreement with data, in general, are suitable to define the corresponding modelling uncertainty. This analysis presents a measurement of charged-particle distributions sensitive to the different CR models in PYTHIA 8 in top-quark pair production. The measurement is based on data collected using the ATLAS detector at the LHC in proton-proton collisions at a centre-of-mass energy of 13 TeV with an integrated luminosity of 139 fb-1.

T 34.5 Tue 17:15 T-H19

Search for flavour-changing photon interactions in top-quark production and decay at the ATLAS experiment — TOMAS DADO, JOHANNES ERDMANN, BENEDIKT GOCKE, •FLORIAN MAUSOLF, OLAF NACKENHORST, and BJÖRN WENDLAND — TU Dortmund University, Department of Physics

In the Standard Model (SM) of particle physics, flavour-changing neutral currents (FCNC) are strongly suppressed, but several theories beyond the SM predict FCNC with much higher rates. In this talk, a search for flavour-changing photon interactions in top-quark production and decay is presented. Protonproton-collision data corresponding to an integrated luminosity of 139 fb⁻¹ are analysed which were taken with the ATLAS detector at a centre-of-mass energy of 13 TeV. It is separately searched for interactions involving the top quark, the up quark and the photon as well as for interactions of the top quark, the charm quark and the photon. Events with one photon, one b-tagged jet, one electron or one muon, and a minimum amount of missing transverse momentum are selected. Contributions from events with objects mis-reconstructed as photons are estimated using data-driven methods. Multiclass deep neural networks are used to separate the signal from the background. The analysis strategy is presented and upper limits on the strength of the FCNC couplings are set.

Studies for the measurement of the production of top-quark pairs in association with a ${\it Z}$ boson decaying to a pair of tau leptons with the ATLAS detector - •SIMON NEUHAUS and THOMAS DADO - TU Dortmund University, Department of Physics

A Study of the associated production of a top-quark pair and a Z boson decaying into a tau-lepton pair is presented. This process allows to test the leptonuniversality prediction of the Standard Model of particle physics in the topquark-sector. Additionally, this process is sensitive to various BSM couplings between the top quarks and tau-leptons. Because only the visible mass of the ditau system is reconstructable, a significant contribution from the off-shell events is expected in the signal region.

The measurement targets decays with one or two light leptons, at least three jets and two hadronically decaying tau leptons. Some of the important background processes include the diboson (ZZ, WZ, WW) processes and processes with misreconstructed tau leptons. Initial studies on the optimization of the event selection using the Monte Carlo simulations for the ATLAS data of the complete LHC Run 2 will be shown. Observables that are interesting for the optimisation of the selection include: the transverse momentum of all three or four leptons, the number of jets and the number of *b*-tagged jets.

T 34.7 Tue 17:45 T-H19

Location: T-H20

Simulation of selected top-quark processes at the FCC-ee and their interpretation in terms of effective field theories - CORNELIUS GRUNWALD¹, KEVIN Kröninger¹, Romain Madar², Stéphane Monteil², and •Lars Röhrig¹ — ¹Department of Physics, Dortmund, Germany — ²Laboratoire de Physique de Clermont, Clermont-Ferrand, France

While in the flavor-physics sector future upgrades of the LHCb detector at CERN and Belle II at KEK are aimed at precision measurements, experiments at future colliders such as the FCC-ee are expected to improve electroweak and topquark physics in an unrivaled way. Since it is interesting to set the measurements by the FCC-ee into a global context, the estimation of the precision of selected top-quark processes is important for estimating the impact on the constraints of dimension-six operators.

In this talk, relevant observables and the impact of dimension-six operators on these observables are presented. The parameterizations as function of the dimension-six operator strength are given. This will allow to set constraints on the strength of dimension-six operators assuming uncertainties for the measurements at the FCC-ee.

T 35: Higgs Boson: Associated Production 1

Time: Tuesday 16:15-18:15

T 35.1 Tue 16:15 T-H20

Higgs Boson Mass Reconstruction in the ttH Multi-lepton Channel Using AT-LAS data — Igor Boyko¹, •Adam Herold², Nazim Huseynov¹, Jan Kybic², André Sopczak², Petr Urban², and Cyrus Walther³ - ¹JINR Dubna -²CTU in Prague — ³TU Dortmund

This study deals with the reconstruction of the Higgs boson mass in the 2ISS + $1\tau_{had}$ channel in ttH production. Based on the reconstructed mass, the goal is to separate the signal from background productions such as the ttZ production. The data created by the full ATLAS detector simulation are used to develop two neural networks. First, a classification neural network that organizes the data by assigning detected particles to corresponding positions in the channel. Second, a regression neural network that reconstructs the mass of the Higgs boson. The developed neural network is tested and is shown to outperform the Missing Mass Calculator technique.

T 35.2 Tue 16:30 T-H20

Investigation of ttH(bb) Events with Very High Higgs Boson Momentum at ATLAS Detector — •Doga Elitez, Lucia Masetti, Eftychia Tzovara, ASMA HADEF, and ALEXANDER BASAN — Johannes Gutenberg-Universität Mainz, Mainz, Deutschland

The coupling of the Higgs boson to the top quark is very sensitive to effects of the physics beyond the Standard Model (BSM) and the most favorable production mode for direct measurement of the top Yukawa coupling is the Higgs production in association with a pair of top quarks $(t\bar{t}H)$. The decay to two bottom quarks $(H \rightarrow b\bar{b})$ has the largest branching fraction of about 58%. This analysis aims at events in which one of the top quarks decays semi-leptonically and produces an electron or a muon plus several jets. The so-called ultra boosted topology targets events containing a Higgs boson produced at very high transverse momentum, which is contained in a single small-R jet. This topology is not included in the current high p_T (boosted) Higgs boson selection and requires a dedicated analysis. In this talk, methods to improve background rejection and event reconstruction to increase the sensitivity above the current p_T range are presented, along with the challenges of combining the different channels.

T 35.3 Tue 16:45 T-H20 Measurement of the $t\bar{t}H$ production cross-section with $H \rightarrow b\bar{b}$ in the boosted topology with the ATLAS detector - •Eftychia Tzovara, Lucia MASETTI, DOGA ELITEZ, ASMA HADEF, and Alexander Basan — JGU Mainz, Germany

Studying the coupling of the Higgs boson to the top quark is of particular interest, since it could be sensitive to effects of physics beyond the SM. The Higgs production in association with a top-quark pair is the most favourable process for a direct measurement of the top Yukawa coupling. The decay to two b-quarks has the largest branching ratio, while it allows for the reconstruction of the Higgs boson kinematics. The analysis presented here aims at events in which one of the top quarks decays semi-leptonically, producing an electron or a muon, and the other one hadronically. In the single-lepton channel, there is a specific boosted region, targeting events with a Higgs boson produced at high transverse momentum p_T .

Due to the highly complex final state and the large SM backgrounds, the reconstruction of the Higgs boson becomes a complicated task. The ultimate goal is to constrain the background events of the boosted channel in order to maximise the statistical significance of the measurement. For this purpose, multivariate techniques are used to discriminate between signal and background events, in particular from $t\bar{t}$ +jets production. For the first time, the signal strength is also measured differentially in bins of the Higgs boson p_T . Finally, the measurement of the $t\bar{t}H(b\bar{b})$ cross-section, using the full LHC run-2 data, as well as further improvements on the boosted channel, will be presented.

T 35.4 Tue 17:00 T-H20

Improvements of the MVA classifiers for the $t\bar{t}H(b\bar{b})$ analysis in the dilepton channel with full Run2 data in the CMS experiment — • ANGELA GIRALDI and MARIA ALDAYA — DESY, Hamburg, Germany

In the Standard Model (SM), the Higgs boson couples to fermions with a Yukawa-type interaction and a strength proportional to the fermion mass. The associated production of a Higgs boson with a top-quark pair $(t\bar{t}H)$ is therefore the best direct probe of the top-Higgs Yukawa coupling, a vital element to verify the SM nature of the Higgs boson. In the SM, the Higgs boson decays into b-quark-antiquark pair with the largest branching fraction, and is thus experimentally attractive as a final state. The dominant background contributions arise from $t\bar{t}$ +jets production, and in particular the $t\bar{t}b\bar{b}$ background is irreducible with respect to $t\bar{t}H, H \rightarrow b\bar{b}$. To better enhance the sensitivity, the signal is extracted exploiting multivariate analysis (MVA) techniques.

This talk focuses on the analysis of the $t\bar{t}H,H \rightarrow b\bar{b}$ process in final states with two leptons using proton-proton data collected by the CMS experiment at the LHC during 2016-2018 at $\sqrt{s} = 13$ TeV. The possibility to critically increase the sensitivity to the $t\bar{t}H$ signal is investigated using machine learning approaches. Detailed studies on the optimization and performance of MVA discriminants trained using Artificial Neural Networks are presented in this final state.

T 35.5 Tue 17:15 T-H20

Adversarial Machine Learning Methods for Modelling Uncertainty Reduction in the Bottom Anti-Bottom Higgs Decay Channel of Higgs-associated Top Quark Pair Production with ATLAS at 13 TeV — ARNULF QUADT, •CHRIS SCHEULEN, and ELIZAVETA SHABALINA — II. Physikalisches Institut, Georg-August Universität Göttingen

The bottom anti-bottom Higgs decay channel of Higgs-associated top quark pair production offers direct access to measurements of the top Yukawa coupling and Higgs- p_T differential cross section, which are sensitive to potential new physics. To incorporate improvements such as developments in *b*-tagging and event simulation, a legacy analysis of the $t\bar{t}H(H \rightarrow b\bar{b})$ process in the full ATLAS Run 2 Dataset of $\mathcal{L} = 139 \, \text{fb}^{-1}$ is currently ongoing.

Modelling differences between Monte Carlo samples of the dominant $t\bar{t}$ + jets background process were found to be one of the most significant sources of uncertainty in previous analysis rounds. Along with investigating and mitigating the source of these modelling differences via generator studies of new $t\bar{t}$ + jets background simulation setups or improving event classification performance to decrease background contamination in signal regions, the usage of adversarial machine learning techniques to select robust features could decrease the impact of background modelling systematics on the fit performance. This talk will present ongoing efforts concerned with developing such adversarial machine learning approaches.

T 35.6 Tue 17:30 T-H20

Performance Tests of tH(bb) Signal and Background Separation Using a Binary Classifier Neural Network with ATLAS Data — Igor Boyko¹, Nazim HUSEYNOV¹, OKSANA KOVAL¹, •MARCEL PATZWAHL², and ANDRÉ SOPCZAK² — ¹JINR Dubna — ²CTU in Prague

The production of a Higgs boson in association with a single top quark is a strongly suppressed process in the Standard Model (SM). In the current ATLAS

T 36: Higgs Boson: Extended Models 1

Time: Tuesday 16:15-18:30

T 36.1 Tue 16:15 T-H21

Search for heavy Higgs bosons decaying to top quark pairs using the CMS experiment — AFIQ ANUAR, ALEXANDER GROHSJEAN, •JONAS RÜBENACH, DO-MINIC STAFFORD, and CHRISTIAN SCHWANENBERGER — DESY, Hamburg, Germany

The discovery of the Higgs boson at the Large Hadron Collider in 2012 marked a major breakthrough for particle physics, as it permits the verification of the Higgs mechanism, a central building block of the Standard Model. However, the Standard Model still lacks explanation for many phenomena we observe throughout the universe, including dark matter. For a great number of proposed extensions, such as the minimal supersymmetric standard model, a key ingredient is the existence of additional Higgs bosons. Using data collected by CMS at the LHC at $\sqrt{s} = 13$ TeV, corresponding to a luminosity of 138 fb⁻¹, a search is performed for scalar and pseudoscalar, electrically neutral bosons decaying predominantly to top quark pairs, which are assumed to further decay dileptonically. The challenges connected to this particular search, such as interference with the standard model background and unknown quantities resulting from neutrino momenta, are tackled by a full reconstruction of the top quark system and the utilization of multi-dimensional distributions arising from mass and spin information.

T 36.2 Tue 16:30 T-H21

Exotic Higgs Decays: ATLAS Search for Higgs Decays to Two Light Scalars — •JUDITH HÖFER, CLAUDIA SEITZ, RICKARD STRÖM, and BEATE HEINEMANN — DESY, Hamburg, Germany

Extensions of the SM Higgs sector featuring one or several singlet scalar fields are realised in many BMS models. While several searches have been performed targeting decays of the SM Higgs boson to two light spin-zero particles of the same mass, the decay to two new scalars of different mass is largely unexplored. The successive decays of these particles can give rise to spectacular high-multiplicity

data set of 140 fb⁻¹, the SM expected production rate is below the experimental sensitivity. Thus, observing such a tH production would indicate new physics. The absolute ttH coupling strength was already measured and the tH process can in addition measure the relative sign of the ttH coupling. Therefore, observing the tH process gives an important additional insight into the physics of the Higgs mechanism. Owing to the low production rate, it is particularly important to enhance the signal sensitivity, and a Neural Network (NN) is used. The resulting significance is studied by varying the NN structure. Based on simulated data, the performances of these different NN structures were tested and results are expressed as area under the ROC curve to quantify the signal and background separation.

T 35.7 Tue 17:45 T-H20

Higgs Boson Mass Reconstruction in the tH Multi-lepton Channel Using AT-LAS Data — Igor Boyko¹, Adam Herold², Nazim Huseynov¹, Lars Kolk³, Jan Kybic², André Sopczak², Petr Urban², and •Cyrus Walther³ — ¹JINR Dubna — ²CTU in Prague — ³TU Dortmund

The Higgs boson mass is reconstructed in single top production in association with a Higgs boson, tH, using a regression neural network approach. The reconstruction of the Higgs boson mass is expected to show discrimination to background processes. A focus lies on the lepton association. For the lepton association, a classification neural network is used. Hyperparameter optimization, as well as feature importance studies, are applied in order to increase the neural network performance. For the Higgs boson mass reconstruction, a hyperparameter optimization is also performed. The performance of the network is tested on tH signal and tZ background simulations.

T 35.8 Tue 18:00 T-H20

Associated production of a Higgs boson and a single top quark from t-channel production (tHq) in channels with hadronically decaying tau leptons at AT-LAS — •TANJA HOLM and IAN C. BROCK — Physikalisches Institut Universität Bonn

Associated Higgs boson production gives us the opportunity to study its couplings to fermions and bosons. An especially interesting but challenging channel is the associated production with a single top quark, as it allows one to probe the relative coupling to both kind of objects. The downside to this is a small predicted cross-section and a complicated final state including jets from light quarks or gluons, jets containing b-hadrons, missing E_T and leptons. The decay into tau leptons which subsequently decay hadronically was chosen as it has a relatively high Higgs decay branching ratio, while having a lower background than hadronic processes with higher branching ratios. This talk will discuss the search for this channel in the Run 2 LHC dataset by ATLAS.

Models

Location: T-H21

collider signatures, including so-called cascade decays, where the heavier of the scalars decays into the lighter one. The talk discusses an analysis searching for scalar decays to multi-b final states with the ATLAS experiment at the Large Hadron Collider, CERN. The analysis focuses on the ZH production mode and the channel where the scalars decay to b-quarks, resulting in a challenging low-pT jet final state. These signatures motivate the use of many novel reconstruction techniques, such as the reconstruction of soft secondary vertices, a newly developed low-pT X->bb tagger, and an event hypothesis neural network to accurately identify the Higgs decay to the light scalars among the reconstructed objects.

T 36.3 Tue 16:45 T-H21

Search for DiHiggs production $H \rightarrow hh_S$ in an extended NMSSM Higgs sector with CMS — •Martin Marz, Felix Heyen, Ulrich Husemann, Nikita Shadskiy, Michael Wassmer, and Roger Wolf — Institute of Experimental Particle Physics (ETP), Karlsruhe Institute of Technology (KIT)

The search for new physics is essential to validate or exclude theoretical models. For this purpose, models including additional non-standard-model-like Higgs bosons are of great interest. The Next-to-Minimal Supersymmetric Standard Model (NMSSM) for example predicts such additional Higgs bosons. Especially the decay of a heavy Higgs boson H into two lighter Higgs bosons (h_s , h), allows for a rich experimental program. Here h is the observed Higgs boson with properties as expected by the SM.

The analysis presented in this talk is intended to use a general rereconstruction of the LHC run 2 dataset from 2018 to profit from the best understanding of physics objects measured and reconstructed with the CMS detector in the search of new physics. The decay channel $h \rightarrow bb$, $h_S \rightarrow \tau \tau$ and the subsequent decay of the tau pair into a muon and hadrons is studied to add to a recently published analysis (JHEP 11 (2021) 057) by the CMS collaboration.

T 36.4 Tue 17:00 T-H21

Verwendung parametrischer neuronaler Netzwerke bei der Suche nach neuer Physik im Rahmen von NMSSM inspirierten Modellen — •RALF SCHMIEDER, Markus Klute, Günter Quast, Roger Wolf, Sebastian Brommer, Ma-XIMILIAN BURKART, FELIX HEYEN UND TIM VOIGTLÄNDER — KIT, Karlsruhe, Deutschland

Ein parametrisches neuronales Netzwerk (pNN) ist äquivalent zu einer Folge einzelner, verwandter NNs, von denen jedes eine eigene Aufgabe erfüllt. Diese Äquivalenz wird erreicht, indem der Raum der Eingangsparameter des pNNs, im Vergleich zu den einzelnen NNs, um zusätzliche Modellparameter erweitert wird. Das pNN erfüllt dann, abhängig von diesen Modellparametern, die Aufgaben der einzelnen NNs. Ein typisches Beispiel für den Einsatz eines pNN in der Teilchenphysik ist die Suche nach einem neuartigen Teilchen mit unbekannter Masse. In diesem Fall wird der aus physikalischen Observablen bestehende Raum der Eingangsparameter der einzelnen NNs um den Modellparameter der Masse des neuen Teilchens für die jeweils zu testende Signalhypothese erweitert.

Dieser Vortrag behandelt Studien zu pNNs im Kontext einer durch das NMSSM inspirierten Analyse der Daten des CMS Experiments. Gesucht wird nach dem Zerfall eines schweren Higgs-Bosons H in zwei leichtere Higgs-Bosonen h und h_S im Endzustand mit zwei *τ*-Leptonen und zwei b-Quarks, $H \rightarrow h(\tau \tau)h_s(bb)$, unter der Annahme von m(h) = 125 GeV. Dieses Problem besitzt zwei unbekannte Massen m(H) und $m(h_S)$, die von der zu testenden Signalhypothese abhängen und beide als Modellparameter in das pNN Training eingehen sollen.

T 36.5 Tue 17:15 T-H21

Search for NMSSM inspired di-Higgs events in bb+ $\tau\tau$ final states — •FELIX Heyen, Ralf Schmieder, Sebastian Brommer, Günter Quast, Roger WOLF, NIKITA SHADSKIY, MARTIN MARZ, and MAXIMILIAN BURKHART - KIT, Karlsruhe

In the next-to-minimal supersymmetric extension of the Standard Model (NMSSM), modifications to the Standard Model Electroweak sector lead to an extended Higgs sector with a total of seven Higgs bosons. The decay of a heavy scalar Higgs boson to a light scalar Higgs boson and a Higgs boson with the properties of the discovered Higgs boson is a promising target of this extension. This talk discusses the physics motivations of the NMSSM and introduces the search for such a decay in $\tau \tau$ + bb final states. Of the possible tau lepton final states that can be considered, this search focusses on $\tau_{\rm h}\tau_{\rm h}$ final state. A simulation of the 2018 CMS data taking period is considered.

T 36.6 Tue 17:30 T-H21

Search for additional MSSM/2HDM H→bb with Run 2 CMS data — •DAINA LEYVA PERNIA — DESY, Hamburg, Germany

Some Beyond Standard Model (BSM) theories, like the Minimal Supersymmetric extension of the Standard Model (MSSM) or the Two-Higgs Doublet Model (2HDM), predict the existence of additional Higgs bosons with an enhanced coupling to bottom quarks. This talk focuses on the search for new neutral Higgs bosons decaying into b-quarks and produced in association with at least one b-quark. The analyzed data were collected by the CMS experiment at a centreof-mass energy of 13 TeV, with the latest data reprocessing. First limits on the MSSM $H \rightarrow bb$ process using these data are shown.

T 36.7 Tue 17:45 T-H21

A 96 GeV Higgs Boson in the 2HDM plus Singlet - •CHENG LI¹, STEVEN PAASch¹, Gudrid Moortgat-pick^{1,2}, Sven Heinemeyer³, and Florian Lika² – ¹DESY, Notkestraße 85, Hamburg, Germany – ²II. Institut für Theoretische Physik, Universität Hamburg, Luruper Chaussee 149, Hamburg, Ger-

In the 2HDM and N2HDM a discrete \mathbb{Z}_2 symmetry is usually present in order

Tuesday

many — ³Instituto de Física Teórica (UAM/CSIC), Universidad Autónoma de Madrid, Cantoblanco, Madrid, Spain

We discuss a \sim 3 σ signal (local) in the light Higgs-boson search in the diphoton decay mode at ~ 96 GeV as reported by CMS, together with a ~ 2 σ excess (local) in the $b\bar{b}$ final state at LEP in the same mass range. We interpret this possible signal as a Higgs boson in the 2 Higgs Doublet Model type II with an additional Higgs singlet, which can be either complex (2HDMS) or real (N2HDM). We find that the lightest \mathscr{CP} -even Higgs boson of the two models can equally yield a perfect fit to both excesses simultaneously, while the second lightest state is in full agreement with the Higgs-boson measurements at 125 GeV, and the full Higgs-boson sector is in agreement with all Higgs exclusion bounds theoretical and experimental constraints. We derive bounds on the 2HDMS and N2HDM Higgs sectors from a fit to both excesses and describe how this signal can be further analyzed at future e^+e^- colliders. We analyze in detail the anticipated precision of the coupling measurements of the 96 GeV Higgs boson at the ILC. We find that these Higgs-boson measurements at the LHC and the ILC cannot distinguish between the two Higgs-sector realizations.

T 36.8 Tue 18:00 T-H21

Dark Matter Phenomenology in Two Higgs Doublet Model with a Complex Singlet — GUDRID MOORTGAT-PICK, JUHI DUTTA, and •JULIA ZIEGLER — II. Institut für Theoretische Physik Universität Hamburg, Luruper Chaussee 149, 22761 Hamburg, Germany

Although the Standard Model is very successful, there are still open problems which it cannot explain. (e.g. dark matter, baryon asymmetry etc.) This has led to various Beyond Standard Model theories, of which Two Higgs Doublet models are very popular, as they are one of the simplest extensions and lead to a rich phenomenology. We consider a Two Higgs Doublet model extended by a complex singlet scalar, where both, the doublets, as well as the singlet obtain a vacuum expectation value (vev). The singlet serves as a dark matter candidate. This model can solve the above mentioned problems and can also provide gravitational wave signals under specific circumstances, respectively. Furthermore one could obtain additional mixing of the dark matter and Higgs sector through the singlet vev. In this work we examine the influence of the parameters of the singlet potential on the dark matter relic density and nuclear scattering cross sections. The results are then compared with constraints from experiments.

T 36.9 Tue 18:15 T-H21

Impact of the different discrete symmetries in the 2HDM and N2HDM on Domain Wall formation and its phenomenological implications - \bullet LUIS HELLMICH¹ and GUDRID MOORTGAT-PICK^{1,2} — ¹Universität Hamburg, Hamburg, Deutschland — 2 DESY, Hamburg

Domain wall formation is a consequence of spontaneously broken discrete symmetries. Stable domain walls are cosmological bad news as they are expected to dominate the energy density of the universe. One way to overcome this domain wall problem are energetically biased vacua, which can render domain walls unstable.

to avoid FCNCs, which is softly broken and hence can produce collapsing domain walls. We find that in the Type I 2HDM and N2HDM there is an additional inherent and not explicitly broken \mathbb{Z}_2 symmetry. Furthermore the N2HDM also exhibits the well-known only spontaneously broken \mathbb{Z}_2' symmetry. We want to discuss the impact of those different discrete symmetries on domain wall formation. In particular we want to show how stable domain walls for the unbroken discrete symmetries may be suppressed and analyze the phenomenological consequences when applying the resulting constraints on the 2HDM and N2HDM.

T 37: Search for New Particles 2

Time: Tuesday 16:15-18:30

T 37.1 Tue 16:15 T-H22

Search for heavy neutral leptons in decays of W bosons using a dilepton displaced vertex in \sqrt{s} = 13 TeV pp collisions with the ATLAS detector. — •CHRISTIAN APPELT and HEIKO LACKER - Humboldt University, Berlin, Germanv

We present the ATLAS search for displaced heavy neutral leptons (\mathcal{N}) using the full integrated LHC-Run 2 luminosity of 139/fb. Adding right-handed Majorana neutrinos, so-called heavy neutral leptons, to the SM Lagrangian can help explain observed phenomena such as neutrino oscillations, matter-antimatter asymmetry, and dark matter. For the first time, we test not only the single-flavor mixing scenario but also multi-flavor mixing scenarios motivated by neutrino flavor oscillation results for normal and inverted neutrino mass hierarchies. The signature involves \mathcal{N} production in W-boson decays, $W \to \mathcal{N} \mu$ or $W \to \mathcal{N} e$, and its decay into two charged leptons and a neutrino forming a displaced vertex. We interpret the search results in the \mathcal{N} coupling versus mass plane.

Location: T-H22

T 37.2 Tue 16:30 T-H22

Search for excited leptons in the contact interaction and Z decay channels with $\text{CMS}- \bullet \text{Fabian}$ Nowotny, Thomas Hebbeker, and Kerstin Hoepfner - III. Physikalisches Institut A, RWTH Aachen

The Standard Model of particle physics does not provide a comprehensive explanation for the observed hierarchy of three generations of fermions, for both leptons and quarks. A possible explanation is delivered by models postulating that quarks and leptons themselves are composite objects. Their constituents are bound by an asymptotically free gauge interaction below a characteristic scale A. Such models of compositeness predict the existence of excited lepton (l^*) and excited quark (q^{*}) states at the characteristic scale Λ of the new binding interaction. The theory allows the production of excited leptons via contact interactions in conjunction with a Standard Model lepton. Furthermore, the leptons can decav into several final states.

This talk focuses on the contact interaction and Z-boson decay channels, both

resulting in $l^* \rightarrow lq\bar{q}$ transitions where *l* represents *e* and μ . Preliminary results are presented on the 2018 proton-proton dataset corresponding to a luminosity of 59.8 fb⁻¹ at a center of mass energy of $\sqrt{s} = 13$ TeV.

T 37.3 Tue 16:45 T-H22

Search for pair-produced leptoquarks decaying into quarks of the third and leptons of the first or second generation with the ATLAS experiment at $\sqrt{s} = 13$ TeV — •VOLKER AUSTRUP and FRANK ELLINGHAUS — Bergische Universität Wuppertal

Motivated by similarities between the quark and lepton sectors, leptoquarks (LQs) are hypothetical bosons assumed to couple to quarks and leptons at the same time. First proposed in the 1980s, the initial models included couplings only within one generation. However, hints at lepton flavor universality violation observed in several B meson decay experiments such as LHCb, BaBar, and Belle have sparked a renewed interest in LQ models, particularly extensions allowing couplings to quarks and leptons of different generations. These models introduce lepton flavor violating processes - strongly suppressed in the Standard Model at tree level, thus modifying rare B meson decays. In this talk, a search for pairproduced scalar and vector LQs decaying into quarks of the third and leptons (neutral and charged) of the first or second generation is presented. The focus of the analysis is on final states with exactly one charged lepton and large amounts of missing transverse momentum. Neural networks are utilized to ensure good separation between signal and background processes across a wide range of the parameter space. Exclusion limits are presented, based on pp-collision data corresponding to 139fb⁻¹ at a centre-of-mass energy of $\sqrt{s} = 13$ TeV collected by the ATLAS experiment at the LHC between 2015 and 2018.

T 37.4 Tue 17:00 T-H22

Leptoquark production in a single τ , charm/bottom and met final state at the ATLAS detector — •PATRICK BAUER, PHILIP BECHTLE, and KLAUS DESCH — Physikalisches Institut Bonn

At B-factories, anomalies were observed in decays of the B-hadrons into $D^{(*)}$ and $K^{(*)}$, which are consistent with the hypothesis of contributions from Leptoquarks in the high GeV to low TeV range.

Therefore, the direct search for leptoquarks (LQ) got once again in focus at high energy collider experiments. So far most searches aimed at the pairproduction via strong interaction, as it is enables a almost model independent approach and is for LQ-masses below 1 TeV expected to be dominating.

However for LQ masses well above 1 TeV the single production mode becomes more relevant. The analysis presented this talk, offers the most direct approach for a search of LQ signature related to the $B \rightarrow D^{(*)} \tau v$ anomaly, as it incorporates essentially the same couplings. Furthermore the process to be investigated could be mediated by a U₁-vector LQ, which is presently widely discussed among theorists, as preferred solution to B-anomalies. It could explain the two observed anomalies within one model. The talk will motivate the analysis and present the ongoing search for vector LQ in single and pair production in final states with one τ , bottom or charm jet and large met.

T 37.5 Tue 17:15 T-H22

The LHC as Lepton–Proton Collider: Searches for Resonant Production of Leptoquarks — •DANIEL BUCHIN, MICHAEL HOLZBOCK, and HUBERT KROHA — Max-Planck-Institut für Physik, München

Searches for leptoquarks constitute an essential part of the physics programme at the ATLAS detector. These hypothetical particles couple to both leptons and quarks and are predicted by many extensions of the Standard Model such as Grand Unified Theories. In particular leptoquarks with couplings to third generation fermions are of general interest since they represent a possible solution to the tensions observed in tests of the lepton flavour universality in $b \rightarrow s$ and $b \rightarrow c$ transitions. By considering the small but non-zero lepton content of the proton due to quantum fluctuations it becomes possible to target the resonant production of leptoquarks at the LHC. Phenomenological studies indicate that this production mode yields competitive sensitivity to existing leptoquark searches.

Resonantly produced leptoquarks give rise to a lepton + jet signature. Interestingly, such a final state is currently not well covered by ATLAS and CMS. Therefore, the analysis of the resonant production is an exciting complementary approach to the current leptoquark searches that consider e.g. pair production. The talk will introduce the leptoquark models of interest for this analysis and outline its general strategy. T 37.6 Tue 17:30 T-H22

Search for Vector Like Quarks in the decay channel to top and Higgs boson with the CMS experiment — •GUILLAUME GREAU — Deutsches Elektronen-Synchrotron DESY, Hamburg, Allemagne

A search for vector like quarks (VLQ) using the Run 2 data collected by the CMS experiment is presented. The VLQs are searched in the decay channel into a top quarks and a Higgs boson, in which the Higgs boson further decays into WW. The channel with two leptons of same sign is studied, as it suppresses considerably the standard model background. First results on distributions to discriminate the signal from the background will be shown.

T 37.7 Tue 17:45 T-H22

Suche nach vektorartigen Quarks in Endzuständen mit einem Lepton, Jets und fehlendem transversalem Impuls am ATLAS Experiment — FRANK EL-LINGHAUS und •JENS ROGGEL — Bergische Universität Wuppertal

Verschiedene Modelle für Physik jenseits des Standardmodells sagen vektorartige Quarks voraus, deren rechts- und links-händige Komponenten gleichartig unter der schwachen Wechselwirkung transformieren.

Die Analyse fokussiert sich auf die Suche nach vektorartigen Top-Quarks aus Paarproduktion mit einem Zerfall in Top-Quark und Z-Boson, wobei das Z-Boson in Neutrinos zerfällt. Die betrachteten Ereignisse werden durch ein Lepton, Jets und einen hohen fehlenden transversalen Impuls im Endzustand gekennzeichnet. Weiter führen die hohen Massen der vektorartigen Quarks zu einem starken Boost der Zerfallsprodukte, was zu einer kollimierten Zerfallstopologie führt. Der Status der Analyse der ATLAS *pp* Daten bei $\sqrt{s} = 13$ TeV wird präsentiert, wobei Ausschlussgrenzen auf die Paarproduktion von vektorartigen Top- und auch Bottom-Quarks in allen Zerfallskanälen des vektorartigen Quarks in ein Boson und ein Quark gesetzt werden.

T 37.8 Tue 18:00 T-H22

Search for long-lived particles within the CMS tracker — •KARIM EL MORA-BIT, LISA BENATO, MELANIE EICH, GREGOR KASIECZKA, and KARLA PENA — Institut für Experimentalphysik, Universität Hamburg

Several theories for physics beyond the standard model (BSM) predict the existence of long-lived particles (LLPs) that have comparably long lifetimes leading to macroscopic flight distances. Higgs-portal models, for example, propose the existence of a dark sector with particles that are neutral under the standard model (SM) gauge groups. In such theories, the SM Higgs boson mixes with a dark partner and acts as a mediator between the SM and the dark sector. The SM Higgs boson could then decay to a pair of dark sector LLPs which subsequently decay to SM particles – predominantly into bottom quark-antiquark (bb) pairs.

This talk discusses searches for LLPs using data recorded with the CMS experiment at a center-of-mass energy of 13 TeV. The searches target events in which the LLPs decay into bb pairs within the CMS tracking system after flight distances ranging from micrometers up to 1 m. The signature of the signal events consists of bb originating from displaced vertices. The searches face different challenges depending on the lifetime of the LLPs. For short lifetimes, the decay products of LLPs need to be distinguished from those of SM particles, while the search for longer lifetimes requires dedicated tracking and vertex reconstruction methods. In both cases the challenges are tackled using machine learning approaches.

T 37.9 Tue 18:15 T-H22

Search for long-lived particles in the CMS calorimeters and muon chambers — •LISA BENATO, JÖRG SCHINDLER, and GREGOR KASIECZKA — Institut für Experimentalphysik, Universtät Hamburg

Many beyond the standard model (BSM) theories predict the existence of longlived particles (LLPs) that have long lifetimes and decay in the outermost parts of a hadron collider experiment, such as the calorimeters and muon chambers of the CMS detector. Very displaced signatures (decay length beyond 1 m) can only be reconstructed with non-standard approaches by using low-level detector information (hits in the muon chambers and scintillation time of the calorimeter crystals). LLP decays in calorimeters are identified as jets, delayed with regards to the proton-proton collision and with a small number of associated tracks. Muon chambers act as sampling calorimeters and LLP decays originate showers of hits in the gas detectors, identified as clusters, with no concurrent activity in the inner layers. No SM process produces this kind of signatures at a relevant rate. The expected background is nearly zero and due to detector noise and non-collision backgrounds. Such a clean environment allows to probe light LLPs with unprecedented sensitivity.

T 38: Search for New Particles 3

Time: Tuesday 16:15-18:00

T 38.1 Tue 16:15 T-H23

Searching for Axion-Photon Couplings — •ROBIN LÖWENBERG, TOM KROKOTSCH, DANIEL KLEIN, GUDRID MOORTGAT-PICK, and KRISZTIAN PETERS — Deutsches Elektronen-Synchrotron (DESY), Hamburg, Germany

The nature of dark matter is not yet known. Extensive studies have been done searching for a weakly interactive massive particle (WIMP) but are still waiting for a positive signal.

Another popular approach is assuming that dark matter consists of light particles rather than WIMPs. These particles are called weakly interactive slender particles (WISPs). One of the most promising candidates is the axion which was proposed in the seventies to solve the so-called strong-CP-problem. It later turned out to be a good dark matter candidate in a certain mass range.

Experiments searching for axions and axion-like particles like ALPS 1 (2007-2010) and ALPS 2 (since 2021) are provided by Deutsches Elektronen-Synchrotron (DESY) in Hamburg with further experiments on the way. Most of them are based on the assumption that axions show photon-like couplings and therefore would interact with electromagnetic fields.

This talk aims to explain the concept of axions and in which mass ranges and coupling strengths they might be a good dark matter candidate. Furthermore, a quick overview will be given of current and possible future experiments promising to detect the particle if it exists.

T 38.2 Tue 16:30 T-H23

The ALPS II experiment at DESY - Status and prospects — •KANIOAR KARAN for the ALPS-Collaboration — DESY / Cardiff University, Hamburg, Germany The Any Light Particle Search II (ALPS II) is a laboratory-based light shining through a wall experiment (LSW) to probe the existence of Axion-Like-Particles (ALPs) with a coupling to electromagnetic fields as low as $g_{a\gamma\gamma}\approx 2\times 10^{-11}{\rm GeV}^{-1}$ that is hinted at some astrophysical anomalies such as stellar evolutions and the TeV transparency of the universe. This LSW experiment is based on the simple idea that a high power laser field that propagates through a static magnetic field can partly oscillate into an ALP field. The ALP field then crosses an opaque wall to a second static magnetic field and can partly re-oscillates into an electromagnetic field which can be detected with a detector. In order to achieve the anticipated sensitivity, two 125m long optical cavities, operated in a vacuum system, are used: one with an expected circulating power of 150kW for the ALPs production and one with an expected power build-up of 40,000 to enhance the regeneration of the electromagnetic field. The circulating field in each cavity is directed through a string of 12 superconducting HERA dipole magnets providing a magnetic field of 5.3T. The ALPS II experiment is located at DESY in Hamburg and is currently in the commissioning phase. In this talk, we will present the current status, challenges and perspectives of the ALPS II experiment with the focus on the optical setup.

T 38.3 Tue 16:45 T-H23

Analysis and simulation of TES data in the ALPS II experiment — •JOSÉ ALEJANDRO RUBIERA GIMENO for the ALPS-Collaboration — Deutsches Elektronen-Synchrotron, Hamburg, Germany

The Any Light Particle Search II (ALPS II) is a Light-Shining-through-a-Wall experiment under construction at DESY, Hamburg. Its goal is to probe the existence of Axion Like Particles (ALPs), a possible candidate for dark matter. In the ALPS II region of study, a rate of photons reconverting from ALPs on the order of 10^{-5} cps is expected. This requires a sensor capable of measuring low-energy photons (1.165 eV) with high efficiency and a low dark count rate. We investigate a tungsten Transition Edge Sensor (TES) system as a photon-counting detector that promises to meet these requirements. This detector exploits the drastic change in its resistance caused by the absorption of a single photon when operated in its superconducting transition region at millikelvin temperatures. In this work, the analysis procedure applied to measured TES pulses, in time and frequency domain, is presented. This analysis allows extracting characteristic parameters used for signal discrimination against backgrounds. The energy resolution computed from data is compared to simulations of electronic noise superimposed with ideal photon pulses.

T 38.4 Tue 17:00 T-H23

Status and Prospects of a TES-based Detector System for ALPS II — \bullet GULDEN OTHMAN for the ALPS-Collaboration — University of Hamburg, Hamburg, Germany

The Any Light Particle Search II (ALPS II) experiment will search for QCD axions and axion-like particles (ALPs) in an important parameter space that is relevant in understanding anomalous astrophysical phenomena, including stellar evolution and dark matter. ALPS II takes advantage of the axion coupling to photons using a Light-Shining-through-a-Wall technique. Photons created using a strong laser may convert into ALPs in the presence of a strong magnetic field. The ALPs can traverse a light-tight barrier, reconvert into photons within a second magnet string, and be subsequently detected. The rate of re-converted photons is extremely low, on the order of 10^{-5} counts/second, and their observation requires the use of sensitive photon detectors with high efficiency and low backgrounds. The first stage of ALPS II, currently under construction at DESY, Hamburg, will use a heterodyne detection method. In the subsequent phase, ALPS II can utilize advances in cryogenic quantum sensing by employing Transition Edge Sensors (TESs). We are currently developing a TES-based detector system that can meet the requirements for ALPS II, offering single-photon detection with high efficiency and low-backgrounds at the 1064 nm (1.165 eV) energy of interest. In this talk, we present the feasibility, challenges, and current status of the TES-based detector system for ALPS II at DESY, Hamburg.

T 38.5 Tue 17:15 T-H23

Monte Carlo based ray tracing for BabyIAXO — •JOHANNA VON OY, KLAUS DESCH, JOCHEN KAMINSKI, TOBIAS SCHIFFER, and SEBASTIAN SCHMIDT — Physikalisches Institut der Universität Bonn

The premise of the International AXion Observatory (IAXO) and its intermediate experimental stage BabyIAXO is to detect the undiscovered particle axion, which can be a good candidate for dark matter. This will be done by utilizing the inverse Primakoff effect to reconvert them into X-rays in the magnetic field of a movable magnet. Following this, they would get focused by an X-ray optic and detected in, for example, a window sealed gaseous detector.

To simulate this whole process, ray tracing based on the Monte Carlo method is a useful tool, as a certain number of axions would be generated and get assigned different probabilities and changes of direction depending on the setup. This talk will focus on the individual steps, starting with the production in the sun and following the axion's path to the detector.

T 38.6 Tue 17:30 T-H23

Anisotropy effects in a dielectric haloscope for dark matter searches — •BERNARDO ARY DOS SANTOS for the MADMAX-Collaboration — RWTH III. Physikalisches Institut A, Aachen, Germany

The MADMAX collaboration intends to build a dielectric haloscope targeted to detect galactic axion dark matter, in the mass range $40 - 400\mu eV$. This experiment consists of a series of dielectric discs and a mirror placed inside a strong homogeneous magnetic field that would produce the emission of coherent electromagnetic radiation with a frequency related to the mass of the axion. One of the current challenges is to simulate this experiment taking into account realistic settings. We present an improved simulation that is able to include the effects of anisotropic dielectric discs in the experiment.

T 38.7 Tue 17:45 T-H23

Axion dark matter searches using superconducting radio frequency cavities — •TOM KROKOTSCH¹, ROBIN LÖWENBERG¹, DANIEL KLEIN¹, GUDRID MOORTGAT-PICK^{1,2}, and KRISZTIAN PETERS² — ¹Universität Hamburg, Hamburg, Germany — ²Deutsches Elektronen-Synchrotron DESY, Hamburg, Germany

Recent proposals of employing superconducting radio frequency (SRF) cavities to detect axions suggest promising sensitivities in previously unexplored parameter space. This includes QCD axions and axion-like dark matter with masses in a range of about eight orders of magnitude.

The setup aims to detect photons generated through the predicted interaction of axions with electromagnetic fields. An advantage of this design is that it allows scanning over a range of axion masses by slightly altering the cavity. Such a detector benefits from the high quality factors achieved in modern SRF technology.

In the talk we discuss the general idea behind the detector, it's most significant parameters to achieve a high signal power and evaluate which cavity geometries and transitions are the most promising. Particular attention must be paid to minimise various noise sources which limit the sensitivity of the setup.

Location: T-H23

T 39: Gaseous Detectors 2

Time: Tuesday 16:15-18:15

T 39.1 Tue 16:15 T-H24

Development of gas and high voltage systems for a small-strip Thin Gap Chamber quadruplet — •JOSÉ ANTONIO FERNÁNDEZ PRETEL, KSENIA SOLOVIEVA, BERNHARD PFEIFER, JÜRGEN TOBIAS, PATRICK SCHOLER, UL-RICH LANDGRAF, and VLADISLAVS PLESANOVS — Albert-Ludwigs Universität Freiburg

In the ATLAS detector, a good performance of the trigger and tracking systems is needed to ensure its physics program. For this purpose, the end-cap muon system has been upgraded by installing the so-called New Small Wheel. The small-strip Thin Gap Chambers (or sTGCs for short, one of its main technologies) are multi-wire proportional counters running with a working mixture of CO_2 :n-pentane (55:45) and a high voltage about 3kV for signal amplification. For the test setup being designed in Freiburg, a gas and high voltage systems are needed to run sTGCs, whose optimal operation needs to be guaranteed in real time via monitoring. In this presentation, the mixing, delivery and storage components of the gas system, the high-voltage and monitoring systems to run sTGCs in Freiburg are discussed.

T 39.2 Tue 16:30 T-H24

Cosmic Test Stand Studies with a small-strip Thin Gap Chamber quadruplet — •KSENIA SOLOVIEVA, JOSE ANTONIO FERNANDEZ PRETEL, PATRICK SCHOLER, VLADISLAVS PLESANOVS, and ULRICH LANDGRAF — Albert-Ludwigs University, Freiburg

The small-strip Thin Gap Chamber (sTGC) technology is being implemented in the New Small Wheel upgrade of ATLAS for improved triggering and tracking in a higher particle rate environment. The sTGC detector readout includes a pad segmentation, which plays a key role in the trigger chain. For the purpose of investigating readout and trigger parameters, a quadruplet was set up in a cosmic muon test stand in Freiburg and read out with the final ATLAS NSW readout system and the final gas mixture. Another quadruplet was also tested in the gamma irradiation facility (GIF++) at CERN with a muon beam. The results of this study will be used for comparison with the results from the local setup. This presentation discusses the goals and challenges of the dedicated setup, as well as presenting the prospective results.

T 39.3 Tue 16:45 T-H24

Study of the Position Resolution of Large Scale Micromegas Detectors — •VLADISLAVS PLESANOVS, PATRICK SCHOLER, ULRICH LANDGRAF, and GREGOR HERTEN — University of Freiburg

During the current Long Shutdown 2 (LS2), the ATLAS muon spectrometer will receive an upgrade with the New Small Wheels (NSW), which consist of two new technologies, the sTGC detector (trigger) and the Micromegas detector (tracking).

The performance parameters of the Micromegas detectors, such as efficiency, gain, and position resolution, have been studied in the past with cosmic muons and test beams. In a recent study using cosmic muons, which forms the basis for this presentation, an attempt was made to improve the position resolution using a charge-weighted mean time correction.

T 39.4 Tue 17:00 T-H24

Clock phase calibration of the readout controller of the NSW — •JONAS ROEMER¹, ANNE FORTMAN², MICHELLE SOLIS³, and JARED STURDY¹ — ¹Department of Physics and Astronomy, University of California, Irvine — ²Harvard University — ³University of Arizona

ATLAS introduces a new muon detector system for the upcoming Run 3 of the LHC: The New Small Wheel (NSW). The NSW features two detector technologies, the Micromegas and the small Thin Gap Chambers. The front-end boards of both technologies use the same readout controller ASIC (ROC). The purpose of the ROC is, among other things, to receive and route trigger, timing and control (TTC) signals, and to collect the channel hit data and transmit it to the readout system.

The ROC samples the TTC stream with a configurable 40 MHz clock. The phase of this clock must be calibrated to interpret the TTC words correctly. Furthermore, the ROC has two 160 MHz clocks for internal functionality and for sampling the hit data from the VMMs. Both must be calibrated relative to each other.

This talk outlines the electronics architecture and calibration procedure and presents the results.

Location: T-H24

T 39.5 Tue 17:15 T-H24

Gas Monitoring Chambers for the T2K Near Detector Upgrade — PHILIP HAMACHER-BAUMANN, INES HANNEN, LEON MANS, THOMAS RADERMACHER, STEFAN ROTH, DAVID SMYCZEK, and •NICK THAMM — RWTH Aachen University - Physics Institute III B, Aachen, Germany

Measurements from the Tokai to Kamioka (T2K) long baseline neutrino oscillation experiment have shown first indications for CP violation in the leptonic sector. To improve these results, a part of the near detector (ND280) of T2K will be replaced by a 3D fine-grained scintillator tracker acting as target (SFGD), two high-angle time projection chambers (HA-TPCs) and a time-of-flight system (TOF). The Aachen group is developing and constructing new gas monitoring chambers for the existing and new TPC systems to be installed in 2023. These chambers are dedicated to the continuous calibration and monitoring of the TPC drift gas properties. This talk gives an overview of the design, simulation, comissioning and series production of the new gas monitoring system.

T 39.6 Tue 17:30 T-H24

Electronics of the New T2K Gas Monitoring Chambers — PHILIP HAMACHER-BAUMANN, THOMAS RADERMACHER, STEFAN ROTH, •DAVID SMYCZEK, and NICK THAMM — RWTH Aachen University - Physics Institute III B, Aachen, Germany

A pair of High Angle Time Projection Chambers (HATs) will be installed during the upgrade of the T2K near detector ND280. For the calibration of the new HATs the gas parameters will be continuously monitored. For this task, new Gas Monitoring Chambers (GMCs) are developed by the Aachen group and currently under construction. The electronics of the GMC consist of a preamplifier for the anode signals, a triggering system using SiPMs, a data acquisition system using VME waveform digitizers and a slow control logging parameters like pressure and temperature. The design and the commissioning of the electronics system will be presented.

T 39.7 Tue 17:45 T-H24

Hydrogen-rich Gases for High Pressure Time Projection Chambers at Neutrino Beamlines — •PHILIP HAMACHER-BAUMANN, THOMAS RADERMACHER, STEFAN ROTH, and NICK THAMM — Physics Institute III B, Aachen, Germany DUNE's near detector complex foresees a magnetized high-pressure gaseous time projection chamber (HPgTPC) as part of its detector suite. The gaseous active volume boasts a very low detection threshold with high particleidentification power and large acceptance for tracking. Especially interactions on the gas itself in the high intensity neutrino beam will be collected with an unmatched rate. For design and development of a pressurized TPC, it is essential to quantify and validate electron drift parameters, to predict performance of the final detector, e.g. HPgTPC. This presentation investigates how electron drift parameters of drift gas mixtures perform at higher than atmospheric pressures. Additionally, a study of hydrogen-rich Argon:Methane gas mixtures for consideration in HPgTPC is presented using measurements from a test chamber.

T 39.8 Tue 18:00 T-H24

Boron-based neutron Time Projection Chamber – •DIVYA PAL¹, JOCHEN KAMINSKI¹, MICHEAL LUPBERGER¹, MARKUS KÖHLI^{1,2}, KLAUS DESCH¹, MARKUS GRUBER¹, SAIME GÜRBUZ¹, and LAURA RODRIGUEZ GÓMEZ¹ – ¹Physikalisches Institut, Universität Bonn – ²Physikalisches Institut, Universität Heidelberg

Thermal neutron detection is crucial in various areas ranging from fundamental physics research to national security, crystallography and medicine. Tradionally, thermal neutron detectors use Helium-3 filled proportional counters. However, due to the supply shortage of Helium-3, leading to a rapid increase in its price, alternative detectors are sought.

In Bonn, the BOron DEtector with Light and Ionization Reconstruction (BODELAIRE) is being developed to provide high spatial and time resolution in thermal neutron detection. The BODELAIRE is based on the principle of a Time Projection Chamber (TPC) with thin layers of Boron-10 neutron converters placed perpendicular to a GridPix readout which will have Timepix3 as ASIC. The trigger is placed along the field cage, consisting of multiple layers: Boron, scintillator and light readout. Thus, the working principle is that the conversion of the neutron with Boron-10 gives two tracks, one giving a trigger signal in the scintillator while the other leaves a track in the gas volume.

The concept and current development status of the BODELAIRE will be presented.

T 40: Pixel Detectors 2

Time: Tuesday 16:15-18:30

T 40.1 Tue 16:15 T-H25

Tangerine: Monte Carlo simulations of MAPS in a 65 nm imaging process — •MANUEL ALEJANDRO DEL RIO VIERA, HÅKAN WENNLÖF, and ADRIANA SIMAN-CAS for the Tangerine-Collaboration — Deutsches Elektronen-Synchrotron (DESY)

The Tangerine project's goal is to develop the next generation of small collection electrode monolithic silicon pixel detectors using the 65 nm CMOS imaging process. This offers a higher logic density and overall lower power consumption compared to previously used processes. In monolithic sensors the sensitive volume and readout are in a single chip, which enables a lower material budget, and reduced cost and production effort compared to hybrid sensors.

In order to understand the processes and parameters that are involved in the developments in the new 65 nm technology, a combination of TCAD and Monte Carlo (MC) simulations are used. Allpix Squared utilizes the realistic electric fields and doping profiles provided by the TCAD simulations and by the use of MC methods, obtains results that can later be compared to experimental data from test beam experiments.

This presentation will cover the design and setup of the Monte Carlo simulations and present the results obtained so far.

T 40.2 Tue 16:30 T-H25

Tangerine project - Studies of MAPS prototypes in CMOS 65 nm technology — •GIANPIERO VIGNOLA for the Tangerine-Collaboration — Deutsches Elektronen-Synchrotron, Hamburg, Germany

CMOS Monolithic Active Pixel Sensors, widely used as imaging devices, represent an attractive alternative to hybrid pixel detectors in High Energy Physics. The Tangerine project at DESY aims to develop a fully integrated 65 *nm* CMOS pixel sensor for future application, from beam-test facilities to Higgs factories. The goal is to achieve time resolutions of the order of 1 ns and spatial resolution below 3 μ m. First 65 *nm* CMOS test chips with 4 pixels of 16 μ m pitch and analog readout have been investigated in beam-test studies and with an Iron-55 source. Results of detailed waveform analysis, characterizing the charge sensitive amplifier, will be reported. The outcome of these studies will be used to improve the sensor layout and its signal-processing circuitry in the next Tangerine prototype.

T 40.3 Tue 16:45 T-H25

Device simulations of a MAPS developed in 65nm CMOS Imaging Technology — •ADRIANA SIMANCAS for the Tangerine-Collaboration — Deutsches Elektronen-Synchrotron, Hamburg, Deutschland

Monolithic CMOS sensors have found their way through imaging technologies into High Energy Physics thanks to multiple advantages in particle detection. Their main characteristic is the integration of an active sensor and readout in the same silicon wafer, which provides a reduction in production effort, costs and material. The Tangerine project aims to develop the next generation of silicon pixel sensors for lepton colliders using a 65 nm CMOS imaging technology with a small collection electrode. It offers a significant improvement in the logic density of the pixels, the power consumption, the material budget and the S/N in comparison to previously studied technologies. Since the electric fields in monolithic sensors are quite complex, device simulations are needed to develop an understanding of this technology and provide important insight into performance parameters of the sensor. TCAD is a very powerful tool that allows to simulate the electrical properties of semiconductors. Herewith, it is possible to optimize the sensor layout and other features to achieve excellent time and spatial resolution. This contribution will present the latest developments in device simulation of a 65 nm CMOS sensor with a small collection electrode using TCAD.

T 40.4 Tue 17:00 T-H25

Simulating Hexagonal Monolithic Pixel Sensors in CMOS Imaging Technology — •LARISSA MENDES for the Tangerine-Collaboration — Deutsches Elektronen-Synchrotron (DESY), Hamburg, Germany

In this research, monolithic pixel sensors in CMOS imaging technologies with small collection electrodes are investigated for fast signal collection and precise timing to be used at future lepton colliders. Integrated monolithic CMOS offers a cost-effective monolithic integration of sensor and electronics. It allows significant reductions in the material budget compared to hybrid pixel detectors, providing an excellent signal-to-noise ratio and position resolution.

When compared to a square grid, the distance between the pixel border and the collection electrode can be reduced by placing the collection electrodes on a hexagonal grid for a given area defined by the circuit functionality. The hexagonal grid reduces charge sharing, increasing the signal in the seed pixel. The studies are being carried out with 3D Synopsys' Technology Computer-Aided Design (TCAD) to create a sensor structure that can optimize the depletion region size and increase the lateral electric field for fast charge collection by drift. The electrode size, electrode size, as well as other parameters of the sensor design, are optimized using 3D TCAD simulations. Location: T-H25

T 40.5 Tue 17:15 T-H25

Guard-ring optimisation for sensors in LFoundry 150nm CMOS technology — •SINUO ZHANG, TOMASZ HEMPEREK, and JOCHEN DINGFELDER — Physikalishes Institut, Rheinishe Friedrich-Welhem Universität Bonn, Nussallee 12, 53115 Bonn, Germany

In high energy physics, the silicon pixel sensors manufactured in commercial CMOS chip fabrication lines have been proven to have a good radiation hardness and spatial resolution. Along with the mature manufacturing techniques and the potential of large throughput provided by the foundries, the so-called "passive CMOS" sensor has become an interesting alternative to standard planar sensors, in particular for large-area applications. High and predictable breakdown behaviour for pre- and post-irradiation is a major design goal for sensors and the guard-ring structure is one factor to optimise. This is especially important for applications that require higher voltages.

We present several concepts of the guard-ring design that can be realised in LFoundry 150nm CMOS technology. As was studied with TCAD simulations, such designs can lead to a higher breakdown voltage by modifying the potential and electric field distribution in the guard-ring area. A number of test structures have been designed for the RD50 MPW-3 and the CMS CROC submission for verifications and further studies.

T 40.6 Tue 17:30 T-H25

Monte Carlo simulations of a beam telescope setup based on the 65nm CMOS Imaging Technology — •SARA RUIZ DAZA for the Tangerine-Collaboration — Deutsches Elektronen-Synchrotron, Hamburg, Germany

Monolithic CMOS sensors enable the development of detectors with a low material budget and a low fabrication cost. Moreover, using a small collection electrode results in a small sensor capacitance, a low analogue power consumption, and a large signal-to-noise ratio. These characteristics have become very attractive in the development of new silicon sensors for charged particle tracking at future experiments. A beam telescope setup consisting of detector prototypes designed in a novel 65 nm CMOS imaging process is being simulated. This contribution describes the first steps and verifications in the design of such a telescope using the Allpix Squared and Corryvreckan frameworks for simulation and analysis.

T 40.7 Tue 17:45 T-H25

The Allpix Squared pixel detector simulation framework — •HÅKAN WENNLÖF, SIMON SPANNAGEL, and PAUL SCHÜTZE — DESY, Notkestrasse 85, 22607 Hamburg, Germany

The Allpix Squared sensor simulation framework is a modular and flexible opensource tool for simulating pixel detectors. The framework has the capability of simulating the full detector chain, from energy deposited by incident particles to signal formation and digitisation. Both single detectors and more complex setups, such as testbeam telescope experiments, can be investigated in great detail. Through its interface to GEANT4, Allpix Squared has access to advanced physics models and particle sources. The framework can also import fields and doping maps from technology computer-aided design (TCAD) simulations, creating a detailed simulation of individual charge carrier behaviour in the investigated sensor models. This yields a powerful combination of TCAD and Monte Carlo simulations, which provides accurate high-statistics results while also accounting for stochastic fluctuations in the involved processes. This method has been used in comparing simulations with testbeam data, and finding a good match between results. Through the detailed electric field inclusion, time-dependent charge pulse formation can also be simulated in the framework.

Allpix Squared is currently used as part of developments of several state-of-the art sensors, for example in the Tangerine project at DESY. This contribution will present the current status of the Allpix Squared framework, and give examples of use cases.

T 40.8 Tue 18:00 T-H25

Development and characterization of a DMAPS chip in TowerJazz 180 nm technology for high radiation environments — IVAN BERDALOVIC², •CHRISTIAN BESPIN¹, JOCHEN DINGFELDER¹, TOMASZ HEMPEREK¹, TOKO HIRONO^{1,3}, FABIAN HÜGGING¹, HANS KRÜGER¹, THANUSHAN KUGATHASAN², CESAR AUGUSTO MARIN TOBON², KONSTANTINOS MOUSTAKAS¹, HEINZ PERNEGGER², WALTER SNOEYS², TIANYANG WANG¹, and NORBERT WERMES¹ — ¹Universität Bonn, Bonn, Deutschland — ²CERN, Genf, Schweiz — ³DESY, Hamburg, Deutschland

The increasing availability of commercial CMOS processes with high-resistivity wafers has fueled the R&D of depleted monolithic active pixel sensors (DMAPS) for usage in high energy physics experiments. One of these developments is a series of monolithic pixel detectors with column-drain readout architecture and small collection electrode facilitating low-power designs: the TJ-Monopix series.

TJ-Monopix is designed in a 180 nm TowerJazz CMOS process and features a

pixel size of 33 um * 33 um. Due to improvements on the front-end electronics and sensor design of the current iteration TJ-Monopix2 the radiation hardness and efficiency could be increased while lowering the threshold and noise. Results from laboratory measurements and test beam campaigns will be presented to discuss the suitability of TJ-Monopix2 for use in high-radiation environments.

T 40.9 Tue 18:15 T-H25

Characterization of depleted monolithic active pixel sensors (DMAPS) designed in 150nm CMOS technology — •LARS SCHALL¹, JOCHEN DINGFELDER¹, CHRISTIAN BESPIN¹, IVAN CAICEDO¹, TOMASZ HEMPEREK¹, TOKO HIRONO^{1,2}, FABIAN HÜGGING¹, HANS KRÜGER¹, PIOTR RYMASZEWSKI¹, TIANYANG WANG^{1,3}, and NORBERT WERMES¹ — ¹Physikalisches Insitut, University of Bonn — ²DESY, Hamburg — ³Zhangjiang Laboratory, China

Monolithic active pixel sensors with depleted substrates are a promising option for pixel tracker detectors in high radiation environments. The use of a highly

T 41: Calorimeters 1

Time: Tuesday 16:15-18:30

T 41.1 Tue 16:15 T-H26

Scintillator studies for the CALICE Analogue Hadronic Calorimeter and the DUNE ND-GAr Electromagnetic Calorimeter — ANDREA BROGNA², PE-TER BERNHARD², VOLKER BÜSCHER¹, KARL-HEINZ GEIB¹, ASMA HADEF¹, •ANTOINE LAUDRAIN¹, LUCIA MASETTI¹, MARISOL ROBLES MANZANO¹, ANNA ROSMANITZ¹, CHRISTIAN SCHMITT¹, ALFONS WEBER¹, and QUIRIN WEITZEL² for the CALICE-D-Collaboration — ¹Institut für Physik, Johannes Gutenberg Universität, Mainz — ²PRISMA Detector Lab, Johannes Gutenberg Universität, Mainz

The CALICE Analogue Hadronic Calorimeter (AHCal) is designed as an imaging calorimeter optimised for particle flow algorithms. It features a high granularity and timing resolution, with plastic scintillator active layers and SiPM readout. The concept of Megatile is developed to ease the assembly of a large scale detector, separating a large piece of scintillator with a glue + TiO_2 mixture and gluing a reflective sheet on both sides instead of individually wrapping smaller tiles in a reflective foil for each channel.

In the DUNE ND-GAr, neutron energy is reconstructed using the time-offlight information provided by the Electromagnetic Calorimeter (ECal). The ECal must therefore provide neutron identification with high granularity and timing, making the AHCal with Megatiles a potential design candidate. Pulse shape discrimination is envisaged as a means to discriminate photon from neutron interaction in the ECal. A setup to study the PSD performance of various plastic scintillators with SiPM readout has been set up and first results will be presented.

T 41.2 Tue 16:30 T-H26

Exploring the intrinsic Time Resolution of the SiPM-on-Tile Technology — •FABIAN HUMMER, LORENZ EMBERGER, and FRANK SIMON for the CALICE-D-Collaboration — Max-Planck-Institut für Physik

The SiPM-on-Tile technology, where small plastic scintillator tiles are directly read out with SiPMs, has been developed for the CALICE Analog Hadron Calorimeter (AHCAL), and has been adopted for parts of the hadronic section of the CMS HGCAL. For future electron-positron colliders, a single cell time stamping on the sub-nanosecond level for energy deposits corresponding to single minimum-ionizing particles is desired to provide background rejection and to support pattern recognition and energy reconstruction with particle flow algorithms. To better understand the intrinsic time resolution achievable with the SiPM-on-Tile technology, detailed measurements have been performed in test beams at DESY, probing different scintillator tile sizes and materials. The study is complemented by laser measurements that provide insights into processes within the scintillator tile relevant for timing. Geant4 simulations allow us to verify our results and to find the correlations between scintillator tile size, light yield and time resolution. In this contribution, we will discuss our measurement methods, the results of our SiPM-on-Tile timing study and the implementation and performance in simulations.

T 41.3 Tue 16:45 T-H26

Latest Tests of CMS HGCAL Tilemodule Prototypes — •MALINDA DE SILVA, MATHIAS REINECKE, KATJA KRÜGER, OLE BACH, and FELIX SEFKOW — Deutsches Elektronen-Synchrotron (DESY), Hamburg, Germany

For the HL-LHC phase, the calorimeter endcap of the CMS detector will be upgraded with a High Granularity Calorimeter (HGCAL), a sampling calorimeter that will use silicon sensors as well as scintillator tiles read out by silicon photomultipliers (SiPMs) as active material (SiPM-on-tile). The complete HGCAL will be operated at -30 °C. The SiPMs will be used in areas where the expected radiation dose during the lifetime of the detector is up to $5 \times 10^{13} n_{eq}/cm^2$. The design of the SiPM-on-tile part is inspired by the CALICE AHCAL.

resistive silicon substrate and short drift paths enhance the radiation tolerance, while a careful guard ring design facilitates high biasing voltages to deplete the sensor.

LF-Monopix2 is the latest prototype of a DMAPS development in 150 nm CMOS technology. It features a fully functional column-drain readout architecture in a 2x1 cm² matrix. A reduced pixel pitch of 50x150 μ m² compared to its predecessor results in a smaller detector capacitance and an improved spatial resolution. Each pixel's digital electronics are integrated within the large collection electrode. Optimization of the pixel layout minimizes potential cross talk from the digital transients into the sensor node.

LF-Monopix2 chips have successfully been thinned to a thickness of 100 μ m while their breakdown voltage remained above 350 V. In this talk, the ongoing characterization and preliminary results of the first test-beam campaign of these sensors are discussed.

Location: T-H26

The basic detector unit in the SiPM-on-tile part is the tile module, consisting of a PCB with one or two HGCROC ASICs, reading out up to 96 tiles with SiPMs. To acquire the data as well as to send the fast and slow control commands, monitor temperature and voltages from the tile modules a dedicated DAQ system has been designed and implemented. This DAQ system was tested alongside the latest generation of tile modules at the October 2021 test beam at DESY as well as tests at -30 °C were conducted using a climate chamber. Results from these tests will be reported.

T 41.4 Tue 17:00 T-H26

Test beam results of the Megatile prototype for the CALICE AHCAL — •ANNA ROSMANITZ for the CALICE-D-Collaboration — Johannes Gutenberg-Universität Mainz

The CALICE collaboration is developing a highly granular Analog Hadronic Calorimeter (AHCAL) for a future e^+e^- linear collider.

The current design for the AHCAL consists of small, separately produced scintillator tiles with a size of 3x3 cm² read out by silicon photomultipliers (SiPM). They are separately wrapped in reflective foil and glued to the boards. In total, the AHCAL is going to consist of around 8 million tiles.

To facilitate the assembly process, a new mechanical design is currently under development. The new concept, called Megatile, is based on a larger scintillator plate, in which the optical separation is obtained via trenches filled with a glue and TiO₂ mixture for reflectivity and optical insulation. The Megatile is enclosed with reflective foil, the edges are covered with varnish.

This talk presents the performance in terms of light yield and cross talk of the latest megatile prototype, based on test beam data from recent campaigns at DESY.

T 41.5 Tue 17:15 T-H26

Timing analysis of testbeam data with the KLauS6 ASIC for a Silicon Photomultiplier on Scintillator Tile Setup — •ERIK WARTTMANN for the CALICE-D-Collaboration — Kirchhoff-Institute for Physics, Heidelberg

The CALICE collaboration is developing highly-granular scintillator-based calorimeters with the need of sophisticated readout solutions. For this purpose, the KLauS ASIC has been developed. It is designed to dissipate very little power and features precise charge measurement optimized for low-gain SiPMs, covering their full dynamic range. Additionally, the sixth iteration of the chip provides timing information via a Phase-Locked-Loop (PLL) driven TDC with a bin size of 200 ps. The chip has been evaluated in a SiPM-on-Tile setup at the DESY testbeam facility, using an arrangement of four layers of scintillating tiles air-coupled to SiPMs. Various calibration measurements have been carried out to ensure precise charge and time measurements. The influence of multiple chip parameters on timing at the single MIP level has been investigated. To study the dependence of timing on the deposited energy, measurements with absorbers have been carried out. For the analysis of single MIP and absorber data, a stable and versatile timewalk correction has been implemented. We present the time measurement capabilities of KLauS6 using the SiPM-on-Tile arrangement, discuss the calibration routines and critical ASIC parameters. The results from single MIP runs and the absorber data regarding the time measurement are presented, which are comparable to the intrinsic resolution of the setup.

T 41.6 Tue 17:30 T-H26

Apply Computer Vision Algorithm to High Granularity Calorimetry — •JULIAN UTEHS and STAN LAI for the CALICE-D-Collaboration — II. Institute of Physics, Georg-August-University Göttingen, Germany

In the course of the development of the new ILC detector, there is extensive research towards high granularity calorimeters. The CALICE collaboration has developed a prototype, the Analog Hadron Calorimeter, which uses SiPM technology to read out highly granular scintillator tiles. Test beam data provides unprecedented opportunities to investigate particle shower reconstruction with highly granular calorimetry.

This talk will concentrate on the opportunities of shower shapes analysis, that are given by the usage of well-known computer vision algorithms. These algorithms, coming from object detection and industry robotics and automation, will be applied on AHCAL calorimetry data. Their application allows for opportunities to reconstruct and resolve sub-shower activity and possibly estimate the electromagnetic fraction of hadronic showers. This can be input into energy reconstruction and particle identification techniques, and possible improvements for these are investigated.

T 41.7 Tue 17:45 T-H26

New shower direction reconstructing calorimeter — •MATEI CLIMESCU and RAINER WANKE — Johannes Gutenberg Universität Mainz

The so-called SplitCAL detector is a new design of a mixed electromagnetichadronic calorimeter which provides both energy reconstruction through layers of scintillating strips, read-out with wavelength shifting fibres and SiPMs, and shower direction information through high-precision layers. This can be used for fixed target experiments which require high geometrical precision and directional reconstruction of photon showers. The development needs to account for low rates but large dynamic range. The whole concept is presented with specific focus on the link of the scintillating fibres to the SiPMs and the readout.

T 41.8 Tue 18:00 T-H26

Convolutional Neural Networks for the Energy Reconstruction of AT-LAS Liquid-Argon Calorimeter Signals — ANNE-SOPHIE BERTHOLD, NICK FRITZSCHE, •CHRISTIAN GUTSCHE, MAX MÄRKER, JOHANN CHRISTOPH VOIGT, and ARNO STRAESSNER — Institut für Kern- und Teilchenphysik, TU Dresden, Germany

In 2027, it is planned to start the High-Luminosity LHC, which will push the possibilities of research in particle physics with ATLAS to a new level. But since a higher trigger rate and more simultaneous collisions imply more pile-up the

readout electronics of the detector will face a new challenge. The signal processing of the LAr Calorimeter is currently using an optimal filter algorithm which will reach its limits in performance with increasing overlapping signals. New approaches for energy-reconstruction are needed, and neural networks are promising candidates for such a task. While it is not hard to build a neural network which reconstructs energies reliably with a lot of trainable parameters, the problem is the limited availability of resources on the FPGAs which are foreseen for the digital signal processing.

In this talk, a possible solution for this task using convolutional neural networks (CNNs) will be presented. It will be shown how CNNs can be structured and trained in such a way that they will fit to the above-mentioned requirements. Special attention will be paid to the energy resolution for signals with a small temporal distance, having the pile-up at the HL-LHC in mind.

T 41.9 Tue 18:15 T-H26

ATLAS Liquid Argon Calorimeter Readiness for LHC Run 3 — • Tom Kresse, Arno Straessner, and Rainer Hentges — Institut für Kern- und Teilchenphysik, Dresden, Deutschland

Liquid argon (LAr) sampling calorimeters are employed by ATLAS for all electromagnetic calorimetry in the pseudo-rapidity region $|\eta| < 3.2$, and for hadronic and forward calorimetry in the region from $|\eta| = 1.5$ to $|\eta| = 4.9$. After detector consolidation during a long shutdown, LHC Run 2 started in 2015 and about 150fb-1 of data at a center-of-mass energy of 13 TeV was recorded. Phase-I detector upgrades began after the end of Run 2. New trigger readout electronics of the ATLAS LAr Calorimeter have been developed. Installation began at the start of the LHC shutdown in 2019 and is expected to be completed in 2021. A commissioning campaign is underway in order to realise the capabilities of the new, higher granularity and higher precision level-1 trigger hardware in Run 3 data taking. This contribution will give an overview of the new trigger readout commissioning, including the first data taken of the recommissioned LAr system in the October 2021 Pilot run, as well as its preparations for Run 3 detector operation.

T 42: Detector Systems 1

Time: Tuesday 16:15-18:00

T 42.1 Tue 16:15 T-H27

R&D of Multidimensional Calorimetry for the SHiP SBT — •FAIRHURST LYONS for the SBT-Collaboration — ALU Freiburg

We present R&D towards a large-area detector for energy reconstruction and limited resolution tracking, which consists of many individual cells filled with liquid scintillator. Each cell is equipped with two wavelength-shifting optical modules (WOMs) that capture scintillation light and transfer it to silicon photomultipliers. This design could serve as the surrounding background tagger (SBT) of the proposed Search for Hidden Particles (SHiP) experiment, a general-purpose detector housed at the CERN SPS accelerator to search for light, feebly interacting particles. SHiP allows probing dark photons, dark (pseudo-)scalars, and heavy neutrinos, as well as the investigation of light dark matter, neutrinos, and flavour physics. The SBT studied here surrounds the vacuum decay vessel of SHiP to detect charged particles either entering the vacuum vessel from outside, or produced in inelastic interactions of muons and neutrinos in the vacuum vessel walls. We present studies of readout electronics and simulated individual detector cells investigating scintillator response.

T 42.2 Tue 16:30 T-H27

Studies to improve the light absorption of wavelength-shifting optical modules in the SHiP experiment — •JAKOB SCHMIDT for the SBT-Collaboration — Humboldt-Universität zu Berlin, Berlin, Germany

The usage of wavelength-shifting optical modules (WOMs) as photon detectors was first proposed for the IceCube large-volume extension and then also for large-area liquid scintillator detectors as the SHiP surrounding background tagger (SBT). A WOM is a light-guiding tube coated with a UV light-absorbing paint that emits secondary photons in the visible spectrum. By total internal reflection inside the tube walls, these photons are guided to the actual photon detector, which in this case is made of a ring array of silicon photomultipliers that is coupled to one end of the tube. The light detection efficiency depends significantly on the light absorption in the wavelength shifter. This talk will present studies of coating methods to improve the light absorption in WOMs. This work is funded by the DFG.

T 42.3 Tue 16:45 T-H27

Measurements of photon exit angles of Wavelength-Shifting Optical Modules used in a large-area liquid-scintillator detector — •FLORIAN REHBEIN for the SBT-Collaboration — RWTH Aachen University, Aachen, Germany

This contribution presents first laboratory measurements of the optical characteristics and the quality of a Wavelength-Sifting Optical Module (WOM) as foreseen for a liquid-scintillator-based large-area detector. Measurements of the photon exit angle distribution have been taken with a DSLR camera on a test stand that was built specifically for this purpose. Further, it will be discussed how systematic measurements and comparisons to simulations will help to examine the properties of the module. WOMS combine a well-designed light guide with a wavelength-shifting coating, presenting a novel optical sensor for numerous applications, first proposed for the large-volume extension of the IceCube detector. WOMs are also foreseen as photon detectors in the Surround Background Tagger (SBT) in SHiP (Search for Hidden Particles), a proposed general-purpose fixed target experiment at the SPS accelerator of the CERN Facility. The SBT acts as a discriminator against external particle interactions and is composed of many cells utilizing liquid scintillator and tube-shaped WOMs made of PMMA to detect traversing particles. The coating of the WOMs absorbs the scintillation photons and re-emits wavelength-shifted photons, which are then detected by an array of SiPMs coupled to one end of the WOM. Supported by the DFG.

T 42.4 Tue 17:00 T-H27

Location: T-H27

Position-dependent detector response of a liquid scintillation detector instrumented with wavelength-shifting optical modules and SiPMs using cosmic muons — •ANDREA ERNST for the SBT-Collaboration — Humboldt-Universität zu Berlin

The usage of wavelength-shifting optical modules (WOMs) as photon detectors was first proposed for the IceCube large-volume extension and then also for large-area liquid scintillator detectors as the SHiP surrounding background tagger (SBT). A WOM is a light-guiding tube coated with a UV light-absorbing paint that emits secondary photons in the visible spectrum. By total internal reflection inside the tube walls, these photons are guided to the actual photon detector, which in this case is made of a ring array of silicon photomultipliers that is coupled to one end of the tube. This contribution shows results of studies on the particle-position-dependent response of a prototype SBT-unit located at HU Berlin using cosmic muons. This work is funded by the DFG.

T 42.5 Tue 17:15 T-H27

Impact of a reflector on the light yield of WOMs — •ALEXANDER VAGTS for the SBT-Collaboration — Humboldt-Universität zu Berlin, Berlin, Deutschland The usage of wavelength-shifting optical modules (WOMs) as photon detectors was first proposed for the IceCube large-volume extension and then also for large-area liquid scintillator detectors as the SHiP surrounding background tagger (SBT). A WOM is a light-guiding tube coated with a UV light-absorbing paint that emits secondary photons in the visible spectrum. By total internal reflection inside the tube walls, these photons are guided to the actual photon detector, which in this case is made of a ring array of silicon photomultipliers that is coupled to one end of the tube. This contribution shows results on whether a reflector at the other end of a WOM tube improves the light yield using cosmic muons traversing a liquid-scintillator detector prototype. This work is funded by the DFG.

T 42.6 Tue 17:30 T-H27

Tracking of charged particles using an FE-I4B pixel telescope and moving emulsion films — •NIKOLAUS OWTSCHARENKO¹, VADIM KOSTYUKHIN¹, CHRISTOPHER BETANCOURT², FABIAN HÜGGING³, DAVID-LEON POHL³, AN-TONIA DI CRESCENZO⁴, ANTONIO IULIANO⁴, and MARKUS CRISTINZIANI¹ — ¹Center for Particle Physics Siegen, Experimentelle Teilchenphysik, Universität Siegen — ²Universität Zürich — ³Physikalisches Institut, Universität Bonn — ⁴Sezione INFN di Napoli

The SHiP collaboration proposes a general purpose fixed-target experiment to search for hidden particles at the new beam-dump facility at CERN SPS. To estimate the charm production cross section in the experiment, which includes hadronic cascade production, several dedicated measurements have been proposed. A first run was performed in summer 2018. Protons from SPS interacted with a thick multilayer target, interleaved with tracking emulsion films. While the emulsion detector offered high spatial resolution, it did not provide timing information. For full event reconstruction a 6-plane telescope made of ATLAS IBL double-chip modules was assembled and placed downstream of the target to provide a high timing resolution. An occupancy limit on the emulsion films made a movement of the target during and in between spills necessary. Recon-

struction of tracks and vertices in the pixel detector as well as matching of track and vertex candidates reconstructed in the moving emulsion detectors are presented.

T 42.7 Tue 17:45 T-H27

The SHiP Surrounding Background Tagger — •Annika Hollnagel for the SBT-Collaboration — JGU Mainz

Within the CERN Physics Beyond Colliders (PBC) initiative, the SHiP fixed-target experiment is a frontrunner proposal for the SPS Beam-Dump Facility (BDF). Making use of the high-intensity SPS beam with 4×10^{19} protons on target per year, the experiment will combine the Search for Hidden Particles (SHiP) of masses up to 200 MeV/ c^2 - such as Heavy Neutral Leptons (HNL) and Light Dark Matter (LDM) - with studies of tau neutrino physics.

The Hidden Sector Decay Spectrometer (HSDS) of the SHiP detector consists of a large evacuated volume followed by a magnetic spectrometer and particle identification system. To enable a background-free study of the decays of feeblyinteracting particles, the reduction of beam-induced background heavily relies on the Surrounding Background Tagger (SBT) that envelops the 50 m-long decay vessel. The current baseline for the SBT is a segmented Liquid Scintillator (LS) detector of LAB and PPO that is instrumented with Wavelength-shifting Optical Modules (WOM) and read out via SiPMs.

Since 2017, several test beam exposures of prototype detector cells have been conducted, supported by laboratory measurements and simulations. This talk will give a general overview on the SBT, summarise the state of the ongoing R&D, and present our plans for the next period of test beam measurements.

T 43: DAQ and Trigger 2

Time: Tuesday 16:15-18:15

T 43.1 Tue 16:15 T-H28

Calibration and Trigger Studies for the OSIRIS pre-detector of JUNO — •RUNXUAN LIU^{1,2}, PHILIPP KAMPMANN³, KAI LOO⁴, LIVIA LUDHOVA^{1,2,3}, ALEXANDRE GÖTTEL^{1,2}, NIKHIL MOHAN^{3,2}, LUCA PELICCI^{1,2}, MARIAM RIFAI^{1,2}, APEKSHA SINGHAL^{1,2}, and CORNELIUS VOLLBRECHT^{1,2} — ¹Forschungszentrum Jülich GmbH, Nuclear Physics Institute IKP-2, Jülich, Germany — ²III. Physikalisches Institut B, RWTH Aachen University, Aachen — ³GSI Helmholtz Centre for Heavy Ion Research, Darmstadt, Germany — ⁴Institute of Physics and EC PRISMA⁺, Johannes Gutenberg Universität Mainz, Mainz

JUNO is a 20 kt liquid scintillator detector under construction in Jiangmen, China, which is expected to start data-taking in 2023. Its main goal is to determine the neutrino mass hierarchy with the measurement of reactor neutrinos from two adjacent nuclear power plants. This requires stringent limits on the radiopurity of the detector. The OSIRIS pre-detector is designed to monitor the liquid scintillator during the several months of filling the large volume of JUNO. OSIRIS will contain 18 ton of scintillator and will be equipped with 76 20-inch PMTs. OSIRIS will utilize two calibration systems: a laser fiber system and an insertion system for an LED or radioactive sources. The data acquisition system will have no global hardware trigger: instead, each PMT will provide a datastream composed of the digitized PMT pulses, each containing a time stamp. Based on the latter, dedicated event builder software will organize these data streams into events. This talk will discuss the several trigger modes that are realized in the event builder software, a coincidence trigger and the calibration trigger.

T 43.2 Tue 16:30 T-H28

The MDT Trigger Processor for the ATLAS HL-LHC Upgrade — •DAVIDE CIERI, MARKUS FRAS, OLIVER KORTNER, and SANDRA KORTNER — Max-Planck-Institut für Physik, Munich, Germany

The novel MDT Trigger Processor (MDTTP) system is a fundamental part of the upgrade of the first-level (L0) muon trigger of the ATLAS experiment at the HL-LHC. The new system will be responsible for improving the muon momentum resolution and thus refining the muon selectivity, using for the first time at L0 the precision tracking information from Monitored Drift Tube (MDT) chambers in addition to the trigger chamber information. The system will also transmit the MDT hit data to the data acquisition (DAQ) system in the event of a trigger accept. A total of 64 MDTTP boards will be installed in ATLAS, one for each MDT trigger sector. The design of the MDTTP is highly challenging, requiring a high number of optical links and high-performance processing units.

We present here the recently designed prototype of the MDTTP. The prototype adopts a ATCA design, composed by two modules: the Service Module responsible for the powering and the infrastructure; and the Command Module, performing the trigger and DAQ processing and communicating with the other components of the ATLAS muon trigger. The Command Module mounts a state-of-the-art Xilinx Virtex Ultrascale+ FPGA, and employs ten 12-channel bidirectional optical transceiver modules with a link speed of up to 14 Gbps. Location: T-H28

T 43.3 Tue 16:45 T-H28

The ATLAS Global Event Processor and the jet energy determination perspectives at level-0 trigger in Run-4 — •FERNADO DEL RIO — Kirchhoff Institut für Physik, Universität Heidelberg

Run 4 of the LHC, scheduled to take place between 2028 and 2032, is planned to deliver never-before-seen conditions for high energy physics measurements, including a center of mass energy of 14 TeV and an average number of interactions per crossing of 200. Under these challenging conditions the triggering algorithms and their performance becomes more important than ever.

Amongst the changes that will be implemented in ATLAS for this Run we highlight the construction and installation of the Global Event Processor (GEP), a single board with the whole detector information made for triggering at 40 MHz with offline-like algorithms for the reconstruction of all objects, such as electrons, photons, muons and jets. In the particular case of jet reconstruction, access to layer information allows for state-of-the-art calibration techniques to be applied at the trigger level. This talk will discuss the potential of having a level 1 version of the Global Sequential Calibration, a part of the jet calibration sequence tailored for improving the resolution in the measurement of jet energy. This improvement in resolution naturally yields a better trigger efficiency and a higher sensitivity to dijet mass resonances, allowing ATLAS to fully exploit the large statistics given by Run4.

T 43.4 Tue 17:00 T-H28

Trigger optimization studies in the search for displaced heavy neutral leptons with the ATLAS detector. • JULIUS EHRSAM, HEIKO LACKER, and CHRISTIAN APPELT — Humboldt University, Berlin, Germany

Adding heavy neutral leptons (HNLs) with masses below the electroweak scale to the SM Lagrangian can help explain observed Beyond-Standard-Model phenomena such as neutrino oscillations, matter-antimatter asymmetry, and dark matter. We study a new trigger option for the ATLAS experiment designed explicitly for the search for HNLs produced in events $pp \rightarrow W + X$, $W \rightarrow \ell + HNL$, with HNL $\rightarrow \mu^+\mu^-\nu$, resulting in a prompt lepton and a displaced vertex due to the expected long HNL lifetime. The proposed trigger uses a muon as the prompt lepton and the angular separation between the two muons from the HNL decay.

T 43.5 Tue 17:15 T-H28

Optimizing the ATLAS b-jet Trigger for the LHC Run 3 — •VICTOR H. RUELAS RIVERA — Humboldt-Universität zu Berlin, Berlin, Germany

The Higgs potential provides a way to experimentally probe and understand the underlying principles of mass generation and electroweak symmetry breaking. The shape of the Higgs potential is proportional to the Higgs self-coupling, λ_{HHH} , which can be probed at the LHC via proton-proton collisions $(pp \rightarrow HH)$. Di-Higgs to 2 pairs of quarks $(HH \rightarrow b\bar{b}b\bar{b})$ is one of the most sensitive decay channels and it relies heavily on b-jet triggers. Triggers select information in real-time from the collisions and help mitigate ATLAS data acquisition getting overwhelmed by QCD jets. However, $HH \rightarrow b\bar{b}b\bar{b}$ is difficult to trigger due to the soft signal kinematics and high thresholds of hadronic triggers.

Hence, more signal can be gained in Run 3 with better b-jet triggers. One of the goals of Run 3 is to improve the $HH \rightarrow b\bar{b}b\bar{b}$ triggers to enhance signal acceptance of SM and Beyond Standard Model (BSM) scenarios. The taggers that will be used for the b-jet trigger in Run 3 exploit multivariate analysis techniques, mainly Deep Neural Networks. This talk presents the optimization of the neural-network based flavour tagging discriminant used by the b-jet trigger. The algorithm is optimized on tracks, jets and vertices reconstructed by the High Level Trigger (HLT) software. The training software, network architecture and simulated events are being shared with the ATLAS offline b-tagging group to address redundancies and combine efforts towards a unified training framework for quick model re-optimization.

T 43.6 Tue 17:30 T-H28

Development of the Topological Trigger for LHCb Run 3 - JOHANNES ALBRECHT¹, GREGORY MAX CIEZAREK², NIKLAS NOLTE³, MIROSLAV SAUR¹, and •NICOLE SCHULTE¹ — ¹TU Dortmund University, Department of Physics 2 CERN — 3 MIT

In Run 3, the LHCb experiment will undergo a significant upgrade including the conversion to a software-only trigger system. This means the trigger software will have to efficiently process a substantially higher amount of data compared to previous years. The largest trigger algorithm for beauty physics in LHCb software is the Topological Trigger, which produces output used for most analyses. It is an inclusive trigger, aiming to select beauty decays based on topological and kinematic properties. Due to its inclusive nature, it can select candidates and trajectories from decays that might not have been discovered yet.

Two contributions to the Topological Trigger are shown. The first contribution will serve as a baseline for the collaboration and is based on a boosted decision tree algorithm. This approach has proven to be efficient in former years of data taking but has been optimized and remodeled for simulation of LHCb's upgrade. One of the major changes for the upgrade is the rise in the number of primary vertices during the interaction. Since primary vertex information is crucial for the Topological Trigger, the algorithm needs to be optimized to the conditions in the upgrade. The second contribution will explore a more experimental technique using a newly developed neural network architecture providing a robust model for the selection. Finally, both performances are compared.

T 43.7 Tue 17:45 T-H28

Study of potential lifetime bias in the LHCb reconstruction software for Run 3 - • PAULA HERRERO GASCÓN and PEILIAN LI - University of Heidelberg, Physikalisches Institut Heidelberg, Germany

The upgraded LHCb experiment will restart data taking in spring 2022 at an instantaneous luminosity of up to 2×10^{33} cm⁻² s⁻¹. To effectively select the interesting b and charm hadron events at these high rates a full software trigger system is required. The entire reconstruction framework and its algorithms have been reimplemented and optimized. To be well-prepared for the coming data taken, it is essential to validate the physics performance for this completely new system. The precise reconstruction of the decay-time of b hadrons is vital for many time-dependent physics measurements at LHCb. This talk will focus on a performance study of the decay-time reconstruction for Run 3 and investigate potential reconstruction-induced biases.

T 43.8 Tue 18:00 T-H28 Performance of Belle II's Level 1 Single Track Trigger - •FELIX Meggendorfer^{1,3}, Christian Kiesling^{1,4}, Elia Schmidt^{1,4}, Kai Lukas Unger², Steffen Bähr², Alois Knoll³, Alexander Lenz³, Sebastian SKAMBRAKS², and JÜRGEN BECKER² for the Belle II-Collaboration - ¹MPI for Physics, Munich, Germany – ²KIT, Karlsruhe, Germany – ³TUM, Munich, Germany — ⁴LMU, Munich, Germany

Belle II is the world record holder for the highest instantaneous luminosity, and the machine is still more than a factor 20 away of what it is capable of. These high collision rates make it mandatory to have an efficient trigger system. The Neurotrigger is a level 1 track trigger using the central drift chamber in the Belle II experiment. It estimates the z vertex and the polar angle θ of the tracks. To suppress the dominating background of events coming from outside of the interaction point, a cut on the track vertices along the beam axis is combined wit a momentum cut. This trigger, the 'STT', operates without a prescale and outperforms all other track triggers in Belle II. It is the first of its kind in high energy physics and most important for events with low charged multiplicity, such as tau pair production and candidates for dark matter searches.

T 44: Experimental Methods (general) 2

Time: Tuesday 16:15-17:45

T 44.1 Tue 16:15 T-H29

Particle identification with fast timing detectors at future Higgs factories - •Bohdan Dudar^{1,2}, Jenny List¹, and Ulrich Einhaus¹ - ¹Deutsches Elektronen-Synchrotron DESY, Hamburg, Germany — 2 Universität Hamburg, Hamburg, Germany

Future e^+e^- Higgs factory collider projects are designed for precision measurements of the Higgs boson and of electroweak observables, thereby utilizing every event to their full potential. The identification of the pions, kaons and protons plays a key role for precision measurements and event reconstruction, especially for the flavour tagging. To improve the identification of charged hadrons at low momentum we can use the time-of-flight method. It relies on current silicon sensor technologies with extremely good time resolution of 10 - 30 ps. This allows to measure the time-of-flight of particles and reconstruct their mass providing additional tool for identification of π^{\pm} , K^{\pm} and p.

We study possible realistic implementation scenarios and potential physics applications of the fast timing silicon sensors into the future Higgs factory detectors using as an example the International Large Detector (ILD) at the International Linear Collider (ILC).

T 44.2 Tue 16:30 T-H29

Charged Kaon Mass Measurement via Time-of-Flight at a Future Higgs Factory – •Ulrich Einhaus¹, Bohdan Dudar^{1,2}, and Jenny List¹ – ¹Deutsches Elektronen-Sychrotron DESY, Notkestraße 85, 22607 Hamburg – ²Universität Hamburg, Germany

The proposed future e^+e^- Higgs factories will allow to study not only the Higgs boson but the entire electroweak sector to an unprecendented precision. In order to utilise each event as well as possible, particle identification (PID) can be a powerful tool, which has been studied with increasing interest in recent years. The development of picosecond-timing detectors has been driven by the demand of background suppression at the LHC, but they can also be used for an effective time-of-flight (TOF) measurement to distinguish different charged hadrons. In addition to flavour physics applications, TOF can also be used to perform a competitive measurement of the charged kaon mass. This value is among the less precisely known ones and the two leading contributions from the early 1990s are at tension by more than 5σ with each other. The kaon mass value, however, is input to precision tracking and decay chain reconstruction, and can also be used to test theory predictions, e.g. from lattice QCD. This presentation discusses the prospects of a measurement of the charged kaon mass at future Higgs factories based on the full detector simulation of the International Large Detector (ILD) concept. This is the first analysis to make use of the recent implementation of TOF in a full-detector simulation for a future Higgs factory. It shows that such

T 44.3 Tue 16:45 T-H29

Location: T-H29

Identification of leptons inside jets at future Higgs factories – •Leonhard Reichenbach^{1,2}, Yasser Radkhorrami^{1,2}, and Jenny List¹ – ¹DESY Hamburg — ²Universität Hamburg

a measurement is feasible, but also highlights requirements and open questions.

One goal of a future Higgs factory is the precise measurement of the 125 GeV Higgs boson properties. As the Higgs boson predominantly decays to bb, the precise reconstruction of heavy flavor jets is crucial. A source of uncertainty for these jets is missing momentum from semi-leptonic decays $b \rightarrow \ell \nu X$. Recent work has shown the possibility of correcting this missing neutrino momentum. For this, the charged lepton from the decay needs to be successfully detected and reconstructed. While particle flow detector concepts with their high granularity offer ideal conditions to identify leptons inside jets, the excellent hardware needs to be matched with corresponding reconstruction algorithms. In this work, we use the detailed simulation of the ILD detector concept to investigate how to exploit the information provided by a particle flow detector to identify single electrons and muons in a dense environment and how this improves the reconstruction of $H \rightarrow b\bar{b}$ decays.

T 44.4 Tue 17:00 T-H29

Detection of Spectra in the Strong-Field QED Regime with LUXE - • JOHN HALLFORD^{1,2} and MATTHEW WING^{1,2} - ¹University College London -²Deutsches Elektronen Synchrotron

Conventional QED's validity breaks down in the presence of an external strong electric field. LUXE (LASER Und XFEL Experiment), in Hamburg, intends to collide a high-intensity LASER pulse with highly boosted electrons and photons, up to 17.5 GeV from the EuXFEL, creating assisted electric fields up to and greater than the Schwinger limit.

This enables a non-negligible probability of non-linear Compton Scattering and Breit-Wheeler interactions - which represents a spontaneous boiling of the vacuum. The rates and kinematics of these interactions will be measured. Detection challenges include low-flux positron detection and tracking in a highradiation environment, GeV-photon spectrometry, and high-flux, high-energy electron energy distribution reconstructions for a variety of spectrum shapes and dynamic ranges.

One of two detection solutions employed for the electron detection is a thin screen of a scintillating material, imaged remotely by optical cameras, and using magnetic deflection to reconstruct with respect to energy. The reconstruction methods and expected results for this detector and its consequences for LUXE are discussed.

T 44.5 Tue 17:15 T-H29

Simulation Studies for the Polarimetry of a LPA Electron Beam – •JENNIFER POPP^{1,2}, SIMON BOHLEN¹, FELIX STEHR^{1,2}, JENNY LIST^{1,2}, GUDRID MOORTGAT-PICK^{2,1}, JENS OSTERHOFF¹, and KRISTJAN PÕDER¹ – ¹Deutsches Elektronen-Synchrotron DESY, Hamburg -²University of Hamburg

Polarized particle beams are a key instrument for the investigation of spindependent processes and Laser Plasma Acceleration (LPA) has become a promising alternative to conventional RF accelerators. However, so far, it has only been theoretically shown that polarized LPA beams are possible.

The LEAP (Laser Electron Acceleration with Polarization) project at DESY aims to demonstrate this experimentally for the first time, using a prepolarized plasma target.

Because it is best suited for the expected energy range, the electron polarization will be measured with photon transmission polarimetry. It makes use of the production of circularly polarized Bremsstrahlung during the passage of the electrons through a suitable target. The photon polarization is then measured

with the aid of the transmission asymmetry related to the magnetization direction of an iron absorber.

In this contribution an overview of the LEAP project will be given and a design for the polarimeter, simulation studies, and requirements on beam and polarimeter parameters for reliable polarization measurements will be presented.

T 44.6 Tue 17:30 T-H29

Calorimeter R&D for LPA Polarimetry — •FELIX STEHR^{1,2}, SIMON BOHLEN¹, Oleksandr Borysov¹, Maryna Borysova^{3,1}, Jennifer Popp^{1,2}, Jenny List¹, Gudrid Moortgat-Pick^{2,1}, Jens Osterhoff¹, and Kristjan Põder¹ – ¹Deutsches Electronen-Synchrotron DESY, Hamburg – ²University of Hamburg — ³Institute for Nuclear Research NASU, Kyiv

The LEAP (Laser Electron Acceleration with Polarization) project at DESY aims to demonstrate the generation of polarized electron beams with a Laser-Plasma-Accelerator (LPA). Due to the expected beam energy of about 50 MeV, photon transmission polarimetry will be used to determine the achieved degree of polarization.

The key observable is an energy asymmetry of photons passing through a magnetized iron absorber. The total transmitted photon energy will be in the order of tens of TeV, which needs to be measured with percent-level accuracy in order to reliably detect asymmetries of a few ten percent. This contribution will discuss the detector requirements derived from detailed Geant4-simulations of the polarimeter and compare them to a first test of a calorimeter prototype operated in the LPA beam.

T 45: Gamma Astronomy 2

Time: Tuesday 16:15-18:15

T 45.1 Tue 16:15 T-H30

The Crab pulsar wind nebula in high-energy gamma-rays and its flaring emission — •Michelle Tsirou¹, Brian Reville¹, Emma de Oña-Wilhelmi², GWENAËL GIACINTI¹, and JOHN KIRK¹ — ¹MPIK, Heidelberg, Germany -²DESY, Berlin-Zeuthen, Germany

The Crab nebula system is one of the brightest gamma-ray sources in the Milky Way, extensively observed across the electromagnetic spectrum. Recent studies purport intermittent flaring events deviating from the continuous flux associated with its synchrotron spectrum below a few hundreds of MeV, straining the theoretical synchrotron burn-off limit in these energy ranges. By analysing available Fermi-LAT data across a thirteen-year-long monitoring, we study the energy-dependence of its flaring behaviour in a few hundred MeV up to a few GeV energy ranges. We explore acceleration mechanisms prone to explain the variability of the observed emission and discuss its implication on our current understanding of this extreme system.

T 45.2 Tue 16:30 T-H30 Machine learning methods for constructing probabilistic Fermi-LAT catalogs - •Ааказн Внат¹ and Dмітку Malysнev² — ¹Dr. Karl-Remeis Strernwarte, Bamberg — ²ECAP Erlangen

Classification of sources is one of the most important tasks in astronomy. Sources detected in one wavelength band, for example using gamma rays, may have several possible associations in other wavebands or there may be no plausible association candidates. In this work, we aim to determine probabilistic classification of unassociated sources in the third and the fourth data release 2 Fermi Large Area Telescope (LAT) point source catalogs (3FGL and 4FGL-DR2) into two classes (pulsars and active galactic nuclei (AGNs)) or three classes (pulsars, AGNs, and other sources). We use several machine learning (ML) methods to determine probabilistic classification of Fermi-LAT sources. We evaluate the dependence of results on meta-parameters of the ML methods, such as the maximal depth of the trees in tree-based classification methods and the number of neurons in neural networks. We determine probabilistic classification of both associated and unassociated sources in 3FGL and 4FGL-DR2 catalogs. We cross-check the accuracy by comparing the predicted classes of unassociated sources in 3FGL that have associations in 4FGL-DR2. We find that in the 2-class case it is important to correct for the presence of other sources among the unassociated ones in order to realistically estimate the number of pulsars and AGNs.

T 45.3 Tue 16:45 T-H30

Classification of Fermi-LAT blazars with Bayesian neural networks - ANJA BUTTER¹, •THORBEN FINKE², FELICITAS KEIL², MICHAEL KRÄMER², and SIL-VIA MANCONI 2 — 1 Institut für Theoretische Physik, Universität Heidelberg, Germany – ²Institute for Theoretical Particle Physics and Cosmology, RWTH Aachen University, Germany

We apply Bayesian neural networks on the classification of y-ray sources within the Fermi-LAT catalog. We focus on blazar candidates and their subclassification into BL Lacertae and Flat Spectrum Radio Quasars. We explore

the correspondence between conventional dense and Bayesian neural networks and the effect of data augmentation. We find that Bayesian neural networks provide a robust classifier with reliable uncertainty estimates and are particularly well suited for classification problems that are based on comparatively small and imbalanced data sets. The results of our blazar candidate classification are valuable input for population studies aimed at constraining the blazar luminosity function and to guide future observational campaigns.

T 45.4 Tue 17:00 T-H30

Location: T-H30

Analysis of the high energy γ -ray emission from HESS J1813-178 with H.E.S.S and Fermi-LAT data — •TINA WACH, VIKAS JOSHI, ALISON MITCHELL, and STE-FAN FUNK — Erlangen Center for Astroparticle Physics

HESS J1813-178 is one of the brightest and most compact objects detected by the HESS Galactic Plane Survey. Within the extent of the TeV emission lies a young Supernova Remnant G12.8-0.02 and a pulsar wind nebula driven by a pulsar with the second highest spin-down luminosity of known pulsars in the Galaxy PSR J1813-1749. The origin of the TeV emission is still not clear. Because of the young age of the system, the pulsar wind nebula and the Supernova Remnant overlap and present a good opportunity to examine how the interactions between these two components influence the acceleration of particles in the system. In previous analyses, a discrepancy in extension has been observed between a point-like component seen in TeV γ -rays measured by H.E.S.S and a extended component in GeV y-ray observations from the Fermi-LAT satellite. We used 3 dimensional map-based analysis with gammapy to do a morphological and spectral analysis of this region in GeV and TeV energy ranges, as well as a joint-analysis of both datasets. We find a new significant extended emission component in the TeV energy range, with a morphology close to the GeV energy range. While the question of the origin of the very high y-ray emission from HESS J1813-178 could not be answered yet, our analysis allows a consistent description and a smooth energy spectrum of the region across five decades of energy.

T 45.5 Tue 17:15 T-H30

Science verification and highlights of the new FlashCam-based camera in the $28m \, telescope \, of \, H.E.S.S. - \bullet {\sf SIMON} \, {\sf Steinmassl} \, for \, the \, H.E.S.S. - Collaboration$ – Max Planck Institut für Kernphysik, Heidelberg, Germany

In October 2019, the 28m telescope in the centre of the H.E.S.S. array was upgraded with a new camera. The camera itself is a prototype based on the Flash-Cam design, which was originally developed for the Cherenkov Telescope Array (CTA). We report here the results of the validation and science verification programme that has been performed after the commissioning period. From that, we conclude that the camera, analysis and simulation pipelines are working up to expectations. Finally, we discuss some highlights of the scientific results obtained during the first two years of science operation with the new camera, as well as possible future science cases.

T 45.6 Tue 17:30 T-H30

Lake Deployment of Southern Wide-field Gamma-ray Observatory Detector Units — •HAZAL GOKSU for the SWGO-Collaboration — Max-Planck-Institut für Kernphysik, Heidelberg, Germany

This contribution is about the lake concept, one of the possible detector designs for Southern Wide-field Gamma-ray Observatory (SWGO). SWGO will be a next-generation high altitude gamma-ray survey observatory in the southern hemisphere consisting of an array of water cherenkov detectors. With its energy range, wide field of view, large duty cycle, and location it will complement the other existing and planned gamma-ray observatories. The lake concept is an alternative to the HAWC-like separate detector unit design, and the LHAASOstyle artificial ponds. Instead of having tanks filled with water, bladders filled with clean water are deployed near the surface of a natural lake, where each bladder is a light-tight stand-alone unit containing one or more photosensors. We will give an overview of the advantages and challenges of this design concept and describe the first results obtained from prototyping.

T 45.7 Tue 17:45 T-H30 Status and first results from TAIGA — •MICHAEL BLANK and MARTIN

ТLUCZYКОNТ — Institut für Experimentalphysik, Universität Hamburg

TAIGA has implemented a new, unique observation technique, based on a combination of the imaging air Cherenkov telescope (IACT) technique, and the HiSCORE timing array concept. TAIGA aims to further explore the celestial gamma-ray sources at energies of a few 10s of TeV to several 100 TeV. This energy range is particularly important to spectrally resolve the cutoff regime of the galactic Pevatrons, the cosmic-ray accelerators to PeV energies. TAIGA currently consists of an array of 121 wide angle (0.6 sr) air Cherenkov timing stations distributed over an area of about 1 km² and two IACTs with a diameter of 4.75 m and a field of view of 9.7°. A third IACT will be completed during this year.

In this presentation, the current status of the experiment and the analysis is discussed and the detection of the Crab Nebula with data of previous seasons, with a smaller size array and only with the first IACT in operation, is shown.

T 45.8 Tue 18:00 T-H30

COMCUBE: Exploring the violent Universe with CubeSat Technology — •JAN LOMMLER for the COMCUBE-Collaboration — Johannes Gutenberg-Universität Mainz

Gamma Ray Bursts are a window into some of the most energetic processes in the Universe. Due to the energy range of the emitted electro-magnetic radiation, measurements have to be performed in space by a network of either dedicated observatories like SWIFT and POLAR or secondary detectors mounted on larger observatories like Fermi GBM. Most detectors only allow the measurement of the burst's energy-spectrum and time evolution, missing out on polarization of the incident photons. Using Compton scattering as main detection channel, Cubesats offer the opportunity to setup a network of small-scale dedicated detectors at relatively low cost that are able to pinpoint GRBs, measure their spectra and temporal evolution while obtaining polarization information. In this talk, we want to present the detector concept of COMCUBE, performance estimates and the status of the balloon prototype.

T 46: Neutrino Astronomy 2

Time: Tuesday 16:15-18:30

T 46.1 Tue 16:15 T-H31

Search for the Galactic Diffuse Neutrino Flux with IceCube — •JONAS HELLRUNG¹, JAKOB BÖTTCHER¹, PHILIPP FÜRST¹, ERIK GANSTER¹, PHILIPP MERTSCH², GEORG SCHWEFER¹, ROMAN SUVEYZDIS¹, and CHRISTOPHER WIEBUSCH¹ for the IceCube-Collaboration — ¹III. Physikalisches Institut B, RWTH Aachen University — ²Institut für Theoretische Teilchenphysik und Kosmologie, RWTH Aachen University

A diffuse flux of high-energy astrophysical neutrinos has been detected with the IceCube Neutrino Observatory. This flux is dominated by neutrinos of extragalactic origin. However, a fraction of these neutrinos is expected to originate from cosmic rays interacting with the interstellar medium in our Galaxy. This flux has not been measured yet, but a detection could be within reach of IceCube's sensitivity. The signature of this flux with respect to the isotropic extragalactic flux is its close correlation with the galactic plane. The detection would contribute to the understanding of propagation and sources of galactic cosmic rays of typically PeV energies. In this talk the principle of an analysis using neutrino-induced through-going muons is presented and its sensitivity is discussed.

T 46.2 Tue 16:30 T-H31

The energy spectrum of the diffuse neutrino flux in a combined fit using 10 years of IceCube data — •RICHARD NAAB¹, ERIK GANSTER², MARKUS ACKERMANN¹, and CHRISTOPHER WIEBUSCH² for the IceCube-Collaboration — ¹DESY, Zeuthen, Germany — ²Physics Institute III B, RWTH Aachen University, Germany

The IceCube Neutrino Observatory has discovered a diffuse astrophysical neutrino flux and measures the energy spectrum and flavor composition in different detection channels. With almost 10 years of data, we aim to combine different detection channels to constrain this energy spectrum and flavor composition with unprecedented precision. The increased statistics require rigorous treatment of systematic uncertainties, which we aim to achieve with the so-called SnowStorm method, recently developed within the IceCube collaboration. This technique involves a continuous variation of systematics parameters during the detector simulation and requires a dedicated analysis approach.

In this work, we present the validation of this method for the purpose of measuring the energy spectrum, using two distinct analysis approaches. Furthermore, we discuss the sensitivity of an upcoming analysis combining the detection channels of tracks and cascades.

T 46.3 Tue 16:45 T-H31

Impact of the Binning in the Likelihood-Analysis of the Diffuse Neutrino Flux in IceCube — •ROMAN SUVEYZDIS, JAKOB BÖTTCHER, PHILIPP FÜRST, ERIK GANSTER, JONAS HELLRUNG, and CHRISTOPHER WIEBUSCH for the IceCube-Collaboration — RWTH Aachen University, Aachen, Germany

The IceCube Neutrino Observatory is measuring a diffuse flux of astrophysical high-energy muon-neutrinos. For this measurement, events are binned in energy and zenith direction, and a flux model can be fitted to the resulting distributions with a likelihood. The current analysis is based on a binning scheme with equidistant bins in logarithmic energy and cosine zenith. Such a binning is suboptimal, because particularly for high energies, the event statistics in the experimental data as well as in the Monte-Carle prediction is low, leading to inherent uncertainties. Here, we explore the effects of non-equidistant and generalized binnings, and investigate how these binnings can improve the analysis.

T 46.4 Tue 17:00 T-H31

Location: T-H31

Studying the Energy Dependent Cosmic Ray Moon and Sun Shadow with IceCube Data. — •JOHAN WULFF and JULIA BECKER TJUS for the IceCube-Collaboration — Institute for theoretical physics IV, Ruhr-University Bochum, Germany

While the Cosmic Ray Moon shadow can be utilised to verify the pointing of detectors such as IceCube, the Sun shadow has proven to be a successful tool for the assessment of solar magnetic field models.

In a previous publication (Aartsen et al, Phys. Rev. D), the relationship between the solar activity and the strength of the Cosmic Ray Sun shadow was investigated by comparing seven years of IceCube data with the solar cycle and solar magnetic field models. By modelling the propagation of Cosmic Rays through the solar magnetic field, two models of the coronal magnetic field were tested by comparing the observed shadow to the predicted one.

For this analysis, an energy reconstruction of Cosmic Ray induced muon events was introduced, allowing for an investigation of the energy dependence of both shadows. Furthermore, magnetic field effects of the Sun shadow can be investigated at different energies and an energy- dependent pointing can be studied with the Cosmic Ray Moon shadow. In this talk, the performance of the machine-learning based energy reconstruction with respect to the IceCube Cosmic Ray Sun and Moon shadow dataset, as well as the extension of the Cosmic Ray Sun shadow analysis up to the minimum of the 24th Solar cycle will be discussed. This project is funded via BMBF Verbundforschung, Project 05A20PC2.

T 46.5 Tue 17:15 T-H31

Seasonal Variations of the Atmospheric Neutrino Flux determined from 10 years of IceCube Data with DSEA+ — •KAROLIN HYMON and TIM RUHE for the IceCube-Collaboration — Astroparticle Physics WG Rhode, TU Dortmund University, Germany

The IceCube Neutrino Observatory is a cubic-kilometer detector array at the South Pole. Beyond the detection of astrophysical neutrinos, the detector measures atmospheric neutrinos at a high rate. These atmospheric neutrinos originate from cosmic ray interactions in the upper atmosphere, mainly from the decay of pions and kaons. The rate of the measured neutrinos, however, is affected by seasonal temperature variations in the Stratosphere, and the variations are expected to increase with the particle's energy. The Dortmund Spectrum Estimation Algorithm (DSEA+) is a novel approach to spectrum unfolding. The ill-posed problem is transferred to a multinomial classification task, in which the energy distribution is estimated from measured quantities by machine learning algorithms. In this talk, the analysis approach to measure the spectral dependence of the seasonal neutrino flux will be presented. Seasonal neutrino energy

spectra are determined by DSEA+, utilizing 10 years of IceCube's atmospheric muon neutrino data. The differences of the unfolded seasonal spectra will be compared to the unfolded annual mean flux.

T 46.6 Tue 17:30 T-H31

NN-based parametrization of muon deflections simulated by PROPOSAL — •PASCAL GUTJAHR and MIRCO HÜNNEFELD — As- troparticle Physics WG Rhode, TU Dortmund University, Germany

Neutrinos are fundamental particles that are nearly massless and not charged. This allows them to traverse the universe along a straight line without deflection. It makes them an excellent candidate for the search of astrophysical neutrino sources. This is achieved using detec- tors such as the IceCube Neutrino Observatory, which is located at the geographic South Pole. In the ice at a depth of 1450 m, Cherenkov radiation from the secondary particles of neutrino interactions is mea- sured. Reconstruction of the daughter particles can then be used to infer the direction of the original neutrino. Due to the high energies in the GeV to PeV range, the deflection of the particles due to stochas- tic energy losses and multiple scattering is assumed to be negligible compared to the directional resolution of IceCube. This assumption may begins to be inaccurate as resolution increases and low energy muons are studied especially with respect to the IceCube upgrade. A study with the lepton propagator PROPOSAL is done to present ex- pected deflections of muons for different energy ranges. The result is parametrized by a neural network.

T 46.7 Tue 17:45 T-H31

Improving Astrophysical Muon-Neutrino Measurements with New Energy Estimators in IceCube – •PHILIPP FÜRST, JAKOB BÖTTCHER, ERIK GANSTER, JONAS HELLRUNG, GEORG SCHWEFER, ROMAN SUVEYZDIS, and CHRISTO-PHER WIEBUSCH for the IceCube-Collaboration — RWTH Aachen University - Physics Institute III B, Aachen, Germany

The IceCube Neutrino Observatory is measuring a diffuse astrophysical neutrino flux with a spectral shape compatible with a single power law at Tev to PeV energies. With neutrinos conveying information directly from their source, their spectral shape is closely connected to open questions about the acceleration and propagation of the hadronic component of cosmic rays. While hints for spectral features beyond a single power law exist, spectral resolution in the muonneutrino detection channel is strongly tied to the achievable energy resolution. We present new methods of energy estimation based on machine learning methods and estimate the expected sensitivity gain for model hypotheses beyond the single power law, exploring the possibility to distinguish different spectral shapes in the future.

T 46.8 Tue 18:00 T-H31

The Pacific Ocean Neutrino Experiment: prototype line development — •MARTIN DINKEL, ELISA RESCONI, and CHRISTIAN SPANNFELLNER for the P-ONE-Collaboration — Technische Universität München

The Pacific Ocean Neutrino Experiment (P-ONE) is a planned large-scale neutrino telescope in the Northeast Pacific Ocean, near the coast of Vancouver Island, British Columbia. The P-ONE collaboration, founded by an initiative of Canadian and German institutes, successfully deployed two pathfinder experiments to probe the site for P-ONE. The next milestone is the construction and deployment of the first line of the Pacific Ocean Neutrino Experiment. This socalled prototype line will comprise optical and calibration modules and follow a novel design approach to minimize interface gaps and allow a scalable deployment approach. We will present the optical and calibration module development status and introduce the overall line design and deployment approach.

T 46.9 Tue 18:15 T-H31

Machine-learning aided experimental design for P-ONE — •JANIK PROT-TUNG, CHRISTIAN HAACK, and ARTURO LLORENTE ANAYA — Technical University Munich, Germany

The Pacific Ocean Neutrino Experiment (P-ONE) is a collaboration of Ocean Networks Canada (ONC) and Universities from Germany, Canada, and the USA to build a large volume neutrino telescope in the Pacific Ocean. Similar to other neutrino telescopes, P-ONE wants to instrument the ocean with photosensors deployed on vertical cables (lines) to detect high-energy neutrino interactions by the Cherenkov light emitted from secondary particles. The design of such telescopes has a variety of free parameters, such as the sensor spacing and sensor density, trigger algorithms and thresholds, or hardware used for signal digitization. These parameters directly impact the physics potential of the telescope and need to be optimized under external constraints (cost, bandwidth, site limitations). These optimization studies typically require expensive Monte-Carlo simulations that limit the explorable parameter phase space. This talk presents a framework that uses graph-neural networks and multi-parameter optimization to comprehensively explore the parameter phase space while reducing the simulation time. The framework facilitates data-driven decisions for the P-ONE design, maximizing the physics potential while minimizing the expenses.

T 47: Cosmic Ray 2

Time: Tuesday 16:15-18:15

T 47.1 Tue 16:15 T-H32

The Underground Muon Detector in The Pierre Auger Observatory: calibration and characterization — •MARINA SCORNAVACCHE^{1,2}, FEDERICO SÁNCHEZ¹, JUAN MANUEL FIGUEIRA¹, and MARKUS ROTH² for the Pierre Auger-Collaboration — ¹Instituto de Tecnologías en Detección y Astropartículas, Comisión Nacional de Energía Atómica, Buenos Aires, Argentina — ²Institut für Astroteilchenphysik, Karlsruher Institut für Technologie, Karlsruhe, Deutschland

The Pierre Auger Observatory was designed to answer the key questions about the origin and composition of ultra-high energy cosmic rays. The Underground Muon Detector (UMD) is optimized to perform a direct measurement of the muon component of the air showers in the ankle-region of the energy spectrum. To estimate the number of muons, one of the most sensitive observables to the mass composition of primary cosmic rays, the UMD works in two complementary ways dubbed as binary (or counting) and ADC (or integrator) modes. The first relies on the amplitude of the signals, the latter on its charge. In this work, we will focus on the integrator mode, where the number of muons can be estimated once the average charge left by a single muon is known. In previous analysis, we showed how to calibrate the integrator based on simulations. Now we study evolution of the calibration data, including monitoring variables involved in the calibration process. We also perform comparisons between data and simulation results and report on most recent developments and interpretations.

T 47.2 Tue 16:30 T-H32 Compatibility of trigger and timing between the non-upgraded and the upgraded electronics of the Pierre Auger Observatory. — •FABIO CONVENGA for the Pierre Auger-Collaboration — Karlsruher Institut für Technologie (IAP), Karlsruhe, Germany

The surface detectors of the Pierre Auger Observatory are being upgraded by adding new detectors and replacing electronics.

The upgraded electronics, dubbed Upgraded Unified Board (UUB), is able to acquire data from the new detectors. It includes an improved GPS receiver, a higher sampling rate, and a more powerful logic capacity. The new features of UUB made it possible to introduce new types of triggers. Despite this, to ensure backward compatibility, pre-upgrade triggers are implemented using digitally filtered and downsampled waveforms to simulate the triggering behavior in the non-upgraded stations.

The logic functionality of the UUB also includes a module for event timing. The fundamental architecture of this module is parallel to that used in nonupgraded electronics. The on-board software that manages the module is similar to that of the non-upgraded electronics, with minor modifications required for the new UUB hardware.

In this talk, we will present the first analysis on the compatibility of the triggering efficiency and timing focusing on two neighboring stations one with the non-upgraded electronics and the other with the UUB.

T 47.3 Tue 16:45 T-H32

Location: T-H32

Performance and Calibration of the Upgraded Surface Detector of the Pierre Auger Observatory — •ALEXANDER STREICH, DAVID SCHMIDT, DARKO VE-BERIC, MARKUS ROTH, and RALPH ENGEL for the Pierre Auger-Collaboration — Karlsruher Institut für Technologie (KIT), Karlsruhe, Deutschland

The AugerPrime upgrade defines the transition to the next measurement stage of the Pierre Auger Observatory and is rolled out among other things with the addition of new scintillation detectors on top of the water-Cherenkov detectors of the Surface Detector. These Surface Scintillator Detectors provide a complementary measurement of the secondary particles enabling the discrimination of the individual air shower components on an event-to-event basis. This leads to an significantly improved determination of the properties of the ultra-high energy cosmic rays, for example their mass composition. Besides the scintillation detectors, the electronics boards of each station will be replaced with an enhanced version with an increased signal resolution, a higher sampling frequency, as well as faster data processing. This presentation focuses on the analyses of the performance of these hardware components under the measurement conditions in the Observatory. Thereby, the general operation stability and hardware requirements are investigated. Additionally, the performance of the air shower reconstruction applied to the first array of AugerPrime stations is analyzed, including the necessary adaptations to the detector calibration procedure and the signal discrimination capability of the hybrid event detection utilizing the scintillation and Cherenkov detectors.

T 47.4 Tue 17:00 T-H32

Test of the front-end electronics of the AugerPrime Radio Detector* — STEPHAN KELLER and •JULIAN RAUTENBERG for the Pierre Auger-Collaboration — Bergische Universität Wuppertal, Gaußstr. 20, 42119 Wuppertal

As part of the AugerPrime extension of the Pierre Auger Observatory each of the 1660 Surface Detector stations will be upgraded for the radio detection of air showers. The corresponding specialized front-end electronics has been developed within the collaboration for optimal performance in the cosmic ray detection. To ensure the quality of the electronics before deployment over the 3000 km² of the Observatory they are temperature cycled between -25 and 75 degree for 48 hours. At four temperatures their performance is repeatedly measured with a defined set of test-signal inputs. The setup to test the 2000 electronic boards will be presented together with the results of the first pre-production boards.

*Gefördert durch die BMBF Verbundforschung Astroteilchenphysik (Vorhaben 05A20PX1).

T 47.5 Tue 17:15 T-H32

CRPropa 3.2 — an advanced framework for high-energy particle propagation in extragalactic and galactic spaces — •JULIEN DÖRNER for the CRPropa-Collaboration — Theoretische Physik IV, Ruhr University Bochum, Bochum, Germany — RAPP-Center at Ruhr University Bochum, Bochum, Germany

The landscape of high- and ultra-high-energy astrophysics has changed in the last decade, largely due to the inflow of data collected by large-scale cosmic-ray, gamma-ray, and neutrino observatories. At the dawn of the multimessenger era, the interpretation of these observations within a consistent framework is important to elucidate the open questions in this field. CRPropa 3.2 is a Monte Carlo code for multimessenger simulations of high-energy particles and their secondaries. With this new version, the framework is now extended to more than extragalactic propagation opening up the possibility to more astrophysical applications, like Galactic cosmic-ray and local source modeling.

In this talk, we will show some of these new aspects that can be applied with CRPropa 3.2. It will include simulations of high-energy particles in diffusion dominated domains and self-consistent, fast modelling of electromagnetic cascades and interactions with customized photon fields. With the new CRPropa 3.2 version several technical updates and improvements were implemented, which will be presented in the talk.

T 47.6 Tue 17:30 T-H32

Testing the Starburst Galaxy and Active Galactic Nuclei correlation result for Pierre Auger Observatory with CRPropa simulations* — •WILSON NAMASAKA — Bergische Universität Wuppertal, Gaußstr. 20, 42119 Wuppertal, Germany. Intermediate scale anisotropies in the distribution of Ultra-High Energy Cosmic Rays (UHECRs) arrival directions can be associated with two prominent classes of extragalactic gamma-ray sources detected by Fermi-LAT. In most recent study, a correlation between the arrival direction of cosmic rays at energies above 38 EeV for starburst galaxies (SBG) and 39 EeV for active galactic nuclei (AGN) was reported by the Pierre Auger Collaboration with a significance of 4.5 σ and 3.1 σ , respectively. The cosmic ray excess models for these sources used an angular smearing parameter to fit the observed arrival direction distribution in an optimization scan.

The viability of this angular smearing using CRPropa is investigated in this research, to test whether the results of the Pierre Auger Observatory can be reproduced by the deflections expected due to magnetic fields. The five strongest gamma-ray sources in both the Fermi-LAT AGN and SBG catalogs based on flux weight have been selected. In this talk, results from the simulations and expected angular deflections shall be presented.

*Funded by DAAD under reference No.91653888.

T 47.7 Tue 17:45 T-H32

Extending CRPropa with hadronic interactions for in-source propagation of UHECRs* — •LEONEL MOREJON — Bergische Universität Wuppertal, Gaußstr. 20, 42119 Wuppertal

CRPropa is a versatile Monte Carlo code for the propagation of Ultra-High Energy Cosmic Rays (UHECRs). Part of its strength is the possibility to combine it with user code to provide additional functionality. For example, it has been employed with post-simulation computations to study sources where hadronic interactions are of interest. However, when such interactions are too important, like may be the case in certain bursting sources, this type of approach is inconsistent.

This contribution reports on the ongoing effort to extend CRPropa with the inclusion of hadronic interactions. The approach involves exposing hadronic modelling softwares (*e.g.* EPOS-LHC, SYBILL, QGSJet) to the main code, and thus, make them available to all users natively (*i.e.* no need for additional coding). The result is a new module within CRPropa that computes hadronic interactions in runtime, along with the other interactions in the code. The performance of this new module is profiled and simple case studies are selected to illustrate its use. * *Funded by the Deutsche Forschungsgemeinschaft through project number* 445990517.

T 47.8 Tue 18:00 T-H32

Advances in parallelization of cosmic rays simulations in CORSIKA 8 — •ANTONIO AUGUSTO ALVES JUNIOR, PRANAV SAMPATHKUMAR, and RALF UL-RICH for the CORSIKA 8-Collaboration — Institute for Astroparticle Physics (IAP) - KIT

Advances in the parallelization of CORSIKA 8, which is being developed in modern C++17 and is designed to run on multi-thread modern processors and accelerators, are discussed.

Aspects such as out-of-the-order calculations, generation of high quality random numbers and fast task scheduling and submission on massively parallel platforms are highlighted, followed by presentation of preliminary performance measurements.

Finally, the design choices and integration into CORSIKA 8 are presented, together with some basic examples.

T 48: Neutrino Physics without Accelerators 3

Location: T-H33

Time: Tuesday 16:15–18:45

T 48.1 Tue 16:15 T-H33

Improved measurement of the neutrino mixing angle θ_{13} with Double Chooz — •PHILIPP SOLDIN, LARS HEUERMANN, ACHIM STAHL, and CHRISTOPHER WIEBUSCH — III. Physics Institute B, RWTH Aachen University

Double Chooz is a reactor neutrino disappearance experiment that was operating between 2011 until 2018. Its primary purpose was the precise measurement of the neutrino mixing angle θ_{13} . The experimental setup consisted of two identical liquid scintillator detectors at average baselines of about 400 m and 1 km to two nuclear reactor cores in Chooz, France. The neutrinos were detected by measuring the inverse beta decay (IBD) signature, which consists of a prompt positron annihilation and a delayed neutron capture signal. The simultaneous measurement of the neutrino energy spectra with two detectors is used in a Poisson-based likelihood fit to obtain the neutrino mixing angle θ_{13} . This fit takes into account the respective neutrino fluxes, their energy spectra, all relevant backgrounds, and correlated uncertainties. The final data set, latest results, and novel ideas are presented in this talk.

T 48.2 Tue 16:30 T-H33

Likelihood based Sterile Neutrino Search with Double Chooz — •LARS HEUERMANN, PHILIPP SOLDIN, ACHIM STAHL, and CHRISTOPHER WIEBUSCH — RWTH Aachen University - Physics Institute III B, Aachen, Germany

The Double Chooz experiment was a reactor electron anti-neutrino disappearance experiment with the purpose of measuring the neutrino oscillation angle $\theta_{13}.$ It used two nearly identical detectors at the respective distances of 400m and 1050m to the reactors of the nuclear power plant in Chooz, France, for this purpose.

Sterile neutrinos are hypothetical neutrino mass states, which do not participate in weak interactions, but still could participate in neutrino oscillations and thus would be measurable with Double Chooz. We present a search which focusses on a profile likelihood approach to test the data with respect to the parameters of a model with a single additional sterile neutrino state (3+1 model). A particular challenge thereby is that Wilks' theorem is not fulfilled and computationally intensive parameter scans have to be used to estimate the test statistics. The talk explains the analysis method and presents the median sensitivity for the sterile mixing parameter θ_{14} in dependence of Δm_{41}^2 , as well as an analysis of the available experimental dataset.

T 48.3 Tue 16:45 T-H33

OSIRIS - a radiopurity detector for JUNO — •CHRISTIAN WYSOTZKI for the JUNO-Collaboration — RWTH Aachen University - Physics Institute III B, Aachen, Germany

The Jiangmen Underground Neutrino Laboratory (JUNO) is a 20 kt liquid scintillator based neutrino observatory, which is currently under construction in southern China.

OSIRIS (Online Scintillator Internal Radioactivity Investigation System), as one of JUNOs subsystems, monitors the radiopurity of the liquid scintillator. Stringent limits regarding the contamination of radioactive isotopes, especially uranium and thorium have been defined to ensure the success of JUNOs physics program.

The assembly and commissioning of OSIRIS is foreseen for 2022. This talk will give an overview of the design philosophy, the detector systems, and the current status.

T 48.4 Tue 17:00 T-H33

Calibration of the JUNO pre-detector OSIRIS — •MORITZ CORNELIUS VOLLBRECHT^{1,2}, ALEXANDRE GÖTTEL^{1,2}, PHILIPP KAMPMANN³, RUNXUAN LIU^{1,2}, LIVIA LUDHOVA^{1,2}, NIKHIL MOHAN^{1,2}, LUCA PELICCI^{1,2}, and MARIAM RIFAI^{1,2} — ¹Forschungszentrum Jülich GmbH, Nuclear Physics Institute IKP-2, Jülich — ²III. Physikalisches Institut B, RWTH Aachen University, Aachen — ³GSI Helmholtzcentre for Heavy Ion Research, Darmstadt

The multi-kton liquid scintillator (LS) detector of the Jiangmen Underground Neutrino Observatory (JUNO) experiment, currently under construction in Southern China, has a vast potential for insights in several fields of (astro-) particle physics. To achieve its goals of determining the neutrino mass ordering, stringent radiopurity levels are required. To ensure these limits, the Online Scintillator Internal Radioactivity Investigation System (OSIRIS) was designed as a predetector for JUNO. In OSIRIS, 76 self-triggering intelligent PMTs (iPMTs) instrument a watershielded 20 ton liquid scintillator target. During the monthslong filling of JUNO, a fraction of each purified scintillator batch passes through OSIRIS and its radiopurity is closely monitored. This enables fast countermeasures on possible contaminations. Multiple calibration systems are employed in OSIRIS. An Automatic Calibration Unit (ACU) of the Daya Bay experiment is used to calibrate energy and vertex event reconstructions as well as iPMT timing and charge responses. A separate laser system is used for the timing calibration of the iPMTs. This presentation will summarize the current status of the work concerning the calibration strategy of OSIRIS.

T 48.5 Tue 17:15 T-H33 Liquid Handling System of the OSIRIS detector — MICHAEL WURM, HANS STEIGER, KAI LOO, ERIC THEISEN, and •OLIVER PILARCZYK — Johannes Gutenberg-Universität Mainz

The Jiangmen Underground Neutrino Observatory (JUNO) is a 20 kt liquid scintillator experiment currently under construction in Jiangmen (China). Its main scientific goal is to determine the neutrino mass ordering by measuring antielectron neutrinos from two nearby nuclear power plants at a distance of 53 km. To achieve this goal the liquid scintillator has to go through several purification plants on site to make sure it meets the optical and radiopurity requirements.

The $20m^3$ OSIRIS pre-detector is the last device behind these purification plants and will be constructed in a underground hall close to the main JUNO detector. Its task is to monitor the radiopurity of the purified scintillator before it is filled in the JUNO detector and it will reach sensitivity levels of $10^{-16}g/g$ on Uranium and Thorium. OSIRIS is expected to be operated in a continuous mode, which means that parts of the scintillator from the main filling line will be redirected into a bypass in which the OSIRIS detector is placed and then being send on back into the main filling line. To make sure every batch of the scintillator stays about 24 h inside the OSIRIS detector a temperature gradient will be established in the detection volume. This talk covers the operation of the Liquid Handling System (LHS) and the included Level Measurement (LM) which control and oversee the operation of OSIRIS. The development is funded by the DFG Research Unit *JUNO* (FOR2319) and the Cluster of Excellence *PRISMA*⁺.

T 48.6 Tue 17:30 T-H33

The Laser calibration system of the JUNO pre-detector OSIRIS - •TOBIAS STERR, LUKAS BIEGER, DAVID BLUM, MARC BREISCH, SRIJAN DELAMPADY, JESSICA ECK, GINA GRÜNAUER, TOBIAS HEINZ, BENEDICT KAISER, FRIEDER KOHLER, TOBIAS LACHENMAIER, AXEL MÜLLER, ALEXANDER TIETZSCH, and JAN ZÜFLE — Eberhard Karls Universität Tübingen, Physikalisches Institut The Jiangmen Underground Neutrino Observatory is a multi-kton, multipurpose Neutrino observatory currently under construction in Kaiping, Southern China. It has a vast potential for progress in several field of (astro-) particle physics (e.g., the neutrino mass ordering, supernova neutrinos, geoneutrinos, etc.). To achieve the goals, a thorough control of the radiopurity levels of the liquid scintillator is required. Therefor the Online Scintillator internal Radioactivity Investigation System (OSIRIS) was introduced as a pre-detector to JUNO. OSIRIS features 76 self-triggering intelligent PMTs (iPMTs) as well as a watershielded, 20-ton liquid scintillator target. During the filling phase of the main detector, OSIRIS will measure the radiopurity of a fraction of each batch of liquid scintillator passed to JUNO. To enable OSIRIS to reach the necessary sensitivity to perform this task, two calibration systems will be available: An automated calibration unit and a laser calibration system. This talk will focus on the laser calibration system which will be used for both, timing and charge calibration of the iPMTs. Additionally, the preliminary calibration strategy of the laser system of OSIRIS will be presented. This work is supported by the Deutsche Forschungsgemeinschft (DFG)

Tuesday

T 48.7 Tue 17:45 T-H33

Final Results of the PoLiDe Experiment — •ULRIKE FAHRENDHOLZ¹, LOTHAR OBERAUER¹, HANS STEIGER^{1,2}, RAPHAEL STOCK¹, DAVID DÖRFLINGER¹, MARIO SCHWARZ¹, and KONSTANTIN SCHWEIZER¹ — ¹Technische Universität München (TUM), Physik-Department, James-Franck-Str. 1, 85748 Garching bei München — ²Institute of Physics and Excellence Cluster PRISMA+, Johannes Gutenberg-Universität (JGU) Mainz, Staudingerweg 9, 55099 Mainz

The Jiangmen Underground Neutrino Observatory (JUNO) aims to determine the neutrino mass ordering by detecting reactor electron-antineutrinos. In the main detection channel, the inverse beta decay (IBD), a positron is produced, which is used to reconstruct the neutrino*s initial energy. If the positron forms a bound state with spin 1, called ortho-Positronium (o-Ps), with an electron prior to its annihilation, the time structure of the event is distorted due to the longer lifetime of o-Ps. The Positron Lifetime Determination (PoLiDe) experiment was built to determine the lifetime and formation probability of o-Ps in liquid scintillators (LSs). In this talk, the final results of the lifetime measurements for different organic LSs, including the JUNO mixture, and a water-based LS sample are presented. This work is supported by the DFG research unit "JUNO", the DFG collaborative research centre 1258 "NDM", and the DFG Cluster of Excellence "Origins.

T 48.8 Tue 18:00 T-H33

Measurements of the attenuation length and the group velocity of liquid scintillators with the CELLPALS method — •JESSICA ECK, LUKAS BIEGER, DAVID BLUM, MARC BREISCH, SRIJAN DELAMPADY, GINA GRÜNAUER, TOBIAS HEINZ, BENEDICT KAISER, FRIEDER KOHLER, TOBIAS LACHENMAIER, AXEL MÜLLER, TOBIAS STERR, ALEXANDER TIETZSCH, and JAN ZÜFLE — Eberhard Karls Universität Tübingen, Physikalisches Institut

The Jiangmen Underground Neutrino Observatory (JUNO) is currently constructed in southern China with the main goal to determine the neutrino mass hierarchy by detecting reactor antineutrinos. The JUNO detector consists of a large spherical vessel filled with 20ktons of highly transparent liquid scintillator. To quantify the transparency, a measurement of the attenuation length is crucial, however, this poses a challenge for attenuation lengths of several tens of meters due to the necessity of a sufficient long light path through the sample.

This talk will present the CELLPALS method to measure the attenuation length of liquid scintillators using an optical cavity to extend the effective light path. In addition, the CELLPALS method also provides the determination of the group velocity of the sample. The experimental setup and the results for different samples will be presented.

This work is supported by the Deutsche Forschungsgemeinschaft.

T 48.9 Tue 18:15 T-H33

Fluorescence Time Profiles of the JUNO and TAO Liquid Scintillators using a Pulsed Neutron Beam in the Energy Range from 3.5 to 5.5 MeV — •MATTHIAS RAPHAEL STOCK¹, HANS TH. J. STEIGER^{1,2}, DAVID DÖRFLINGER¹, STEFAN SCHOPPMANN³, ULRIKE FAHRENDHOLZ¹, LOTHAR OBERAUER¹, LUCA SCHWEIZER¹, KORBINIAN STANGLER¹, and DORINA ZUNDEL² — ¹Physik-Department, TU München, James-Franck-Str. 1, 85748 Garching — ²JGU Mainz, Cluster of Excellence PRISMA+, Staudingerweg 9, 55128 Mainz — ³University of California, Department of Physics, Berkeley, CA 94720-7300, USA

We simultaneously performed two liquid scintillator (LS) characterization experiments using a pulsed neutron beam at the CN accelerator of INFN Laboratori Nazionali di Legnaro. At energies ranging from 3.5 MeV to 5.5 MeV, one experiment measures the quenching factor of recoil protons while the one other measures the scintillation time profile of recoil protons. This talk is about the time profile experiment, where we show results of LS mixtures for the upcoming neutrino experiments JUNO and TAO in China as well as for future projects e.g., Theia, which will use a water-based LS. Differences in the time profiles after gamma and neutron excitation allows to perform pulse shape discrimination and therefore advances the ability to distinguish the neutrino signal from background. This work is supported by the BMBF Verbundforschung 05H2018 "R&D Detektoren (Szintillatoren)", the DFG research unit "JUNO", the DFG CRC 1258 "NDM" and the DFG Clusters of Excellence "PRISMA+" and "Origins".

T 48.10 Tue 18:30 T-H33

Quenching Factor Studies for Organic Liquid Scintillators with a Pulsed Neutron Beam — •HANS TH. J. STEIGER^{1,2}, M. RAPHAEL STOCK², DAVID DÖRFLINGER², STEFAN SCHOPPMANN³, MANUEL BÖHLES¹, ULRIKE FAHRENDHOLZ², LOTHAR OBERAUER², LUCA SCHWEIZER², KORBINIAN STANGLER², and DORINA ZUNDEL¹ — ¹JGU Mainz, Cluster of Excellence PRISMA+, Staudingerweg 9, 55128 Mainz, Germany — ²Physik-Department, TU München, James-Franck-Str. 1, 85748 Garching, Germany — ³UC Berkeley, Department of Physics, CA 94720-7300, USA

Leading current multi-ton-scale neutrino experiments rely on the successful application of liquid scintillators (LS). Therefore, detailed understanding of the detectors and their detection media are crucial to construct realistic simulations and predictions of phenomena. With our measurements at the CN accelerator of the INFN Laboratori Nazionali di Legnaro, we conduct precision characterizations of LSs in terms of proton quenching and fluorescence spectra. In these experiments neutron interactions are distinguished from beam correlated gammas by time-of-flight measurements. During several beamtimes in the past two years we studied samples of scintillators from the experiments JUNO, TAO, Borexino, Theia and SNO+. Precise data about the quenching factors of proton and 12C-recoils for all measured liquid scintillators were gained for the first time. This work is supported by the Cluster of Excellence PRISMA+, the Bundesministerium für Bildung und Forschung (Verbundprojekt 05H2018: R&D Detectors and Scintillators) and the DFG Research Unit JUNO.

T 49: Neutrino Physics without Accelerators 4

Time: Tuesday 16:15-18:40

Group Report

T 49.1 Tue 16:15 T-H34 Direct neutrino mass measurement with the Project 8 experiment: status and outlook — •LARISA THORNE and MARTIN FERTL for the Project 8-Collaboration - Institute of Physics and Excellence Cluster PRISMA+, Johannes Gutenberg University Mainz, 55099 Mainz, Germany

There have been significant gains in characterizing neutrino properties in recent decades, however the absolute neutrino mass scale continues to be elusive. The Project 8 experiment seeks to probe this quantity directly via kinematic analysis of atomic tritium single beta decay, using the novel CRES (cyclotron radiation emission spectroscopy) technique. CRES employs a frequency-based approach to measure the differential tritium beta decay spectrum in the endpoint region, where the spectral shape is most sensitive to distortions from a non-zero neutrino mass. Here we present a roadmap of Project 8's milestones towards a neutrino mass measurement with a final design sensitivity of 40 meV. This includes recent results from a successful first-time demonstration of the CRES technique with molecular tritium, as well as status updates on the components comprising the experiment's future full-scale version.

T 49.2 Tue 16:35 T-H34 Atomic hydrogen beam monitor for Project 8 — • CHRISTIAN MATTHE and SE-BASTIAN BÖSER for the Project 8-Collaboration — Institute of Physics and Excellence Cluster PRISMA+, Johannes Gutenberg University Mainz, 55099 Mainz, Germany

The Project 8 collaboration aims to determine the absolute neutrino mass with a sensitivity of 40 meV by measuring the tritium decay spectrum around the endpoint energy. For this level of precision it is necessary to use atomic tritium, since molecular tritium sensitivity is limited by the molecular final state distribution to about 100 meV. We anticipate using an atomic tritium flux of $\approx 10^{19}$ atoms/s from a source to inject a beam with $\approx 10^{15}$ atoms/s of the proper state and temperature into the detection volume.

For monitoring this beam, we are developing a detector that uses a wire with a micrometer-scale diameter intersecting the beam on which a small fraction of the beam's hydrogen atoms recombine into molecules. The energy released heats the wire and produces a measurable change in its resistance. Using either a grid of wires or a sweep with a single wire the beam profile will be determined. Such a detector is suitable for both development work and for minimally disruptive online monitoring in the final experiment. In this talk I will present first results using such a detector in the beam of the Mainz atomic hydrogen setup.

T 49.3 Tue 16:50 T-H34

Atom-source development for Project 8 — •ALEC LINDMAN and SEBASTIAN BÖSER for the Project 8-Collaboration — Institute of Physics and Excellence Cluster PRISMA+, Johannes Gutenberg University Mainz, 55099 Mainz, Germany

The Project 8 experiment aims to make a direct measurement sensitive to much of the unexplored range of neutrino masses. Past experiments used molecular tritium, which has a large energy smearing from its final states. Project 8 will use atomic tritium to reach $m_{\beta} \leq 40$ meV. This requires $\mathcal{O}(10^{20})$ T atoms held at tens of mK in a several-cubic-meter magnetic trap. The efficiencies of cooling the atoms and their trapped lifetime require $>10^{19}$ atoms/s at the source. Phase III of Project 8 will include an Atomic Tritium Demonstrator to confirm we can produce, cool, and trap sufficient atomic T for the final Phase IV experiment.

I will discuss work at the University of Mainz to develop a high-flux tritiumcompatible atom source. Our tests extend to a hydrogen flow of 20 sccm, 20 times the previously-published values for this type of source. Recent progress includes a redesign that boosted the atomic signal 100-fold and separation of the atom signal from background via low-energy ionization. Upgrades are underway to definitively determine if the present atom source provides sufficient atomic flux for Project 8's neutrino mass sensitivity. Designs for a higher-output source, if needed, and subsequent cooling and trapping stages are in progress and will be tested in due course.

T 49.4 Tue 17:05 T-H34

Project 8 Free Space CRES Demonstrator: Signal characteristics and matched-filter detection — • René Reimann, Florian Thomas, and Martin FERTL for the Project 8-Collaboration — Institute of Physics and Excellence Cluster PRISMA+, Johannes Gutenberg University Mainz, 55099 Mainz, Germany

Location: T-H34

The Project 8 collaboration aims to measure the neutrino mass with a sensitivity of 40 meV by measuring the endpoint region of the atomic tritium beta decay spectrum using the new technology of Cyclotron Radiation Emission Spectroscopy (CRES). While the measurement principle of CRES has been successfully demonstrated using the enclosed volume of a microwave guide filled with molecular tritium or krypton, one major challenge is to scale up the source volume in order to increase the overall statistics. One promising approach is to leave the confined space of the microwave guide and detect the cyclotron radiation emitted by the electron in free space using an array of antennas. To investigate the CRES technique in free space the so called Free Space CRES Demonstrator (FSCD) is under development. Because the signal is diluted into 4π an antenna array with a high number of readout channels is required, which drastically increases the data rate. Therefore real-time processing, triggering and reconstruction are required in the FSCD. The signal characteristics are mainly dominated by the magnetic field and the antenna response. In this talk we present the influence of the magnetic field on the signal spectrum and present how matched filtering can be used to detect and reconstruct CRES signals.

Group Report

T 49.5 Tue 17:20 T-H34

The SNO+ experiment: current status and future prospects - •JOHANN DITTMER and KAI ZUBER — IKTP, TU Dresden, Deutschland SNO+ is a large liquid scintillator based experiment reusing the infrastructure

of the successful Sudbury Neutrino Observatory (SNO). Located 2 km underground in a mine near Sudbury, Ontario, Canada, the detector consists of 12 m diameter acrylic vessel which is filled with 780 tonnes of a liquid scintillator. For the main goal, the search for the neutrinoless double beta decay $(0\nu\beta\beta)$ of 130 Te, the scintillator will be doped by 0.5% natural Tellurium. Since SNO+ was designed as a general purpose neutrino detector, it is also possible to measure neutrinos from different sources (reactor, geo, solar, Supernova, etc.). After a commissioning water phase which was ended in 2018, the scintillator fill was completed in April 2021.

In this talk the recent results and broad physics program will be presented. SNO+ is supported by the German Research Foundation (DFG).

T 49.6 Tue 17:40 T-H34

Column Density Determination for the KATRIN Neutrino Mass Measurement — •Christoph Köhler¹, Fabian Block², and Alexander MARSTELLER² for the KATRIN-Collaboration - ¹Technical University of Munich/Max Planck Institute for Physics — ²Karlsruhe Institute of Technology

The KATRIN experiment aims to model-independently probe the effective electron anti-neutrino mass with a sensitivity of 0.2 eV (90 % CL) by investigating the endpoint region of the tritium beta decay spectrum. To achieve this goal the gas quantity of the windowless gaseous tritium source, characterized by the column density, has to be known with great accuracy.

We present in this talk the principle of measuring the column density with an angular resolved photoelectron source and report on the monitoring accuracy of the column density achieved with dedicated activity monitoring devices in the first five measurement campaigns of KATRIN. The influence of the column density uncertainty on the neutrino mass determination is then discussed in light of KATRIN's world-leading direct upper limit on the neutrino mass and the ongoing further data-taking.

This work is supported by the Technical University of Munich, the Max Planck Society, the Helmholtz Association (HGF), the Ministry for Education and Research BMBF (05A17PM3, 05A17PX3, 05A17VK2, and 05A17WO3), the Helmholtz Alliance for Astroparticle Physics (HAP), the GRK 1694, and the Helmholtz Young Investigator Group (VH-NG-1055).

T 49.7 Tue 17:55 T-H34

^{83m}Kr N-line spectrum measurement at KATRIN — MATTHIAS BÖTTCHER¹, MORITZ MACHATSCHEK², MAGNUS SCHLÖSSER², and •JAROSLAV STOREK² for the KATRIN-Collaboration — 1 Institute of Nuclear Physics, University of Münster — ²Institute for Astroparticle Physics, Karlsruhe Institute of Technology

The KArlsruhe TRItium Neutrino experiment currently provides the best neutrino mass upper limit of 0.8 eV/ c^2 (90% C. L.) in the field of direct neutrinomass measurements. This result has been obtained with only 5% of the anticipated total measurement time. However, reaching the target sensitivity of 0.2 eV/c^2 at 90% C. L. not only requires the full measurement time, but also the detailed study of systematic measurement uncertainties. Several of them can be studied by measuring a shape distortion of the ^{83m}Kr intrinsic electron conversion N-lines which creates high demands on precise knowledge of the undistorted spectrum. Results of a dedicated measurement of the intrinsic ^{83m}Kr Nspectrum conducted at unprecedented precision at KATRIN will be presented in this talk.

This work is supported by the Ministry for Education and Research BMBF (05A17PM3, 05A17PX3, 05A17VK2, and 05A17WO3) and the Doctoral School "Karlsruhe School of Elementary and Astroparticle Physics: Science and Technology (KSETA)" through the GSSP program of the German Academic Exchange Service (DAAD).

T 49.8 Tue 18:10 T-H34

Unmodeled features in the KATRIN spectrum as a hint for unaccounted systematic effects * — •KAROL DEBOWSKI for the KATRIN-Collaboration — Bergische Universität Wuppertal, Wuppertal, Deutschland

The Karlsruhe Tritium Neutrino (KATRIN) Experiment is designed and operated to determine the mass of the electron-antineutrino with a final sensitivity of 200meV (90% C.L.) using the radioactive beta decay spectrum of tritium. In order to achieve the design sensitivity, a precise knowledge of all systematic effects is needed.

Simulations of the energy spectrum take all known systematic effects into account which are also used in the neutrino mass analysis. Using the measured slow control variables as inputs to the simulation, the measured and simulated spectra should agree with each other except for statistical fluctuations in the measured data. Any significant features exceeding those statistical fluctuations can be potential hints towards systematic effects which are not considered (correctly) in the simulation and hence also in the analysis. By intentionally introducing unprecisely modeled systematics into simulations, the nature and origin of potential unknown effects can be estimated for better understanding and modeling of the measured spectrum in the final analysis. This work aims on a simulation based proof of concept for upcoming analyses to find unaccounted systematic effects in the experiment.

* This work is supported by the Ministry for Education and Research BMBF (05A20PMA, 05A20PX3, 05A20VK3)

T 49.9 Tue 18:25 T-H34

Increasing KATRIN's luminosity by an enlarged acceptance angle — •EMANUEL WEISS, JAN BEHRENS, FERENC GLÜCK, and STEPHANIE HICKFORD for the KATRIN-Collaboration — Institute for Astroparticle Physics and Institute of Experimental Particle Physics, Karlsruhe Institute of Technology

The KATRIN collaboration aims to determine the neutrino mass with a sensitivity of 0.2 eV/ c^2 (90% CL). This will be achieved by measuring and fitting the endpoint region of the tritium β -electron spectrum. The integral spectrum is measured by a MAC-E filter, which features a high acceptance angle for electrons emitted by a high-luminosity, isotropically emitting tritium source, $\Delta\Omega/2\pi = 1 - \cos\theta_{\rm max}$.

One approach to improve the statistical uncertainty of the experiment is to further the acceptance angle θ_{max} , which depends on the ratio of source and maximum magnetic field. This can be achieved by keeping the source magnetic field at standard setting and scaling down the magnetic fields in the rest of the beam line. The changed electromagnetic conditions lead to increased β -electron statistics and influence several systematic effects. These effects, as well as the gain in statistics compared to the standard magnetic field settings, are evaluated by simulations and measurements that are presented in this talk.

This work is supported by the Helmholtz Association (HGF), the Ministry for Education and Research BMBF (05A17PM3, 05A17PX3, 05A17VK2, and 05A17WO3), the Helmholtz Alliance for Astroparticle Physics (HAP), and the Helmholtz Young Investigator Group (VH-NG-1055).

T 50: Search for Dark Matter 2

Time: Tuesday 16:15-18:00

T 50.1 Tue 16:15 T-H35

Optimisation of the TPC aspect ratio for the DARWIN observatory — •SEBASTIAN VETTER — Institute for Astroparticle Physics (IAP), Karlsruhe Institute of Technology (KIT)

DARWIN is a proposed next-generation astroparticle physics observatory focused on the direct detection of WIMP Dark Matter. It will use 40 tonnes of natural xenon inside a dual-phase Time Projection Chamber (TPC) following the technology of the successful detectors that use noble elements as active material. Due to the elusive nature of Dark Matter and the envisaged unprecedented discovery potential, DARWIN requires an ultra-low background level.

The final design of many components of the detector is not decided yet. One of the central parameters for optimization is the aspect ratio (AR) of the TPC. A higher AR leads to better light collection and shorter electron drift lengths, but also comes with different fractions of sensor and cryostat materials, being a source of external background.

In this talk we present results based on detailed 3-dim Monte Carlo simulations on the influence of the AR on the S1 signal detection efficiency for WIMPs as well as on background components such as dark count rates and neutrons from the TPC materials.

T 50.2 Tue 16:30 T-H35 **The Freiburg DARWIN Demonstrator** – •JULIA MÜLLER – University of Freiburg

Liquid xenon (LXe) time projection chambers (TPCs) are the leading detector technology for searches for dark matter in form of WIMPs. DARWIN will be the ultimate LXe-based dark matter detector covering the entire accessible parameter space for WIMP masses above a few GeV/c^2 , superseding current detectors in size and sensitivity. The technical realization of its central low-background TPC with a diameter of about 2.6m will be very challenging due to the size of the detector, the low-temperature operation, and the required radiopurity levels. The DARWIN detector test platform PANCAKE at the University of Freiburg will be used to develop and test flat detector components with diameters up to the DARWIN-scale. We will present an overview of the platform that can accomodate up to 400 kg of xenon gas and present first results from the commissioning phase.

T 50.3 Tue 16:45 T-H35

Full MonteCarlo simulations of the cosmogenic background for the DAR-WIN observatory at different underground locations — •JOSE CUENCA-GARCÍA for the DARWIN-Collaboration — Institute for Astroparticle Physics (IAP), Karlsruhe Institute of Technology (KIT)

The DARWIN observatory is a proposed next-generation experiment focused on the direct detection of Dark Matter. It will use 40 tonnes of natural xenon inside

a dual-phase Time Projection Chamber (TPC) being the evolution of the detectors that use noble elements as active material. The final design of the detector and its location are not decided yet. Besides the direct detection of Dark Matter a large variety of science channels can be studied, as for example 0vbb, Axion Like Particles or solar neutrinos, among others. To fully exploit these physics goals an ultra-low background level is required. Although this type of experiments is located in underground laboratories to shield them against the cosmic radiation, muons and their induced secondary particles can still contribute significantly to the background. This *in situ* production cannot be suppressed and further veto systems are needed. We present here the simulations of the cosmogenic background for several underground laboratories. In particular, we focus on the production of some nuclei that can potentially affect the science channels of interest.

T 50.4 Tue 17:00 T-H35

Location: T-H35

DARWIN background estimations through multi-scatter separation — •MAIKE DOERENKAMP, ANTOINE CHAUVIN, ANDRII TERLIUK, and STEPHANIE HANSMANN-MENZEMER — Physikalisches Institut, Universität Heidelberg

The DARWIN experiment is a future multi-ton dual-phase xenon TPC, whose primary goal will be the search for WIMPs through nuclear recoil. One of the major backgrounds in WIMP-nucleus interactions are radiogenic neutrons. A single nuclear recoil caused by a neutron is indistinguishable from one caused by a WIMP. However, due to their much shorter mean-free path, more than 90% of neutrons scatter multiple times within the detector. This can be exploited for background rejection. This talk will describe a method to separate single- and multi-scatter events in a dual-phase xenon TPC and how this translates to the expected neutron background rates in DARWIN.

T 50.5 Tue 17:15 T-H35

Estimation of electronic recoil leakage into nuclear recoil signal for DAR-WIN — •ANTOINE CHAUVIN, MAIKE DOERENKAMP, ANDRII TERLIUK, and STEPHANIE HANSMANN-MENZEMER — Physikalisches Institut, Universität Heidelberg

The DARWIN experiment is a proposed future Direct Dark Matter detector which aims to detect WIMPs through WIMP-nucleus interactions, in a multiton liquid xenon target. Is goal is to become the most sensitive experiment to WIMP-nucleus interaction. To estimate this sensitivity, good models for signal and background generation, and of the detection processes, are fundamental. Electronic Recoil (ER) processes are the dominant background. Thus a good rejection of ER background, and an estimation of the ER leakage in Nuclear Recoil (NR) signal, is fundamental to achieve a high sensitivity. In this talk, I will report on the setup of a simulation of the DARWIN detection process, and its use to estimate the ER leakage fraction.

T 50.6 Tue 17:30 T-H35

Charge detection via proportional scintillation in a single-phase liquid xenon TPC - • FLORIAN TÖNNIES for the DARWIN-Collaboration — Albert-Ludwigs-Universität Freiburg, Deutschland

Dual-phase liquid/gas xenon TPCs are a well-established detector technology to search for WIMP Dark Matter. Nevertheless, the homogenious detection of the charge signal via proportional scintillation will be challenging at the scale of the next-generation detectors due to the size of the TPCs. The detection of the charge signal in the liquid phase of a single-phase TPC might be an option to circumvent this issue. In Freiburg we successfully operate a single-phase TPC demonstrator which exploits proportional scintillation in the strong electric field around thin wires. Some of the most recent results will be presented in this talk. The MonXe Radon Emanation Chamber — •DANIEL BAUR for the DARWIN-Collaboration — Albert-Ludwigs-Universität Freiburg

Liquid xenon-based experiments are currently leading the search for WIMP dark matter. Their electronic recoil background in the energy region of interest is dominated by the naked (i.e., not accompanied by the coincident emission of a gamma-ray) beta decays of ²¹⁴Pb, a progeny of ²²²Rn which is emanated from all material surfaces. Consequently, the reduction of ²²²Rn emanation is mandatory for the success of next-generation dark matter experiments with multi-ton xenon targets such as DARWIN.

ter y for the detects of head generation and the second se

T 51: Experimental Techniques in Astroparticle Physics 2

Time: Tuesday 16:15-18:20

T 51.1 Tue 16:15 T-H36

Comparison of Sky Models of the Galactic Radio Background for the Calibration of Radio Arrays — •MAX BÜSKEN for the Pierre Auger-Collaboration — Institut für ExperimentelleTeilchenphysik, Karlsruher Institut für Technologie, Karlsruhe, Deutschland

The Pierre Auger Observatory is the largest ground-based experiment for the detection of ultra-high energy cosmic rays. New radio antennas will be installed on each of the surface detector stations as part of the AugerPrime upgrade. This will allow to study the mass composition of cosmic rays arriving with large inclination angles.

Performing an accurate calibration and having a good understanding of its uncertainties is crucial for any physics analysis. Conducting a calibration campaign in the field with a reference antenna is not feasible on this large scale. Therefore the absolute calibration of the radio antennas will be performed using the diffuse galactic radio emission as an absolute reference, as it is the most dominant source of background. I will present a comparison of sky models that predict the galactic emission received from the whole sky. I will show how large the uncertanties on these predictions are and illustrate, what this means for radio experiments relying on this calibration method.

T 51.2 Tue 16:30 T-H36

Verbesserung des externen Triggers von AERA für ausgedehnte Luftschauer am Pierre-Auger-Observatorium — •RUKIJE UZEIROSKA für die Pierre Auger-Kollaboration — Bergische Universität Wuppertal

Das Pierre-Auger-Observatorium ist das größte Observatorium für kosmische Strahlung der Welt. Sein Auger Engineering Radio Array (AERA) besteht aus mehr als 150 Antennenstationen, die eine Fläche von etwa 17 km² abdecken, und dient der Erfassung von Radiosignalen, die von ausgedehnten Luftschauern emittiert werden. Diese Messungen werden verwendet, um die Eigenschaften der primären kosmischen Strahlung zu rekonstruieren, die die Luftschauer verursacht. Die Datennahme von AERA wird insbesondere extern durch die zentrale Datennahme des Observatoriums getriggert. Seine Funktionsweise führt dazu, dass viele Ereignisse getriggert werden, die für AERA nicht relevant sind. Zudem werden nicht alle interessanten Ereignisse als solche erkannt. Um den externen Trigger zu verbessern wird eine neue, vereinfachte Rekonstruktionsmethode für die Bestimmung der Richtung eines ausgedehnten Luftschauers vorgestellt. Die Zuverlässigkeit dieser Methode wird anhand von rekonstruierten Luftschauern getestet. Auf Basis der entwickelten Rekonstruktionsmethode werden verschiedene Trigger-Bedingungen definiert und getestet. Die optimale Trigger-Bedingung erreicht eine Effizienz von 99,87% für relevante AERA-Ereignisse, während sie die Gesamt-Datenrate um 49,98 % reduziert. Dies stellt eine substantielle Verbesserung gegenüber dem aktuellen AERA-Trigger dar.

T 51.3 Tue 16:45 T-H36

Depth of Maximum of Air-Shower Profiles at the Pierre Auger Observatory - •THOMAS FITOUSSI for the Pierre Auger-Collaboration — Karlsruhe Institute of Technology, Institute for Astroparticle Physics, Karlsruhe, Germany

The Pierre Auger Observatory is the largest ultra-high energy cosmic rays observatory in the world. Using a hybrid technique (fluorescence telescopes and surface detectors) it is possible to estimate the mass composition of cosmic rays. The main mass-sensitive observable measured with fluorescence telescopes is the depth of maximum of air-shower profiles called X_{max} .

In this presentation, we will present the analysis of the most recent dataset with a special focus on results at low energies down to $\sim 10^{17}$ eV. These low energy measurements are preformed with the High Elevation Auger Telescope (HEAT) and they allow to study the energy region where the transition between Galactic and extragalactic cosmic rays is expected.

Location: T-H36

T 51.4 Tue 17:00 T-H36

Separation of muonic and electromagnetic signals using the upgraded detectors of the Pierre Auger Observatory — •ALLAN MACHADO PAYERAS^{1,2}, ANDERSON CAMPOS FAUTH², DARKO VEBERIC¹, and DAVID SCHMIDT¹ — ¹Karlsruhe Institute of Technology, Karlsruhe, Germany — ²University of Campinas, Campinas, Brazil

The Pierre Auger Observatory detects extensive air showers (EAS), produced by high-energy cosmic rays. Its surface detector (SD) is composed of 1660 water-Cherenkov detectors (WCD) disposed in a triangular grid with a spacing of 1500 m between nearest neighbours. At the moment, the Observatory is being upgraded with the main addition of a surface scintillation detector (SSD), which will be installed on top of each WCD. The main goal of the upgrade is to obtain data sensitive to the composition of the primary cosmic rays, which is necessary to understand the astrophysical origin of these particles. In this work, we have studied a method to obtain a separation of signals due to the electromagnetic and muonic components of EAS, using the responses of the WCD and SSD. Such separation is the key to obtain composition-sensitive information from the new dataset. The signals of each of the components were modelled for the detectors and, using Monte Carlo simulation of both EASs and of the detector responses, we studied the reconstruction of the components for different distances to the shower axis, energies and zenith angles of the primaries. We assessed the reconstruction precision of the different components not only for the total signal, but also for its time structure.

T 51.5 Tue 17:15 T-H36

Pointing accuracy of the roving laser system for the energy calibration of the Pierre Auger Observatory* — •ALINA NASR ESFAHANI — Bergische Universität Wuppertal, Gaußstr. 20, Wuppertal, Germany

The Fluorescence Detector (FD) of the Pierre Auger Observatory provides a nearly model independent measurement of the energy of primary cosmic rays. This FD energy measurement sets the energy scale of the Surface Detector, its precision thereby factors into the systematic uncertainties of practically all scientific results from the Observatory.

By firing a laser with known energy output in front of the FD telescopes the energy calibration obtained from other methods can be cross-checked independently. The camera response to the laser closely resembles its response to a real cosmic ray shower providing a valuable end-to-end calibration. The laser system includes components to expand and depolarize the laser beam to optimize the scattering in the atmosphere and detection by the FD camera. We built the roving laser system by utilizing a telescope mount as a carrier for the laser. Angular precision measurements show that this greatly improves the pointing accuracy, which has been a significant source of uncertainty in previous campaigns. The precision requirements for a sufficient reduction of systematic uncertainties compared to previous systems are based on the analysis of laser simulations. *Supported by the BMBF Verbundforschung Astrotelichenphysik (Vorhaben

*Supported by the BMBF Verbundforschung Astroteilchenphysik (Vorhaben 05A20PX1).

T 51.6 Tue 17:30 T-H36

Applications of the high-energy lepton and photon propagator PRO-POSAL – •JEAN-MARCO ALAMEDDINE¹, JAN SOEDINGREKSO¹, and ALEXAN-DER SANDROCK² – ¹Astroparticle Physics WG Rhode, TU Dortmund University, Germany – ²University of Wuppertal, Germany

PROPOSAL is a simulation library, usable in both C++ and Python, which provides 3D Monte Carlo simulations of charged leptons and high-energy photons. One key concept of PROPOSAL is to offer a trade-off between simulation precision and performance for each individual use case. Due to its customizable and modular structure, PROPOSAL is used for a wide range of applications, for

example in the simulation chain of the IceCube Neutrino Observatory or as an electromagnetic interaction model in the shower simulation framework COR-SIKA 8.

In this contribution, an introduction to the simulation framework as well as an overview of its current and possible future applications, including muography, are presented.

T 51.7 Tue 17:45 T-H36 Effects of unresolved particles in the counts estimated by a segmented detector — •FLAVIA GESUALDI^{1,2} and DANIEL SUPANITSKY¹ — ¹Instituto de Tecnologías en Detección y Astropartículas (CNEA, CONICET, UNSAM), Centro Atómico Constituyentes, B1650KNA San Martín, Buenos Aires, Argentina — ²Karlsruhe Institute of Technology, Institute for Astroparticle Physics (IAP), 76021 Karlsruhe, Germany

Segmented particle counters are part of many astroparticle physics detectors and are used to estimate particle densities. For instance, measuring the density of muons in air showers is key for composition analyses, which in turn help to elucidate the origin of cosmic rays. Technically, the goal of a segmented particle counter is to provide an accurate estimate of the impinging number of particles from the measured number of hit segments. If two particles hit a same segment within a time interval smaller than the time resolution, the two particles are counted as one. This undercounting effect, referred to as pile-up, is larger when the number of segments is small with respect to the number of impinging particles, and when the time resolution is poor compared to the characteristic duration of a single-particle signal. In this work, we develop a new pile-up-correction

method that makes use of the whole temporal structure of the signal. We compare its performance against methods in literature. We show that the method of this work performs well when considering typical air-shower signals, and that it is also the only one that extends well to long or double-bump-like signals.

Group Report

T 51.8 Tue 18:00 T-H36 E.S.S. telescopes — •NAOMI VO-

Intensity interferometry campaign at the H.E.S.S. telescopes — •NAOMI VOGEL, ANDREAS ZMIJA, GISELA ANTON, STEFAN FUNK, DMITRY MALYSHEV, THILO MICHEL, FREDERIK WOHLLEBEN, and ADRIAN ZINK — Friedrich-Alexander-Universität Erlangen-Nürnberg, ECAP

Intensity interferometry (II) enables high angular resolution (~milliarcsecond) astronomical observations in the optical band by measuring the photon fluxes of at least two telescopes with varying baselines and correlating them. It has already been applied by VERITAS and MAGIC with excellent results. Imaging Atmospheric Cherenkov Telescopes are suitable for performing intensity correlations because of their very large collecting areas. We are planning an upcoming II campaign at the H.E.S.S. telescopes in Namibia. Our developed II setup, which includes a 2nm interference filter, is designed to fit on the lid of the telescopes and to handle the high photon count rates expected from the stars. This is achieved by photomultipliers whose photo currents are measured and then correlated. As preparation for our campaign lab measurements were carried out to achieve low background and good signal-to-noise ratios. In this contribution we will present our technical setup and results from our lab measurements as precursor for II at the H.E.S.S. telescopes.

T 52: Outreach Methods 2

Time: Tuesday 16:15-17:45

T 52.1 Tue 16:15 T-H37

Wie viele Farben hat ein Quark? Eine Messung mit Daten des Belle II Experimentes für die Teilchenphysik-Masterclasses — FLORIAN BERNLOCHNER, JO-CHEN DINGFELDER, •SVENJA GRANDERATH, HENRIK JUNKERKALEFELD, SEBAS-TIAN LÜLSDORF, FLORIN MARTIUS, BARBARA VALERIANI-KAMINSKI UND CHRIS-TIAN WESSEL für die Netzwerk Teilchenwelt-Kollaboration — Universität Bonn Bei den Teilchenphysik-Masterclasses des bundesweiten Projektes "Netzwerk Teilchenwelt" bekommen Jugendliche einen Einblick in die Grundlagen und Forschungsmethoden der Teilchenphysik sowie in die Arbeitswelt von Wissenschaftler:innen. 2021 wurde an der Universität Bonn von der Belle II-Arbeitsgruppe und von Lehramtsstudierenden eine neue Masterclass entwickelt, die auf einer Messung mit Daten des Experimentes in Japan basiert. Schüler:innen untersuchen dabei Bilder von Teilchenkollisionen, bei denen ein Teilchen-Antiteilchen-Paar entstanden ist, und lernen, die Quark-Antiquarkund Lepton-Antilepton-Paare anhand geeigneter Selektionskriterien zu klassifizieren. Anschließend bestimmen sie den experimentellen Wert des R-Wertes sowie die daraus resultierende Anzahl der Farbladungen und vergleichen dann diesen Wert mit der theoretischen Erwartung, die sie während der Masterclass selber berechnet haben. In dem Vortrag werden das didaktische Konzept und die Materialien der Masterclass sowie die Erfahrungen präsentiert, die bei Durchführung der Messungen sowohl online als auch in Präsenz gemacht wurden.

T 52.2 Tue 16:30 T-H37

Outreach Modules for a New Particle Search Using the ATLAS Forward Proton Detector and Higgs Boson Physics — IVAN DEMCHENKO, MARTIN KUPKA, ANDRÉ SOPCZAK, •ANTOINE VAUTERIN, and PETER ZACIK — CTU in Prague We present two modules as part of the Czech Particle Physics Project (CPPP). These are intended as learning tools in masterclasses aimed at high-school students (aged 15 to 18). The first module is dedicated to the detection of an Axion-Like-Particle (ALP) using the ATLAS Forward Proton (AFP) detector. The second module focuses on the reconstruction of the Higgs boson mass using the Higgs boson golden channel with four leptons in the final state. The modules can be accessed at the following link: http://cern.ch/cppp

T 52.3 Tue 16:45 T-H37

Ten years of the Outreach Project "International Cosmic Day" $- \cdot$ NORA FEIGL, CAROLIN SCHWERDT, and HEIKE PROKOPH for the Netzwerk Teilchenwelt-Collaboration — DESY Zeuthen

Since 2012 students become researchers in astroparticle physics for one day of the year in November: the International Cosmic Day (ICD). Participants from more than 60 institutions such as schools, universities and research institutes in 17 countries performed experiments, discussed results and learned from scientists about the latest research in their field on ICD 2021.

There are many ways to participate, such as carrying out cosmic particle experiments, analyzing publicly available cosmic ray data and much more. Communication between the groups is arranged via up to 10 video meetings that take place throughout the day.

Location: T-H37

The ICD allows the students to have a first-hand experience of working in science, shows the students how international collaborations work and how science functions as a connecting element across national borders, language barriers and cultural differences.

In this talk we will give insights into the organization of the day and present the different possibilities of participating in the International Cosmic Day - from small projects in high school physics classes to more time-intensive elaborate experiments.

T 52.4 Tue 17:00 T-H37

IceCube Masterclass - ein Onlinekonzept — •JANNES BROSTEAN-KAISER, MA-REN VOITZ-WIEDAU, CAROLIN SCHWERDT, LEANDER FISCHER und NORA FEIGL für die Netzwerk Teilchenwelt-Kollaboration — DESY Zeuthen Im Rahmen einer Bachelorarbeit an der HU Berlin und bei DESY, Standort

Zeuthen, wurde die IceCube Masterclass des Netzwerk Teilchenwelt von einer Präsenz- zu einer digitalen Veranstaltung umstrukturiert. Dabei wurden die einzelnen Veranstaltungsbereiche auf Kompatibilität mit dem Rahmenlehrplan Berlins geprüft, modular und mit Fokus auf den Kompetenzbereich Kommunizieren gestaltet. Diese Masterclass konzentriert sich inhaltlich auf drei Schwerpunkte. Zunächst wird das Standardmodell der Teilchenphysik vorgestellt, um das theoretische Fundament und die Relevanz des Detektors selbst zu geben. Anschließend werden die Funktionsweise des IceCube Detektors, sowie einfache Analysen und Identifikationsmöglichkeiten von Teilchen präsentiert. Und zuletzt werden noch am Beispiel der Punktquellensuche statistische Methoden in der Physik erklärt. Alle Bereiche werden von einem/einer Vermittler:in mit einem klassischen Vortrag gestartet, sowie mit Hands-On Übungen für die Schüler:innen begleitet. Der Ablauf der Einheiten orientiert sich am Basismodell 4 Begriffs- und Konzeptbildung nach Oser. Den Abschluss der Masterclass bildet ein Vortrag eines wissenschaftlichen Mitarbeitenden, welche:r den Detektor am Südpol besucht hat, sowie eine allgemeine Fragerunde zum wissenschaftlichen Arbeiten und dem Werdegang eines Forschenden.

T 52.5 Tue 17:15 T-H37

A Michelson interferometer as a demonstrator for gravitational wave detection in outreach activities — •DAVID KOKE and ALEXANDER KAPPES — WWU Münster, Münster, Deutschland

Gravitational waves are one of the most exciting phenomena in astrophysics and have given us new insights into our universe since their first direct detection in 2015. In order to easily demonstrate the basic principles of gravitational wave detection in outreach activities, a demonstration experiment based on a Michelson interferometer was created in the framework of a master thesis. The subject of this talk is the presentation of the current status of the project, with a special focus on the technical realization and the challenges involved.

T 52.6 Tue 17:30 T-H37

Escape Radon: Entwicklung eines digitalen Escape Rooms für den Physikunterricht — •HANNES NITSCHE — Technische Universität Dresden Digitale Spiele werden über die letzten Jahre vermehrt zu Lehrzwecken genutzt und sollen Lernkonzepte auf spielerische Art und Weise erweitern. Eine der außergewöhnlicheren Spielformen, die ihren Weg in die Bildung findet, ist die des digitalen Escape Rooms. Grundlage dieses Vortrags ist eine wissenschaftliche Arbeit, in der der didaktische Mehrwert dieses Spielformats für den Physikunterricht untersucht wurde. Dazu wurde eine digitale Escape Story entwickelt, welche sich inhaltlich mit der Radonbelastung in Deutschland auseinandersetzt

und dabei kernphysikalische Grundlagen vermittelt. Im Vortrag wird die Escape Story 'Escape Radon' sowie die Ergebnisse ihrer Erprobung und Evaluation vorgestellt. Des Weiteren wird erörtert, welche Gestaltungselemente von digitalen Escape Rooms das Interesse der Lernenden am Lehrinhalt fördern können und wieweit sich die Methode für Lehrzwecke adaptieren lässt.

T 53: Data Analysis, Information Technology and Artificial Intelligence 3

Time: Tuesday 16:15-18:30

T 53.1 Tue 16:15 T-H38

Improved selective background Monte Carlo simulation at Belle II with graph attention networks and weighted events — •BOYANG YU, NIKOLAI HARTMANN, and THOMAS KUHR — Ludwig-Maximilians-Universität München

When measuring rare processes at Belle II, a huge luminosity is required, which means a large number of simulations are necessary to determine signal efficiencies and background contributions. However, this process demands high computation costs while most of the simulated data, in particular in case of background, are discarded by the event selection. Thus filters using graph neural networks with attention mechanisms are introduced after the Monte Carlo event generation to save the resources for the detector simulation and reconstruction of events discarded at analysis level. Merely filtering out events will however inevitably introduce biases. Therefore statistical methods including sampling and reweighting are invested to deal with this side effect.

T 53.2 Tue 16:30 T-H38

Analysis Specific Filters for Selective Background Monte Carlo Simulations at Belle II — •LUCA SCHINNERL, BOYANG YU, NIKOLAI HARTMANN, and THOMAS KUHR — Ludwig Maximilians University Munich, Munich, Germany

The Belle II experiment is expected to accumulate a data sample of 50 ab-1 in its lifetime. For rare processes, strong background suppression is needed to precisely measure these types of events. Because of this, an extremely large number of simulated background events is necessary for an effective analysis. However, a significant portion of the simulated data is discarded trivially in the first stage of analysis, demanding a better method of simulation to keep up with the amount of data. For this purpose a neural network is implemented to select the relevant data after the Monte Carlo event generation and then only run the costly detector simulation and reconstruction for selected events. Existing methods have shown good success with graph neural networks. However, the total speedup of simulations is limited when considering generic selections with a retention rate of 4.25%. Here a maximum speedup of 2.1 was reached. In this work we iteratively introduce analysis specific filters to the training of the neural networks, which can greatly increase efficiencies. For the rare process B -> K*vv this methodology has been successful in significantly improving simulation speed.

T 53.3 Tue 16:45 T-H38

Preparing transformer-based dose predictions: Performance of encoder/decoder structures for CT- and dose-sequence encoding — \bullet Piet Hoffmann, Kevin Kröninger, Armin Lühr, Florian Mentzel, and Jens Weingarten — TU Dortmund

In radiotherapy, fast dose predictions based on CT images are useful as they reduce the need for computing-intensive Monte Carlo simulations and thus can speed up treatment planning. A new approach to these fast dose predictions consists of interpreting the CT to dose conversion as a sequence translation task and making use of a transformer machine learning model.

For this, the CT data is disected perpendicularly to the beam into a sequence of 2D slices, if not already aquired in this direction. Before these slices are fed into the translation architecture it is useful to first encode them to reduce their dimensionality and concentrate the contained information. After translation the data then has to be decoded into a dose prediction slice.

For the whole model to work properly, the structure of such encoder and decoder is important and thus in this talk different approaches are compared with respect to their performance. Properties of the resulting encoded data space, like smooth transitioning between data points and the density distribution of data points, and their potential benefits are discussed.

T 53.4 Tue 17:00 T-H38

Progressive Generative Adversarial Networks for High Energy Physics Calorimeter Simulations — •SIMON SCHNAKE^{1,2}, KERSTIN BORRAS^{1,2}, DIRK KRÜCKER¹, FLORIAN REHM^{2,3}, and SOFIA VALLECORSA³ — ¹DESY, Hamburg, Germany — ²RWTH Aachen, Germany — ³CERN openlab, Geneva, Swiss The simulation of particle showers in calorimeters is a computational demanding process. Deep generative models have been suggested to replace these computations. One of the complexities of this approach is the dimensionality of the data produced by high granularity calorimeters. One possible solution could

be progressively growing the GAN to handle this dimensionality. In this study,

electromagnetic showers of a (25x25x25) calorimeter in the energy range of 10

- 510 GeV are used to train generative adversarial networks. The resolution of the calorimeter data is increased while training. First results of this approach are shown.

T 53.5 Tue 17:15 T-H38

Location: T-H38

Generative Models For Hadron Shower Simulation — •ENGIN EREN — Deutsches Elektronen-Synchrotron, Notkestrasse 85, 22607 Hamburg, Germany Simulations provide the crucial link between theoretical descriptions and experimental observations in particle physics. It describes fundamental processes or the interactions of particles with detectors. The high computational cost associated with producing precise simulations in sufficient quantities, e.g. for the upcoming data-taking phase of the Large Hadron Collider (LHC) or future colliders, motivates research into more computationally efficient solutions. However, the simulation of realistic showers in a highly granular detector remains a hard problem due to a large number of cells, values spanning many orders of magnitude, and the overall sparsity of data.

This contribution advances the state of the art in two key directions: Firstly, we present a precise generative model for the fast simulation of hadronic showers in a highly granular hadronic calorimeter. Secondly, we compare the achieved simulation quality before and after interfacing with a so-called particle-flow-based reconstruction algorithm. Together, these bring generative models one step closer to practical applications.

T 53.6 Tue 17:30 T-H38

Deep Set Generation of Collider Events — •Екік Винмалл — Institut für Experimentalphysik, Universität Hamburg

With current and future high-energy collider experiments' vast data collecting capabilities comes an increasing demand for computationally efficient simulations. Generative machine learning models allow fast event generation, yet are largely constrained to fixed data and detector geometries.

We introduce a novel autoencoder setup for generation of permutation invariant point clouds with variable cardinality - a flexible data structure optimal for collider events. Our model is simple, lightweight and purely set based without exploiting additional graph structures. We show that our model scales well to large particle multiplicities and achieves good performance on various data sets.

T 53.7 Tue 17:45 T-H38

Angular Conditioning of Generative Models for Fast Calorimeter Shower Simulation — • PETER MCKEOWN — Deutsches Elektronen-Synchrotron DESY, Hamburg, Germany

Modern high energy physics experiments fundamentally rely on accurate simulation- both to characterise detectors and to bridge observed signals and underlying theory. Traditional simulation tools are reliant upon Monte Carlo methods which, while powerful, require significant computational resources, and are projected to become a major bottleneck at the high luminosity stage of the LHC and for future colliders. Calorimeter showers are particularly computationally intensive to simulate, due to a large number of particle interactions with the detector material.

A potential solution based on deep generative models promises to provide drastic reductions in compute times. Previous work in our group has demonstrated the ability of various generative models to accurately reproduce key physical properties of showers in highly granular calorimeters. While this work has focused on the specific case of a particle incident perpendicular to the face of the calorimeter, a practical simulator must be able to correctly simulate arbitrary angles of incidence. In this talk, efforts to add conditioning on the incident angle of the particle will be addressed.

T 53.8 Tue 18:00 T-H38

Refinement of jet simulation with generative adversarial networks — •Shruthi Janardhan^{1,2}, Sven Harder¹, Patrick Connor¹, Peter Schleper¹, Daniel Ruprecht², and Sebastian Götschel² — ¹Universität Hamburg — ²Technische Universität Hamburg

In High Energy Physics, the interaction of particles with matter at the detectors are best simulated with the GEANT4 software. Alternatively, less precise but faster simulations are sometimes preferred to reach higher statistical precision. We present recent progress of refinement of fast simulations with ML techniques to enhance the quality of such fast simulations. We demonstrate the

use of generative adversarial networks in the context of jet simulation using a Wasserstein loss function. The architecture consists of two opposing networks, Refiner and Critic. The Refiner, refines the distribution of the energy of the jets obtained with the fast simulation. The Critic is used to effectively differentiate between the distributions of refined energy and the distribution obtained by the GEANT4 simulation. The Refiner can be used solely to obtain a fast but refined jet simulation.

T 53.9 Tue 18:15 T-H38 Using ML to analytically model the CMS detector response to jets - •NILS GERBER, SAMUEL BEIN, and PETER SCHLEPER — Universität Hamburg

T 54: Invited Topical Talks 3

Time: Wednesday 11:00-12:40

Invited Topical Talk T 54.1 Wed 11:00 T-H15 Hunting XYZ Beasts at Belle and Belle II - •ELISABETTA PRENCIPE for the Belle II-Collaboration — Justus-Liebig-University of Giessen, Giessen, Germany

The search for conventional and non-conventional charmonium states has gained a lot of attention over the past twenty years. It has undoubtedly been shown since 2003 that there are more complex structures than mesons and baryons, and for several of those a non-unique interpretation has been provided, mainly due to the lack of statistics. We refer to these resonant states as X, Y, Z, depending on their properties. Recently the PDG has renamed those that are well established, trying to provide a better understanding. Undoubtedly B factories such as Belle and BaBar have in the past made a notable contribution in filling in the missing blocks of the charmonioum spectrum; however, the limited statistics did not allow to search further for these exotic XYZ states.

LHCb collected huge data sets, which allowed to conduct interesting analyses in the *i.e.* search for pentaquarks and the properties of the X(3872). An important contribution to the field will come from the Belle II data, once the planned integrated luminosity is reached, in particular in the search for exotic states in radiative decays, ISR and $\Upsilon(nS)$ transitions, which represent unique physics cases.

Recent spectroscopy results with the complete Belle data sets are discussed, and a summary of results in charmonium and bottomonium spectroscopy is provided with current Belle II data. Future plans with Belle + Belle II combined data sets are then presented.

Invited Topical Talk T 54.2 Wed 11:25 T-H15 Precision tests of the Standard Model using CP violation in B meson decays

•THIBAUD HUMAIR for the Belle II-Collaboration — Max Planck Institute for Physics, Föhringer Ring 6, 80805 Munich

One of the main goals of the Belle II physics program is to test the flavour sector of the Standard Model to very high precision by measuring the parameters of the CKM triangle, which governs quark mixing. This talk focuses on the measurements of two of these parameters: the angles alpha and beta.

The angle alpha requires to measure direct CP violation in various charmless, rare, modes. These modes require excellent performances in background reduction and in the reconstruction of neutrals. The angle beta is accessed through Many applications in particle physics require an accurate modelling of the energy response of detectors to individual particles as well as jets. For example, unfolding, fast detector simulation, as well as certain background estimation techniques, all require some input related to the jet response. While typical models are constructed from established functional forms such as Crystal Ball functions fit to data distributions of the response, an alternative approach is explored where a DNN classifier is employed in order to arrive at a model which takes into account correlated dependencies in the response on the true jet energy, pseudorapidity, jet flavour, and other factors.

Location: T-H15

time-dependent measurements of CP violation. These measurements require a very good B vertex resolution, excellent flavour tagging capabilities and a perfect understanding of the detector response. I will present recent Belle II results in both types of measurements, and discuss the future prospects for analyses of highest possible precision.

Invited Topical Talk T 54.3 Wed 11:50 T-H15 Back to the top: charting the bounds of the standard model — •AFIQ ANUAR - Deutsches Elektronen Synchrotron (DESY), Notkestraße 85, D-22607 Hamburg

The top quark, the most massive member of the standard model, is unique in that it is the only known fermion with a Yukawa coupling of order one. In addition, its short lifetime provides us with the only opportunity to study a quark prior to hadronization. These advantages make it among our best probes in searches for physics beyond the standard model. At the same time, the stream of negative results in searches of specific extensions of the standard model from the LHC makes the use of effective approaches increasingly attractive. In this talk, experimental analyses where such approaches are employed will be discussed, ranging from interpretations of precise standard model measurements to direct constraints on effective operators through the use of advanced statistical methods.

Invited Topical Talk

T 54.4 Wed 12:15 T-H15

Dark matter from spin-2 mediators — • STEFAN VOGL — University of Freiburg Dark matter interacting with massive spin-2 mediators is an intriguing possibility. However, due to the high energy behavior of longitudinal modes of spin-2 particles, the rate of DM annihilation into the mediators exhibits a tremendous growth as soon the channel is kinematically open. To have a consistent effective theory for the spin-2 particle, we analyze an extra-dimensional model such that the mediator(s) are the Kaluza-Klein (KK) modes of the 5D graviton. We find that including the full KK-tower in the computation reduces the annihilation rate by an order of magnitude or more. This casts some doubt on the universal applicability of previous studies with spin-2 mediators within an EFT framework and indicates that a careful consideration of UV physics is required to accurately capture the phenomenology.

T 55: Invited Topical Talks 4

Time: Wednesday 11:00–12:40

Invited Topical Talk

T 55.1 Wed 11:00 T-H16 Machine Learning for LHC Theory — • ANJA BUTTER — Institut für Theoretische Physik, Heidelberg, Germany

Over the next years, measurements at the LHC and the HL-LHC will provide us with a wealth of data. The best hope of answering fundamental questions like the nature of dark matter, is to adopt machine learning techniques for particle experiment and theory. LHC physics relies at a fundamental level on our ability to simulate events efficiently from first principles. In the coming LHC runs, these simulations will face unprecedented precision requirements to match the experimental accuracy. Neural networks can overcome limitations from the calculation of amplitudes and event generation. Generative networks can achieve high-precision in simulations while maintaining control over training stability and associated uncertainties. Since networks in the form of normalizing flows can be inverted, they also open new avenues in LHC analyses. The access to the density of the generated distribution enables new methods for anomaly detection, while their interpretation in terms of probability densities leads to new methods for multi-dimensional unfolding.

Invited Topical Talk

Towards high-precision deep learning for astroparticle physics

•CHRISTOPH WENIGER — University of Amsterdam, Netherlands Observational data relevant for astroparticle physics and astrophysical searches for dark matter becomes increasingly complex and detailed. We are in a situation where often what we can learn from new observations is limited not by the amount of data, but by the sophistication of our analysis tools and the quality and detail of our physical models. Classical statistical techniques, like Markov Chain Monte Carlo, severely limit model realism and complexity, due to their high simulation requirements and limitation on the number of free parameters. Neural simulation-based inference algorithms have the capability to break through these barriers in surprising ways. However, using these new classes of algorithms without compromising the precision and accuracy of statistical inference results remains challenging. I will present both successful examples and discuss typical pitfalls related to the application of neural simulation-based inference algorithms to dark matter searches with astrophysical data.

134

Location: T-H16

T 55.2 Wed 11:25 T-H16

Invited Topical Talk T 55.3 Wed 11:50 T-H16 The quest for the mechanism behind the matter-antimatter asymmetry —

•JULIA HARZ — Technische Universität München, München, Germany Our own existence is still a mystery, as some yet unknown mechanism had to generate an excess of matter over antimatter during the evolution of the Universe. After an introduction on why physics beyond the Standard Model is needed in order to explain the observed matter-antimatter asymmetry, I will give an overview of different theoretical mechanisms that are potentially able to explain such an asymmetry. Hereby, I will highlight interesting possible connections to neutrino physics and dark matter. Moreover, I will discuss the challenges of probing baryogenesis models and review promising experimental strategies.

Invited Topical TalkT 55.4Wed 12:15T-H16Towards the lightest dark matter in direct searches- • BELINA VON KROSIGK— Karlsruhe Institute of Technology, Institute for Astroparticle Physics,Eggenstein-Leopoldshafen, Germany

T 56: Flavour Physics 4

Time: Wednesday 16:15-18:30

T 56.1 Wed 16:15 T-H15

Bs mixing in scalar Leptoquark models — •JORDI FOLCH EGUREN¹, JAVIER VIRTO², and ANDREAS CRIVELLIN³ — ¹University of Barcelona/TU Dortmund — ²University of Barcelona — ³PSI

Leptoquarks provide viable solutions to the flavour anomalies, i.e. they can explain the tensions between the measurements and the Standard Model predictions of the anomalous magnetic moment of the muon as well as b-s and b- $c\tau v$ processes.

However, LQs also contribute to other flavour observables, such as F = 2 processes, at the loop-level. In particular, Bs mixing provides a crucial bound in setups addressing b- $c\tau\nu$ data, often excluding a big portion of the parameter space that could otherwise account for it.

In this work, we first derive the complete leading order matching, including all five scalar LQ representations, for D0, K0 and Bs mixing (at the dimension-six level). We then calculate the next-to-leading order α s matching corrections to these F = 2 processes in generic scalar leptoquark models.

We find that the two-loop corrections increase the effects in F = 2 processes by 5-10% and significantly reduce the matching scale uncertainty.

T 56.2 Wed 16:30 T-H15

Testing the Standard Model with CP-asymmetries in flavour-specific nonleptonic decays — TIM GERSHON¹, ALEXANDER LENZ², •ALEKSEY RUSOV², and NICOLA SKIDMORE³ — ¹Department of Physics, University of Warwick, Coventry, CV4 7AL, UK — ²Center for Particle Physics Siegen, Theoretische Teilchenphysik, Universität Siegen, Walter-Flex-Str. 3, 57068 Siegen, Germany — ³University of Manchester, Schuster Building, Manchester, M13 9PL, UK

Motivated by recent indications that the rates of colour-allowed non-leptonic channels are not in agreement with their Standard Model expectations based on QCD factorisation, we investigate the potential to study CP asymmetries with these decays. In the Standard Model, these flavour-specific decays are sensitive to CP violation in $B_{(s)}^0 - \bar{B}_{(s)}^0$ mixing, which is predicted with low uncertainties and can be measured precisely with semileptonic decays. If there are beyond Standard Model contributions to the non-leptonic decay amplitudes, there could be significant enhancements to the CP asymmetries. Measurements of these quantities therefore have potential to identify BSM effects without relying on Standard Model predictions that might be affected by hadronic effects. We discuss the experimental prospects, and note the excellent potential for a precise determination of the CP asymmetry in $\bar{B}_s \rightarrow D_s^+ \pi^-$ decays by the LHCb experiment.

T 56.3 Wed 16:45 T-H15

Endpoint divergences in QED corrections to $B_s \rightarrow \mu^+ \mu^- - \bullet$ Nicolas Seitz, Thorsten Feldmann, Tobias Huber, and Nico Gubernari — Center for Particle Physics Siegen, Theoretische Teilchenphysik, Universität Siegen

We consider leptonic B-meson decays of the form $B_s \rightarrow \mu^+ \mu^-$. These are mediated in the standard model by operators of the effective weak Hamiltonian for $|\Delta B| = |\Delta S| = 1$ transitions. At leading order, only the semi-leptonic operator O_{10} contributes. If one calculates the QED corrections mediated by the operator O_7 , the suppression by the fine-structure constant α is amplified by a factor $1/\lambda^2 = m_b/\Lambda_{\rm QCD} \gg 1$ Here, one additionally obtains a quadratic logarithmic amplification $\propto \ln^2 \lambda^2$, which comes from endpoint divergences in regions of small momenta. The aim of our project is to investigate the interaction of QCD corrections on the hadronic side and the endpoint divergences of the muon propagator. An important technical tool here is the method of regions in the calculation of the loop integrals that occur. In the last decades, astronomical observations have consistently indicated that most of the matter in the Universe remains hidden to even the most sensitive telescopes because it is nonluminous - because it is dark. Observing the respective dark matter particles became one of the most tantalizing endeavors of modern physics. A new generation of large exposure direct search experiments is at the ready to observe weak-scale dark matter particles, with their successors already in the planning. At the same time a new era has begun towards a direct detection of ever lighter dark matter candidates. Novel detector designs are reaching ultra-low detection thresholds with which new detection channels can be exploited and unprecedentedly low dark matter masses can be probed. Stateof-the-art direct detection searches most sensitive to light dark matter will be reviewed together with an outlook on where the near future is expected to take us in this quest towards dark matter discovery in the laboratory.

Location: T-H15

T 56.4 Wed 17:00 T-H15

BSM effects in lifetimes of *B* **mesons** — •JAKOB MÜLLER, ALEXANDER LENZ, MARIA LAURA PISCOPO, and ALEKSEY RUSOV — Center for Particle Physics Siegen, Theoretische Teilchenphysik, Universität Siegen

We study the impact of potential BSM contributions to non-leptonic *b* quark decays on observables like $\tau(B^+)/\tau(B_d)$ and $\tau(B_s)/\tau(B_d)$. These observables are measured with a precision of the order of several per mille. The corresponding theory predictions are obtained within the framework of the Heavy Quark Expansion.

T 56.5 Wed 17:15 T-H15

NLO QCD corrections to inclusive $b \rightarrow c \ell \bar{v}$ decay spectra up to $1/m_b^3$ — Thomas Mannel, •Daniel Moreno, and Alexei A. Pivovarov — Center for Particle Physics Siegen, Theoretische Physik 1, Universität Siegen, 57068 Siegen, Germany

We present analytical results for higher order corrections to the decay spectra of inclusive semileptonic heavy hadron weak decays, using the heavy quark expansion (HQE). We describe the analytical computation of the spectrum of the leptonic invariant mass for $B \rightarrow X_c \ell \bar{\nu}$ up to terms of order $1/m_b^3$ within the HQE at next-to-leading order (NLO) in α_s . The full dependence of the differential rate on the mass of the final-state quark is taken into account. We discuss the implications of our results for the precision determination of the CKM matrix element $|V_{cb}|$.

T 56.6 Wed 17:30 T-H15

Towards completion of the four-body contributions to $\bar{B} \rightarrow X_s \gamma$ at NLO — TOBIAS HUBER and •LARS-THORBEN MOOS — Center for Particle Physics Siegen, Theoretische Teilchenphysik, Universität Siegen

The inlusive radiative $\overline{B} \to X_s \gamma$ decay constitutes an important pillar in the indirect search for new physics and allows to constrain the parameter space of many models.

In this talk we present the ongoing efforts in the computation of four-body contributions to the process $\overline{B} \to X_s y$, namely those of $b \to sggy$ at NLO in the strong coupling and the necessary complementing 5-particle cuts of the gluon-bremsstrahlung $b \to sggy + g$.

Although these corrections are expected to be small, this computation formally completes the NLO contributions to $\bar{B} \rightarrow X_s \gamma$.

Since the anomalous dimensions are already computed to a sufficient order, the main tasks are the systematic generation of the 1-loop amplitude, the automation of the phase space integration, the infrared-regularization and finally the renormalization of the diagrams including the operator mixing.

The results obtained so far are shown and the further structure of the calculation is outlined.

T 56.7 Wed 17:45 T-H15

Improved theory determination of $|V_{ub}|$ **from inclusive B-decays** — •KEVIN OLSCHEWSKY — Center for Particle Physics (CPPS), Siegen University

Inclusive B-meson decays into light final state particles like the semileptonic $\bar{B} \rightarrow X_u \ell \bar{v}$ are of great importance for the precise determination of the Cabbibo-Kobayashi-Maskawa (CKM) matrix element $|V_{ub}|$. With the unprecedented amount of experimental data, it becomes more and more important to have sound theoretical predictions with small and controllable uncertainties.

In order to obtain the $\bar{B} \to X_u \ell \bar{\nu}$ decay rate, the much larger $b \to c$ background has to be removed by appropriate kinematical cuts. Theoretically the calculation of partial decay rates in this region of phase space where $\bar{B} \to X_c \ell \bar{\nu}$ decays are suppressed requires the introduction of a non-perturbative distribution function; the "shape function" (SF).

In this talk I will present an update for the BLNP framework, which is based on the soft-collinear effective field theory framework. This includes updates for all known perturbative quantities as well as new parameterizations for the SF. Our systematic approach in modelling the SF allows us to provide a sound error analysis based on even higher orders in the Heavy Quark Expansion than before.

T 56.8 Wed 18:00 T-H15

Dispersive bounds for local form factors of $\Lambda_b \rightarrow \Lambda$ — Thomas Blake¹, Ste-FAN MEINEL², •MUSLEM RAHIMI³, and DANNY VAN DYK⁴ – ¹Department of Physics, University of Warwick, UK - ²Department of Physics, University of Arizona, USA — ³Center for Particle Physics Siegen, Theoretische Physik 1, Universität Siegen, Germany — ⁴Physik Department T31, Technische Universität München, Germany

We investigate the 10 form factors relevant to the b-baryon decay $\Lambda_b \to \Lambda \ell^+ \ell^$ by combining information of Lattice QCD and dispersive bounds. To this end, we use a parametrization of the local form factors in terms of orthonormal polynomials with respect to the dispersive integral kernel. Our approach provides

T 57: Flavour Physics 5

Time: Wednesday 16:15–18:15

T 57.1 Wed 16:15 T-H16

Search for $B^0_{(s)} \to p\bar{p}\mu^+\mu^-$ decays with the LHCb experiment — JOHANNES ALBRECHT, •MAIK BECKER, LUKAS CALEFICE, and VITALII LISOVSKYI — Experimentelle Physik 5, TU Dortmund

In 2019 the LHCb collaboration reported the first observation of the decays $B_{(s)}^0 \rightarrow J/\psi p\bar{p}$. The branching fraction of the B_s^0 mode was measured to be $(3.6 \pm 0.4) \times 10^{-6}$, which was much larger than the theoretically expected value of $\mathcal{O}(10^{-9})$ at that time. For the B^0 mode, however, the branching fraction was in agreement with theoretical predictions.

The question arises whether the corresponding non-resonant decays are also observable with the full data set of 9 fb⁻¹ collected by the LHCb experiment. For the B_s^0 mode the leading-order Feynman diagram is similar to the one for the $B_s^0 \rightarrow \mu^+ \mu^-$ decay, but includes an additional $p\bar{p}$ pair from gluon radiation, lifting the helicity suppression. For the B^0 mode Cabibbo-suppressed $b \rightarrow d\mu^+\mu^$ transitions dominate.

In this talk an ongoing analysis of $B^0_{(s)} \rightarrow p\bar{p}\mu^+\mu^-$ decays using data from the LHCb experiment will be presented. In particular, the selection and studies on the resonant control channels are shown. The search aims at intensifying the efforts of the LHCb collaboration to study rare decays with leptons and baryons in the final state.

T 57.2 Wed 16:30 T-H16 Test of lepton universality with $\Lambda_b \rightarrow pKl^+l^-$ decays at LHCb — Johannes Albrecht, Vitalii Lisovskyi, and •Jannis Speer — Experimentelle Physik 5, TU Dortmund

In recent measurements of *b*-hadron decays a pattern of consistent tensions with the SM predictions is observed. This includes decays with $b \rightarrow sl^+l^-$ transitions, which play an important role in lepton flavor universality tests such as R_K and $R_{K^{*0}}$. Complementary to *b*-meson decays, lepton flavor universality can also be tested in b-baryon decays which come with partly orthogonal experimental uncertainties. The first measurement of the ratio of branching fractions of the decays $\Lambda_b \to pKe^+e^-$ and $\Lambda_b \to pK\mu^+\mu^-$, R_{pK}^{-1} , was published by the LHCb Collaboration using proton-proton collision data corresponding to 4.7 fb⁻¹. In the dilepton mass-squared range $0.1 < q^2 < 6.0 \text{ GeV}^2/c^4$ and the *pK* mass range $m(pK) < 2600 \text{ MeV}/c^2$ the ratio of branching fractions was measured to be $R_{pK}^{-1} = 1.17^{+0.18}_{-0.16} \pm 0.077$. The legacy measurement of R_{pK}^{-1} tries to reduce the uncertainties by analysing the full 9 fb⁻¹ dataset of LHCb experiment and improving the selection.

In this talk the first study of the data recorded in the years 2017 and 2018 is presented. Furthermore the ongoing improvements in the signal selection requirements are discussed.

T 57.3 Wed 16:45 T-H16

Isospin asymmetries in rare B decays — JOHANNES ALBRECHT, •FABIO DE VEL-LIS, and VITALII LISOVSKYI - Experimentelle Physik 5, TU Dortmund

Isospin symmetry is a fundamental property of the Standard Model. It predicts a branching fraction that is almost the same for decays which differ only by one spectator quark, like $B^0 \to K^0 \mu^+ \mu^-$ and $B^+ \to K^+ \mu^+ \mu^-$. The same is true for the decays $B^0 \to K^{*0} \mu^+ \mu^-$ and $B^+ \to K^{**} \mu^+ \mu^-$. For these decays a quantity which describes differences in branching fraction, namely the asymmetry, can be defined. This is particularly convenient since it is theoretically clean and it allows to cancel some experimental systematics. Previous measurements on these decays from LHCb and Belle, despite being compatible with expectations, suggested coherent deviations that could be interpreted as statistical fluctuations, or unaccounted theoretical uncertainties, or as a sign of New Physics. In this

control over the form factor uncertainties due to truncation of the series expansion and extrapolation to the region of low momentum transfer, which is of great phenomenological interest.

T 56.9 Wed 18:15 T-H15

B-meson decay into a proton and dark antibaryon from QCD light-cone sum $rules - {\it Alexander Khodjamirian and \bullet Marcel Wald} - {\it Center for Particle}$ Physics Siegen, Theoretische Teilchenphysik, Universität Siegen

Recently, a B-Mesogenesis scenario was suggested to simultaneously solve the baryon asymmetry and relic dark matter abundance problems. In this scenario, decays of B-mesons into a baryon and dark antibaryon in the final state are expected with an appreciable branching fraction within the reach of modern Bfactories. We suggest to apply QCD light-cone sum rules to the decay mode $B^+ \to p\Psi$, where Ψ is a dark antibaryon. With this method we obtain the $B \to p$ hadronic matrix element of the three-quark effective operator in terms of the nucleon light-cone distribution amplitudes and estimate the partial width.

Location: T-H16

talk an update of the asymmetry measurement with the full LHCb dataset is presented. This means that data corresponding to an integrated luminosity of 6 fb⁻¹ are added to the dataset used in previous Run 1 analysis. This analysis also aims to give an update to the differential branching fraction measurement of the above-mentioned decays.

T 57.4 Wed 17:00 T-H16 Measurement of the ratio $R_{K\pi\pi}$ with the LHCb experiment – CHRISTOPH LANGENBRUCH, •JOHANNES HEUEL, and STEFAN SCHAEL — I. Physikalisches Institut B, RWTH Aachen University

In the Standard Model (SM) of particle physics, the coupling of electroweak gauge bosons to all leptons is universal. Stringent tests of this Lepton Flavour Universality (LFU) are possible by measuring ratios of rare $b \rightarrow s\ell\ell$ decays with different leptons in the final state. These decays are loop-suppressed in the SM and therefore sensitive to new heavy particles beyond the SM.

The LHCb experiment is ideally suited for the study of rare b hadron decays due to its large acceptance, the high trigger efficiencies and the excellent tracking and particle identification. Recent measurements of $b \rightarrow s\ell\ell$ ratios published by the LHCb Collaboration show tensions with the SM predictions of up to 3.1 standard deviations. Therefore, further studies of LFU tests using other rare B decay channels are crucial.

The current status of the ongoing measurement of the ratio $R_{K\pi\pi}$ of the branching fractions of the decays $B^+ \to K^+ \pi^- \pi^- \mu^+ \mu^-$ and $B^+ \to K^+ \pi^+ \pi^- e^+ e^$ is presented. The measurement is experimentally challenging as the hadronic system is measured inclusively.

T 57.5 Wed 17:15 T-H16

Probing multilepton decays with the LHCb experiment - JOHANNES AL-BRECHT and •VITALII LISOVSKYI - Experimentelle Physik 5, TU Dortmund In the recent years, a number of tensions has been observed in rare decays of B hadrons to a lighter hadron and two leptons. With the large dataset collected by the LHCb experiment, it becomes possible to study even higher-order processes. For instance, in the Standard Model, radiation of a virtual photon from the initial state or the final state can create an additional dilepton pair, leading to a final state with four leptons. In theories beyond the Standard Model, there are alternative mechanisms to reach such final state, which makes such decays excellent probes in searches for New Physics. In this talk, decays of beauty hadrons and quarkonia to final states with four leptons will be discussed. In particular, a search for the decay $B^+ \rightarrow K^+ \mu^+ \mu^- \mu^+ \mu^-$ with the dataset collected by the LHCb experiment will be presented. I will discuss the experimental challenges and sources of background, as well as estimate the expected sensitivity.

T 57.6 Wed 17:30 T-H16

Search for the $B^0 \rightarrow D^0 \bar{D}^0$ decay with the LHCb experiment. — •JONAH BLANK and SOPHIE HOLLIT — TU Dortmund

With precise measurements of B meson decays the LHCb experiment can test the integrity of the Standard Model of particle physics. Especially $B \rightarrow D D$ decays are interesting to examine CP violation and further constrain the unitarity triangle. While decays to charged D^{\pm} mesons have already been well measured, the $B^0 \rightarrow D^0 \overline{D}^0$ decay channel has not yet been observed by any experiment. In this analysis, data collected by the LHCb experiment at $\sqrt{s} = 7,8$ TeV and 13 TeV corresponding to an integrated luminosity of 9 fb^{-1} is used to search for the $B^0 \to D^0 \bar{D}^0$ decay channel. The $B^0 \to \bar{D}^0 \pi^+ \pi^-$ decay channel is utilized as a normalisation mode to cancel most uncertainties. An update of the current status of the analysis will be presented.

T 57.8 Wed 18:00 T-H16

T 57.7 Wed 17:45 T-H16

Measurements of strangeness production with the upgraded LHCb detector — •Lukas Calefice^{1,2}, VITALII LISOVSKYI¹, JOHANNES ALBRECHT¹, and VLADIMIR GLIGOROV² — ¹Experimentelle Physik 5, TU Dortmund — ²CNRS/LPNHE, Sorbonne Université, Paris

The LHCb experiment is currently undergoing a major upgrade of its detector to enable running at a five times higher instantaneous luminosity with respect to the previous data taking. Among other things the upgrade comprises the removal of the hardware trigger stage, a complete re-design of the software trigger, replacing the front-end readout electronics of all sub-detectors and an entire new set of tracking detectors. Validating the performance and data quality of the newly configured detector is a crucial task for the beginning of the next data taking period.

Due to their very large production cross-sections at the LHC strange hadrons such as Λ^0 and K_S^0 can be studied with only few days of data taking. Therefore, these are used to investigate the alignment of tracking detectors, to check the PID performace and validate the simulation of the upgraded detector. Finally, a measurement of the strangeness production cross-sections at 13 and 13.5 TeV will be performed with the early data after restarting the LHC.

This talk focuses on the preparations of the detector validation with data from the previous data takings.

Search for ³He ions at LHCb — •HENDRIK JAGE, GEDIMINAS SARPIS, VALERY ZHUKOV, and STEFAN SCHAEL — I. Physikalisches Institut B, RWTH Aachen University

In recent presentations, AMS-02 has reported the observation of several antihelium candidates in cosmic rays. In 2020, it has been suggested by M. Winkler and T. Linden that dark matter annihilation into *b*-quarks could produce a detectable ³He flux in cosmic rays via $\overline{\Lambda}_b^0$ decays.

The LHCb detector at CERN is an experiment dedicated to the study of *b*-hadrons, which are abundantly produced in the proton-proton collisions at the Large Hadron Collider (LHC). Therefore, the large sample of Λ_b^0 decays, collected by LHCb until 2018, provides a unique opportunity to study the potential displaced production of ³He via Λ_b^0 decays.

While prompt ³He from proton-proton collisions has already been observed at the LHC by the ALICE Collaboration in the central region (|y| < 0.5), prompt and displaced ³He has not yet been searched for at LHCb ($2 < \eta < 5$). In this talk, the possibility of identifying ³He at LHCb is discussed and the status of the on-going analysis is presented.

T 58: QCD (Exp.) 2

Time: Wednesday 16:15-18:00

T 58.1 Wed 16:15 T-H17

Measurements of the total charm and beauty cross sections with the CMS detector — \bullet Josry Metwally, Achim Geiser, Nur Zulaiha Jomhari, and Yewon Yang — DESY, Hamburg, Germany

The aim of this project is the determination of the total cross section for inclusive charm and beauty production at the LHC with different center-of-mass energies down to very low transverse momentum, and the comparison with QCD predictions in next-to-next-leading order of perturbation theory. The measurement of the cross sections for the production of heavy quarks at the LHC is one important test of QCD, and can, as has already happened in the case of top quark production, be used for a measurement of the quark masses.

Other experiments as ATLAS and ALICE covered only small fractions of the available phase space while the LHCb experiment fully covered the forward region, 2.0 < y < 4.5. For this project, we measure cross sections in the full phase space complementary to LHCb of prompt D mesons, and D mesons from b hadron decays through the decays $B \rightarrow D^*X \rightarrow D^0\pi_s X \rightarrow K\pi\pi_s X$ and $B \rightarrow D^0 X \rightarrow K\pi X$. One of the challenges is the separation of prompt D mesons and D mesons from b hadron decays near the production threshold. In this talk, the details of this separation and resulting cross sections including a comparison with theory are presented in the accessible phase space of CMS for different center-of-mass energies and, where it can be performed, a comparison with other experiments is shown.

T 58.2 Wed 16:30 T-H17

Low mass Drell-Yan measurement in p-p collision at $\sqrt{s} = 13$ TeV using the ATLAS detector at the LHC — •ALESSANDRO GUIDA — DESY (Hamburg) High energy physics experiments are performed at the Large Hadron Collider at CERN colliding bunches of protons at energies up to 13 TeV. The ATLAS experiment, with its multipurpose detector, studies the products of these collisions and compares the experimental measurements with the predictions of the Standard Model. This talk presents the study of the process $Z/\gamma^* \rightarrow \mu\mu$ at low invariant mass of the di-muon pair, in the region between 7 GeV and 60 GeV, below the the Z boson resonance mass peak ($m_Z = 91.2 \text{ GeV}$). The single and double differential cross sections $d\sigma/dm_{\mu\mu}$, $d^2\sigma/dm_{\mu\mu} d|y_{\mu\mu}|$ and $d^2\sigma/dm_{\mu\mu} dp_T^{Z/\gamma^*}$ of the process are measured in 13 TeV proton-proton collisions at the LHC, using the ATLAS detector. The measurement explores an extreme region of the phase space and is sensitive to resummation results in the theoretical prediction. The analysis exploits the good resolution of the ATLAS detector in reconstructing low momentum muons. The main difficulties come instead from the high background component that enters in the event selection, the triggering of events and the modelling of some key physical quantities.

The main features of the analysis, the studies done to overcome the main challenges, as well as the first results and comparison to theory predictions are presented in the talk.

T 58.3 Wed 16:45 T-H17

Studies on Monte Carlo tuning using Bayesian Analysis — •Salvatore La Cagnina¹, Andrii Verbytski², Johannes Erdmann¹, Kevin Kröninger¹, and Stefan Kluth² — ¹TU Dortmund, Fakultät Physik — ²Max-Plank-Institut für Physik, München

Monte Carlo (MC) simulations are an essential aspect of data analysis at the LHC. One aspect of MC event generation involves hadronisation and parton shower models. Since these models are based on approximations, they introduce a number of parameters. These parameters cannot be inferred from first principles. Therefore, their values have to be optimized using numerical tools and experimental data (MC tuning). Generally, MC tuning is performed by choosing observables that are sensitive to the parameters. Afterwards, a fit of the parameters to data using a simplified MC response function derived from fits to MC events is performed. Though state-of-the-art methods for MC tuning exist, uncertainties are usually treated as uncorrelated. In this talk, MC tuning using a Bayesian approach will be discussed. The EFTfitter tool is used for fitting, which enables the implementation of correlations for different sources of uncertainties. First results using this method on a MC tune with LEP data will be presented.

T 58.4 Wed 17:00 T-H17

Location: T-H17

LHCb for astroparticle physics: Prompt production of charged particles – Johannes Albrecht¹, •Julian Boelhauve¹, Hans Dembinski¹, and Michael Schmelling² – ¹TU Dortmund University, Dortmund, Germany – ²Max Planck Institute for Nuclear Physics, Heidelberg, Germany

A long-standing issue in the field of cosmic-ray research is the discrepancy in the number of muons produced in high-energy air showers between observations and simulation, which is referred to as the Muon Puzzle. Precision measurements of hadron production in the forward region are required in order to validate and improve the hadronic-interaction models used in the simulation of air showers, aiming at solving the Muon Puzzle. For this, measuring the differential cross-section of prompt production of long-lived charged particles as a function of transverse momentum and pseudorapidity is of great importance.

An analysis in which this differential cross-section has recently been determined in proton-proton collisions recorded with the LHCb experiment at a centre-of-mass energy of 13 TeV is presented in this talk. Moreover, extensions of the analysis towards a measurement of prompt production of identified hadrons are described.

T 58.5 Wed 17:15 T-H17

Potential of Common Data-Taking of the ATLAS, AFP, ZDC and LHCf Detectors in Run 3 of the LHC — •YUSUF CAN ÇEKMECELIOĞLU, CLARA ELISABETH LEITGEB, and ÇIĞDEM IŞSEVER — DESY, Zeuthen, Germany

Studies of air showers induced by highly energetic cosmic particles depend heavily on models for the soft hadronic interactions. Perturbative QCD cannot be applied to these interactions due to the low momentum exchange between particles. Instead, phenomenological models that take inputs from the (ultra-)forward regions of collider experiments are used to better understand these processes. The LHC with a collision energy of $\sqrt{s} = 13.6$ TeV in run 3 can generate such events and provide data to reduce the large uncertainties for hadronic models.

This talk will target the potential of a common data-taking of several forward detectors (so far used independently) that are located at both sides of the ATLAS detector, namely: The ATLAS Forward Proton detector (AFP), the ATLAS Zero Degree Calorimeters (ZDC), and the LHC forward (LHCf) calorimeters. The analysis focuses on the determination and optimisation of the common acceptance between detectors for simulated single diffractive (SD) events at (preliminary) run 3 beam conditions. SD events allow AFP to tag the intact proton and

Wednesday

the LHCf and ZDC calorimeters to detect the neutral particles from the dissociated proton. Together with pseudorapidity gap measurements in the central region provided by the ATLAS detector, a joint data-taking between these detectors could improve the identification and kinematic reconstruction of such events.

T 58.6 Wed 17:30 T-H17

Study of the X(3915) at Belle — •YAROSLAV KULII¹, THOMAS KUHR¹, and BORIS GRUBE² — ¹Ludwig-Maximilians-Universität München — ²Technische Universität München

Charmonium states consist of a charm and anti-charm quark. Detailed theoretical predictions of the charmonium excitation spectrum agree well with the experimental data.

However, in recent years experiments discovered a growing number of charmonium-like states that do not fit into the predicted charm-anticharm excitation spectrum. One such state is the X(3915). It has been discovered by the BaBaR and Belle collaborations in the reaction $e^+e^- \rightarrow e^+e^-X(3915) \rightarrow e^+e^-J/\psi\omega$, where the final-state electron and positron are not detected. The analysis of projections of the decay angular distribution preferred the $J^{PC} = 0^{++}$ hypothesis, but other quantum numbers, in particular $J^{PC} = 2^{++}$, could not be excluded.

Because of this the *X*(3915) was initially identified as the $\chi_{c0}(2P)$ charmonium

state, although its mass and decay width were not in good agreement with the theory predictions. Following the Belle discovery of the X^* (3860), which agrees much better with the $\chi_{c0}(2P)$ hypothesis, opinions shifted towards interpreting the X(3915) as an exotic state. It could be, for example, a meson molecule or a so-called hybrid meson.

We will present the current state of measuring of the spin and parity of the X(3915) at Belle and discuss the prospects of studying the X(3915) using the Belle II data.

T 58.7 Wed 17:45 T-H17

Location: T-H18

Partial wave analysis of the $\tau \rightarrow 3\pi v_{\tau}$ **decay at Belle** — •ANDREI RABUSOV, DANIEL GREENWALD, and STEPHAN PAUL — TUM, Munich, Germany

The COMPASS collaboration observed a potential new particle, the $a_1(1420)$, that doesn't fit the quark model. An independent study of the existence of this particle, as well as the studies of the light axial and pseudoscalar resonances, can be done in the tauon decay to three pions and a tau neutrino. The latest such study was published by the CLEO II collaboration in 1999 by analyzing 51000 data events. That study can be significantly improved at B-factories, which collected tens of millions events of this decay. We present data selection criteria, acceptance studies, and partial wave analysis of the $\tau \rightarrow 3\pi v_{\tau}$ decay with the Belle detector.

T 59: Neutrino Physics with Accelerators 1

Time: Wednesday 16:15-17:50

T 59.1 Wed 16:15 T-H18

Particle Identification and Reconstruction with the DUNE ND-GAr Near Detector — •LORENZ EMBERGER and FRANK SIMON — Max-Planck-Institut für Physik

The Near Detector (ND) of the Deep Underground Neutrino Experiment (DUNE) will play an important role in the search of CP violation in the neutrino sector. Additionally, as a standalone complex, it will be an excellent laboratory to study a wide range of neutrino interactions and BSM models. The ND design consists of three independent sub-detectors, placed downstream of the neutrino production target. One of these detectors, called ND-GAr, consists of a magnetized high pressure gaseous Argon Time Projection Chamber (TPC), surrounded by an electromagnetic calorimeter (ECAL) and a magnet yoke. One key aspect of the ECAL is the reconstruction of neutral particles such as neutral pions and potentially neutrons. The ECAL also extends the detector's separation capability of muons and pions, which is further enhanced by a muon tagger in the magnet yoke. We present a simulation study of the detector system featuring a highly granular electromagnetic calorimeter inspired by the SiPM-on-Tile technology developed by the CALICE collaboration. We will introduce the detector design considerations, as well as the potential physics program. Furthermore, we will discuss the separation of muons and pions using the ECAL and study the impact of different possible muon tagger layouts. A simulation study on time-of-flight reconstruction of the kinetic energy of neutrons will also be presented.

T 59.2 Wed 16:30 T-H18

Studies on the DUNE ND-GAR ECAL Design — •SEBASTIAN RITTER¹, PE-TER BERNHARD², ANDREA BROGNA², VOLKER BÜSCHER¹, KARL-HEINZ GEIB¹, ASMA HADEF¹, ANTOINE LAUDRAIN¹, LUCIA MASETTI¹, MARISOL ROBLES MANZANO¹, ANNA ROSMANITZ¹, CHRISTIAN SCHMITT¹, ALFONS WEBER¹, and QUIRIN WEITZEL² — ¹Johannes-Gutenberg Universität Mainz — ²PRISMA+ Detector Lab

The Deep Underground Neutrino Experiment (DUNE) aims to unlock the mystery of neutrinos. One of the major goals is to measure the CP-violating phase of the neutrino mixing matrix for which the DUNE near detector (ND) is crucial. A leading role in measuring neutrino interactions in the ND will be filled by the high-pressure gaseous argon TPC. A sampling ECAL based on plastic scintillators with SiPM readout is surrounding the TPC (ND-GAr). In this talk, optimized geometries are considered for the ND ECAL motivated by external boundary conditions and the beam-on-target nature of DUNE. A second focus will be on the readout of the ECAL's scintillator strips trying to efficiently use the available space and optimizing the light output.

T 59.3 Wed 16:45 T-H18 **Plastic Scintillator and Light Guide Research and Development** — •PATRICK DEUCHER — Johannes Gutenberg Universität Mainz

Plastic scintillators are broadly used in physics experiments for the detection of particles and electromagnetic radiation. With tunable emissive properties and a fluorescent decay time of a few ns, plastic scintillators are a solid option that can be tailored to individual applications. First steps for the production and optimization of polystyrene based plastic scintillators have been taken at the Johannes Gutenberg University in Mainz. This includes the purification of styrene, the addition of different fluorophores and thermal polymerization. In cooperation with Tübingen, we develop dedicated active light guides for use with large SiPM arrays. Moreover, plastic scintillators with optimized capability for pulse-shape discrimination are investigated for use in the ECAL of the DUNE Near Detector. This talk will present the progress on plastic scintillator production and characterization including absorption, emission and lifetime measurements of first samples. This work is supported by funds of the Excellence Cluster PRISMA+.

Group ReportT 59.4Wed 17:00T-H18ANNIE: The Accelerator Neutron Neutron Interaction Experiment-•MARC BREISCH for the ANNIE-Collaboration — Physikalisches Institut, Eberhard Karls Universität Tübingen

The Accelerator Neutrino Neutron Interaction Experiment (ANNIE) is a 26ton gadolinium doped water Cherenkov detector situated on-axis of the Booster Neutrino Beam (BNB) at FermiLab. Its main goal is to measure the final state neutron multiplicity of neutrino-nucleus interactions to improve the systematic uncertainties of next-generation long baseline neutrino experiments. An additional milestone will be the deployment of the first Large Area Picosecond Photodetectors (LAPPD). These novel detectors will feature a time resolution less than 100 picoseconds and a spatial accuracy of a few millimetres, thus improving the track reconstruction capabilities of the detector. This talk will give a general overview of ANNIE including an update on the currently running Phase Two as well as an upcoming expansion using Water based Liquid Scintillator (WbLS).

T 59.5 Wed 17:20 T-H18

Water-based Liquid Scintillators in ANNIE — DANIELE GUFFANTI², DAVID MAKSIMOVIC¹, •MICHAEL NIESLONY¹, and MICHAEL WURM¹ for the ANNIE-Collaboration — ¹Johannes Gutenberg-Universität Mainz, Germany — ²Università degli Studi di Milano Bicocca, Italy

The Accelerator Neutrino Neutron Interaction Experiment (ANNIE) is a Gadolinium doped water Cherenkov detector located in the Booster Neutrino Beam at Fermilab with the primary goal of measuring the final state neutron multiplicity of neutrino-nucleus interactions. A future phase of the experiment will explore the benefits of using the novel detection medium of Gadolinium-doped Water-based Liquid Scintillators (GdWbLS) in a neutrino beam environment by placing a small vessel within the current detector. GdWbLS aims to combine the advantages of liquid scintillator and water Cherenkov detectors by accessing directional information from the Cherenkov light and simultaneously using the scintillation signal to infer additional calorimetric event properties, especially with respect to lower-energy hadronic recoil signals. The following talk will highlight the expected improvements for the neutrino energy reconstruction in beam events for GdWbLS as a target material in comparison to a more conventional water detection volume based on simulation studies.

T 59.6 Wed 17:35 T-H18

Overview of the ESSv**SB Conceptual Design** – •TAMER TOLBA – Institut für Experimentalphysik, Universität Hamburg, Hamburg - Germany

In the search for the CP-violation in the leptonic sector, crucial information has been obtained from neutrino experiments. The measurement of the third neutrino mixing angle, θ_{13} , opened the possibility of discovering the Dirac leptonic CP violating angle, δ_{CP} , with intense "super" neutrino beam experiments. In the

Location: T-H19

light of these new findings, an urgent need has arisen to improve the detection sensitivity of the current long-baseline detectors, considering proton driver at MW scale with MegaTon scale detector, with a key modification to place the far detectors at the second, rather than the first, oscillation maximum.

The European Spallation Source neutrino Super Beam (ESSvSB) aims to benefit from the high power of the ESS, LINAC in Lund-Sweden, to produce the world's most intense second-generation neutrino beam, enabling measurement to be made at the second oscillation maximum. Assuming a ten-year exposure

T 60: Top Quarks: Decay and CP Violation and Mixing Angles

Time: Wednesday 16:15-18:30

sensitivity to small coupling constants.

T 60.1 Wed 16:15 T-H19

Tagging of Boosted Leptonically Decaying Top Quarks Using Convolutional Neural Networks — •HALA ELHAG, SOHAM BHATTACHARYA, and ISABELL MELZER-PELLMANN — DESY, Hamburg, Germany

The study of boosted top quarks is very important for probing a wide variety of new physics models. The use of machine learning techniques for tagging leptonically decaying boosted top quarks has not yet been explored as extensive as the hadronic decay channel. In this study, we utilize an image based machine learning technique for tagging highly boosted leptonically decaying top quarks. Jet images – representing the energies of jet constituents displayed in the form of a grid of pixels – are used as inputs to our convolutional neural network (CNN) based tagger. This talk will discuss the details of the jet formation, jet preprocessing and the CNNs, and a few promising preliminary results will be shown.

T 60.2 Wed 16:30 T-H19

Studies for the search for $t \rightarrow Zc$ transitions via interference effects — •LUCAS CREMER, JOHANNES ERDMANN, RONI HARNIK, JAN LUKAS SPÄH, and EMMANUEL STAMOU — TU Dortmund University, Department of Physics Exclusion limits for anomalous flavour-changing neutral currents (FCNC) are typically set by searching for two-body decays of the top quark, which are quadratic in the new physics (NP) coupling. The limits are constantly improved by new data, but the sensitivity to small couplings could be enhanced by an alternative approach. This approach takes advantage of the interference between

the FCNC and the Standard Model (SM) contributions. The focus of this work is to access the experimental feasibility of this idea. Concretely we investigate the three-body decay $t \rightarrow b\bar{b}c$ in the presence of an anomalous t - Z - c coupling. In the SM, this process proceeds through the emission of a W boson, while the FCNC diagram contains an intermediate Z boson. The two contributions interfere. The dominant contribution of the interference is in the restricted kinematic region in which both intermediate bosons are onshell. In this region, both the SM and the pure FCNC contribution are suppressed by the small width of the gauge bosons, which enhances the impact of the interference contribution. Because the interference scales linearly with the NP coupling constant, while the pure FCNC contribution scales with the coupling constant squared, an analysis of events in this region can potentially improve the

T 60.3 Wed 16:45 T-H19

Search for FCNC-couplings between the top-quark and the Higgs-boson in dilepton final states — •MARVIN GEYIK¹, WOLFGANG WAGNER¹, OLIVER THIELMANN¹, ABHISHEK SHARMA², FREDERIC DELIOT³, CHARLES GRANT⁴, PAUL JACKSON⁴, PETER ONYISI⁵, KYUNGEON CHOI⁵, and MARC TOST⁵ — ¹Bergische Universität Wuppertal — ²Columbia University — ³Universite Paris-Saclay — ⁴University of Adelaide — ⁵University of Texas at Austin

Flavor-changing neutral current interactions are strongly suppressed in the Standard Model. Still, some extensions of the Standard Model predict tree-level FCNC-couplings between the top quark, other up-type quarks and neutral bosons, including the Higgs boson. These anomalous couplings can be parameterised in the framework of effective field theories (EFT). The presented analysis searches for the production of a single top-quark in association with a Higgs boson and for top-quark-antiquark production with one of the top quarks decaying to an up quark or a charm quark and a Higgs boson. Higgs decays to WW*, ZZ* and two taus leading to leptonic final states are considered in the event selection. Two analysis channels are defined: one with two leptons (electrons or muons) of the same electric charge and a second channel with three leptons. This talk focuses on advancements in the dilepton final state. The sensitivity of the analysis in setting limits to relevant coefficients of EFT operators will be presented.

T 60.4 Wed 17:00 T-H19

Search for flavour-changing neutral current couplings between the top-quark and the Higgs boson in the $H \rightarrow b\bar{b}$ decay channel and the tri-lepton final state with the ATLAS detector at the LHC — •OLIVER THIELMANN¹, GeoFFREY GILLES⁴, WOLFGANG WAGNER¹, MARVIN EMIN GEYIK¹, DOMINIC HIRSCHBÜHL¹, KYUNGEON CHOI², FREDERIC DELIOT³, CHARLES MICHAEL with five-years running in neutrino- and five-years in antineutrino-mode, CPviolation could be established with a significance of 5σ over more than 70% of all values of δ_{CP} . With the current design-study program of the experiment is coming to its successful end, with the production of the CDR, an overall status of the project will be presented. The technical aspects on the current design study programs running within the collaboration and the physics potential of the experiment will be presented, as well.

GRANT⁵, PAUL JACKSON⁵, PETER ONYISI², ABHISHEK SHARMA⁶, and MARC TOST² – ¹Bergische Universität Wuppertal – ²Austin – ³Saclay CEA – ⁴Nikhef – ⁵Adelaide – ⁶Columbia

A search for flavour-changing neutral current (FCNC) couplings between the top-quark and the Higgs boson in the $H \rightarrow b\bar{b}$ decay channel and the tri-lepton final state is presented. The search for FCNC couplings in the top-quark-Higgs-boson sector is a promising search for a theory beyond the SM. Proton-proton collision data produced by the LHC at a centre-of-mass energy of $\sqrt{s} = 13$ TeV and collected by the ATLAS experiment during 2015, 2016, 2017 and 2018, and corresponding to an integrated luminosity of 139 fb^{-1} , are used. Data is analysed in different final states, characterised by the number of jets where either three (for $H \rightarrow b\bar{b}$) or one (for tri-lepton final state) of them are identified as b-jets. A machine learning analysis based on neural networks is conducted to improve the discrimination between the signal and the backgrounds. Preliminary results, interpreted in the context of an effective field theory for FCNC, are presented, where additional exclusion limits on the qtH effective coupling are derived.

T 60.5 Wed 17:15 T-H19

Search for charged lepton flavour violation in top-quark production and decay with the ATLAS experiment at 13 TeV — MARKUS CRISTINZIANI¹, WILLIAM GEORGE², •GABRIEL GOMES¹, CARLO GOTTARDO³, CHRIS HAWKES², JACOB KEMPSTER², ALEXIOS STAMPEKIS², and MIRIAM WATSON² — ¹Center for Particle Physics Siegen, Experimentelle Teilchenphysik, Universität Siegen — ²University of Birmingham — ³NIKHEF

In the Standard Model (SM) with massless neutrinos, the flavour of charged leptons cannot be altered in weak interactions. However, the observed neutrino oscillations allow for charged lepton flavour violating (cLFV) processes, even though highly suppressed. Hence, experimental evidence of such rare processes would provide signs of new physics beyond the SM.

Investigations targeting a direct search for cLFV will be presented using proton–proton collision data collected by the ATLAS detector between 2015 and 2018 at $\sqrt{s} = 13$ TeV. Decays of a top quark into a pair of opposite-sign different-flavour (OSDF) leptons and an up-type quark, as well as single top-quark production in association with an OSDF dilepton pair, are examined. Thus, be-sides the top-quark decay channel, the single top-quark production channel is included, providing additional sensitivity. For signal-discrimination purposes, a multivariate discriminant, namely a boosted decision tree, is implemented and optimised.

T 60.6 Wed 17:30 T-H19 Measurement of *CP* violation in $B_s^0 \rightarrow D_s^+ D_s^-$ and $B^0 \rightarrow D^+ D^-$ decays with the LHCb experiment — •LOUIS GERKEN, PHILIPP IBIS, and ANTJE MÖDDEN — Experimentelle Physik 5, TU Dortmund

At the LHCb experiment, time-dependent measurements of *CP* violation are performed to test the Standard Model of particle physics. The decays $B_s^0 \rightarrow D_s^+ D_s^$ and $B^0 \rightarrow D^+ D^-$ give access to the *CP* violation parameters ϕ_s and $\sin(2\beta)$. In these decays of neutral mesons, *CP* violation arises in the interference of the direct decay and the decay after mixing. Due to the similarities of the decays, a time-dependent *CP* violation measurement is performed in parallel for both decays.

In this talk, the current status of these measurements will be presented. The analysis uses data collected by the LHCb detector during 2015 to 2018 at a centre-of-mass energy of 13 TeV corresponding to an integrated luminosity of 6 fb⁻¹.

T 60.7 Wed 17:45 T-H19 Search for $B_s^0 \rightarrow D^{*+}D^{*-}$ and *CP* violation studies in $B_d^0 \rightarrow D^{*+}D^{*-}$ with the LHCb experiment — SOPHIE HOLLITT, PHILIPP IBIS, •JAN LANGER, and ANTJE MÖDDEN — Experimentelle Physik 5, TU Dortmund

At the LHCb experiment, precision measurements are performed to search for physics beyond the Standard Model. For this purpose, e.g. searches for unobserved decays and measurements of their branching fractions or measurements of CP violation in decays of neutral B mesons are carried out.

The primary aim of this analysis is to observe the decay $B_s^0 \rightarrow D^{*+}D^{*-}$. Besides, the branching fraction is measured relative to the decay $B_d^0 \to D^{*+}D^{*-}$. By measuring the relative branching ratio, dominant systematic uncertainties cancel out. Further, an angular and decay-time dependent CP violation measurement is performed in the $B_d^0 \to D^{*+} D^{*-}$ decay, which allows the measurement of the parameter $sin(2\beta)$.

In this talk, the current status of both analyses is presented using the full data set of the LHCb experiment corresponding to an integrated luminosity of 9 fb^{-1} .

T 60.8 Wed 18:00 T-H19 Measurement of the CKM mixing angle γ with $B_s^0 \rightarrow D_s^{\mp} K^{\pm}$ decays at the LHCb experiment — • QUENTIN FÜHRING and KEVIN HEINICKE — Experimentelle Physik 5, TU Dortmund

In the Standard Model of particle physics the quark mixing matrix is expected to be unitary. To test this unitarity, the properties of the unitarity triangles are constrained. The mixing angle y is such a property of interest.

To constrain the mixing angle y, precise decay-time-dependent measurements of *CP* violation in $B_s^0 \to D_s^{\mp} K^{\pm}$ decays can be used. With an excellent decay-time resolution and a large number of B_s^0 decays, the LHCb experiment provides the necessary data for this measurement.

T 61: Higgs Boson: Decay in Bosons

Time: Wednesday 16:15-18:15

T 61.1 Wed 16:15 T-H20

Effective Field Theory interpretation of the $pp \rightarrow H \rightarrow 4\ell$ Higgs boson decay measurements with the ATLAS detector — •ALICE REED, SANDRA KORTNER, and Нивегт Ккона — Max Planck Institut für Physik (Werner-Heisenberg-Institut), München

An important process for the measurement of the Higgs boson properties is the Higgs boson decay into two Z bosons, which subsequently decay into a $\mu^+\mu^-$ or an e^+e^- pair, $pp \to H \to 4\ell$. In the Standard Model (SM), the Higgs boson is predicted to be a spin-0 particle with a positive CP quantum number. This hypothesis is also favoured by the Run-1 data at the LHC. Still, small admixtures of anomalous and possibly also CP-violating couplings with non-SM tensor structure are not yet excluded.

Such deviations from the SM can be described within the effective field theory (EFT) framework in which the SM is extended by the addition of higherdimensional operators. In this talk, the EFT interpretation of the measured Higgs boson properties in the 4-lepton decay channel is presented, allowing constraints on several EFT parameters to be determined. Particular emphasis is given to the impact of the these EFT parameters on the acceptance of the fourlepton event selection criteria, which needs to be taken into account in addition to the EFT effects on the production cross section and branching ratio.

T 61.2 Wed 16:30 T-H20

Measurement of $H \rightarrow WW^*$ Decays in the ℓvqq Final State with a Large-R Jet — •JOHANNES HINZE, KARSTEN KÖNEKE, and BENEDICT WINTER — Universität Freiburg

The talk presents a study of $H \rightarrow WW^*$ decays at large transverse momenta (pT(H) > 200 GeV) with one leptonic $(W \rightarrow \mu \nu \text{ or } W \rightarrow e\nu)$ and one hadronic W boson decay, where the experimental signature of the hadronic W boson decay is a large-R jet. The lepton provides means to efficiently trigger event candidates and to eliminate background events in particular from multijet events. Further background events, primarily from W+jets events, can be suppressed via W-boson taggers for large-R jets. The measurement benefits from the larger branching fraction in comparison with $\ell \nu \ell \nu$ final states, and from the reduced background levels for large transverse momenta. The measurement will contribute significantly in an area of the phase space that is considered particularly sensitive to possible BSM effects.

T 61.3 Wed 16:45 T-H20

Multivariate Techniques for Measurement of Higgs Bosons in $H \rightarrow WW^* \rightarrow$ evµv Decays at ATLAS — •AHMED MARKHOOS, BENEDICT WINTER, and KARSTEN KÖNEKE — University of Freiburg

Since its discovery, the Higgs Boson has been studied in detail at the LHC. The $H \rightarrow WW^* \rightarrow e \nu \mu \nu$ channel offers sizeable signal and moderate background yields enabling accurate measurements of the total cross-section and of differential cross-sections. The measurements for gluon-fusion production are generally dominated by systematic uncertainties except in the sparsely populated regions of the phase space such as at large transverse momenta.

The talk showcases deep neural networks (DNN) that can enhance the signal purity with respect to the current cut-based selection reducing systematic uncertainties from backgrounds and statistical uncertainties. Additionally, a regression DNN is presented that determines the Higgs Boson transverse momentum, which is elusive due to the presence of neutrinos, and required to measure simplified template cross-sections (STXS).

In this talk a decay-time-dependent analysis aiming to measure *y* is presented. Data corresponding to an integrated luminosity of 6 fb⁻¹ recorded by the LHCb experiment from 2015 to 2018 at a centre-of-mass energy of 13 TeV are used for this analysis.

T 60.9 Wed 18:15 T-H19

Measurement of CP violation in $B^0 \rightarrow \psi K_S^0$ decays with the LHCb experiment — VUKAN JEVTIC, PATRICK MACKOWIAK, and •GERWIN MEIER — Experimentelle Physik 5, TU Dortmund

Precision measurements of parameters of the Standard Model are important methods for tests of the Standard Model. One excellent parameter to measure is the CKM angle β , where the golden mode is $B^0 \rightarrow J/\psi K_S^0$ due to the dominant contributions of tree-level amplitudes. With new reconstruction types of the $K_{\rm S}^0$ and the combination of different decay channels it is possible to increase the statistical sensitivity in the most precise measurement of this quantity to date.

In this talk the current status of the time-dependent $sin(2\beta)$ measurement in the decays $B^0 \to J/\psi(\to \ell\ell)K^0_S(\to \pi^{\pm}\pi^{\mp})$ with $\ell = e, \mu$ and $B^0 \to \psi(2S)(\to \pi^{\pm}\pi^{\mp})$ $\mu\mu$) $K_{S}^{0}(\rightarrow \pi^{\pm}\pi^{\mp})$ will be presented, where the full LHCb Run II dataset from 2015 to 2018 corresponding to 6 fb^{-1} is used.

Location: T-H20

T 61.4 Wed 17:00 T-H20

Search for Di-Higgs production in the $bb\gamma\gamma$ final state with the ATLAS detector - •FLORIAN BEISIEGEL, JOCHEN DINGFELDER, and TATJANA LENZ -Physikalisches Institut, Uni Bonn

The discovery of the Higgs boson in 2012 was a great success of modern particle physics since it served as a proof of the Higgs mechanism introduced in 1964. One focus of the current particle physics experiments at the LHC is the measurement of the Higgs properties, such as its coupling strengths to fundamental particles. In addition to the coupling of the Higgs boson to fermions and gauge bosons, the Higgs mechanism predicts Higgs self-coupling. The triple-Higgs self-coupling can be measured in di-Higgs (non-resonant) production. Di-Higgs analyses also facilitate the search for new heavy particles that decay to two Higgs bosons (resonant production).

This talk presents a search for di-Higgs production in the bbyy final state using 139 $\rm fb^{-1}$ of proton-proton collisions at 13 TeV recorded with the ATLAS detector. The analysis aims to measure the non-resonant SM di-Higgs production cross section and the Higgs self-coupling as well as search for resonant di-Higgs production. The focus is put on studies to improve the limits on the non-resonant production cross-section using a 2D fit in m_{yy} and m_{bb} .

T 61.5 Wed 17:15 T-H20

Search for non-resonant Higgs boson pair production in the bbWW final state with leptonic W boson decays at the CMS experiment - MARTIN ERDMANN, •PETER FACKELDEY, BENJAMIN FISCHER, and DENNIS NOLL — III. Physikalisches Institut A, RWTH Aachen University

The measurement of the Higgs boson pair production is a direct test of the electroweak symmetry breaking in the standard model of particle physics (SM) with direct access to the shape of the Higgs potential.

The cross section of the Higgs boson pair production is about a factor of a thousand smaller than that of a single SM Higgs boson, making it a highly challenging search. Physics-inspired deep learning techniques are leveraged for the signal extraction and the control over overwhelming backgrounds, mainly from the top pair production and Drell-Yan processes.

The expected sensitivity of the search for $HH \rightarrow bbW_{lep}W_{lep}$ is presented for the data-taking periods 2016, 2017, and 2018 of the CMS experiment.

T 61.6 Wed 17:30 T-H20

Search for non-resonant di-Higgs production in the semi-leptonic bbWW decay channel at the CMS experiment — MARTIN ERDMANN, PETER FACKELDEY, BENJAMIN FISCHER, and •DENNIS NOLL — III. Physikalisches Institut A - RWTH Aachen University

A measurement of the di-Higgs boson production can directly determine the trilinear Higgs coupling and probe the structure of the Higgs potential.

We present a search for Higgs boson pair production with one Higgs boson decaying into b quarks and the other Higgs boson decaying into W bosons, with one W boson decaying leptonically.

The central challenge of this analysis is a tiny signal among a large amount of background. We approach this task with a Deep Neural Network driven Physics Process Multi-Classification. It utilises a physics motivated architecture, the Lorentz-Boost Network, in conjunction with a Residual Neural Network.

We present expected limits corresponding to the data recorded at the CMS experiment in Run 2.

T 61.7 Wed 17:45 T-H20

Search for non-resonant Higgs boson pair production in the $b\bar{b}b\bar{b}$ final state with the CMS Experiment — MARTIN ERDMANN, PETER FACKELDEY, •BENJAMIN FISCHER, and DENNIS NOLL — III. Physikalisches Institut A, RWTH Aachen University

The non-resonant Higgs boson pair production enables probing the shape of the Higgs potential, in particular the triple Higgs self coupling λ_{hhh} . The decay channel with the highest branching ratio of ~ 1/3 has a four-*b*-quark final state.

This phase space is dominated by QCD-processes, which are challenging to model using Monte Carlo samples. A data-driven modeling is implemented through a Neural Network based reweigthing from a sideband region into the signal region. Through a Neural Network based multi-classification both Di-Higgs production modes, gluon- and vector-boson-fusion, are separated from background processes for the statistical inference.

T 61.8 Wed 18:00 T-H20 Search for non-resonant Higgs boson pair production in lepton+jets final states of the bbWW decay mode at CMS — •MATHIS FRAHM, JOHANNES HALLER, MATTHIAS SCHRÖDER, and ARTUR LOBANOV — Institut für Experimentalphysik, Universität Hamburg The Higgs boson self-coupling is an important parameter of the Standard Model, since it is related to the shape of the Higgs potential. At the LHC, this parameter can be probed by measuring the Higgs boson pair production (HH) cross section. In the Standard Model, HH production occurs in processes via Higgs-boson self-coupling and in processes with a fermion loop. Due to destructive interference of these two contributions, the resulting production cross section is small, amounting to only 33 fb at 13 TeV.

In this talk, a search for HH production in lepton+jets finals states of the bbWW decay mode is presented. The analysis is performed on data recorded by the CMS experiment during LHC Run 2 at a center-of-mass energy of 13 TeV, which corresponds to an integrated luminosity of 137.2 fb^{-1} . The analysis utilizes a deep neural network to classify between signal and different background categories. Exclusion limits on the production cross section are derived as a function of the Higgs boson self-coupling strength to set constraints on this parameter.

T 62: Higgs Boson: Extended Models 2

Time: Wednesday 16:15–18:30

T 62.1 Wed 16:15 T-H21 Combined measurements of Higgs boson production and interpretation in the context of two Higgs doublet models at the ATLAS experiment — •BIRGIT STAPF — Universität Hamburg/DESY, Hamburg, Germany

The discovery of the Higgs boson in 2012 is the latest big success story of the Standard Model of particle physics (SM). Although this discovery formally completes the SM, it is not an end to all questions on the matter. There are many observations and phenomena that the SM is unable to explain, and it is clear that there must be Beyond the Standard Model (BSM) physics. However, the lack of discoveries of (BSM) particles since 2012 puts particle physics at a cross-roads: there are many ways forward, posed by many different BSM theories, but it is unclear which one is likely to be successful as a description of reality. High precision measurements of the Higgs bosons' properties and its couplings help to guide the way: any measured deviation from the SM predictions indicate the existence of BSM physics and may even point at specific theories. To unlock the full potential of such measurements, the results from analyses of several different Higgs production processes and decays are combined. This talk covers the latest Higgs combination results from ATLAS using up to 139 fb⁻¹ of *pp*-collision data with \sqrt{s} =13 TeV, focussing in particular on the measurement of inclusive production cross-sections and the interpretation in terms of specific BSM models, such as two Higgs doublet models.

T 62.2 Wed 16:30 T-H21

Search for charged Higgs bosons in the $H^+ \to Wh \to qqbb$ decay channel — •Shubham Bansal, Jochen Dingfelder, and Tatjana Lenz — Physikalisches Institut, Universität Bonn

After the discovery of the Higgs boson at a mass of 125 GeV, the last missing piece of the Standard Model (SM) might be found. On the other hand, various theories beyond the SM predict additional Higgs bosons, one of which could be the found Higgs boson at 125 GeV. One such example is the two-Higgs-Doublet Model (2HDM) that features an extended scalar sector including the existence of charged Higgs bosons (H^+). The observation of such a charged scalar particle would clearly indicate physics beyond the SM. The H^+ production mechanism depends on its mass (m_{H^+}) and for $m_{H^+} > m_t + m_b$, the leading H^+ production mode is the associated production with a top and a bottom quark via $gg \rightarrow tbH^+$. In the alignment limit for 2HDM, the dominant decay mode is $H^+ \rightarrow tb$. However, in models like N2HDM and the Georgi-Machacek (GM) model, it is possible to obtain a sizable branching ratio for $H^+ \rightarrow Wh$.

This talk presents a search for charged Higgs bosons in $H^+ \rightarrow Wh \rightarrow qqbb$ decays. The recent developments of the analysis strategy will be discussed, which include the use of boosted decision trees to reconstruct the H^+ , data-driven corrections to improve the modeling of the main background, $t\bar{t}$, the definition of signal-enriched and -depleted regions, and a first estimate of the expected sensitivity using the full Run-2 ATLAS dataset.

T 62.3 Wed 16:45 T-H21

Search for a charged Higgs Boson decaying to *cs* in the low mass region with the ATLAS detector at $\sqrt{s} = 13$ TeV — JOCHEN DINGFELDER, TATJANA LENZ, and •CHRISTIAN NASS — Physikalisches Institut, Universität Bonn, Deutschland In the Standard Model (SM) electroweak symmetry breaking (EWSB) is introduced by a single complex scalar field. The consequence is the prediction of a scalar, neutrally charged particle, the Higgs Boson, which was discovered at the

Location: T-H21

LHC in 2012 at the LHC. A simple extension of the SM is to introduce EWSB through two complex scalar fields. Such two-Higgs doublet models (2HDM) are attractive because they offer the opportunity to include additional CP violation in the SM, which is needed for explaining baryogenesis. 2HDMs feature 3 neutral and 2 charged Higgs bosons. Observation of such a charged scalar would be a striking signal for physics beyond the SM.

In the low mass region, i.e. $m_H^{\pm} < m_t$, the dominant production mode is by a $t\bar{t}$ pair with one *t*-quark decaying to $H^{\pm}b$. At low masses, the search for $H^{\pm} \rightarrow cs$ decays is promising, as suggested in several theory papers. This talk will present background estimates, data-driven MC corrections and usage of c-tagging to define signal enriched and depleted regions as well as the first estimate of the expected sensitivity for the $H^{\pm} \rightarrow cs$ search with the full Run-2 ATLAS dataset recorded at a center-of-mass energy of 13 TeV.

T 62.4 Wed 17:00 T-H21

Search for $A \rightarrow ZH \rightarrow v\bar{v}b\bar{b}$ at $\sqrt{s} = 13$ TeV with the ATLAS detector — •ILIA KALAITZIDOU, TETIANA MOSKALETS, and SPYRIDON ARGYROPOULOS — University of Freiburg

The extension of the scalar SM Higgs sector, as described in the Two Higgs Doublet Models (2HDMs), could lead to a cosmological first order electroweak phase transition, which is necessary to explain the origin of the matter-antimatter asymmetry in the early Universe. The existence of a second Higgs doublet results in five physical scalar fields, two charged (H^{\pm}), a CP-odd (A) and two CP-even (h and H) neutral fields. A strong electroweak phase transition favours a heavy CP-odd scalar state A, together with a large mass splitting between the CP-odd A and CP-even H scalars. In this scenario, the $A \rightarrow ZH$ decay becomes dominant. In the present work, the $A \rightarrow ZH$ decay is investigated, with the H boson decaying to a pair of b-quarks and the Z boson decaying to neutrinos. The $Z \rightarrow v\bar{v} \bar{v}$ decay is examined, because of the expected increased sensitivity for large A masses. The optimisation studies for the $A \rightarrow ZH \rightarrow v\bar{v}b\bar{b}$ analysis are presented, along with the expected exclusion in the $m_H - m_A$ plane, covering a previously unexplored region.

T 62.5 Wed 17:15 T-H21

Search for $A \rightarrow ZH \rightarrow \ell\ell t\bar{t}$ at $\sqrt{s} = 13$ TeV with the ATLAS detector — •ROMAN KUESTERS, TETIANA MOSKALETS, and SPYROS ARGYROPOULOS — University of Freiburg, Freiburg im Breisgau, Germany

The generation of the matter-antimatter asymmetry in the universe is one of the biggest open questions that require physics beyond the Standard Model. An attractive explanation is provided by the electroweak baryogenesis models, which require the addition of a second Higgs doublet, giving rise to five Higgs bosons: a light (heavy) *CP*-even Higgs h(H), a *CP*-odd one (*A*) and two charged ones (H^{\pm}) . A necessary requirement for baryogenesis is a large mass splitting between the heavy *CP*-odd and *CP*-even Higgs bosons, which makes the $A \rightarrow ZH$ decay dominant.

This talk presents a search for the $A \rightarrow ZH \rightarrow \ell\ell t\bar{t}$ process, targeting the phase space with $m_H > 350$ GeV, which has not been explored so far. The talk will discuss the optimisation of the event selection and the sensitivity expected to be achieved with the full Run 2 ATLAS data.

T 62.6 Wed 17:30 T-H21

Search for heavy Higgs bosons in the $Zt\bar{t}$ final state with CMS — •DANIEL CHRISTIAN HUNDHAUSEN, KSENIA DE LEO, YANNICK FISCHER, JOHANNES HALLER, and MATTHIAS SCHRÖDER — Institut für Experimentalphysik, Universität Hamburg

Since its discovery in 2012 the properties of the Higgs boson at 125 GeV have been studied extensively, confirming a standard model like behavior. However, the observed resonance might well be part of an extended Higgs sector, which is predicted in various scenarios of new physics beyond the standard model. Two Higgs Doublet Models (2HDM) provide a generic description of the phenomenology arising in models with a second Higgs doublet. In this talk we will investigate the hypothetical decay of a CP odd heavy Higgs boson A decaying into a CP even heavy Higgs boson H and a Z boson, with the H decaying further into a pair of top quarks. This decay channel is particularly relevant in the high mass and low tan(β) regime. We will present the strategy and status of our analysis, targeting the fully hadronic $t\bar{t}$ decay with the Z boson decaying to $\mu^+\mu^-$.

T 62.7 Wed 17:45 T-H21

Search for heavy Higgs bosons in the $Z + t\bar{t}$ final states with CMS — •YANNICK FISCHER, KSENIA DE LEO, JOHANNES HALLER, DANIEL HUNDHAUSEN, and MATTHIAS SCHRÖDER - Institut für Experimentalphysik, Universität Hamburg Since its discovery in 2012, the properties of the Higgs boson at 125 GeV have been studied in numerous measurements. Within the uncertainties all results suggest a Standard Model like behaviour. However, the observed boson might well be part of an extended Higgs sector, which is predicted in various scenarios of new physics beyond the standard model. Two Higgs Doublet Models (2HDM) provide a generic description of the phenomology arising in models with a second Higgs doublet. In this talk we will investigate the hypothetical decay of a CP odd heavy Higgs boson A decaying into a CP even heavy Higgs boson H and a Z boson, with the H decaying further into a pair of top quarks. This decay channel is particularly relevant in the high mass and low $\tan(\beta)$. We will give an overview about the parameter space with detailed phenomenological studies, the results of the investigations of the kinematic properties of the decay products and present first expected exclusion limits in the decay channel $Z \rightarrow \mu^+ \mu^-$ and $t\bar{t} \rightarrow$ jets.

T 62.8 Wed 18:00 T-H21

Search for light pseudoscalar boson pairs produced from decays of the 125 GeV Higgs boson in final states with tau leptons — •LAKSHMI PRAMOD — DESY

An extended Higgs sector is well-motivated in several Beyond the Standard Model theories. A vast set of models containing two Higgs doublets plus one

additional Higgs singlet complex field (2HD+1S) are consistent with SM measurements, constraints from searches for additional Higgs bosons and supersymmetry, as well as with the measured properties of the H(125) boson. The Higgs sector of the 2HD+1S models contains seven physical states: three CP-even, two CP-odd and two charged bosons. In the context of these models, the H(125) boson can decay into a pair of light pseudoscalar bosons (a1), which can subsequently decay to pairs of Standard Model particles. There exist scenarios where a_1 can have an enhanced decay rate to a pair of τ leptons. A search for a pair of light bosons, a1, produced from decays of 125 GeV Higgs boson, each decaying to a pair of τ leptons will be presented. The search is based on proton-proton collision data collected by the CMS experiment during Run 2 at a centre of mass energy of 13 TeV, corresponding to an integrated luminosity of 138 fb⁻¹. Modelindependent upper limits at 95% confidence level on the 125 GeV Higgs boson production cross-section times the branching fraction into the studied final state, relative to the SM H(125) production cross-section, are set. Model-specific upper bounds obtained as constraints on the parameter space of the different benchmark scenarios within the 2HDM+S will also be presented.

T 62.9 Wed 18:15 T-H21

Search for a light CP-odd Higgs boson decaying into a pair of taus with AT-LAS — •JANNIK FRIESE, TOM KRESSE, MAX MÄRKER, WOLFGANG MADER, and ARNO STRAESSNER — IKTP, Dresden, Germany

Even though theoretical predictions of the SM are corresponding to experimental results to an incredible degree, there are still some phenomena unexplained, for example the deviation of the measured anomalous magnetic moment, g - 2, of the muon from SM calculations. This deviation could be explained by the flavor-aligned two-Higgs-doublet model (2HDM). The introduction of a second Higgs doublet leads to four additional Higgs bosons, one of which being CP-odd and electrically neutral. The muon g - 2 deviation is best explained with a light CP-odd Higgs boson which couples nearly exclusively to top quarks and tau leptons. This talk presents the search of such a light CP-odd Higgs boson produced via gluon fusion. The decay into two tau leptons is analyzed by looking at one electron and one muon in the final state. The mass of the light Higgs boson is assumed to be in the range between 40 GeV and 90 GeV. The search is based on 139 fb⁻¹ of data collected by the ATLAS experiment at 13 TeV center of mass energy. Before unblinding the signal region, background control regions are analyzed to verify a good description of the data distributions. In particular, the transverse momentum spectrum of the electrons and muons receives contributions from QCD background which is not modeled well by Monte Carlo simulations. This talk focuses on the estimation of this source of background using a data driven fake-factor method.

T 63: Search for New Particles 4

Time: Wednesday 16:15-18:30

T 63.1 Wed 16:15 T-H22

Search for a long-lived particle in $b \rightarrow s$ transitions at Belle II — •SASCHA DREYER¹ and TORBEN FERBER² for the Belle II-Collaboration — ¹DESY, Hamburg, Germany — ²ETP, KIT, Karlsruhe, Germany

The Belle II experiment at the asymmetric e^+e^- SuperKEKB collider in Tsukuba, Japan provides an ideal test bench for searches for light dark sectors, due to a clean collision environment leading to low backgrounds.

A hypothetical new long-lived particle could serve as a portal to dark sectors. This particle could be produced in B-meson decays via $b \rightarrow s$ quark transitions and decay to pairs of charged Standard Model particles. The displaced vertex signature can be reconstructed in case the particle decays within the tracking detectors.

This talk gives an overview of the search for such a new long-lived particle at Belle II. The sensitivity for different lifetime and mass scenarios will be shown together with work towards validating long-lived particle performance in data using control samples.

T 63.2 Wed 16:30 T-H22

Study of $e^+e^- \rightarrow D_{s}^{\pm}D_{s0}^{*}(2317)^{\mp}A$ process at Belle+BaBar — •DMYTRO MELESHKO¹, ELISABETTA PRENCIPE^{1,2}, JENS SOEREN LANGE¹, and ASHISH THAMPI² — ¹Justus-Liebig Univ. Giessen — ²IKP-1, Forschungszentrum Juelich The present analysis is focused on the study of the $e^+e^- \rightarrow D_s^{\pm}D_{s0}^{*}(2317)^{\mp}A$ process in the continuum (A = anything else) combining the BaBar and Belle data sets to cure problems of insufficient statistics. The main goal of this work is the analysis of the $D_s^+ D_{s0}^*(2317)^-$ invariant mass system to look for possible resonant states with $c\bar{c}s\bar{s}$ quark content, and cross-section measurement. At the current stage of the analysis, a preliminary study over MC samples and Belle data has already been performed. The acquired results contain intriguing sign of a possible state seen in the invariant mass distribution of the $D_s^+D_{s0}^*(2317)^{\mp}$ system. In addition, the analysis is opened to a potential perspective of measuring the $D_{s0}^*(2317)^\pm$ width upper limit, which plays an important role in understanding the nature of the $D_{s0}^*(2317)^\pm$ itself.

T 63.3 Wed 16:45 T-H22

Location: T-H22

Search for inelastic Dark Matter with a Dark Higgs at Belle II — •PATRICK ECKER, TORBEN FERBER, PABLO GOLDENZWEIG, and JONAS EPPELT for the Belle II-Collaboration — Karlsruhe Institute of Technology, Karlsruhe, Germany

Although the Standard Model (SM) of particle physics describes most of the phenomena observed in our universe very well, there are still some observations where the SM lacks to provide an explanation. One of these observations is the presence of Dark Matter which is very well motivated. Nevertheless, it is still not clear which particles make up this Dark Matter.

This talk will present a sensitivity study based on Monte Carlo simulations for a search for an inelastic Dark Matter model which involves the presence of a Dark Higgs boson. This model has a signature of up to two displaced vertices, one from the resonant decay of the Dark Higgs and another non-resonant one emerging from the decay of the involved Dark Matter particles.

T 63.4 Wed 17:00 T-H22

Search for $B^{\pm} \to K^{\pm}a(a \to \gamma\gamma)$ with promptly decaying ALPs - A Monte Carlo study — •LUCAS WEIDEMANN, PABLO GOLDENZWEIG, and TORBEN FERBER — Institute of Experimental Particle Physics, Karlsruhe, Germany

Axion-Like-Particles (ALPs) are a well motivated extension to the Standard Model. They are light, pseudoscalar particles which are dominantly interacting with Standard Model gauge bosons. As a result ALPs can be produced in a flavor-changing neutral current transition by its coupling g_{aWW} to W bosons. Subsequently, the ALP dominantly decays into two photons, which makes $B^{\pm} \rightarrow K^{\pm}a(a \rightarrow \gamma\gamma)$ a promising process for measuring g_{aWW} .

We use Monte Carlo simulation for analyzing this process by reconstructing the *B* meson and examining the invariant di-photon mass distribution. The study

An outline of our analysis as well as our current status are going to be presented in this talk.

Sensitivity study in the Search for $B^{\pm} \to K^{\pm}a$ (displaced $a \to \gamma\gamma$) Decays at Belle II — •ALEXANDER HEIDELBACH, PABLO GOLDENZWEIG, and TORBEN FERBER — Institute of Experimental Particle Physics, Karlsruhe Institute of Technology, Karlsruhe, Germany

In a set of extensions of the Standard Model, Axion Like Particles (ALPs) arise as (pseudo) Nambu-Goldstone bosons of an additional spontaneously broken U(1) symmetry. Due to constraints in flavour-changing processes, involving the coupling of ALPs directly to Standard Model fermions and gluons, direct couplings of ALPs to the electroweak gauge bosons are particularly interesting. As a result of coupling to *W* bosons, ALPs can emerge in flavour-changing-neutral-current $b \rightarrow s$ transitions. Depending on the ALP model, mass and the coupling to photons possible signatures become probable in $B \rightarrow Ka$ transitions which can be studied at e^+e^- colliders like the Belle II experiment. I present a search for $B^{\pm} \rightarrow K^{\pm}a$, $a \rightarrow \gamma \gamma$ with long-lived ALPs. We investigate the mentioned decay based on a full Belle II Monte Carlo study. In this talk, I will show sensitivity estimates established by the reconstruction of the *B* meson, an optimised candidate selection and a scan of the invariant di-photon mass spectrum for different ALP lifetimes.

T 63.6 Wed 17:30 T-H22

Dark Photon Searches at Future e^+e^- **Colliders** — •SEPIDEH HOSSEINI^{1,2}, JENNY LIST², MIKAEL BERGGREN², and GUDRID MOORTGAT-PICK^{1,2} — ¹Universität Hamburg, Hamburg, Germany — ²Deutsches Elektronen-Synchrotron (DESY), Hamburg, Germany

The dark photon (A_D) is a hypothetical particle that can be possibly produced through its kinetic mixing with the ordinary, visible photon. The existence of kinetic mixing means that the two gauge bosons can transform into each other as they propagate and this provides a link between the dark and visible sectors. The decay modes of the dark photon to the standard model charged fermions motivate to look at $A_D \rightarrow \mu^+ \mu^-$ as signal. The dark photon will have a specific mass and hence, the invariant mass of the muon pair is the main observable to look for the dark photon in the presence of standard model background. For this, we evaluate the prospects to detect the dark photon and to determine the mixing parameter for the example of the International Large Detector (ILD) concept at the International Linear Collider (ILC).

T 63.7 Wed 17:45 T-H22

Investigation of Collider Effects of Flavour Anomalies Using EFTs and Simplified Models — PHILIP BECHTLE, KLAUS DESCH, CHRISTIAN GREFE, and •MURILLO VELLASCO — Rheinische Friedrich-Wilhelms-Universität Bonn The Standard Model of particle physics (SM) is undoubtedly one of the most successful scientific theories in history. Despite its overwhelming success, several tensions with its predictions have been discovered in recent years, including the measurements of the muon g-2 at Fermilab and rare B-decays at the LHC. The natural next step is to build even more powerful future colliders, but exactly which combination of future experiments is best to investigate further these hints is non-trivial.

The most challenging possibility is a scenario where the energy scale of New Physics is out of reach of direct discovery at future colliders. Therefore, our approach is to use the formalism of the Standard Model Effective Field Theory (SMEFT), performing scans over several operators which could perhaps explain the hints of experimental deviations from the SM. Kinematic observables of collider signatures can be directly affected by the presence of these operators, and the optimal combination of future experiments would be the one that optimizes the observation of these kinematic deviations. In this talk, we will discuss a possible strategy to choose the best combination of experiments given all current measurements in the electroweak and the flavour sectors. Ideally, any strategy for future colliders should aim for a "no-lose" scenario, analogous to the situation for the approval of the LHC.

T 63.8 Wed 18:00 T-H22

Heavy neutrino search at LHCb — MARTINO BORSATO, REBECCA GARTNER, and •MAURICE MORGENTHALER — Physikalisches Institut - Universität Heidelberg

Neutrino masses could be explained by the existence of heavy neutrinos. These elusive particles might have escaped detection at previous experiments due to their long lifetimes and low production rate. If their mass is in the range of a few GeV, heavy neutrinos are produced copiously from the weak decays of beauty hadrons and can be searched effectively at the world-brightest source of these hadrons: the Large Hadron Collider (LHC). In this talk I will present an ongoing search for heavy neutrinos produced in (semi)leptonic beauty-hadron decays using the dataset collected by the LHCb experiment at the LHC. The search strategy relies on the heavy-neutrino macroscopic lifetime and its decay to a pion and a muon. The sensitivity is maximised by targeting all beauty hadron species (including strange and charmed B mesons) and using a partial reconstruction of the decay. The sensitivity of the search in comparison to other experiments will be discussed.

T 63.9 Wed 18:15 T-H22

Search for Heavy Majorana Neutrinos in same-sign W Boson Scattering — •JONAS NEUNDORF — Deutsches Elektronen-Synchrotron, Notkestraße 85, 22607 Hamburg

Among the open question of particle physics is the origin of neutrino masses. While they are predicted to be zero by the Standard Model, oscillation measurements have shown that at least two of the three neutrino flavours observed in nature are massive. These masses can be explained by the "Seesaw Mechanism", which introduces Majorana neutrinos with a mass on the TeV scale. For the first time at the LHC, a search for Heavy Majorana Neutrinos produced via same-sign W boson scattering is performed. This talk will discuss the discovery potential and outline the analysis design.

T 64: Search for New Particles 5

Time: Wednesday 16:15-18:30

T 64.1 Wed 16:15 T-H23

Model Unspecific Search in CMS (MUSIC) — •LORENZO VIGILANTE¹, ARND MEYER¹, THOMAS HEBBEKER¹, and SARANYA SAMIK GHOSH² — ¹III. Physikalisches Institut A, RWTH Aachen, 52074 Aachen, Germany — ²CERN, Experimental Physics Department, Geneva, Switzerland

The Model Unspecific Search in CMS (MUSiC) is a search for new physics beyond the standard model (BSM). The analysis looks for significant deviations from the standard model (SM) expectation in LHC data. The method of the analysis is to compare kinematic distributions of the data with the SM expectation in hundreds of different final states, using an automated procedure. This strategy allows MUSiC to search for new phenomena in final states that are not all covered by dedicated analyses in CMS. In this talk the general method and its implementation will be discussed and results of the MUSiC analysis using 35.9 fb⁻¹ of data collected by the CMS experiment during proton-proton collisions at a center of mass energy of 13 TeV will be presented. The current status of the analysis and a possible new search method, based on machine learning technique, will be described.

T 64.2 Wed 16:30 T-H23

Search for high mass lepton flavour violating processes with CMS — •SEBASTIAN WIEDENBECK, THOMAS HEBBEKER, ARND MEYER, and SWAGATA MUKHERJEE — III. Physikalisches Institut A, RWTH Aachen University Lepton flavour is a conserved quantity in the standard model of particle physics, but it does not follow from an underlying symmetry. Neutrino oscillations imply that lepton flavour is not conserved in the neutral sector. Lepton flavour violating processes are common in several models of physics beyond the standard model (e.g. supersymmetry with R-parity violation, black hole production, and leptoquarks). Some models predict objects at the TeV mass scale that can decay into two standard model leptons of different flavours: electron + muon, muon + tau, or electron + tau. The challenges in a search for such phenomena are to achieve a high mass resolution, good rejection of standard model backgrounds, and efficient lepton identification at the same time. The status of the analysis is presented, based on the latest CMS data taken in Run 2.

T 64.3 Wed 16:45 T-H23

Location: T-H23

Search for new physics in the τ +MET final state with CMS — •Christoph Schuler, Kerstin Hoepfner, Thomas Hebbeker, and Swagata Mukherjee — III. Physikalisches Institut A, RWTH Aachen University

A search for new physics in the τ +missing transverse momentum (MET) channel is presented based on proton-proton collisions measured with the CMS detector at the LHC, using the full Run-2 CMS data set recorded at a center of mass energy of 13 TeV. The analysis strategy is discussed and the results are interpreted in the context of various models predicting enhancements to the Standard Model in the high mass region.

T 64.4 Wed 17:00 T-H23 Interpreting the ATLAS missing-energy-plus-jets measurement for New Physics — •MARTIN HABEDANK and PRISCILLA PANI — Deutsches Elektronen-Synchrotron (DESY) The large dataset of 139 fb⁻¹ of proton–proton collisions at 13 TeV recorded with the ATLAS detector allows for a precise measurement of events with large missing transverse momentum and at least one jet. In the Standard Model, this final state can mostly be attributed to Z bosons being produced in association with jets, with subsequent decay of the Z boson into neutrinos. By correcting for detector effects, the measurement can thus make an important contribution to our Standard Model understanding and modelling.

However, various questions that cannot be answered within the framework of the Standard Model give rise to models predicting New Physics. Many of those – in particular when incorporating a Dark Matter candidate – also give rise to events with missing transverse momentum and one or more jets, rendering the mentioned measurement a powerful handle in constraining these models.

This talk gives insight into the approach taken by the mentioned measurement and progress on interpreting it with regard to New Physics models.

$T~64.5 \quad Wed~17:15 \quad T-H23$ **Identification of highly boosted Z** \rightarrow e⁺e⁻ decays with the ATLAS detector — Dominik Duda, •Florian Kiwit, Sandra Kortner, and Hubert Kroha — Max-Planck-Institut für Physik

The identification of W, Z and Higgs bosons with large transverse momenta is crucial in many searches for new heavy resonances. Thus far, the development of algorithms for the tagging of boosted bosons focuses on the reconstruction and identification of hadronic boson decays, while no dedicated algorithm to identify boosted $Z \rightarrow e^+e^-$ decays exists. The performance of the standard electron reconstruction and identification algorithms degrades with decreasing angular separation between the e^+e^- pairs and will eventually vanish once the angular separation between the e^+e^- pairs is too small to construct individual clusters in the calorimeter. To improve the reconstruction and identification of such highly boosted $Z \rightarrow e^+e^-$ decays, a dedicated algorithm for $Z \rightarrow e^+e^-$ tagging is being developed using a deep neural network. This talk presents first results of this development.

T 64.6 Wed 17:30 T-H23

Search for charged Higgs bosons in $H^+ \to Wh \to l\nu bb$ decays with the AT-LAS detector — DOMINIK DUDA, •SIMON GREWE, SANDRA KORTNER, and HUBERT KROHA — Max-Planck-Institut für Physik, München

Various theories predicting an extended Higgs sector predict also the existence of at least one set of charged Higgs bosons. The main production mode of these new particles depends on their mass. For charged Higgs boson masses larger than the sum of the top and the bottom quark mass, the dominant production mode is expected to be in association with a top quark and a bottom quark (tbH^+).

In the alignment limit of the two-Higgs-Doublet Model (2HDM), heavy charged Higgs bosons with m(H⁺)>m(t)+m(b) decay almost exclusively via $H^+ \rightarrow tb$. However, in other models such as the Next-to-two-Higgs-Doublet Model (N2HDM), the three-Higgs-Doublet model (3HDM) or in Higgs triplet models (e.g. Georgi-Machacek model), significant branching ratios for $H^+ \rightarrow W^+h$ are possible. The latter decay mode has so far been covered neither by ATLAS nor CMS.

We present first studies on the search for $H^+ \to W^+ h \to \ell \nu b b$ decays in final states with the resolved topology containing five or more jets, one charged lepton and missing transverse momentum. A multiclass classifier is used to separate the semileptonic $H^+ \to l \nu b b$ and fully hadronic $H^+ \to q q b b$ decay modes from the dominant background processes. The reconstruction of the charged Higgs boson decay is performed using boosted decision trees (BDTs).

T 64.7 Wed 17:45 T-H23 Status of the FASER Experiment — •Markus Prim and Florian Bern-

LOCHNER — Rheinische Friedrich-Wilhelms-Universität Bonn

FASER, or the Forward Search Experiment, is a new experiment at CERN designed to complement the LHC's ongoing physics programme, extending its discovery potential to light and weakly-interacting particles that may be produced copiously at the LHC in the far-forward region. New particles targeted by FASER, such as long-lived dark photons or dark scalars, are characterised by a signature with two oppositely-charged tracks or two photons in the multi-TeV range that emanate from a common vertex inside the detector. The experiment is composed of a silicon-strip tracking-based spectrometer using three dipole magnets with a 20-cm aperture, supplemented by four scintillator stations and an electromagnetic calorimeter to allow for energy measurements. The full detector was successfully installed in March 2021 in an LHC side-tunnel 480 meters downstream from the interaction point in the ATLAS detector. FASER is planned to be operational for the upcoming LHC Run 3. We will discuss the physics reach of FASER, present the current status of the experiment and results from the ongoing in situ commissioning.

T 64.8 Wed 18:00 T-H23

Results from First Simulation Studies for a Dark Photon Search Experiment at the ELSA Electron Accelerator — Philip Bechtle, Klaus Desch, Oliver FREYERMUTH, MATTHIAS HAMER, •JAN-ERIC HEINRICHS, and MARTIN SCHÜR-MANN — Rheinische Friedrich-Wilhelms-Universität Bonn

The true nature of Dark Matter (DM) has long been of interest for scientists worldwide. Previous searches have so far been unsuccessful in finding proposed DM particles. A promising and not well explored parameter space of light DM particles up to 1 GeV remains to be subjected to intense experimental testing. Mainly two approaches are investigated by the community, namely beam dump and fixed targets experiments.

This talk highlights the future prospects of a fixed target experiment aimed at finding evidence for a dark sector, which couples to the standard model through a dark photon. The underlying theory and the resulting experimental challenges and strategy will be explained. The possibility of building a corresponding experiment at the ELSA electron accelerator in Bonn is highlighted. This talk covers first steps towards a Geant 4 simulation. Special emphasis will be put on some first simulation results concerning the layout and technology for an electromagnetic calorimeter based on requirements on radiation hardness and response times.

T 64.9 Wed 18:15 T-H23

Neutrino and Muon induced background studies - SHiP experiment — •ANUPAMA REGHUNATH, HEIKO LACKER, ANDREW CONABOY, and JAKOB SCHMIDT for the SBT-Collaboration — Institut für Physik, Humboldt-Universität zu Berlin, Berlin, Germany

The Search for Hidden Particles (SHIP) experiment is proposed to be constructed at a dedicated beam-dump facility at the CERN Super Proton Synchrotron (SPS) aiming to search for new feebly interacting particles generated in the decay of heavy flavour hadrons or through interactions of photons inside the beam-dump target. During a period of five years, SHIP aims to collect data from $2 \cdot 10^{20}$ 400 GeV/c protons on target. The experimental design is optimised to maximise the production of charm and beauty mesons with zero background events to have the best sensitivity to hidden-sector particles. Simulation studies with emphasis on the neutrino and muon induced background arising from inelastic interactions and its rejection using kinematical and topological requirements as well as information dedicated veto detectors, such as the Surround Background Tagger, will be discussed.

T 65: Silicon Strip Detectors

Time: Wednesday 16:15–18:15

T 65.1 Wed 16:15 T-H24

Optical inspection of silicon strip sensors for the Phase-2 Upgrade of the CMS Tracker — ANNA BECKER¹, CHRISTIAN DZIWOK², LUTZ FELD¹, PATRICK JURASCHITZ¹, KATJA KLEIN¹, MARTIN LIPINSKI¹, ALEXANDER PAULS¹, OLIVER POOTH², •NICOLAS RÖWERT¹, and TIM ZIEMONS² — ¹I. Physikalisches Institut B, RWTH Aachen University, Germany — ²RWTH Aachen University - Physics Institute III B, Aachen, Germany

During the Long Shutdown 3 the LHC will be upgraded to the High Luminosity LHC with a planned instantaneous luminosity of at least $5 \cdot 10^{34}$ cm⁻²s⁻¹. For this purpose the current strip tracker of the CMS experiment will be entirely replaced by a new system consisting of innovative silicon detector modules. RWTH Aachen University as one of the assembly centers is responsible for manufacturing around 1000 so-called 2S modules, which are equipped with two vertically stacked strip sensors having an active area of $10 \cdot 10$ cm². During module production the inspection of the silicon sensors is crucial in order to detect damage or contamination caused by the assembly process which

mega pixel camera equipped with a macro lens and an xy-stage. This talk presents the setup, the approach to detect damages, and results obtained from images taken in the course of the assembly of a 2S module prototype featuring preproduction sensors.

could seriously affect the successful operation of the module. For this purpose an automated optical inspection setup has been developed. It consists of a 24

T 65.2 Wed 16:30 T-H24

Location: T-H24

Testing results of the latest Service Hybrid prototypes for CMS silicon strip modules — Christian Dziwok², Lutz Feld¹, Katja Klein¹, Martin Lipinski¹, Daniel Louis¹, •Alexander Pauls¹, Oliver Pooth², Matej Repik¹, Nicolas Röwert¹, Michael Wlochal¹, and Tim Ziemons² — ¹1. Physikalisches Institut B, RWTH Aachen — ²3. Physikalisches Institut B, RWTH Aachen
The CMS Collaboration is developing silicon strip modules for the second phase of the CMS tracker upgrade. This upgrade will enable the CMS experiment to utilize the high luminosity provided by the future HL-LHC. The modules' Service Hybrids are responsible for the sensor bias voltage and low voltage distribution on the module and the data transmission via optical links to the back-end electronics. Multiple batches of Service Hybrid prototypes have been produced. The latest use the final ASIC set with its most recent available versions and materials and geometries as foreseen in the detector. The prototyping campaign and present measurements of the hybrid performance are summarized. The measurements were performed with setups similar to the foreseen production test system, which is also presented.

T 65.3 Wed 16:45 T-H24

Vermessung von Streifenmodulen für das CMS Phase-2 Tracker Upgrade bei Betriebstemperatur — Christian Dziwok², Lutz Feld¹, •Patrick JURASCHITZ¹, KATJA KLEIN¹, MARTIN LIPINSKI¹, ALEXANDER PAULS¹, OLIVER POOTH², NICOLAS RÖWERT¹, MICHAEL WLOCHAL¹ und TIM ZIEMONS² — ¹1. Physikalisches Institut B, RWTH Aachen — ²RWTH Aachen University - Physics Institute III B, Aachen, Germany

Im Zuge des HL-LHC Upgrades im Long-Shutdown 3 (2025-2027) muss der Silizium-Spurdetektor des CMS Experiments ausgetauscht werden. Das Phase-2 Outer Tracker Upgrade umfasst neue Module, welche aus zwei übereinander parallel ausgerichteten Sensoren bestehen. Diese ermöglichen es durch die Analyse der Teilchendurchgänge bereits im Auslesechip eine Abschätzung des transversalen Impulses zu erhalten. Die relative Ausrichtung der Sensoren zueinander beeinflusst dabei maßgeblich die Qualität dieser Abschätzung. Die Alignierung wird mittels doppelseitiger Metrology gemessen, wobei das Modul von zwei Kameras betrachtet wird.

Der Phase-2 Spurdetektor soll bei einer Temperatur von -35°C betrieben werden. Daher ist das Verhalten der Module im Hinblick auf die Ausrichtung bei Betriebstemperatur von besonderem Interesse.

Es wird ein Verfahren präsentiert, welches es ermöglicht, Module mit einer Kühlmaschine zu kühlen und im Anschluss zu vermessen. Die Ergebnisse werden mit Messungen bei Raumtemperatur verglichen, um mögliche Auswirkungen der Temperatur auf die Ausrichtung der Sensoren zueinander nachzuvollziehen.

T 65.4 Wed 17:00 T-H24

Stress testing optical readout components for CMS 2S modules — •CHRISTIAN DZIWOK², LUTZ FELD¹, KATJA KLEIN¹, MARTIN LIPINSKI¹, ALEXANDER PAULS¹, OLIVER POOTH², NICOLAS RÖWERT¹, and TIM ZIEMONS² — ¹I. Physikalisches Institut B, RWTH Aachen University, Aachen, Germany — ²RWTH Aachen University - Physics Institute III B, Aachen, Germany

For the upcoming CMS Phase-2 Outer Tracker upgrade, new detector modules will be installed. There are two general types of modules, one consisting of two co-planar silicon strip sensors (2S) and one of a macro pixel and a strip sensor (PS). The communication and the auxiliary support are supplied by a so called SErvice Hybrid (SEH) in case of a 2S module. At the RWTH Aachen University the SEHs are qualified regarding power and communication stability in a so called test board setup, where the SEHs will undergo additional thermal cycling while being tested. This talk will focus on the data tests of this setup.

T 65.5 Wed 17:15 T-H24

Functional Testing of Silicon Sensor Modules for the CMS Experiment using Infrared LED Arrays — •ROLAND KOPPENHÖFER¹, TOBIAS BARVICH¹, BERND BERGER¹, JUSTUS BRAACH², ALEXANDER DIERLAMM¹, ULRICH HUSEMANN¹, MARKUS KLUTE¹, GANI KÖSKER¹, STEFAN MAIER¹, THOMAS MÜLLER¹, MAR-IUS NEUFELD¹, HANS JÜRGEN SIMONIS¹, PIA STECK¹, and FLORIAN WITTIG¹ — ¹Institute of Experimental Particle Physics (ETP), Karlsruhe Institute for Technology (KIT) — ²CERN

In the context of the Phase-2 Upgrade of the CMS experiment, the complete CMS tracker will be replaced. The new CMS Outer Tracker will consist of two types of silicon sensor modules (PS and 2S modules). These modules are built and tested for their full functionality at different assembly centers. The ETP is one of the assembly centers for 2S modules. A dedicated test station for the electrical characterization of 2S modules has been designed and built at ETP. Using

infrared LEDs it is possible to generate charge in the silicon sensors and test every module channel. This talk will present the test station developed at ETP and summarize the functional test results of the newest 2S module prototypes.

T 65.6 Wed 17:30 T-H24

Assembly and Test Procedures of 2S modules for the future Outer Tracker of the Phase-2 Upgrade of the CMS Experiment — • STEFAN MAIER, TOBIAS BARvich, Bernd Berger, Alexander Dierlamm, Alexander Droll, Umut Eli-CABUK, JAN-OLE MÜLLER-GOSEWISCH, ULRICH HUSEMANN, ROLAND KOPPEN-HÖFER, MARKUS KLUTE, GANI KÖSKER, THOMAS MÜLLER, MARIUS NEUFELD, HANS JÜRGEN SIMONIS, PIA STECK, LEA STOCKMEIER und FLORIAN WITTIG -Institute of Experimental Particle Physics, Karlsruhe Institute of Technology In preparation for the High Luminosity LHC, the entire tracker of the CMS experiment will be exchanged within the Phase-2 Upgrade until 2027. The new outer tracker will be made of approximately 13000 silicon sensor modules called 2S modules (consisting of two parallel silicon strip sensors) and PS modules (one pixel and one strip sensor combined in a module). These modules provide tracking information to the Level-1 trigger by correlating the hit information of both sensor layers and, thus, allowing to discriminate charged particles by their transverse momentum. To guarantee successful operation of the CMS detector at the HL-LHC, the production of the outer tracker modules has to fulfil strict requirements. The talk will summarize the various assembly and test concepts for the large scale production of 2S modules and will present the latest module prototypes.

T 65.7 Wed 17:45 T-H24

Integration Tests with 2S Module Prototypes for the Phase-2 Upgrade of the CMS Outer Tracker — •Lea Stockmeier, Tobias Barvich, Bernd Berger, Alexander Dierlamm, Alexander Droll, Umut Elicabuk, Ulrich Huse-Mann, Markus Klute, Roland Koppenhöfer, Stefan Maier, Thomas Müller, Jan-Ole Müller-Gosewisch, Marius Neufeld, Hans Jürgen Simonis, Gani Kösker, Pia Steck, and Florian Wittig — Institut für Experimentelle Teilchenphysik (ETP), Karlsruher Institut für Technologie (KIT)

To deal with the increased luminosity of the HL-LHC, the CMS experiment will be upgraded until 2027. During this Phase-2 Upgrade the CMS Outer Tracker will be equipped with modules each consisting of two silicon sensors. Depending on the position in the tracker, these silicon sensors are pixel or strip sensors. The modules with two strip sensors are called 2S modules. In the barrel region, they are placed on mechanical structures called ladders. A fully equipped ladder contains twelve modules.

Within the prototyping phase of the 2S modules, laboratory and integration tests are performed. This talk summarizes the first integration test on a ladder performed with four modules at Institut Pluridisciplinaire Hubert CURIEN (IPHC) in cooperation with KIT and laboratory tests performed at KIT.

T 65.8 Wed 18:00 T-H24

Location: T-H25

Commissioning of a Burn-In Setup for PS and 2S Detector Modules for the Upgrade of the CMS Outer Tracker — •ANA VENTURA BARROSO, KATERINA LIPKA, and PAUL SCHÜTZE — DESY, Hamburg, Germany

The high luminosity LHC Upgrade will increase the instantaneous luminosity by a factor of five. The CMS detector will be upgraded in the so called Phase-2 Upgrade in order to meet the new requirements, among others the level of radiation tolerance and coping with larger pileup and thus higher data rates, as well as to add triggering capabilities. The entire silicon tracker will be replaced. The Outer Tracker (OT), consisting of macro-pixel and strip detectors, will be based on silicon modules that must operate at low temperatures (-33° C) due to the exposition at high radiation levels. The probability for defective electronic components to fail is higher after few hours of operation. Moreover, temperature cycles can induce mechanical stress. Therefore a burn-in procedure as well as thorough quality control is needed to ensure the correct operation of each of the OT modules before installation.

For this, a burn-in system is being commissioned at DESY. This setup will perform thermal cycles from room to operation temperature and key measurements to ensure the good performance of the modules. In this talk, the status of the DESY burn-in setup as well as first tests will be presented.

T 66: Semiconductor Detectors: Radiation Hardness, new Materials and Concepts 2

Time: Wednesday 16:15-18:30

T 66.1 Wed 16:15 T-H25

Study of the self-heating in SiPMs — •CARMEN VICTORIA VILLALBA PETRO, ERIKA GARUTTI, ROBERT KLANNER, STEPHAN MARTENS, and JÖRN SCHWANDT — Universität Hamburg, Luruper Chaussee 149, 22761 Hamburg, Deutschland. A dramatic increase of the dark current is the main effect of radiation damage in SiPMs. The power dissipated, if not properly cooled, heats the SiPM, whose performance parameters depend on temperature. Therefore, the knowledge of the SiPM temperature is necessary to understand the changes of its parameters with irradiation.

The self-heating studies were performed with a KETEK SiPM, $15x15 \text{ mm}^2$ pixel size, mounted on an Al_2O_3 substrate 0.6 mm thick, which was either directly connected to the temperature controlled chuck of a probe station, or through layers of material with well-known thermal resistance. The SiPM was illuminated by a LED operated in DC-mode. SiPM current was measured at dif-

ferent voltages, LED currents, chuck temperatures, and thermal resistivities for a number of measurement cycles. The data are used to determine the steady-state temperature as a function of dissipated power and thermal resistance, as well as the time dependencies for heating and cooling. This information could be used to correct the parameters determined for radiation-damaged SiPM for the effects of self-heating.

The presentation describes the experimental layout, the data taking, the analysis methods, the results obtained and a comparison to thermal simulations. The application of the method for the study of radiation damaged SiPMs and its use in actual experiments is discussed.

T 66.2 Wed 16:30 T-H25

PeakOTron: A Python Module for Fitting Charge Spectra of Silicon Photomultipliers — •JACK ROLPH, ERIKA GARUTTI, and JOERN SCHWANDT — Institute for Experimental Physics, University of Hamburg, Luruper Chaussee 149, 22761 Hamburg, Germany

An automated software method is developed to characterise the pulse-height spectra of SiPM obtained in the dark or response to low-intensity light illumination. The method is based on a fit utilising a published SiPM response model. It aims to provide SiPM performance parameters for single measurements, mass production characterisation, or high granular detector calibration. The fit considers dark count rate, the average number of detected photons, crosstalk and after-pulsing, electronics noise and gain fluctuations. Due to the sensitivity of the fit to the initial values assigned to these parameters, the software must perform careful estimation before fitting the model to data, which has been designed to be fast, accurate and robust to fluctuations. First, the model*s accuracy is validated against simulation and then tested to experimental data from various SiPMs.

T 66.3 Wed 16:45 T-H25

Silicon Photomultiplier characterization for a space-borne optical instrument in Low Earth Orbit — •Lucas $FINAZZI^{1,2,4}$, FEDERICO GOLMAR^{2,4}, Andreas Haungs¹, Thomas Huber¹, and Federico Izraelevitch^{2,3,4} – ¹Institut für Astroteilchenphysik (IAP), Karlsruher Institut für Technologie (KIT), Germany – ²Universidad de San Martín (UNSAM), Argentina ³Comisión Nacional de Energía Atómica (CNEA), Argentina — ⁴Consejo Nacional de Investigaciones Científicas y Técnicas (CONICET), Argentina

The S4 (Space-borne SiPM-based Single-photon Sensor) is an instrument designed for single-photon detection in Low Earth Orbit (LEO) using Silicon Photomultiplier (SiPM) technology. The instrument, currently under developed by the LabOSat group (UNSAM), is expected to launch on Q2 2023 as a secondary payload of a commercial Earth-observation satellite. Beside the increase of the Technology Readiness Level (TRL) of SiPM sensors for space applications, the S4 instrument will be a pathfinder for Ground-LEO single-photon communications and fundamental Physics experiments that require single-photon detection.

In this talk we will describe the S4 instrument, focusing on the Analog-Front End (AFE) and SiPMs subsystem. Particularly, we will present the characterization performed at the SPOCK (Single PhOton Calibration stand at KIT) laboratory at the Institute of Astroparticle Physics (KIT), including studies on the performance of the AFE board and sensors at different temperatures, SiPM overvoltages and illumination conditions.

T 66.4 Wed 17:00 T-H25

Timing measurements of a tip avalanche photodiode (TAPD) - A nearinfrared enhanced silicon photomultiplier based on spherical depletion -•WOLFGANG SCHMAILZL^{1,3}, JONATHAN PREITNACHER¹, ERIKA GARUTTI², and WALTER HANSCH 1 — 1 Bundeswehr University Munich, Neubiberg, Germany -²University of Hamburg, Hamburg, Germany – ³Broadcom Inc., Regensburg, Germany

The gain of attention in LiDAR technology also pushed the development of sensors for time of flight measurements. Such a measurement requires a fast and very sensitive receiver. Often a wavelength in the near-infrared (NIR) region is chosen if the emitter is part of the system while a high sensitivity over a broad range is an advantage for deviating applications. Silicon photomultiplier (SiPM) with enhanced sensitivity in the NIR region provide a good combination of performance and cost-efficiency. In our investigations we simulated and fabricated spherical junctions for the single-photon avalanche diodes and a photodetection efficiency of 22% at 905nm was achieved. We present timing measurements of this new NIR SiPM to provide an overview and to put the performance into the context of possible applications with very low and high light intensities. Compared to their blue enhanced counterparts, the single photon time resolution of these devices is lower where some of this performance loss is related to the required depletion depth for longer wavelengths while the current design is still not at its limit. We want to outline some of these intrinsic limits and show where improvements can be made.

T 66.5 Wed 17:15 T-H25

Characterisation and Radiation Hardness of Tip-Avalanche PhotoDiodes - •Julius Römer¹, Erika Garutti¹, Wolfgang Schmailzl^{2,3}, Jörn Schwandt¹, and Stephan Martens¹ - ¹Universität Hamburg, Luruper

Chaussee 149, 22761 Hamburg, Deutschland – ²Bundeswehr Universität München, Neubiberg, Deutschland — ³Broadcom Inc., Regensburg, Deutschland

Silicon Photomultipliers (SiPM) are the photon detectors of choice for many applications. Development of SiPMs with conventional single photon avalanche diode (SPAD) built around a planar pn-junction reaches a trade-off between photon detection efficiency (PDE) and dynamic range. A challenge in designing efficient red-sensitive SiPMs lies in the requirement for large depletion depths.

A novel design design featuring a quasi-spherical pn-junction called Tip Avalanche Photodiode (TAPD) tackles both problems. For a pixel pitch of 15 $\mu \mathrm{m},$ the SiPM prototype reaches a PDE of 73 % at 600 nm and 22 % at 900 nm, a high dynamic range, as well as a recovery time below 4 ns. The aim of this study is to characterize this novel type of SiPM. In particular the question of radiation hardness of TAPD is addressed.

After irradiation with reactor neutrons with fluences up to $1 \cdot 10^{12} \mbox{ cm}^{-2} \ 1 \mbox{ MeV}$ neutron equivalent, characterisation with I - V and C - V - f measurements show that despite the increased bulk volume, TAPD-SiPMs show similar loss of response and increase in dark count rate (DCR) with irradiation compared to planar SiPM devices, enhancing the usability in high energy particle experiments.

T 66.6 Wed 17:30 T-H25

Investigation of the Time Resolution of LGADs and 3D Sensors - •LEENA DIEHL, MONTY KING, DENNIS SPERLICH, ULRICH PARZEFALL, MARC HAUSER, and Christina Schwemmbauer — Universität Freiburg

Future collider experiments as the high-luminosity LHC or the FCC will increase the demands of the detectors used for tracking. Sensors will not only face fluences of up to $1 \cdot 10^{17} n_{eq}/\text{cm}^2$, but also high pile-up scenarios. Thus sensors are needed which have a high radiation tolerance, but also an excellent time resolution while still providing a good spatial resolution. Currently Low Gain Avalanche Diodes (LGADs) are the prime candidate when it comes to timing, reaching a resolution of below 30 ps. However, 3D sensors are promising candidates as well, as they have not only a good time resolution but also a superior radiation hardness. In order to investigate the time resolution of both LGADs and 3D sensors thoroughly, timing measurements were performed using either a beta-source or a laser with infrared wavelength. The timing-TCT measurements allow to measure the position-dependence of the time resolution, which is interesting especially for the 3D-sensors, where the time walk is an important component of the resolution. This talk will present some initial results of the measurements, including the calibration of the timing-TCT setup and first measurements with were 3D-sensors as well as LGADs.

T 66.7 Wed 17:45 T-H25

Time resolution comparison between LGADs and 3D silicon detectors -•Montague King, Christina Schwemmbauer, Leena Diehl, Dennis Sper-LICH, MARC HAUSER, and ULRICH PARZEFALL — Albert-Ludwigs Universität, Freiburg, Germany

For the planned high luminosity upgrade of the LHC it is vital to develop detector systems with both excellent spatial and timing resolution in the inner tracking layers to be able to distinguish separate collisions in the same bunch crossing. This has to be achieved even after severe irradiation with expected fluences of up to 2.6 x 10^{16} n_{eq}/cm².

It has been shown that a timing resolution of 30 ps can be achieved with Low Gain Avalanche Diodes (LGADs). However, these devices have drawbacks, especially regarding spatial resolution, effective fill factor and radiation hardness. A second detector type, 3D-detectors with electrodes extending deeply into the silicon bulk, are expected to be able to achieve such a high timing resolution while having a superior spatial resolution and radiation hardness.

In this talk, the calibration of a setup developed specifically for fast timing measurements at the University of Freiburg will be presented. This setup utilises a ⁹⁰Sr electron source along with a reference sensor and scintillator to measure fluctuations in the time between signals from reference and tested sensor. Furthermore, first results measured for LGADs before and after irradiation as well as unirradiated 3D sensors will be discussed.

T 66.8 Wed 18:00 T-H25

Boron removal effect in silicon sensors induced by 6 MeV electrons -•Chuan Liao, Eckhart Fretwurst, Erika Garutti, and Joern schwandt - Universität Hamburg, Luruper Chaussee 149, 22761 Hamburg, Deutschland In the frame of the CERN_RD50 collaboration, the acceptor removal effect is investigated in p-type material, used for example in pixel sensors or Low Gain Avalanche Detectors (LGADs). The suspected cause is the displacement of substitutional Boron (Bs), being negatively charged, by incident particles or other recoil atoms into an interstitial position (Bi). This is followed by Bi migration and being captured by Oxygen atoms and forming complex defects of interstitial Boron and interstitial Oxygen (BiOi) with a positive charge. This is the boron removal effect. For lower radiation fluence, this has one main consequence: The maximum electric field at a given reverse bias will decrease, causing e.g. a decrease of the LGAD gain. In this presentation, the Thermally Stimulated Techniques including TS-Current (TSC) and TS-Capacitance (TS-Cap) have been used to study the properties of the radiation-induced BiOi defect complex by 6 MeV electrons. Two different types of diodes manufactured on p-types epitaxial-(EPI) and Czochralski(CZ) silicon with a resistivity of about 10 Wcm were irradiated with fluence values in the range between 1 x 1015 cm-2 and 6 x 1015 cm-2. The results will be presented and compared with those gained from sensors exposed to 23 GeV protons.

T 66.9 Wed 18:15 T-H25 Physics case for a forward timing disc covering 3< η <4 of the CMS detector at **HL-LHC** – •Anna Albrecht¹, Anna Benecke², Andreas Hinzmann¹, Ben KILMINSTER³, and STEFANOS LEONTSINIS³ - ¹Universität Hamburg - ²UC Louvain — ³University of Zurich

Time: Wednesday 16:15-18:40

Group Report T 67.1 Wed 16:15 T-H26

DT upgrade activities during LHC long shutdown II and readiness status for Run 3 — • Archana Sharma, Thomas Hebbeker, Kerstin Hoepfner, Hans REITHLER, MARKUS MERSCHMEYER, and DMITRY ELISEEV — III. Physikalisches Institut A, RWTH Aachen University

After delivering an integrated luminosity of more than 160/fb until the end of Run 2, at the beginning of 2019, LHC was shut down until the end of 2021 (LS2) in order to get its accelerator-chain and detectors upgraded for the HL-LHC phase, expected to deliver an instantaneous luminosity 5 times higher with respect to the present value. During this LS2, the Compact Muon Solenoid (CMS) experiment worked to upgrade its electronics and detector performance to improve the data taking and a precise reconstruction of all the particles in high pile-up conditions of HL-LHC. Drift Tubes (DT) detectors equip the CMS muon system barrel region, serving both as offline tracking and triggering devices. An upgrade of the current readout and trigger electronics is also planned in order to withstand event rates and integrated doses far beyond the initial design specification expected in HL-LHC. During LS2, prototypes of the new electronics were installed in four DT chambers with the same azimuthal acceptance, instrumenting a demonstrator of the HL-LHC DT upgrade (DT slice-test). This report briefly summarises the commissioning activities performed during LS2, along with the status of the slice-test and its performance with cosmic -ray events.

T 67.2 Wed 16:35 T-H26

A tester tool for the new read-out electronics of the MDT detectors of the AT-LAS muon spectrometer in the Phase-II upgrade — •THORBEN SWIRSKI, GIA Кногіаиli, and Raimund Ströhmer — Universität Würzburg

To make the ATLAS monitored drift tube detectors (MDT) ready for the new high rate environment in the High-Luminosity LHC, new read-out electronics, including new mezzanine cards, are being developed. These mezzanine cards feature a triggerless read-out mode to allow the MDT stations to be used in the trigger decisions. A tester-tool has been designed for mass quality testing of the new mezzanine cards. The tester-hardware is ready since summer 2021 and is now being programmed to allow for the testing of all functionalities of the mezzanine cards. The work allows the use of the tester tool not only for the mass testing of already produced cards, but also adds valuable insights during the ongoing prototyping effort. This talk will present the tester-tool, give an overview over the test-regime, show how the gathered data will be archived in a dedicated database and present examples of how the use of the tester tool has helped the development of the mezzanine cards.

Group Report

T 67.3 Wed 16:50 T-H26 The small-diameter Drift Tube (sMDT) Chambers for ATLAS at High-Luminosity LHC and for Future Colliders - DAVIDE CIERI, GREGOR EBER-WEIN, OLIVER KORTNER, HUBERT KROHA, PATRICK RIECK, MARIAN RENDEL, and •ELENA VOEVODINA — Max-Planck-Institut für Physik, München, Deutschland

The small-diameter Drift Tube (sMDT) detector technology with 15 mm tube diameter has proven to be an excellent candidate for precision muon tracking detectors in experiments at future hadron colliders like HL-LHC and FCC-hh where unprecedentedly high background rate capabilities are required. The upgrades of the inner barrel layer of the ATLAS Muon Spectrometer in the LS2 (2019-2021) and LS3 (2026-2028) shutdowns of the LHC machine use combinations of the sMDT detectors and RPC trigger chambers. We present the test results of the performance of the sMDT chambers under construction for the operation at HL-LHC and the measurements of the behavior of the detectors and their new readout electronics under very high background irradiation rates like at FCC-hh performed at the CERN Gamma Irradiation Facility (GIF++).

The LHC will be upgraded to a collider with 10 times higher luminosity, the high luminosity (HL)-LHC. One main challenge arising from the upcoming high luminosity, is the large amount of interactions that occur in one proton-proton bunch crossing, and therefore the separation of the interaction of interest from the additional ones (pileup). The insertion of a new timing layer in the upgraded CMS experiment is planned, to use timing as an additional discrimination variable between signal and pileup.

One interesting channel to probe at the HL-LHC is vector boson fusion (VBF) Higgs pair production that has two characteristic jets in the forward region of the detector. The separation of this signal from pileup is extremely challenging. In this physics case study, the performance of an extension of the timing layer from $\eta < 3$ to $\eta < 4$ is estimated by using the Delphes simulation software.

T 67: Myon Detectors

Location: T-H26

T 67.4 Wed 17:10 T-H26

Quality Control in the Construction of new small-diameter Muon Drift Tube (sMDT) Chambers for the ATLAS Muon Spectrometer at the HL-LHC -•DANIEL BUCHIN, MARIAN RENDEL, ALICE REED, PATRICK RIECK, OLIVER KOктиев, and Hubert Kroha — Max-Planck-Institut für Physik, München

In order to improve the muon trigger efficiency and the rate capability of the AT-LAS muon detectors for operation at the high luminosity upgrade of the Large Hadron Collider (HL-LHC), the Monitored Drift Tube (MDT) tracking chambers in the inner barrel layer of the ATLAS Muon Spectrometer will be replaced by small-diameter Muon Drift Tube (sMDT) chambers integrated with new thingap RPC trigger chambers. The sMDT chambers fit, together with the RPCs, into the very tight available space and provide an order of magnitude higher background rate capability compared to the current detectors.

The sMDT chambers are in serial production since January 2021. In the talk, the measures implemented to assure the quality of the chambers will be explained, starting from the results of the tests and inspections each individual assembled drift tube is undergoing. The dedicated quality control database and monitoring web interface will also be discussed.

T 67.5 Wed 17:25 T-H26

Construction of small-diameter Monitored Drift Tube (sMDT) chambers for the ATLAS Muonspectrometer at the HL-LHC — •MARIAN RENDEL, DANIEL BUCHIN, ALICE REED, OLIVER KORTNER, HUBERT KROHA, PATRICK RIECK, and ELENA VOEVODINA - Max-Planck-Institut für Physik (Werner-Heisenberg-Institut)

For the high luminosity upgrade of the Large Hadron Collider, the Monitored Drift Tube (MDT) tracking chambers in the inner barrel layer of the ATLAS muon spectrometer will be replaced by small-diameter Monitored Drift Tube (sMDT) chambers integrated with new thin-gap RPC trigger chambers. The sMDT chambers fit, together with the RPCs, into the very tight available space and provide an order of magnitude higher background rate capability compared to the current MDT chambers.

Since January 2021 the sMDT chamber serial production is proceeding at a steady pace of one chamber every two weeks as planned. In this talk, the sMDT chamber design and the construction procedures will be explained as well as the technology transfer to the second production site at the University of Michigan.

T 67.6 Wed 17:40 T-H26

Measuring the geometry of the new small-diameter Monitored Drift Tube (sMDT) chambers constructed for the HL-LHC upgrade of the ATLAS Muonspectrometer — •Alice Reed, Daniel Buchin, Marian Rendel, Patrick RIECK, ELENA VOEVODINA, OLIVER KORTNER, and HUBERT KROHA - Max-Planck-Institut für Physik (Werner-Heisenberg-Institut), München

In order to improve the muon trigger efficiency and the rate capability of the AT-LAS muon detectors for operation at the high-luminosity upgrade of the Large Hadron Collider (HL-LHC), the Monitored Drift Tube (MDT) chambers in the inner barrel layer of the ATLAS Muon Spectrometer will be replaced by smalldiameter Muon Drift Tube (sMDT) chambers integrated with new thin-gap RPC trigger chambers. The sMDT chambers fit, together with the RPCs, into the very tight available space and provide provide an order of magnitude higher background rate capability compared to the current detectors.

The sMDT chambers have to provide a sense wire positioning accuracy of better than 20 μ m and this requires the geometry of the chambers to be measured with a high precision. The measurement procedure and results are discussed for the chambers in serial production.

T 67.7 Wed 17:55 T-H26 Fabrication of Resitive Plate Chambers — OLIVER KORTNER, HUBERT KROHA, DANIEL SOYK, and •TIMUR TURKOVIC - Max-Planck-Institut für Physik, Föhringer Ring 6, 80805 München, Germany

Resistive plate chambers (RPCs) with electrodes of high-pressure phenolic laminate (HPL) and small gas gap widths down to 1 mm provide large area tracking at relatively low cost in combination with high rate capability and fast response with excellent time resolution of better than 500 ps. They can be operated up to y background count rates of up to 10 kHz/cm2 which is five times the maximum rate these RPCs will encounter in the innermost layer of the barrel muon spectrometer of the ATLAS detector where they will be installed in the phase-II upgrade for the HL-LHC operation. A cost-effective production procedure which is compliant with industrial techniques was worked out and tested by a production of prototype thin-gap RPCs in the laboratory. Cosmic ray muons were used to check the proper functioning of these prototypes. The new production procedure has been transfered to several companies for the production of test samples. We will report about the RPC production in the laboratory and the technology transfer to industry.

T 67.8 Wed 18:10 T-H26

The CMS Muon upgrade and the commissioning of the first GEM station -•Francesco Ivone, Thomas Hebbeker, Kerstin Hoepfner, Giovanni Mo-CELLIN, and SHAWN ZALESKI — III. Physikalisches Institut A, RWTH Aachen University

The LHC will undergo a major upgrade to deliver ~10 times more proton-proton collisions in the next two decades, which has been named High-Luminosity LHC (HL-LHC). To cope with the higher event rates and to improve the trigger capabilities in the forward region, the Compact Muon Solenoid (CMS) experiment will undergo several upgrades. These include the installation of an additional set of muon detectors based on the Gas Electron Multipliers (GEM) technology. The triple-GEM detectors named GE1/1 have already been installed during

Time: Wednesday 16:15–18:15

T 68.1 Wed 16:15 T-H27

Design and construction of a neutron imaging detector based on a neutron sensitive MCP on Timepix3 ASICs — •SAIME GÜRBÜZ¹, MARKUS GRUBER¹, Jochen Kaminski¹, Markus Köhli², Michael Lupberger¹, Divya Pal¹, LAURA RODRÍGUEZ GÓMEZ¹, and KLAUS DESCH¹ – ¹Physikalisches Institut, Rheinsiche Friedrich-Wilhlems-Universität Bonn, Bonn – ²Physikalisches Institut, Ruprecht-Karls-Universität Heidelberg, Heidelberg

The non-invasive method of neutron imaging has applications from medicine to engineering. With its unique capabilities, we can image elements that can not be distinguished via X-ravs.

At the University of Bonn, we have been developing neutron detectors with different configurations and specifications. One of the neutron imaging detectors that has been studied is based on neutron sensitive Microchannel plate (nMCP) in combination with a Timepix3 ASIC. Previous studies show that neutron sensitive 10 B and Gd enriched MCPs have up to 10 μ m spatial resolution and neutron detection efficiency of up to 70%. The readout will be achieved using four Timepix3 ASICs with 55 μ m pixels covering (28 x 28) mm² active area. The readout has a timing resolution of up to 1.5 ns.

By combining the efficiency of the MCP with the good time and spatial resolution of Timepix3, we aim to achieve excellent instrument performance in both domains. Such a detector can be used for many neutron imaging applications such as neutron tomography or time-of-flight imaging. In this talk, the current status as well as the principles of the detector will be presented.

T 68.2 Wed 16:30 T-H27

Wavelength-shifter coated polystyrene as a low-cost plastic scintillator detector — Alessia Brignoli¹, Andrew Conaboy¹, •Doramas Jimeno^{1,2}, Неіко Lacker¹, Christian Scharf¹, and Jakob Schmidt¹ – ¹Humboldt-Universität zu Berlin, Berlin, Germany — ²Universitat de Barcelona, Barcelona, Spain

Plastic scintillator detectors are widely used in particle physics for detecting charged particles crossing the scintillating material, converting the excitation energy into fluorescence radiation. We studied the light yield of pure polystyrene plates uncoated, and coated with a wavelength-shifting dye, coupled to a photomultiplier, using beta particles from a Sr-90 source. The results with the coated polystyrene plate show around four times higher photoelectron yield compared to the uncoated plate. To estimate the fraction of Cherenkov radiation, we compared it to the light yield of an uncoated and a coated plate of a non-scintillating material (PMMA). These results motivate future studies for the development of easy-to-build, low-cost, polystyrene-based plastic scintillator detectors.

T 68.3 Wed 16:45 T-H27

Simulation of the Beam Conditions Monitor for the upgrade of the LHCb experiment — JOHANNES ALBRECHT, ELENA DALL'OCCO, and •DAVID ROLF — Experimentelle Physik 5, TU Dortmund

2020. Two more stations, GE2/1 and GE0, will adopt the same technology during subsequent shutdowns. The GE1/1 system consists of 72 chambers made by two layers of Triple-GEM detectors, this will improve the muon triggering and tracking capabilities of CMS.

We report on the commissioning of the GE1/1 detectors, focusing on the performance measured during the first CMS commissioning runs which include cosmic-ray muons and collision events. We discuss on the strategy implemented to identify the optimum working point and the system performance stability.

T 67.9 Wed 18:25 T-H26

A first look at CMS gas electron multiplier data and certification — THOMAS HEBBEKER, KERSTIN HOEPFNER, FRANCESCO IVONE, and •SHAWN ZALESKI -III Physikalisches Institut A, RWTH Aachen University

To accommodate the increased radiation doses expected during the upcoming Run 3 at the CERN LHC, one of the primary experiments there, the Compact Muon Solenoid (CMS) will be upgraded. The first part of this upgrade, the installation of the gas electron multiplier (GEM) GE1/1 detector has been completed recently to help CMS cope with the radiation levels. The GEM chambers that comprise the GE1/1 detector have passed a rigorous battery of quality checks and are currently being commissioned in the end caps of the CMS muon system.

Initial data has been recorded by the GE1/1 detector using cosmic ray muon as well as proton-proton collision data. Dedicated data collection runs with and without the CMS 3.8 T magnetic field were performed. Certification procedures for this new GEM system have been developed and are being integrated with the existing procedures from the other CMS muon subsystems.

T 68: Detector Systems 2

The LHCb experiment is currently undergoing a major upgrade, preparing for the next period of data taking (Run 3) starting in 2022. The instantaneous luminosity in Run 3 will increase by a factor of five with respect to the previous runs. To ensure a safe operation of the experiment, simulation studies are performed for the Beam Conditions Monitor (BCM), responsible of protecting the LHCb detector against possible damage from the beam. Should an adverse beam scenario occur, the BCM will request a beam dump, thus preventing the beam from damaging the detector.

In this talk the impact of the changed beam conditions and configurations of the LHCb subdetectors on the BCM settings is evaluated, by studying the BCM response under both nominal and failing beam scenarios. Under nominal conditions the currents of the diamond sensors of the BCM detector are simulated and new settings for the beam dump logic are determined.

The effectiveness of the new proposed BCM settings are then investigated in critical beam scenarios. Simulations are performed under a beam scraping scenario to ensure the BCM will dump the beam before the beam drifts too far from its nominal position and starts to damage the sensitive detector parts.

T 68.4 Wed 17:00 T-H27

Location: T-H27

Development of a real-time track reconstruction for the proposed LumiTracker detector — Johannes Albrecht, Lukas Calefice, Elena DALL'OCCO, •LUKAS ROLF, and HOLGER STEVENS - Experimentelle Physik 5, TU Dortmund

Measuring the luminosity is a vital task performed at the LHCb experiment. The luminosity is used as feedback for the LHC and as input for many analyses. The main goal of the proposed LumiTracker detector is to provide a measurement of luminosity by operating independently from the rest of the LHCb experiment. The LumiTracker measures luminosity by reconstructing and counting tracks and to provide an online measurement of luminosity per bunch every few seconds, the reconstruction needs to be performed in real-time. This contribution will present the current status of the pattern recognition and track fitting algorithms developed. The overall performance of the track reconstruction will be compared between a CPU and a GPU implementation, with the latter bringing the advantage of greater parallelism.

T 68.5 Wed 17:15 T-H27

Commissioning of the SND@LHC detector — •ANDREW PICOT CONABOY — Humboldt-Universität zu Berlin, Berlin, Deutschland

The Scattering and Neutrino Detector at the LHC (SND@LHC) is a compact experiment installed 480 m from the ATLAS interaction point. SND@LHC allows for a novel investigation of all three neutrino flavours in the pseudo-rapidity range 7.2 < η < 8.6, with energies from 100 GeV to the TeV scale. Crucial for the reconstruction of neutrino charged current events, is the identification of the outgoing lepton: SND@LHC will reconstruct muons using iron blocks interleaved with planes of plastic scintillators coupled to silicon-photomultipliers. This technology will also be used to perform hadronic calorimetry. In the CERN north area, the muon system was installed for pion and muon test beams of energies in the 100 GeV to 300 GeV range. Presented in this contribution is the current status of muon system analyses from CERN test beams.

T 68.6 Wed 17:30 T-H27

Commissioning and Results of a Scintillator Based Beam Abort and Machine Protection System at SuperKEKB — •IVAN POPOV, HENDRIK WINDEL, and FRANK SIMON for the Belle II-Collaboration — Max-Planck-Institut für Physik, München, Deutschland

The asymmetric-energy collider SuperKEKB started its physics operation in March 2019. The usage of the nano-beam scheme enables collisions of electrons and positrons at record-breaking luminosities, but requires continuous particle injections at a rate of 50 Hz. These injections result in periods of high backgrounds, which can negatively affect the operation of Belle II subdetectors. In order to monitor and mitigate such backgrounds, the CLAWS detector system, consisting of scintillator tiles read out by silicon photomultipliers, has been in operation in various forms since 2016. Beginning with the first physics run in 2019, 32 sensors have been distributed along the final focusing magnets. Over the course of SuperKEKB's run time in 2020 they have proven to reliably observe disturbances in the particle beam which can result in catastrophically high backgrounds and quenches of the final focusing magnets. An electronics upgrade together with the implementation of a smart trigger logic enables the generation of a beam abort trigger within 200 ns after the occurrence of excessive background, thus ensuring the safe operation of the experiment. The CLAWS have been operating as a beam abort system since May 2021. In this report, the commissioning of the system and results achieved are discussed, and an outlook on plans for its further expansion is given.

T 68.7 Wed 17:45 T-H27

Detector system and simulation of the 155 MeV Hydro-Møller polarimeter at MESA — •MICHAIL KRAVCHENKO for the P2-Collaboration — PRISMA+ Cluster of Excellence and Institute of Nuclear Physics, Johannes Gutenberg University Mainz

The Mainz Energy-recovering Superconducting Accelerator (MESA) is an elec-

tron accelerator, which is currently under construction at the Johannes Gutenberg University Mainz. One aim for the MESA is the precise measurement of the weak mixing angle $sin^2\theta_w$, an important parameter of the Standard Model, with a relative uncertainty of 0.14%. The measurement will be performed by the P2 experiment by measuring the parity-violating asymmetry in elastic electronproton scattering at low momentum transfer Q^2 . MESA will provide a 150 μA beam of alternatingly polarized 150 MeV electrons with excellent beam stability. In order to achieve the goal of the P2 experiment, the beam polarization must be measured online with a very low systematic error (< 0.5% relative). The 155 MeV Møller polarimeter, as proposed by V. Luppov and E. Chudakov opens the opportunity for achieving these requirements. The current design of the detector system for the Hydro-Møller polarimeter and the results of the simulation with Geant4 are presented.

T 68.8 Wed 18:00 T-H27 **Fast neutron and gamma tomography with a stilbene-based multi-pixel de tector** — •NINA HÖFLICH and OLIVER POOTH — RWTH Aachen University -Physics Institute III B, Aachen, Germany

The neutron detectors group at the Physics Institute III B, RWTH Aachen University, develops a multi-pixel detector for a compact fast neutron imaging setup. Since the interactions of fast neutrons in matter differ from those of X-rays and gamma rays, imaging with fast neutrons in addition to X- or gamma ray imaging can provide complementary information about the object of interest.

Our current detector prototype uses cuboids of the organic scintillator stilbene as active material, coupled to a SiPM array. The pixel size is $6 \times 6 \text{ mm}^2$. The usage of stilbene allows to distinguish neutron- and gamma-induced signals in the detector.

An Americium-Beryllium neutron source delivers fast neutrons of up to 11 MeV as well as gamma rays of 4.44 MeV for our measurements.

The talk will focus on recent results from tomographic test measurements of simple objects. Strategies for image reconstruction and image processing will be presented. Additionally, possible improvements will be discussed, based on Geant4 simulation results.

T 69: DAQ and Trigger 3

Time: Wednesday 16:15-18:30

T 69.1 Wed 16:15 T-H28

Development of an FPGA Implementation of Convolutional Neural Networks for Signal Processing for the Liquid-Argon Calorimeter at ATLAS — ANNE-SOPHIE BERTHOLD, •NICK FRITZSCHE, MARKUS HELBIG, RAINER HENT-GES, ARNO STRAESSNER, and JOHANN CHRISTOPH VOIGT — Institut für Kernund Teilchenphysik (IKTP), TU Dresden, Germany

The Phase-II upgrade of the ATLAS detector will prepare for the high-luminosity phase of the LHC, where the number of proton-proton collisions occurring at the same time will increase significantly. This leads to higher requirements for the data processing, since the rate of detected particles in one detector cell will increase. New machine learning solutions are under development to better reconstruct the energy deposited in the calorimeter and its timing information than the current optimal filter approach.

This talk introduces the implementation of convolutional neural networks for FPGA hardware. The application of time division multiplexing is discussed, which is necessary to cover the high number of detector readout channels and to reuse the network for multiple input streams. The latest performance results in terms of FPGA resource usage, achievable operation frequency and latency are presented. To verify the hardware implementation, a software reference model was created and the precision of the calculation results was analyzed. At last, first preparations for tests on hardware are shown.

T 69.2 Wed 16:30 T-H28 gger of SuperCDMS — •HANNO

Towards the next generation Level-1 Trigger of SuperCDMS — •Hanno Meyer zu Theenhausen — Karlsruher Institut für Technologie

The SuperCDMS SNOLAB dark matter search experiment targets sensitivity towards nuclear and electron recoils down to energies of a few eV. At the lowest energies, the detector sensitivity is limited by thermal and electronic noise. To extract signals from the noise with high efficiency and resolution, SuperCDMS employs a Level-1 trigger system implemented on an FPGA on custom-hardware detector readout cards. Therein, digitized input traces are analyzed in a complex trigger architecture with a finite-impulse-response (FIR) filter at its heart. The FIR is configured as an optimal filter (OF) which can cause pathological "echo triggers" in the presence of large pulses. This presentation reports on the Level-1 trigger architecture, the OF FIR design and how echo triggers are circumvented making use of the complex trigger logic. Furthermore, an outlook on the performance of a neural network trigger as a potential trigger upgrade is given. Location: T-H28

T 69.3 Wed 16:45 T-H28

The ATLAS Forward Feature Extractor for the Phase-II trigger upgrade — •Adrian Alvarez Fernandez, Julian Blumenthal, Stefan Tapprogge, Ulrich Schaefer, and Bruno Bauss — Institut für Physik, Johannes Gutenberg Universität

The ATLAS detector will undergo many upgrades to account for the more challenging running conditions of the High Luminosity LHC. Some of these Phase-II upgrades will be focused on improving the trigger system, a crucial part to deal with the higher data rates and pile-up. Phase-I upgrades for Run 3 introduced the Feature EXtractors for a more refined processing of the calorimeter information and to better discriminate between jets, photons, electrons and taus. A Forward Feature EXtractor (fFEX) will be developed for the HL-LHC that will provide more flexible algorithms for the objects in the forward region ($|\eta| > 2.5$ for electrons and photons and $|\eta| > 3.2$ for jets). In contrast to the first level calorimeter trigger before HL-LHC, this new system will have access to the full detailed calorimeter granularity. The status of the preliminary design work on the fFEX will be presented and possible technology options will be discussed.

T 69.4 Wed 17:00 T-H28

Neural network based primary vertex reconstruction with FPGAs for the upgrade of the CMS level-1 trigger system — •MATTHIAS KOMM — Desy, Hamburg

A major challenge of the high-luminosity upgrade of the CERN LHC is to single out the primary interaction vertex of the hard scattering process from the expected 200 pileup interactions that will occur each bunch crossing. To meet this challenge, the upgrade of the CMS experiment comprises a complete replacement of the silicon tracker that will allow for the first time to perform the reconstruction of charged particle tracks and the primary interaction vertex at the hardware-based first level of the event trigger system (L1). Knowledge of the primary interaction vertex is a central component for distinguishing tracks and calorimeter clusters belonging to the hard scattering process from pileup interactions, which subsequently improves the energy estimate and resolution of physics objects such as jets and the missing transverse momentum. This talk will focus on the reconstruction of the primary vertex from tracks at L1 within the stringent time requirements of O(100ns) while being additionally restricted by the FPGA resource usage and latency. To optimally exploit and pass-on the available information at each stage of the vertex reconstruction, an algorithm based on an neural network model has been developed that possesses simultaneous knowledge about all stages and hence enables end-to-end optimization. Future plans for operating and tuning the algorithm on real data during data-taking will also be outlined.

T 69.5 Wed 17:15 T-H28

Development of machine-learning based topological selection algorithms for the upgraded L1 trigger system of the CMS detector — •IHOR KOMAROV, JO-HANNES HALLER, FINN LABE, ARTUR LOBANOV, and MATTHIAS SCHRÖDER — Institut für Experimentalphysik, Universität Hamburg

Future data-taking periods at the LHC bring a major increase of the instantaneous luminosity. To cope with the large detector occupancy within the bandwidth constraints, significant improvements of the trigger systems of the experiments are needed. The upgraded Level-1 trigger system of the CMS experiment will allow the execution of complex algorithms, such as neural networks, on fieldprogrammable gate arrays (FPGAs).

In this talk, a first proof-of-concept study on fast neural-network-based selection algorithms for the L1 trigger system of CMS will be presented. The algorithms were benchmarked with top-quark and Higgs pair production signals. Preliminary results show significant performance improvements compared to existing algorithms.

T 69.6 Wed 17:30 T-H28

Anomaly detection as a new strategy for the CMS Trigger — •Sven Bollweg, KARIM EL MORABIT, GREGOR KASIECZKA, and ARTUR LOBANOV — University of Hamburg, Germany

There exist strong hints for the existence of physics beyond the standard model (BSM). To search for such BSM processes at the LHC, potential candidate events first need to be selected. At CMS, the first selection step is the Level 1 (L1) trigger, which decides whether an event is stored for further analysis. The trigger decision is usually based on criteria motivated by specific models. Another strategy uses the idea that BSM events differ from events originating from standard model (SM) processes. A trigger decision could then utilize this difference to detect anomalous event properties.

This talk discusses such an anomaly detection trigger based on neural networks. An autoencoder (AE) network is trained to reproduce SM events. Using the AE to reproduce BSM events with anomalous properties leads to a reduced quality which can be used for the trigger decision. Since the L1 trigger has a very limited time for the decision, the AE needs to be deployed on dedicated hardware in the form of field programmable gate arrays which presents additional challenges.

T 69.7 Wed 17:45 T-H28

Implementation of tracking algorithms for live reconstruction using AI processors — •PATRICK SCHWÄBIG¹, JOCHEN KAMINSKI¹, MICHAEL LUPBERGER¹, KLAUS DESCH¹, and STEPHEN NEUENDORFFER² — ¹Physikalisches Institut, Universität Bonn, Deutschland — ²Xilinx Research Labs, San Jose, USA

For years, data rates generated by modern detectors and the corresponding readout electronics exceeded by far the limits of bandwidth and data storage space available in many experiments. Using fast triggers to discard uninteresting and irrelevant events is a solution used to this day. FPGAs, ASICs or even directly the readout chip are programmed or designed to apply a fixed set of rules based on low level parameters for an event pre-selection. Up until the last few years, live track reconstruction for triggering was rarely possible due to a conflict between processing time and the required trigger latency. With the emergence of novel fast and highly parallelized processors, targeted mainly at AI inference, attempts to sufficiently accelerate tracking algorithms become viable. The Xilinx Versal AI Series Adaptive Compute Acceleration Platform (ACAP) is one such technology and combines traditional FPGA

ory access. In this talk AI and non-AI algorithms for track reconstruction and especially their implementation on the Xilinx VCK190/Versal VC1902 Evaluation Kit for a dark photon experiment at the ELSA accelerator in Bonn will be shown and the expected performance will be discussed.

and CPU resources with dedicated AI cores and a network on chip for fast mem-

A hardware-based solution to filter interesting events more efficiently was explored for the planned increase in luminosity during the ATLAS Phase-II Upgrade. The plan is to install a filter based on the particles' tracks, which have to be evaluated in real-time. This task is well suited for implementation in hardware since, on the one hand, it requires only basic arithmetics and comparisons, and, on the other hand, it can be implemented in highly parallel hardware architecture, e.g., by exploiting FPGAs. This approach utilizes linear approximations to the helix parameters of trajectories and a database of simulated possible trajectories. A Track Fitter is implemented on the Stratix X FPGA from Intel with an integrated High Bandwidth Memory (HBM) for storing the simulated constants. In this talk, the final state of the firmware will be discussed and the implementation in hardware will be compared to simulation results and expectancies.

T 69.9 Wed 18:15 T-H28

Jet Tagging in the Level-1 Trigger of CMS for the HL-LHC — •PHILIPP RINCKE, KARIM EL MORABIT, GREGOR KASIECZKA, and ARTUR LOBANOV — Institut für Experimentalphysik, Universität Hamburg

With the upcoming upgrade of the High-Luminosity LHC, triggering in the CMS experiment will become more challenging as more particles will be present in each event. A possible solution to the increased complexity could be to employ trigger algorithms that use inputs from all sub-detectors that will become available in future upgrades and the increased computing power of the FPGAs on which the algorithms of the Level-1 Trigger (L1T), the first trigger level in CMS, are implemented.

At the L1T track parameters and particle identification of some of the jet constituents will be available, making it possible to evaluate the jet substructure. Many high-energy physics analyses require jet flavour identification, for which the substructure information can be exploited. By attempting this in the L1T a higher fraction of interesting events could be recorded or thresholds could be lowered. One big challenge is that neural networks, often used in jet tagging, are not straightforward to deploy on FPGAs.

Besides the tagging performance, strict timing and resource limitations need to be considered, which results in a compromise between network architecture and size. In this talk we present studies on how jet tagging can be used in the L1T. We consider a simple multilayer perceptron architecture as well as graphinspired network architectures.

T 70: Experimental Methods (general) 3

Time: Wednesday 16:15-18:30

T 70.1 Wed 16:15 T-H29 Calibration of the b-tagging mis-tag rate for charm jets based on W+c events at $\sqrt{s} = 13$ TeV with the ATLAS experiment — JOHANNES ERDMANN and •BENEDIKT GOCKE — TU Dortmund University, Department of Physics

Identifying jets containing heavy-flavour hadrons is needed for many analyses in the ATLAS experiment. This is done using flavour tagging algorithms. Since these algorithms need to be calibrated, it is important to also calibrate the mistag rate for non-heavy flavour jets. These calibrations are done by matching their performance in simulation to data.

The mis-tag rate calibration for *c*-jets using W+c events is presented, which allows for a relatively pure sample of *c*-jets. For this process, a prompt muon and a muon inside a jet (soft muon) are expected in the final state. Therefore, the usage of an event selection which selects events in regions defined by the electric charge of the expected muon pair in the final state is shown. Furthermore, the construction of a signal region to reduce the background contributions, which is based on the expected opposite-sign-charged lepton pair for the signal process is illustrated. The calculation of the efficiencies for the used heavy flavour tagger is shown. Finally, the resulting data-to-simulation scale factors are discussed.

Location: T-H29

T 70.2 Wed 16:30 T-H29

Convolutional Networks and Deep Learning at the Belle II Experiment — •JOHANNES BILK, SÖREN LANGE, KATHARINA DORT, STEPHANIE KÄS, and TIMO SCHELLHAAS — Justus-Liebig-Universität, Gießen,Germany

The Belle II pixeldetector (PXD) has a trigger rate of up to 30 kHz for 8 M pixels. Its proximity to the interaction point allows it to detect exotic highly ionizing particles such as antideuterons, magnetic monopoles, stable tetraquarks or pions with small transverse momenta < 100 MeV. Those particles leave no tracks in the outer parts of the Belle II detector, and thus their pixel data may be deleted online as part of background suppression. In this contribution, we evaluate the performance of a machine learning algorithm to identify slow pions only on the basis of pattern recognition of pixel cluster structures. We employ convolutional neural networks with different kernel configurations and use images of 9x9 pixel matrices as input. On the long term such image recognition techniques could provide a rescue mechanism for the pixel data before they are erased. Results on accuracy and sensitivity are presented.

T 70.3 Wed 16:45 T-H29

b-Tagging studies for the ATLAS experiment - •Eleonora Loiacono Deutsches Elektronen-Synchrotron DESY, Hamburg and Zeuthen, Germany At the LHC, among the different Higgs boson production mechanisms, the associated production with a vector boson $VH \rightarrow \bar{b}b$ is considered as the golden channel for the measurement of the Higgs coupling to b-quarks, due to its high sensitivity. In the VH(bb) analysis, the available statistics of the simulated Monte Carlo (MC) samples is largely reduced by the b-tagging algorithms, especially for background processes, resulting in large statistical fluctuations in the MC templates. The first study that will be presented, is related to the reduction of this statistical uncertainty. In order to avoid being dominated by this un- certainty a technique based on Truth Tagging is performed, exploiting a Graph Neural Network to make an optimal use of multidimensional information associated to jets. The second study relies on improving of the connection between Flavour Tagging scale factors measured on data and those corresponding to MC to MC variations. b- tagging is very important for VH(bb) and the calibration procedures provide the use of the Scale Factors (SF). The so-called smoothing procedure is currently applied to these SF, which allows to have SF that are not dependent on the binning of the histograms used to do the calibration. In particular, there will be a focus on reviewing the currently smoothing procedure and also on including more statistical tests on procedure currently adopted for data SF.

T 70.4 Wed 17:00 T-H29

Extrapolation of flavour tagging calibrations to high transverse momenta — ARNULF QUADT, ELIZAVETA SHABALINA, and •SREELAKSHMI SINDHU — II. Physikalisches Institut, Georg-August-Universität Göttingen

Identifying jets containing heavy flavoured hadrons can be very beneficial for a variety of analyses and this can be done using flavour tagging algorithms. Currently, these algorithms are calibrated by matching their performance on data to simulation. However, for jets with transverse momentum greater than a few hundred GeV these calibrations do not exist due to statistical limitations. Flavour-tagging information for jets with high transverse momenta can be quite useful for analysis such as search for heavy resonances. Hence, a Monte Carlo based extrapolation of the data based flavour tagging calibrations is done to extend these up to 3 TeV. The extrapolation uncertainties are calculated by propagating the relevant modelling, tracking and jet uncertainties through the tagging algorithm. In this talk, the extrapolation norecdure will be explained and the results from my work on the extrapolation uncertainties for charm and light jets will be presented for the b-tagging algorithms (DL1, DL1r) that are used for the analysis of the Run 2 data by the ATLAS Collaboration.

T 70.5 Wed 17:15 T-H29 Utilizing muons to tag b-jets in ATLAS — •FREDERIC RENNER, CLARA ELISA-

ветн Leiтgeв, and Cigdem Issever — DESY, Zeuthen, Germany

Several Beyond Standard Model theories predict Higgs boson pair production with much larger cross sections compared to the Standard Model. While Higgs boson pair production has not yet been observed at the LHC due to the rarity of the process, it may come within reach with more data in upcoming runs with improved statistics and improved particle reconstruction. A Higgs boson pair decays predominantly into four b-quarks, which makes the identification of jets originating from b-quarks particularly important. B-quarks have a longer lifetime than lighter quarks, leading to a second displaced vertex separated from the primary vertex which is mainly exploited in b-tagging algorithms. However, when using the results in a neural network to distinguish the flavor origin of jets, the network still misidentifies a lot of jets originating from lighter quarks as bquark jets. About 20% of b-quarks decay semileptonically with a so-called soft muon in the final state. They are soft because they are generally less energetic than particles originating from the primary vertex. Utilizing the muon information, the rejection of fake b-jets can be improved substantially at a given selection efficiency.

T 70.6 Wed 17:30 T-H29

Signal efficiency corrections for boosted $X \rightarrow b\bar{b}$ tagger using $Z \rightarrow b\bar{b}$ events with the ATLAS experiment — •DAARIIMAA BATTULGA, ARELY CORTES GON-ZALEZ, and CIGDEM ISSEVER — Institut für Physik, Humboldt-Universität zu Berlin

It is of utmost interest to efficiently identify a new heavy resonant particle or to study boosted Higgs boson decaying into a pair of b-quarks in the ATLAS experiment. Particles produced with high transverse momentum (p_T) will have very collimated decay products in the final state. This boosted topology makes it particularly challenging to distinguish two b-jets in the calorimeter. To overcome this, these b-jet fragmentations are clustered within a large radius R=1.0 jet, and its associated track jets are b-tagged. However, even with this approach, the double b-tagging efficiency decreases at the large p_T . The boosted $X \rightarrow b\bar{b}$ tagger improves the efficiency of the b-tagging at the higher p_T region. This double b-tagger is based on the neural network that uses the kinematic distributions of the large-R jet and flavour information of variable radius track jets. In order to apply this $X \rightarrow b\bar{b}$ tagger to the physics analysis, it needs a dedicated calibration.

Hence, this talk presents the in situ signal efficiency calibrations of a new $X \rightarrow b\bar{b}$ tagger using $Z \rightarrow b\bar{b}$ events. We have derived the data-to-simulation scale factors using full Run 2 *pp* collision data collected by the ATLAS experiment at the center of mass energy of $\sqrt{s} = 13$ TeV with an integrated luminosity of $\mathcal{L} = 139 \text{ fb}^{-1}$. The signal efficiency corrections covering the soft (hard) p_T region are derived using $Z \rightarrow b\bar{b}$ events with a recoiling photon (jets).

T 70.7 Wed 17:45 T-H29

A machine-learning based method to improve isolation variables for photon identification with the ATLAS detector — JOHANNES ERDMANN, OLAF NACK-ENHORST, and •MICHAEL WINDAU — TU Dortmund University, Department of Physics

The study of photons is crucial for finding and measuring many processes at colliders. Predominantly, prompt photons, which are created during the collisions, play an important role and have to be distinguished from hadrons decaying into photons. Different methods are used to distinguish this signal from the background. One of these is the use of isolation variables. These are based on track measurements and information from the calorimeters, where they are defined by the activity in a cone around the candidate object. They are currently built in ATLAS by discrete cuts.

In this talk, studies on improving isolation variables using deep neural networks will be presented.

T 70.8 Wed 18:00 T-H29

Data-driven corrections to shower shape variables for photon identification at the ATLAS experiment with 13 TeV *pp* collision data — •JAN LUKAS SPÄH, BJÖRN WENDLAND, and JOHANNES ERDMANN — Technische Universität Dortmund, Fakultät Physik

Measurements of Standard Model processes, searches for new particles or processes forbidden in the Standard Model with photons in the final state play an important role in the physics programme of the ATLAS experiment. At hadron colliders, studies of photons are particularly challenging, as large background contributions arise from jets that can be misidentified as photons. This requires an identification algorithm that provides high efficiency for genuine photons while ensuring an excellent background rejection for misreconstructed objects.

Currently, this method relies on rectangular cuts on so-called shower shape variables, which capture relevant information about the shape and evolution of the electromagnetic shower and the possible leakage into the hadronic calorimeter. While the longitudal shower development through the calorimeter layers is modelled well, residual mismodelling is observed for lateral shower shape distributions. Therefore, the simulated distributions are corrected with a data-driven approach.

In this talk, studies of univariate first- and second-order corrections obtained from the full Run 2 dataset are discussed and recent improvements are highlighted.

T 70.9 Wed 18:15 T-H29

Towards tuning electromagnetic shower properties to data with AtlFast3 – •JOSHUA BEIRER^{1,2}, MICHAEL DUEHRSSEN¹, and STAN LAI² – ¹CERN – ²Georg-August-Universität Göttingen

AtlFast3 is the next generation of high precision fast simulation in ATLAS and encompasses a parametrised and a machine-learning approach based on Generative Adversarial Networks (GANs). With respect to its predecessor, AtlFast3 significantly improves in physics performance while retaining the benefit of a considerably faster simulation in comparison to Geant4.

A precise simulation of electromagnetic (EM) shower properties in the ATLAS calorimeter is crucial for the identification of particle showers originating from electrons and photons. While AtlFast3 precisely simulates the properties of EM showers, it inevitably inherits any mismodelling of the full Geant4 simulation, upon which its parametrization is based. Differences between the Geant4 simulation and data collected by the ATLAS detector are well known but insufficiently understood. Traditionally, these discrepancies are corrected using ad hoc methods such as the applications of shifts to the central values of the corresponding distributions, a procedure known as fudging.

In this talk, a brief overview of fast simulation in ATLAS is given. Furthermore, the development of different models directly embedded within the simulation framework used to tune EM shower properties directly to data are described and it is shown that AtlFast3 can be modified in a way that the shower shapes observed in data are accurately reproduced by the simulation.

T 71: Neutrino Astronomy 3

Time: Wednesday 16:15-18:30

T 71.1 Wed 16:15 T-H30

Up-going high energy showers with the fluorescence detector of the Pierre Auger Observatory — IOANA ALEXANDRA CARACAS and •KARL-HEINZ KAM-PERT for the Pierre Auger-Collaboration — Bergische University Wuppertal, Gaußstr. 20, Wuppertal, Germany

The ANITA observations of two steeply up-going cosmic ray like showers with energies above 10^{17} eV remain unexplained. The Fluorescence Detector (FD) of the Pierre Auger Observatory is also sensitive to such phenomena, given its wide field of view and substantial operation time. Using 14 years of available FD data, the post-selection exposure to up-going induced showers exceeds the one of ANITA by a factor of at least 10, as indicated from dedicated studies. Therefore a search for up-going induced air showers with the FD can be used to either refute or confirm the occurrence of such intriguing events.

We have conducted a generic search for upward cosmic ray like induced air showers using the FD of the Pierre Auger Observatory. Dedicated Monte Carlo simulations of both signal and expected background, together with the usage of a 10% burn data sample, have been used in order to apply selection criteria and calculate the resulting FD exposure. The unblinding of the data indicates no excess found above background expectations. As a result, preliminary upper limits are set on the flux of up-going cosmic ray like induced air showers.

*Gefördert durch die BMBF Verbundforschung Astroteilchenphysik (Vorhaben 05A20PX1)

T 71.2 Wed 16:30 T-H30

Constraining BSM scenarios producing up-going τ induced air showers with the Pierre Auger Observatory^{*} — •IOANA ALEXANDRA CARACAS for the Pierre Auger-Collaboration — Bergische University Wuppertal, Gaußstr. 20, Wuppertal, Germany

High energy steeply up-going air showers as observed by ANITA can't be explained by Standard Model (SM) physics and require Beyond Standard Model (BSM) scenarios. τ -leptons decaying in the atmosphere represent the main primary candidates for such showers.

The Pierre Auger Observatory has set strict upper limits on the flux of upgoing cosmic ray like air showers. The generic search is recast here in terms of BSM particles producing up-going τ -leptons. In any such BSM scenario a significantly reduced cross section of the hypothetical particle with matter is required, allowing them to propagate through the Earth with sufficiently low interaction probability. Interactions close to the surface could result in the creation of τ leptons that escape into the atmosphere to induce up-going showers. The optimum BSM cross section for this to happen is found to be $\sigma_{\rm BSM} \approx 10^{-2} \sigma_{\rm v}$.

Using both the Surface Detector and the Fluorescence Detector data, combined upper limits are set on particles creating up-going τ -leptons for different BSM scenarios.

*Gefördert durch die BMBF Verbundforschung Astroteilchenphysik (Vorhaben 05A20PX1)

T 71.3 Wed 16:45 T-H30

Search for a high energy neutrino flux from Gamma Ray Bursts using the Pierre Auger Observatory* — •TOBIAS HEIBGES for the Pierre Auger-Collaboration — Bergische Universität Wuppertal, Gaußstraße 20 42119 Wuppertal, Germany

Gamma Ray Bursts (GRBs) are among the most violent explosions known in the Universe. A characteristic feature is a very high flux of gamma rays produced in these explosions which can be observed and located by satellites, such as the Swift and Fermi satellites.

A high energy gamma ray flux can be interpreted as an indication for the acceleration of charged particles up to the highest energies. Therefore, Gamma Ray Bursts are among the prime candidates to be sources of ultra high energy cosmic rays (UHECRs). High energy neutrinos are commonly regarded as a smoking gun indicator of UHECR acceleration and as they are not deflected by magnetic fields they can be easily traced back to their source and thereby contribute to unraveling the mystery about the origin of UHECRs.

The Pierre Auger Observatory is sensitive to high energy neutrinos with energies beyond $\sim 10^{17}$ eV. In this talk the non-observation of any high energy neutrino events is used to set an upper limit on the high energy neutrino flux seen on earth, produced by GRBs.

* Supported by BMBF Verbundforschung, grant 05A20PX1

T 71.4 Wed 17:00 T-H30

Event selection for new triggers used for neutrino detection at the Pierre Auger Observatory* — •SRIJAN SEHGAL for the Pierre Auger-Collaboration — Bergische Universität Wuppertal, Wuppertal, Germany

The Pierre Auger Observatory with its large array of Surface Detector (SD) stations can be used to detect highly inclined neutrino-induced extensive air showers. Two new SD triggers, time-over-threshold-deconvolved (ToTd) and multiplicity of positive steps (MoPS), installed in 2014 were shown to vastly increase the detection capability for the neutrino-induced air showers in the lower energy $(E < 10^{19} \text{eV})$ regime with a little to no change in background events.

This talk analyzes the event selection procedure on data and neutrino-induced showers simulated with CORSIKA both reconstructed using the Auger software framework. Events in the zenith angle range of $60^{\circ} < \theta < 75^{\circ}$ and energies below 10^{19} eV are selected to investigate the low-energy performance of the new triggers. The main point of focus is the effect of the new triggers on efficiency and purity of an improved neutrino selection.

*Gefördert durch die BMBF Verbundforschung Astroteilchenphysik (Vorhaben 05A20PX1)

T 71.5 Wed 17:15 T-H30

Reconstruction performance using RNO-G — •SJOERD BOUMA for the RNO-G Collaboration-Collaboration — Erlangen Centre for Astroparticle Physics (ECAP), Erwin-Rommel-Str. 1, D-91058 Erlangen

RNO-G is an in-ice radio detector in Greenland which aims to detect Extremely High Energy (EHE) neutrinos through the Askaryan effect. Deployment started in Summer 2021, with the first 3 out of a planned 35 detector stations now built and taking data. As the first production-scale in-ice radio neutrino detector, RNO-G both complements as well as helps to inform the design of the planned radio extension of IceCube-Gen2.

One important aspect of RNO-G and potential future radio neutrino detectors, aside from their effective volume, is the ability to reconstruct the properties of detected (neutrino-induced) radio shower signals. An accurate reconstruction of the neutrino direction is crucial in order to identify potential sources of EHE neutrinos. We will present a brief overview of the current state of reconstruction performance using algorithms developed for the open-source NuRadioMC software package by the RNO-G collaboration, and provide an outlook for future improvements.

T 71.6 Wed 17:30 T-H30

Data reduction for the Radio Neutrino Observatory Greenland — •ZACHARY MEYERS for the RNO-G Collaboration-Collaboration — DESY, Platanenallee 6, 15738 Zeuthen, Germany — Erlangen Center for Astroparticle Physics (ECAP), Friedrich- Alexander-University Erlangen-Nuremberg, 91058 Erlangen, Germany

Continuing the search for utra-high energy neutrinos (> 10 PeV) beyond the range of optical detection methods, the Radio Neutrino Observatory Greenland (RNO-G) is now online after a successful first season of deployment. Total data taken during the shortened 2021 campaign from the three operational stations amounts to nearly ten million recorded events, requiring more than 330GB of storage. While this could be considered a manageable sum, next year another 7 stations are planned to come online, while the complete array will consist of 35 total. And for future experiments, requiring hundreds of similar stations, the data volumes rapidly increase to a level where it is no longer feasible to run direction and energy reconstruction algorithms on the entire dataset. Low level cuts must be made early in the data processing stages (or even onboard the detector itself in real time) in order to be computationally efficient. In an attempt to discriminate between thermal noise fluctuations, anthropogenic noise and neutrino-like signal, we show the potential effectiveness of deep learning approaches, specifically convolutional neural networks (CNNs), in both the time and frequency domains. When compared and combined with more traditional methods such as matched filtering, a comprehensive strategy for post trigger filtering can be established.

T 71.7 Wed 17:45 T-H30

Cosmic ray detection efficiency and implications for in-ice radio detectors for high-energy neutrinos — •LILLY PYRAS for the RNO-G Collaboration-Collaboration — DESY, Platanenallee 6, 15738 Zeuthen, Germany — Erlangen Center for Astroparticle Physics (ECAP), Friedrich-Alexander-University Erlangen-Nuremberg, 91058 Erlangen, Germany

A promising technique to measure neutrinos above 10 PeV is the detection of radio signals generated by the Askaryan effect. The emission is caused by neutrinoinduced particle cascades in dense media e.g. ice. Since 2021 the Radio Neutrino Observatory Greenland (RNO-G) is being deployed, consisting of in-ice strings of antennas down to 100 m and antennas closer at the surface. One of the main challenges of the data analysis is distinguishing between background stemming from cosmic rays e.g. high energy muons and a real neutrino event. By building the detector with surface antennas we can use the established method of radio detection of air showers to identify incoming muons and use these signals as veto mechanism in the neutrino detection. An efficient veto trigger will lend higher confidence in identifying neutrinos and prevent the false positive neutrino detections caused by muons. This report presents the development of tagging incoming air showers as veto and analyses its performance.

T 71.8 Wed 18:00 T-H30

Search for periodic low energy neutrino sources — •MAXIMILIAN EFF for the ANTARES-KM3NET-ERLANGEN-Collaboration — ECAP, Friedrich-Alexander Universität Erlangen-Nürnberg (FAU), Erlangen, Germany

Pulsars are rotating neutron stars that emit beams of electromagnetic radiation. Neutrino emission from pulsars has been the subject of phenomenological models during the last decades. So far, experimental data has not shown any significant neutrino emission at high energies. This contribution reports about the development of a novel approach that aims at identifying low-energy neutrinos from periodic sources with a neutrino telescope. This is done by applying a Fast Fourier Transformation to the PMT counting rate time series.

T 71.9 Wed 18:15 T-H30

PLEnuM: A world-wide monitoring system of high-energy astrophysical neutrinos — •LISA SCHUMACHER¹, MATTEO AGOSTINI², MAURICIO BUSTAMANTE³, FOTEINI OIKONOMOU⁴, and ELISA RESCONI¹ — ¹ECP, TU Munich, GER — ²UCL, London, UK — ³NBI, Copenhagen, DEN — ⁴NTNU, Trondheim, NOR

T 72: Cosmic Ray 3

Time: Wednesday 16:15-18:30

T 72.1 Wed 16:15 T-H31

Using in-ice muons for Cosmic-Ray composition analysis at IceCube Observatory — •PARAS KOUNDAL for the IceCube-Collaboration — Institute for Astroparticle Physics, KIT Karlsruhe, Germany

Understanding the dynamics of astrophysical sources is a pursuit that is very dear to many astrophysicists. Cosmic-Rays (CRs), charged particles from these astrophysical accelerators, provide us with a unique opportunity to discern the fundamental properties and behavior of such sources. IceCube Neutrino Observatory, concealed deep under the South-Pole Antarctic ice, detects the particles from these astrophysical sources. The integrated operation of the in-ice array of IceCube (primarily a neutrino detector), with its surface array, IceTop, affords us unique three-dimensional detection capabilities for CR-induced air showers.

The talk will discuss the work done to use the in-ice shower-footprint primarily caused by high-energetic muons in cosmic-ray air-showers, for improving cosmic-ray composition estimation at IceCube Observatory. The work will introduce new composition-sensitive parameters with minimal dependence on hadronic-interaction models. Hence, the work provides a suitable solution for detailed composition analysis while reducing systematic effects of choosing a hadronic-interaction model for interpretation of observed real data.

T 72.2 Wed 16:30 T-H31

Unfolding the Atmospheric Muon Spectrum Using Stopping Muons in Ice-Cube — •LUCAS WITTHAUS, KAROLIN HYMON, JOHANNES WERTHEBACH, JAN-INA BOLLES, and JAN SOEDINGREKSO for the IceCube-Collaboration — Astroparticle Physics WG Rhode, TU Dortmund University, Germany

The IceCube Neutrino Observatory is a cubic-kilometer neutrino detector located in the ice sheet close to the geographical South Pole. However, the majority of detected events is caused by atmospheric muons created in cosmic ray induced air showers in the upper layers of the atmosphere. Upon entering the antarctic ice, they lose energy in interactions with the surrounding matter, resulting in a limitation of their propagation length.

This talk presents the unfolding of the stopping muon depth intensity by means of a maximum likelihood approach. It is conducted on a subset of events, comprising single muons, which stop inside the IceCube detector. Deep neural networks are used to perform the event classification and reconstruction tasks.

T 72.3 Wed 16:45 T-H31

Towards the Energy Spectrum of Cosmic Rays using Atmospheric Stopping Muons in IceCube — •JANINA BOLLES, KAROLIN HYMON, JOHANNES WERTHE-BACH, LUCAS WITTHAUS, and JAN SOEDINGREKSO for the IceCube-Collaboration — Astroparticle Physics WG Rhode, TU Dortmund University, Germany

In the IceCube Neutrino Observatory the main type of detected events are muons being produced by cosmic ray particles interacting with the Earth's atmosphere. In the context of neutrino analyses, these muons are the dominant background. In case of cosmic ray physics the energy losses of the muons within the detector can be used as an indicator to reconstruct the cosmic ray energy spectrum.

In this work, muon events stopping inside the detector are selected to use the range to the stopping point as a proxy for the muon energy. This approach takes advantage of the high statistics of atmospheric muons. Strict cuts on the reconstruction can be applied to obtain an event sample of single muons with high resolution. The reconstructed range of the muons can be used to estimate the cosmic ray energy spectrum. First results of the unfolded cosmic ray flux are presented.

The discovery of high-energy astrophysical neutrinos by IceCube has shaped neutrino astronomy in the recent years. However, the observation rate of astrophysical neutrinos in the TeV-PeV energy range remains small, such that various questions about high-energy neutrinos and their astrophysical origin remain open. This situation will improve when new neutrino telescopes will come online in the next years: KM3NeT, Baikal-GVD and P-ONE in the Northern Hemisphere, as well as IceCube-Gen2 as extension of IceCube in the Southern Hemisphere. In order to answer our open questions, we propose the Planetary Neutrino Monitoring System (PLEnuM), a concept for a combined repository of world-wide high-energy neutrino observations. PLEnuM can reach more than four times the exposure available today by combining the exposures of the current and future neutrino telescopes distributed around the world. Depending on the declination, spectral index, and flavor, PLEnuM will improve the sensitivity to astrophysical neutrinos by up to two orders of magnitude. We present first estimates on the capability of PLEnuM to discover Galactic and extragalactic sources of astrophysical neutrinos and to characterize the diffuse flux of highenergy neutrinos in unprecedented detail.

Location: T-H31

T 72.4 Wed 17:00 T-H31

An updated model of galactic diffuse neutrinos for future IceCube searches — •GEORG SCHWEFER^{1,2}, PHILIPP FÜRST², ERIK GANSTER², PHILIPP MERTSCH¹, and CHRISTOPHER WIEBUSCH² — ¹RWTH Aachen University - Institute for Theoretical Particle Physics and Cosmology, Aachen, Germany — ²RWTH Aachen University - Physics Institute III B, Aachen, Germany

Diffuse galactic neutrinos are produced in interactions of hadronic cosmic rays with the interstellar medium in the Milky Way. This flux is a practically guaranteed signal for high-energy neutrino observatories. It has not been identified yet, but recent searches indicate that a discovery might be in reach within the next few years.

Because of the large background from atmospheric neutrinos, these searches require detailed modelling of the signal. These models also relate the (non-)observations to the propagation and injection properties of galactic cosmic rays at PeV energies.

In this talk, we present an updated model for the galactic diffuse neutrino flux tuned to the latest direct cosmic ray and diffuse gamma ray measurements, and discuss its systematic dependencies. We also show sensitivity estimates for future IceCube galactic plane searches with this model.

T 72.5 Wed 17:15 T-H31

Neural networks for cosmic ray simulations — • PRANAV SAMPATHKUMAR, AN-TONIO AUGUSTO ALVES JUNIOR, TANGUY PIEROG, and RALF ULRICH for the CORSIKA 8-Collaboration - Institute for Astroparticle Physics (IAP) - KIT Simulating cosmic ray showers at high energies is very memory and time intensive. Current model-dependent hybrid techniques are constrained by our ability to model from known physics. This contribution discusses novel machine learning techniques in order to bypass explicit simulations, and extract features which can't be modeled easily from first principles. The potential of Generative Adversarial Neural Networks (GANs) in learning and emulating cosmic ray simulations is discussed, along with a presentation of preliminary attempts in using a GAN in generating universal electron-positron distributions associated to showers with varying primaries and energies. The applicability and potential pitfalls in using a neural network based approach for cosmic ray simulations is also discussed. Finally, a CONEX (hybrid simulations using cascade equations) inspired Recurrent Neural network (RNN) model is presented. Preliminary results obtained from training an RNN using a cosmic ray simulation dataset for electromagnetic cascades generated using CORSIKA8 are summarized.

T 72.6 Wed 17:30 T-H31

Extrapolation uncertainty of meson-air cross-sections in UHECR air shower simulations — •MAXIMILIAN REININGHAUS^{1,2}, RALF ULRICH¹, RALPH ENGEL¹, and TANGUY PIEROG¹ — ¹Karlsruher Institut für Technologie, Karlsruhe, Deutschland — ²Instituto de Tecnologías en Detección y Astropartículas, Buenos Aires, Argentina

The interaction cross-sections of long-lived hadrons with air nuclei are an important ingredient in the simulation of air showers initiated by high energy cosmic rays. For protons they are tightly constrained by LHC measurements. For other species, in particular pions, which are the most abundant hadrons in air showers, however, precise measurements are available only at low energies. Since there exists significant leeway in a large energy range up to the highest energies, hadronic interation models differ in their extrapolations by up to 30 %.

In this contribution, we study the impact of this extrapolation uncertainty on air shower phenomenology by introducing ad hoc, energy-dependent factors to scale the cross-sections for each species independently. Using a hybrid setup with

153

CORSIKA 8 and CONEX, we simulate UHECR air showers with these modified cross-sections and study the effect on muon content, shower maximum and muon production depth. We find that the longitudinal development is sizeably affected, while the particle content changes only to a minor degree.

T 72.7 Wed 17:45 T-H31

Simulating radio emission from air showers with CORSIKA8 — •NIKOLAOS KARASTATHIS¹, REMY PRECHELT², TIM HUEGE^{1,3}, and JUAN AMMERMAN-YEBRA⁴ for the CORSIKA 8-Collaboration — ¹Institute for Astroparticle Physics, Karlsruhe Institute of Technology, Karlsruhe, Germany — ²Department of Physics and Astronomy, University of Hawaii Manoa, Hawaii, USA — ³Astrophysical Institute, Vrije Universiteit Brussel, Brussels, Belgium — ⁴Instituto Galego de Física de Altas Enerxías, Universidade de Santiago de Compostela, Santiago de Compostela, Spain

CORSIKA 8 (C8) is a new framework for air shower simulations implemented in modern C++17, based on past experience with existing codes like CORSIKA 7. It is a project structured in a modular and flexible way that allows the inclusion and development of independent modules that can produce a fully customizable air shower simulation. The calculation of radio emission from the simulated particle showers is incorporated as an integral module of C8, including signal propagation and electric field calculation at each antenna location using the "Endpoints" and ZHS formalisms simultaneously. Due to C8's flexibility, the radio functionality can be used both to validate other physics modules and to investigate specific physical scenarios. In this talk, we are going to present air shower simulations generated with C8 and compare their predicted radio emission with corresponding air showers simulated with CORSIKA 7 and ZHAireS. Radio calculation validation, a comparison of the "Endpoints" and ZHS formalisms along with the future steps of radio in C8 are also going to be shown.

T 72.8 Wed 18:00 T-H31 Energy Reconstruction using a Template Method for Radio Signal of Air Showers recorded by the Prototype Station of the IceCube Surface Enhancement — •ROXANNE TURCOTTE for the IceCube-Collaboration — Institut für Astroteilchenphysik, Karlsruher Institut für Technologie (KIT) The IceTop array, located at the surface of the IceCube Neutrino Observatory, is currently used as a veto for the in-ice neutrino detection as well as a cosmic-ray detector. Over the years, snow accumulated on the IceTop detector leading to a reduction of its sensitivity and resolution. In order to improve the detector, an enhancement of IceTop is planned in the coming years which consists of an array of scintillation panels and radio antennas. The radio antennas will lead to a better resolution of the energy and the depth of shower maximum (X_{max}) around the second knee region of the cosmic-ray energy spectrum. Eventually, hybrid detection will enable a better estimation for the mass of the primary cosmic ray.

In January 2020, a prototype station comprising three antennas and eight scintillation panels was deployed at the South Pole. We developed the tools necessary to use a template-matching method for energy reconstruction and applied it to some of the radio events recorded. This template method uses Monte-Carlo simulations and compares it to recorded data. For this, a set of simulations is created using the reconstruction by IceTop as input to CORSIKA/COREAS. In this talk, we will present the method and the preliminary results.

T 72.9 Wed 18:15 T-H31

IceAct Upgrade Status - SiPM Based Compact Imaging Air-Cherenkov Telescopes for IceCube — •HANNAH ERPENBECK, THOMAS BRETZ, LARS HEUER-MANN, CENGIZ KURUOGLU, FRANK MALOWSKI, MARK MEYERS, FLORIAN RE-HBEIN, MERLIN SCHAUFEL, and CHRISTOPHER WIEBUSCH for the IceCube-Collaboration — RWTH Aachen University - Physics Institute III B, Aachen, Germany

IceAct is an array of compact imaging air Cherenkov telescopes that are optimized for the harsh conditions of the South Pole. Since January 2019 two IceAct telescopes, featuring 61 SiPM pixels and a Fresnel lens based optics, operate at the surface above IceCube in the center of IceTop. By hybrid measurements of cosmic rays together with the IceTop and the IceCube detectors, they enable improved cosmic ray studies and cross calibrations. Six new telescopes are currently being assembled as an upgrade for IceAct. To ensure high instrument reliability, each of the telescopes is tested individually including field tests and strict quality assurance of all components. This talk will report on the project status as well as on the construction and the testing results of the new telescopes.

T 73: Cosmic Ray 4

Time: Wednesday 16:15-18:15

T 73.1 Wed 16:15 T-H32

Probing magnetic fields in the Galactic halo and studing their effects on arrival direction of cosmic rays — •VASUNDHARA SHAW, ANDREW TAYLOR, and ARJEN VAN VLIET — Deutsches Elektronen-Synchrotron (DESY), Zeuthen, Germany

The Galactic halo in the past was less explored than other regions of the Galaxy. However, in the last decade, this has started to change with the observation of the Fermi bubbles and the latest eROSITA bubbles, we know that there is much to unravel in this region.

These large extended Galactic halo bubbles can play a key role in the magnetic field structure of the Galaxy. The magnetic fields for the bubble region have so far been largely masked out in the models however, the strength and geometry of magnetic fields in this region can be fundamental not only in understanding the Galactic magnetic fields but also the deflection of extra-Galactic cosmic rays.

In this talk, I will try to motivate the reason behind our toy model magnetic field for the Galactic halo and highlight how it compares with radio observation data. The second part of this talk will focus on the effect arrival directions of cosmic rays from our toy model and compare it with other existing magnetic field models.

T 73.2 Wed 16:30 T-H32

Incorporating the Galactic magnetic field into the propagation effects of cosmic rays in the transition region* — •ALEX KÄÄPÄ — Bergische Universität Wuppertal, Gaußstr. 20, 42119 Wuppertal

In the energy range signifying the transition from Galactic to extragalactic cosmic rays (GCRs and EGCRs), propagation in the Galactic magnetic field (GMF) changes from diffusive to ballistic. This leads to a range of observable effects vital to understanding the respective contributions of GCRs and EGCRs to the total flux. GCRs more readily leak out of the Galaxy with increasing energy and, hence, the flux arriving at Earth is suppressed. EGCRs experience two competing effects, shielding from as well as confinement in the Galactic plane, both of which weaken with energy. These effects have been re-confirmed to cancel exactly in the case of isotropic injection. Flux modifications can occur in the case of an anisotropic EGCR flux into the Galaxy. Their nature depends both on the type and direction of the anisotropy.

In this talk, we present the propagation effects that the GMF imposes on the flux of GCRs and EGCRs. We incorporate these into minimal, experimentally and theoretically motivated injection spectra of GCRs and EGCRs. With this incorporation, we seek to retrieve a more realistic picture of the expected flux arriving at Earth, and to better estimate the nature and degree of possible additional contributions to the injected flux.

* Gefördert durch die BMBF Verbundforschung Astroteilchenphysik (Vorhaben 05A20PX1).

T 73.3 Wed 16:45 T-H32

Location: T-H32

Cosmic-ray transport in astrophysical environments is often dominated by diffusion in a magnetic field with a turbulent component. The diffusion properties of charged particles directly influence observable properties, such as the spectrum of cosmic rays and their secondaries produced in interactions. In many diffusion scenarios, the simplified assumption of fully resonant Kolmogorov diffusion in the quasi-linear limit results in a parallel diffusion coefficient $D \propto E^{1/3}$. A quantitative investigation of the scattering regimes, however, shows that the diffusion coefficient tensor can deviate significantly from this behaviour. In this talk, the complex dependencies of charged particle diffusion on the turbulence level of the magnetic field are presented. Examples of how this affects observational signatures will be shown in the context of galaxies or the transient sky, i.e., flaring Blazars.

T 73.4 Wed 17:00 T-H32

Perception of arrival direction maps of cosmic rays — •Edyvania Emily Pereira Martins¹, Markus Roth², and Darko Veberič² — ¹Institut für Experimentelle Teilchenphysik - Karlsruher Instituts für Technologie, Karlsruhe, Germany — ²Institut für Astroteilchenphysik - Karlsruher Instituts für Technologie, Karlsruhe, Germany

The processing of visual information in the human brain is guided by identifying colors and patterns formed by same-color areas. In cosmic-ray research, two main interests are to relate detected events to their sources and to identify excess regions in the sky. In this pursuit, maps of arrival directions are a commonly used tool. Depending on the choice of the smoothing applied to the maps, the plotting can render different locations of the perceived flux-excess regions on the map, represented in a different color to the flux-deficit regions. In addition, the most commonly used smoothing fabricates structures that are not real nor significant, and can lead to misinterpretation. An alternative to the standard, top-hat smoothing is presented, which facilitates the interpretation of data.

T 73.5 Wed 17:15 T-H32

Self-confinement of low-energy cosmic rays around supernova remnants -•HANNO JACOBS, PHILIPP MERTSCH, and VO-HONG MINH PHAN — TTK RWTH Aachen

Supernova Remnants have long been considered as promising candidate sources for cosmic rays. However, modelling the transport around these sources is difficult due to its nonlinear nature. The strong overdensity in the near source region leads to the production of plasma turbulence, upon which the particles scatter. To calculate this mechanism, called self-confinement, requires the numerical solution of two coupled differential equations describing the transport of particles and waves. Here, this formalism is extended to energies below 10 GeV, where energy losses become relevant. Particles around 100 MeV are found to be confined for in between 300 kyr and 1 Myr, depending on the interstellar medium. The diffusion coefficient is initially suppressed by up to two orders of magnitude. Interestingly, the spectrum outside the supernova flattens below 1 GeV at later times, similar to the spectral behaviour observed by Voyager. Furthermore, the grammage accumulated in the near source region is found to be non-negligible, which could be important for precision fitting cosmic ray spectra.

T 73.6 Wed 17:30 T-H32

No longer ballistic, not yet diffusive-the formation of cosmic ray smallscale anisotropies — •MARCO KUHLEN, VO HONG MINH PHAN, and PHILIPP MERTSCH — TTK Institut, RWTH Aachen

In the standard picture of cosmic ray transport the propagation of charged cosmic rays through turbulent magnetic fields is described as a random walk with cosmic rays scattering on magnetic field turbulence. This is in good agreement with the highly isotropic arrival directions as this diffusion process effectively isotropizes the cosmic ray distribution. However, high-statistics observatories like IceCube and HAWC have observed significant deviations from isotropy down to very small angular scales. This is in strong tension with this standard picture of cosmic ray propagation. By relaxing one of the assumptions of quasilinear theory and explicitly considering the correlations between the fluxes of cosmic rays from different directions, we show that power on small angular scales

is a generic feature of particle propagation through turbulent magnetic fields. We present a first analytical calculation of the angular power spectrum assuming a physically motivated model of the magnetic field turbulence and find good agreement with numerical simulations. We argue that in the future, the measurement of small-scale anisotropies will provide a new window to testing magnetic turbulence in the interstellar medium.

T 73.7 Wed 17:45 T-H32

Non-thermal ion acceleration at highly oblique non-relativistic shocks -•NAVEEN KUMAR and BRIAN REVILLE - Max-Planck Institute for Nuclear Physics Heidelberg, Germany

Non-thermal acceleration of particles (both electrons and ions) at an oblique, non-relativistic shock is demonstrated by using one-dimensional large-scale particle-in-cell simulations. Our results show the generation of non-thermal ions at highly oblique shocks with acceleration efficiencies of ~ 5% measured at the end of simulation runs. These results have important implications for understanding the non-thermal radiation generation at astrophysical sites such as supernova remnants.

T 73.8 Wed 18:00 T-H32

Can superbubbles accelerate PeV protons? — • THIBAULT VIEU and BRIAN REvILLE - Max-Planck-Institut für Kernphysik, Postfach 10 39 80, 69029 Heidelberg, Germany

The local cosmic-ray spectrum and recent gamma-ray observations suggest the existence of Galactic sources able to accelerate protons up to at least several PeV. These sources are still to be identified. Standard scenarios of particle acceleration at isolated supernova remnants struggle to reach PeV bands. However, most massive stars are not isolated but clustered. Clustered stars heat their surrounding medium, which inflates a cavity called a superbubble. In the superbubble, the stellar feedback creates multiple shocks, a turbulent environment, and amplifies the magnetic fields. These are ideal conditions for particle acceleration and superbubbles have long been thought to accelerate PeV protons. While it is indeed expected that an extended and strongly turbulent source could accelerate protons up to tens of PeV, it is yet unclear how the different acceleration processes can act collectively in superbubbles.

In this work we estimate the maximum energy of protons accelerated in superbubbles, considering various detailed scenarios. We derive under which circumstances PeV protons are expected. The forward shock of the superbubble barely accelerates particles up to 100 TeV. Supernova remnants expanding in the interior, or the collective wind termination shock which forms around a compact cluster, barely accelerate PeV protons. We show that protons of several PeV are only expected within loose and extended stellar clusters.

T 74: Neutrino Physics without Accelerators 5

Time: Wednesday 16:15-18:20

Group Report

T 74.1 Wed 16:15 T-H33 CEvNS and new neutrino physics searches with the CONUS experiment •EDGAR SANCHEZ GARCIA for the CONUS-Collaboration — Max-Planck-Institut für Kernphysik (MPIK), Heidelberg

The CONUS experiment (COherent elastic NeUtrino nucleus Scattering) aims to detect coherent elastic neutrino-nucleus scattering (CEvNS) of reactor antineutrinos on germanium nuclei in the fully coherent regime. The CONUS experiment - operational since April 2018 - is located at a distance of 17m from the 3.9 GWth core of the Brokdorf nuclear power plant (Germany). The possible CEvNS signature is measured by four 1 kg point-contact high-purity germanium (HPGe) detectors, which provide a sub keV energy threshold with background rates in the order of 10 events per kg, day and keV. The analysis of the first CONUS data set allows to establish the current best limit on CEvNS from a nuclear reactor with a germanium target. Moreover, competitive limits on neutrino physics beyond the standard model can be set such as on non-standard neutrino interactions or on the neutrino electromagnetic properties. These results will be presented in this talk together with the CONUS long-term operation.

T 74.2 Wed 16:35 T-H33

Novel constraints on neutrino physics beyond the standard model from the CONUS experiment — • THOMAS RINK for the CONUS-Collaboration — Max-Planck-Institut für Kernphysik (MPIK)

The detection of coherent elastic neutrino-nucleus scattering (CEvNS) opens up new opportunities for neutrino physics within and beyond the standard model. As a leading reactor experiment, CONUS aims for an observation in the fully coherent regime with antineutrinos emitted from the powerful $3.9\,\mathrm{GW}_{\mathrm{th}}$ reactor of the nuclear power plant in Brokdorf (Germany). In particular, the application of ultra-low threshold, high-purity germanium detectors within a compact shield design in close vicinity to a nuclear reactor core describes the next milestone towards high-statistics neutrino physics. The acquired and future CONUS

data sets allow investigations of yet undetected neutrino channels and electromagnetic properties. Moreover, measurements of the Weinberg angle with neutrinos at the MeV-scale and analyses of the high energy part of a reactor's antineutrino emission spectrum become possible with CEvNS. This talk deals with constraints on beyond the standard model neutrino phenomenology that are obtained from data collected between April 2018 and June 2019. Bounds on nonstandard neutrino-quark interactions of vector and tensor type from CEvNS are presented. Further, the parameter space of simplified scalar and vector mediators that is probed by CEvNS and elastic neutrino-electron scattering is discussed. Finally, limits on an effective neutrino magnetic moments and an effective neutrino millicharge are given.

T 74.3 Wed 16:50 T-H33 Pulse Shape Discrimination for the CONUS experiment — • JOSEF STAUBER for the CONUS-Collaboration - Max-Planck-Institut für Kernphysik (MPIK), Heidelberg, Germany

The CONUS experiment, located at the nuclear power plant in Brokdorf, Germany, aims at the detection of coherent elastic neutrino nucleus scattering (CEvNS) in the fully coherent regime. Four 1kg-sized point-contact germanium detectors are used for this purpose. The suppression of the background and a very low energy threshold are crucial for the successful detection of CEvNS. The pulse shape discrimination PSD offers a tool to reduce the background by analysing the shapes of the individual events. The data acquisition module (DAQ) can alter the pulse shape (electric feedback) and add electrical noise to the signal. In this talk the concept of the PSD will be presented with special focus on the DAQ feedback and the impact of electrical noise in the low energy regime.

Location: T-H33

T 74.4 Wed 17:05 T-H33

Shield and detector optimization for the CONUS experiment - •JANINA HAKENMÜLLER for the CONUS-Collaboration — Max-Planck-Institut für Kernphysik, Saupfercheckweg 1, 69115 Heidelberg, Germany

The CONUS experiment is looking for coherent elastic neutrino nucleus scattering (CEvNS) with four low energy threshold point contact high-purity Ge spectrometers. The CONUS Collaboration is exploring options for a new reactor site outside Germany to continue and extend its scientific program. In the talk, the possibilities of improvement regarding the shield and detector design will be explored. The strong neutrino flux required for a detection is provided by a nuclear power plant, where only a small overburden to shield against cosmic rays is available. The current CONUS shield employs 25 cm of lead to shield against environmental gamma rays. Cosmic ray muons create secondaries inside the lead especially neutrons, that can induce CEvNS like recoils in the detector. Even though the background is suppressed by an active muon veto, it still contributes significantly to the measured background. With the help of MC simulations improvements and alternatives of the shield design are examined. Additionally, also the impact of a larger crystal size on the background as well as potentially enhancements of detector-specific background will be discussed.

T 74.5 Wed 17:20 T-H33

Simulation of Pulse Shapes for Low Background Germanium Spectrometers in CONUS — • JANINE HEMPFLING for the CONUS-Collaboration — Max-Planck Institut für Kernphysik (MPIK), Heidelberg

The CONUS experiment aims to detect coherent elastic neutrino nucleus scattering (CEvNS). For this goal four 1 kg point-contact high-purity germanium detectors are operated near the $3.9\,\mathrm{GW}_{\mathrm{th}}$ core of the Brokdorf nuclear power plant. A very good background suppression is crucial for the success of the experiment, achieved by an elaborate shield. A new opportunity for additional background reduction is offered by pulse shape analysis of the detector signals. To verify this analysis a pulse shape simulation (PSS) for the CONUS experiment is developed based on the software package SigGen. Additionally, investigations of the correlation between the signal shape and the interaction position as well as the fraction of single-site events and multi-site events are possible with the PSS. This talk presents the requirements needed to set up a PSS, starting from the input signal generation with the GEANT4 framework MaGe to the modeling of the response of the electronics up to the final output pulses. Furthermore, a comparison between the obtained simulation results and the measured signals will be discussed.

T 74.6 Wed 17:35 T-H33

The CRAB Experiment: a New Calibration Technique for CEvNS Detectors in the 100 eV Regime - • VICTORIA WAGNER for the CRAB-Collaboration Physik-Department, Technische Universität München, D-85748 Garching, Deutschland

Searches for light dark matter (DM) and studies of coherent elastic neutrinonucleus scattering (CEvNS) imply the detection of nuclear recoils in the 100 eV range. However, an absolute energy calibration in this regime is yet missing. The CRAB project proposes a method based on nuclear recoils induced by the emission of an MeV-gamma following thermal neutron capture. Detailed feasibility studies show that this method yields distinct nuclear recoil calibration peaks at 112 eV and 160 eV for tungsten. In the first phase, the CRAB project foresees to perform a nuclear recoil calibration of cryogenic CaWO₄ detectors read-out by TES. The low power TRIGA reactor in Vienna provides a clean beam of thermal neutrons well suited for such a measurement. In the second phase, additional

tagging of the photons produced in the de-excitation process will allow to extend the calibration method to even lower energies and to a wider range of detector materials, such as Ge. Combined with an electron recoil calibration, CRAB will allow to measure energy quenching in the sub-keV regime. With its novel idea, CRAB provides a direct and accurate calibration of nuclear recoils in the ROI of light DM and future CEvNS experiments, which is essential for studying new physics.

This work is supported by the DFG through the Excellence Cluster ORIGINS and the SFB1258.

T 74.7 Wed 17:50 T-H33

Development oft the cryogenic detector for observing coherent elastic neutrino nucleus scattering with NUCLEUS $10g - \cdot$ Nicole Schermer¹, Andreas Erhart¹, Dieter Hauff^{1,2,3}, Margarita Kaznacheeva¹, Tobias $Ortmann^1, Luca Pattavina^1, Johannes Rothe^1, Raimund Strauss^1, Victoria Wagner^1, and Alexander Wex^1 — ^1Technische Universität München,$ Physik Department Lehrstuhl E15, James-Franck-Straße 1, D-85748 Garching

²Max-Planck-Institut für Physik, Föhringer Ring 6, D-80805 München — ³Universität Tübingen, Physikalisches Institut, Auf der Morgenstelle 14, D-72076 Tübingen

The study of coherent elastic neutrino nucleus scattering (CEvNS) offers the opportunity to explore fundamental neutrino properties and to search for physics beyond the Standard Model. The NUCLEUS experiment aims to precisely measure the CEvNS cross-section of reactor-antineutrinos produced by the Chooz nuclear power plant in France at low energies with gram-scale cryogenic detectors with ultra-low energy thresholds of O(10eV). The first science phase, NU-CLEUS 10g, will deploy two detector modules, each containing nine cryogenic target detectors embedded in an inner veto. I will present the development of the first versions of the NUCLEUS 10g detector components as well as further strategies towards the full cryogenic detector, which will consist of the two detector modules, a LED calibration system and a cryogenic outer veto. The research was supported by the DFG through the Excellence Cluster ORIGINS and the SFB1258, and the ERC Starting Grant 2018 "NU-CLEUS".

T 74.8 Wed 18:05 T-H33

Exploring Coherent Elastic Neutrino-Nucleus Scattering at a nuclear reactor with the NUCLEUS experiment — •JOHANNES ROTHE for the NUCLEUS-Collaboration - Physik-Department, Technische Universität München, D-85748 Garching, Germany

The NUCLEUS collaboration is working towards the first detection of reactor antineutrinos via Coherent Elastic Neutrino-Nucleus Scattering (CEvNS) using cryogenic detectors operating at 15mK temperature. This observation will open a new window to study fundamental neutrino properties at low energy with a high-flux source.

The first physics phase will employ a 10g target made of Al₂O₃ and CaWO₄ crystals read out with superconducting transition edge sensors, surrounded by cryogenic infrastructure, passive shielding and active anti-coincidence veto systems. Assembly and commissioning of all components of the experimental setup is planned at TUM in 2022. Afterwards, the experiment will be moved to the Chooz Nuclear Power Plant in France. I will present updates on the design and simulation of the setup, experimental work towards the 10g target detector and the physics goals of NUCLEUS-10g. This research was supported by the DFG through the Excellence Cluster ORIGINS and the SFB1258, and the ERC Starting Grant 2018 "NU-CLEUS".

T 75: Neutrino Physics without Accelerators 6

Time: Wednesday 16:15-18:35

Group Report

T 75.1 Wed 16:15 T-H34 Overview of LEGEND and the Commissioning Status of LEGEND-200 -•PATRICK KRAUSE for the LEGEND-Collaboration — Physik-Department, Technische Universität München, Garching

The discovery that neutrinos are Majorana fermions would have profound implications for particle physics and cosmology. The Majorana character of neutrinos would make neutrinoless double-beta $(0\nu\beta\beta)$ decay, a matter-creating process without the balancing emission of antimatter, possible. The LEGEND Collaboration pursues a phased, ⁷⁶Ge-based double-beta decay experimental program with discovery potential covering the inverted hierarchy. The first phase, LEGEND-200, will have a discovery potential of 10^{27} years and a background index of 0.6 cts/(ROI t yr). The second phase, LEGEND-1000, will deploy 1000 kg of enriched germanium and will have a discovery sensitivity beyond 10²⁸ years. This talk will give an overview of LEGEND and will report on the currently ongoing commissioning work in LEGEND-200. This research is supported in part by the BMBF through the Verbundforschung 05A2020, the MPG, the DFG through

the Excellence Cluster ORIGINS and the SFB1258, and the through the ERC Advanced Grant 786430 - GemX

T 75.2 Wed 16:35 T-H34

Location: T-H34

ASIC based readout electronics for high-purity Germanium detectors in LEGEND-1000 — •FLORIAN HENKES, FRANK EDZARDS, SUSANNE MERTENS, and MICHAEL WILLERS for the LEGEND-Collaboration — Technische Universität München, München, Deutschland

The Large Enriched Germanium Experiment for Neutrinoless $\beta\beta$ Decay (LEG-END) is a ton-scale, ⁷⁶Ge-based, neutrinoless double-beta ($0\nu\beta\beta$) decay experimental program with discovery potential at half-lifes greater than 10²⁸ years.

Signal readout electronics are essential in order to achieve the experiment's sensitivity on $0\nu\beta\beta$ -decay. The close proximity to the detectors poses unique challenges to balance electronic performance with radiopurity requirements. In LEGEND-200, the readout system consists of a charge sensitive amplifier realised from discrete components with an ultra radiopure first stage close to the detectors and a second stage from less radiopure commercial components. In LEGEND-1000, the use of Application-Specific Integrated Circuit (ASIC) technology would allow to implement the entire charge sensitive amplifier into a single low-mass chip with ultimate electronic noise performance and signal fidelity while ideally further reducing backgrounds.

In this contribution, the current status of the LEGEND-1000 ASIC based readout development at the Technical University of Munich will be presented and prospects for future developments of ASIC based charge sensitive amplifiers for high-purity germanium detectors will be discussed.

T 75.3 Wed 16:50 T-H34 Constraining the $^{77(m)}$ Ge Production with GERDA Data and Implications for LEGEND-1000 — •MORITZ NEUBERGER¹, LUIGI PERTOLDI¹, STEFAN SCHÖNERT¹, and CHRISTPH WIESINGER^{1,2} for the GERDA-Collaboration — ¹Physik-Department E15, Technische Universität München — ²Max-Planck-Institut für Physik (Werner-Heisenberg-Institut) Föhringer Ring 6 80805 München

The delayed decay of ^{77(m)}Ge, produced by neutron capture on ⁷⁶Ge, is a potential background for the next-generation neutrino-less double-beta decay experiment LEGEND-1000, especially when considering the alternative LNGS site. Based on Monte Carlo simulations, various mitigation strategies and suppression techniques have been proposed to tackle this background [1,2]. In this talk we will present first results on ^{77(m)}Ge searches in the full GERDA data. Given the very similar configuration - bare germanium detectors in liquid argon - it serves as a benchmark for our LEGEND-1000 predictions. This research was supported by the BMBF through the Verbundforschung 05A20WO2 and by the DFG through the SFB1258 and Excellence Cluster ORIGINS.

[1] C. Wiesinger et al., Eur. Phys. J. C (2018) 78: 597

[2] LEGEND-1000 pCDR, arXiv 2107.11462

T 75.4 Wed 17:05 T-H34

First light in LEGEND-200 — •ROSANNA DECKERT, PATRICK KRAUSE, LASZLO PAPP, and STEFAN SCHÖNERT — Technische Universität München

LEGEND (Large Enriched Germanium Experiment for Neutrinoless $\beta\beta$ Decay) is a ton-scale experiment to search for neutrinoless double beta ($0\nu\beta\beta$) decay using high-purity germanium detectors enriched in ⁷⁶Ge. An observation of $0\nu\beta\beta$ decay would prove the existence of lepton number violation and provide insight into the nature of neutrino masses. The first phase of the experiment LEGEND-200 will deploy 200 kg of enriched material and aims for a sensitivity of 10^{27} years on the $0\nu\beta\beta$ half-life. To achieve this, the germanium detectors will be operated in liquid argon (LAr), instrumented as an active detector to detect the scintillation light produced by backgrounds from trace radioactive contaminants.

Commissioning of the LAr instrumentation, consisting of wavelength-shifting fibers, a wavelength-shifting reflector and silicon photomultiplier arrays, started in August 2021 at the Laboratori Nazionali del Gran Sasso. In this talk, the analysis of the first LAr commissioning data for LEGEND-200 will be presented.

This research is supported by the BMBF through the Verbundforschung 05A20WO2 and by the DFG through the Excellence Cluster ORIGINS and the SFB1258.

T 75.5 Wed 17:20 T-H34

New Limit on the radiative 0ν ECEC of 36 Ar from GERDA Phase II Data — •MICHELE KOROSEC¹, ELISABETTA BOSSIO¹, and CHRISTOPH WIESINGER² for the GERDA-Collaboration — ¹Physik-Department, Technische Universität München — ²Max-Planck-Institut für Physik, München

Neutrinoless double electron capture (0vECEC) is a theoretically possible decay that would prove lepton number violation, which is forbidden by the Standard Model of Physics, and therefore provide evidence for the Majorana nature of neutrinos. The GERmanium Detector Array (GERDA) experiment offers the possibility to search for 0vECEC of ³⁶Ar which is one of the isotopes that theoretically allows this rare decay.

A search for neutrinoless double electron capture of ³⁶Ar was conducted based on Phase II data from the GERDA experiment, located at the Gran Sasso Laboratory of INFN, Italy. A simultaneous fit to multiple datasets has been performed in which no signal for the decay has been observed. However, a new, preliminary experimental lower limit on the half-life of 0*v*ECEC in ³⁶Ar has been established with the CLs method at $T_{1/2} > 1.22 \cdot 10^{22}$ yr (90% C.L.) which will take over from the previous best limit of $T_{1/2} > 3.6 \cdot 10^{21}$ years (90% C.I.) [1] which was found in GERDA Phase I.

This research is supported by the BMBF through the Verbundforschung 05A20WO2 and by the DFG through the Excellence Cluster ORIGINS and the SFB1258. [1] GERDA Collaboration, Eur.Phys.J.C 76 (2016) 12, 652

T 75.6 Wed 17:35 T-H34

In-situ characterization of germanium detectors from ³⁹Ar decays for lowenergy data modeling in GERDA and LEGEND — •LUIGI PERTOLDI for the GERDA-Collaboration — TU München, Germany

A reliable estimate of the active volume of high-purity germanium (HPGe) detectors, defined as the internal volume in which charge collection efficiency (CCE) reaches its maximum, is a fundamental piece of a detector's response model. Typical HPGe detectors feature a null or incomplete CCE in correspondence with the lithium-doped, high-voltage bias contact. In this contribution, a new method for determining the active volume of HPGe detectors immersed in liquid argon (LAr), will be presented. The method exploits the shape of the low-energy distribution of 39 Ar β^- decays, naturally occurring in atmospheric LAr and recorded by the detectors, which strongly depends on the CCE profile. The technique is applied to physics data by the GERDA experiment and used to characterize the deployed detectors in-situ. As a consequence, a first model of the low-energy data spectrum recorded by the experiment will also be shown. The developed technique will be useful for the LEGEND experiment, which aims to perform searches of new-physics phenomena at low energies. Moreover, by using these novel ³⁹Ar-based active volume estimates, we aim to obtain a precise and unbiased estimate of the two-neutrino double-beta decay rate of ⁷⁶Ge. This research is supported by the BMBF through the Verbundforschung 05A20WO2 and by the DFG through the Excellence Cluster ORIGINS and the SFB1258.

T 75.7 Wed 17:50 T-H34

New results on the ⁷⁶Ge double-beta decay with neutrinos and exotic decay modes from GERDA Phase II — •ELISABETTA BOSSIO for the GERDA-Collaboration — Physik Department, Technische Universität München, Garching, Germany

Two-neutrino double beta $(2\nu\beta\beta)$ decays are amongst the rarest nuclear processes ever observed. Precision studies of the electron sum energies require ultralow background and an excellent understanding of the experiment's response. Both are key features of the Germanium Detector Array (GERDA) experiment, which searched for neutrino-less double beta $(0\nu\beta\beta)$ decay with enriched high purity germanium detectors in Liquid Argon at Laboratori Nazionali del Gran Sasso (LNGS) in Italy. The measurement of the Standard Model $2\nu\beta\beta$ decay half-life of ⁷⁶Ge was performed with unprecedented precision, profiting from the high signal-to-background ratio and the small systematic uncertainties. It provides essential inputs for nuclear structure calculations, that benefit the interpretation of $0\nu\beta\beta$ decay results. Furthermore, the search for distortions of the $2\nu\beta\beta$ decay spectrum allows exploring new physics, like $0\nu\beta\beta$ decay with Majorons emission, Lorentz invariance, or search for sterile neutrinos. The new results of the ⁷⁶Ge $2\nu\beta\beta$ decay half-life and improved limits on exotic decay modes will be presented in this talk.

This research is supported by the BMBF through the Verbundforschung 05A20WO2 and by the DFG through the Excellence Cluster ORIGINS and the SFB1258.

T 75.8 Wed 18:05 T-H34

Double weak decays of $^{124}\rm Xe$ and $^{136}\rm Xe$ in XENON1T and XENONNT — •CHRISTIAN WITTWEG for the XENON-Collaboration — Physik-Institut, Universität Zürich

In recent years xenon-based dark matter direct detection experiments have reached large enough target masses and low enough background levels to also probe rare double weak decays. Among these decays are the two-neutrino double electron capture (2 ν ECEC) of ¹²⁴Xe as well as the neutrinoless double beta decay (0 $\nu\beta\beta$) of ¹³⁶Xe. Observation of the hypothetical neutrinoless decay would provide definite proof of the neutrino's Majorana nature and indicate lepton number violation. The measurement of the Standard Model 2 ν ECEC – first detected by XENON1T in 2018 – provides nuclear structure information that is a crucial input for the nuclear models used to interpret 0 $\nu\beta\beta$ experiments. This contribution will present the ¹²⁴Xe 2 ν ECEC results and search for 0 $\nu\beta\beta$ of ¹³⁶Xe in XENON1T. Moreover, the sensitivity projection for a ¹³⁶Xe 0 $\nu\beta\beta$ search in XENONnT will be outlined.

T 75.9 Wed 18:20 T-H34

Latest results from XENON1T and prospects for XENONnT for $0\nu\beta\beta$ – •Tim Michael Heinz Wolf for the XENON-Collaboration — MPIK, Heidelberg, Deutschland

XENON1T was a dual-phase xenon time projection chamber (TPC) looking mainly for the direct detection of WIMP dark matter at energy depositions of up to tens of keV. This talk will focus on the high energy part of the spectrum (above hundreds of keV) where we can search for neutrinoless double beta decay ($0\nu\beta\beta$). This is a hypothetical process implying two beta decays without emitting any neutrinos. The natural isotope ¹³⁶Xe (abundance approximately 8.9%) is a candidate isotope to look for $0\nu\beta\beta$ which allows to carry out a search in LXe TPCs. This process tests lepton flavour conservation and it is potentially able to shed light on the nature of neutrinos (Dirac or Majorana particles) or give hints on the mass ordering of neutrinos. We present for the first time, a search for $0\nu\beta\beta$ where a limit on the $0\nu\beta\beta$ half-life of ¹³⁶Xe with XENON1T data is derived, and give prospects on the performance of XENONnT data.

T 76: Search for Dark Matter 3

Time: Wednesday 16:15–18:30

T 76.1 Wed 16:15 T-H35

Search for dark matter production in association with a single-top and wboson with the ATLAS experiment — •ALVARO LOPEZ SOLIS — DESY-Zeuthen, Zeuthen, Germany

Measurements at large scales suggest that Dark Matter (DM) constitutes around 27% of all the energy available in the Universe and around 85% of all the available mass. However, its nature remains a mystery. Several theories try to address this problem by suggesting the existence of new weakly interacting particles that would constitute most of this new type of energy. This talk will present a search using the ATLAS experiment at LHC for these weakly interacting particles. It is motivated by the hypothesis that DM particles would couple to the Standard Model (SM) particles via a pseudo-scalar mediator within an extended two-Higgs-doublet model (2HDM+a). Among all the possible signatures predicted by this model, this talk will present the search for DM production in association to a single-top quark and a W-boson in channels where both, top and W-boson, are assumed to decay hadronically (0L channel) or either one of them decay leptonically (1L channel).

T 76.2 Wed 16:30 T-H35

Search for Dark Matter at the ATLAS detector with a W-boson and a topquark in the final state — •PAUL MODER¹, ALVARO LOPEZ SOLIS², BEN BRÜERS², and CLAUDIA SEITZ¹ — ¹DESY Hamburg — ²DESY Zeuthen

The Standard Model (SM) is one of the most robust models in particle physics containing all observed elementary particles and their interactions. Over the years, its predictions were tested and proven in a number of experiments. However, there are still observations that can not be explained by the SM with one of the most prominent ones being the existence of Dark Matter (DM). While the existence of DM was first theorised through astronomical observations, extensions of the SM allow for a search of DM at the Large Hadron Collider (LHC) as well. Since DM can not be detected directly, final states analysing its existence at the LHC are always designed around high missing transverse energy. This talk will present such a search at the ATLAS detector based on an extended two-Higgs-doublet model (2HDM+a) where the pseudo-scalar mediator allows the production of DM in the final state. In addition to the DM, a top-quark and a Wboson are produced in the final state, where the W-boson of this signal process can be expected to have high momentum. This allows for a unique technique by tagging these W-bosons through large-radius jets increasing the sensitivity of the signal process. This talk presents the cut-based definitions for an analysis with zero leptons in the final state as well as the most recent results.

T 76.3 Wed 16:45 T-H35

Search for Dark Matter in a tW+MET signature with the ATLAS experiment — •BEN BRÜERS — Deutsches Elektronen Synchrotron DESY, Zeuthen, Germany

Dark Matter (DM) remains one of the unrevealed mysteries of the universe. Even though it constitutes ~ 80% of the matter, considerably little is known about DM, despite it significantly influences the dynamics of galaxies and the expansion of the universe. The search for DM at colliders, probing mainly a particle nature of the unknown matter, marks an important pillar in exploring all possible realisations of DM. This talk will present a search for DM with the ATLAS experiment, where the DM is coupled to the Standard Model (SM) via a pseudo-scalar mediator within an extended two-Higgs-doublet model (2HDM+a). The associated production of DM with a W-boson and a top-quark is considered. As the DM deposits no energy in the detector, the experimental signature includes high missing transverse energy ($E_{\rm T}^{\rm miss}$). To reconstruct highly energetic W-bosons, expected for signals with a heavy H⁺, large-radius jets are employed. The talk will give an introduction to the analysis and present the most recent results.

T 76.4 Wed 17:00 T-H35

Search for dark matter produced in association with two top quarks and missing energy in the final state using ATLAS 13 TeV pp collision data — •MARCO RIMOLDI — DESY, Hamburg, Germany

The hypothesis of the existence of non-baryonic dark matter (DM) comes from gravitational evidence across a wide range of astrophysical and cosmological systems. Of the many types of DM candidate proposed, weakly interacting massive particles (WIMP) are believed to be a theoretically convincing candidate. WIMPs must interact weakly with electromagnetic radiation and be consistent with the expected DM density. If WIMPs are the manifestation in nature of DM, then it may be possible to produce it directly at the LHC.

Results of the combination of four analyses are presented, selecting final state events with two top quarks and invisible particles. Proton proton collisions data collected by the ATLAS experiment at a centre-of-mass energy of 13 TeV during the Run-2 data-taking are used. Results are interpreted in terms of dark matter simplified models considering a spin-0 mediator to dark sector. Upper limits on the Higgs boson invisible branching ratio, where the Higgs is produced according to the Standard Model in association with a pair of top quarks are also reported.

T 76.5 Wed 17:15 T-H35

Search for new physics in $t\bar{t}$ + \mathbf{E}_T^{miss} final states in pp collisions at 13 TeV with the ATLAS experiment. - •SIMRAN GURDASANI — Albert-Ludwigs-Universität, Freiburg, Germany

This talk will present the developments of an ongoing search for Beyond Standard Model (BSM) signatures that can be probed using the $t\bar{t}+E_T^{miss}$ final state at the Large Hadron Collider (LHC). Neural Networks are used for the search which is performed on data collected with the ATLAS detector between 2015 and 2018, corresponding to 139 fb^{-1} of pp data at 13 TeV. Models specifically targeted include DM production via scalar or pseudo-scalar mediators, SUSY stop pair production and Higgs decays to new invisible particles. A two-fold implementation of neural nets is designed, where the first step aims to efficiently reconstruct the hadronically decaying top quarks in a given event. This is designed to specifically target mid-pt range tops decaying to resolved jets. The second step aims to exploit full kinematic correlations of the $t\bar{t}+E_T^{miss}$ system and tag a given event to one of the targeted BSM processes while providing background rejection against both major (ttbar and Wjets) and non-major SM processes. The talk will give an overview of the strategy developed and the status of ongoing optimization studies.

T 76.6 Wed 17:30 T-H35

Combining Dark Matter searches with top quarks with the ATLAS detector — •MARIANNA LIBERATORE — Deutsches Elektronen-Synchrotron DESY, Hamburg and Zeuthen, Germany

A motivation to WIMP dark matter (DM) searches at the LHC, and in particular in ATLAS, is the especially promising possibility that interactions between ordinary matter and DM are mediated by new spin-0 particles. Such particles would extend the SM with a potential dark sector, to which DM particles belong. Similarly to the Higgs boson, these new mediators interact strongest with the heaviest particles via Yukawa-type couplings, making them more prone to associated production with heavy-flavour quarks.

To test those models, two recently released search channels are considered within ATLAS: DM with top quark pairs[1] or a single top quark[2], with a focus on the two charged leptons final states. This talk will motivate how the statistical combination of these two results in simplified models could significantly enhance the sensitivity to DM signals, and the first results of these combined studies will be presented.

[1] JHEP04(2021)165

[2] Eur.Phys.J.C(2021)81:860

T 76.7 Wed 17:45 T-H35

Searching for Dark Matter in top quark production with the CMS experiment — DANYER PEREZ ADAN, AFIQ ANUAR, ALEXANDER GROHSJEAN, LAURIDS JEPPE, JONAS RÜBENACH, CHRISTIAN SCHWANENBERGER, •DOMINIC STAFFORD, and NICOLE STEFANOV — Deutsches Elektronen-Synchrotron DESY, Hamburg, Germany

Astronomical observations provide strong evidence that a large proportion of the matter in the universe is "Dark Matter" (DM), which is currently not included in the Standard Model (SM) of particle physics. Furthermore, many cosmological models suggest Dark Matter should couple to the SM around the 100 GeV scale, and hence may be produced at the LHC, appearing as missing transverse momentum. We present a search for Dark Matter produced in association with top quarks, via a spin-0 mediator, with a focus on the dileptonic channel. This analysis will be part of the upcoming CMS result with the full Run-2 dataset, and will be the first to combine the top quark pair + DM and single top + DM processes for dileptonic, semileptonic and full hadronic final states, which greatly aids sensitivity to the highest mediator masses in the search.

The dileptonic channel poses an interesting challenge due to a large amount of missing transverse momentum in the SM $t\bar{t}$ background, and an irreducible $t\bar{t}Z(Z \rightarrow \nu\nu)$ background. This analysis therefore uses novel variables and machine learning techniques in the signal extraction, and new control regions to constrain the irreducible backgrounds.

T 76.8 Wed 18:00 T-H35

Performance of different MET reconstruction methods in a monotop DM analysis — •JOST VON DEN DRIESCH¹, SEBASTIAN WIELAND¹, MICHAEL WASSMER¹, NIKITA SHADSKIY¹, ULRICH HUSEMANN¹, MATTEO CREMONESI², LINDSEY GRAY³, and YIHUI LAI⁴ — ¹Institute of Experimental Particle Physics (ETP), Karlsruhe Institute of Technology (KIT) — ²University of Notre Dame (ND) — ³Fermilab — ⁴University of Maryland (UMD)

Missing transverse momentum (MET) is an important quantity in many analyses at hadron colliders. Especially Dark Matter (DM) analyses often make use of this

Location: T-H35

T 76.9 Wed 18:15 T-H35

quantity as DM particles leave the detector without interactions and therefore create large amounts of MET. However, due to its origin from non-detectable particles, MET cannot be measured directly, but must be estimated from the transverse momentum of all reconstructable particles.

Over the years, various MET reconstruction methods have been developed and applied at CMS. The latest approaches use machine learning methods, e.g. Convolutional Neural Networks (DeepMET) or Graph Neural Networks (GraphMET). Monte Carlo studies show an improvement of MET reconstruction performance by these novel reconstruction methods compared to the older ones. Yet, it remains unclear how large this effect will be in a full analysis.

This talk will introduce the aforementioned MET reconstruction methods and compare their expected impact on a monotop analysis, aimed at the search for Dark Matter in events with a single top quark and large MET.

Search for axion-like particles (ALPs) at Belle II experiment. — •AWAIS BIN ZAHID, PABLO GOLDENZWEIG, and TORBEN FERBER — Karlsruhe Institute of Technology (KIT), Karlsruhe, Germany

One possible extension of the Standard Model (SM), which may contribute in solving the mystery of Dark Matter (DM) and/or explain some astrophysical anomalies, are Axion-Like Particles (ALPs). The model taken into consideration in this search is of an ALP interacting with SM photons with a coupling strength $g_{a\gamma\gamma}$ and having mass m_a . The search for the direct production of such an ALP via the process (ALP-Strahlung) $e^+e^- \rightarrow \gamma a(a \rightarrow \gamma\gamma)$, is performed in the mass range $0.2 < m_a < 9.8 \text{ GeV}/c^2$. Given that the final state of the $e^+e^- \rightarrow \gamma a(a \rightarrow \gamma\gamma)$ process is fully neutral, being made up of three photons, a proper kinematic fit with neutral particles is a powerful tool to improve signal resolution. In this talk, I will present the status of sensitivity evaluation based on Monte Carlo simulation which corresponds to the data of almost 500 fb^{-1} that will be collected by Belle II at the end of summer 2022.

T 77: Search for Dark Matter 4

Time: Wednesday 16:15-18:15

background events caused by bursts.

T 77.1 Wed 16:15 T-H36

Light dark matter search using SuperCDMS single-charge-sensitive devices and anticoincident tagging - •ALEXANDER ZAYTSEV for the SuperCDMS-Collaboration - Karlsruher Institut für Technologie, Karlsruhe, Germany As a part of its R&D program, the SuperCDMS collaboration has been developing cryogenic gram-scale eV-resolution (HVeV) detectors that utilize Neganov-Trofimov-Luke amplification by applying a voltage bias across the Si crystals. During the two previous above-ground dark matter (DM) searches, each collecting data with only one HVeV detector, competitive constraints were obtained for DM-electron scattering, as well as for dark photon and axion-like particle absorption. However during the second HVeV run, a background characterized by bursts of single electron-hole pair events was observed, which may originate from luminescence in SiO2 - the primary component of the detector holder material. Single-pulse events from the tails of such bursts degrade the sensitivity of the HVeV DM searches in the entire mass range of interest. In the latest underground HVeV DM search (Run 3), we have collected O(10) gram-days of exposure using three HVeV detectors, operated simultaneously within a shared housing. We present the current progress of the respective DM search analysis, which is expected to surpass the previous HVeV DM limits by utilizing an interdetector anticoincidence event selection that considerably suppresses the rate of

T 77.2 Wed 16:30 T-H36 Characterizing bursts of single-electron-hole-pair events using a Super-CDMS R&D device and a sodium-22 source — •MATTHEW WILSON for the SuperCDMS-Collaboration — Karlsruher Institut für Technologie

Recently, R&D facilities within the SuperCDMS collaboration have developed and employed cryogenic, high-voltage, eV-scale (HVeV) detectors that have single-charge sensitivity. When a bias voltage is applied across these gram-sized, silicon detectors, the charge signals are amplified in the form of phonons, making the detectors sensitive to low-energy electron interactions. HVeV detectors have been previously utilized for two separate above-ground dark matter (DM) searches to set competitive constraints on low-mass DM candidates. However, these constraints are limited by the presence of an unknown, low-energy background, a large component of which appears to be bursts of single-electron-holepair events. One hypothesis is that these events originate from the photoluminescence of SiO₂, a primary component of the detector holder material. A sodium-22 source has been placed near an HVeV detector to determine whether such burst events are induced by the high-energy gammas emitted by a radioactive source, which would support this hypothesis. This presentation shows the latest results of the investigation and characterization of this low-energy background.

T 77.3 Wed 16:45 T-H36

SuperCDMS detector testing at the Cryogenic Underground TEst (CUTE) facility — •SUKEERTHI DHARANI for the SuperCDMS-Collaboration — Universität Hamburg

SuperCDMS SNOLAB is an upcoming direct dark matter search experiment using silicon and germanium detectors operated at cryogenic temperatures. The experiment is planned to start data taking in 2023 at SNOLAB which is located 2 kilometers underground in the Creighton mine in Canada. With a low background from cosmic sources, SNOLAB is ideal for rare event searches. The Cryogenic Underground TEst (CUTE) is a well-shielded test facility operating at SNOLAB with a measured background rate of ~ 7 events/keV/kg/day. It acts as a testbed for the SuperCDMS detectors and facilitates performing early science runs. In this talk, an overview of the CUTE facility's features, ongoing activities, and applications for the SuperCDMS experiment will be presented. Location: T-H36

T 77.4 Wed 17:00 T-H36

Current Status of the BRASS-P Experiment — •FAYEZ BAJJALI¹, LE HOANG NGUYEN¹, DIETER HORNS¹, ANDREI LOBANOV^{1,2}, ARTAK MKRTCHYAN¹, SVEN DORNBUSCH², CHRISTOPH KASEMANN², MARTIN TLUCZYKONT¹, and MARKO EKMEDŽIĆ¹ — ¹Institute of Experimental Physics - University of Hamburg — ²Max-Planck-Institute for Radio Astronomy - Bonn

Axions and Hidden Photons (HPs) are among the best motivated candidates for explaining the enigmatic nature of the dark matter. These weakly interacting slim particles (WISPs) have a small mass and can be detected via electromagentic (EM) radiation arising from their interaction with normal matter, photons, and magnetic field. The concept for Broadband Radiometric Axion/ALP Searches (BRASS) provides a pioneering experimental setup for WISP searches in the range of 10-10000 μ eV. The prototype setup BRASS-P is currently being constructed at the University of Hamburg. It combines permanently magnetized conversion panels producing the EM signal from passages of WISPs, a parabolic mirror focusing the EM signal, a cryogenic 12-18 GHz heterodyne receiver, and a broadband digitizing backend DBBC3 for detecting and processing the signal.

The structure of the conversion panels and the measurement of the static magnetic field will be presented. The setup and calibration procedures employed for the 12-18 GHz receiver and the DBBC3 digitizer will be discussed. Finally, preliminary results from the first science run carried out for searching for HPs in the frequency range of 12-16 GHz will be presented.

T 77.5 Wed 17:15 T-H36

Axion simulation in various geometry — •JOHANNES ULRICHS¹, LE HOANG NGUYEN¹, DIETER HORNS¹, and ANDREI LOBANOV^{1,2} — ¹Institut für Experimentalphysik, Universität Hamburg, Hamburg, Deutschland — ²Max-Planck-Institut für Radioastronomie, Bonn, Deutschland

Using commercial FEM software (COMSOL TM), we solve the Axion-Maxwell equation in the geometrical context of experiments that search for axion and axion-like-particles (ALPs) dark matter. Firstly, the BRASS-p is the pilot experiment that search for axion/alps in the frequency range of 12 - 18 GHz (49.63 - 74.4 μ eV). The multiphysics simulation (AC/DC and RF modules) is used to explore the realistic magnetic field of the magnet panels and the axion-induced radiation. Accompanied with further studies concerning the efficiency and coherence effect of the overall setup. Secondly, we consider the possibility of detecting the skin current induced by the low mass axion dark matter (few kHz to 3MHz, 4.14 peV - 12.4 neV) using a novel solenoid magnet. The theoretical foundation, simulation result is discussed. Followed by the proposed approaches to pickup the signal using High Impedance Amplifier (HIA) and SQUIDS receiver.

T 77.6 Wed 17:30 T-H36

Low Temperature MMC-based X-ray Detectors for IAXO — •DANIEL UNGER, ANDREAS ABELN, DANIEL BEHREND-URIARTE, DANIEL HENGSTLER, ANDREAS FLEISCHMANN, CHRISTIAN ENSS, and LOREDANA GASTALDO — Kirchhoff Institute for Physics, Heidelberg University

Axion helioscopes search for evidence of axion-like particles (ALPs) produced in the Sun. Via the generic ALP-photon coupling, a strong magnetic field would convert ALPs into photons which could then be detected by low background and high efficiency X-ray detectors. Having also detectors with good energy resolution and low energy threshold would in addition in case of discovery allow to investigate ALP properties and generation mechanisms in the Sun. We propose to use low temperature metallic magnetic calorimeters (MMCs) for the International Axion Observatory (IAXO). We present the current state of our detector system developed for IAXO containing a two-dimensional 64-pixel MMC array covering an active area of 16 mm² with a fill factor of 93 %. We achieve an average energy resolution of 6 eV FWHM allowing for energy thresholds well below 100 eV. The results obtained during experiments with different experimental configurations show a background reduction in the case of low-Z material directly surrounding the active part of the detector. In the future, active and passive shields will be used to reduce the background further. The obtained results highlight that MMC-based arrays are a suitable technology for helioscopes to discover and study ALPs.

T 77.7 Wed 17:45 T-H36

Indirect dark matter search with IceCube — •LI RUOHAN, STEPHAN MEIGHEN-BERGER, and ANJA BRENNER — Technische Universität München, James-Franck-Straße 1, 85748, Garching, Germany

Dark Matter annihilation can generate standard particle pairs in primary and decay into neutrinos at the final state. Its spectrum can have a line shape in case of direct annihilation into neutrinos pair. IceCube neutrino observatory is a powerful instrument for indirect dark matter search because of its sensitivity to neutrinos energy of TeV to PeV. Its planned Upgrade can improve the dark matter nucleons interaction cross-section limits of one magnitude at lower energy. This talk will show a potential approach to test IceCube's line-spectrum detection ability and estimate the conservative cross-section using both spin-independent and -dependent effective fields theory.

T 77.8 Wed 18:00 T-H36

EXCESS workshop: a collaborative investigation of the sub-keV backgrounds observed in various rare event search experiments — •MARGARITA KAZNACHEEVA¹, ALEXANDER FUSS^{2,3}, FLORIAN REINDL^{2,3}, and FELIX WAGNER² — ¹Physik-Department E15, Technische Universität München, D-85748 Garching, Germany — ²Institut für Hochenergiephysik der Österreichischen Akademie der Wissenschaften, 1050 Wien, Austria — ³Atominstitut, Technische Universität Wien, 1020 Wien, Austria

After having lowered the energy thresholds down to O(10eV), various dark matter and coherent elastic neutrino-nucleus scattering experiments observe an unexpected exponential rise of the event rate towards low energies. This excess signal caused by an as yet unknown origin currently provides the main limitation for further sensitivity improvement. A collective initiative to share experimental observations and compare the measured excess signals was started. I will report the outcomes of the dedicated EXCESS workshop that took place in June 2021 as a joint effort of 10 collaborations and lead to an in-depth discussion within the community. Presented measurements were taken by cryogenic, CCD, and gaseous ionization detectors, under and above ground, with different levels of shielding and a wide range of operating temperatures. In the scope of the workshop, a publicly accessible data repository was created that allows studying the sub-keV excess signals measured by the participating collaborations. A summary paper of the workshop is expected to be published in early 2022 and further meetings are already planned.

T 78: Experimental Techniques in Astroparticle Physics 3

Time: Wednesday 16:15-18:00

T 78.1 Wed 16:15 T-H37

Monoenergetic electronic recoil calibration of LXe TPCs with 37Ar (XENON1T/nT) — •CHRISTOPHER HILS for the XENON-Collaboration — Institut für Physik & Exzellenzcluster PRISMA+, Johannes Gutenberg-Universität Mainz, 55099 Mainz, Germany

Large multi-ton LXe time projection chambers like XENON1T/nT set the most stringent constraints on the interaction cross-section between nucleons and Dark Matter in form of Weakly Interacting Massive Particles. The large active volume and the excellent self shielding properties of liquid xenon make the use of internal calibration sources a necessity to understand detector responses. In the past these calibrations were mainly based on gaseous 83mKr and 220Rn isotopes diluted into the liquid xenon and distributed equally into the active volume. In the last science run of XENON1T we introduced a new low-energy calibration source, the Argon isotope 37Ar, with calibration lines at energies of 2.8 keV and 270 eV. In this talk we will present the results of the XENON1T calibration in form of a study of the detector response at these ultra low energies. We also show that the isotope can be efficiently removed by cryo distillation in the XENON1T distillation column originally designed for krypton removal, which made this isotope suitable as a regular calibration source despite ist long halflife time of 35 d. In this regard, a first calibration was already performed In XENONnT at the end of 2021 with first results about to come.

T 78.2 Wed 16:30 T-H37

Measuring the liquid xenon scintillation pulse shape and its electric field dependence — •DOMINICK CICHON¹, GUILLAUME EURIN^{1,2}, FLORIAN JÖRG¹, TERESA MARRODÁN UNDAGOITIA¹, and NATASCHA RUPP¹ — ¹Max-Planck-Institut für Kernphysik, Heidelberg, Germany — ²IRFU, CEA, Université Paris-Saclay, Gif-sur-Yvette, France

In the search for new physics, such as particle dark matter and the neutrinoless double-beta decay, liquid xenon (LXe) detectors play an important role and have provided highly competitive results over the past years. As an example, XENON1T, which utilized a LXe time projection chamber (TPC), constrained the cross-section for interactions between weakly interacting massive particles (WIMPs) and nucleons to values below $4.1 \cdot 10^{-47}$ cm² at a WIMP mass of 30 GeV/c². To push LXe detector technology to its limits and achieve even better sensitivities, a detailed understanding of the microphysics processes responsible for signal generation in LXe is necessary. One avenue to investigate such processes is the pulse shape of the prompt scintillation signal caused by excitation of LXe via particle interactions.

This talk presents measurements of the LXe scintillation pulse shape after excitation by either electrons from the isomeric transition of 83m Kr or alpha particles from 222 Rn chain decays. For both sources, the pulse shape has been characterized at more than 25 different electric field configurations between ~ 0 V/cm and ~ 1200 V/cm. The results are compared to previoulsy published data and interpreted in the context of the involved microphysics processes. Location: T-H37

T 78.3 Wed 16:45 T-H37

Gaseous xenon measurements with APIMS and gas chromatography — •VERONICA PIZZELLA and HARDY SIMGEN — Max-Planck-Institut für Kernphysik, Saupfercheckweg 1, 69117, Heidelberg

The latest generation of dual phase gas-liquid xenon TPC detectors for rare event searches employ several tonnes of xenon. It is crucial for the success of these experiments that the chemical impurities in LXe are below ppb level. Some chemical impurities of concern are: electronegative molecules such as oxygen, since they trap the electrons; radioactive impurities such as H-3, since they increase the background in the ROI.

In this presentation, a method to measure chemical impurities below ppb level is presented. The measurement is performed using Atmospheric Pressure Ionization Mass Spectrometry (APIMS), with a commercial instrument from Thermo Scientific. This instrument uses a corona discharge to ionize helium gas at atmospheric pressure, which in turn ionizes mixed trace impurities very efficiently. The setup uses a custom chromatography setup to separate the impurities from the xenon and mix them with helium. Some of the challenges of oxygen and hydrogen quantification are illustrated and some adapted solutions are outlined. A first measurement of the xenon from the gas phase of the XENONnT experiment is reported.

T 78.4 Wed 17:00 T-H37

Scintillation and optical properties of xenon-doped liquid argon — •CHRISTOPH VOGL, MARIO SCHWARZ, XAVER STRIBL, JOHANNA GRIESSING, PATRICK KRAUSE, and STEFAN SCHÖNERT — Chair for Astroparticle Physics, Department of Physics, Technical University Munich, Garching, Germany

Liquid argon (LAr) is widely employed as a scintillator in rare-event searches. Its optical and scintillation properties, as well as the impact of impurities, are being studied extensively by many groups world-wide. LAr scintillation light exhibits a main emission wavelength of 128 nm, which makes propagation and detection challenging because of short attenuation lengths and low quantum efficiencies of photo sensors in the VUV spectral range. The addition of small amounts of xenon to LAr shifts the emission wavelength towards 175 nm and reduces the overall scintillation time. Here, we present our latest study of xenon-doped LAr with focus on the primary photon yield, the effective triplet lifetime and attenuation length, with xenon concentrations ranging from 3 ppm to 300 ppm. The scintillation and optical properties were measured simultaneously with the LLAMA instrument operated inside SCARF, a 1 ton LAr test stand, and the xenon concentrations using IDEFIX, a dedicated mass spectrometer setup. This research is supported by the DFG through the Excellence Cluster ORIGINS and the SFB1258.

T 78.5 Wed 17:15 T-H37 **Characterization of Wavelength Shifters for LAr Instrumentation Using VUV Spectrofluorometry** — •ANDREAS LEONHARDT¹, GABRIELA R. ARAUJO², PATRICK KRAUSE¹, LASZLO PAPP¹, TINA R. POLLMANN³, and STEFAN SCHÖNERT¹ — ¹Physik Department, Technische Universität München, Garching, Germany — ²Physik-Institut, Universität Zürich, Zurich, Switzerland — ³Nikhef National Institute for Subatomic Physic, Amsterdam, Netherlands Experiments searching for dark matter or neutrinoless double-beta decay commonly use liquid argon (LAr) as a target or instrumented shielding medium. Particle interactions in the LAr produce vacuum-ultraviolet (VUV) light flashes peaking at 128 nm, which are converted to longer wavelengths by wavelength shifters (WLSs). Due to the short LAr scintillation wavelength and low LAr temperature, the characterization of WLSs requires VUV optics and a cooling system in vacuum. We present the developed custom spectrofluorometer, which enables us to characterize WLSs at VUV excitation and low temperatures. The setup consists of a high-intensity deuterium light source coupled to a VUV monochromator and a vacuum-tight sample chamber. The wavelength shifting material can be mounted on a cryocooler coldhead to measure the wavelength-resolved and wavelength-integrated photoluminescence light yield at the relevant LAr temperature. We describe the characterization campaign of the wavelength-shifting reflector of the LEGEND-200 experiment with the VUV spectrofluorometer and summarize the results. This research is supported by the DFG through the Excellence Cluster ORIGINS and the SFB1258.

T 78.6 Wed 17:30 T-H37

Development of an Organic Plastic Scintillator based Muon Veto Operating at Sub-Kelvin Temperatures for the NUCLEUS Experiment — •ANDREAS ERHART^{1,2}, VICTORIA WAGNER¹, LUDWIG KLINKENBERG¹, THIERRY LASSERRE², DAVID LHUILLIER², CLAUDIA NONES², TOBIAS ORTMANN¹, LUCA PATTAVINA¹, RUDOLPH ROGLY², JOHANNES ROTHE¹, VLADIMIR SAVU², NICOLE SCHERMER¹, RAIMUND STRAUSS¹, and MATTHIEU VIVIER² — ¹Physik-Department, Technische Universität München, D-85748 Garching — ²IRFU, CEA, Université Paris Saclay, F-91191 Gif-sur-Yvette

The NUCLEUS experiment aims to measure coherent elastic neutrino nucleus

scattering of reactor anti-neutrinos using cryogenic calorime- ters. Operating at an overburden of 3m.w.e., muon-induced back- grounds are expected to be dominant. It is therefore essential to de- velop an efficient muon veto, with a detection efficiency of more than 99%. A novel concept has been investigated, featuring a plastic scin- tillator based muon veto operating inside the NUCLEUS cryostat at sub-Kelvin temperatures. The required investigation of the detector's low temperature behavior led to the first reported measurements of organic plastic scintillators at sub-Kelvin temperatures. The function- ality of the principal scintillation process has thereby been confirmed. A disc-shape muon veto equipped with wavelength shifting fibers and a silicon photomultiplier has been developed. The research was sup- ported by the DFG through the Excellence Cluster ORI-GINS and the SFB1258, and the ERC Starting Grant 2018 "NU-CLEUS".

T 78.7 Wed 17:45 T-H37 Vibration Decoupling System for the NUCLEUS Experiment — •Alexander Wex, Raimund Strauss, Johannes Rothe, Luca Pattavina, Nicole Scher-Mer, Andreas Erhart, Tobias Ortmann, Victoria Wagner, and Mar-Garita Kaznacheeva — Technische Universität München, Physik Department Lehrstuhl E15, James-Franck-Straße 1, D-85748 Garching

The coherent neutrino nucleus scattering experiment NUCLEUS deploys a newgeneration dry dilution refrigerator. Vibrations induced by the cryostat's pulse tube cooler are a challenge for stable detector operation. To achieve detector performance undisturbed by pulse tube operation, a dedicated spring-decoupling system is being developed for NUCLEUS. Recent results and benchmark measurements for the design of this cryogenic vibration decoupling system are presented. The research was supported by the DFG through the Excellence Cluster ORIGINS and the SFB1258, and the ERC Starting Grant 2018 "NU-CLEUS".

T 79: Data Analysis, Information Technology and Artificial Intelligence 4

Time: Wednesday 16:15–18:30

T 79.1 Wed 16:15 T-H38 K_S^0 tracking efficiency studies at Belle II — •ELISABETTA PRENCIPE¹, OLEK-SANDR SKORENOK², and JENS SOEREN LANGE¹ — ¹JLU-Giessen, Giessen, Germany — ²TSNU-Kyiv, Kyiv, Ukraine

Reconstruction efficiency of low momentum tracks in particle physics is a very important issue. Here we report about a study conducted at Belle II, with MC samples and Phase-3 data, and present a dedicated study of the efficiency to correctly reconstruct $K_S^0 \rightarrow \pi^+ \pi^-$, whose daughter tracks can have a different efficiency due to their displacement from the primary event origin.

A significant number of analyses in Belle II involve the reconstruction of $K_S^0 \rightarrow \pi^+\pi^-$. The reconstruction efficiency of the K_S^0 daughters depends on the K_S^0 transverse momentum, p_T , polar angle, θ_{LAB} and transverse (XY) flight distance, d_{XY} , which is computed as the distance between the primary vertex of the event and the refitted K_S^0 decay vertex.

The general strategy is to subdivide the data and MC events into a large number of samples by choosing an appropriate binning in these variables, determine the number of K_S^0 in each bin, in data and MC samples, and for each of the momentum and polar angle ranges, normalize the ratio of its value in the first bin in d_{XY} , where all tracking effects are well understood. The results here presented are acquired by studying the $B \to K^+ K^- K_S^0$ and $B \to \pi^+ \pi^- K_S^0$ decay channels at Belle II. They will help in understanding systematic effects of analysis where K_S^0 are involved.

T 79.2 Wed 16:30 T-H38

Clustering Energy Depositions in the Electromagnetic Calorimeter at Belle II using Graph Neural Networks — •FLORIAN WEMMER, PABLO GOLDENZWEIG, and TORBEN FERBER for the Belle II-Collaboration — Karlsruher Institut fuer Technologie

Electromagnetic calorimeters in particle detectors like at the Belle II Experiment consist of almost ten thousand sensitive crystals providing detailed energy deposition information in space. The correct assignment of energy depositions in those crystals to clusters originating from a distinct particle imposes a huge challenge especially in the presence of beam induced backgrounds, electronic noise and overlapping clusters. Graph Neural Networks (GNNs) allow for a machine learning algorithm to unrestrictedly and elegantly learn a feature space best suited to solve a problem. Using readily available Monte Carlo data we apply a GNN to try and cluster crystalwise energy information as well as distinguishing physics signals from beam background in the Belle II electromagnetic calorimeter. As a starting point to the development of more capable algorithms the - in actuality complex - detector data is simplified to two possibly overlapping clusters and beam background. We give insight to possible loss functions and metrics of the GNN as well as presenting first results of the clustering process. Location: T-H38

T 79.3 Wed 16:45 T-H38 Fast Particle Reconstruction in the Belle II Experiment with Graph Neural

Networks — •ISABEL HAIDE, PABLO GOLDENZWEIG, and TORBEN FERBER for the Belle II-Collaboration — Karlsruhe Institute of Technology

The correct clustering and reconstruction of particles in electromagnetic calorimeters are vital to many analyses to ensure a correct reconstruction of the actual event. This clustering poses difficulties such as an unknown number of particles in the calorimeter and the existence of background events and promotes the use of machine learning (ML) algorithms. Due to the irregular geometry of such detectors, graph neural networks (GNNs) are most suitable to provide an improvement over standard algorithms. GNNs do not depend on regular geometries to learn detector-hit representations and have already been successfully applied to simulated data of a simplified calorimeter model. Extending this application to the geometry of real detectors, such as the Belle II electromagnetic calorimeter (ECL), while reconstructing an unknown number of clusters with possible overlap and additional background events, is the goal of this study. In this talk, the concept of using object condensation with GNNs to reconstruct particles in the ECL and the current status of this development is shown. The evaluation method, which is the separation of the signature of the hypothetical dark photon process $e^+e^- \rightarrow A'\gamma, A' \rightarrow e^+e^-$ to the signature of radiative Bhabha scattering $e^+e^- \rightarrow e^+e^-\gamma$, is also explained.

T 79.4 Wed 17:00 T-H38

Identification of pions and muons with the Belle II calorimeter using convolutional neural network — •ABTIN NARIMANI CHARAN¹ and TORBEN FERBER² for the Belle II-Collaboration — ¹Deutsches Elektronen-Synchrotron (DESY), Hamburg, Germany — ²Karlsruhe Institute of Technology (KIT), Karlsruhe, Germany

The Belle II experiment is located at the asymmetric SuperKEKB e^+e^- collider in Tsukuba, Japan. The electromagnetic calorimeter (ECL) in Belle II is designed to reconstruct neutral particles. Additionally, the ECL can identify charged particles e.g. electrons, muons and pions. Identification of low-momenta muons and pions in the ECL is crucial if they do not reach the outer muon detector.

This talk presents an application of a convolutional neural network (CNN) to separate muons and pions in the ECL using energy deposition patterns of 7x7 crystal images. Due to the high level of beam background, the performance of the CNN in samples with different beam background levels is studied. Moreover, the impact of adding pulse-shape and timing information in addition to the energy information is investigated and compared with the previous results in different momentum ranges. Finally, the performance of the network is investigated with data control samples of muons and pions.

T 79.5 Wed 17:15 T-H38

Anomaly detection in the searches for inelastic dark matter at Belle II -•JONAS EPPELT, PATRICK ECKER, PABLO GOLDENZWEIG, and TORBEN FERBER for the Belle II-Collaboration - Karlsruhe Institut for Technology, Karlsruhe, Germany

Inelastic Dark Matter (IDM) is a rather complex model containing two Dark Particles and two Mediators. Its expected signatures include one resonant and one non-resonant decay.

In this talk current efforts are presented to approach searches for IDM at Belle II with anomaly detection via machine learning. This approach aims to train algorithms on (simulated) background events in order to recognize previous unknown signals. It seeks to reduce limitations due to the chosen model and parameter space in searches for physics beyond the Standard Model.

The current status on employing anomaly detection in an IDM search at Belle II will be given.

T 79.6 Wed 17:30 T-H38

Implementing a graph-based approach for semi-inclusive tagging in Belle II - FLORIAN BERNLOCHNER, •AXEL HEIMEROTH, WILLIAM SUTCLIFFE, and ILIAS TSAKLIDIS — Physikalisches Institut, University of Bonn, Germany

In the Belle II experiment pairs of B-mesons are produced from electronpositron collisions. The clean experimental environment allows for constraining kinematically the second B-meson if the other one is correctly reconstructed. The current tagging algorithm used in Belle II, called Full Event Interpretation, has relatively low tagging efficiency since all the intermediate decays of the second B-meson must be explicitly reconstructed. This can be critical for searches of extremely rare decays. A semi-inclusive approach, where one reconstructs only partially the tag-side, instead of an exclusive one, can be used in order to increase the overall efficiency. In this work we explore how graph neural networks can improve the purity and efficiency of a semi-inclusive approach, where only a charmed hadron, instead of a B-meson, is reconstructed. In this talk I present a proof of concept on a generic phasespace dataset and a realistic application on $B \rightarrow D^* l v$ decays from the official Belle II Monte Carlo.

T 79.7 Wed 17:45 T-H38

Performance portability for the Physics Object Reconstruction Software of the CMS Experiment — • WAHID REDJEB — RWTH University, III. Physikalisches Institut A, Aachen, Germany

The High Luminosity upgrade of the LHC will pose unprecedented challenges for the offline and online computing. The higher luminosity and pileup will require larger processing power, not achievable with the current CPUs. Heterogeneous computing will play a fundamental role in the physics object reconstruction software to fully exploit the reach of the HL-LHC. Several computing architectures are available for the CMS software, but specialized implementations for each of them is not sustainable in terms of development, maintenance and validation. Performance Portability libraries allow performance portability across different hardware architectures with a single code basis. In this talk, we present the last results of the first usage of the Alpaka performance portability library on a standalone version of the reconstruction of tracks and vertices in the CMS silicon pixel detector. Porting the pixel tracks and vertices reconstruction

to Alpaka demonstrates the possibility of writing a single source code that can be executed on different devices with different parallelization strategies, achieving similar performance with respect to the native implementations.

T 79.8 Wed 18:00 T-H38

Designing VQE ansatz circuits for track reconstruction with Quantum Computers at LUXE — ARIANNA CRIPPA¹, LENA FUNCKE³, TOBIAS HARTUNG⁴, Beate Heinemann^{1,2}, Karl Jansen¹, Annabel Kropf¹, Stefan Kühn⁵, Fed-ERICO MELONI¹, •DAVID SPATARO¹, CENK TÜYSÜZ¹, and YEE CHINN YAP¹ ¹Deutsches Elektronen Synchrotron DESY — ²Albert-Ludwigs-Universität Freiburg — ³Massachusetts Institute of Technology — ⁴University of Bath -⁵CaSToRC, The Cyprus Institute

The recently proposed Laser Und XFEL Experiment (LUXE) enables the study of Quantum Electrodynamics (QED) in the strong-field regime, where QED becomes non-perturbative. In this regime, the production of electron-positron pairs by field-induced tunneling out of the vacuum is realised in the interaction of a high energy electron beam from the European XFEL and a high power laser.

Positron track reconstruction on a silicon pixel tracking detector becomes a demanding combinatorial problem at high laser intensity. It is expected to measure up to 10⁶ positrons on the four consecutive detector layers. A Quadratic Unconstrained Binary Optimization (QUBO) formulation enables the use of quantum computers and a Variational Quantum Eigensolver (VQE) to reconstruct tracks. For this, designing a suitable ansatz circuit which maps the track candidates to qubits is an important part of the VQE heuristic. Results are compared to common hardware efficient ansatzes. In addition, the final track reconstruction efficiency is compared to a classical approach.

T 79.9 Wed 18:15 T-H38 Benchmarking Variational Quantum Algorithms for track reconstruction **at LUXE** — ARIANNA CRIPPA¹, LENA FUNCKE³, TOBIAS HARTUNG⁴, BEATE HEINEMANN^{1,2}, KARL JANSEN¹, •ANNABEL KROPF¹, STEFAN KÜHN⁵, FED-ERICO MELONI¹, DAVID SPATARO¹, CENK TÜYSÜZ¹, and YEE CHINN YAP¹ ¹Deutsches Elektronen-Synchrotron DESY — ²Albert-Ludwigs-Universität Freiburg - ³Massachusetts Institute of Technology - ⁴University of Bath -⁵CaSToRC, The Cyprus Institute

The primary aim of the recently proposed LUXE experiment is to investigate the transition into the non-perturbative regime of Quantum Electrodynamics. For this, the interaction of photons with electrons, or photons with photons is measured at field strengths where couplings to charges become non-perturbative. In these interactions, up to 10⁶ positrons are produced that then impinge on a four-layered silicon pixel tracking detector. The accurate reconstruction of the positrons' trajectories from a set of hits is a combinatorial problem challenging for a classical computer to solve. For LUXE, a novel approach is explored that expresses pattern recognition as a quadratic unconstrained binary optimisation, allowing the algorithm to be mapped onto a quantum computer. Variational quantum algorithms provide a promising approach to solve combinatorial optimisation problems on noisy quantum devices. Here, we benchmark the accuracy of two such algorithms, the Variational Quantum Eigensolver and the Quantum Approximate Optimisation Algorithm, against classical tracking using data from an idealised LUXE detector set-up.

T 80: Invited Talks 3 (joint session T/EP)

Time: Thursday 11:00-12:30

Invited Talk

T 80.1 Thu 11:00 T-H15

Borexino looks in the direction of solar neutrinos - •LIVIA LUDHOVA for the Borexino-Collaboration — Forschungszentrum Jülich, Jülich, Germany -RWTH Aachen University, Aachen, Germany

Borexino is a 280-ton liquid scintillator detector located at the LNGS in Italy. Characterized by an unprecedented radio-purity, it has succeeded in providing several milestone measurements of MeV-scale neutrinos, with the main focus on solar neutrinos. The latter are the only direct probe of the Hydrogen-to-Helium fusion powering our Sun. The European Physical Society awarded the 2021 Giuseppe and Vanna Cocconi Prize to the Borexino Collaboration for the ground-breaking observation of solar neutrinos from the pp chain and CNO cycle that provided unique and comprehensive tests of the Sun as a nuclear fusion engine. Borexino has developed a new method, Correlated and Integrated Directionality (CID), to exploit the sub-dominant directional Cherenkov light in a liquid scintillator detector. This technique can disentangle the solar neutrino signal, correlated with the known position of the Sun, from the isotropic background. In the region of interest dominated by the signal from 0.862 MeV Be-7 solar neutrinos, the no-solar neutrino hypothesis has been excluded with ${>}5\sigma$ C.L. This novel method is readily applicable to next generation experiments. The talk will focus on the recent Borexino solar neutrino results, including the motivation, analysis details, as well as their interpretation.

Invited Talk

T 80.2 Thu 11:30 T-H15

Location: T-H15

Gravitational waves - a new probe of the early Universe — •VALERIE DOMCKE - CERN, Geneva, Switzerland

Due to their extremely weak interactions with the matter content of the Universe, gravitational waves generated right after the Big Bang can traverse the Universe basically unperturbed, carrying information about their production processes and the expansion history of our Universe. This makes them a unique probe of BSM physcis at very high energies. I will talk about possible next steps in this field, including the search for the stochastic gravitational wave background and new ideas for searching for gravitational waves at ultra-high frequencies.

Invited Talk

T 80.3 Thu 12:00 T-H15 Gravitational wave detectors - current and future challenges - •MICHÈLE

HEURS - Leibniz Universität Hannover Since the first direct detection in 2015, gravitational wave signals have been enriching the field of multi-messenger astronomy with insights into formerly "invisible" regimes of the universe. Despite their mind-boggling sensitivities, the current (second) generation of ground-based gravitational wave detectors are limited by various noise sources in their detection band, in particular quantum noise, thermal noise, and seismic noise. Next-generation detectors (e.g. Einstein Telescope, Cosmic Explorer) aim for sensitivities one or two orders of magnitude better even, making innovative techniques for noise reduction or mitigation a requirement. This talk will present challenges and technical developments on the

Location: T-H15

Time: Thursday 14:00-15:40

Invited Topical Talk

T 81.1 Thu 14:00 T-H15 LND - A ("Made in Germany") Radiation Monitor Operating at the far Side of the Moon — •Sönke Burmeister¹, Shenyi Zhang², Jia Yu¹, Zigong Xu¹, Stephan Böttcher 1 , and Robert Wimmer-Schweingruber $^1-{}^1$ Institut für Experimentelle und Angewandte Physik, Uni Kiel — ²NSSC, Chinese Academy of Science

Space Radiation is one of the major concerns in human space flight. Of course, this also applies to human exploration of the Moon. On the lunar surface, this consists of chronic exposure to galactic cosmic rays and sporadic solar particle events. The interaction of this radiation field with the lunar soil leads to a third component that consists of neutral particles, i.e., neutrons and gamma radiation. Chang'E 4 is the Chinese mission that landed on the far side of the Moon on January 3rd, 2019. It consists of a lander, a rover, and a relay spacecraft. The LND (Lunar Lander Neutrons and Dosimetry) instrument that was built by CAU is located inside the lander under an opening lid. It consists of a stack of ten segmented Si solid-state detectors (SSDs), which form a particle telescope to measure charged particles (electrons from 0.5 MeV to several MeV, protons 8-35 MeV, and heavier nuclei 17-75 MeV/nuc). A special geometrical arrangement allows observations of fast neutrons (and gamma-rays) that are also important for dosimetry purposes. Thermal neutrons are measured by using a very thin Gd conversion foil sandwiched between two SSDs. The Lunar Lander Neutrons and Dosimetry experiment aboard China's Chang'E 4 lander has made the first ever measurements of the radiation exposure to both charged and neutral particles on the lunar surface.

Invited Topical Talk T 81.2 Thu 14:25 T-H15 Energetic Pulsar Environments and the Origins of Galactic Cosmic Rays -•ALISON MITCHELL — Erlangen Centre for Astroparticle Physics, FAU, Erlangen, Germany

Cosmic Rays - and their origins - have fascinated Physicists for over a hundred years. Within our Milky Way Galaxy, particles are known to reach energies beyond the so-called Cosmic Ray *knee*, a spectral break at ~3 PeV in the all particle cosmic ray spectrum. However, evidence for the particle accelerators reach PeV energies - PeVatrons - has proven elusive. Only within the last five years have astrophysical sources of gamma-rays above 100 TeV been identified; gamma-rays produced through interactions of particles with PeV energies. Many of these sources are associated with known energetic pulsars.

In this talk, I will review the current census of PeVatrons and discuss implications for our understanding of pulsar environments. There are several open

questions to grapple with: Which particle species are being accelerated - leptonic or hadronic? How are the particles transported through the surrounding medium? What is the maximum energy limit for particle acceleration in pulsar environments? In the near future, data from current and forthcoming facilities will help us to address these questions.

Invited Topical Talk T 81.3 Thu 14:50 T-H15 Looking forward to exciting physics with FASER — \bullet Felix Kling — DESY Physics searches and measurements at high-energy collider experiments traditionally focus on the high-pT region. However, if particles are light and weaklycoupled, this focus may be completely misguided: light particles are typically highly collimated around the beam line, allowing sensitive searches with small detectors, and even extremely weakly-coupled particles may be produced in large numbers there. The FASER experiment will use the opportunity and extend the LHC's physic potential by searching for long-lived particles and studying neutrino interactions at TeV energies. In this talk, I will present the physics potential of FASER for new physics searches, neutrino physics and QCD and astro-particle physics.

T 81.4 Thu 15:15 T-H15 **Invited Topical Talk** Astroparticle physics at the LHC: Exploring the forward region with crosssection measurements — •HANS DEMBINSKI — Fakultät Physik, Technische Universität Dortmund, Dortmund, Germany

Astroparticle physics is the study of the non-thermal universe with gamma rays, neutrinos, and cosmic rays. Cosmic rays are abundantly produced in cosmic accelerators, like supernova remnants. Some gamma rays and neutrinos are produced indirectly in interactions of cosmic rays with matter in the source, and cosmic rays interact with Earth's atmosphere to produce air showers, which are observed by ground-based cosmic ray observatories and contribute the main background to gamma ray and neutrino observatories. QCD cross-sections for the forward production of hadrons with light and heavy flavor are therefore needed to interpret astroparticle measurements. The experiments at the Large Hadron Collider (LHC) have powerful instruments to measure forward production, but data are more sparse compared to central production. I will summarize the stateof-the-art of forward cross-section measurements at the LHC from the point of view of astroparticle physics and give an outlook into the opportunities in near future with the upcoming run of the LHC and the planned pilot run with oxygen beams.

T 82: Invited Topical Talks 6

T 81: Invited Topical Talks 5 (joint session T/EP)

Time: Thursday 14:00-15:40

Invited Topical Talk T 82.1 Thu 14:00 T-H16 Searches for new scalar particles at the LHC - • DOMINIK DUDA - Max-Planck-Institut für Physik

The Higgs boson discovery by the ATLAS and CMS experiments at the Large Hadron Collider was a great success. Ever since, numerous studies have been performed to establish whether it is a Standard Model particle or rather the first observed physical state of an extended scalar sector beyond the Standard Model.

Extended scalar sectors are well motivated as they can modify the electroweak phase transition and facilitate baryogenesis, enhance vacuum stability, provide a dark matter candidate or provide a solution to the strong CP problem (i.e. predict axions). In short, extending the scalar sector provides solutions to some of the questions the Standard Model fails to answer.

Various theories beyond the Standard Model predict the existence of new Higgs bosons in addition to the already discovered one. E.g., the introduction of a second Higgs doublet field in the minimal supersymmetric extension of the Standard Model leads to the prediction of three neutral and two charged Higgs bosons, while an additional Higgs triplet field e.g. in models with a type-II seesaw mechanism would result in seven scalars in total. The discovery of such new scalar particles would be a direct evidence of new physics.

In this presentation, the latest searches for additional neutral and charged scalars performed with the ATLAS and CMS experiments will be reviewed.

Invited Topical Talk

Location: T-H16

T 82.2 Thu 14:25 T-H16 Novel approaches to search for new physics in rare charm decays — • DOMINIK STEFAN MITZEL — TU Dortmund University, Germany

Recent studies of rare *b*-hadron decays have revealed a coherent pattern of deviations from Standard Model predictions in $b \rightarrow s\ell^+\ell^-$ transitions, known as flavour anomalies. Rare charm decays are sensitive to $c \rightarrow u \ell^+ \ell^-$ flavourchanging neutral-current processes and offer the unique and complementary opportunity to search for anomalies in the up-type quark sector that has hardly been explored in the past. For long, rare charm decays have been considered as less promising due to difficulties in the description of its low energy dynamics. During this talk, I will discuss how exact or approximate symmetries in the charm system allow to construct clean null-test observables, yielding an excellent road to the discovery of New Physics.

Invited Topical Talk

T 82.3 Thu 14:50 T-H16 Constraining the Higgs-charm Yukawa coupling with the CMS experiment - •LUCA MASTROLORENZO — RWTH, Aachen, Germany

In this talk, an overview of the most recent results of the direct search for the VH, H→cc process with the CMS experiment is presented. The search targets Higgs bosons produced in association with a vector boson (W, Z) exploiting the full Run-2 data set. The analysis is carried out in mutually exclusive channels selecting specific leptonic decays of the vector bosons: $Z \rightarrow \ell \ell, Z \rightarrow \nu \nu, W \rightarrow \ell \nu$, with ℓ =electron or muon. To fully exploit the topology of the Higgs boson decay in the different regimes of the Higgs boson transverse momentum, two strategies have been adopted aiming to reconstruct the Higgs boson candidate through two distinct Ak4 jets or via a unique Ak15 "fat-jet". Remarkable improvements have been brought to the analysis techniques with respect to the previous public results: from new and more efficient algorithms to tag charm-initiated jets to dedicated jet energy and mass regression techniques, conceived exploiting advanced machine learning methods. The analysis strategy has been extensively validated by observing the VZ, Z \rightarrow cc process for the first time at a hadron collider experiment. The results represent the world's most stringent limit on the VH, H \rightarrow cc process and on the Higgs-charm Yukawa coupling.

Invited Topical TalkT 82.4Thu 15:15T-H16Characterization of H boson events in the $\tau\tau$ decay channel with the full CMSRun-2 data set — •SEBASTIAN WOZNIEWSKI — Georg-August-Universität, Göttingen, Germany

The LHC Run-2 data set of proton-proton collisions provides first deeper insights into the properties and production of Higgs bosons. Besides the verification of the assumed coupling structure, which remains a challenge, also differential investigations of Higgs boson events are important tests of the Standard Model (SM) Higgs sector. Moreover, models based on supersymmetry allow for modifications of the couplings of the SM-like Higgs boson to (down-type) fermions, which puts particular interest on the decay channel into tau leptons. The analysis of di-tau events of the full LHC Run-2 data set, taken by the CMS experiment, in the STXS framework for differential cross sections measurements is presented. It is based on modern technologies in terms of object identification, data-driven background modeling, and neural-network based multiclass eventclassification. The full granularity of the differential STXS measurement, with twelve signal components, is reflected by the neural networks, in addition to the major background contributions.

T 83: Astroteilchen: Von der Quelle zum Detektor (contributed talks) (joint session EP/T)

Convenor: withAbstracts

Time: Thursday 16:15-18:30

T 83.1 Thu 16:15 EP-H1

Multi-messenger studies with gravitational waves and neutrinos — •TISTA MUKHERJEE for the IceCube-Collaboration — Institute for Astroparticle Physics (IAP), Karlsruhe Institute of Technology (KIT), Karlsruhe, Germany

IceCube is a neutrino observatory located in Antarctica. Since its discovery of a high-energy neutrino (IC170922A) from the blazar TXS0506+056 in 2017, neutrino astronomy has been established as a viable option to probe the high-energy Universe. Neutrinos can carry undistorted information about their respective astrophysical sources, thus can serve as a cosmic 'messenger' to us. There are other potential messengers as well, e.g. gravitational waves (GW) and cosmic rays other than the traditional photons of various wavelengths. Combining interesting signals of such messengers available from different observatories leads us towards multi-messenger searches, allowing us to address many of the so far unanswered questions about the fundamentals of this Universe, such as the origin of ultra-high-energy cosmic ray sources. So far, we have the knowledge of detecting electromagnetic signal in multiple wavelengths, spatially and temporally correlated with GW and high-energy neutrinos, as two separate events. However, there is still a missing link as we have not been able to correlate GW with neutrino signals. The aim of my work is to contribute in this aspect, searching for Virgo detected GW counterparts of neutrino events detected by IceCube, including low-energy neutrinos as well as sub-threshold GW events in our analysis. The work plan and initial results will be discussed in this talk.

T 83.2 Thu 16:30 EP-H1

Multi-messenger characterization of Mrk501 during historically low Xray and gamma-ray activity — •LEA HECKMANN¹, DAVID PANEQUE¹, SAR-GIS GASPARYAN², MATTEO CERRUTI³, NAREK SAHAKYAN², and AXEL ARBET-ENGELS¹ for the Multi-wavelength collaborators and the MAGIC and Fermi-LAT-Collaboration — ¹Max-Planck-Institut für Physik, D-80805 München, Germany — ²ICRANet-Armenia, Marshall Baghramian Avenue 24a, Yerevan 0019, Armenia — ³Institut de Ciències del Cosmos (ICCUB), Universitat de Barcelona (IEEC-UB), Martí i Franquès 1, 8 E08028 Barcelona, Spain

Blazars are the most numerous very-high-energy (>0.2 TeV, VHE) gamma-ray emitters, due to their continuous and very luminous emission; but they are far from being understood.

In this contribution we describe the multi-wavelength behavior of Mrk501, one of our closest and therefore brightest blazars, from 2017 to 2020. Alongside the dense monitoring campaign over this four years, three long observations with NuSTAR were conducted displaying various low-activity flux levels for the hard X-ray emission. This very comprehensive data set reveals a historically low X-ray and VHE gamma-ray emission period lasting two years. Using the low-activity broadband spectral energy distribution (SED) and data published by IceCube, we investigate the nature of the low state. Additionally, we try to explain the evolution of the broadband SED data during the low state to evaluate its potential of being the baseline emission of Mrk501 that is usually outshone by more dominant and variable components.

T 83.3 Thu 16:45 EP-H1

Hadronic models of active galaxies to constrain cosmic-ray acceleration — •XAVIER RODRIGUES — DESY Zeuthen

In a new era of multi-messenger observatories, numerical models can help shed light on what are the sources of the astrophysical neutrinos and the ultra-highenergy cosmic rays. In this talk I discuss recent results on active galactic nuclei (AGN) as multi-messenger sources, based on numerical simulations of photonuclear cosmic-ray interactions. Assuming AGN jets can reaccelerate cosmic rays up to the EeV regime, I will show that an AGN population may in fact dominate the observed flux and chemical composition of ultra-high-energy cosmic rays. Under certain conditions, the accompanying neutrino flux may be observable by future EeV neutrino telescopes, while respecting the current IceCube limits at PeV energies.

T 83.4 Thu 17:00 EP-H1

Location: EP-H1

Neutrino Emission during Supermassive and stellar mass Binary Black Hole Mergers — •ILJA JAROSCHEWSKI¹, JULIA BECKER TJUS¹, and PETER L. BIERMANN^{2,3} — ¹Theoretische Physik IV, Ruhr-Universität Bochum — ²MPI for Radioastr., Bonn — ³Dept. of Phys. & Astron., Univ. Alabama, Tuscaloosa, AL, USA

Ever since the discovery of a diffuse astrophysical neutrino flux by IceCube, the question arose which sources contribute most. With several neutrino-blazar associations since the first high-probability association of such a neutrino to the blazar TXS 0506+056 in 2017, there is an indication that at least a non-negligible part of this diffuse neutrino flux emerges from blazars.

As over ninety stellar mass binary black hole mergers were already detected via gravitational waves (GWs), with more to come, there are strong indications that supermassive black holes (SMBHs) in galaxy centers, and thus blazars, also merge and have had at least one merger in their lifetime. Such a merger is almost always accompanied by a change of observable jet direction, leading to interactions of a preceding jet with surrounding molecular clouds and neutrino productions.

By creating a connection between neutrinos and GWs, we set limits on how much energy can be emitted in form of neutrinos in each merger of binary SMBHs and stellar mass black holes and estimate their contributions to the diffuse neutrino flux that is measured by IceCube. As neutrino production is directly connected to high energy cosmic ray interactions, the contribution of these sources to the cosmic ray injection rate is established.

T 83.5 Thu 17:15 EP-H1

Search for high-energy neutrinos from blazars with IceCube – •CRISTINA LAGUNAS GUALDA for the IceCube-Collaboration — DESY Zeuthen

The IceCube Neutrino Observatory is the world's largest neutrino telescope, instrumenting one cubic kilometre of Antarctic ice. IceCube started its operation with full configuration in 2011 and a diffuse flux of neutrinos was discovered in 2013. To this day the sources of those neutrinos are still largely unknown. One of the most promising neutrino source candidates is blazars, Active Galactic Nuclei with jets aligned towards Earth.

In 2018 IceCube reported the first clearly identified observation of an astrophysical high-energy neutrino, IC170922A, in spatial and temporal coincidence with blazar TXS 0506+056. Other examples of coincidences that have been observed with lower significance are, but not limited to, IC190730A with blazar PKS 1506+012 and IC141209A with blazar GB6 J1040+0617. What these have in common is that they involve a blazar and a high-energy neutrino with a high probability of being astrophysical in origin (neutrino alert). These coincidences can be combined to calculate a global p-value by performing a stacking analysis. Here we present the results obtained with the Fourth Catalog of Active Galactic Nuclei detected by Fermi-LAT (4LAC-DR2, for Data Release 2) and neutrinos detected by IceCube between 2011 and 2020 that would have passed the neutrino alert criteria.

T 83.6 Thu 17:30 EP-H1 Comparison of Models for Predicting Periodic Gamma-Ray & Neutrino Emissions From Blazars — •ARMIN GHORBANIETEMAD, ILJA JAROSCHEWSKI, and JULIA BECKER TJUS — Theoretische Physik IV, Ruhr-Unviersität Bochum There are several indications that electromagnetic emissions from blazars have quasi-periodic variability, ranging from minutes to years. The long-term peri-

odicity in the span of years is particularly evident in gamma-ray observations with the Fermi LAT instrument. Two separate high probability associations of neutrinos, detected by IceCube in 2014/15 and 2017, to the blazar TXS0506+056 further indicate that blazars are neutrino emitters. These two flares can be interpreted as a possible periodicity. It is the aim of this work to develop a general set of models that can explain the periodic gamma-ray and neutrino emissions from blazars.

In this talk, we present models with single supermassive black holes as well as supermassive binary black hole mergers at the centers of blazers. Our focus lies on supermassive binary black hole mergers, due to them radiating gravitational waves which could be detectable by the Laser-interferometer Space Antenna (LISA). The binary systems are characterized by the change of jet direction accompanied by jet precession close to an imminent merger. This allows predictions of possible neutrino and gravitational wave emissions from blazars with quasi-periodic behavior.

T 83.7 Thu 17:45 EP-H1

First science results from the X-ray telescope STIX on Solar Orbiter -•Alexander Warmuth, Frederic Schuller, and Gottfried Mann — Leibniz -Institut für Astrophysik Potsdam (AIP)

The ESA mission Solar Orbiter was successfully launched in 2020, with the main goal of improving our understanding of how the Sun creates and controls the heliosphere. The Spectrometer/Telescope for Imaging X-rays (STIX) is one of six remote-sensing instruments on board and provides imaging spectroscopy of solar flares in the energy range of 4 to 150 keV. Thus, STIX is able to measure quantitatively both the parameters of the hot flare plasma and the characteristics of the accelerated electrons. Together with the other instruments on Solar Orbiter as well as with other space-borne and ground based observational assets, STIX studies energy release and particle acceleration in solar flares. This talk will be focused on the first science results obtained during the cruise phase of Solar Orbiter (2020 and 2021). This includes observations of microflares, constraints on flare energetics, collaborative studies of gamma-ray flares together with Fermi, and the investigation of flare-associated solar energetic particle events.

T 83.8 Thu 18:00 EP-H1

Unfolding the muon neutrino energy spectrum from 10 years of IceCube data with DSEA+ — •LEONORA KARDUM, KAROLIN HYMON, JOHANNES WERTHE-BACH, PASCAL GUTJAHR, TIM RUHE, and JEAN-MARCO ALAMEDDINE for the IceCube-Collaboration - Astroparticle Physics WG Rhode, TU Dortmund University, Germany

T 84: Flavour Physics

Time: Thursday 16:15-18:30

T 84.1 Thu 16:15 T-H15

Tagged analysis of $B \rightarrow X_{\mu} \ell \nu$ at Belle — •ARMINDOKHT AFSHARIPOUR, Florian Bernlochner, Jochen Dingfelder, Svenja Granderath, Peter LEWIS, and RIEKA RITTSTEIGER — Physikalisches Institut, Universität Bonn We present a study of the exclusive charmless semileptonic decays, $B \rightarrow X_u \ell v$, where $X_{\mu} = \pi^{0}, \eta, \eta'$ and ℓ is an electron or a muon, with the Belle experiment at the SuperKEKB collider in Japan. In the Belle experiment, electrons and positrons are collided at the center-of-mass energy equal to the mass of the $\Upsilon(4S)$ resonance, which decays to pairs of B mesons. One of the two B mesons can be fully reconstructed in a hadronic decay mode (hadronic B tagging) using the Full Event Interpretation algorithm. The signal B meson is then reconstructed from the remaining particles (hadrons and leptons) formed from the unassigned tracks and neutral clusters in the event. In this talk, the general analysis strategy and status of the $B \rightarrow X_u \ell v$ analysis using hadronic B-tagging in Belle is presented.

T 84.2 Thu 16:30 T-H15

Untagged Analysis of $B \rightarrow \pi \ell \bar{\nu}_{\ell}$ and extraction of $|V_{ub}|$ at Belle II — FLO-RIAN BERNLOCHNER, JOCHEN DINGFELDER, •SVENJA GRANDERATH, and PETER LEWIS for the Belle II-Collaboration — Physikalisches Institut der Rheinischen Friedrich-Wilhelms-Universität Bonn

One of the puzzles of current research in flavor physics is the persisting discrepancy between the results of exclusive and inclusive measurements of the CKM matrix element $|V_{ub}|$. The charmless semileptonic decay $B \to \pi \ell \bar{\nu}_{\ell}$ is one of the most accessible and powerful channels for determining $|V_{ub}|$ in exclusive modes. Using data from the Belle II experiment, a new precision measurement of $|V_{ub}|$ can be performed. In preparation for this, an untagged measurement method for extracting $B \to \pi \ell \bar{\nu}_{\ell}$ events is developed using Belle II data. An untagged measurement allows for sufficiently large samples of this rare decay already with the current Belle II dataset. In order to increase the signal purity, boosted decision trees are employed to suppress continuum and $B\overline{B}$ backgrounds. This talk will discuss the current status of the analysis and $|V_{ub}|$ extraction.

Neutrinos, the most elusive particles in the Standard Model, can travel tremendous distances unaffected by magnetic fields or encountered particles from distant sources in the Universe. As this makes them perfect information carriers, many attempts at uncovering their properties are made. The IceCube Neutrino Observatory, a cubic kilometer detector embedded in the South Pole ice, is capable of detecting neutrinos from several GeV up to PeV energies enabling precise reconstruction of the neutrino spectrum. Determining the accurate spectrum is of great importance to neutrino physics, especially in differentiating the three predicted components - prompt, conventional, and astrophysical, of which only the latter two have been detected so far. The Dortmund Spectrum Estimation Algorithm (DSEA+) is a novel approach to unfolding the energy spectrum from measured experimental quantities that effectively translates ill-posed problems to multinomial classification solvable using readily available machine learning tools. The current status of applying DSEA+ on 10 years of IceCube data will be presented.

T 83.9 Thu 18:15 EP-H1

Recent solar and geoneutrino results from Borexino - •SINDHUJHA KU-MARAN for the Borexino-Collaboration — Forschungszentrum Jülich - Institute for Nuclear Physics, IKP-2, Jülich, Germany - III. Physikalisches Institut B, RWTH Aachen University, Aachen, Germany

Borexino is a 280-ton liquid scintillator experiment which ran from May 2007 until October 2021 at the Laboratori Nazionali del Gran Sasso (LNGS), Italy. The main goals of Borexino include the measurement of solar neutrinos and geoneutrinos. The extreme radiopurity and thermal stability of the detector have proven to be valuable assets in achieving these goals. Borexino has not only performed a complete spectroscopy of the dominant pp-chain solar neutrinos but has also provided the first direct experimental evidence of the rare CNO-cycle neutrinos. These measurements have several implications for solar and stellar Physics and further improvements are envisioned using the full dataset. In addition, it has recently presented the first directional measurement of sub-MeV solar neutrinos using the sub-dominant Cherenkov light, through a novel technique called Correlated and Integrated Directionality (CID). This method can be further combined with a typical spectral fit and can also prove valuable for next-generation liquid scintillator detectors. The latest geoneutrino measurement from Borexino included a substantial improvement in the precision as well as the rejection of the no-mantle signal with a high significance. This group report will summarize all these recent solar and geoneutrino results of Borexino.

Location: T-H15

T 84.3 Thu 16:45 T-H15 Exclusive $B \rightarrow X_{\mu} \ell \nu_{\ell}$ Decays with Hadronic Tagging in Belle II Data — •MORITZ BAUER, PABLO GOLDENZWEIG, and TORBEN FERBER for the Belle II-Collaboration - KIT, Karlsruhe, Germany

There has been a longstanding 3σ tension between inclusive and exclusive measurements of the magnitude of the CKM matrix element $|V_{ub}|$. Semileptonic decays involving $b \rightarrow u$ transitions present a unique opportunity to measure $\left|V_{ub}\right|$ with the current Belle II dataset due to their comparatively high branching fraction.

We present analyses of the semileptonic processes $B \to \pi \ell \nu_{\ell}$ and $B \to \rho \ell \nu_{\ell}$ in Belle II data as steps towards the extraction of this matrix element from exclusive decays. These analyses are conducted with hadronic tagging, utilizing a new tagging algorithm, the Full Event Interpretation.

Tagged analysis of $B \rightarrow \rho^0 \ell \nu$, $B \rightarrow \rho^+ \ell \nu$ and $B \rightarrow \omega \ell \nu$ – Ar-MINDOKHT AFSHARIPOUR, FLORIAN BERNLOCHNER, JOCHEN DINGFELDER, SVENJA GRANDERATH, PETER LEWIS, and •RIEKA RITTSTEIGER for the Belle-Collaboration - Physikalisches Institut der Rheinischen Friedrich-Wilhelms-Universität Bonn

We present a study of the charmless semileptonic decays $B \to \rho^0 \ell \nu, B \to \rho^+ \ell \nu$ and $B \rightarrow \omega \ell \nu$, where ℓ is an electron or a muon, with the Belle experiment at KEK in Tsukuba, Japan. Belle took data from 1999 until 2010 at a center-of-mass energy corresponding to the mass of the $\Upsilon(4S)$, which predominantly decays into a pair of B mesons. The Full Event Interpretation is used to fully reconstruct one B meson in a hadronic decay mode. The signal B meson is reconstructed from a lepton and the respective hadron (ρ , ω) using the remaining tracks and neutral clusters in the event. The precise knowledge of the tag-side B decay leads to a good signal-to-background ratio but also a very small efficiency.

In this talk the analysis strategy and status of the $B \to \rho^0 \ell \nu, B \to \rho^+ \ell \nu$ and $B \rightarrow \omega \ell \nu$ analysis using hadronic B-tagging at Belle are presented.

T 84.5 Thu 17:15 T-H15

Measurement of the photon energy spectrum in the fully-inclusive hadronic-tagged $B \rightarrow X_s \gamma$ decays at the Belle II experiment — •HENRIKAS SVIDRAS — DESY, Hamburg

Belle II is an experiment at the next-generation *B* factory SuperKEKB located at KEK in Tsukuba, Japan. It aims to probe heavy flavour physics at a higher precision than its predecessors, namely BaBar and Belle. The goal is to collect 50 ab^{-1} of data during its run: more than 50 times that of Belle. One of the particularly promising decay channels to study is the inclusive radiative $B \rightarrow X_s y$ decay, where X_s denotes any possible decay products containing an *s* quark and *y* is a high-energetic photon. This decay can provide constraints for beyond-SM theories, for example by measuring *CP* asymmetries, and be used to extract important parameters such as the *b* quark mass. The analysis presented in this talk focuses on the hadronic-tagged fully-inclusive approach, where one of the daughter *B* mesons of the $\Upsilon(4S) \rightarrow B\bar{B}$ decays into hadrons. The extraction of the photon energy spectrum of the $B \rightarrow X_s y$ is one of the goals of the analysis. The talk presents the analysis setup and main challenges of this measurement at Belle II.

T 84.6 Thu 17:30 T-H15

Measurement of the q^2 moments in semi-leptonic B meson decays at Belle II — •MAXIMILIAN WELSCH, FLORIAN BERNLOCHNER, and JOCHEN DINGFELDER — Physikalisches Institut der Rheinischen Friedrich-Wilhelms-Universität Bonn

The determination of inclusive $|V_{cb}|$ from $b \rightarrow c\ell v$ decays relies on the Heavy Quark Expansion (HQE) involving coefficients and associated non-perturbative matrix elements, which can be expressed in terms of a number of expansion parameters. The moments of the kinematic distribution of the decay can be computed in a similar manner and are dependent on the same HQE parameters. Consequently, measurements of such moments can be used to better constrain the expansion parameters and, thereby, more precisely determine $|V_{cb}|$. In this talk, we present the first measurement of the q2 moments of $B \rightarrow X_c \ell v$ decays with 62.8 fb⁻¹ of Belle II data. The q2 moments of the $b \rightarrow c\ell v$ transition are particularly powerful for constraining the HQE expansion as they can be expressed in terms of a reduced set of non-perturbative parameters due to reparametrization invariance. In addition, we present an preliminary determination of $|V_{cb}|$.

T 84.7 Thu 17:45 T-H15

Untagged $\overline{B}^0 \to D^{*+} \ell^- \overline{\nu}_{\ell}$ **studies with Belle II** — FLORIAN BERNLOCHNER¹, LU CAO^{1,2}, JOCHEN DINGFELDER¹, and •CHAOYI LYU¹ — ¹Physikalisches Institut der Rheinischen Friedrich-Wilhelms-Universität Bonn — ²Deutsches Elektronen-Synchrotron DESY

The precise determination of the CKM matrix element $|V_{cb}|$ and semileptonic form factors in *B* meson decays are important for carrying out precision tests of the flavour sector of the Standard Model and to search for new physics. The decay of $\overline{B}^0 \to D^{*+} \ell^- \overline{\nu}_{\ell}$ is particularly well suited to determine $|V_{cb}|$ due to its large branching fraction, small backgrounds and the availability of lattice data

to describe the form factors. In this talk, we will present the current status of establishing an untagged measurement of the $\overline{B}^0 \to D^{*+} \ell^- \overline{\nu}_{\ell}$ branching fraction and form factors using early Belle II data.

T 84.8 Thu 18:00 T-H15

Measuring kinematic distributions of $B \to D^* \ell \nu_\ell$ with hadronic tagging at Belle II — FLORIAN BERNLOCHNER, JOCHEN DINGFELDER, •MICHAEL ELIACHEVITCH, MICHAEL HEDGES, PETER LEWIS, MARKUS PRIM, and WILLIAM SUTCLIFFE — Physikalisches Institut der Rheinischen Friedrich-Wilhelms-Universität Bonn

The analysis of semileptonic B meson decays is one of the main pillars of the physics program of the Belle II experiment, since their theoretical cleanliness enables precise theoretical predictions which can be compared with measurements for tests of the Standard Model of particle physics (SM).

This talk presents early results and analysis plans for the $B \rightarrow D^* \ell \nu_{\ell}$ decay with ℓ denoting the light leptons e and μ . Its high branching fraction and ease of reconstruction allow to use available Belle II data for differential measurements of the shapes describing the decay kinematics. These are sensitive to the $|V_{cb}|$ CKM matrix-element and to the form-factors describing the interactions of the hadronic current. Based on these shapes we can further measure angular observables such as the forward-backward asymmetry A_{FB} to probe the SM. In the presented analysis the other *B* meson originating from the Y(4S) is fully reconstructed in hadronic decay modes via the *Full Event Interpretation* tagging algorithm, providing the full four-momentum of the signal *B* meson. Due to the resulting high purity and good resolution this serves as an important cross-check of similar measurements using an inclusive tagging approach.

T 84.9 Thu 18:15 T-H15

Studies of $B \to D^{**}\ell v$ at Belle II — ARIANE FREY, •NOREEN RAULS, and BENJAMIN SCHWENKER — II. Physikalisches Institut, Georg-August-Universität Göttingen, Friedrich-Hund-Platz 1, 37077 Göttingen, Deutschland

The semileptonic decay $B \rightarrow D^{**}\ell\nu$ is one of the main background modes for the determination of the ratio $R(D^{(*)})$, which is used to probe the Standard Model. One of the aims of this analysis is to understand the decay $B \rightarrow D^{**}\ell\nu$ as background contribution to the $R(D^{(*)})$ measurement. For the reconstruction of the daughters of the D^{**} various different hadronic D and D^* modes are used. The B meson is reconstructed using these D^{**} decay modes as well as light and charged leptons.

B mesons are always produced in pairs on the $\Upsilon(4S)$ resonance at the Belle II experiment in Japan. One of these *B* mesons is reconstructed using the decay mode stated above. The other *B* meson uses the hadronic Full Event Interpretation (FEI) for its reconstruction. The FEI is an algorithm, which is based on the hierarchical reconstruction of final-state, intermediate particles and *B* mesons using multivariate classifiers.

This talk will give a first insight on the reconstruction of the B meson in these channels using Belle II Monte Carlo samples. Furthermore, a brief outlook will be shown.

T 85: Beyond the Standard Model (Theory) 2 and QFT and Lattice Gauge Theory 1

Time: Thursday 16:15-18:30

T 85.1 Thu 16:15 T-H16

Four-top final states as a probe of Two-Higgs-Doublet models — •STEVEN PAASCH and HENNING BAHL — Deutsches Elektronen-Synchrotron DESY

Using a CMS measurement of four top $(t\bar{t}t\bar{t})$ production in proton-proton collisions we constrain the parameter space of BSM scalar models. We study these effects for models with a generic scalar X with couplings to W-bosons and to top-quarks. We use Monte-Carlo simulators and fast detector simulations to recast the CMS analysis in order to obtain upper limits on the cross section times branching fraction for the production modes $pp \rightarrow (t\bar{t}, tW, t) + X$ with $X \rightarrow t\bar{t}$, where X is a new heavy Higgs H, a pseudoscalar A or mixed CP-state. Furthermore we study the impact on Two-Higgs-Doublet models where four top production places constraints on the low tan β region which is of special interest for Baryogenesis.

T 85.2 Thu 16:30 T-H16

Flavour and LHC constraints in the 2HDM — OLIVER ATKINSON¹, •MATTHEW BLACK², CHRISTOPH ENGLERT¹, ALEXANDER LENZ², ALEKSEY RUSOV², and JAMES WYNNE³ — ¹SUPA, School of Physics and Astronomy, University of Glasgow, Glasgow, UK — ²Physik Department, Universität Siegen, Walter-Flex-Str. 3, Siegen, Germany — ³IPPP, Department of Physics, University of Durham, UK We present comprehensive studies of the Two Higgs Doublet Model with Type I and Type II Yukawa couplings. To find bounds on the mass spectrum of these models, contributions to flavour, Higgs, and electroweak observables from the new scalars and couplings are considered, using theoretical constraints to inform these fits. We compare the results of these fits to those from direct searches at the

LHC, finding regions of parameter space allowed by both flavour and collider. In addition, we test the consequences of our results on electroweak baryogenesis and the possibility of generating a strong first order phase transition in the 2HDM.

T 85.3 Thu 16:45 T-H16

Location: T-H16

Benchmarking Di-Higgs Production in Various Extended Higgs Sector Models — •DUARTE AZEVEDO^{5,6}, HAMZA ABOUABID¹, ABDESSLAM ARHRIR¹, JAOUAD EL FALAKI³, PEDRO M. FERREIRA^{2,4}, MARGARETE MÜHLLEITNER⁵, and RUI SANTOS^{2,4} — ¹AbdelMalek Essaadi University, Faculty of Sciences and Techniques B.P 416, Tangier, Morocco — ²ISEL - Instituto Superior de Engenharia de Lisboa, Instituto Politécnico de Lisboa 1959-007 Lisboa, Portugal — ³EPTHE, Physics Department, Faculty of Science, Ibn Zohr University, Faculty of Sciences Agadir, Morocco — ⁴Centro de Física Teórica e Computacional, Faculdade de Ciências, Universidade de Lisboa, Campo Grande, Edifício C8 1749-016 Lisboa, Portugal — ⁵Institute for Theoretical Physics, Karlsruhe Institute of Technology, 76128 Karlsruhe, Germany — ⁶Institute for Astroparticle Physics, Karlsruhe Institute of Technology, 76344 Karlsruhe, Germany

We study di-Higgs production in various extended Higgs sector models such as the real and complex 2HDM, the 2HDM with a real singlet and the next-to-Minimal Sypersymmetric Standard Model. We study the process $p_{t,j} = h_{j}$ with $h_{i=h_{j}} = h_{j}$ the 125 GeV scalar and the case where $h_{i} = h_{j}$ and/or h_{j} are new scalars predicted by the models. When performing the parameter scan, we consider all relevant constraints. The di-Higgs production cross section in these models can exceed the Standard Model rate by more than one order of magni-

tude. Furthermore we presenting our results for extended Higgs sector considered and list benchmark scenarios which exhibit important di-Higgs production rates at the LHC.

T 85.4 Thu 17:00 T-H16 Phenomenology of unusual top partners in composite Higgs models - GI-ACOMO CACCIAPAGLIA¹, THOMAS FLACKE², •MANUEL KUNKEL³, and WERNER POROD³ - ¹Université Lyon 1, Villeurbanne, France - ²Center for AI and Natural Sciences (KIAS), Seoul, Korea — ³Institut für Theoretische Physik und Astrophysik, Universität Würzburg, Germany

We consider a particular composite Higgs model which contains SU(3) color octet top partners besides the usually considered triplet representations. Moreover, color singlet top partners are present as well which can in principle serve as dark matter candidates. We investigate the LHC phenomenology of these unusual top partners. Some of these states could be confused with gluinos predicted in supersymmetric models at first glance.

T 85.5 Thu 17:15 T-H16

HiggsTools: a toolbox for BSM scalar phenomenology – •STEVEN PAASCH¹, CHENG LI¹, JONAS WITTBRODT¹, THOMAS BIEKOETTER¹, HENNING BAHL¹, GEORG WEIGLEIN¹, and SVEN HEINEMEYER² – ¹Deutsches Elektronen-Synchrotron DESY – ²Instituto de Fisica Teorica (UAM/CSIC), Universidad Autonoma de Madrid

The codes HiggsBounds and HiggsSignals compare model predictions of BSM models with extended scalar sectors to searches for additional scalars and measurements to the 125GeV Higgs boson. We present a unification and extension of the functionalities provided by both codes into the new HiggsTools framework. The codes have been re-written in modern C++ with a native python interface for easy interactive use. We discuss the user interface for providing model predictions, now part of the new sub-library HiggsPredictions, which also provides access to many tabulated cross sections and BRs in reference models such as the SM. HiggsBounds now implements experimental limits purely through json data files and can better handle clusters of BSM particles of similar mass, even for complicated search topologies. In HiggsSignals, the treatment of different types of measurements has been unified, both in the χ^2 computation and in the data file format used to implement experimental results.

T 85.6 Thu 17:30 T-H16

First-order strong-field QED processes including the damping of particle states — • TOBIAS PODSZUS and ANTONINO DI PIAZZA — Max Planck Institute for Nuclear Physics, Heidelberg, Germany

Strong field QED considers electrodynamic processes in the presence of an electromagnetic background field. Here 'strong' refers to the intensity of the background field, which is so high that the interactions of electrons and positrons with the background field have to be taken into account exactly in the calculations. This is done by implementing the background field in the quantization procedure of the fermion field. Exact solutions of the corresponding Dirac equation in the presence of an arbitrary plane wave field are the so called Volkov states. However, Volkov states, as well as free photon states, are not stable in the presence of the background plane-wave field but 'decay' as electrons/positrons can emit photons and photons can transform into electron-positron pairs. By using the solutions of the corresponding Schwinger-Dyson equations within the locally constant field approximation, we compute the probabilities of nonlinear single

Compton scattering and nonlinear Breit-Wheeler pair production by including the effects of the decay of electron, positron, and photon states. As a result, we find that the probabilities of these processes can be expressed as the integral over the light-cone time of the known probabilities valid for stable states per unit of light-cone time times a light-cone time-dependent exponential damping function for each interacting particle.

T 85.7 Thu 17:45 T-H16

QCD equation of state via the complex Langevin method - •FELIPE ATTANASIO¹, BENJAMIN JÄGER², and FELIX ZIEGLER^{2,3} — ¹Institute for Theoretical Physics, Universität Heidelberg, Philosophenweg 16, D-69120, Germany -²CP3-Origins & Danish IAS, Department of Mathematics and Computer Science, University of Southern Denmark, Campusvej 55, 5230 Odense M, Denmark — ³School of Physics and Astronomy, The University of Edinburgh, EH9 3FD Edinburgh, United Kingdom

The equation of state of hadronic matter is of high importance for many fields, ranging from heavy-ion collisions to neutron stars. Non-perturbative methods to simulate QCD encounter difficulties at finite chemical potential μ due to the so-called sign problem. We employ the complex Langevin method to circumvent this problem and carry out simulations at a variety of values for temperature and μ . We present results on the pressure, energy and entropy equations of state, as well as a numerical observation of the Silver Blaze phenomenon.

T 85.8 Thu 18:00 T-H16

High-energy diffractive processes without and with soft photon radiation and Low's theorem — •PIOTR LEBIEDOWICZ¹, OTTO NACHTMANN², and ANTONI $\mathsf{Szczurek}^1-{}^1\mathsf{Institute} \text{ of Nuclear Physics Polish Academy of Sciences, Krakow,}$ Poland — 2 Institut für Theoretische Physik, Universität Heidelberg, Heidelberg, Germany

Exclusive high-energy reactions at small momentum transfer without and with soft photon radiation are discussed. The tensor-pomeron model describing such reactions involving hadrons is introduced. The tensor-pomeron model is for hadronic high-energy reactions and has its origin in general investigations of the soft, nonperturbative, pomeron in QCD. The emission of soft radiation provides a fundamental probe of the consistency of the underlying Quantum Field Theory. As an example pion-pion scattering without and with soft photon emission is considered. The "exact" model results are compared to soft photon approximations based on Low's theorem. The term of order (photon energy)⁰ in the expansion of the photon emission amplitude needs to be changed compared to Low's result.

T 85.9 Thu 18:15 T-H16

Location: T-H17

Decay constants of B, B_s , and B_c mesons — •MATTHEW BLACK and OLIVER WITZEL - Physik Department, Universität Siegen, Walter-Flex-Str. 3, Siegen, Germany

We present the status of our ongoing work extracting B-meson decay constants f_B with $B = B_0, B^+, B_s, B_c$. Our calculation is based on $N_f = 2 + 1$ dynamical flavour gauge field ensembles generated by the RBC/UKQCD collaborations using domain wall fermions and the Iwasaki gauge action. Using domain wall light, strange, and charm quarks and relativistic b-quarks we obtain results at three different lattice spacings $a \approx 0.11, 0.08, 0.07$ fm and multiple valence quark masses. We perform a global fit to obtain phenomenologically relevant results in the continuum and at physical quark masses.

T 86: Neutrino Physics (Theory) 1 and Theoretical Astroparticle Physics and Cosmology 1

Time: Thursday 16:15-18:00

T 86.1 Thu 16:15 T-H17

Majoron Models without Domain Walls — •TIM BRUNE — TU Dortmund In the singlet Majoron model, neutrino masses arise due to the spontaneous violation of lepton number. Apart from providing a viable dark matter candidate, the model can also explain the generation of a baryon asymmetry via Leptogenesis. However, the Majoron model suffers from the existence of cosmological domain walls, contradicting observations. I present extensions of the Majoron model by additional right-handed doublets and triplets that render it domain wall free. Besides avoiding the existence of domain walls, the additional particles have a non-trivial effect on the Sphaleron process that converts the lepton asymmetry to a baryon asymmetry in Leptogenesis models.

T 86.2 Thu 16:30 T-H17

Search for Fermionic Dark Matter using Astrophysical Neutrinos and Quantum Gravitational Decoherence — HEINRICH PÄS, •DOMINIK HELLMANN, and ERIKA RANI — TU Dortmund, Dortmund, Deutschland

A model involving quantum gravitationally induced decoherence is proposed to investigate on the properties of fermionic dark matter using astrophysical neutrinos.

The main assumption of the model is that interactions of particles with the spacetime foam violate global quantum numbers such as lepton number and only conserve unbroken gauge quantum numbers. Hence, if ${\it N}$ hypothetical fermionic dark matter species exist transforming as a singlet under $SU(3)_c \times U(1)_{EM}$, quantum gravity interactions cannot distinguish between neutrinos and these unknown degrees of freedom.

Applying this phenomenological 3 + N flavor model to systems of high energy neutrinos shows that these effects lead to a uniform flavor distribution over all neutral fermionic species in an initially pure neutrino beam after sufficiently long distances. Therefore, fluxes of neutrinos from astrophysical origin are expected to differ drastically from the standard expectation depending on the number of additional dark matter fermions present.

Consequently, future neutrino experiments could provide new clues about the fermionic dark sector.

T 86.3 Thu 16:45 T-H17

CPT Violation in Neutrino Oscillations with Quantum-Gravitational De**coherence** – •ERIKA RANI^{1,2}, HEINRICH PÄS¹, and DOMINIK HELLMANN¹ -¹TU Dortmund, Lehrstuhl für Theoretische Physik III — ²UIN Maulana Malik Ibrahim Malang, Indonesia

Quantum gravity effects can lead to the violation of fundamental symmetries, including global symmetries such as flavor, baryon and lepton number and discrete symmetries such as CPT. In this talk, we investigate CPT violation in neutrino oscillations in the context of quantum-gravitational decoherence. By employing the Ellis, Hagelin, Nanopoulos, and Srednicki (EHNS) formalism we study a two-flavor scenario and discuss the size of the effect in different parametrizations.

T 86.4 Thu 17:00 T-H17 "Magnetic" Mass of the Neutron - •MANFRED GEILHAUPT - HSN, Mönchengladbach, Germany

In Quantum Physics, the Spin of an elementary particle is defined to be an intrinsic, inherent property. The same to the magnetic moment (μ) due to the spin of charged particles - like Electron (me) and Proton (mp). So the intrinsic spin (S=1/2h-bar) of the electron entails a magnetic moment because of charge (e). However, a magnetic moment of a charged particle can also be generated by a circular motion (due to spin) of an electric charge (e), forming a current. Hence the orbital motion (of charge around a mass-nucleus) generates a magnetic moment by Ampère's law. This concept must lead to an alternative way calculating the neutrino mass (mv) while looking at the beta decay of a neutron into fragments: proton, electron, neutrino and corresponding kinetic energies. The change of neutrons magnetic moment (μn) during the decay process is a fact based on energy and spin and charge conservation, so should allow to calculate the restmass of the charge-less neutrino due to a significant change of: μ e= -9.2847647043(28)E-24J/T down to μ ev= -9.2847592533(28)E-24J/T (while assuming mv=0.30eV to be absorbed and if (g-2)/2 from QED is independent of mass). As always the last word has the experiment

T 86.5 Thu 17:15 T-H17 Impact of bound states on non-thermal dark matter production - •JULIAN Bollig — Albert-Ludwigs-Universität, Freiburg

In this talk I will discuss the influence of non-perturbative effects, namely Sommerfeld enhancement and bound state formation, on the cosmological production of non-thermal dark matter (DM). For this purpose, I will focus on a class of simplified models with t-channel mediators. These naturally combine the requirements for large corrections in the early Universe, i.e. beyond the Standard Model states with long range interactions, with a sizable new physics production cross section at the LHC.

I will show that the dark matter yield of the superWIMP mechanism is suppressed considerably due to the non-perturbative effects under consideration, which leads to a significant shift in the cosmologically preferred parameter space of non-thermal dark matter in these models. By revisiting the implications of LHC bounds on long-lived particles associated with non-thermal dark matter, I will conclude that testing this broad class of DM models at the LHC and its successors is a bigger challenge than previously anticipated.

T 86.6 Thu 17:30 T-H17

Constraining dark matter annihilation with cosmic ray antiprotons using neural networks — Felix Kahlhoefer¹, Michael Korsmeier², Michael $Krämer^1$, SILVIA MANCONI¹, and •KATHRIN NIPPEL¹ — ¹Institute for Theoretical Particle Physics and Cosmology (TTK), RWTH Aachen University, D-52056 Aachen, Germany — ²The Oskar Klein Centre for Cosmoparticle Physics, Department of Physics, Stockholm University, Alba Nova, 10691 Stockholm, Sweden

The interpretation of data from indirect detection experiments searching for dark matter annihilations requires computationally expensive simulations of cosmic-ray propagation. We present new methods based on Recurrent and Bayesian Neural Networks that significantly accelerate simulations of secondary and dark matter Galactic cosmic ray antiprotons by at least two orders of magnitude compared to conventional approaches while achieving excellent accuracy. This approach allows for an efficient profiling or marginalisation over the nuisance parameters of a cosmic ray propagation model in order to perform parameter scans for a wide range of dark matter models. We present resulting constraints using the most recent AMS-02 antiproton data on several models of Weakly Interacting Massive Particles.

T 86.7 Thu 17:45 T-H17

Constraining Inflationary Dynamics with the SKA - TANMOY MODAK¹, TILMAN PLEHN¹, •LENNART RÖVER¹, and BJÖRN MALTE SCHÄFER² — ¹Institut für Theoretische Physik, Universität Heidelberg, Germany- $^2\mathrm{Astronomisches}$ Recheninstitut, Universität Heidelberg, Germany

The SKA allows to map the distribution of neutral hydrogen in the Universe over a vast redshift range. The three-dimensional 21cm power spectrum found through this map can be used to perform precision tests not only for several astrophysical phenomena but also for early Universe cosmology. Even considering only a small redshift range it allows to significantly improve current constraints on the Hubble slow-roll parameters when combined with the CMB anisotropies measurement of Planck.

T 87: Electroweak Interactions (Exp.) 2

Time: Thursday 16:15-18:20

Group Report

T 87.1 Thu 16:15 T-H18 Latest results on vector boson scattering studies from CMS - •ANKITA MEHTA, ANDREAS HINZMANN, and STEFFEN ALBRECHT - UHH, Hamburg, Germany

Vector boson scattering (VBS) is a key production process to probe the electroweak symmetry breaking sector of the standard model and to look for new physics phenomena beyond the standard model indirectly. VBS measurements have entered into a precision era with availability of larger dataset from LHC runII and recent developments in the field of machine learning. A summary of the latest VBS cross section measurements in different channels from the CMS experiment will be presented. A future outlook for the measurement of longitudinally polarized scattering of the W and Z bosons in the HL-LHC scenario and a combination of ATLAS and CMS measurements will also be presented.

T 87.2 Thu 16:35 T-H18

Measurement of same-sign WW Production at 13 TeV with the ATLAS Experiment — •SHALU SOLOMON — Albert-Ludwigs University of Freiburg, Germany The electroweak production of same-sign W boson pairs is one of the pivotal processes to experimentally probe the electroweak symmetry breaking mechanism. With the Standard Model Higgs boson playing a key role in regularising the scattering amplitude of longitudinally polarised W bosons, and thanks to the large electroweak to strong production mode ratio, measuring same-sign WW production is an important tool to understand the electroweak sector.

The analysis of 2015-2016 LHC data at $\sqrt{s} = 13$ TeV by the ATLAS experiment resulted in the observation of electroweak WW production with a signal significance of 6.5 σ . With the entire Run 2 dataset, corresponding to an integrated luminosity of 139 fb⁻¹, the signal event yield has increased approximately by a factor of four, which gives the potential for the first differential cross-section measurement of this process.

The talk aims to present a summary of the analysis. The estimations of all the major backgrounds along with their uncertainties and behaviour in different validation regions are discussed. Fitting and unfolding methods including closure tests are presented. The expected event yields and the largest uncertainties on the cross-section are also shown.

T 87.3 Thu 16:50 T-H18

Location: T-H18

Study of polarization fractions in same-sign W boson scattering — $\bullet \texttt{Prasham}$ JAIN, BEATE HEINEMANN, and OLEG KUPRASH — Albert-Ludwigs-Universität Freiburg, Freiburg, Germany

Polarized same-sign W boson pair production is a crucial process to examine the electroweak symmetry breaking mechanism. A measurement of the fraction of longitudinally polarized W bosons, $W_L^{\pm}W_L^{\pm}$, directly probes the unitarization mechanism of the vector boson scattering amplitude through Higgs boson contributions, and is sensitive to potential new physics effects. This talk presents the validation of Monte Carlo samples for polarized WW production at the LHC and shows first results of applying machine learning methods for extracting the longitudinal polarization fraction.

T 87.4 Thu 17:05 T-H18

Polarized same-sign W boson scattering at the CMS experiment - THORSTEN CHWALEK¹, NILS FALTERMANN¹, ABIDEH JAFARI², THOMAS MÜLLER¹, and •Komal Tauqeer¹ — ¹Institut für Experimentelle Teilchenphysik (ETP), Karlsruher Institut für Technologie (KIT) - ²Deutsches Elektronen-Synchrotron (DESY), Hamburg

Polarized vector boson scattering (VBS) provides an opportunity for testing the Higgs mechanism in the electroweak sector of the standard model. At the LHC, the scattering of the weak gauge bosons can reveal the actual process by which they get their masses. In particular, the longitudinal polarized state of these bosons can reveal new information about the goldstone bosons of the electroweak symmetry breaking sector.

The most promising VBS channel for this type of study is same-sign WW scattering, which has a good balance between signal and backgrounds. In particular, the semi-leptonic decay channel provides a larger cross section than the fully leptonic decay channel; however, this channel faces large background contributions from V + jets process.

To increase the signal extraction, our study aims for advancements in the boosted W-jet tagging techniques to identify the W jet decaying into hadrons along with its charge and polarization. In this talk, I will discuss about the ongoing work to identify the W jet charge via ParticleNet algorithm.

T 87.5 Thu 17:20 T-H18

A Search for anomalous couplings in the hadronic decay channel of Vector Boson Scattering at the CMS experiment — STEFFEN ALBRECHT², THORSTEN CHWALEK¹, NILS FALTERMANN¹, THOMAS MÜLLER¹, and •MAX NEUKUM¹ — ¹Institut für Experimentelle Teilchenphysik (ETP), Karlsruher Institut für Technologie (KIT) — ²Institut für Experimentalphysik, Universität Hamburg

Vector boson scattering (VBS) is the dominating process to investigate the quartic vertex of electroweak theory at the LHC. Additionally, cancellations with contributions from the Higgs boson are required to ensure unitarity. New physics in the Higgs sector may thus alter the cross section noticeably even if it is currently out of reach of direct measurements.

Deviations of couplings at high energies are formulated in an effective field theory, a bottom-up approach, which parametrizes a multitude of explicit theories. Limits on introduced parameters allow to draw conclusions regarding the strength and energy scale of new physics.

A search for anomalous couplings in the hadronic decay channel of VBS based on proton-proton collisions at a center-of-mass energy of 13 TeV is presented. Jet substructure techniques are used to distinguish between signal and background events and a three-dimensional fit supresses contributions from QCD events.

T 87.6 Thu 17:35 T-H18

Search for $\gamma\gamma \rightarrow WW(jjl\nu)$ coupling in *pp* collisions at $\sqrt{s} = 13$ TeV with the ATLAS detector — •VARSIHA SOTHILINGAM — Kirchhoff-Institut für Physik, Heidelberg

Due to the non-abelian nature of the electroweak sector of the Standard Model of Particle Physics (SM), direct interactions between gauge couplings is possible. Measurements of its cross section allow for validation of the SM and potential deviations opens possibilities for physics beyond the SM. This talk is focused on the coupling between W bosons and photons where the W bosons decay semileptonically. They interact via the triple (γW^+W^-) and quartic $(\gamma \gamma W^+W^-)$ gauge couplings of the SM. This process can be produced via centrally exclusive production at the LHC, which provides the signal while keeping the protons intact.

These protons can be detected using the ATLAS Forward Proton (AFP) spec-

trometers, which are located around 200m away from the ATLAS detector, on both sides. Using the central detector to measure the WW decay and protons measured using AFP, the four momenta of the final state is known. This talk gives a deeper overview of this rare process and the methods used to define a mass signal region of interest to perform the searches for this process. Additionally it will provide an overview of the pile up and background challenges which have to be managed in this analysis.

T 87.7 Thu 17:50 T-H18

Automated Resummation of Electroweak Sudakov logarithms using SCET_{EW} — •STEFAN RODE — Julius-Maximilians-Universität Würzburg

The search for new physics at the LHC necessitates the precise comparison between theoretical predictions in the Standard Model and experimental data. The calculation of electroweak radiative corrections, required to reach a sub-percent accuracy on the theory side, yields potentially large logarithmic corrections of the form $\alpha \log(s/M_W^2)$ to each order in perturbation theory. While at the LHC these corrections can already reach several tens of percent, this problem will be even more relevant at future colliders, where a resummation of these corrections will be inevitable to maintain the predictive power of the calculation.

We present a framework for the automated calulation of observables including the resummation of large logarithmic corrections using a generalisation of Soft-Collinear Effective Theory (SCET) to spontaneously broken gauge theories. We present results for *W*-boson pair production at the LHC and compare with the well-known results obtained using fixed-order methods.

T 87.8 Thu 18:05 T-H18

Upper limits on tri nucleon decays from GERDA — •SEAN SULLIVAN for the GERDA-Collaboration — MPIK, Heidelberg, Germany

Tri-nucleon decay is a hypothetical process violating baryon number conservation allowed for in the standard model with an extension for light neutrino masses [1,2]. The GERDA search for neutrinoless double-beta decay employs HPGe detectors enriched in ⁷⁶Ge which could undergo tri-nucleon decay to unstable daughter nuclei. This work investigates the tri-nucleon decay chain of ⁷⁶Ge to ⁷³Ga and the subsequent beta decay and gamma cascade to the stable daughter ⁷³Ge. This allows setting limits on the tri-nucleon decay channels without the need to consider the exotic physics of the initial decay.

[1]S.I. Alvis et al. (Majorana Collaboration), 2019. Search for Tri-Nucleon Decay in the Majorana Demonstrator Phys. Rev. D 99, 072004

[2] Babu, K.S., Gogoladze, I. and Wang, K., 2003. Gauged baryon parity and nucleon stability. Physics Letters B, 570(1-2), pp.32-38.

T 88: Top Quarks: Production (Exp.) 3

Time: Thursday 16:15-18:30

T 88.1 Thu 16:15 T-H19

Studies of modern generators for top-quark pair-production — •DOMINIC HIRSCHBÜHL, JENS ROGGEL, and WOLFGANG WAGNER — Bergische Universität Wuppertal. Deutschland

The precise simulation of $t\bar{t}$ processes is crucial for precision tests of the Standard Model and in the search for new physics at the Large Hadron Collider. Several important new event generators for the simulation of $t\bar{t}$ production have been released in the last years. The "*bb4ℓ*" generator is a NLO matrix-element generator for $pp \rightarrow b\bar{b}\ell^+\ell^-' v\bar{v}'$ final states implemented in the POWHEG-BOX. It includes theoretical improvements in the simulation of $t\bar{t}$ processes, which allows the production of $t\bar{t}$ and tW events including interference, off-shell effects, and top-quark decays at NLO. In this talk comparisons of these predictions with unfolded ATLAS data as well as with predictions for different DR and DS models implemented in MG5_aMC@NLO are shown. A second major improvement is the MINNLOPS implementation of t $t\bar{t}$ process in the POWHEG-BOX, which can be used to generate the $t\bar{t}$ production at NNLO. Studies of several parameter variations and an optimised setup for the NNLO generator matched with Pythia8 are presented.

T 88.2 Thu 16:30 T-H19

Machine learning approaches for parameter reweighting in MC samples of top quark production — •VALENTINA GUGLIELMI, KATERINA LIPKA, and SI-MONE AMOROSO — DESY, Hamburg, Germany

In high-energy particle physics, complex Monte Carlo simulations are needed to connect the theory to measurable quantities. Often, the significant computational cost of these programs becomes a bottleneck in physics analyses.

In this contribution, we evaluate an approach based on a Deep Neural Network to reweight simulations to different models or model parameters, using the full kinematic information in the event. This methodology avoids the need for simulating the detector response multiple times by incorporating the relevant variations in a single sample. Location: T-H19

We test the method on Monte Carlo simulations of top quark pair production, that we reweight to different SM parameter values and to different QCD models.

T 88.3 Thu 16:45 T-H19

Modelling $t\bar{t}$ + jets with the Sherpa fusing approach — •LARS FERENCZ¹, JU-DITH KATZY¹, FRANK SIEGERT², STEFAN HÖCHE³, CHRIS POLLARD⁴, and JO-HANNES KRAUSE² — ¹DESY, Hamburg, Germany — ²IKTP, Dresden, Germany — ³Fermilab, Chicago, United States — ⁴University of Oxford, Oxford, United Kingdom

Analyses interested in measuring the production of a top quark pair in association with a Higgs boson decaying into a $b\bar{b}$ final state rely on the MC modelling or irreducible backgrounds like $t\bar{t} + b\bar{b}$, $t\bar{t} + c\bar{c}$ and $t\bar{t}$ +light jets. Recently MCs with massive b-quarks in the $t\bar{t}b\bar{b}$ ME@NLO (so called 4FS) predictions became available and are used for the $t\bar{t}b\bar{b}$ related backgrounds. However, in order to predict the background for the analysis, they need to be combined with calculations of $t\bar{t} + c\bar{c}$ and $t\bar{t} + \text{light jets}$ taken from a different calculation based on massless quarks (5FS). This talk will present a method for embedding a $t\bar{t}b\bar{b}$ 4FS prediction into a multi-leg tt+jets 5 FS prediction using the Sherpa MC generator. This method is called "Sherpa Fusing". This approach would make it possible to predict the various $t\bar{t}$ + jets components like $t\bar{t}$ + light jets, $t\bar{t}$ + c-jets and $t\bar{t}$ + b-jets in a single sample, while ensuring the coverage of the full phase space. This talk will introduce the fusing method and present results comparing the fused predictions to standalone $t\bar{t}$ + jets and $t\bar{t}b\bar{b}$ predictions.

T 88.4 Thu 17:00 T-H19

Studies of $t\bar{t}$ production with additional heavy flavour jets in *p*-*p* collision with the ATLAS detector — •LUCAS KLEIN, MAHSANA HALEEM, and RAIMUND STRÖHMER — Julius-Maximilians-Universität Würzburg

The production of $t\bar{t}$ -pairs with additional jets provides a strong test of quantum chromodynamics (QCD) predictions at high orders. Furthermore, this production represents as a significant background to rare SM processes (e.g. $t\bar{t}H$, $t\bar{t}t\bar{t}$),

as well as to processes beyond the standard model. The additional jets consisting of *b*-quarks originating from gluon splitting are particularly interesting in constraining uncertainties in the prediction of the process.

In this talk, we will show studies of $t\bar{t}$ -pair production with additional *b*-jets in the dileptonic top decay channel using full Run 2 ATLAS data from protonproton collision at $\sqrt{s} = 13$ TeV. Events are chosen by requiring an oppositelycharged $e\mu$ -pair and at least two *b*-jets in the final state as a baseline selection. The backgrounds originating from $t\bar{t}$ events with additional light- or *c*-flavour jets ($t\bar{t}l$, $t\bar{t}c$) misidentified as *b*-jets in exclusive 3 *b*-tagged jet and ≥ 4 *b*-tagged jet regions are estimated using a data-driven method.

The fiducial cross-sections for a phase space with at least one and at least two additional *b*-jets at particle-level will be shown and compared. Differential cross-section distributions for several variables in $\geq 3 b$ -jet and $\geq 4 b$ -jet regions will also be shown and compared to several theoretical predictions.

T 88.5 Thu 17:15 T-H19

Measurement of the associated production of top-antitop pair with charm jets using the DL1r b-tagging algorithm in the dileptonic final state — •MORITZ HABBABA, ARNULF QUADT, and ELIZAVETA SHABALINA — II. Physikalisches Institut, Georg-August-Universität Göttingen

Top-antitop pair production production accompanied by a quark-antiquark pair is an important background for many measurements and searches. Top- antitop pair production in association with a bottom-antibottom pair has been measured with relatively good precision inclusively and differentially using powerful b-tagging tools. The b-tagging algorithms have a long and stable history within the analysis tools of particle physics. Dedicated tagging algorithms for charm quarks on the other hand have not been widely used. This work investigates the possibilities of using existing b-tagging tools for c-tagging. The expected signal purity that can be achieved by using DL1r b-tagging algorithm for the purpose of c-jet tagging is investigated. The optimal signal regions in the dilepton ttbar final state are defined and expected uncertainty on the tt + cc cross section is evaluated.

T 88.6 Thu 17:30 T-H19

Results of differential tī+bb measurement in the lepton+jets channel at the CMS experiment — •JAN VAN DER LINDEN, ULRICH HUSEMANN, and EMANUEL PFEFFER — Institute of Experimental Particle Physics (ETP), Karlsruhe Institute of Technology (KIT)

Due to the large mass difference between the light bottom quarks and the heavy top quarks, the modeling of the associated production of bottom quarks with a pair of top quarks ($t\bar{t}$ + $b\bar{b}$) is very challenging and is still associated with large uncertainties today. Past measurements of the inclusive cross section of that process also showed discrepancies between the predicted and measured cross sections, which can be attributed to the challenging modeling. Furthermore, different Monte Carlo event generation methods show significant differences in the modeling of this process, for example in the different treatment of heavy flavors in the parton density functions or the simulation of the additional QCD radiation via parton showering or matrix element simulation.

Hence a measurement of the inclusive cross section, as well as the absolute and normalized differential cross section of $t\bar{t}+b\bar{b}$ production is performed at the CMS experiment in the lepton+jets decay channel of the $t\bar{t}$ system. The measurement will provide important input for the development and tuning of future Monte Carlo event generators, to describe the physics of the $t\bar{t}+b\bar{b}$ process more accurately.

In this talk results of this analysis, using data from the full Run-2 period of the LHC, are discussed.

T 88.7 Thu 17:45 T-H19

Simultaneous measurement of tt+X(bb) processes in the semileptonic channel at the CMS experiment — •RUFA RAFEEK, EMANUEL PFEFFER, JAN VAN DER LINDEN, MICHAEL WASSMER, and ULRICH HUSEMANN — Institute of Experimental Particle Physics (ETP), Karlsruhe Institute of Technology (KIT)

Top quark anti-quark pairs $(t\bar{t})$ are produced in association with other particles (X) where X can be the Higgs boson, Z/W boson or QCD-initiated heavy flavour jets $(b\bar{b}/c\bar{c})$. The measurement of $t\bar{t} + X$ is a direct probe of the coupling of standard model particles like the Higgs and Z boson to the top quark and may reveal new physics effects in modifications of these couplings.

The analysis is challenging as these processes, particularly when the bosons decay into heavy flavour quarks, like for example, $t\bar{t} + H(H \rightarrow bb)$ and $t\bar{t} + bb$ or $t\bar{t} + Z(Z \rightarrow bb)$, share the same signature and kinematic features. These high jet multiplicity final states create ambiguities in the reconstruction and identification of these processes and thus, it is hard to differentiate them from each other. Due to this challenge, an attempt to simultaneously measure these $t\bar{t} + X$ processes is made by exploring multivariate analysis strategies.

In this talk, an overview of the ongoing analysis, designed with the full Run-2 data of the LHC using the single lepton channel, is given.

T 88.8 Thu 18:00 T-H19

Graph neural network applications in tī+heavy flavor studies at the CMS experiment — •EMANUEL PFEFFER, MAX ERHART, ULRICH HUSEMANN, RUFA RAFEEK, JAN VAN DER LINDEN, AND MICHAEL WASSMER — Institute of Experimental Particle Physics (ETP), Karlsruhe Institute of Technology (KIT)

Graph neural networks are a promising and novel method of artificial intelligence. In contrast to other classes of machine learning, graphs may be better at mapping relationships and dependencies among objects such as jets and at learning interconnections. The application of these networks is particularly interesting in processes consisting a pair of top quarks produced in association with bottom guarks (tīt+bb) and other indistinguishable processes such as tīt+H with H \rightarrow bb or tīt+Z with Z \rightarrow bb. Initial studies show promising results in identifying the b jets not originating from the decay of the top quarks in such processes. The graph neural network studies show better results than previous ones based on conventional deep neural networks.

This talk gives an overview of the applications and results of jet assignments in $t\bar{t}$ +heavy flavor processes at the CMS experiment using graph neural networks.

T 88.9 Thu 18:15 T-H19

Studies on $t\bar{t}+c\bar{c}/t\bar{t}+b\bar{b}$ **separation** — •Max Erhart, Ulrich Husemann, Jan van der Linden, Emanuel Pfeffer, and Rufa Rafeek — Institute of Experimental Particle Physics (ETP), Karlsruhe Institute of Technology (KIT)

Bottom jets and charm jets have very similar signatures and their identification techniques (b-tagging, c-tagging) use similar approaches, thus misidentification between these jets is common. Therefore the production of a top quark pair $(t\bar{t})$ in association with heavy flavor jets, i.e. charm and bottom $(t\bar{t}+c\bar{c}, t\bar{t}+b\bar{b})$, is an important, irreducible background to analyses targeting top quark pairs with additional bottom quark jets such as associated Higgs boson production with a subsequent Higgs boson decay into bottom quarks $(t\bar{t}+H(b\bar{b}))$. The modeling of the additional bottom or charm jets also suffers from large uncertainties in theroetical calculations due to the difference in mass scale for the large top quark mass and the relatively light bottom or charm quarks.

Due to the similarities of these processes a simultaneous measurement of the production cross section for heavy flavor jets in association with top quarks in the dileptonic final state of the top quarks is being conducted at the CMS experiment.

In this talk an overview of the ongoing analysis and the analysis strategies to separate the $t\bar{t}+b\bar{b}$, $t\bar{t}+H(b\bar{b})$, $t\bar{t}+Z(b\bar{b})$ and $t\bar{t}+c\bar{c}$ processes from one another is given.

T 89: Higgs Boson: Decay in Fermions 2

Location: T-H20

Time: Thursday 16:15-18:00

T 89.1 Thu 16:15 T-H20

Search for decays of boosted Higgs bosons to pairs of charm quarks with the CMS Experiment — •ANDRZEJ NOVAK, LUCA MASTROLORENZO, XAVIER COUBEZ, SPADAN MONDAL, ANDREY POZDNYAKOV, and ANDRZEJ NOVAK — Physics Institute III A, RWTH Aachen

The Higgs boson decay into charm quarks has the highest branching fraction of the yet unobserved decays. Moreover, it is predicted to be the strongest coupling to the second generation of fermions which as of now remains unconfirmed. This talk presents a search for the Higgs boson in the gluon fusion production mode with high Lorentz boosts, decaying to a pair of charm quarks. The analysis is modelled on a previous analysis of decays to pairs of bottom quarks and is enabled by recent developments in deep learning based tools for jet identification in such topologies. Probing this channel is not only important for completeness, but it could also be sensitive to potential beyond Standard Model corrections.

T 89.2 Thu 16:30 T-H20

Search for Higgs boson decay to a pair of charm quarks in a two-jets topology at CMS with full Run-2 data. — •ANDREY POZDNYAKOV¹, BJORN BURKLE², XAVIER COUBE2^{1,2}, ALENA DODONOVA¹, LUCA MASTROLORENZO¹, SPANDAN MONDAL¹, ANDRZEJ NOVAK¹, and ALEXANDER SCHMIDT¹ — ¹RWTH University, Germany — ²Brown University, USA

Full Run-2 data of the CMS experiment has been analyzed in order to obtain the most sensitive result to date on the measurement of the Higgs boson coupling to charm quarks. The coupling of the Higgs boson to charm is probed in a direct search for the H \rightarrow cc decay, when the Higgs boson is produced in association with a W or Z boson. The full analysis is carried out in two topologies : boosted, where the two jets from a Higgs boson candidate are merged into one large-cone jet, and resolved, where two small-radius jets are reconstructed. In this talk we present a detailed overview of the resolved part of the full analysis. Major developments are introduced compared to the previously published CMS result, based on a partial data set. Those include an improved jet flavor tagging for charm jets based on DNNs, a dedicated jet-energy regression, a "kinematic fit" that constrains momenta of the jets using leptons from an associated Z boson decay. All these improvements lead to a stringent constraint on $|\kappa_c|$, which pose a limit for BSM models with large κ_c variations.

T 89.3 Thu 16:45 T-H20

Deep Sets for the ttH(H->bb) cross section measurement — •JOSÉ MANUEL CLAVIJO COLUMBIÉ and JUDITH KATZY — Deutsches Elektronen-Synchrotron DESY, Notkestraße 85, D-22607 Hamburg

We present our analysis strategy for the measurement of the cross-section of the ttH(H->bb) process differential in Higgs pT. We use a newly developed Deep Set neural network to split the events into separate regions which are enriched in signal or any of the main background processes. This allows the signal measurement together with good control of the background normalization factors and their associated uncertainties. In addition, we split the signal-like events into different Higgs pT ranges for the differential cross-section measurement. We perform an Asimov fit which gives the expected uncertainties of the measurement, and a fit of the background regions to real collision data to measure our ability to control the background processes.

T 89.4 Thu 17:00 T-H20 Effective Field Theory interpretations in Higgs boson pair production studies and constraints on the Higgs boson self-coupling — •CHRISTINA DIMITRIADI^{1,2}, JOCHEN DINGFELDER¹, ARNAUD FERRARI², TATJANA LENZ¹, and SERHAT ÖRDEK² — ¹Physikalisches Institut Universität Bonn, Germany — ²Uppsala University, Sweden

After the discovery of the Higgs boson in 2012, an important test of electroweak symmetry breaking would be to establish evidence of the Higgs boson self-coupling (λ_{HHH}), which can be achieved through a measurement of Higgs boson pair production. In the Standard Model (SM), di-Higgs events are dominantly produced in gluon-gluon fusion processes, e.g. involving the Yukawa coupling to top quarks (top quark loops) or via the Higgs boson self-coupling. These two production mechanisms interfere destructively, which leads to a very small di-Higgs production cross-section. However, deviations in couplings of the Higgs boson from the SM expectation could lead to a significant enhancement of the di-Higgs production rate.

A re-interpretation of the search for non-resonant Higgs boson pair production in the $bb\tau\tau$ channel, which is one of the most sensitive for probing the Higgs self-coupling, is presented. A scan of the self-coupling modifier, $\kappa_{\lambda} = \lambda_{HHH}/\lambda_{HHH}^{SM}$, is performed and limits on κ_{λ} are set. Projected sensitivity results at the High-Luminosity LHC are also discussed. Finally, preliminary studies for Higgs Effective Field Theory interpretations of the existing analysis are shown.

T 89.5 Thu 17:15 T-H20

Search for Higgs Boson Pair Production in Multi-Lepton Final States with the ATLAS Detector — Volker Büscher, Antoine Laudrain, Christian Schmitt, •Niklas Schmitt, and Duc Bao Ta — Johannes Gutenberg-Universität, Mainz

After the discovery of the Higgs boson in 2012 at the LHC, many of its properties

Thursday

have already been determined precisely using data of an integrated luminosity of 139 fb⁻¹. However, one of the biggest challenges in this field remains the measurement of the coupling of the Higgs boson to itself. It allows for a deep insight into the real shape of the Higgs potential and hence has a big impact on the understanding of fundamental interactions not only at the electroweak scale. In order to constrain the trilinear self-coupling, the Di-Higgs production cross section is measured. While decay modes including *b*-quarks typically have larger branching fractions, leptonic final states are generally much cleaner and have less SM background. Accordingly, probing this channel as a complement to $b\bar{b}$ analyses will be very promising.

Because of the small branching ratio and the large number of different SM backgrounds, it is difficult to investigate every leptonic *HH* decay mode individually. For this reason, dedicated neural networks in the 2,3 and 4 lepton final states have been trained to distinguish all re- levant signal processes against the sum of all backgrounds. This talk will introduce the analysis strategy and give an overview on the performance of the multi-lepton channel compared to other decay modes.

```
T 89.6 Thu 17:30 T-H20
```

Search for Higgs-boson pair production in the $bb\ell\ell + E_T^{miss}$ final state with the ATLAS detector — •Benjamin Rottler, Benoit Roland, and Markus Schumacher — Albert-Ludwigs-Universität Freiburg

The determination of the triple Higgs-boson self-coupling λ is one of the key goals of the physics program at current and future colliders. It will allow to reconstruct the Higgs potential. The self-coupling can be accessed via non-resonant Higgs-boson pair production, which can happen at the LHC via the destructively interfering top-loop and Higgs self-interaction diagrams.

The goal of this analysis is to measure the cross-section of the non-resonant Higgs-boson pair production σ_{HH} and the Higgs-boson self-coupling λ using the full Run-2 dataset collected by the ATLAS experiment, corresponding to an integrated luminosity of ~139 fb⁻¹ at $\sqrt{s} = 13$ TeV. This is done by considering the $bb\ell\ell + E_T^{\text{miss}}$ final state, which combines the high branching ratio of the $H \rightarrow bb$ decay and the good efficiency of lepton triggers. Our focus is on a combined search for the $HH \rightarrow bb(WW \rightarrow 2\ell 2\nu)$, $HH \rightarrow bb(\tau \tau \rightarrow 2\ell 4\nu)$ and $HH \rightarrow bb(ZZ \rightarrow 2\ell 2\nu/2\ell 2q)$ processes.

A multi-class deep neural network (NN) is used to separate signal and background processes on top of a loose preselection. The shape of the NN output distribution will be used in the statistical analysis. I will present the results of the σ_{HH} and λ measurements. I will also discuss the application of parametrized NNs (pNN) for the λ measurement, which allows to train a combined NN for multiple λ values.

T 89.7 Thu 17:45 T-H20

 $HH {\rightarrow} bb\tau\tau\tau$ Run 3 Trigger Studies — • Andrés Melo, Jason Veatch, and Stan Lai — II. Physikalisches Institut, Georg-August-Universität Göttingen

The HH \rightarrow bb $\tau\tau$ search is performed using data taken from the ATLAS experiment at the LHC, and allows us to probe the Higgs self-coupling (for non-resonant decays) and the existence of heavier particles from which both Higgs would have been produced (for resonant decays). Since the Higgs boson is identified by its decay products, tau triggering is an important and crucial part of the search.

In order to ensure optimal trigger efficiency for this final state signature in Run-3, the efficiency at the high level trigger (HLT) with respect to the Level-1 trigger (L1) is studied. A characterization of various kinematic variables and quantities of the events that do not pass the HLT triggers is performed. This study allows for an improvement of the trigger efficiency, minimizing event loss.

T 90: Higgs Boson: Associated Production 2

Time: Thursday 16:15-18:15

T 90.1 Thu 16:15 T-H21 Search for the standard model Higgs boson in associationwith a bottomquark pair (bbH) — •MARYAM BAYAT MAKOU — CMS-DESY, Hamburg, Germany

One of the main goals of the LHC experiment is the precise measurement of the Higgs boson production mechanisms to clarify its coupling structure. In the Standard Model of particle physics, the coupling of the Higgs boson to fermions is introduced via the Yukawa interaction. Up to now the Yukawa coupling to b-quarks (y_b) was measured only in the decay process, and not yet in the production mechanism due to the low cross-section and the overwhelming background processes.

This measurement aims at measuring the b-associated Higgs production (bbH) using data collected by the CMS experiment during Run 2. The study covers events where the Higgs boson is produced through the bbH channel and further decays into two tau leptons, subsequently fully leptonically ($\tau_e \tau_\mu$) or fully

Location: T-H21

hadronically ($\tau_h \tau_h$). A machine learning approach has been used to classify the events into two Higgs signal classes and several background classes. First results on the sensitivity on the bbH production channel will be shown.

T 90.2 Thu 16:30 T-H21

Higgsstrahlung with H→bb decay at NNLO+PS — •SILVIA ZANOLI¹, MAURO CHIESA², EMANUELE RE^{3,4}, MARIUS WIESEMANN¹, and GIULIA ZANDERIGHI^{1,5} — ¹Max-Planck-Institut für Physik, Föhringer Ring 6, 80805 München, Germany — ²Dipartimento di Fisica, Università di Pavia, and INFN, Sezione di Pavia, Via A. Bassi 6, 27100 Pavia, Italy — ³INFN, Sezione di Milano - Bicocca, and Università di Milano - Bicocca, Piazza della Scienza 3, 20126 Milano, Italy — ⁴LAPTh, Université Grenoble Alpes, Université Savoie Mont Blanc, CNRS, 74940 Annecy, France — ⁵Physik-Department, Technische Universitä München, James-Franck-Strasse 1, 85748 Garching, Germany

I present the computation of HZ and HW production followed by the Higgs bo-

son decay into a pair of bottom quarks within the MiNNLO_{PS} framework. Both the production and the decay of the Higgs boson are evaluated at NNLO+PS accuracy, consistently combined and then matched to the parton shower by means of a vetoed shower. This MiNNLO_{PS} calculation supersedes the previous NNLOPS results, obtained through MiNLO'+reweighting. I will show and discuss phenomenological results for LHC collisions at 13 TeV.

T 90.3 Thu 16:45 T-H21

Back to the Formula - LHC Edition — ANJA BUTTER¹, TILMAN PLEHN¹, •NATHALIE SOYBELMAN², and JOHANN BREHMER³ — ¹Institut für Theoretische Physik, Heidelberg, Deutschland — ²Weizmann Institute of Science, Rehovot, Israel — ³New York University, New York, USA

While neural networks offer an attractive way to numerically encode functions, actual formulas remain the language of theoretical particle physics. We use symbolic regression trained on matrix-element information to extract, for instance, optimal LHC observables. This way we invert the usual simulation paradigm and extract easily interpretable formulas from complex simulated data. We introduce the method using the effect of a dimension-6 coefficient on associated ZH production. We then validate it for the known case of CP-violation in weakboson-fusion Higgs production, including detector effects.

T 90.4 Thu 17:00 T-H21

Deep-Learning driven Signal Extraction in the Associated VH-Production with the CMS Experiment — •NICLAS EICH, SVENJA DIEKMANN, and MARTIN ERDMANN — RWTH Aachen

The Higgs Boson production in association with a Vector Boson known as "Higgs-Strahlung" (VH) is one of the four main production modes of the Higgs at the LHC. The VH-production can be divided into two modes, being produced by quark-annihilation and gluon-fusion for a Z boson respectively. In our analysis, we aim to measure the gluon-fusion process in the final states with the Higgs decaying to a b-quark pair and the W/Z decaying leptonically. We make use of the symmetries between the quark and gluon induced processes and deploy Deep Learning techniques to maximize the sensitivity of the measurement. Finally, we present first results towards the signal extraction using the data-taking period 2017 of the CMS experiment.

T 90.5 Thu 17:15 T-H21

Extracting the Gluon Fusion Component of the Associated ZH Production with the CMS Experiment — •SVENJA DIEKMANN, NICLAS EICH, and MARTIN ERDMANN — III. Physikalisches Institut A, RWTH Aachen University

The gluon fusion production mechanism of the associated ZH production $(gg \rightarrow ZH)$ is a yet unmeasured Standard Model process sensitive to various new physics scenarios. The considered final state of two leptons and two b-jets is not only populated by large backgrounds arising from other processes, but also by the dominant quark initiated ZH production $(q\bar{q} \rightarrow ZH)$. In order to separate these two production mechanisms, the total ZH production can be utilised in combination with the WH production to extract the gluon fusion component by analysing the ratio of their cross sections. The strategy of this analysis to extract the gluon fusion component of the ZH production is demonstrated using the data-taking period 2017 of the CMS experiment.

T 90.6 Thu 17:30 T-H21 Extraction of the gluon-initiated component of the associated production of the Higgs boson and a vector boson with the CMS experiment - •ALENA DODONOVA¹, ALEXANDER SCHMIDT¹, XAVIER COUBEZ^{1,2}, LUCA

Time: Thursday 16:15–17:45

T 91.1 Thu 16:15 T-H22

Search for the Higgs Boson decay to a Z boson and a photon — •MING-YAN LEE — RWTH III. Physikalisches Institut A, Aachen, Germany

The results of the search for the Higgs boson decays to $Z\gamma$ will be presented. In the search for this rare Higgs boson decay, the leptonic channel is most promising as it has relatively low background and can be fully reconstructed in the CMS detector. A few analysis techniques such as multivariate analysis methods, kinematic fit and final state radiation recovery are introduced to improve the sensitivity. The combination of the Run 2 data results in an observed (expected) upper limit on the signal strength of 4.1 (1.8) at 95% confidence level.

T 91.2 Thu 16:30 T-H22 **Higher order QCD corrections and effective operators in Higgs boson pair production** — GUDRUN HEINRICH¹, •JANNIS LANG¹, and LUDOVIC SCYBOZ² — ¹Karlsruhe Institute of Technology (KIT) — ²University of Oxford, UK We present results for Higgs boson pair production in gluon fusion including both, NLO (2-loop) QCD corrections with full top quark mass dependence as well as anomalous couplings related to operators describing effects of physics Associated Higgs boson production with a Z boson (ZH) contains quark- and gluon-initiated components. The gluon-initiated component ($gg \rightarrow ZH$) could be a good probe for the physics beyond the Standard Model (SM) since the effects of the new physics for the loop-induced processes would be of the same order as the SM process. Due to destructive interference between box and triangle contributions at the leading order, this component is suppressed with respect to the dominant quark-initiated contribution to ZH production.

In this talk, I will present the prospects to measure the upper limit on the $gg \rightarrow ZH$ component in the $H \rightarrow b\bar{b}$ decay channel using multivariate analysis. The study is performed with the full Run 2 dataset collected with the CMS detector at the LHC at $\sqrt{s} = 13 \ TeV$.

T 90.7 Thu 17:45 T-H21

Simplified Template Cross Section Measurement of $pp \rightarrow WH \rightarrow WWW \rightarrow lvlvlv - \bullet$ Moritz Hesping, Volker Büscher, Ralf Gugel, and Christian Schmitt - Johannes Gutenberg Universität Mainz

The measurement of the couplings of the Higgs boson is of great scientific interest, since it has the potential of testing possible extensions to the Standard Model. The decay of a Higgs boson into a pair of W bosons after production in association with a W boson is especially useful, since in this process the Higgs boson exclusively couples to W bosons.

In addition to the inclusive analysis of the full run 2 dataset of the ATLAS experiment, the $pp \rightarrow WH \rightarrow WWW \rightarrow lvlvlv$ process was measured in the scheme of Simplified Template Cross Sections (STXS). The STXS mesurement of this process requires access to the transverse momentum of the associated W boson, which due to the presence of three neutrinos in the final state cannot be fully reconstructed. In this talk the analysis strategy will be presented, with a focus on the regression neural network for the W momentum reconstruction and the multiclassifier network for signal-background separation.

T 90.8 Thu 18:00 T-H21 **Prospects of measuring the Higgs self-coupling at the ILC** – •JULIE TORNDAL^{1,2}, JENNY LIST¹, and YASSER RADKHORRAMI^{1,2} – ¹Deutsches Elektronen-Synchrotron DESY, Hamburg – ²Universität Hamburg, Hamburg, Germany

The Higgs-self coupling is a key feature in understanding the Higgs potential and hence the underlying mechanism that provides mass to the elementary particles. Any deviations from the Standard Model (SM) could indicate new physics beyond the SM.

Experimentally, the measurement of the Higgs self-coupling poses many challenges due to the very small production cross sections and multi-jet final states. The International Linear Collider offers a clean experimental environment at energies above the threshold of double Higgs-strahlung. Since the Higgs selfcoupling was last studied (5-8 years ago) there have been tremendous progress for high-level reconstruction tools which could lead to a large improvements in the sensitivity.

In this contribution, we revisit the projections made for the Higgs selfcoupling measurement at the ILC at a center-of-mass energy of 500 GeV and study the effects of improvements made in the high-level reconstruction tools.

-

beyond the Standard Model. The latter can be realized in non-linear (HEFT) or linear (SMEFT) Effective Field Theory frameworks. We show results for both and discuss the impact of different approximations within the SMEFT description.

T 91.3 Thu 16:45 T-H22

Location: T-H22

Employing Matrix Elements with Neural Networks to Search for Higgs Selfcoupling — •CHRISTOPH AMES, OTMAR BIEBEL, and LARS LINDEN — Ludwigs-Maximillians-Universität, Munich

The Higgs boson was discovered in 2012 as predicted by the Standard Model, however, not all of its predicted couplings have been measured yet. One such coupling is the Higgs self-coupling, in which a Higgs boson decays into two further Higgs bosons. By integrating over all possible initial states and by using the details of the end state, the matrix element method evaluates the weight of an event for the specific production cross section. In this work, machine learning is combined with the matrix element method to search for $HH \rightarrow b\bar{b}W^+W^-$ using simulated data. A neural network is trained to calculate the matrix element weight of an event and to use this to determine whether the event contains a signal or a background decay.

T 91: Higgs Boson: Rare Decays

T 91.4 Thu 17:00 T-H22

Investigating Sensitive Observables as Input Variables for Neural Networks in the Search for Higgs Self-Interaction — •LARS LINDEN, CHRISTOPH AMES, and OTMAR BIEBEL — Ludwig-Maximilians-Universit\"at, M\"unchen

A precise measurement of Higgs boson self-interaction is important to determine the properties and the shape of the Higgs potential. Deviations of the expected shape may possibly hint to new physics phenomena. However, the crosssection of the Higgs self-interaction process is very small. So, neural networks are employed to enhance the experimental sensitivity for this process. These networks require a training using specific input observables sensitive to the Higgs self-interaction final states. This talk presents such observables for Higgs boson pair production via the process \$gg \to HH\$ with a special focus on \$HH\to b\bar{b}W^+W^-\$ final states.

T 91.5 Thu 17:15 T-H22

Search for Higgs boson pairs decaying to multilepton final states with the CMS experiment — •TOBIAS KRAMER, TORBEN LANGE, MARCEL RIEGER, and PETER SCHLEPER — Institut für Experimentalphysik, Universität Hamburg In this presentation, the first CMS analysis searching for Higgs boson pairs decaying into multilepton final states is presented. It uses the full Run 2 dataset corresponding to 138 fb⁻¹ recorded by the CMS experiment at a center of mass energy $\sqrt{s} = 13$ TeV. Several scenarios for producing events with SM Higgs boson pairs are considered, such as the decay of heavy resonances as well as non-

resonant production via the SM as well as EFT modified couplings. The talk

focuses on Higgs decays to pairs of W bosons and tau leptons. The analysis provides limits on the yet to be discovered trilinear Higgs self-coupling λ as well as cross section limits for different BSM scenarios, especially at high values for the coupling strength modifier κ_{λ} and low resonance masses.

T 91.6 Thu 17:30 T-H22

Search for di-Higgs Production with the CMS Experiment using the full Run 2 Dataset — •MARCEL RIEGER, PETER SCHLEPER, and TOBIAS KRAMER — Institut für Experimentalphysik, Universität Hamburg

After the discovery of the Higgs boson and measurements of its couplings to vector bosons as well as third generation fermions, the search for processes involving more than one Higgs boson will constitute one of the main goals of the global particle physics program in the coming years. The study of trilinear self-coupling will eventually give rise to the structure of the Higgs potential and can lead to profound theoretical consequences. Thereby, di-Higgs searches can gauge our understanding of electroweak symmetry breaking and probe a variety of scenarios that reach beyond the Standard Model.

The full Run 2 collision dataset recorded by the CMS experiment is analyzed in a wide range of potential di-Higgs final states to maximize the coverage of the available phase space. This talk highlights aspects of these searches with focus on final states involving two bottom quarks and two tau leptons. Prospects of the combination of channels, as far as available, are presented and constraints on the trilinear self-coupling as well as the coupling of two Higgs and two vector bosons are reported.

T 92: Higgs Boson: Extended Models 3

Time: Thursday 16:15-18:00

T 92.1 Thu 16:15 T-H23

Suche nach unsichtbaren Zerfällen des Higgs-Bosons in Ereignissen mit einem hadronisch zerfallenden Vektorboson mit dem ATLAS-Detektor — •JOHANNES BALZ, VOLKER BÜSCHER, DUC BAO TA und KIRA KÖHLER — Institut für Physik, Johannes Gutenberg-Universität Mainz

Eines der gegenwärtig größten Ziele für das ATLAS Experiment ist neben der präzisen Vermessung des Standardmodells (SM) die Suche nach Physik jenseits des SM.

In diesem Vortrag geht es um die Suche nach unsichtbaren Zerfällen des Higgs-Bosons jenseits des Standardmodells. Beim untersuchten Kanal wird das Higgs-Boson über die assoziierte Produktion mit einem Vektorboson erzeugt, wobei das beteiligte Vektorboson im weiteren Verlauf hadronisch zerfällt und das Higgs-Boson in für den Detektor unsichtbare Teilchen, zum Beispiel Dunkle Materie, zerfällt. Daher umfasst die Selektion Ereignisse mit hohem fehlendem Transversalimpuls und Jets, die kompatibel mit einem hadronisch zerfallenden Vektorboson sind. Dadurch kann der Hauptuntergrund $Z \rightarrow vv$ bereits stark unterdrückt werden. Die höchste Sensitivität liegt bei hohen Transversalimpulsen des Vektorbosons, das dann als ein großflächiger Jet rekonstruiert wird. Dessen Jetsubstrukturvariablen ermöglichen eine weitere Untergrundunterdrückung.

Im Vortrag wird der aktuelle Stand der Analyse bei einer Schwerpunktsenergie von $\sqrt{s}=13$ TeV vorgestellt.

T 92.2 Thu 16:30 T-H23

Higgs Decay into Dark Matter in the CxSM at Next-to-Leading Order •FELIX EGLE¹, MARGARETE MÜHLLEITNER¹, RUI SANTOS^{2,3}, and JOÃO VIANA $^2 - {}^1$ Institute for Theoretical Physics, Karlsruhe Institute of Technology, Wolfgang-Gaede-Str. 1, 76131 Karlsruhe, Germany — ²Centro de Física Teórica e Computacional, Faculdade de Ciéncias, Universidade de Lisboa, Campo Grande, Edifício C8 1749-016 Lisboa, Portugal — ³ISEL - Instituto Superior de Engenharia de Lisboa, Instituto Politécnico de Lisboa 1959-007 Lisboa, Portugal The Complex Singlet extension of the Standard Model (CxSM) is one of the simplest models to extend the scalar sector of the Standard Model (SM) and is also able to provide a suitable Dark Matter (DM) candidate. A possibility to probe the DM candidate at the LHC is given by the decay of the 125 GeV SM-like Higgs boson into a pair of DM particles. In order to match the experimental accuracy, higher-order corrections to this process have to be considered. We will present the computation of the complete next-to-leading order electroweak corrections to this decay process. In particular, we will describe the renormalization procedure for the CxSM, compare different renormalization schemes, discuss theoretical and experimental constraints on the input parameters and compare the results with current exclusion bounds.

T 92.3 Thu 16:45 T-H23

Dark Matter Phase Transitions in 'CP in the Dark' — •LISA BIERMANN, MAR-GARETE MÜHLLEITNER, and JONAS MÜLLER — Institute for Theoretical Physics, Karlsruhe Institute of Technology, 76128 Karlsruhe, Germany

We study the possibility of a strong first-order electroweak phase transition (SFOEWPT) in the extended scalar sector model 'CP in the Dark'. 'CP in the

Dark' provides a Dark Matter (DM) candidate as well as explicit CP violation in the dark sector. A global minimization of the one-loop daisy-resummed effective potential at finite temperature is performed with the C++ code BSMPT. We find a broad viable parameter space for an SFOEWPT within the reach of XENON1T and future invisible decay searches. 'CP in the Dark' also offers SFOEWPT points that display spontaneous CP violation at finite temperature. Having not only an SFOEWPT that provides the necessary departure from thermal equilibrium, but also a source of additional non-standard CP violation, opens a promising gate towards enabling the creation of the BAU in an electroweak baryogenesis scenario.

T 92.4 Thu 17:00 T-H23

Location: T-H23

Combination of Higgs boson measurements using Simplified Template Cross Sections with interpretations in the κ-framework at the ATLAS experiment — •JOSHUA CLERCX — Universität Hamburg/DESY, Hamburg, Germany

In the past, the standard model (SM) has been very successful in explaining and predicting a wide range of phenomena, but currently there are clear indications that not everything is described by the SM. Experimental measurements of certain high energy physics parameters could show deviations from the theoretical predictions, which would indicate the existence of physics beyond the standard model (BSM). Depending on where these deviations are found, it also gives some insight into which BSM physics theories are interesting to further investigate. Measurements in the Higgs sector are especially interesting, as there are many opportunities to detect BSM effects here. The most precise measurements in the Higgs sector are obtained by combining measurements of cross sections of different Higgs boson production processes and decay channels. This is typically done in the Simplified Template Cross Sections (STXS) framework: measurements of cross sections of mutually exclusive regions of phase space, defined per production process, are combined. The measurements can be interpreted in the κ -framework as a test of the SM coupling values of the Higgs boson to other SM particles. What will be presented are results from the most recent combination from the autumn of 2021, which is based on analyses of 13 TeV data with an integrated luminosity of up to 139 fb^{-1} .

T 92.5 Thu 17:15 T-H23

 Z_2 Non-Restoration and Composite Higgs: Singlet-Assisted Baryogenesis w/o Topological Defects — Andrei Angelescu, Florian Goertz, and •Aika Tada — Max-Planck-Institut für Kernphysik, Heidelberg

Simple scalar extensions of the Standard Model with a spontaneously broken Z_2 symmetry allow for a strongly first order electroweak phase transition, as sought in order to realize electroweak baryogenesis. To circumvent the emergence of phenomenologically problematic domain walls often encountered in this context, in 2112.12087 (A. Angelescu, F. Goertz, AT), a scalar singlet framework featuring a thermal history which does not restore Z_2 in the early universe is proposed. This can be realized in a low energy effective theory with D>4 operators. A possible UV completion is provided by a SO(6)/SO(5) Composite Higgs model with fermions in a symmetric 20' of SO(6), where the potential and the Yukawa terms are obtained by spurion analysis and a CP violating term arises.

Matching the two models and exploring them numerically shows that this scenario can fulfil all Sakharov criteria needed for electroweak baryogenesis.

T 92.6 Thu 17:30 T-H23

The question why an excess of matter over antimatter was produced shortly after the Big Bang is one of the greatest unsolved problems of modern physics. The Standard Model of particle physics cannot explain the amount of \mathscr{CP} violation needed for the observed baryon asymmetry of the universe. Additional \mathscr{CP} violation may be found in the Higgs sector where a mixed \mathscr{CP} state of the 125 GeV Higgs boson is not ruled out experimentally by current search limits.

In this talk, the Higgs Characterization model is presented. It is parameterized by factors c_i and \tilde{c}_i which modify the scalar and pseudoscalar part of Higgs couplings to other SM particles, respectively. The program HiggsSignals (HS) is used to calculate the resulting Higgs signal rates and compare them to available Run 1 and Run 2 data from the Large Hadron Collider (LHC) at CERN.

The signal rate measurements in HS are complemented by a dedicated \mathscr{CP} analysis in the $H \rightarrow \tau \tau$ decay from CMS. Furthermore, constraints from the

current best limit on the electron electric dipole moment are examined to complement the constraints from the LHC. The amount of baryogenesis reachable within the allowed parameter space regions is computed in the vev insertion approximation with optimal parameters.

T 92.7 Thu 17:45 T-H23

Studies of di-Higgs production at the FCC-hh in the bbZZ(llvv) final state — •KEVIN LAUDAMUS — DESY, Hamburg, Germany

The FCC-hh is a proposed circular hadron collider at an energy of 100 TeV. The total integrated luminosity is expected to be around $30 \ ab^{-1}$. With such a large dataset, 400 times more double-Higgs events are expected than with the full HL-LHC dataset, allowing to measure the Higgs self-coupling with high precision. As a consequence, also rarer final states, which are not within reach of the (HL)-LHC, have good prospects at the FCC-hh. One such final state is the bbZZ(llvv) channel, which is studied in this work. A multivariate analysis of bbllvv events is implemented and upper limits on the di-Higgs production cross-section are derived in order to assess the potential of this channel. Moreover, it is investigated in how far specific kinematic regions, such as at high Higgs transverse momentum, can be exploited. In these studies, particular attention needs to be extremely challenging at the FCC-hh due to the very high pile-up environment.

T 93: Search for New Particles 6

Time: Thursday 16:15-18:45

T 93.1 Thu 16:15 T-H24

Symmetries, Safety, and Self-Supervision — BARRY M. DILLON¹, GREGOR KASIECZKA², HANS OLISCHLÄGER¹, TILMAN PLEHN¹, PETER SORRENSON^{1,3}, and •LORENZ VOGEL¹ — ¹Institut für Theoretische Physik, Universität Heidelberg, Germany — ²Institut für Experimentalphysik, Universität Hamburg, Germany — ³Heidelberg Collaboratory for Image Processing, Universität Heidelberg, Germany

Collider searches face the challenge of defining a representation of highdimensional data such that physical symmetries are manifest, the discriminating features are retained, and the choice of representation is data-driven and newphysics agnostic. We introduce JetCLR (Contrastive Learning of Jet Representations) to solve the mapping from low-level jet constituent data to optimized observables through self-supervised contrastive learning. Using a permutationinvariant transformer-encoder network, physical symmetries such as rotations and translations are encoded as augmentations in a contrastive learning framework. As an example, we construct a data representation for top and QCD jets and visualize its symmetry properties. We benchmark the JetCLR representation against other widely-used jet representations, such as jet images and energy flow polynomials.

T 93.2 Thu 16:30 T-H24

Searching for Jet Pairs with Anomalous Substructure in CMS — GREGOR KASIECZKA, LOUIS MOUREAUX, TOBIAS QUADFASEL, and •MANUEL SOMMER-HALDER — Institut für Experimentalphysik, Universität Hamburg

Despite compelling experimental and theoretical motivation as well as extensive new physics searches at the Large Hadron Collider, there have been no discoveries of physics beyond the standard model (BSM) so far. One potential reason for this is that the common search strategy relies on selecting BSM signal candidate events based on specific signal and background models. Such a dedicated search cannot be performed for every possible BSM theory and phase space region. Thus, model-independent anomaly detection methods are an important addition to existing search methods. These algorithms aim to select signal candidates in a data-driven manner based on anomalous phase space signatures.

One such anomaly detection method is CATHODE. It detects resonant signal peaks by combining neural density estimation in a sideband region with a weakly supervised classification task of disinguishing real data from synthetic background-like samples. We present the first application of CATHODE in a search for BSM physics in the CMS experiment targeting a dijet final state.

T 93.3 Thu 16:45 T-H24

Autoencoders and k-Means for unsupervised anomaly detection in high energy physics — THORBEN FINKE, MICHAEL KRÄMER, ALESSANDRO MORANDINI, ALEXANDER MÜCK, and •IVAN OLEKSIYUK — Institute for Theoretical Particle Physics and Cosmology (TTK), RWTH Aachen University, D-52056 Aachen, Germany

Unsupervised anomalous jet tagging based on low-level observables has recently gained popularity in the high energy physics community. The main goal here is to be as efficient and model-independent as possible. We scrutinize a widely used anomaly detection method based on the reconstruction loss of a deep autoencoder to show its capabilities, but also its limitations. Although we reproduce the positive results from the literature, we show that the standard autoencoder setup cannot be considered as a model-independent anomaly tagger by inverting the task: the autoencoder fails to tag QCD jets if it is trained on top jets. We improve the capability of the autoencoder to learn non-trivial features of jet images, such that it is able to achieve both top jet tagging and QCD jet tagging with the same setup. We propose an alternative machine learning approach using k-Means and Gaussian Mixture Model to construct anomaly scores. We show that these methods, albeit simple, have several benefits and may also be regarded as promising anomaly detection tools.

T 93.4 Thu 17:00 T-H24

Location: T-H24

How to Identify Anomalies at the LHC — •THORSTEN BUSS¹, BARRY DILLON¹, THORBEN FINKE², MICHAEL KRÄMER², ALESSANDRO MORANDINI², ALEXANDER MÜCK², IVAN OLEKSIUK², and TILMAN PLEHN¹ — ¹Institut für Theoretische Physik, Universität Heidelberg, Germany — ²Institute for Theoretical Particle Physics and Cosmology (TTK), RWTH Aachen, Germany

Searches for anomalous events are the main motivation for the LHC and define key analysis steps, including triggers. We discuss how LHC anomalies are defined through probability density estimates evaluated through an appropriate latent space. Different approaches, like invertible networks, and Dirichlet latent spaces are illustrated for the especially challenging scenario of dark matter jets. Finally, we present benchmark results for unsupervised top vs QCD jet tagging.

T 93.5 Thu 17:15 T-H24

Better latent spaces for better autoencoders — BARRY DILLON¹, TILMAN PLEHN¹, CHRISTOF SAUER², and •PETER SORRENSON^{1,3} — ¹Institut für Theoretische Physik, Universität Heidelberg, Germany — ²Physikalisches Institut, Universität Heidelberg, Germany — ³Heidelberg Collaboratory for Image Processing, Universität Heidelberg, Germany

Autoencoders as tools behind anomaly searches at the LHC have the structural problem that they only work in one direction, extracting jets with higher complexity but not the other way around. To address this, we derive classifiers from the latent space of (variational) autoencoders, specifically in Gaussian mixture and Dirichlet latent spaces. In particular, the Dirichlet setup solves the problem and improves both the performance and the interpretability of the networks.

T 93.6 Thu 17:30 T-H24

Development of a new trigger for exotic particle searches with IceCube — •TIMO STÜRWALD for the IceCube-Collaboration — Bergische Universität, Wuppertal, Deutschland

The IceCube Neutrino Observatory is a cubic kilometer scale Cherenkov light detector that also searches for signatures of particles beyond the standard model. The upcoming IceCube Upgrade and IceCube-Gen2 extension will improve the sensitivity for these searches due to an increased and partly denser instrumented sensitive volume. The better sensitivity allows for the detection of signatures of exotic particles including fractionally charged particles, which directly and indirectly produce light.

In this talk results of adjusted IceCube standard triggers applied on simulated fractionally charged particles are presented. Furthermore, the development of a new trigger is presented. This new trigger includes the analysis of isolated single hits that so far are not included in any IceCube trigger, because a large fraction of them originates from well understood noise sources. For simulated faint exotic

signatures these isolated single hits become the dominant hit type. *Funded by BMBF-Verbundforschung Astroteilchenphysik

T 93.7 Thu 17:45 T-H24

Simulation of the WOM in IceCube Gen2 for the detection of exotic particles — •NICK JANNIS SCHMEISSER¹, ANNA POLLMANN¹, and JOHN RACK-HELLEIS² for the IceCube-Collaboration — ¹Bergische Universität Wuppertal — ²JGU Mainz

The IceCube Neutrino Observatory is a cubic kilometer scale Cherenkov light detector located at the geographic South Pole. Besides the detection of neutrinos, it is used for searches for particles beyond the Standard Model. One kind of these exotic particles are fractionally charged particles which carry a fraction of the electron charge. The IceCube detector will be enhanced first by the IceCube Upgrade and then by IceCube Gen2 in the next years. The Wavelength-shifting Optical Module (WOM) is a newly developed sensor that is going to be deployed in the IceCube Upgrade amongst other sensors. The WOM achieves an improved signal to noise ratio and UV-sensitivity in comparison to the sensors deployed in the original IceCube detector via wavelength shifting and light guiding. In this presentation the results of simulations with the WOM in IceCube Gen2 in vestigating the efficiency of the WOM in detecting fractionally charged particles in comparison to the optical modules used in the original IceCube detector is shown. These simulations show that the WOM has a higher efficiency than the original sensors.

T 93.8 Thu 18:00 T-H24

Search for heavy neutral lepton production and decay with the IceCube Neutrino Observatory — •LEANDER FISCHER for the IceCube-Collaboration — DESY Zeuthen

We investigate the ability of IceCube DeepCore to reconstruct low-energy (GeV) double-cascade topologies, which can be produced through Beyond Standard Model interactions. In particular, we consider Heavy Neutral Leptons (HNLs) in the mass range of 0.1-3 GeV that are produced from up-scattering of atmospheric tau-neutrinos. The sensitivity to HNL interactions, where the production and subsequent decay happen inside the detector volume, is investigated using 8 years of atmospheric data from IceCube DeepCore.

T 93.9 Thu 18:15 T-H24 Search for Sub-Relativistic Magnetic Monopoles in IceCube — •CHRISTIAN DAPPEN¹, JAKOB BÖTTCHER¹, SUKEERTHI DHARANI², and CHRISTOPHER WIEBUSCH¹ for the IceCube-Collaboration — ¹RWTH Aachen University -Physics Institute III B, Aachen, Germany — ²Universität Hamburg The IceCube Neutrino Observatory detects high energy neutrinos through their interaction in the Antarctic ice while also searching for more exotic particles such as magnetic monopoles. These hypothetical particles are predicted by Grand Unified Theories as relics from the very early Universe. For masses on the GUTscale ($10^{14}\,\text{GeV}$ - $10^{17}\,\text{GeV}$) those monopoles would move at sub-relativistic speeds ($\beta < 10^{-2}$) through IceCube. A subrelativistic monopole in matter may catalyze nucleon decays via the Rubakov-Callan effect. This results in Cherenkov light from small particle showers along the trajectory of the monopole with separations of centimeters up to tens of meters. This pattern is recorded by a dedicated slow particle trigger at a rate of ≈ 10 Hz. For the separation of signal from background events, we have developed a chain of boosted decision trees (BDTs) which are trained with simulated monopole signal and data-driven background events. In each level of the BDT, a background rejection of about 99% is achieved which allows a more efficient training of the subsequent BDT for rare backgrounds. Based on the final selection, the sensitivity is estimated and the analysis is evaluated with an experimental dataset of five months.

T 93.10 Thu 18:30 T-H24

Searching for axion-like particles via ultra-high-energy photons — •CHIARA PAPIOR, MARKUS RISSE, and PHILIP RUEHL — Center for Particle Physics Siegen, Experimentelle Astroteilchenphysik, Universität Siegen

As axion-like particles (ALPs) provide solutions for several questions that current models leave unanswered, it is of great interest to study them experimentally. A potential coupling of ALPs to photons can be probed by large scale cosmic ray detectors like the Pierre Auger Observatory which are sensitive to ultra-highenergy (UHE) photons. After the hypothetical production of high energy ALPs in the environment of e.g. a flaring blazar or a binary neutron star merger, ALPs may propagate over cosmological distances through the intergalactic medium without attenuation. Back-conversion into UHE photons may happen via the Primakoff effect in the magnetic field of the local cluster or the Galaxy itself. A high-confidence detection of an UHE photon from a transient source well beyond the megaparsec scale would be a strong indicator for the presence of ALPs due to the opaqueness of the cosmic microwave background towards photons at these energies. The probability of the ALP-photon conversion depends on several parameters such as the magnetic field in the intergalactic medium, the distance of propagation and the mass and coupling of the ALPs themselves. The work presented in this contribution aims to evaluate the possibility to probe the phase space of UHE ALPs by using the functionality of the gammaALPs python package.

T 94: Silicon Strip Detectors 2

Time: Thursday 16:15-18:00

T 94.1 Thu 16:15 T-H25

Status of the CMOS Strip Detector Project in Freiburg — •Niels Sorgenfrei, Leena Diehl, Marc Hauser, Cedric Hönig, Karl Jakobs, Sven Mägdefessel, Ulrich Parzefall, Arturo Rodriguez, and Dennis Sperlich — Albert-Ludwigs Universität, Freiburg, Germany

Due to the increased use of large area silicon detectors in current and future particle detectors and only very few vendors available, which are capable of silicon production and processing fulfilling quality and size requirements, the need for reliable, fast and cost efficient production processes arises. As part of a CERN market survey, CMOS sensors in pixel and strip geometries were developed.

The idea is to utilise the already existing industry infrastructure of the CMOS process. However, typical CMOS foundries are usually only equipped to process smaller sensors compared to what is required in e.g. the strip region of the ATLAS Inner Tracker. Therefore, the process of stitching has to be used. By employing wafer masks where the sensor structure is divided up into different parts, the individual parts can be imprinted multiple times side by side on the wafer resulting in coherent areas larger than the reticles themselves. The effect of stitching on charge collection, electric field strength and configuration, detection efficiency and radiation hardness has to be investigated.

In this talk measurements on passive CMOS strip sensors, produced by LFoundry in a 150 nm process, will be discussed. Three different strip designs are investigated and the results of IV, 90 Sr-source and edge-TCT measurements will be presented.

T 94.2 Thu 16:30 T-H25 **Humidity Studies on ATLAS ITk sensors** — NAIM BORA ATLAY², INGO BLOCH¹, HEIKO LACKER², •ILONA STEFANA NINCA¹, and CHRISTIAN SCHARF² — ¹Deutsches Elektronen-Synchrotron (DESY), Zeuthen, Germany — ²Humboldt-Universität zu Berlin, Berlin, Germany

This study aims to investigate electrical breakdown in the periphery of silicon sensors, with a focus on humid conditions. The study will take advantage of the large number of silicon strip sensors produced for the ATLAS Upgrade in order

to probe their behavior. The sensors are investigated using the transient current technique (TCT). The TCT setup available at Deutsches Elektronen-Synchrotron in Zeuthen employs both red and infrared laser light. This setup can be used to study the transport of free charge carriers in the sensors at the surface and inside the bulk. Additionally, to better understand the electric properties of the sensors, simulations using Technology Computer Aided Design were produced. ATLAS ITk silicon miniature diodes were investigated in a humid environment. We are planning to use a dedicated camera to localize the critical spots where breakdown occurs on the surface of the diodes through luminescence of hot carriers. Using the information provided by the TCT measurements, the humidity dependence of the electric field at the breakdown point is planned to be studied. Understanding these dependencies will help us propose better geometries for future sensors and potentially allow us to develop improved operation scenarios for the future ATLAS ITk strips detector.

T 94.3 Thu 16:45 T-H25

Location: T-H25

Recent results from the End-of-Substructure card for the ATLAS Strip Tracker Upgrade — •Rickard Stroem¹, Artur Boebel¹, Harald Ceslik¹, Mogens Dam², Sergio Diez Cornell¹, Peter Goettlicher¹, Ingrid Gregor¹, James Keaveney³, Max Nikoi Van Der Merwe³, Jan Oechsle², Stefan Schmitt¹, Marcel Stanitzki¹, and Jane Wyngaard³ — ¹DESY, Germany — ²Niels Bohr Institute, Denmark — ³University of Cape Town, South Africa

The silicon tracker of the ATLAS experiment will be upgraded for the High-Luminosity Upgrade of the LHC. The main building blocks of the new strip tracker are modules of silicon sensors and hybrid PCBs hosting the read-out ASICs. The modules are mounted on rigid carbon-fibre substructures, known as staves in the central barrel region and petals in the end-cap regions, that provide services to all the modules. At the end of each stave/petal, a so-called End-of-Substructure (EoS) card facilitates the transfer of data, power, and control signal between the modules and the off-detector systems. The module front-end ASICs transfer data to the EoS card on 640 Mbit/s differential lines. The EoS connects T 94.4 Thu 17:00 T-H25

Beam Test Studies with Silicon Sensor Modules for the CMS Experiment — •Florian Wittig¹, Tobias Barvich¹, Bernd Berger¹, Alexander DIERLAMM¹, ALEXANDER DROLL¹, UMUT ELICABUK¹, ULRICH HUSEMANN¹, Markus Klute¹, Gani Kösker¹, Roland Koppenhöfer¹, Stefan Maier¹, Thomas Müller¹, Jan-Ole Müller-Gosewisch¹, Andreas Nürnberg², Marius Neufeld¹, Hans Jürgen Simonis¹, Pia Steck¹, and Lea Stock¹ — ¹Institute of Experimental Particle Physics (ETP), Karlsruher Institute of Technology (KIT) — ²Deutsches Elektronen-Synchrotron (DESY)

In the context of the Phase-2 Upgrade of the CMS Experiment, the whole tracker will be replaced. The new CMS Outer Tracker will be equipped with two different types of silicon sensor modules called PS and 2S modules. During the LHC runtime, the modules and especially the silicon sensors will accumulate radiation damage, which lowers the signal measured by the module. In order to ensure the full functionality of the 2S modules during the entire operation time of CMS, modules in the advanced prototyping phase have been built. One of the prototypes features sensors that have been irradiated to fluences that are expected at the end of the CMS runtime. These unirradiated and irradiated prototypes have been tested at the DESY beam test facility. This talk summarizes the results gathered at the beam tests.

T 94.5 Thu 17:15 T-H25 test beam analysis of a silicon-strip module for the CMS phase-II tracker upgrade — •CHUN CHENG — DESY, Hamburg, Germany

The foreseen Large Hadron Collider upgrade is expected to deliver an integrated luminosity that is one order of magnitude larger after 2027. Rare processes and new phenomena may be observed in this high luminosity era. The Phase-II Outer Tracker upgrade of the CMS experiment is required to surmount higher radiation and increased event rate. Transverse momentum (PT) discrimination is introduced in the design and will contribute to the Level-1 Trigger. A CMS 2S silicon strip module with PT discrimination concept was built by the DESY Outer Tracker group and has undergone a test beam experiment at the DESY test beam facility.

The talk will briefly summarize the assembly of the DESY 2S module, sensor studies and the data acquisition scheme during the beam test. The main focus will be on the results from recent test beam measurements. The analysis will be done based on corryvreckan framework, a modular concept of test beam recon-

T 95.1 Thu 16:15 T-H26

T 95: Pixel Detectors 3

Effects of gamma radiation on DEPFET pixel sensors for the Belle II exper**iment** – •Georgios Giakoustidis¹, Jochen Dingfelder¹, Ariane Frey², Botho Paschen¹, Benjamin Schwenker², and Marike Schwickardi² for the Belle II-Collaboration — ¹University of Bonn, Germany — ²University of Göttingen, Germany

Time: Thursday 16:15-18:15

For the Belle II experiment at KEK (Tsukuba, Japan) the KEKB accelerator was upgraded to deliver e^+e^- collisions at a center of mass energy of E_{CM} = 10.58 GeV with an instantaneous luminosity of $8 \cdot 10^{35}$ cm⁻²s⁻¹. As the innermost part of the Belle II detector, the PiXel Detector (PXD), based on DEpleted P-channel Field Effect Transistor (DEPFET) technology, is most exposed to radiation from the accelerator. Prototypes as well as a module from the final Belle II production batch were irradiated with X-rays to doses up to 20 Mrad, corresponding to the expected lifetime exposure. The performance of the DEPFET sensors and front-end electronics will be presented and the results of two recent campaigns will be compared to previous results.

T 95.2 Thu 16:30 T-H26

CMOS Upgrade of the Belle II Vertex Detector - •MARCO VOGT, CHRIStian Bespin, Ivan Dario Caicedo Sierra, Jochen Dingfelder, Tomasz HEMPEREK, FABIAN HÜGGING, HANS KRÜGER, and NORBERT WERMES -Physikalisches Institut der Universität Bonn

The Belle II experiment at KEK in Japan will have an opportunity for upgrades of its detector components during a long shutdown in 2027, when several upgrades of the SuperKEKB $e^+ e^-$ collider are planned. At the expected high instantaneous luminosity of $6x10^{35}$ cm⁻², SuperKEKB will generate a high rate of background particles, especially in the inner detector layers, the vertex detector. Here, hit rates will exceed 100 $MHz \ cm^{-2}$, inducing radiation levels of 100 Mrad TID and fluences reaching $5x10^{14}$ neq cm⁻². Such a high level of beam background will

struction chain. Hit efficiency of the sensors under a bias scan, the performance over the large module area, and on-module PT discrimination functionality will be presented.

T 94.6 Thu 17:30 T-H25

Beam test of 2S module prototypes for the Phase-2 CMS Outer Tracker -Christian Dziwok², Lutz Feld¹, Katja Klein¹, Martin Lipinski¹, Alexan-DER PAULS¹, OLIVER POOTH², NICOLAS RÖWERT¹, and •TIM ZIEMONS² -¹I. Physikalisches Institut B, RWTH Aachen University, Germany — ²RWTH Aachen University - Physics Institute III B, Aachen, Germany

The CMS detector will be upgraded in the Phase-2 Upgrade for the operation at the HL-LHC. Among others, the silicon tracking system will be completely replaced by a new system providing an extended acceptance, an improved granularity and the feature to include tracking information into the level-1 trigger. The new Outer Tracker will consist of 2S modules with two strip sensors and PS modules with a macro-pixel sensor and a strip sensor, specialized detector modules with on board $p_{\rm T}$ discrimination.

The functionality of current generation prototype 2S modules has been tested at the test beam facility at DESY Hamburg in November 2019. With a 4 GeV electron beam, various studies are performed like efficiency scans at different positions of the module or at varying inclination angles to mimic different $p_{\rm T}$ particles. In this talk, efficiency studies are presented.

T 94.7 Thu 17:45 T-H25

A new high rate electron beam line at DESY II — • DOHUN KIM — DESY, Notkestraße 85, 22607 Hamburg

The R-Weg is a former transfer beam line from the DESY II synchrotron to DORIS. Recently, it has been refurbished to serve as a high-rate electron beam line. The full DESY II beam with up to several $10^{10}e^{-1}$ can be dumped at a rate of 12.5 Hz. The available rates allow many detector tests that require high particle rates, but this also allows to use the beam line has a facility for electron irradiation.

Before the R-Weg is put into full operation, it is necessary to understand the beam parameters and the radiation field in detail. Therefore, the R-Weg has been simulated and studied using FLUKA, which is a MC simulation framework for the interaction and transport of particles in materials.

The beam divergence, stability, beam profile etc. have been simulated. To verify the results, a suite of measurements has been prepared and compared. In addition, the neutron and gamma background from the beam dump are studied to ensure safe operation and to enable the use as a electron irradiation facility. This presentation is going to explain details of the R-Weg and present the simu-

lation result. Finally, an outlook into future measurements at the R-Weg is given.

Thursday

create new challenges and requirements that can be met by a new vertex detector.

A performant and robust vertex detector upgrade (VTX) is currently being defined by the Belle II collaboration. The baseline design foresees an all-layer pixel detector system based on the TJ-Monopix2 fully depleted monolithic CMOS sensor. For the inner layers of the VTX, a new ultra-thin all-silicon ladder concept is being developed and tested.

In this talk, the proposed monolithic CMOS upgrade will be presented. Technological aspects, system integration and performance estimations will be discussed.

T 95.3 Thu 16:45 T-H26

Location: T-H26

Irradiation burst study of Belle II PXD module components - FLORIAN Bernlochner¹, Jochen Dingfelder¹, Georgios Giakoustidis¹, Matthias HOEK², BOTHO PASCHEN¹, and •JANNES SCHMITZ¹ for the Belle II-Collaboration ¹University of Bonn, Germany — ²University of Mainz, Germany

The Belle II detector started recording collision data in spring 2019. During physics runs, unexpected irradiation burst events occurred, which exposed the inner detectors and especially the PXD (PiXel Detector) to unwanted levels of prompt irradiation. Dedicated measurement campaigns were carried out at the Mainz Microtron (MAMI), which aimed to reproduce the observed effects of irradiation bursts on the PXD in Belle II. To this end, a focused high intensity (up to 10uA) pencil beam of 855 MeV electrons was used to irradiate full system demonstrators in several spatially confined fiducial regions. During first campaigns in 2020 the observed failure mode could be reproduced and restricted to vulnerable regions in one specific module component. In this talk, the results of the latest measurement campaign in December 2021 will be presented, focussing on possible protective measures against the impact of irradiation bursts on the PXD modules installed inside Belle II.

T 95.4 Thu 17:00 T-H26

KEK Total Ionizing Dose Measurement of PXD Modules and sensor effects at Belle II — Ariane Frey, Benjamin Schwenker, Yannik Buch, and •Marike Schwickardi — II. Physikalisches Institut, Georg-August-Universität Göttingen

The Belle II experiment at the Japanese B-factory SuperKEKB has started data taking in early 2019 and the peak luminosity will be ramped up to $8 \cdot 10^{35} \text{ cm}^{-2} \text{ s}^{-1}$, which is 40 times higher than the previous luminosity delivered to the Belle experiment. It was therefore equipped with a new DEpleted P-channel Field Effect Transistor (DEPFET) based silicon PiXel Detector (PXD) for vertex detection to cope with higher beam backgrounds.

The monitoring of radiation effects on the new PXD is important throughout operation, since sensor settings have to be adjusted to ensure efficient operation. To compare observed effects to preliminary studies an accurate measurement of the total ionizing dose (TID) is needed. One way of dose monitoring is the diamond control units, that measure the radiation conditions for the Belle II detector. The diamonds are placed at different angles and positions close to the beam pipe and continuously take data. However, the diamonds do not measure the dose that the PXD received during beam times correctly, therefore the approach was chosen to estimate the TID of the PXD, using its occupancy. The drawback of this is that the PXD only sends out data while it is powered. To compensate that, the dose rate deposited when PXD DAQ is disconnect is estimated by scaling the dose rate of nearby diamonds with the ratio of pxd dose to diamond dose.

T 95.5 Thu 17:15 T-H26

Analog to Digital Converter ASICs scan on irradiated modules of the Pixel Vertex Detector at the Belle II experiment — •TOMMY MARTINOV and ARTHUR BOLZ — DESY, Hamburg, Germany

The Belle II detector placed at the SuperKEKB collider in Japan aims at studying heavy flavour physics through electron-positron collisions at a center of mass energy of approximately 10.6 GeV. The innermost element of the detector is the Pixel Vertex Detector (PXD). The pixels are based on the DEPFET technology (depleted p-channel field effect transistor). The PXD is composed of 10 ladders representing a total of approximately 3.6 * 10^6 pixels. The PXD detects charged particles with a transverse momentum higher than 40 MeV. On each module, four Drain Current Digitizers (DCDs) measure the drain currents of all DEPFET pixels. In each DCD, 256 ADCs convert the analog currents to digital signals. The characteristics of the ADCs are influenced by several components and reference voltages in the DCD. An ADC scan needs to be performed in order to calibrate the ADCs and optimize the number of channels considered as "good". The Belle II detector has been taking data for three years and it has become vital to evaluate the ageing of the different components of the detector. This presentation focuses on the ADC calibration as a way to assess the operation quality of the PXD DCDs in 2022 after three years of data taking.

T 95.6 Thu 17:30 T-H26 Challenges of offset calibration in irradiated modules of the Pixel Vertex Detector at the Belle II experiment — •Maria Konstantinova and Arthur Bolz — DESY Hamburg

The Pixel Vertex Detector (PXD) is situated as the innermost subdetector of the Belle II experiment at the SuperKEKB collider in Japan. Each PXD module in-

Thursday

cludes a matrix of 192'000 pixels based on the Depleted P-channel Field-Effect Transistor (DEPFET) technology. The matrices are read out in rolling shutter mode such that 1000 channels are digitized in parallel by four custom Drain Current Digitizer (DCD) ASICs. For a consistent response to transversing charged particles during operation, homogeneous pedestal currents must be subtracted for each pixel to obtain as much room as possible for analogue-to-digital conversion of the signal current. Therefore, a narrow pedestal spread is achieved by adding a 2-bit DAC offset current to every pixel. The offset currents show non-linear behaviour which is dependent on the hardware architecture and may change during module operation in a harsh radiation environment in the interaction region. In this talk those effects are analyzed and the challenges they pose to a good offset calibration of PXD are discussed.

T 95.7 Thu 17:45 T-H26

Characterizatoin of DEPFET Pixel Modules for PXD2 — •LARISSA VON JASIENICKI, JANNES SCHMITZ, PATRICK AHLBURG, GEORGIOS GIAKOUSTIDIS, BOTHO PASCHEN, FLORIAN BERNLOCHNER, and JOCHEN DINGFELDER for the Belle II-Collaboration — University of Bonn, Germany

The SuperKEKB electron-positron collider in Japan has reached unprecedented luminosities. The Belle II experiment operating at the SuperKEKB collider is equipped with a PiXel Detector (PXD) based on the Depleted P-channel Field Effect Transistor (DEPFET) technology, which serves as the innermost two layers of the VerteX Detector (VXD) of Belle II and was designed to cope with the large particle rates. The current PXD is, however, incomplete, since (mostly) only the innermost of the two layers is installed. A new full-scale PXD (*PXD2*) is currently being built and is expected to be installed in 2023(?).

This talk highlights improvements in the lab characterization of individual PXD2 modules. In particular, the biasing requirements of the sensors and investigations of the intrinsic properties like transconductance and gain of the DEPFET sensor will be discussed.

T 95.8 Thu 18:00 T-H26

Assembly and tests of the first TRISTAN detector modules — •DANIEL SIEG-MANN for the KATRIN-Collaboration — Max-Planck Institute for Physics, Munich, Germany

The TRISTAN (TRitium Invetigations of STerile to Active Neutrino mixing) project aims to search for the signature of a keV sterile neutrino in the tritium beta decay spectra by upgrading the detector system of the KATRIN experiment. This extension of the experiment will be performed after its neutrino mass survey.

To reach a high sensitivity to the sterile neutrino mixing angle the strong activity of the KATRIN tritium source is required. The resulting high electron rate is one of the greatest challenges for the TRISTAN project. It will be approached by distributing the rate among 3500 pixels, resulting in count rates of 100 kcps per pixel. To resolve the kink-like signature of the keV sterile neutrino signal the detector needs to maintain an excellent energy resolution of 300 eV (FWHM) at 20 keV and a low energy threshold.

The TRISTAN detector is segmented into 21 identical modules, each hosting 166 independent pixel. The development and tests of the first detector modules will be presented in this talk.

This work is supported by the Max Planck society and the TU Munich ("Chair for Dark Matter, Susanne Mertens").

T 96: Detector Systems 3

Time: Thursday 16:15-18:15

T 96.1 Thu 16:15 T-H27

Status of the Mu3e tile detector — •HANNAH KLINGENMEYER for the Mu3e-Collaboration — Kirchhoff-Institut für Physik, Universität Heidelberg

The aim of the Mu3e experiment is the search of the lepton-flavour violating decay $\mu \rightarrow eee$ down to a sensitivity of 10^{-16} . It will be operated at the Paul Scherrer Institute (PSI) in Switzerland and consists of dedicated tracking and timing detectors, which will provide precise spatial and time measurements in order to suppress any background mimicking the signal decay.

One of the timing systems is the Mu3e tile detector, which allows precise timing of individual electrons with a resolution below 100 ps. It uses plastic scintillator tiles and silicon photomultipliers that are read out by a custom-designed ASIC, and is currently in the pre-production phase. This talk will give an overview of the current status and development of the tile detector, as well as of the performance of individual detector components. The tile matrix production and quality assurance measurements will be discussed along with advancements in the readout electronics, the cooling system, and the integration into the full experiment. Furthermore, an outlook on the full detector production will be given.

Location: T-H27

T 96.2 Thu 16:30 T-H27

Simulation studies on the Mu3e tile detector - Time alignment & clustering — •ERIK STEINKAMP and MAXIMILIAN KÖPER for the Mu3e-Collaboration — Kirchhoff Institute for Physics, Heidelberg University

The Mu3e experiment aims to detect the charged lepton flavor violating decay $\mu \rightarrow eee$ with a target sensitivity of 10^{-16} , improving the existing limit by four orders of magnitude. A successful observation would be a strong indicator for physics beyond the Standard Model. Precise timing information is needed to correctly identify the vertices of the three decay electrons and to suppress background from internal conversion decays and combinatorics. The tile detector, which utilizes scintillator tiles and SiPMs, aims to provide this precise time measurement of better than 100ps. To achieve this precision on the detector system level, a time calibration scheme using different event topologies to determine time offsets for every tile, is required. The obtained single-tile timestamps are clustered and matched to tracks from the tracking detectors. We present simulation studies of the time calibration routine, as well as clustering and track-tile matching.

T 96.3 Thu 16:45 T-H27

Irradiation studies for the Mu3e tile detector — •TIANCHENG ZHONG for the Mu3e-Collaboration — Kirchhoff-Institut für Physik, Universität Heidelberg, Heidelberg, Germany

The Mu3e experiment is designed to search for the decay $\mu^+ \rightarrow e^+ e^+ e^-$ with a sensitivity of 10^{-16} , which would be a clear signal for new physics beyond the Standard Model. To reduce the combinatorial background from muon decays while efficiently identifying 3-electron final states, a scintillating-tile detector with a required timing resolution < 100 ps and efficiency close to 100% is under development.

Irradiation damage and effects on Silicon Photomultiplier (SiPM) used in the tile detector were investigated by exposing the sensors to the decay electrons from stopped muons at the PiE5 beamline at PSI. For the SiPMs irradiated with a dose up to 1.57×10^{11} 1 MeV n_{eq}/cm^2 , corresponding to 70% of the maximum dose of the Mu3e Phase I run, the dark current increased by a factor up to 1000. We will report on the irradiation campaign performed, measurements of dark current and impact on annealing at different temperatures. The timing performance after irradiation was investigated in testbeam campaigns at DESY and will also be discussed.

T 96.4 Thu 17:00 T-H27

The detector system for the Stopping Target Monitor of the Mu2e experiment at Fermilab — •ANNA FERRARI, STEFAN E. MUELLER, OLIVER KNODEL, and REUVEN RACHAMIN for the Mu2e-Collaboration — Helmholtz-Zentrum Dresden-Rossendorf, Dresden, Germany

The Mu2e experiment, which is currently under construction at the Fermi National Accelerator Laboratory near Chicago, will search for the neutrinoless conversion of muons to electrons in the field of an aluminum nucleus, with the aim to reach a sensitivity four orders of magnitude better than previous experiments . Observations of a signal would be an example of Charged Lepton Flavor Violation, which would require physics beyond the Standard Model.

In order to normalize the result, a stopping target monitor will measure the number of stopped and captured muons in the aluminum target. The detector system includes a HPGe and a Lanthanum Bromide detector, which with different capabilities will measure x- and gamma-ray lines up to 1809 keV, in presence of high rate Bremsstrahlung and other backgrounds. At the Helmholtz-Zentrum Dresden-Rossendorf, we use the Bremsstrahlung photon beamline at the ELBE radiation facility to study the detector system performance in the pulsed conditions similar to that expected in Mu2e.

In the presentation, after an overview of the design and status of the Mu2e experiment, the main results of the ELBE campaign will be presented and discussed.

T 96.5 Thu 17:15 T-H27

Track reconstruction for the Mu3e experiment — •ALEXANDR KOZLINSKIY for the Mu3e-Collaboration — Institut für Kernphysik, Johannes Gutenberg-Universität Mainz

The *Mu3e* experiment is designed to search for the lepton flavor violating decay $\mu^+ \rightarrow e^+ e^- e^+$. The ultimate aim of the experiment is to reach a branching ratio sensitivity of 10^{-16} . The experiment is located at the Paul Scherrer Institute (Switzerland) and an existing beam line providing 10^8 muons per second will allow to reach a sensitivity of a few 10^{-15} in the first phase of the experiment. The muons with a momentum of about 28 MeV/c are stopped and decay at rest on a target. The decay products (positrons and electrons) with energies below 53 MeV are measured by a tracking detector consisting of two double layers of

50 μ m thin high-voltage monolithic active pixel sensors. The high granularity of pixel detector with a pixel size of 80×80 μ m together with the small material budget allows for a precise track reconstruction. The track reconstruction is based on 3-dimensional multiple scattering fit and uses special methods to remove incorrectly reconstructed tracks, which is made possible by high efficiency and low noise rate of the pixel detector. This talk will present the details of the track reconstruction including the methods that are used to reduce the number of fake tracks.

T 96.6 Thu 17:30 T-H27

Commissioning of the LHCb Scintillating Fibre Tracker — SEBASTIAN BACH-MANN, DANIEL BERNINGHOFF, XIAOXUE HAN, BLAKE LEVERINGTON, ULRICH UWER, and •LUKAS WITOLA — Universität Heidelberg, Heidelberg, Deutschland The LHCb detector underwent a major upgrade in the past years. The modifications enable the detector to operate at an increased instantaneous luminosity and to read out data at the LHC bunch crossing rate of 40MHz. The new operating conditions required the replacement of the complete tracking system. The main tracking stations are replaced by the SciFi Tracker, a large high granular scintillating fibre tracker readout by arrays of silicon photomultipliers (SiPMs). A custom ASIC is used to digitise the SiPM signals at 40MHz. Further digital electronics perform clustering and data-compression before the data is sent via optical links to the DAQ system.

The detector modules together with the readout electronics and all services are mounted on so-called C-Frames. The serial assembly and commissioning of frames is in its last stages before the start of the LHC in early 2022. The talk will give an overview of the detector and present experiences from the serial production and the latest commissioning results.

T 96.7 Thu 17:45 T-H27 Alignment studies of the LHCb SciFi Tracker — •NILS BREER¹ and SOPHIE HOLLITT² — ¹TU Dortmund, experimentelle Physik 5 — ²TU Dortmund, experimentelle Physik 5

As part of the LHCb upgrade, the Scintillating Fibre Tracker (SciFi) will replace the previous Outer and Inner Tracker detectors. It is crucial to understand which constraints and which parts of the SciFi have the most impact on the alignment quality.

In order to align the SciFi, several configurations of degrees of freedom and alignment constraints are studied. Further analysis is used to search for possible weak modes and confirm that the alignment configuration produces stable results. In this talk, we will present results from misalignment tests using the current best configuration.

T 96.8 Thu 18:00 T-H27

Monitoring alignment performance for LHCb's Scintillating Fibre Tracker — NILS BREER and •SOPHIE HOLLITT — Experimentelle Physik 5, TU Dortmund During the LHC upgrade period, the LHCb experiment has replaced the majority of its subdetectors and extensively upgraded its trigger system. Before the physics data collection period late in 2022, careful commissioning of the full system is required. The Scintillating Fibre Tracker (SciFi) is the first detector after the LHCb magnet, and its alignment will be crucial for the final mass and momentum resolution of the upgraded experiment.

In this talk, we discuss how misalignment effects in the SciFi detector can be monitored via tracking and physics performance, and give an overview of the alignment and calibration procedures for the commissioning period of the detector.

T 97: Electronics 1

Time: Thursday 16:15-18:00

T 97.1 Thu 16:15 T-H28

Instrumentation of a GEM-Based Neutron Detector — •LAURA RODRÍGUEZ GÓMEZ¹, SAIME GÜRBÜZ¹, JOCHEN KAMINSKI¹, MARKUS KÖHLI², MICHAEL LUPBERGER¹, DIVYA PAL¹, and KLAUS DESCH¹ — ¹Universität Bonn — ²Universität Heidelberg

The VMM chip, originally developed for the ATLAS New Small Wheel Upgrade, was implemented in the multi-purpose readout system of the RD51 collaboration over the last years. Within this so-called Scalable Readout System (SRS) the frontend board is called VMM hybrid as it holds two frontend chips as well as an FPGA and other electronic components to handle data and powering. This system provides a complete readout chain for a large variety of particle detectors.

For the developement of a multilayer Gas Electron Multiplier (GEM) - based neutron detector, a larger system of hybrids is planned, set up and tested. This includes not only the high voltage protection of the readout electronics and the power stability of all hybrids, but also the design of a cooling system and a mechanical suspension. Location: T-H28

The detector concept, VMM hybrid test results and measurements with a test layer are presented. The application of VMM hybrids and GEMs in multilayer neutron detectors as a technology transfer is discussed.

T 97.2 Thu 16:30 T-H28

Development of a DC-DC converter for powering the Mu3e detector — •SOPHIE GAGNEUR for the Mu3e-Collaboration — Institut für Kernphysik, JGU Mainz

The Mu3e experiment under construction at the Paul Scherrer Institute, Switzerland, aims to search for the lepton flavour violating decay of a muon into one electron and two positrons with an ultimate sensitivity of one in 10¹⁶ muon decays. The Mu3e detector consists of a tracker based on High-Voltage Monolithic Active Pixel Sensors (HV-MAPS) combined with scintillating tile and fibre timing detectors. The detector ASICs need a supply voltage of around 2V. This voltage is generated from the 20V external supply voltage via switching DC-DC converters. These buck converters must be able to operate within a magnetic field and provide a constant output voltage with a ripple of less than 10mV to guarantee a proper operation of the pixel sensors and timing detectors. The second version of the Mu3e DC-DC converter has been designed, produced and already tested successfully in the laboratory, implementing features such as a secondary output filter and a temperature interlock for the pixel sensors. The final version is currently being designed and integrated into the experiment to be used during the upcoming commissioning runs.

T 97.3 Thu 16:45 T-H28

Powering Scheme of the Tracking Detector of the P2 Experiment at MESA — •LARS STEFFEN WEINSTOCK — PRISMA+ Cluster of Excellence and Institute of Nuclear Physics, Johannes Gutenberg University Mainz

P2 is a precision experiment planned for the Mainz Energy recovering Superconducting Accelerator (MESA) currently under construction. The goal of P2 is to determine the electroweak mixing angle at low energy scales with unprecedented precision by measuring the parity violating asymmetry in proton-electron scattering at low momentum transfer. A key parameter for the analysis, the electron momentum transfer during scattering, is measured by the P2 tracker, which is placed inside the 0.6 T solenoid spectrometer. The tracker consists of eight identical modules utilising a total of 4320 novel High Voltage Monolithic Active Pixel Sensors (HV-MAPS) for precise track reconstruction. With each HV-MAPS drawing about 1W the tracker front-end requires more than 5kW of power, which is supplied to the tracker using a remote-sense technique. This talk presents the current state of the powering scheme of the P2 tracking detector, key design parameters, and technical details of the implementation including first test measurements.

T 97.4 Thu 17:00 T-H28

A Simulation Framework to Optimize Signal Processing for Particle Detectors — •FLORIAN RÖSSING¹, ANDRÉ ZAMBANINI¹, CHRISTIAN GREWING¹, and STEFAN VAN WAASEN^{1,2} — ¹ZEA-2, Forschungszentrum Jülich GmbH — ²NTS, University of Duisburg-Essen

Particle Detectors evolve to ever higher performance, both in terms of sensitivity and channel density. This increases the amount of data to be handled. As transmitting this raw data is often not a viable option, data reduction has to be employed. To achieve this, the individual channel signals are converted, and the data is processed close to the sensor, extracting observable parameters of the signal. Recent developments often rely on low-level, analog blocks and simple digitizers as signal converters, which are tailored to the specific sensor used in the detector. This limits reusability, making a repeated design effort necessary. The design of generic readout electronics based on digital data processing could overcome this issue. In a pursuit to build such a generic detector readout, part of the necessary work is the design of a single channel signal conversion and data handling, both to be used for a wide range of detectors with different sensors. For this, MatLab and Simulink are used to study and evaluate signal and data processing chains. This includes shaping, different digitization approaches (e.g. TDC, ADC) and data processing algorithms. This contribution will describe the models used as input signals for simulations, the architecture of the simulation software, and introduce first algorithm implementations.

T 97.5 Thu 17:15 T-H28

FPC design and prototype for the ATLAS High Granularity Timing Detector Demonstrator — •MARIA DE LA SOLEDAD ROBLES MANZANO¹, PE-TER BERNHARD², ANDREA BROGNA², ATILA KURT², KARL-HEINZ GEIB¹, LUCIA MASETTI¹, BINH PHAM², STEFFEN SCHOENFELDER², and QUIRIN WEITZEL² — ¹Institut für Physik, Johannes-Gütenberg Universität Mainz — ²PRISMA⁺ Detector Lab, Johannes-Gütenberg Universität Mainz ATLAS forward region, providing a time resolution of about 30 ps per track. The active area consists of 2-double-sided disks per end-cap. The HGTD basic unit, so-called module, is made up of two 2x2 cm² Low Gain Avalanche Detectors bump-bonded to two ASICs and glued to a flexible PCB. The modules are connected to the Peripheral Electronics Boards, surrounding the active area, via a Flexible Printed Circuit (Flex tail) that serves as interconnection for power, communication signals and HV bias. As part of the HGTD R&D phase, a demonstrator is proposed to test the functionality and assembly of a subset of components of the full detector. The design and tests of a Flex tail prototype in the context of an overall description of the demonstrator activities are presented.

The ATLAS detector requires upgrades to face the challenges of the new HL-

LHC, mainly the increase of pile-up interactions. The High-Granularity Timing

Detector (HGTD) will be built in order to mitigate the effects of pile-up in the

T 97.6 Thu 17:30 T-H28 Prototyping Serial Powering with RD53A - KLAUS DESCH, MATTHIAS HAMER, •FLORIAN HINTERKEUSER, FABIAN HÜGGING, HANS KRÜGER, CHAR-LOTTE PERRY, and LARS SCHALL — Physikalisches Institut, Universität Bonn The high luminosity upgrade for the Large Hadron Collider at CERN requires a complete redesign of the current inner detectors of ATLAS and CMS. These new inner detectors will consist of all-silicon tracking detectors. A serial powering scheme has been chosen as baseline for the pixel detector to cope with the higher number of modules and the higher power consumption of the new front-end chip, spatial constraints and the need to minimize the tracker's material budget. This new powering scheme provides challenges for the electrical and mechanical design. In order to verify this new powering scheme and its implications on the detector integration, efforts have been ongoing to set up a prototype for serial powering using modules based on the RD53A chip, a half-size prototype for the new Pixel front-end chip, developed by the RD53 collaboration. In particular, a serial powering stave consisting of up to 8 RD53A quad chip modules has been set up in Bonn. The results from the activities with RD53A chips are presented. Emphasis is put on the electrical characterization of an RD53A serial powering chain, using representative services and power supplies. The setup, measurement goals and characterization of the serial powering chain will be discussed.

T 97.7 Thu 17:45 T-H28

Test Results of the New ASD Chips for Phase-II Upgrade of the ATLAS MDT Chambers at HL-LHC – •KATRIN PENSKI, OTMAR BIEBEL, STEFANIE GÖTZ, VITALIY HAVRYLENKO, RALF HERTENBERGER, CHRISTOPH JAGFELD, MAXIMILIAN RINNAGEL, CHRYSOSTOMOS VALDERANIS, and FABIAN VOGEL – LMU München

The Phase-II Upgrade of the ATLAS Muon Spectrometer to the High Luminosity LHC (HL-LHC) requires an efficient trigger and readout system for the Monitored Drift Tube (MDT) chambers. For this purpose, new front-end electronics have been developed including an 8-channel amplifier shaper discriminator (ASD) chip built in 130 nm GF CMOS technology. Using pre-production chips, this presentation discusses the overall performance of these chips as well as the dependence on programmable parameters. Moreover, the uniformity between preproduction chips and the first batch of production chips is shown. These test results are used to define the acceptance criteria for the series testing which is planned for winter 2021. For this series testing a new tester is necessary. Its behavior is studied by retesting all chips with this tester and comparing the corresponding results with those of the previous tester.

T 98: Experimental Methods (general) 4

Time: Thursday 16:15-18:05

Group ReportT 98.1Thu 16:15T-H29High-D: F&E für hochsegmentierte mehrdimensionale Detektoren für zu-
künftige Experimente — •ERIKA GARUTTI für die High-D-Kollaboration — U
Hamburg

Zukünftige Experimente für Higgs-Präzisionsmessungen, die Suche nach Physik über das Standardmodell hinaus, sowie für die Untersuchung des Quark-Gluon-Plasmas und die Erforschung des QCD-Phasendiagramms, verlangen eine neue Generation von Hochpräzisionsdetektoren mit beispielloser räumlicher, zeitlicher und energetischer Auflösung. Die Anforderungen an solche 5-dimensionale (5D) Messungen können nur durch die Kombination von Detektoren mit extremer Granularität und neuartigen Rekonstruktionsmethoden erreicht werden. Eine höhere Segmentierung wird durch neu zu entwickelnde mikroelektronische Technologien, Halbleiterdesigns, Segmentierungskonzepte und Ausleseelektronik möglich werden. Diese Forschung auf der Detektorseite muß von neuartigen Algorithmen begleitet werden, die die bereitgestellte 5D-Information effektiv nutzt. Sie geht darin weit über einen einzelnen Detektor hinaus, indem sich alle Komponenten von einem Detektorsystem ergänzen, um eine optimale Rekonstruktionspräzision zu gewährleisten. High-D ist ein neuer vom BMBF geförderter Verbund, in dem die Gemeinschaften der Elementarteilchen-, Kernund Hadronenphysik erstmalig miteinander gemeinsam an der Entwicklung verschiedener grundlegender Technologien zu solchen 5D-Detektoren zusammenarbeiten. Der Vortrag gibt einen Überblick über die geplanten Arbeiten und Projekte.

T 98.2 Thu 16:35 T-H29 The characterisation of non collisions background events in the ATLAS detector during Run-2 data taking. — •SERGIO GRANCAGNOLO — Humboldt-Universität, Berlin, Germany

Understanding events from proton interactions with residual gas in the beam pipe, with collimators, or from cosmic rays is of primary importance to identify

Location: T-H29

potential risk of damage to the accelerator and experiments. In addition, these events represent one of the main background on non-conventional physics signatures based on tracks not pointing to the interaction point, out-of-time energy deposits, or displaced decay vertices might come from signals released by longliving heavy particles. The characteristics of these non-collision backgrounds are illustrated in detail in order to identify, estimate and reject them by using all the ATLAS detector.

T 98.3 Thu 16:50 T-H29

BGNet: A neural network for beam background prediction for SuperKEKB — •YANNIK BUCH, ARIANE FREY, and BENJAMIN SCHWENKER — II. Physikalisches Institut, Georg-August-Universität Göttingen, Friedrich-Hund-Platz 1, 37077 Göttingen, Deutschland

In pursuit to understanding the observed CP-violation in our universe, the Belle II detector investigates the b-sector by measuring the decays of the $\Upsilon(4S)$ resonance. These resonances are produced by the SuperKEKB accelerator at KEK in Tsukuba, Japan. The goal of SuperKEKB is to achieve an instantaneous luminosity of 6.5×10^{35} cm⁻²s⁻¹. Currently, a luminosity of 5×10^{34} cm⁻²s⁻¹ is reached, showing that considerable improvements to the beam focusing and increases of the ring currents are still necessary. A key component to achieve the design luminosity is the nano beam scheme. At the same time, however, the Belle II detector must not be damaged or its performance compromised by extensive radiation and hit rates.

The beam backgrounds at Belle II are mostly composed of storage backgrounds, luminosity-based backgrounds and injection backgrounds of both rings due to the top-up injection scheme. BGNet is trained to predict the overall hit rates and their composition in terms of background source for various Belle II sub-detectors. The input data for BGNet are 1 Hz time series of variables describing the state of the SuperKEKB accelerator. This helps to monitor and mitigate the beam backgrounds during future operation.

T 98.4 Thu 17:05 T-H29 Low-background poly(ethylene naphthalate) as active structural material for the LEGEND-200 $0\nu\beta\beta$ experiment — •FELIX FISCHER for the PEN-Collaboration — Max-Planck-Institut für Physik, München, Deutschland

Poly(ethylene naphthalate), PEN, is a widely used industrial polyester which intrinsically scintillates blue light and has very good mechanical properties, also at cryogenic temperatures. This makes PEN an ideal candidate for self-vetoing structural materials in the close surrounding of ultra low background detectors for the search of extremely rare events like $0\nu\beta\beta$ decay. The process from procurement of commercially available PEN pellets to an optically active low background support-structure to be used in the next generation $0\nu\beta\beta$ -decay search experiment LEGEND-200 will be presented.

T 98.5 Thu 17:20 T-H29

Setup for a ground parameter measurement for the radio detection at the Pierre Auger Observatory* — •JANNIS PAWLOWSKY for the Pierre Auger-Collaboration — Bergische Universität Wuppertal, Gaußstraße 20, 42119 Wuppertal, Deutschland

The Pierre Auger Observatory is the largest observatory in the world measuring ultra high energy cosmic rays (UHECR). With the on-going AugerPrime upgrade, the hybrid detection of the 27 Fluorescence Telescopes and 1660 Water Cherenkov Detectors (WCD) is complemented by the Radio Detector (RD) on top of each WCD. The mounted antenna detects the radio emission of the air shower.

The reconstruction of the UHECR properties with radio signals is dependent on environmental parameters like atmospheric (weather) conditions and

(DOMs), attached to 86 cable strings, which are embedded in the Antarctic ice in a depth between 1.5 to 2.5 kilometers. The reconstruction of neutrino events is

Physics Institute III B, Aachen, Germany

a depth between 1.5 to 2.5 kilometers. The reconstruction of neutrino events is based on the arrival times of Cherenkov light at the position of these DOMs. Currently, these positions are only known with a precision of about one meter, because of the uncertainty during the hole drilling process. For several calibration purposes IceCube has produced data-sets in which flashes from LEDs installed inside the DOMs are detected by the surrounding DOMs. The transit times of these signals can be converted into distances between each pair of DOMs. We have developed a multilateration algorithm which fits the positions of all DOMs to these distances in a global maximum-likelihood analysis. The here presented algorithm can be applied beyond IceCube also to the acoustic calibration systems that are foreseen in the IceCube Upgrade and IceCube-Gen2.

T 99: Neutrino Astronomy 4

Time: Thursday 16:15-18:35

Group Report T 99.1 Thu 16:15 T-H30 Status of the KM3NeT experiment and contributions from ECAP

- •RODRIGO GRACIA-RUIZ for the ANTARES-KM3NET-ERLANGEN-Collaboration — Erlangen Centre for Astroparticle Physics, Erlangen, Germany There are exciting times ahead for neutrino physics and neutrino astronomy! Despite the pandemic the KM3NeT detectors keep growing at two locations in the depth of the Mediterranean Sea. They currently host a total of 18 detection units, 10 for the ORCA detector in France and 8 for the ARCA detector off the coast of Sicily, with a total of more than 10000 photomultiplier tubes in 324 optical modules. Further significant extensions of the detector arrays are under way. This already enables sensitive investigations of GeV-scale atmospheric and TeV-to-PeV-scale cosmic neutrinos during the ongoing commissioning phase. We will report on the status of the KM3NeT detectors, first results on flavour oscillations, and the contributions from the ECAP team to these scientific endeavours. soil composition. The ground conditions are especially important for inclined air showers, where a non-negligible fraction of the radio signal is reflected off ground prior to being measured by the antenna. Knowledge of the ground parameters, namely permittivity, ε , and conductivity, σ , is therefore essential for a precise reconstruction.

The usage of constant reference values for ε and σ is not applicable for the extended RD grid. With an area of approximately 3000 km^2 it consists of many different soil types and has distinct weather conditions. In this talk we present a campaign for a measurement of the different soil types and a permanent stable and cost efficient setup in order to detect relative changes due to varying weather conditions.

*Gefördert durch die BMBF Verbundforschung Astroteilchenphysik (Vorhaben 05A20PX1)

T 98.6 Thu 17:35 T-H29

Rydberg Background reduction in KATRIN experiment using THz Radiation* — •SHIVANI RAMACHANDRAN and ENRICO ELLINGER for the KATRIN-Collaboration — Bergische Universitaet Wuppertal, Germany

One of the key requirement for the KATRIN experiment to reach its goal sensitivity in measuring the neutrino mass is minimal background. In order to achieve this and eliminate some known contributors, several background suppression methods have already been implemented. Presently the most prominent contribution to the background in the measured signal are electrons produced by thermal ionization of Rydberg atoms which originate from the walls of the main spectrometer.

A plausible method is using THz and microwave radiation which can lead to a reduced lifetime of Rydberg atoms and allow for dedicated stimulated deexcitation. The influence of THz light source in the main spectrometer along with the state and spatial evolution of the Rydberg atoms is studied via simulations. The transition and ionization rates due a light source depend mainly on its frequency, intensity and spectral width. A range of frequencies in the THz regime with different intensities are tested for background reduction. Influence of currently available broadband and narrowband THz sources are also studied. The results show important parameters namely the radiative power, frequency range and number of sources required for effective background reduction with this method. *Gefördert durch die BMBF-Verbundforschung Astroteilchenphysik

Geometry Calibration of IceCube - A new Likelihood-based Multilatera-

tion Approach — •SASKIA PHILIPPEN, CHRISTOPH GÜNTHER, and CHRISTO-PHER WIEBUSCH for the IceCube-Collaboration — RWTH Aachen University -

The IceCube Neutrino Observatory consists of 5160 Digital Optical Modules

T 98.7 Thu 17:50 T-H29

Location: T-H30

T 99.2 Thu 16:35 T-H30

Studying optical water properties with atmospheric muon events in KM3NeT/ORCA — •MARTIN SCHNEIDER for the ANTARES-KM3NET-ERLANGEN-Collaboration — ECAP, Friedrich-Alexander-Universität Erlangen-Nürnberg

KM3NeT-ORCA is an underwater neutrino detector featuring a dense configuration of optical modules, designed for the detection of atmospheric neutrinos down to the low-GeV energy regime. Located in a deep-sea environment, the detector performance depends on the optical water properties. In this talk, a study of the optical properties of the sea water surrounding the ORCA detector is presented. For this, a sample of atmospheric muons is compared to sets of Monte Carlo simulations obtained by varying the light attenuation length.
T 99.3 Thu 16:50 T-H30

Energy spectrum unfolding for Supernova burst neutrinos in JUNO — •THILO BIRKENFELD for the JUNO-Collaboration — RWTH Aachen University Since the detection of neutrinos emitted by the supernova SN 1987A, no neutrinos from other supernovae have been observed to date. The Jiangmen Underground Neutrino Observatory (JUNO) will measure the neutrino burst from a galactic supernova explosion. High statistics, a low detection threshold, and an excellent energy resolution will strongly constrain the details of the neutrinodriven supernova mechanism. JUNO will be sensitive to signals from all neutrino flavours via different detection channels. The reconstruction of their respective energy spectra requires an effective event classification, whose preliminary results will be presented in this talk. A subsequent bayesian-based energy spectrum unfolding method for reconstructing the initial neutrino energy distribution will also be presented.

T 99.4 Thu 17:05 T-H30

Constraining neutrino mass using black hole formation during supernova neutrino emission — •GEORGE PARKER and MICHAEL WURM — Johannes Gutenberg Universität Mainz, Mainz, Germany

In this work, we study how neutrino emission from supernovae collapsing to black holes could be used constrain the absolute neutrino mass. In the case where a black hole forms during a core-collapse supernova, it would lead to a sharp cut-off in the neutrino flux. An abrupt drop-off in the neutrino emission offers a clear-cut stage to look for neutrino time-of-flight effects, allowing stricter constraints to be set on the neutrino mass compared to previous estimates. We focus on the possibility that supernova neutrinos are detected with the Jiangmen Underground Neutrino Experiment (JUNO), a next-generation neutrino experiment with enhanced flavour sensitivity, exceptional energy resolution and high statistics. Using three-dimensional core-collapse supernova simulations, the sensitivity of JUNO to the absolute neutrino mass is evaluated.

T 99.5 Thu 17:20 T-H30

Studies on Solar Be7 Neutrino Measurements and Applications in JUNO — •SEBASTIAN ZWICKEL^{1,2}, LOTHAR OBERAUER¹, SIMON CSAKLl¹, CARSTEN DITTRICH¹, DAVID DÖRFLINGER¹, ULRIKE FAHRENDHOLZ¹, FLORIAN KÜBELBÄCK¹, MATTHIAS MAYER¹, VINCENT ROMPEL¹, LUCA SCHWEIZER¹, KONSTANTIN SCHWEIZER¹, KORBINIAN STANGLER¹, and RAPHAEL STOCK¹ for the JUNO-Collaboration — ¹Technische Universität München — ²Helmholtz Zentrum Dresden Rossendorf

Besides its major physics goal, the determination of the neutrino mass ordering, the upcoming Jiangmen Underground Neutrino Observatory (JUNO) will have a rich physics program. One part of this are solar neutrinos, where JUNO benefits the most from its large target mass of 20 kt liquid scintillator. In this talk the results of studies on searching for periodic flux variation, e.g. caused by solar g-modes, in the solar (Be7) neutrino flux, as well as the possible use of solar Be7 neutrinos for detector monitoring will be presented.

This work is supported by the DFG research unit "JUNO", the DFG collaborative research center 1258 "NDM", and the DFG Cluster of Excellence "Origins.

T 99.6 Thu 17:35 T-H30

First Comparison of Ballistic and Diffusive Propagation in Flares of Blazar Jets - Implications for Neutrino Emission Models — •MARCEL SCHROLLER¹, JULIA BECKER TJUS¹, PATRICK REICHHERZER^{1,2}, and MARIO HÖRBE^{1,3} — ¹Ruhr-Universität Bochum, Theoretische Physik IV — ²IRFU, CEA, Université Paris-Saclay — ³Oxford Astrophysics, University of Oxford

Active galactic nuclei (AGN), and the accompanied jets, are some of the most luminous objects in the observable Universe. Both the active cores and their jets are candidates for the engine of cosmic rays, gamma rays, and neutrinos with the highest energies measured at Earth. In 2017, IceCube recorded an extragalactic high energy neutrino event with a strong hint of a directional coincidence with the position of a known jetted AGN TXS0506+056. A deep understanding of the processes related to jets will fuel the field of high energy cosmic rays, fundamental plasma, astro, and particle physics. The physical and mathematical modelling of an AGN jet is challenging, with ambiguous signatures that need to be understood by numerical simulations of cosmic ray transport and interactions. Based on the work of Hoerbe et al. (MNRAS 2020), a simulation framework for hadronic constituents and their interactions inside of a plasmoid, propagating along the AGN jet axis, was made. For this talk, we tested several state-of-the-art simulation setups from the literature in this field with our framework to analyse the assumptions about propagation behaviour both ballistically and diffusively. We present the results and point out, where those assumptions cannot hold in a realistic setup.

T 99.7 Thu 17:50 T-H30

A new code for the modeling of multimessenger flares from blazars — •LEANDER SCHLEGEL^{1,2}, MARCEL SCHROLLER^{1,2}, MARIO HÖRBE^{1,2,3}, and JULIA BECKER TJUS^{1,2} — ¹Theoretische Physik IV, Ruhr Universität Bochum, Bochum, Germany — ²RAPP-Center at Ruhr Universität Bochum, Bochum, Germany — ³University of Oxford, Oxford Astrophysics, Denys Wilkinson Building, Keble Road, Oxford, OX1 3RH, United Kingdom

Since their discovery over a century ago, the origin of cosmic rays of the highest energies is still widely uncertain. While in the past much attention was paid to analyzing steady state source models, bursting sources that appear in quiescent and flaring states, like the class of Active Galactic Nuclei (AGN) seem to be a promising candidate for possible sources of ultra-high-energy cosmic rays (UHECR). The goal of this work is trying to understand the detailed behaviour of bursting sources by simulating the time resolved propagation of a plasma blob inside the jet of an AGN. For this purpose, a tool for cosmic-ray propagation in relativistic plasmoids of AGN jets has been developed and implemented into the open-source code CRPropa 3.1. With this framework, we will predict the multimessenger signatures of flaring sources, aiming to contribute to a more complete picture of the UHECR sky including the bursting sources and therefore also to a deeper understanding of the origin of the highest energetic charged particles. First results of the flaring behaviour from relativistic plasmoids are being presented.

T 99.8 Thu 18:05 T-H30

Investigation of the effect of elliptical orbits in supermassive binary black holes at the example of the neutrino lightcurve of the blazar TXS0506+056 — •JOHANNES JUST, JULIA BECKER TJUS, and ILJA JAROSCHEWSKI — Theoretische Physik IV, Ruhr-Universität Bochum

IceCube detections from 2014/15 and 2017 show two possible high-energy neutrino correlations with the blazar TXS0506+056, making blazars promising candidates for high-energy neutrino emission. Those neutrinos can be produced in pp or py interactions of cosmic rays, making bazars possible sources of high energy cosmic rays. Two separate detections might imply a periodicity of the neutrino flux from TXS0506+056 at Earth.

Such a periodicity can be explained by a precession of the heavier super massive black hole jet in a merger of a super massive binary black hole (SMBBH), caused by a Spin-Flip of the Jet. Considering the post newtonian mechanics up to the 2.5 order, the Spin-Flip-Phenomenon is described with the Spin (and therefore the Jet) slowly aligning with the total angular momentum.

Assuming that TXS0506+056 is a SMBBH merger, this work predicts the upcoming neutrino flux as well as the observability of the emitted gravitational waves with LISA, taking different eccentricities of the SMBBH orbit into account. Several eccentricities, leading to differing periodicities and shrinking timescales, are discussed.

T 99.9 Thu 18:20 T-H30

A novel Machine Learning-approach for the detection of the DSNB — •DAVID MAKSIMOVIĆ¹, MICHAEL NIESLONY², and MICHAEL WURM³ — ¹Johannes Gutenberg-Universität Mainz — ²Johannes Gutenberg-Universität Mainz — ³Johannes Gutenberg-Universität Mainz

The Diffuse Supernova Neutrino Background (DSNB) is the faint signal of all core-collapse supernovae explosions on cosmic scales. A prime method for detecting the DSNB is finding its IBD signatures in Gadolinium-loaded large water Cherenkov detectors like Super-Kamiokande(SK-GD). While the enhanced neutron tagging capability of Gadolinium greatly reduces single-event backgrounds, correlated events mimicking the IBD coincidence signature remain a potentially harmful background. Especially in the low-energy range of the observation window, Neutral-Current (NC) interactions of atmospheric neutrinos dominate the DSNB signal, which leads to an initial signal-to-background (S:B) ratio inside the observation window of about 1:10.

Here, we report on a novel machine learning method based on Convolutional Neural Networks (CNNs) that offer the possibility for a direct classification of the PMT hit patterns of the prompt events. Based on the events generated in a simplified SK-GD-like detector setup, we find that a trained CNN can maintain a signal efficiency of 96 % while reducing the residual NC background to 2 % of the original rate, corresponding to a final signal-to-background ratio of about 4:1. This provides excellent conditions for a DSNB discovery.

This work has been funded by the Excellence Cluster PRISMA+.

T 100: Cosmic Ray 5

Time: Thursday 16:15-18:30

T 100.1 Thu 16:15 T-H32

Low-Energy Cosmic Ray Composition with IceCube and IceTop – •JULIAN SAFFER for the IceCube-Collaboration — Institut für Experimentelle Teilchenphysik, Karlsruher Institut für Technologie (KIT)

IceTop is the surface component of the IceCube Neutrino Observatory at the geographic South Pole and dedicated to the indirect detection of cosmic rays (CRs). Studying the primary CR spectrum and mass composition around the knee requires a dedicated IceTop trigger for smaller air showers initiated by lower-energy primaries as well as the combination of surface (predominantly electromagnetic) and corresponding in-ice (muonic) signals. Monte-Carlo simulation data of air showers at IceCube ranging down to $E_0 = 10^5$ GeV have been used to train boosted decision trees for the reconstruction of shower core position, zenith angle, primary energy and mass of the incoming CR particles.

This talk presents the input features fed into the different machine learning models, the chosen model architectures and reconstruction results for four primary mass types. Additionally, plans towards an enhancement of the reconstruction utilizing a set of convolutional neural networks are discussed.

T 100.2 Thu 16:30 T-H32

Simulation Study of Atmospheric Muons with IceCube-Gen2 — •JONATHAN MESSNER, AGNIESZKA LESZCZYNSKA, and ANDREAS HAUNGS for the IceCube-Collaboration — Institute for Astroparticle Physics, Karlsruhe Institute of Technology (KIT), Karlsruhe, Germany

The next generation of the IceCube Neutrino Observatory called IceCube-Gen2 will extend the detector's capabilities in both, neutrino and cosmic-ray measurements. In particular, the combination of the in-ice optical modules and the large array of surface detectors will enhance the understanding of extensive air showers and the studies of cosmic rays. Muons produced in air showers can deliver relevant information not only about incoming cosmic rays but also about properties of the air showers. Conventional atmospheric muons are produced by decays of pions and kaons, while prompt muons originate mainly from decay of charmed and unflavoured mesons. This prompt component is expected to dominate the muon flux at higher energies. Due to larger aperture for coincident measurements, with surface and in-ice arrays, IceCube-Gen2 has the potential to measure this prompt component in relation to the properties of parent cosmic rays. In this contribution high energy muons, especially prompt muons, will be studied based on air shower simulations in order to better understand the capabilities of IceCube-Gen2.

T 100.3 Thu 16:45 T-H32

Improving gamma-hadron separation for air showers at the IceCube Neutrino Observatory — •FEDERICO BONTEMPO for the IceCube-Collaboration — Institute for Astroparticle Physics, Karlsruhe Institute of Technology (KIT), Karlsruhe, Germany

The IceCube Neutrino Observatory is an experiment located at the geographic South Pole. It is composed of two detectors: an optical array deep in the ice and an array of ice-Cherenkov tanks at the surface called IceTop. The combination of the two detectors can be exploited for the study of cosmic rays. The in-ice detector measures the high-energy muonic component of air showers, whereas the signal strength on IceTop is dominated by the electromagnetic component. The aim of this work is to discriminate between photon and hadron initiated air showers. This discrimination was already attempted using a machine learning technique named Random Forest. Here, a different approach is presented which uses both Random Forests and deep learning techniques, in particular, supervised learning techniques that predicts unknown data after studying labeled data. The physics quantities used for this study are the charges measured by the in-ice detector, the reconstructed zenith angle, the in-ice containment of the shower, the reconstructed energy and a likelihood estimator that captures both the presence of individual muons and charge fluctuations in the surface array.

Furthermore, the planned enhancement of IceTop, comprised of surface radio antennas and scintillator panels, will contribute to the improvement of the gamma-hadron separation.

T 100.4 Thu 17:00 T-H32

Mass-sensitive parameter with the IceTop surface array — •DONGHWA KANG for the IceCube-Collaboration — Karlsruher Institut für Technologie (KIT) IceTop, the surface component of the IceCube Neutrino Observatory at the South Pole, measures the air showers of cosmic ray with energies from PeV up to EeV. By means of the charge signal measurements only with the IceTop surface array, a parameter sensitive to the muon content was defined and estimated event by event. In this contribution, the estimated mass-sensitive parameter and its dependencies on the hadronic interaction models will be presented. In addition, the applicability of energy and mass composition reconstruction of cosmic rays will be discussed. Location: T-H32

T 100.5 Thu 17:15 T-H32

Sensitivity of IceCube-Gen2 for High-Energy Cosmic Ray Anisotropy Studies — •WENJIE HOU and DONGHWA KANG for the IceCube-Collaboration — Institute for Astroparticle Physics, Karlsruhe Institute of Technology (KIT), Karlsruhe

At which energy the transition from galactic to extra-galactic cosmic rays (CRs) takes place is one of the major unresolved issues of cosmic ray physics. Although the sources of high-energy cosmic rays remain unknown, one expects to identify them by studying the anisotropy in their arrival directions. Recently, the cosmic ray anisotropy measurements in TeV to PeV energy range were updated from IceCube using its nine years of data. IceCube-Gen2 is designed to have a detector volume about 8 times larger than IceCube and will achieve an increased exposure to cosmic rays by a factor of 30. This improvement could allow us to obtain high-quality measurements of cosmic ray air showers and investigate the cosmic ray anisotropy with higher sensitivity.

The sensitivity of IceCube-Gen2 to anisotropy is purely a matter of statistics. Hence, based on the simulation of IceCube-Gen2, the first attempt is to make an exposure map taking into account detection efficiency and resolution. We will build a Monte Carlo toy simulation model of IceCube-Gen2. In this case, the expected maps will be generated with random events, resulting in the angular power spectrum. Eventually, we can determine under what conditions IceCube-Gen2 could achieve the highest sensitivity to observe cosmic ray anisotropy. In this contribution, the current studies on the anisotropy with the simulation will be discussed.

T 100.6 Thu 17:30 T-H32 First results of the IceCube Surface Array Enhancement Prototype – •MARIE

OEHLER for the IceCube-Collaboration — KIT, Karlsruhe, Germany The IceCube Observatory is a cubic-kilometer neutrino detector installed in the ice at the geographic South Pole. To increase the efficiency of detecting astrophysical neutrinos the upgrade IceCube-Gen2 is under development. To also boost the sensitivity of the surface array, IceTop, an enhancement consisting of

a hybrid scintillation-detector and radio-antenna array is planned. An optimized prototype station, consisting of eight scintillation detectors and three radio antennas, was deployed in January 2020. Both, scintillation detectors and radio antennas, are read out by a central hybrid data acquisition system (DAQ). The scintillation detectors transfer digitized integrated signals to the DAQ to minimize the amount of transmitted data and trigger the radio antennas. The radio waveforms are transferred as analog signals to the central DAQ and are digitized and read out, when triggered by the scintillation detectors. In this contribution the first measurement results will be shown.

T 100.7 Thu 17:45 T-H32

Measurements with the IceCube Surface Array Enhancement prototype — •HRV0JE DUJMOVIC for the IceCube-Collaboration — Institut für Astroteilchenphysik, Karlsruher Institut für Technologie (KIT)

IceTop, the surface array of the IceCube Neutrino Observatory, currently consists of 162 ice Cherenkov tanks distributed over an area of 1 km. IceTop is used for cosmic-ray air shower detection and as a veto for the in-ice neutrino detector. The science case of IceTop will be greatly improved by complementing the existing detectors with an array of radio antennas and scintillator panels. The enhancement array will cover the same footprint as IceTop and will consist of 32 stations. One such station, consisting of 3 radio antennas and 8 scintillator panels, was deployed at the South Pole in January 2020. In this talk, we will present the measurements with the prototype station. The results obtained from the prototype station will help us better understand the full capabilities and physics potential of the IceCube*s surface enhancement.

T 100.8 Thu 18:00 T-H32

Status of the R&D and production of the scintillation detectors for the Surface Array Enhancement — •SHEFALI SHEFALI for the IceCube-Collaboration — Institut für Astroteilchenphysik, Karlsruher Institut für Technologie (KIT), Karlsruhe, Germany

The IceCube Neutrino Observatory is a cubic kilometer scale detector deployed in the Antarctic ice and is involved in cosmic ray physics. The surface array of IceCube, IceTop, operates as a veto for the astrophysical neutrino searches and calibration detector for the IceCube in-ice instrumentation. Despite its contribution, the snow accumulation on top of these detectors results in an increased energy uncertainty in the detected particles and consequently, the shower reconstruction. Moreover, the enhancement of IceTop will lead to a better measurement of the extensive air showers and improve the astrophysics of the highenergy cosmic-ray sky.

Enhancing IceTop with a hybrid array of scintillation detectors and radio antennas will lower the energy threshold for air-shower measurements, provide more efficient veto capabilities, enable the separation of the electromagnetic and muonic shower components and improve the detector calibration by compensating for snow accumulation. Following the success of the first complete prototype station consisting of three radio antennas and eight scintillation detectors deployed at the South Pole in 2020, the R&D and production of detectors for a total of 32 stations is ongoing. The production challenges, deployment status, and calibration methods of the scintillation detectors will be discussed in this contribution.

T 100.9 Thu 18:15 T-H32

An IceCube Surface Array Enhancement station for deployment at Telescope Array — •NOAH GOEHLKE for the IceCube-Collaboration — Institut für Astroteilchenphysik, Karlsruher Institut für Technologie (KIT), Karlsruhe

The IceTop array, located on the surface of the IceCube Neutrino Observatory, will be enhanced with hybrid radio and scintillator stations. The DAQ of each station is housed in a FieldHub. In January 2020 a full prototype station was deployed and is successfully operating and taking data. For the planned IceCube-

Gen2 facility, the DAQ of the surface array and the in-ice array will be combined, using a modified FieldHub. The development of this FieldHub will be performed by the University of Utah, which is also contributing to the Telescope Array (TA), an air-shower detector array located in Utah.

By deploying a prototype station at TA, the Univerity of Utah is provided with the preliminary hardware of the future surface array, which is needed to design the new FieldHub. In addition, it can serve as a testing platform for IceCube-Gen2 and it enables cross calibration with TA. Since the environment and infrastructure in Utah and the South Pole differ significantly, adjustments of the prototype station are in development. As example, the detectors have to be able to measure air-shower particles at much higher ambient temperature and humidity levels as found at the South Pole.

In this contribution the adapted design of the prototype station as well as experiments done to investigate the detectors behavior at higher temperatures will be presented.

T 101: Cosmic Ray 6

Time: Thursday 16:15-18:30

T 101.1 Thu 16:15 T-H33

Implications of turbulence dependent diffusion on Galactic cosmic ray — •JULIEN DÖRNER^{1,2}, PATRICK REICHHERZER^{1,2,3}, JULIA BECKER TJUS^{1,2}, and HORST FICHTNER^{1,2} — ¹Theoretische Physik IV, Ruhr University Bochum, Bochum, Germany — ²RAPP-Center at Ruhr University Bochum, Bochum, Germany — ³IRFU, CEA, Université Paris-Saclay, F-91191 Gif-sur-Yvette, France The motion of Galactic cosmic rays is dominated by spatial diffusion. Therefore, in order to describe their transport, a detailed knowledge of the diffusion tensor $\hat{\kappa}$ is necessary. This tensor depends on the particle energy, the structure of the local background field \vec{B} , and its turbulent component \vec{b} . Recent numerical analyses of diffusion coefficients in three-dimensional, isotropic turbulence show a discrepancy between its energy scaling for intermediate turbulence levels b/Band the corresponding quasi-linear prediction.

In this talk we report about probing different models for the diffusion tensor and its dependence on energy and turbulence level. We compare the results with observations of the gradient in the cosmic-ray density and of the spectral energy behavior in the Milky Way by Fermi-LAT.

T 101.2 Thu 16:30 T-H33

Optimization of the Numerical Calculation of the Field Line Random Walk **Diffusion Tensor** — •JAN-NIKLAS BOHNENSACK¹, PATRICK REICHHERZER^{1,2}, JULIA BECKER TJUS^{1,2}, LEANDER SCHLEGEL^{1,2}, and JULIEN DÖRNER^{1,2} — ¹Theoretical Physics IV: Plasma-Astroparticle Physics, Faculty for Physics & Astronomy, Ruhr-Universität Bochum, D-44780 Bochum, Germany – ²Ruhr Astroparticle And Plasma Physics Center (RAPP Center), Bochum, Germany The goal of this talk is to calculate the Field Line Random Walk (FLRW) diffusion coefficient. FLRW is the random movement of the magnetic fieldlines in a turbulent field, particles with low rigidities follow those lines and diffuse accordingly. The FLRW diffusion coefficient only depends on the turbulence of the magnetic fields and is particle-energy independent. That makes their calculation more universal and computationally more efficient. Therefore, the coefficient is of high relevance for providing astrophysical simulations for cosmic-ray transport with a fundamental description of the diffusion tensor. These can be applied in environments such as the Milky Way, but also relativistic plasmoids in jets of active galaxies. To check the agreement of the fieldlines to the corresponding particle trajectories, a GPU based visualization software based on the software Vispy was developed in this bachelor thesis. With the optimized FLRW software, we will present first results and interpretations of the FLRW diffusion tensor for different turbulence levels. Furthermore, we will show examples of 3D plotted fieldlines and particle trajectories with our visualization tool.

T 101.3 Thu 16:45 T-H33

Proton event reconstruction with the MAGIC experiment — •ALICIA FAT-TORINI and MAXIMILIAN NOETHE for the MAGIC-Collaboration — Astroparticle Physics WG Elsässer, TU Dortmund University, Germany

Air showers induced by cosmic protons and heavier nuclei form the dominant background for very high energy gamma-ray observations with Imaging Air Cherenkov Telescopes. Even for strong very high energy gamma-ray sources the signal-to-background ratio in the raw data is typically less than 1:5000, so a very large statistic of cosmic proton and heavier nuclei induced events are available as a byproduct of gamma-ray source observations. In this contribution, we present the reconstruction of the particle type of primary events and the energy reconstruction of the events classified as protons. For this purpose, we used a random forest method trained and tested by using Monte Carlo simulations from the MAGIC telescopes, for energies above 70 GeV. We use the aict-tools framework, which includes machine learning methods for the particle type classification and energy reconstruction. The open-source Python project aict-tools was developed at TU Dortmund University and its reconstruction tools are based on scikit-learn predictors. Finally, an unfolding taking into account the back-ground is performed to compensate for the typical bias of the random forest results. Here we report on the performance of the proton event reconstruction using the well-tested and robust random forest approach.

T 101.4 Thu 17:00 T-H33 Search for heavy antimatter with AMS — • ROBIN SONNABEND — 1. Physikalisches Institut B, RWTH Aachen

The Alpha Magnetic Spectrometer (AMS-02) on the International Space Station has been performing precision measurements of cosmic rays in the GeV to TeV energy range since 2011. The search for heavy antimatter (Z>=2) requires advanced methods for the suppression of instrumental background which arises from the mis-reconstruction of the charge sign. I will present a set of dedicated multivariate estimators for different event topologies designed to achieve this goal.

T 101.5 Thu 17:15 T-H33

Location: T-H33

Galactic gamma-ray and neutrino emission from interacting cosmic-ray nuclei — •MISCHA BREUHAUS¹, JAMES ANTHONY HINTON¹, VIKAS JOSHI², BRIAN REVILLE¹, and HARM SCHOORLEMMER^{1,3} — ¹Max-Planck-Institut für Kernphysik, Saupfercheckweg 1, 69117 Heidelberg, Germany — ²Friedrich-Alexander-Universität Erlangen-Nürnberg, ECAP, Erwin-Rommel-Str. 1, D 91058 Erlangen, Germany — ³IMAPP, Radboud University Nijmegen, Nijmegen, The Netherlands

We present a study of the expectations for very/ultra high energy (VHE/UHE) gamma-ray and neutrino emission from interacting cosmic rays in our Galaxy and a comparison to the latest results for the Galactic UHE diffuse emission. We demonstrate the importance of properly accounting for the mixed cosmic-ray composition as well as gamma-ray absorption. We adopt the wounded-nucleon model of nuclei interaction and provide parameterisations of the resulting gamma-ray and neutrino production. Nucleon shielding due to clustering inside nuclei is shown to have a measurable effect on the production of gamma-rays and is particularly evident close to breaks and cut-offs in mixed composition particle spectra. The change in composition around the 'knee' in the cosmic ray spectrum has a noticeable impact on the diffuse neutrino and gamma-ray emission spectra. We show that current and near-future detectors can probe these differences in the key energy range from 10 TeV to 1 PeV, testing the paradigm of the universality of the cosmic ray spectrum and composition throughout the Galaxy.

T 101.6 Thu 17:30 T-H33

Vacuum-Cherenkov radiation in UHE air showers: a way of probing Lorentz violation — •FABIAN DUENKEL, MARCUS NIECHCIOL, and MARKUS RISSE — Center for Particle Physics Siegen, Experimentelle Astroteilchenphysik, Universität Siegen

In extensive air showers induced by ultra-high energy (UHE) cosmic rays, secondary electrons are expected to be produced at energies far above those accessible by other means. Those high energies can be used to search for new physics, in particular we study the effects of isotropic, nonbirefringent Lorentz violation in the photon sector. In the case of a photon velocity which is larger than the maximum attainable velocity of standard Dirac fermions, vacuum-Cherenkov radiation becomes possible, which can lead to significant changes of the shower development. Implementing this Lorentz-violating effect in air shower simulations, we present first results on the impact on the shower development, specifically on the average atmospheric depth of the shower maximum $\langle X_{\max} \rangle$ and its shower-to-shower fluctuations $\sigma(X_{\max})$.

This work is supported by the Deutsche Forschungsgemeinschaft (DFG).

T 101.7 Thu 17:45 T-H33

COREAS simulation for the GRAND project — •CHAO ZHANG, TIM HUEGE, TANGUY PIEROG, MARKUS ROTH, ANDREAS HAUNGS, FRANK SCHROEDER, and RALPH ENGEL — Institut fuer Astroteilchenphysik, Karlsruher Institut fuer Technologie-Campus Nord, Post-fach 3640, 76021 Karlsruhe, Germany

The GRAND project starts deploying antennas this year, which will give birth to its first stage, GP300 with 300 antennas in the near future. A new version of CORSIKA7 has been adapted and validated to simulate upward-going air showers for radio detection which will be the main detection channel in this project. A library of air showers is made with CoREAS by applying the best knowledge of GRAND including the atmospheric model and magnetic field of the site. A detailed analysis of the new patterns induced by inclined air showers leads to a better understanding of the scenario of their radio emission from higher to lower air density.

T 101.8 Thu 18:00 T-H33

Diffractive and radiative corrections to muon energy loss cross-sections — •ALEXANDER SANDROCK — Bergische Universität Wuppertal

High-energy muons can travel large distances before reaching underground detectors, for example cosmic-ray detectors and neutrino telescopes. The accurate simulation of muon transport through matter is therefore especially important for underground experiments. The dominant energy loss processes are ionization and at higher energies pair production, bremsstrahlung and inelastic interaction with nuclei. To reduce uncertainties in the simulation of muon transport, the calculation of higher-order corrections to these cross-sections is necessary. In this contribution, diffractive and radiative corrections to the cross-sections of bremsstrahlung and pair production are discussed.

T 101.9 Thu 18:15 T-H33

Results from a Pilot Study on Measurement of Fragmentation of Intermediate Mass Nuclei with NA61/SHINE at CERN — •NEERAJ AMIN for the NA61/SHINE-Collaboration — Institute for Astroparticle Physics, Karlsruhe Institute of Technology, Karlsruhe, Germany

Cosmic-ray propagation in the Galaxy can be constrained by modeling the secondary-to-primary cosmic-ray flux ratios, like the Boron-to-Carbon flux ratio that reaches Earth. While these fluxes are currently measured with high precision (<5%) by space-based detectors like AMS-02 and CALET, insufficient knowledge of nuclear fragmentation cross sections hinders the precision with which we can constrain cosmic-ray propagation. Therefore, new laboratory measurements of fragmentation cross-sections above 10 A GeV/c are needed.

In this talk, we report on the analysis of pilot data on fragmentation taken in 2018 with the NA61/SHINE experimental facility at CERN with ¹²C projectiles at a $p_{\text{beam}} = 13.5 \text{ A GeV/c}$. The main aim of this pilot run was to demonstrate the ability of NA61/SHINE to measure nuclear fragmentation cross sections in C+p interactions. Two fixed targets, polyethylene (C₂H₄) and graphite were used to achieve this. We present a preliminary measurement of Boron production in C+p interactions including the contribution from the short-lived 'ghost nucleus' ¹¹C.

T 102: Neutrino Physics without Accelerators 7

Time: Thursday 16:15-18:35

T 102.1 Thu 16:15 T-H34

Sensitivity Study with Theia Detector — •WEI-CHIEH LEE, CAREN HAGNER, and BJÖRN WONSAK — Institut für Experimentalphysik, Universität Hamburg, 22761 Hamburg, Germany

With new technologies exploiting the advantages of both the Cherenkov and scintillation lights, a new type of neutrino detector, Theia, is able to determine the direction and species of incoming particles while still having a good energy resolution and low threshold. In consequence, Its ability to reach high levels of background rejection would allow us to improve the precision in the measurement of oscillation parameters when employed at a long baseline neutrino experiment. For this, the realization of Theia is planned at the far site of DUNE, whose near detector also provides information to further improve the measurement of neutrino oscillation. In this experiment, Theia can have enough sensitivity to the determination of mass hierarchy and CP phase of neutrino oscillation, the two missing pieces of our knowledge in the topic. In order to simulate such experiments, a software tool called GLoBES is developed for describing the detector properties and doing analysis easily, and is used in this sensitivity study.

T 102.2 Thu 16:30 T-H34

On the road to Theia: current status of the Mainz WbLS test cell DISCO – •MANUEL BÖHLES¹, DANIELE GUFFANTI¹, HANS STEIGER^{1,2}, and MICHAEL WURM¹ – ¹Johannes Gutenberg-Universität Mainz, Staudinger Weg 7, 55124 Mainz, Germany – ²Technische Universität München, James-Franck-Str. 1, 85748 Garching b. München, Germany

The detection of neutrinos using water-based liquid scintillators (WbLS) is a promising method in the field of detector development. Its strength lies in combining high-resolution energy determination with a low energy threshold through the use of scintillation light and in the directional reconstruction with the help of Cherenkov radiation. The spectrum of potential applications is broad, ranging from long-baseline oscillation experiments to the measurement of lowenergy solar neutrinos. The key point of this new technique is the discrimination between scintillation and Cherenkov photons, which can be achieved by exploiting the different chromatic features, time behaviour and angular emission. In order to characterise this innovative medium and to prove whether scintillation and Cherenkov radiation can be distinguished, we have built a test cell equipped with 16 ultrafast photomultipliers that will provide useful insights towards a new generation of detectors. In addition, complementary ultrafast photodetection systems (SiPM array, LAPPD) can be investigated in future studies. This work is supported by the BMBF Verbundprojekt 05H2018: R&D Detectors and Scintillators.

T 102.3 Thu 16:45 T-H34

Location: T-H34

Characterisation measurement of LAPPDs for *v***-detectors** — •BENEDICT KAI-SER, LUKAS BIEGER, DAVID BLUM, MARC BREISCH, SRIJAN DELAMPADY, JESSI-CA ECK, GINA GRÜNAUER, TOBIAS HEINZ, FRIEDER KOHLER, TOBIAS LACHEN-MAIER, AXEL MÜLLER, TOBIAS STERR, ALEXANDER TIETZSCH UND JAN ZÜFLE — Universität Tübingen, Physikalisches Institut, Auf der Morgenstelle 14, 72076 Tübingen

Large Area Picosecond Photodetectors (LAPPDs) are novel photodetectors suitable for use in upcoming neutrino detection experiments. LAPPDs incorporate a square multi-alkali photocathode, a chevron pair of microchannel plates (MCPs) for photoelectron multiplication and multiple anode strips for readout, all in a hermitically sealed package.

The design of the LAPPDs results in an unprecedented time resolution better than 70 ps and a spatial resolution of 2.5 mm and 0.8 mm in x- and y-direction respectively at uniform gains of 10^6 to 10^7 over a large active detector area of more than 370 cm².

Currently, we are commissioning a setup to test LAPPDs for their performance and key characteristics. This talk will outline the working principle as well as characteristics of an LAPPD and measurement results of the first LAPPD received from the manufacturer will be discussed.

Group Report T 102.4 Thu 17:00 T-H34 Neutrino mass determination with ¹⁶³Ho-loaded MMCs – the ECHo experiment — •ARNULF BARTH for the ECHo-Collaboration — Kirchhoff-Institute for Physics, Heidelberg University

The Electron Capture in ¹⁶³Ho experiment, ECHo, is a running experiment for the determination of the neutrino mass scale via the analysis of the end point region of the ¹⁶³Ho electron capture spectrum. In the first phase, ECHo-1k, about 60 MMC pixels enclosing ¹⁶³Ho ions for an activity of about 1 Bq per pixel have been operated for several months. The goal of this first phase is to reach a sensitivity on the effective electron neutrino mass below $20 \text{ eV}/c^2$ by the analysis of a $^{1\acute{6}3}\mathrm{Ho}$ spectrum with more than 10^8 events. We discuss the characterization of the single pixel performance and the stability over the measuring period. Results from the analysis of the acquired data will be presented with focus on data reduction efficiency and on the procedures to obtain the final high statistics spectrum. A preliminary analysis of the ¹⁶³Ho spectral shape will be described and the expected sensitivity on the effective electron neutrino mass on the basis of the properties of the presented spectrum will be discussed. In conclusion, we will present how the performance obtained by the MMC arrays used during the first phase of the ECHo experiment have led to the design of the MMC arrays for the second phase, ECHo-100k. In this phase, about 12000 MMC pixels each hosting ¹⁶³Ho for an activity of 10 Bq will be simultaneously operated thanks to the microwave SQUID multiplexing readout. Operating these arrays for three years will allow for reaching a sensitivity on the electron neutrino mass at the $1 \,\mathrm{eV}/c^2$ level.

T 102.5 Thu 17:20 T-H34

From Temperature pulses to high statistic Ho-163 spectrum: Analysis Algorithms for the ECHO Experiment — •MARKUS GRIEDEL, ARNULF BARTH, ROBERT HAMMANN, DANIEL HENGSTLER, NEVEN KOVAC, FEDERICA MAN-TEGAZZINI, ANDREAS FLEISCHMANN, and LOREDANA GASTALDO — Kirchhoff-Institute for Physics, Heidelberg University

The goal of the Electron Capture in Ho-163 (ECHo) experiment is the determination of the effective electron neutrino mass by analysing the electron capture (EC) spectrum of Ho-163. The ECHo experiment uses Metallic magnetic calorimeters (MMCs) operating at millikelvin temperatures, in which the Ho-163 has been implanted. In order to reliably infer the energy of Ho-163 events and discard triggered noise or pile-up events, fast and robust analysis algorithms are required. For this, algorithms based on filters acting on the trigger time of the events and on filters using pulse shape information were developed. To convert the measured temperature pulses into an energy spectrum, further steps are takes, as corrections for temperature shifts and energy calibration.

We describe the steps we took for the reduction of the data acquired during the first phase of the ECHo experiment, ECHo-1k; as well as the process to build a high statistic Ho-163 spectrum from data acquired with several single Ho-163 implanted MMCs.

T 102.6 Thu 17:35 T-H34

From ECHo-1k to ECHo-100k: Optimisation of the high-resolution metallic magnetic calorimeters with embedded ¹⁶³Ho for the determination of the electron neutrino mass — •NEVEN KOVAC¹, FEDERICA MANTEGAZZINI¹, LOREDANA GASTALDO¹, ARNULF BARTH¹, MARKUS GRIEDEL¹, ANDREAS FLEISCHMANN¹, MATTHEW HERBST¹, DANIEL HENGSTLER¹, ANDREAS REIFENBERGER¹, CHRISTIAN ENSS¹, CHRISTOPH DÜLLMANN², HOLGER DORRER², TOM KIECK³, NINA KNEIP³, and KLAUS WENDT³ — ¹Kirchhoff-Institut für Physik, Universität Heidelberg — ²Department of Chemistry -TRIGA Site, Johannes Gutenberg-Universität Mainz — ³Institute of Physics, Johannes Gutenberg-Universität Mainz

Large arrays of metallic magnetic calorimeters have been selected for the ECHo experiment due to the excellent energy resolution, the fast response time and the almost linear detector response which allows for a reliable energy calibration. Based on the performance achieved with the detector array developed for the first phase of the ECHo experiment, ECHo-1k, the design of the ECHo-100k arrays has been conceived. This new design features an optimized single pixel geometry, upgrade of the on-chip thermalisation layout and a high operational flexibility. First wafers with ECHo-100k arrays have been fully characterized. We summarise the performance achieved with the ECHo-1k and the newly developed ECHo-100k arrays in comparison with the design performance. We discuss how these results are important for achieving the goals defined for the ECHo-100k experiment.

T 102.7 Thu 17:50 T-H34 **The MONUMENT Experiment; ordinary muon capture as a benchmark for** $0\nu\beta\beta$ **decay nuclear structure calculations** — ELISABETTA BOSSIO¹, ELIZA-BETH MONDRAGON¹, STEFAN SCHÖNERT¹, •MARIO SCHWARZ¹, and CHRISTOPH WIESINGER^{1,2} for the MONUMENT-Collaboration — ¹Physik-Department, Technische Universität München, Garching — ²Max-Planck-Institut für Physik, München Extracting particle physics properties from neutrinoless double-beta $(0\nu\beta\beta)$ decay requires a detailed understanding of the involved nuclear structures. Still, modern calculations of the corresponding nuclear matrix elements (NMEs) differ by factors 2-3. The high momentum transfer of Ordinary Muon Capture (OMC) provides insight into highly excited states similar to those that contribute virtually to $0\nu\beta\beta$ transitions. The precise study of the *y*'s following the OMC process makes this a promising tool to validate NME calculations. The MONU-MENT collaboration is performing a series of explorative OMC measurements involving typical $\beta\beta$ decay daughter isotopes such as ⁷⁶Se and ¹³⁶Ba, as well as other benchmark isotopes. In this talk the experiment carried out at the Paul Scherrer Institute and first results from the beamtime in 2021 will be presented. This research is supported by the DFG Grant 448829699.

T 102.8 Thu 18:05 T-H34

Antineutrino Monitoring with Liquid Organic Time Projection Chambers — JOHANNES BOSSE, SARAH FRIEDRICH, MALTE GÖTTSCHE, HELGE HAVERESCH, •THOMAS RADERMACHER, STEFAN ROTH, GEORG SCHWEFER, and HAGEN WEIGEL — RWTH Aachen University - Physics Institute III B, Aachen, Germany For the first few hundreds of years the dominant radioactivity of nuclear waste comes from long-lived beta-decaying elements that are emitting antineutrinos in the low-energy region below 5 MeV. In a newly envisioned application for nuclear monitoring purposes, we want to use these antineutrinos to monitor the content of nuclear waste repositories. We are investigating a time projection chamber filled with an organic liquid aiming at full reconstruction of inverse beta decay events. In the low energy region, the direction of the neutron in IBD is strongly correlated to the direction of the incoming antineutrino. For this we study to which extend the neutron director medium. This talk gives an overview on our project and the progress of our simulation studies.

T 102.9 Thu 18:20 T-H34

Location: T-H35

T 103.2 Thu 16:30 T-H35

Development of the comprehensive analysis tools for the Supernova neutrino detectors — •VSEVOLOD OREKHOV and MICHAEL WURM — Institute of Physics and Cluster of Excellence PRISMA+, JGU Mainz, Germany

A galactic Supernova explosion is a unique neutrino source: detecting the neutrinos from deep inside the star will help us understand both the physics of the core collapse and properties of the neutrino themselves. If a SN neutrino burst arrived at Earth today or in the near future, it would be detected by a variety of ton to kiloton scale neutrino detectors based on different technologies and target media. By combining the analysis of the possible explosion in multiple next generation neutrino experiments, one could significantly improve the precision of determining the neutrino spectra parameters such as the mean energy and spectral index. In this contribution it is shown what one could achieve by doing a simultaneous fit of the energy spectra of JUNO, DUNE and LecCube-like detectors assuming a common flavour-dependent neutrino signal. This work was supported by funds of the DFG.

T 103: Neutrino Physics without Accelerators 8

Time: Thursday 16:15-18:30

T 103.1 Thu 16:15 T-H35

Status of tau appearance sensitivities and measurements with KM3NeT/ORCA — •NICOLE GEISSELBRECHT for the ANTARES-KM3NET-ERLANGEN-Collaboration — Friedrich-Alexander-Universität Erlangen-Nürnberg, ECAP

KM3NeT/ORCA is a water Cherenkov detector currently under construction in the Mediterranean Sea. The primary goals are the study of neutrino oscillations and the determination of the neutrino mass ordering. One of the main physics objectives in the early phase of ORCA is a measurement of the appearance of tau neutrinos.

Tau neutrinos detected by ORCA are a pure product of neutrino oscillations since ORCA is optimized for atmospheric neutrinos that are almost exclusively produced as electron and muon neutrinos. Even though they can only be measured indirectly as an excess of shower-like events compared to the nonoscillation scenario, their detection will offer valuable information.

One of the first requirements for this study is a reliable particle identification (PID). Therefore, existing PID algorithms need to be optimized with regard to charged current tau neutrino interactions. This talk will report about the first efforts and results using classical machine learning techniques based on boosted decision trees

Neutrino decoherence effects with KM3NeT — •NADJA LESSING for the ANTARES-KM3NET-ERLANGEN-Collaboration — Friedrich-Alexander-Universität Erlangen-Nürnberg

Quantum decoherence of neutrino states is an effect that is proposed in different theories of quantum gravity. It is envisaged to emerge from interactions of the neutrino as a quantum system with the environment and could modify the probabilities of neutrino flavour oscillations in various ways. Therefore, neutrino telescopes such as KM3NeT, that are sensitive to different flavours and to oscillations in a wide range of neutrino energies, can ideally probe this effect. ORCA and ARCA are water Cherenkov detectors that are currently under construction by the KM3NeT Collaboration in the Mediterranean Sea. While ARCA is primarily designed to detect high energy cosmic neutrinos, ORCA aims at the precise measurement of atmospheric neutrino oscillations. This contribution reports on the decoherence sensitivity for both detectors using a phenomenological model in a three-family framework including matter effects. It is shown that, considering different energy dependencies of the phenomenon, either ORCA or ARCA might be capable of improving current bounds on the strength of decoherence effects.

T 103.3 Thu 16:45 T-H35

GIBUU based neutrino interaction simulations in KM3NeT — •JOHANNES SCHUMANN for the ANTARES-KM3NET-ERLANGEN-Collaboration — Friedrich-Alexander-Universität, Erlangen, Germany

The characteristics of the primary neutrino interaction and the subsequent secondaries determine the reconstruction of the primary neutrino properties in neutrino detection experiments. As part of the evaluation of the detector performances, neutrino interactions are simulated via so-called neutrino generators. In order to reduce the computational complexity, these use different approximations which in turn lead to systematic uncertainties on the science output of the experiments. The use of different neutrino generators can therefore help to understand and reduce the systematic uncertainties associated with the simulation of neutrino interactions. GiBUU is a generator that utilises the Boltzmann-Uehling-Uhlenbeck equation in order to propagate the secondary particles through the nucleus. The KM3BUU software package has been developed to adapt GiBUU simulations to the geometry and data format of the KM3NeT neutrino telescope, which is under construction in the Mediterranean Sea. The current status of KM3BUU and first results obtained with this new software package will be presented.

T 103.4 Thu 17:00 T-H35

Modelling Seasonal Variations of Atmospheric Muon Neutrinos using MCEq — •JAKOB BÖTTCHER, HANNAH ERPENBECK, PHILIPP FÜRST, and CHRISTOPHER WIEBUSCH for the IceCube-Collaboration — RWTH Aachen, III. Physikalisches Institut b

The IceCube Neutrino Observatory measures the flux of atmospheric muon neutrinos with unprecedented statistics. These muon neutrinos are produced in cosmic-ray air-showers in the atmosphere, and their flux depends on meteorological quantities such as air temperature and density. The analysis of the resulting seasonal variations improves the understanding of the production of atmospheric neutrinos and provides a novel method for testing hadronic interaction models for air-showers. This talk compares the results of analysing four years of IceCube neutrino data to predictions using the numeric cascade equation solver MCEq. The predictions are based on detailed daily calculations of the atmospheric neutrino flux. These use latitude and longitude dependent vertical temperature profiles of the atmosphere as provided by the AIRS instrument on NASA's Aqua satellite.

T 103.5 Thu 17:15 T-H35

Neutrino oscillation sensitivity studies with IceCube Upgrade - • MARTIN HA MINH for the IceCube-Collaboration — Technische Universität München IceCube is a kiloton-scale neutrino telescope embedded in the Antarctic ice of the South-Pole and it's equipped with over 5000 optical modules. IceCube has delivered world-leading measurements on the neutrino oscillation parameters and plans to further constrain the parameter space with the IceCube Upgrade. The IceCube Upgrade is an augmentation of the detector consisting of 7 detector additional detector strings with new optional modules to improve the directional resolution and lower the energy threshold of particle detection. The construction of this extension will be started in the coming years. In the past, sensitivity curves on the oscillation parameters were based on assumptions of how we expect future event reconstruction resolutions, as we did not have a fully operational algorithms at that time. Now however we developed a Graph Neural Network-based reconstruction algorithm, which allows us to make more realistic predictions about the oscillation analysis performance. In this work we present projections on the sensitivity of the IceCube Upgrade on neutrino oscillation parameters, such as the mixing angle θ_{23} , the squared mass difference m_{23}^2 , and the ratio of expected and recorded v_{τ} flux.

T 103.6 Thu 17:30 T-H35

Precision self-monitoring calibration module for the IceCube Upgrade — •TOBIAS ANDREAS PERTL and FELIX HENNINGSEN — Technische Universität München

The IceCube observatory is a large-volume neutrino observatory at the geographic South Pole. The IceCube Upgrade aims to improve the low-energy and oscillation physics sensitivities as well as re-calibrate the existing detector. This upgrade consists of seven new densely instrumented strings with various different optical and calibration modules. A novel type of precision optical calibration module – or POCAM – for large-volume detectors has been developed and will be deployed as part of the IceCube Upgrade. We report on the design, calibration and production status.

T 103.7 Thu 17:45 T-H35

Optimization of selection cuts for the directional analysis of sub-MeV solar neutrinos in Borexino — •ANTONIA WESSEL^{1,3}, ALEXANDRE GÖTTEL^{2,3}, SINDHUJHA KUMARAN^{2,3}, LIVIA LUDHOVA^{2,3}, LUCA PELICCI^{2,3}, ÖMER PENEK^{2,3}, and APEKSHA SINGHAL^{2,3} — ¹GSI Helmholtzcentre for Heavy Ion Research, Darmstadt, Germany — ²Forschungszentrum Jülich GmbH, Nuclear Physics Institute IKP-2, Jülich, Germany — ³III. Physikalisches Institut B, RWTH Aachen University, Aachen, Germany

Borexino, a liquid scintillator (LS) detector at the Gran Sasso National Laboratory in Italy, has measured all solar neutrinos thanks to its low energy threshold, high energy resolution, and unprecedented radio-purity. Although LS detectors generally cannot obtain any directional information as opposed to Water Cherenkov detectors, Borexino now achieved the first directional measurement of sub-MeV solar neutrinos. The method is based on the early Cherenkov photons which are emitted before the dominant scintillation light. The first photons of the events are then correlated to the Sun's direction. The resulting angles are summed up to form an angular distribution, which is used to measure the total number of solar neutrinos in the selected energy region of interest. For this analysis, high statistics and an excellent signal-to-background ratio are necessary. These criteria can only be met by optimizing the data selection cuts to find the best combination of the fiducial volume and the definition of the energy interval. This talk will describe the strategy for the selection cut optimization and present the resulting fiducial volume and energy region.

T 103.8 Thu 18:00 T-H35

First directional detection of sub-MeV solar neutrinos in Borexino — •JOHANN MARTYN for the Borexino-Collaboration — Johannes Gutenberg-Universität Mainz

Borexino is a 280 t liquid scintillator detector at the Laboratori Nazionali del Gran Sasso (LNGS), Italy. Its main goal is the precision spectroscopy of solar neutrinos down to energies of 0.19 MeV and for this task it features an unprecedented high radio-purity and a high light yield of ~10000 scintillation photons per 1 MeV deposited energy.

In this talk we present the first measurement of sub-MeV solar neutrinos around the ⁷Be edge, using their associated Cherenkov photons in a liquid scintillation detector. In Borexino electrons with E>0.16 MeV produce Cherenkov photons but the ratio of Cherenkov photons from the neutrino scattered electrons is estimated to be < 0.5% for all PMT hits, so a typical reconstruction of the event direction is not possible. Therefore we look instead at the so called "Correlated and Integrated Directionality" (CID), where the known position of the Sun is correlated with the photon hit direction, given by the reconstructed event vertex, and integrated over all selected events. In this way it is possible to measure an angular distribution that shows the statistical contribution of Cherenkov photons from the neutrino recoil electrons. The number of solar neutrinos is then inferred from the measured angle distribution with probability density functions produced by the Geant4-based Borexino Monte Carlo simulation. This work is supported by the Cluster of Excellence No. 2118 PRISMA+, funded by the German Research Foundation (DFG).

T 103.9 Thu 18:15 T-H35

Analysis strategies used in directional analysis of sub-MeV solar neutrinos in liquid scintillator detector — •APEKSHA SINGHAL^{1,3}, ALEXANDRE GÖTTEL^{1,3}, SINDHUJHA KUMARAN^{1,3}, LIVIA LUDHOVA^{1,3}, LUCA PELICCI^{1,3}, ÖMER PENEK^{1,3}, and ANTONIA WESSEL^{2,3} — ¹Forschungszentrum Jülich GmbH, Nuclear Physics Institute IKP-2, Jülich, Germany — ²GSI Helmholtzcentre for Heavy Ion Reseach, Darmstadt, Germany — ³III. Physikalisches Institut B, RWTH Aachen University, Aachen, Germany

The sub-MeV solar neutrinos are measured in a liquid scintillator detector via their elastic scattering off electrons, which induce isotropically emitted scintillation photons that are detected by PMTs. Borexino, located at the LNGS in Italy, is liquid scintillator detector that performed real time spectroscopy of solar neutrinos from pp chain and CNO fusion cycle of the Sun. For first time, it is possible with Borexino to disentangle sub-MeV solar neutrinos detected in liquid scintillator using few Cherenkov photons emitted at early times, and in direction of scattered electrons with given energy threshold. The directional solar neutrino signal is statistically discriminated from isotropic background events by correlating well known position of the Sun and the direction of the first two time-offlight subtracted hits of each event, with respect to the reconstructed vertex. This results in angular distribution of data, fitted with signal and background distributions from Monte Carlo simulations. This talk will describe analysis strategy used to disentangle sub-MeV solar neutrino signal in data in a liquid scintillator detector. The future scope of this method will also be discussed.

T 104: Search for Dark Matter 5

Location: T-H36

Time: Thursday 16:15-18:05

Group Report T 104.1 Thu 16:15 T-H36 Status of the COSINUS Experiment at Gran Sasso — •MARTIN STAHLBERG for the COSINUS-Collaboration — Max-Planck-Institut für Physik, 80805 München

The upcoming COSINUS experiment aims at clarifying the origin of the modulation signal reported by the DAMA-LIBRA collaboration since many years, which is in strong contrast to null-results from other direct dark matter search experiments. Construction of the COSINUS facility has started at Laboratori Nazionali del Gran Sasso (LNGS) in 2021, and first results are expected in 2024. COSINUS cryogenic NaI calorimeters will feature two-channel readout of heat and scintillation signal, and recoil energy thresholds of only few keV. We will give an overview on the status of the experimental setup, as well as outlook on the physics reach, and present first results from remoTES prototype detector measurements.

T 104.2 Thu 16:35 T-H36

remoTES sensors: Development of a novel detector design for NaI cryogenic calorimeters for the COSINUS dark matter experiment — •MUKUND BHARADWAJ for the COSINUS-Collaboration — Max Planck Institute for Physics, Fohringer Ring 6, Munich, Germany - 80805

The COSINUS experiment is an upcoming low-threshold, cryogenic experiment being setup at the Laboratori Nazionali Del Gran Sasso, Italy. It aims to provide a model independent cross-check of the DAMA/LIBRA claim of a potential darkmatter like modulating signal, the result of which is contrary to data reported by other direct dark matter experiments over the past few decades.

COSINUS utilizes a dual-channel readout system based on transition edge sensors(TESs) that allows for particle discrimination. It consists of ultra-pure scintillating sodium iodide (NaI) crystals enclosed by a silicon detector which function as the phonon and light channels respectively. The physical and chemical properties of NaI prevent the direct deposition of a TES on it's surface. In order to overcome this, a new prototype detector design dubbed **remoTES** has been developed. It utilizes a gold pad coupled to the absorber crystal as the primary interface to transmit the phonon signal to the TES, which is fabricated on a separate wafer substrate. The first preliminary results from above ground R&D measurements are reported in this talk.

T 104.3 Thu 16:50 T-H36

Operation of low threshold cryogenic calorimeters in a dry dilution refrigerator in the COSINUS experiment — •MORITZ KELLERMANN, KARL-HEINZ ACK-ERMANN, HENRIK ANSORGE, MUKUND BHARADWAJ, TORSTEN FRANK, KARO-LINE SCHÄFFNER, ROBERT STADLER, MARTIN STAHLBERG, and VANESSA ZEMA — Max-Planck-Institut für Physik, Föhringer Ring 6, 80805 München

For 25 years the DAMA/LIBRA dark matter search measures an annually modulated signal, using sodium iodide (NaI)-scintillators at room temperature. The COSINUS experiment aims to give a model-independent cross-check of the DAMA/LIBRA results by operating pure NaI absorber crystals as cryogenic scintillating calorimeters at mK-temperatures.

COSINUS will be among the first experiments to operate low-threshold calorimeters in a dry dilution refrigerator, being sensitive to temperature changes on a micro-Kelvin level. A pulse tube cooler will be used to arrive at 3K, trading simplified handling for an increased mechanical vibration noise level in the acoustic frequency range. In order to maintain the thermal stability to operate low-threshold calorimeters, it will be necessary to decouple the detectors from possible noise sources. For COSINUS, a spring-based passive decoupling system is planned and tested using piezo-based accelerometers at room temperature. This talk will focus on the design and test results of the in-house developed decoupling system.

T 104.4 Thu 17:05 T-H36

Direct dark matter search with CRESST-III experiment — •LUCIA CANON-ICA for the CRESST-Collaboration — Max-Planck-Institut für Physik, D-80805 München, Germany

CRESST (Cryogenic Rare Event Search with Superconducting Thermometers) is a direct dark matter search experiment located at the Gran Sasso Underground Laboratory (Italy) that uses scintillating cryogenic calorimeters as a target material for elastic DM-nucleus scattering.

The current phase of the experiment, CRESST-III, is optimized for low-energy nuclear recoil detection. It has reached an unprecedented value of 30 eV for nuclear recoil energy thresholds on a CaWO₄ target, allowing the exploration of low-mass dark matter candidates down to 0.16 GeV/c². At higher masses the sensitivity is currently limited by a rising event rate (from threshold up to few hundreds of eV) from a so-far unknown origin.

Currently dedicated measurements with upgraded detectors (including different target materials) are being performed at the Gran Sasso Underground Laboratory, with the goal of investigating and identifying the origin of the event excess.

In this contribution, the current stage of the CRESST-III experiment, together with the most recent dark matter results, will be presented.

T 104.5 Thu 17:20 T-H36 **Development of a Cryogenic Alpha Screening Facility at TUM** – •ANGELINA KINAST¹, ANDREAS ERB^{1,2}, ANDREAS ERHART¹, FIONA HAMILTON¹, MAR-GARITA KAZNACHEEVA¹, ALEXANDER LANGENKÄMPER¹, TOBIAS ORTMANN¹, LUCA PATTAVINA^{1,3}, WALTER POTZEL¹, JOHANNES ROTHE¹, NICOLE SCHERMER¹, STEFAN SCHÖNERT¹, RAIMUND STRAUSS¹, VICTORIA WAGNER¹, and ALEXANDER WEX¹ – ¹Physik-Department E15, Technische Universität München, D-85748 Garching, Germany – ²Walther-Meißner-Institut für Tieftemperaturforschung, D-85748 Garching, Germany – ³INFN, Laboratori Nazionali del Gran Sasso, I-67100 Assergi, Italy

A precise measurement of the radio-purity levels of the CaWO₄ crystals used for Dark Matter search with the CRESST experiment and CEvNS measurements with the NUCLEUS experiment is fundamental for a better background understanding. The sensitivity of HPGe detectors is not sufficient to measure the excellent radio-purity levels of the CaWO₄ crystals produced in-house at the Technische Universität München (TUM). I report on a cryogenic alpha-screening facility developed at TUM, which is currently being commissioned, and will provide a method to determine the radiopurity of our CaWO₄ crystals by measuring the alpha-decays with high precision in the unique experimental environment of the shallow underground laboratory (UGL) at TUM. The research was supported by the DFG through the Excellence Cluster ORIGINS, the SFB1258 and the BMBF: 05A17WO4 and 05A17VTA.

T 104.6 Thu 17:35 T-H36

Development of a new generation of beaker modules for CRESST — •FIONA HAMILTON¹, GODE ANGLOHER², ANTONIO BENTO², ANNA BERTOLINI², LUCIA CANONICA², NAHUEL FERREIRO², DOMINIK FUCHS², ABHIJIT GARAI², DIETER HAUFF^{1,2,3}, MARGARITA KAZNACHEEVA¹, ANGELINA KINAST¹, ALEXANDER LANGENKÄMPER¹, MICHELE MANCUSO², ATHOY NILIMA², TOBIAS ORTMANN¹, LUCA PATTAVINA¹, FEDERICA PETRICCA², WALTER POTZEL¹, FRANZ PRÖBST², FRANCESCA PUCCI², JOHANNES ROTHE¹, KAROLINE SCHAEFFNER², STEFAN SCHÖNERT¹, MARTIN STAHLBERG², LEO STODOLSKT², RAIMUND STRAUSS¹, and VANESSA ZEMA² — ¹Technische Universität München, Physik Department Lehrstuhl E15, James-Franck-Straße 1, D-85748 Garching — ²Max-Planck-Institut für Physik, Föhringer Ring 6, D-80805 München — ³Universität Tübingen, Physikalisches Institut, Auf der Morgen- stelle 14, D-72076 Tübingen

The CRESST experiment is leading the direct search for nuclear recoils induced by light dark matter. The CRESST "beaker modules", using a silicon beaker as a light detector, provide a complete surface anti-coincidence veto. For the next generation of beaker modules, the target and beaker sizes were scaled down in order to further improve energy resolution, allowing background suppression down to energy thresholds below 100 eV. A status update on the research and development of the new generation of beaker modules at TUM is presented. The research was supported by the DFG through the Excellence Cluster ORIGINS and the SFB1258, and the BMBF: 05A17WO4 and 05A17VTA.

T 104.7 Thu 17:50 T-H36

Investigation of Production Techniques for Sputtered Tungsten Thin Films — •Tobias Ortmann¹, Andreas Erhart¹, Margarita Kaznacheeva¹, Angelina Kinast¹, Alexander Langenkämper¹, Luca Pattavina¹, Walter Potzel¹, Johann Riesch², Johannes Rothe¹, Nicole Schermer¹, Stefan Schönert¹, Raimund Strauss¹, Victoria Wagner¹, and Alexander Wex¹ — ¹Technische Universität München, Physik Department Lehrstuhl E15, James-Franck-Straße 1, D-85748 Garching — ²Max-Planck-Institut für Plasmaphysik, Boltzmannstraße 2, D-85748 Garching bei München

Cryogenic rare event searches like the CRESST and the NUCLEUS experiments use TES (Transition Edge Sensors) as phonon sensors to read out their target crystals. This type of sensors utilizes the superconducting phase transition of tungsten to measure the energy deposited in the absorbers. The most established method of production for these films is electron beam physical vapor deposition. For future large scale production the application of argon DC-magnetron sputtering is investigated in terms of film quality and reproducibility. The most recent results of these investigations are presented. The research was supported by the DFG through the Excellence Cluster ORIGINS and the SFB1258, and the BMBF: 05A17WO4 and 05A17VTA.

T 105: Search for Dark Matter 6

Time: Thursday 16:15-18:45

T 105.1 Thu 16:15 T-H37

The MAgnetized Disk and Mirror Axion eXperiment — •CHRISTOPH KRIEGER for the MADMAX-Collaboration — Institut für Experimentalphysik, Universität Hamburg, Luruper Chaussee 149, 22761 Hamburg

The axion is a viable and natural candidate for (cold) dark matter. The mass range for the discovery of axions, favored by some models, 40 to $400 \,\mu eV$, can be investigated using a dielectric haloscope. The MAgentized Disc and Mirror Axion eXperiment is the first axion haloscope based on this approach, utilizing the axion photon conversion at dielectric surfaces in a strong magnetic field. By combining many surfaces, the conversion can be boosted significantly using constructive interference and resonances.

MADMAX will feature a booster system consisting of up to 80 dielectric discs with more than a meter diameter which can be precisely positioned at cryogenic conditions and inside a 9 T magnetic field created by a superconducting dipole magnet with a large warm bore. To prototype this challenging apparatus, a scaled down version (reduced number of discs with 300 mm diameter) is in development. It will be commissioned first at the Universität Hamburg in a dedicated cryostat. It is planned to conduct a first axion-like particle search utilizing the MORPURGO magnet at CERN.

In this presentation, the concept of MADMAX will be presented and an overview will be given on the status and development of MADMAX and its prototype.

T 105.2 Thu 16:30 T-H37

Measurements of dielectric properties of single crystals of Lanthanum Aluminate (LaAlO3) and Sapphire (Al2O3) for the axion dark matter search experiment, MADMAX — •ERDEM OEZ for the MADMAX-Collaboration — RWTH, Aachen

The magnetized disk and mirror axion (MADMAX) experiment will search for axions as cold dark matter candidate in the range of microwave frequencies from 10 to 100 GHz. Multiple parallel dielectric discs are planned to be used to boost the axion generated RF signal by 3 orders of magnitude compared to a single metal surface of the same area. Precise knowledge of the dielectric loss and the dielectric constant of the disc materials is crucial for understanding the predicted axion signal. At these RF frequencies the measurement of low loss materials are especially challenging. Here we present cryostatic measurements of LaAlO3 and Al2O3 which are the best candidate materials for MADMAX. The measurements were done in the 10 to 40 GHz range using a microwave resonator.

T 105.3 Thu 16:45 T-H37

Dielectric Disk Production for the MADMAX — •DOMINIK BREITMOSER — Institut für Experimentalphysik, Universität Hamburg, Luruper Chaussee 149, 22761 Hamburg

The MAgentized Disk and Mirror Axion eXperiment (MADMAX) is an upcoming experiment to search for dark matter axions in the unexplored mass range of 40-400 μ eV. The QCD axion is a solution to the strong CP problem and simultaneously an excellent cold dark matter candidate. In a strong magnetic field, axion-induced photons would be emitted at dielectric interfaces. MADMAX uses the dielectric haloscope approach to boost such a signal by combining up to 80 dielectric disks with 1.25 m diameter and precisely adjustment of the disk distances. The axion to photon conversion is enhanced through interference and resonance effects.

To reach the required sensitivity the disks need a large dielectric constant whilst having low dielectric losses. Experimental constraints demand a planarity below 10 μ m, surface roughness below 10 μ m and a thickness of 1 mm. A favorable material could be lanthanum aluminate. However the material is only available in the size of 3" wafers. Thus, the disks need to be produced from small tiles which are glued together. This talk presents the method of manufacturing a prototype disk (Ø 300 mm), the studies for optimizing production parameters, and explains the measurement system used for quality control.

T 105.4 Thu 17:00 T-H37

Calibrating a Dielectric Haloscope – •JACOB EGGE for the MADMAX-Collaboration – Institut für Experimentalphysik, Universität Hamburg, Luruper Chaussee149, 22761 Hamburg

Dielectric haloscopes like the MAgnetized Disk and Mirror Axion eXperiment aim to detect axions, a dark matter candidate, from the galactic halo by resonant conversion to photons in a strong magnetic field. A movable stack of parallel dielectric disks can amplify the axion-photon conversion probability by several orders of magnitude. This amplification depends on the spatial variation of the axion, photon, and magnetic field. While the axion and magnetic fields are largely static and homogeneous, the frequency-dependent photon field inside the dielectric stack and thus the overall amplification is difficult to characterize.

In this talk, a new calibration method based on non-resonant perturbation theory is presented. It provides a promising way to experimentally constrain the photon field inside the dielectric stack. By perturbing the position of each dielectric disk and measuring the resulting change in reflectivity, one can infer the photon field configuration at each dielectric interface. This then allows computing the amplified axion-photon conversion probability independent of many parameters like dielectric loss and disk geometry. The validity and feasibility of this method are demonstrated with FEM simulations and first measurements on a 5 disks setup.

T 105.5 Thu 17:15 T-H37

Investigating a symmetric booster setup for MADMAX's dielectric haloscope — •LOLIAN SHTEMBARI for the MADMAX-Collaboration — Max Planck Institute for Physics, Munich, Germany

The MADMAX experiment aims to directly detect galactic dark matter axions using the axion-induced emission of electromagnetic waves from boundaries between materials of different dielectric constants placed in a strong magnetic field. Carefully spacing many dielectric disks, their combined emission can be significantly enhanced (boosted) using constructive interference and resonances. In an attempt to reduce the complexity of the system and to gain an understanding of the flexibility and frequency response of the booster, we investigate the performance of a configuration made up of repeating symmetric sections of dielectric disks.

T 105.6 Thu 17:30 T-H37

Loss mechanisms of the MADMAX minimal booster setup — •ANTONIOS GARDIKIOTIS for the MADMAX-Collaboration — Institut für Experimentalphysik,Universität Hamburg, Luruper Chaussee 149, 22761 Hamburg

The axion is a promising candidate to solve the strong CP problem in the SM of particle physics. Its existence could also explain the observations in dark matter (DM) problem. To date, no axions have been detected. An envisioned technique for axion detection in the mass range of 40-400 μ eV is the dielectric haloscope. The MADMAX (MAgentized Disc and Mirror Axion eXperiment) haloscope employs axion conversion into photons on surfaces within a strong magnetic field.

MADMAX uses a booster system consisting of multiple dielectric discs in front of a metal mirror to enhance the tiny axion converted power. Different booster setups have been built to examine the mechanical feasibility and electromagnetic behaviour. Project200 is a proof of principle setup with only one disc in front of a mirror. This simplified version of the MADMAX booster setup can compare basic radio frequency measurements with corresponding simulations.

In this presentation, the study of disc flatness, tilts of the disks and the antenna properties on a minimal setup that can help to better understand the loss mechanisms of the MADMAX booster setup will be discussed.

T 105.7 Thu 17:45 T-H37

Piezoelectric driven dielectric discs for the MADMAX haloscope - • DAGMAR KREIKEMEYER-LORENZO for the MADMAX-Collaboration — Max Planck Institute for Physics, Munich, Germany

Axions are hypothetical particles conceived to explain the strong CP problem of the Standard Model. Simultaneously, axions are an excellent candidate for cold dark matter. MADMAX (MAgnetized Disk and Mirror Axion eXperiment) aims to detect axions in the mass range between 40 and 400 μ eV. For that, it exploits axion to photon conversion at surfaces with different dielectric constants in the presence of a strong magnetic field. The MADMAX haloscope or booster setup will consist of several parallel and adjustable dielectric discs. The separation between discs will be optimised for the axion mass using very precise piezoelectric motors. Here we present the first investigation of the performance of such piezo motors, developed by the company JPE, under realistic conditions relevant for MADMAX. First results at room temperature and at 4.2 K will be shown.

T 105.8 Thu 18:00 T-H37

MADMAX: Dealing with Motor Failures — •DAVID LEPPLA-WEBER for the MADMAX-Collaboration — Institut für Experimentalphysik, Universität Hamburg, Luruper Chaussee 149, 22761 Hamburg

The **MA**gnetized **D**isk and **M**irror **A**xion **eX**periment is a dielectric haloscope aiming to detect axions from the galactic halo by resonant conversion to photons in a strong magnetic field. It uses a stack of dielectric disks, called booster, that amplifies the axion-photon conversion probability over a significant mass range dependent on the position of the adjustable disks. In the planned prototype, depending on the axion mass, amplifications of a factor of ~ 10⁴ can be achieved. The booster is positioned in a cryostat to achieve the low noise levels necessary for the detection of the tiny power stemming from the axion-photon conversion. In case a repair is needed, the cryostat has to be warmed up and cooled down again resulting in significant shutdown time of the experiment. To explore the case where one disk motor of the MADMAX prototype fails, an investigation for the maximum possible scan range with one disk stuck is presented. It is found that one stuck disk reduces the amplification by no more than 25 % over a range

T 105.10 Thu 18:30 T-H37

of 1.5 GHz. This enables the experiment to continue operating for a significant amount of time without having to interrupt the axion search and warm up the A novel approach to simulate axion-induced electrodynamics utilizing deep booster for repairs. $learning \ techniques \ in \ the \ MADMAX \ experiment - Dominik \ Bergermann,$

T 105.9 Thu 18:15 T-H37 A closed booster prototype for MADMAX: CB100 - •CHANG LEE for the MADMAX-Collaboration - Max Planck Institute for Physics, Munich, Germany

CB-100 is a closed dielectric haloscope that aims to verify the dielectric haloscope concept. We use the setup to understand the reflectivity and thermal noise of a dielectric haloscope, which is a prerequisite for a good calibration. The MADMAX collaboration plans to use the system for its axion-like particle dark matter search using the MORPURGO magnet at CERN. We present the concept, design, and construction of CB-100. We also present the reflectivity measurement at room temperature and 4K.

T 106: Experimental Techniques in Astroparticle Physics 4

Deutschland

thereby assist in its optimisation.

Time: Thursday 16:15-18:30

T 106.1 Thu 16:15 T-H38

Automation of the PMT Acceptance Tests for the IceCube Upgrade mDOMs – •Lasse Halve¹, Hannah Erpenbeck¹, Maja Freienhofer², Konstantin MROZIK², JOËLLE SAVELBERG¹, JOHANNES WERTHEBACH², and CHRISTOPHER WIEBUSCH¹ for the IceCube-Collaboration - ¹RWTH Aachen University - Physics Institute III B, Aachen, Germany - ²Astroparticle Physics WG Rhode, TU Dortmund University, Germany

The IceCube Upgrade will extend the IceCube Neutrino observatory with seven additional cable-strings of instrumentation. More than 400 multiple-PMT Digital Optical Modules (mDOMs), with 24 3" Photomultiplier Tubes (PMTs) each, will be deployed. We are testing more than 10.000 PMTs for compliance with manufacturer specifications before the integration into the final modules. A dedicated software for steering the test facilities at RWTH Aachen University and TU Dortmund University for a fully automated operation and online analysis has been developed. We present the design principles and specific solutions for full automization of the test procedures and analyses of PMT data.

T 106.2 Thu 16:30 T-H38

First Results of the PMT Acceptance Tests for the IceCube Upgrade mDOMs - •JOHANNES WERTHEBACH¹, HANNAH ERPENBECK², MAJA FREIENHOFER¹, LASSE HALVE², KONSTANTIN MROZIK¹, and CHRISTOPHER WIEBUSCH² for the IceCube-Collaboration — ¹Astroparticle Physics WG Rhode, TU Dortmund University, Germany — ²III. Physikalisches Institut B, RWTH Aachen Universitv

For the IceCube Upgrade seven new strings will be deployed in the centre of the IceCube Neutrino Observatory. Each string contains several types of modules and in total more than 400 multiple-PMT Digital Optical Modules (mDOMs), with 24 3" Photomultiplier Tubes (PMTs) each, will be frozen into the glacial ice at the South Pole. Testing these PMTs for compliance with manufacturer specifications is crucial before integration into the final mDOM. Utilizing two test facilities at RWTH Aachen University and TU Dortmund University with fully automated operation allows for mass testing of all PMTs. Here, we present the results of the characterization tests for the first batch of PMTs.

T 106.3 Thu 16:45 T-H38

Acceptance Tests for 10,700 PMTs of the mDOMs of the IceCube Upgrade - •Joëlle Savelberg¹, Hannah Erpenbeck¹, Maja Freienhofer², Lasse HALVE¹, KONSTANTIN MROZIK², JOHANNES WERTHEBACH², and CHRISTOPHER WIEBUSCH¹ for the IceCube-Collaboration — ¹RWTH Aachen University -Physics Institute III B, Aachen, Germany — ²Experimentelle Physik 5, TU Dortmund University

The currently prepared IceCube Upgrade will add seven new detector-strings to the central region of the IceCube detector, with the goal of improving the photondetection and lowering the energy threshold. Part of the new instrumentation are more than 400 multi-PMT Digital Optical Modules (mDOMs), each containing 24 3" Photomultiplier Tubes (PMTs) of type Hamamatsu R15458-20. Prior to the assembly of the mDOMs, the 10,700 required PMTs need to be tested for compliance with set specifications. These tests are carried out at dedicated testing facilities at RWTH Aachen University and TU Dortmund University, with setups that have been optimized for large throughput during the production phase. This talk will focus on the design of these setups and development of optimized testing procedures of PMTs in these large quantities.

T 106.4 Thu 17:00 T-H38 Testing the multi-PMT digital optical modules for IceCube Upgrade •NORA FEIGL for the IceCube-Collaboration — DESY Zeuthen

The IceCube Upgrade will enhance IceCube's capabilities at low and high energies. An important part of the Upgrade is the multi PMT approach: the new optical detector module, the multi-PMT digital optical module (mDOM), promises a large sensitive area, homogeneous solid angle coverage and the possibility of multiplicity triggering within a single module.

TIM GRAULICH, •ALEXANDER JUNG, ANDRZEJ NOVAK, ALI RIAHINIA, and

ALEXANDER SCHMIDT — III. Physikalisches Institut A RWTH Aachen, Aachen,

Promising concepts for the search for axion dark matter are dielectric haloscopes,

such as the MAgnetized Disk and Mirror Axion eXperiment (MADMAX). The

realisation of the experiment strongly depends on an accurate and reliable sim-

ulation leading to ever-increasing demands on computing resources, due to the

complexity of simulations. A proof of principle is presented that modern deep

learning techniques can be used in the simulation of the axion haloscope, and

In the past year the mDOM was tested and characterized to verify the design is up to requirements. During the mDOM Design Verification Test (DVT) phase, all the most basic features of the mDOM mainboard, the PMT bases and the calibration systems for the first DVT modules were tested.

The next step will be the large-scale integrated mDOM Final Acceptance Testing (FAT) to verify the previous measurements and the functionality of all subsystems while undergoing temperature cycles.

In this talk the structure and the current state of the mDOM Testing will be presented. Some results of the Design Verification Tests will be shown as well as a short outlook for the upcoming Final Acceptance Testing.

T 106.5 Thu 17:15 T-H38

Location: T-H38

LOM - A multi-PMT optical sensor for IceCube-Gen2. — MARKUS DITTMER, •BERIT SCHLÜTER, ALEXANDER KAPPES, and Lew CLASSEN for the IceCube-Collaboration — WWU Münster

With a smaller diameter and 4-inch PMTs, the eLongated Optical Module (LOM) combines lessons learned from the development of mDOM and DEgg for IceCube Upgrade with gel pads as a new element for optical coupling. The gel pads are a key component here and offer several advantages over previously used approaches. However, they also pose a challenge for the design of the internal mechanical components and subsequent volume production.

This presentation will provide an overview of the proposed LOM design, gel pad studies, and highlight measures taken to ensure consistent quality of the modules under the harsh conditions in the deep ice at the South Pole.

T 106.6 Thu 17:30 T-H38

Reconstruction of simulated muons in a water basin with a multi-PMT optical module - •FRANCISCO JAVIER VARA CARBONELL, MARTIN ANTO-NIO UNLAND ELORRIETA, MARKUS DITTMER, LEW CLASSEN, and ALEXANDER KAPPES for the IceCube-Collaboration — Institut für Kernphysik, Westfälische Wilhelms-Universität Münster, Münster, Germany

The IceCube detector, currently the largest neutrino detector in the world, is scheduled to undergo two upgrades that will be accompanied by new and improved optical modules. These new modules include the LOM and the mDOM, which feature a larger number of PMTs in a pressure vessel. Compared to the old modules, they have a larger effective photosensitive area and nearly uniform angular coverage. In addition, the sensitive area is now fragmented, resulting in intrinsic directional sensitivity of each module. The zenith angle resolution of both types of modules for atmospheric muons in water was studied with Geant4 simulations, using machine learning for reconstruction.

T 106.7 Thu 17:45 T-H38

Studien zum Zeitverhalten eines Photomultipliers mit COMSOL Multiphysics — • Janis Averbeck, Markus Dittmer, Martin Antonio Unland Elor-RIETA, LEW CLASSEN und ALEXANDER KAPPES für die IceCube-Kollaboration -WWU Münster, Münster, Deutschland

Zur Detektion der charakteristischen Tscherenkow-Strahlung werden beim IceCube-Neutrinoteleskop Photomultiplier eingesetzt. Die Photonen lösen an der Photomultiplier-Kathode durch den photoelektrischen Effekt Photonen aus, die anschließend in einem elektrischen Feld beschleunigt werden und ein Dynodensystem mit steigendem Potential durchlaufen. Dabei werden bei jeder Dynodenkollision mehrere Sekundärelektronen ausgelöst, die letztlich zu einem messbaren Signal führen. Für den Einsatz in IceCube ist u.a. eine sehr genaue Kenntnis der zeitlichen Verzögerung zwischen Auslösung an der Photokathode und Ankunft des Elektronensignals an der Anode (Transit-Time) erforderlich. Die Transit-Time hängt dabei nicht nur von der Größe der Beschleunigungsspannung, sondern auch vom Entstehungsort der Elektronen auf der kugelförmigen Kathode, der Anfangsenergie der Elektronen sowie deren Geschwindigkeitsvektor ab. Um den Einfluss dieser drei Parameter auf die Transit-Time zu untersuchen wurde damit begonnen, einen Photomultiplier in COMSOL Multiphysics nachzubauen. Der Vortrag präsentiert erste Ergebnisse der Studien.

T 106.8 Thu 18:00 T-H38

Studies of the LED emission profile in the mDOM with a Geant4 simulation — •ANNA-SOPHIA TENBRUCK, ALEXANDER KAPPES, MARTIN ANTONIO UN-LAND ELORRIETA, LEW CLASSEN, and CRISTIAN JESÚS LOZANO MARISCAL for the IceCube-Collaboration — Institut für Kernphysik, Westfälische Wilhelms-Universität Münster, Münster

The multi-PMT digital optical modules (mDOMs) used in IceCube Upgrade are not only designed to detect neutrinos at low energies, but are also expected to greatly improve the understanding of the detector by installing calibration devices such as LED flashers. LED flashers are versatile devices successfully used in IceCube to measure the ice properties, sensitivity, and timing of optical modules and their positioning. For this purpose, the emission profile of the LEDs must be accurately characterized after they are installed in an mDOM. To study this, the LEDs were simulated in detail in a Geant4 simulation, and the influence of the other module components on the emission profile as well as systematics resulting from the uncertainties in position and inclination of the LED in the

T 106.9 Thu 18:15 T-H38

The Acoustic Module for the IceCube Upgrade — JÜRGEN BOROWKA, •CHRISTOPH GÜNTHER, DIRK HEINEN, MAVERICK SCHÖNELL, CHRISTOPHER WIEBUSCH, and SIMON ZIERKE for the IceCube-Collaboration — RWTH Aachen University - Physics Institute III B, Aachen, Germany

One major goal of the IceCube Upgrade is improved calibration by deploying additional calibration devices in the center of IceCube. Amongst these devices are ten stand-alone Acoustic Modules, capable of receiving and sending acoustic signals. Additionally, these signals are detected by compact acoustic sensors inside some of the optical sensor modules. The positions of emitters and receivers are determined by means of trilateration of the acoustic propagation times. With this system we aim for the calibration of the detector's geometry with a precision better than a few 10 cm. In view of the future IceCube-Gen2 detector, this system will provide an important proof of principle for the reliable geometry calibration on distance scales of a few hundred meters. The design of the acoustic modules and the status of the development are presented in this talk.

T 107: Data Analysis, Information Technology and Artificial Intelligence 5

module were investigated.

Time: Thursday 16:15-18:15

T 107.1 Thu 16:15 T-H39

Introducing novel order statistic tests based on spacings and their applications — **•**LOLIAN SHTEMBARI — Max Planck Institute for Physics, Munich The use of spacings between ordered real-valued numbers is very useful in many areas of science. In particular, either unnaturally small or large spacings can be a signal of an interesting effect. We introduce new statistical tests based on the observed spacings of ordered data. These statistics are sensitive to detect non-uniformity in random samples, or short-lived features in event time series. Under some conditions, these new test can outperform existing ones, such as the well known Kolmogorov-Smirnov or Anderson-Darling tests, in particular when the number of samples is small and differences occur over a small quantile of the null hypothesis distribution. A detailed description of the test statistics is provided including examples and proposed applications for the analysis of neutrino experiments.

T 107.2 Thu 16:30 T-H39

Cosmic ray composition measurement with Graph Neural Networks in KM3NeT/ORCA – •STEFAN RECK for the ANTARES-KM3NET-ERLANGEN-Collaboration — Friedrich-Alexander-Universität Erlangen-Nürnberg (FAU), ECAP

KM3NeT/ORCA is a water-Cherenkov neutrino detector, currently under construction in the Mediterranean Sea at a depth of 2450 meters. The project's main goal is the determination of the neutrino mass hierarchy by measuring the energy- and zenith-angle-resolved oscillation probabilities of atmospheric neutrinos traversing the Earth. Additionally, the detector observes atmospheric muons, which can be used to study the properties of extensive air showers and cosmic ray particles.

This contribution will present a deep-learning based approach to analyse the signatures of muon bundles traversing the detector using graph convolutional networks. Even though the detector is still in an early stage of construction, this reconstruction can already be used to measure the composition of cosmic ray primary particles.

T 107.3 Thu 16:45 T-H39

Tau neutrino selection with Graph Neural Networks for KM3NeT/ORCA - •LUKAS HENNIG for the ANTARES-KM3NET-ERLANGEN-Collaboration Friedrich-Alexander-Universität Erlangen-Nürnberg (FAU), Erlangen Centre for Astroparticle Physics, Erwin-Rommel-Straße 1, 91058 Erlangen, Germany One of the goals of the KM3NeT collaboration is to constrain the PMNS matrix elements associated with the tau neutrino flavor. The data needed to perform this task is taken by KM3NeT's ORCA detector, a water Cherenkov neutrino detector currently under construction in the deep Mediterranean Sea. To constrain the matrix elements, one needs to measure the tau neutrino flux produced by atmospheric muon and electron neutrinos oscillating into tau neutrinos. Selecting the tau neutrino events from the full neutrino event dataset is a notoriously difficult task because the final states of a tau neutrino interaction on a nucleon or nucleus look very similar to those produced in other CC and NC neutrino interactions. This classification problem is addressed by using Graph Neural Networks, a type of neural network architecture that has shown promising results, e.g., in the related task of jet tagging. This presentation will explain how GNNs are applied to neutrino telescope data and report the first results concerning the classification performance.

T 107.4 Thu 17:00 T-H39

Location: T-H39

Conditional Invertible Neural Network for the inference of properties of ultra-high-energy cosmic ray sources — TERESA BISTER, MARTIN ERD-MANN, NATALIE NAST, JOHANNA SCHAFMEISTER, and •JOSINA SCHULTE — III. Physikalisches Institut A, RWTH Aachen

A core challenge in physics data analyses is to estimate model parameters from corresponding measurements, often without the possibility to formulate an explicit inverse function. In a Bayesian framework, parameters are usually estimated using posterior distributions providing not only point estimates but also enabling the determination of uncertainties and correlations. To assess the posterior distributions, we use a so-called conditional Invertible Neural Network (cINN) which is based on the concept of normalizing flows. We apply this new method to a simulated scenario from astroparticle physics to gain information on the parameters of a source model of ultra-high-energy cosmic rays. The corresponding simulated measurements are the energy spectrum, depth of shower maximum distributions and the arrival directions as measured on Earth with similar statistics as data from the Pierre Auger Observatory. We present and evaluate the performance of the cINN on this scenario.

T 107.5 Thu 17:15 T-H39

Deep-Learning-Based Reconstruction of Cosmic-Ray Masses from Extensive Air Shower Measurements with AugerPrime — Martin Erdmann, Jonas GLOMBITZA, BERENIKA IDASZEK, •NIKLAS LANGNER, and DOMINIK STEINBERG — III. Physikalisches Institut A, RWTH Aachen University

Ultra-high-energy cosmic rays (UHECRs) that penetrate the Earth's atmosphere induce extensive air showers. At the Pierre Auger Observatory, showers are measured from the ground using the fluorescence detector (FD) and the surface detector (SD) consisting of water Cherenkov detectors (WCDs). Currently, the SD is extended with scintillators (SSDs) as part of the AugerPrime upgrade.

Actual measurements of the UHECR mass composition are based on FD observations of the depth of shower maximum X_{max} . Using deep learning, X_{max} was successfully extracted using only the SD, exploiting the full statistics of the observatory. However, the precision of X_{max} as a mass estimator at the event level is limited. The new SD upgrade offers the possibility to measure individual components of the shower, potentially improving the reconstruction of the mass composition.

We introduce our network to extract the properties of air showers by analyzing the signals of water Cherenkov detectors as well as the SSDs. We show that the mass-separation power when using only the observable $X_{\rm max}$ is already fully exploited using only WCDs. Thus, we investigate additional observables, novel network architectures and new reconstruction strategies to increase the mass sensitivity with the combination of WCDs and SSD measurements.

T 107.6 Thu 17:30 T-H39

PID with Recurrent Neural Networks in the ATLAS Transition Radiation Tracker for Run 3 — •LENA HERRMANN, CHRISTIAN GREFE, PHILIP BECHTLE, and KLAUS DESCH — Physikalisches Institut, University of Bonn

The measurement of transition radiation effects by the ATLAS transition radiation tracker (TRT) is a key ingredient to the electron identification, especially at low momenta. A recurrent neural network (NN) was developed to combine hitand track-level information into a single classifier, which significantly improves

Particle Physics Division (T)

the particle identification capabilities provided by the TRT.

Since the gas configuration in the TRT will change for the upcoming Run 3 data taking period, separate RNNs have to be trained. The optimisation and training of the RNN will be presented and differences between the Run 2 and Run 3 networks will be discussed. Furthermore, the RNN response on real data taken during Run 2 will be compared to its performance in simulation.

T 107.7 Thu 17:45 T-H39

Usage of neural networks in photon identification in ATLAS — •FLORIAN KIR-FEL — Physikalisches Institut der Universität Bonn

Precise photon identification is crucial for many ATLAS analyses. Currently, photons are selected using a set of cuts on calorimeter variables which characterise the shape of electromagnetic showers. These cuts were optimized using Monte Carlo simulations of photons and jets. Due to the simulations not being ideal, the selection efficiency must be corrected to match data. However, the measurement technique used to determine the identification efficiency in the data requires the hadronic activity around the photon candidate and the photon identification efficiency to be independent. In this work, neural networks are employed to improve over cut-based photon identification. In addition, they are constrained to keep the classification independent of the isolation using the distance correlation. This allows a simplified setup comparing to alternatives such as the adversarial neural network.

T 107.8 Thu 18:00 T-H39

A super-resolution GAN for photon identification at collider experiments — JOHANNES ERDMANN, FLORIAN MENTZEL, OLAF NACKENHORST, and •AARON VAN DER GRAAF — TU Dortmund University, Department of Physics

Many processes in proton-proton collisions contain prompt photons in the final state, ranging from Standard Model (SM) measurements, such as $H \rightarrow \gamma \gamma$, to searches for physics beyond the SM. In order to measure these processes with a high precision, a good photon identification efficiency while retaining a high background rejection rate is important. Most misidentified photons arise from hadron decays, such as $\pi^0 \rightarrow \gamma \gamma$, which could possibly be rejected better with a higher calorimeter granularity. This motivates the idea of artificially increasing the calorimeter granularity by training a super-resolution Generative Adversarial Network (SRGAN) with simulated low and high resolution calorimeter images. As a proof of concept, mono-energetic photons and neutral pions are simulated in an electromagnetic calorimeter and only the second calorimeter layer is used for the SRGAN training. In this presentation, the used SRGAN model, the results of the SRGAN training and the predicted super-resolution images are presented. It is shown that the predicted super-resolution images contain additional information that increases the pion rejection rate compared to the low resolution images.

T 108: General assembly - Particle Physics Division (for DPG members)

Time: Thursday 19:30-21:00

General Assembly - Mitgliederversammlung

T 109: Invited Talks 4

Time: Friday 11:00-12:30

Invited TalkT 109.1Fri 11:00T-H15Ten years of Higgs boson measurements: what we know and what we don't
know — •CHRISTIAN GREFE — Physikalisches Institut, Universität Bonn

One decade after the discovery of the Higgs boson, the ATLAS and CMS experiments continue to publish more and more precise measurements of the Higgs sector, so far confirming the expectations of the Standard Model. With the start of LHC Run 3 ahead – which will double the existing pp-collision dataset – we will review the current knowledge of the fundamental properties of the Higgs boson.

In addition to the Higgs boson couplings, understanding the width and the CP properties of the Higgs boson is crucial to shed light on the open questions in particle physics: Is there CP violation in the Higgs sector? And are there invisible Higgs boson decays which would allow it to couple to a dark sector beyond the Standard Model? We will review the currently available measurements, their limitations and what to expect from Run 3 and beyond.

Invited Talk

Future of Silicon Tracking Detectors: LHC Upgrades and Beyond — •GEORG STEINBRÜCK — Institut für Experimentalphysik, Universität Hamburg In this presentation I will review silicon detector technologies for particle track-

T 109.2 Fri 11:30 T-H15

ing in current and future hadron collider experiments. Significant improvements were needed to reach the requirements for the Phase

2 upgrades of the LHC experiments. I will report on these challenges and the technological solutions with a focus on the upgrades of the CMS and ATLAS tracking detectors. While the collaborations are moving towards production, they are already looking into the future. An overview of further upgrade ideas for the LHC tracking detectors will be given.

The demands on silicon detectors for planned future hadron colliders are even

Location: T-H15

Location: T-MV

more extreme with 1 MeV neutron equivalent fluences up to 10^{17} cm⁻² and huge particle rates. The limiting factors for the use of silicon will be discussed, as well as what is relevant in extrapolating current technologies to the future.

Invited TalkT 109.3Fri 12:00T-H15The dawn of high energy neutrino astronomy- ELISARESCONITechnical University of Munich, Department of Physics, James-Franck-Straße 1, 85748Garching bei MünchenGarching bei MünchenGarching bei München

Current knowledge of the Universe is based on information carried by electromagnetic radiation, gravitational waves, neutrinos, and cosmic rays. For over a century, scientists have observed cosmic rays, but the understanding of their place of production is limited. As a product of cosmic ray interaction, neutrinos can shed light on the extreme part of the Universe. IceCube Neutrino Observatory has been leading neutrino astronomy research over the last ten years and is the only observatory with the exposure to detect high-energy neutrinos beyond Earth*s atmosphere. This presentation will highlight the IceCube observations, including new recent results. Despite the exiting times, with IceCube operating alone and limited by the South Pole location and cubic-km scale, the neutrino astronomy efforts have yet to advance the field past infancy. It is clear that more observatories and larger telescopes, ultimately linked via a global network, are needed to advance fundamental discoveries in astro and particle physics. In this direction, a new opportunity has emerged over the last years to construct a new large volume neutrino telescope, the Pacific Ocean Neutrino Experiment (P-ONE), which will be based on the first time, within an existing oceanographic infrastructure. I will summarize how we have established a scientific relationship with Ocean Networks Canada to pioneer their global network as a testbed infrastructure and identified the optimal location and prepared the ground for first case deployment.

Physics Education Division Fachverband Didaktik der Physik (DD)

Susanne Heinicke Westfälische Wilhelms-Universität Münster Fachbereich Physik Institut für Didaktik der Physik Wilhelm-Klemm-Straße 10 48149 Münster susanne.heinicke@uni-muenster.de

Die Tagung findet als virtuelle Konferenz statt. Die Meetingräume und Poster sind auf der Tagungsplattform verfügbar.

Übersicht über Hauptvorträge, Fachsitzungen und Workshops

(Hörsäle DD-H08, DD-H09, DD-H10, DD-H11 und DD-H12; Poster P)

Hauptvorträge

DD 1.1	Mon	9:00-10:00	DD-H8	Von Naturphänomenen und solchen jenseits unserer Wahrnehmung — • MICHAEL
				Vollmer
DD 21.1	Tue	11:00-11:45	DD-H8	Warum Lehrbuchdarstellungen der Physikgeschichte so schlecht sind – und was wir
				daraus lernen können — • Oliver Passon

Hauptvorträge im gemeinsamen Symposium The Nature of Science (SYNS)

Siehe SYNS für das komplette Programm des Symposiums.

SYNS 1.1	Tue	14:00-14:30	Audimax	The Role of Nature of Science Education for Science Media Literacy – •DIETMAR
				Ноттеске
SYNS 1.2	Tue	14:30-15:00	Audimax	What kinds of identities are deemed in/our of place in physics? — •LUCY AVRAAMI-
				DOU
SYNS 1.3	Tue	15:00-15:30	Audimax	Some thoughts on the status of theoretical physics — •DANIEL HARLOW

Preisträgervorträge im Preisträgersymposium (SYAW) einschließlich Georg-Kerschensteiner-Preisvortrag von Horst Schecker

Siehe SYAW für das komplette Programm des Symposiums.

SYAW 1.1	Wed	14:10-14:40	Audimax	Wie überprüft man die Ziele der Lehramtsausbildung Physik? — •Horst Schecker
SYAW 1.2	Wed	14:40-15:10	Audimax	Astronomy at Highest Angular Resolution - Adaptive Optics, Interferometry and
				Black Holes — •FRANK EISENHAUER
SYAW 1.3	Wed	15:10-15:40	Audimax	Turbulence in one dimension — •ALEXANDER M. POLYAKOV

Gemeinsame Sitzung DD/GP

DD 15.1-15.3	Mon	13:30-15:00	GP-H7	Communicating Physics and its History (joint session GP/DI))
--------------	-----	-------------	-------	--	----

Workshop

DD 51.1 Wed 16:00–17:30 DD-H8 **Konsequenzen aus 3 Jahren Studienreformforschung** — •Stefan Brackertz, Amr El Miniawy, Janette Gehlert, Daniela Kern-Michler, Manuel Längle

Fachsitzungen

DD 1.1–1.1	Mon	9:00-10:00	DD-H8	Hauptvortrag 1
DD 2.1–2.3	Mon	10:15-11:15	DD-H8	Physikdidaktik und Inklusion
DD 3.1-3.3	Mon	10:15-11:15	DD-H9	Neue / digitale Medien – Konzeption
DD 4.1-4.3	Mon	10:15-11:15	DD-H10	Lehr-Lernforschung – Schülervorstellungen Science
DD 5.1–5.3	Mon	10:15-11:15	DD-H11	Hochschuldidaktik – Studieneingangsphase Fachwissen
DD 6.1-6.3	Mon	10:15-11:15	DD-H12	Quantenphysik – Experimente
DD 7.1-7.5	Mon	11:30-12:30	P	Postersession I: Anregungen aus dem Unterricht für den Unterricht
DD 8.1-8.2	Mon	11:30-12:30	P	Postersession 1: Bildung für nachhaltige Entwicklung
DD 9.1-9.1	Mon	11:30-12:30	P	Postersession 1: Geschichte der Physik / Nature of Science
DD 10.1-10.12	Mon	11:30-12:30	P	Postersession 1: Lehr- und Lernforschung
DD 11.1-11.9	Mon	11:30-12:30	P	Postersession 1: Lehreraus- und -weiterbildung
DD 12.1–12.8	Mon	11:30-12:30	P	Postersession I: Neue Konzepte
DD 13.1–13.3	Mon	11:30-12:30	Р	Postersession 1: Physikdidaktik und Inklusion
DD 14.1–14.1	Mon	11:30-12:30	Р	Postersession 1: Sprache und Physikunterricht
DD 15.1–15.3	Mon	13:30-15:00	GP-H7	Communicating Physics and its History (joint session GP/DD)
DD 16.1–16.4	Mon	15:30-16:50	DD-H8	Neue / digitale Medien – Schule
DD 17.1–17.4	Mon	15:30-16:50	DD-H9	Neue / digitale Medien – AR
DD 18.1–18.4	Mon	15:30-16:50	DD-H10	Lehr-Lernforschung – Schülervorstellungen fachbezogen
DD 19.1–19.4	Mon	15:30-16:50	DD-H11	neue Konzepte – Physikunterricht
DD 20.1–20.4	Mon	15:30-16:50	DD-H12	außerschulisches Lernen – Konzepte
DD 21.1–21.1	Tue	11:00-11:45	DD-H8	Hauptvortrag 2
DD 22.1–22.3	Tue	12:00-13:00	DD-H8	Lehreraus- und Lehrerfortbildung – neue Ansätze
DD 23.1–23.3	Tue	12:00-13:00	DD-H9	Neue / digitale Medien – VR
DD 24.1-24.3	Tue	12:00-13:00	DD-H10	Physikdidaktik und Inklusion – Experimentieren
DD 25.1-25.3	Tue	12:00-13:00	DD-H11	Hochschuldidaktik – Studieneingangsphase Studierendenperspektive
DD 26.1-26.3	Tue	12:00-13:00	DD-H12	Quantenphysik – Konzepte
DD 27.1–27.2	Tue	16:15-16:55	DD-H8	außerschulisches Lernen – Metaperspektive
DD 28.1-28.2	Tue	16:15-16:55	DD-H9	Geschichte der Physik und NoS
DD 29.1-29.2	Tue	16:15-16:55	DD-H10	Lehr-Lernforschung – Lernermerkmale
DD 30.1-30.2	Tue	16:15-16:55	DD-H11	Lehreraus- und Lehrerfortbildung – neue Ansätze
DD 31.1-31.2	Tue	16:15-16:55	DD-H12	Praktika und neue Praktikumsversuche
DD 32.1-32.2	Tue	17:00-18:00	Р	Postersession 2: Astronomie
DD 33.1-33.4	Tue	17:00-18:00	Р	Postersession 2: Außerschulisches Lernen
DD 34.1-34.12	Tue	17:00-18:00	Р	Postersession 2: Hochschuldidaktik
DD 35.1-35.13	Tue	17:00-18:00	Р	Postersession 2: Neue / digitale Medien
DD 36.1-36.5	Tue	17:00-18:00	Р	Postersession 2: Praktika und neue Praktikumsversuche
DD 37.1-37.2	Tue	17:00-18:00	Р	Postersession 2: Präsentation von Experimenten
DD 38.1-38.2	Tue	17:00-18:00	Р	Postersession 2: Quantenphysik
DD 39.1-39.3	Tue	17:00-18:00	Р	Postersession 2: Sonstige
DD 40	Tue	18:00-19:30	DD-MV	Mitgliederversammlung des Fachverbands Didaktik der Physik
DD 41.1-41.3	Wed	10:45-11:45	DD-H8	Neue / digitale Medien – Experimente
DD 42.1-42.3	Wed	10:45-11:45	DD-H9	Lehr- und Lernforschung – Repräsentatonsformen
DD 43.1-43.3	Wed	10:45-11:45	DD-H10	BNE – Lernendenperspektive
DD 44.1-44.3	Wed	10:45-11:45	DD-H11	Hochschuldidakitk – neue Konzepte
DD 45.1-45.3	Wed	10:45-11:45	DD-H12	Astronomie
DD 46.1-46.3	Wed	12:00-13:00	DD-H8	BNE - Konzepte
DD 47.1-47.3	Wed	12:00-13:00	DD-H9	Lehreraus- und -weiterbildung – Lehrkonzepte
DD 48.1-48.3	Wed	12:00-13:00	DD-H10	Lehr- und Lernforschung – Methodik
DD 49.1-49.3	Wed	12:00-13:00	DD-H11	Lehreraus- und -weiterbildung – digitale Medien
DD 50.1-50.3	Wed	12:00-13:00	DD-H12	Praktika und neue Praktikumsversuche
DD 51.1-51.1	Wed	16:00-17:30	DD-H8	Workshop: Konsequenzen aus drei Jahren Studienreformforschung

Mitgliederversammlung des Fachverbands Didaktik der Physik

Dienstag 18:00-19:30 DD-MV

- Genehmigung der TagesordnungGenehmigung des Protokolls vom 23.03.2021
- Bericht des Vorstands
- Berichte aus den Arbeitsgruppen
- Termine
- Verschiedenes

Sessions

- Invited Talks, Workshops, Contributed Talks, and Posters -

DD 1: Hauptvortrag 1

Time: Monday 9:00-10:00

Invited Talk

DD 1.1 Mon 9:00 DD-H8 Von Naturphänomenen und solchen jenseits unserer Wahrnehmung — •MICHAEL VOLLMER — University of Applied Sciences Brandenburg Phänomene in unserer Umwelt, insbesondere der Natur, sind besonders geeignet

Interesse für Physik bei Schüler*innen und Student*innen und eigentlich allen an unserer Umwelt Interessierten zu wecken und zu vertiefen. Viele der faszinierendsten Wahrnehmungen gewinnen wir durch unsere Augen, allerdings weisen sie leider auch Beschränkungen auf. Überwinden dieser Begrenzungen führt zu neuen vielfältigen Entdeckungen. Anhand der Variation dreier die Augenwahrnehmung beeinflussenden Variablen (Raum, Zeit und Wellenlänge) wird die große Vielfalt sich neu erschließender Phänomene an ausgewählten Beispielen exemplarisch diskutiert. Physikalisch wird der Bogen gespannt von Mikroskopie zu Satellitentechnik und unserer maximalen Sichtweite, von Zeitraffer zu Highspeedaufnahmen sowie von Röntgen und UV Aufnahmen bis hin zu preiswerten IR Smartphonekameras.

DD 2: Physikdidaktik und Inklusion

Time: Monday 10:15-11:15

DD 2.1 Mon 10:15 DD-H8

Naturwissenschaftlicher Unterricht in Straßenschulen: Zu Lernvoraussetzungen von Straßenjugendlichen — •MATTHIAS FISCHER und MANUE-LA WELZEL-BREUER — Pädagogische Hochschule Heidelberg, Heidelberg, Deutschland

Laut einer Schätzung von Hoch (2017) gibt es in Deutschland etwa 37.000 Straßenjugendliche bis einschließlich 27 Jahre, die entweder wohnungs- oder obdachlos sind. Häufig sind die ausschlaggebenden Ursachen für den Beginn von Straßenkarrieren in ihren Herkunftsfamilien zu finden, in denen sie beispielsweise Gewalt sowie Vernachlässigung erfahren haben. Gleichzeitig birgt auch die Straße als primärer Sozialisationsort vielfältige Herausforderungen. Straßenjugendliche befinden sich somit oftmals in prekären Lebenslagen. So erschwert eine begonnene Straßenkarriere häufig die erfolgreiche Fortführung der eigenen Schulkarriere und folglich finden sich in dieser Gruppe überdurchschnittlich viele Schulabbrüche. Als Abbruchsgründe werden unter anderem Mobbing, fehlende Unterstützung von Lehrkräften und Probleme mit dem Sozialsystem Schule genannt. Die bisherigen Erfahrungen in Herkunftsfamilie, Schule und auf der Straße führen meist zu besonderen Lernvoraussetzungen, die im Regelschulsystem nicht berücksichtigt werden können. Der Aufgabe, Straßenjugendlichen einen Schulabschluss unter Berücksichtigung ihrer Lebenslagen und Lernvoraussetzungen zu ermöglichen, haben sich die sogenannten Straßenschulen verschrieben. In unserer Forschung haben wir mit Interviews untersucht, welche besonderen Lernvoraussetzungen Straßenjugendliche für den naturwissenschaftlichen Unterricht in Straßenschulen mitbringen.

DD 2.2 Mon 10:35 DD-H8

Der Energiebegriff im inklusiven Unterricht — DAVID KOLKENBROCK¹, AN-DREAS KISSENBECK², STEFAN BRACKERTZ¹, RENÉ SCHRÖDER¹ und •ANDREAS $SCHULZ^1 - {}^1$ Universität zu Köln, Köln, Deutschland — 2 Gesamtschule Köln-Holweide, Köln, Deutschland

Energie kann als der Zentralbegriff sowohl im Universum als auch im menschlichen Leben angesehen werden. Daher ist Energie ein vorrangiges Unterrichtsthema zur umfassenden Behandlung gerade am Übergang von der Mittel- hin zur Oberstufe weiterführender Schulen. Hier wird ein Konzept zur Behandlung im inklusiven Unterricht der Klassenstufe 10 entworfen und vorgestellt. Dazu wird für den inklusiven Unterricht eine Differenzierungsmatrix nach Sasse ent-

wickelt, die es den Schüler*innen ermöglicht, beim Lernen am für alle gleichen Gegenstand ihren Schwierigkeitsgrad, aber auch die Art ihres Zugangs (z.B. Einbeziehung von Energieerhaltung) bei den Aufgaben selbst zu wählen. Dabei wird zunächst die Energiebereitstellung thematisiert und besonders auf die Problematik fossiler und nicht-fossiler Energieformen sowie fächerübergreifend auf Energieanwendungen und -Umwandlungen eingegangen und auch das Problem der Energieentwertung behandelt. Besonderer Wert wird auf verantwortungsvollen Umgang mit den verschiedenen Energieformen im Alltag gelegt.

Das Konzept wird in der Gesamtschule Köln-Holweide in 5 UE erprobt und evaluiert. Für die Untersuchung stehen neben der Versuchsgruppe auch Vergleichsgruppen zur Verfügung. Die Durchführung ist zur Tagung abgeschlossen und Ergebnisse werden präsentiert.

DD 2.3 Mon 10:55 DD-H8

Location: DD-H9

Gestaltung von inklusiven Aufgabenformaten am Beispiel des Flaschentutens - •LISA STINKEN-RÖSNER und ELISABETH HOFER — Leuphana Universität Lüneburg

Aufgaben sind ein wesentliches Element im Physikunterricht, da sie Lernende dazu anregen, sich mit Fachinhalten auseinanderzusetzen und ihr Wissen und ihre Fähigkeiten auf neue Kontexte zu übertragen. Klassische textbasierte Aufgabenformate, die auf einen reinen Fachwissenserwerb bzw. dessen Anwendung abzielen, stellen jedoch aufgrund ihrer monotonen Gestaltung für viele Lernende Barrieren dar - was insbesondere aus Sicht der inklusiven Pädagogik problematisch ist (Stinken-Rösner & Abels, 2021). Zudem beschränkt sich ein zeitgemäßer (inklusiver) Physikunterricht nicht einzig auf das Lernen von Fachwissen, sondern verfolgt das Ziel, dass alle Lernenden darüber hinaus an der 'Auseinandersetzung mit naturwiss. Kontexten', dem 'Betreiben von Erkenntnisgewinnung' sowie dem 'Lernen über die Naturwissenschaften' partizipieren können (Hodson, 2014, Stinken-Rösner et al., 2020). Am Beispiel des 'Flaschentutens' werden mögliche Herausforderungen bei der Auseinandersetzung mit einem Kontext theoretisch diskutiert und Möglichkeiten aufgezeigt, wie neue Zugänge durch alternative Aufgabenformate geschaffen werden können (Stinken-Rösner & Hofer, 2022). Im Sinne des Design-Based-Research wurden die verschiedenen Aufgabenformate in drei aufeinanderfolgenden Jahren im Rahmen einer Lehrveranstaltung mit über 100 Studierenden erprobt und kontinuierlich weiterentwickelt.

DD 3: Neue / digitale Medien - Konzeption

Time: Monday 10:15-11:15

DD 3.1 Mon 10:15 DD-H9 Der physikalische Adventskalender "PiA - Physik im Advent" — •ARNULF

QUADT — II. Physikalisches Institut, Georg-August-Universität Göttingen "PiA - Physik im Advent" ist ein physikalischer Adventskalender für Kinder und Jugendliche. Vom 1. bis zum 24. Dezember werden kleine physikalische Experimente mit haushaltsüblichen Materialien per Videofilm vorgeführt. Die TeilnehmerInnen machen sie nach, beobachten das Phänomen, reichen ihre Antwort auf eine gestellt Frage zum Experiment auf der PiA-Webseite ein und erhalten am nächsten Tag die Lösung wieder als Videofilm, evtl. auch einen Punkt. Zu Weihnachten erhalten alle TeilnehmerInnen individuelle Urkunden. Unter den besten TeilnehmerInnen werden Preise in den Kategorien Einzelperson, Schulklasse oder Schule verlost. Im Jahr 2021 hat das Projekt mit rund 65.000 registrierten TeilnehmerInnen und über 2.4 Millionen Besuchen einen neuen Rekord erzielt.

Im Jahr 2022 feiert "PiA - Physik im Advent" 10-jähriges Jubiläum. Das Projekt selbst sowie eine Auswertung der statistischen Daten, konkret das Antwortverhalten differenziert nach Geschlecht, Alter, Bundesland, Land oder anderen Parametern wir der Uhrzeit der Antwort werden vorgestellt und diskutiert.

Location: DD-H8

Location: DD-H10

Location: DD-H11

DD 3.2 Mon 10:35 DD-H9

Eye-Tracking-basierte Gestaltung und Evaluation von Mixed Reality Experimentierumgebungen — •Dörte Sonntag und Oliver Bodensiek — Abt. Physik und Physikdidaktik, Institut für Fachdidaktik der Naturwissenschaften, TU Braunschweig

Mixed Reality (MR) nimmt eine immer prominentere Rolle als Lerntechnologie in den Naturwissenschaften ein, da sie insbesondere beim Experimentieren unterstützen kann. Durch die direkte Integration von virtuellen Elementen in die reale Experimentierumgebung kann das Verständnis der Lernenden verbessert werden. Aktuelle brillenbasierte MR-Hardware bietet ein integriertes Eye-Tracking, das zur Gestaltung und Evaluation von MR-Experimentierumgebungen sinnvoll genutzt werden kann. Basierend auf Eye-Tracking-Daten aus einer realen Experimentierumgebung wird das Vorgehen der Lernenden untersucht und eine MR-Experimentierumgebung kreiert, die insbesondere Novizen derart unterstützen soll, dass sie Problemlösestrategien anwenden, die sonst eher von Experten angewandt werden. Der Erfolg im experimentellen Problemlösen der Probanden in der MR-Umgebung ist dabei signifikant höher als bei Probanden ohne MR und wird von signifikanten visuellen Aufmerksamkeitsverschiebungen begleitet. Weiterhin kann gezeigt werden, dass Novizen mit Unterstützung von MR im experimentellen Problemlösen deutlich strukturierter vorgehen.

DD 3.3 Mon 10:55 DD-H9 Die Rolle räumlicher Kontiguität beim Lernen am Experiment - • PAUL Schlummer¹, Adrian Abazi², Jonas Lauströer³, Rasmus Borkamp³, Reinhard Schulz-Schaeffer³, Wolfram Pernice², Carsten Schuck², Stefan Heusler¹ und Daniel Laumann¹ — ¹Institut für Didaktik der Physik, WWU Münster — ²Center for Nanotechnology, WWU Münster — ³Department Design, HAW Hamburg

Multimediales Lernen wird oft in Zusammenhang mit sogenannten neuen oder digitalen Medien diskutiert und häufig im Rahmen der Cognitive Theory of Multimedia Learning (CTML) modelliert. Diese ist für viele klassische Repräsentationsformen (Text & Bild) in verschiedenen medialen Umsetzungen untersucht worden. Aus physikdidaktischer Sicht decken die im Rahmen der CTML diskutierten Medien- und Repräsentationsformen jedoch nur einen Teil der fachspezifischen Lernsituationen ab und insbesondere die wichtige Rolle von Experimenten als Lernmedium wird meist nicht einbezogen. Der Vortrag präsentiert die Ergebnisse einer Studie, die die Übertragbarkeit des räumlichen Kontiguitätsprinzips multimedialen Lernens der CTML auf Lernsituationen am realen Experiment untersucht. Hierzu wurde eine Augmented-Reality (AR)-Lernumgebung zum Thema Polarisation und Verschränkung entwickelt. Durch Variation der Positionen der zusätzlichen Visualisierungen wurde der Einfluss des Grades an räumlicher Kontiguität zwischen Experiment und Modellvisualisierungen in einer quasi-experimentellen Vergleichsstudie untersucht. Ergebnisse bezüglich Lernzuwachs & kognitiver Belastung werden vorgestellt.

DD 4: Lehr-Lernforschung - Schülervorstellungen Science

Time: Monday 10:15-11:15

DD 4.1 Mon 10:15 DD-H10 Certain about uncertainty - Students' ability to compare data sets - •KAREL

Кок and Burkhard Priemer — Humboldt-Universität zu Berlin Measurement uncertainties are an essential part of a measurement result. It reflects the quality of the data and allows measurement results to be compared. The topic is, however, rarely addressed in secondary school education and students struggle a lot with variance in data sets. The aim of this Ph.D. project was twofold: to find out what conceptual knowledge about measurement uncertainties students need to correctly compare data sets and to see at what grade level the topic can be introduced. To do this, we have developed a Digital Learning Environment (DLE) on the topic of measurement uncertainties. In our study with 154 participants of grades 8 through 11 in Germany, we have used the DLE in a pre-post design.

The pre and post-test consist of a data comparison problem and a competency test. The main focus of the talk will be on the analysis of the justifications from the data comparison problem. We have coded the justifications in terms of the quantity that participants compare and the deciding criterion. The code distributions give a quick and fine-grained overview of what students do when comparing data sets and, hence, indicate their understanding of measurement uncertainties.

Results show that the DLE has a very positive effect on participants' ability to compare data sets and dramatically improves the quality of justifications. Also, we find that the topic can be successfully introduced as early as 8th grade.

DD 4.2 Mon 10:35 DD-H10

Die Förderung des Naturwissenschaftsverständnisses am außerschulischen Lernort - • JESSICA OERTEL und CORNELIA DENZ - Westfälische Wilhelms-Universität Münster, Institut für Angewandte Physik - MExLab Physik, Corrensstr. 2-4, 48149 Münster

Am Experimentierlabor MExLab Physik der WWU Münster werden neben vielfältigen Angeboten für Schüler*innen auch Projektkurse in Kooperation mit Münsteraner Schulen durchgeführt. Das Konzept des Projektkurses bietet die Möglichkeit extracurriculare Themen aus mindestens zwei Fächern für Schüler*innen aufzubereiten und so auch wissenschaftstheoretische Aspekte zu thematisieren. Am Experimentierlabor MExLab Physik wird das Konzept des Projektkurses seit mehreren Jahren erfolgreich durchgeführt und weiterentwickelt. Ziel ist es, die teilnehmenden Schüler*innen durch die Kombination der Fächer Physik und Philosophie bzw. Religion anhand spannender und aus dem Schulunterricht unbekannter Experimente zum kritischen Denken anzuregen und das Wesen der Naturwissenschaften zu entdecken. Der Fokus liegt auf der experiment-basierten Auseinandersetzung mit Fragestellungen aus dem Bereich der Nichtlinearen Physik: Die Phänomene der Strukturbildung und des Chaos bieten besondere Einblicke in den Zusammenhang zwischen Ursache und Wirkung in der Natur und ermöglichen ein vertieftes Verständnis für die Natur der Naturwissenschaft. Mit welchem Verständnis die teilnehmenden Schüler*innen in den Projektkurs einsteigen, wird in mehreren Durchgängen mithilfe von Interviews erhoben und qualitativ ausgewertet

DD 4.3 Mon 10:55 DD-H10 Wissenschaftswerkstatt: Entwicklung von Schülervorstellungen zur Person des Naturwissenschaftlers in der Sekundarstufe 1 - •CLAUDIA HAAGEN-SCHÜTZENHÖFER und THOMAS SCHUBATZKY — Institut für Physik, Universität Graz, Graz, Austria

Verschiedenste Forschungsergebnisse zeigen, dass Schüler:innen stereotype Vorstellungen über Naturwissenschaftler:innen und deren Tätigkeitsfelder haben bzw. entwickeln, die ihr eigenes Verhältnis zu Naturwissenschaften nachhaltig beeinflussen. Das Projekt Wissenschaftswerkstatt der gemeinnützigen Stiftung Kaiserschild, das vom Fachdidaktikzentrum Physik der Universität Graz wissenschaftlich begleitet wird, setzt einen Fokus auf Aspekte des Wissenschaftsverständnisses und dessen Entwicklung bei Schüler:innen der Sekundarstufe 1. Im Projekt werden Mittelschulklassen über 4 Jahre an vier Halbtagen pro Jahr von Lehramtsstudierenden besucht und es wird zu einem MINT Thema gemeinsam gearbeitet und experimentiert. Zudem ist es ein Ziel, die Sichtweise der Schüler:innen auf bestimmte Aspekte von NOS zu schärfen. Die Entwicklung von Schülervorstellungen bezogen auf die Vorstellungen über die Person der/des Wissenschaftler:innen wurde u.a. durch die Draw-a-Scientist Methode erhoben. Hierbei zeigen sich spannende Entwicklungsverläufe über die Projektjahre hinweg.

DD 5: Hochschuldidaktik - Studieneingangsphase Fachwissen

Time: Monday 10:15-11:15

DD 5.1 Mon 10:15 DD-H11 Physikalisches Vorwissen in Physik-Nebenfachveranstaltungen — •KEVIN

SCHMITT und VERENA SPATZ — Technische Universität Darmstadt

Aktuelle Forschungsergebnisse u.a. im Zusammenhang mit den tendenziell hohen Misserfolgs- bzw. Abbruchquoten in Physik belegen immer wieder, dass besonders mathematische aber auch physikspezifische (Vor-)Kenntnisse erheblichen Einfluss auf den Studienerfolg haben können. Dabei konzentrieren sich die bisherigen Erhebungen vorwiegend auf Studierende im Hauptfach (z. B. Müller et. al 2018), während für die Gruppe der Physik-Nebenfachstudierenden, die im Hinblick auf voruniversitäre Bildungsgänge besonders heterogen ist, kaum empirische Erkenntnisse vorliegen.

Vor diesem Hintergrund wurde ein Vorwissenstest entwickelt, um zunächst das physikalische Vorwissen von Studierenden in Physik-Nebenfachveranstaltungen zu untersuchen. Basierend auf der theoretischen Grundlage nach Hailikari, wird das physikalische Wissen dabei in verschiedene Wissensbereiche und Inhaltsfelder segmentiert. Es konnte auf bereits bestehende

195

Messinstrumente (ALSTER-Projekt: Binder et al. 2020) zurückgegriffen werden, die zielgruppengerecht adaptiert wurden.

Der Vorwissenstest wurde im Wintersemester 21/22 in vier Lehrveranstaltungen an der TU Darmstadt pilotiert. Neben einer kurzen Erläuterung der Testkonstruktion werden im Beitrag ausgewählte Ergebnisse der Pilotierung und die daraus folgenden Konsequenzen für die Testkonstruktion präsentiert.

DD 5.2 Mon 10:35 DD-H11

Mindestanforderungskatalog Physik — •Hanno Käss¹, Tilmann Berger², Manuela Boin³, Kim Fujan⁴, Marc Güssmann⁵, Edme Hardy⁶, Flori-MANUELA BOIN", KIM FUJAN', MARC GUSSMANN", EDME HARDY", FLORI-AN KARSTEN⁷, GERRIT NANDI⁸, RONNY NAWRODT⁹, CARSTEN RAUDZIS¹⁰, INA RIECK¹¹, FLORIAN SCHIFFERER¹², STEFAN SCHWARZWÄLDER¹³ und STEFANIE WALZ¹⁴ — ¹Hochschule Esslingen — ²Gymn. Renningen — ³TH Ulm — ⁴Gewerbl. Schule Ehingen — ⁵Lessing-Gymn. Winnenden — ⁶MINT-Kolleg KIT Karlsruhe — ⁷Seminar Stuttgart — ⁸DHBW Heidenheim — ⁹Univ. Stuttgart — ¹⁰Hochschule Reutlingen — ¹¹Grafenbergschule Schorndorf — ¹²Gewerbl. Schule Cömpingen — ¹³Geal Engleg Schule Acateguage — ¹⁴Contraud Ludenge Schule Göppingen — ¹³Carl-Engler-Schule Karlsruhe — ¹⁴Gertrud-Luckner-Gewerbeschule Freiburg

Erstsemester im Bereich der WiMINT-Studiengänge (Wirtschaft, Mathematik, Informatik, Naturwissenschaft, Technik) an den Hochschulen für Angewandte Wissenschaften (HAW) in Baden-Württemberg haben zu Studienbeginn sehr heterogene Kenntnisse in Mathematik und Physik. Dies erschwert den Übergang Schule-Hochschule. Die Arbeitsgruppe cosh Mathematik hat hier 2014 mit ihrem Mindestanforderungskatalog Pionierarbeit geleistet. Er beschreibt mathematische Kenntnisse und Fertigkeiten, die Erstsemester zur erfolgreichen Aufnahme eines WiMINT-Studiums besitzen sollten. In Analogie dazu wurde nun 2019 eine paritätisch aus den Bereichen Schule und Hochschulen zusammengesetzte Arbeitsgruppe cosh Physik gegründet. Die Vorarbeiten der HAWs an einem Mindestanforderungskatalog Physik wurden von ihr weitergeführt und im Oktober 2021 zu einem Abschluss gebracht. Der Beitrag stellt den damit erreichten Stand vor.

DD 5.3 Mon 10:55 DD-H11

Lösungsbeispiele für die Studieneingangsphase Physik – •KAI CARDINAL¹, Andreas Borowski³, Julia Franken², Philipp Schmiemann² und Heike Theyssen¹ — ¹Universität Duisburg-Essen, Didaktik der Physik — ²Universität Dusiburg-Essen, Didaktik der Biologie – ³Universität Potsdam, Didaktik der Physik

In der Studieneingangsphase Physik spielt das fachspezifische Wissen eine zentrale Rolle für den Studienerfolg. Es konnte gezeigt werden, dass in Physik neben dem Konzeptverständnis insb. die Fähigkeit zur Wissensanwendung, d.h. das Finden eines geeigneten Ansatzes und die Ausarbeitung der Lösung unter Nutzung allgemeiner Rechenfähigkeiten, den Studienerfolg vorhersagt. Im Verbundprojekt EASTER (Einfluss der Förderung spezifischer Wissensarten auf Studienerfolg in Biologie und Physik) sollen deshalb diese Fähigkeiten gezielt mit Hilfe von Lösungsbeispielen gefördert werden. Die Lösungsbeispiele orientieren sich strukturell an dem physikalisch-mathematischen Modellierungskreislauf von Trump. Inhaltlich beziehen sie sich auf Themen des ersten Fachsemesters, insb. die Mechanik. Die Lösungsansätze lassen sich den Basiskonzepten aus den Bildungsstandards im Fach Physik für die gymnasiale Oberstufe zuordnen. Im Vortrag wird anhand konkreter Beispiele die systematische Konzeption der Lösungsbeispiele diskutiert.

DD 6: Quantenphysik - Experimente

Time: Monday 10:15-11:15

DD 6.1 Mon 10:15 DD-H12

Wirkung eines quantenphysikalischen Realexperiments auf die physikalische Argumentation — • MORITZ WAITZMANN, RÜDIGER SCHOLZ und SUSANNE WESSNIGK — Leibniz Universität Hannover

Einzelne Photonen erzeugen im Interferometer Interferenzen, sind aber am Strahlteiler nicht teilbar. Dieses Phänomen ist mit klassischer Physik nicht widerspruchsfrei erklärbar. Die Erklärung bedarf quantenphysikalischer Argumente. Lernende verwenden jedoch häufig das semiklassische Argument des Welle-Teilchen Dualismus: Quantenobjekte seien zugleich Welle und Teilchen, erst das Experiment entscheide, welches Merkmal vorliegt. Eine Möglichkeit, die Sichtweise experimentell infrage zu stellen, ist die Betrachtung eines Experiments mit der Kombination eines Einzelphotonen-Strahlteils mit einem Einzelphotonen-Interferometer: Unteilbarkeit und Interferenzfähigkeit sind gleichzeitig beobachtbar. Die notwendige quantentheoretische Erklärung basiert auf drei Grundprinzipien der Quantenphysik: Probabilistik, Superposition und Interferenz (PSI). Inwieweit die Diskussion dieses Experiments und seiner Ergebnisse zu einer Veränderung der Argumentation der Lernenden führt, ist bisher unbekannt. Im Vortrag werden erste Ergebnisse einer Studie mit 80 Studierenden (2. Semester Physik) im Mixed-Methods Design vorgestellt.

DD 6.2 Mon 10:35 DD-H12

Ein spielerischer Einstieg in die Quantenprogrammierung mit QuantumVR — •FRANZISKA GREINERT¹, TOBIAS VOSS², RAINER MÜLLER¹, LINUS KRIEG², GOWTHAM MUTHUSAMY¹, FRANZISKA RÜCKER³ und Klaus Bock-Müller³ — ¹TU Braunschweig, Institut für Fachdidaktik der Naturwissenschaften, Germany — ²TU Braunschweig, Institut für Halbleitertechnik, LENA, Germany — ³SZENARIS GmbH, Bremen, Germany

Im Projekt QuantumVR entwickeln wir ein VR-Spiel für den Einstieg in die gatterbasierte Quantenprogrammierung. Quantencomputing hat in den letzten Jahren deutlich an Bekanntheit gewonnen, auch in der breiten Bevölkerung. Im Rahmen der *Quantum aktiv* Outreach-Initiative sollen Hemmungen gegenLocation: DD-H12

über dieser neuen, scheinbar rätselhaften, vielleicht sogar beängstigenden Technologie abgebaut und Interesse geweckt werden.

Unser Ansatz ist ein Spiel mit Escape-Elementen in virtueller Realität (VR). Kleine Quantenalgorithmen müssen durch Platzieren einfacher Quantengatter (X, H und CX) gelöst werden, um Tiere zu befreien. Eingesetzt werden soll das Spiel bei Events wie Hochschulinformationstagen, aber auch in Workshops mit anschließender Aufbereitung der Inhalte. Die Zielgruppe sind hauptsächlich Schülerinnen und Schüler der Oberstufe, aber auch der Einsatz in der Hochschullehre für den Einstieg in die Quantengatter ist denkbar. Vorgestellt werden das Spiel und erste Erfahrungen aus dem Einsatz.

DD 6.3 Mon 10:55 DD-H12

Location: P

Entwicklung von Analogie-Experimenten zum quantenmechanischen Messprozess — •Stefan Aehle¹, Philipp Scheiger^{1,2} und Holger Cartarius¹ - ¹AG Fachdidaktik der Physik und Astronomie, Friedrich-Schiller-Universität Jena, 07743 Jena — ²Physik und ihre Didaktik, Universität Stuttgart

Trotz steigender Relevanz der Quantenphysik und -technologien für Wirtschaft und Gesellschaft fehlt es noch immer an vielfältigen Lehr-Lern-Materialien zur Unterstützung der Quantum Education - sei es für den schulischen Physikunterricht oder die Hochschullehre. Analogie-Experimente können hilfreich sein, um den nichtklassischen Charakter der Quantenphysik zu veranschaulichen. Im Rahmen dieser Arbeit werden verschiedene Anwendungen entwickelt, die quantenmechanische Phänomene modellieren und dabei aufzeigen, dass Quantenobjekte sich vollkommen anders als klassische verhalten, indem aufgedeckt wird, wie in das Verhalten der klassischen Objekte eingegriffen werden muss, um z.B. die korrekte Statistik einer quantenphysikalischen Messung zu reproduzieren. Ein erster Aufbau aus 3d-Druck, Arduino-Mikrocontrollern und NFC-Chips zur Klärung von Grundbegriffen wie Zustand, Präparation und Messung wird hier vorgestellt und diskutiert. Darauf aufbauend sollen in Zukunft weiterführende Analogversuche zu Themen wie Verschränkung, versteckte Parameter und Quantenkryptographie entstehen.

DD 7: Postersession 1: Anregungen aus dem Unterricht für den Unterricht

Time: Monday 11:30-12:30

DD 7.1 Mon 11:30 P

Nachbau eines Termitenhügels als Projekt zur Wärmelehre - •SARA WILHELM¹ und THOMAS WILHELM² — ¹Universität Würzburg — ²Institut für Didaktik der Physik, Goethe-Universität Frankfurt

Klimawandel und Ressourcenverbrauch stellen die Menschheit vor große technische Herausforderungen. Die Bionik beschäftigt sich damit, die Natur als Vorbild für die Technik zu nehmen. Gerade im Baubereich gilt es, CO2 und Energie einzusparen. Die Baubionik beschreibt das Übertragen von Phänomenen aus der Natur auf technische Funktionen, die für die Architektur und Funktionen eines Gebäudes relevant sein können. Damit liefert sie auch interessante Anwendungen für den Physikunterricht.

Ein bekanntes Beispiel aus der Baubionik ist der Termitenhügel. Die Termiten mögen es in ihrem Bau bei niedriger und vor allem konstanter Temperatur und brauchen im Innern genügend Sauerstoff. Dies erreichen sie u.a. durch einen geschickten Einsatz von Konvektion.

Auf dem Poster wird der Nachbau eines Termitenhügels mit einfachen Mit-

teln gezeigt. Daran können mit einfachen Messungen Inhalte der Wärmelehre veranschaulicht werden. Dazu gehört die Wärmespeicherfähigkeit und Wärmeleitfähigkeit, aber auch Konvektion und der Kamineffekt zur passiven Kühlung. Das Poster zeigt den Nachbau eines Termitenhügels sowie aufgenommene Messwerte und Wärmebilder mit der FLIR ONE-App eines Smartphones.

DD 7.2 Mon 11:30 P

Auseinandersetzungen mit Idealisierungen im Physikunterricht — •FABIAN RAMME¹ und JAN WINKELMANN² — ¹Goethe Universität Frankfurt am Main — ²Pädagogische Hochschule Schwäbisch Gmünd

Im Physikunterricht liegen Idealisierungen stets dem Modellieren und Experimentieren zugrunde. Inwiefern Lehrkräfte die Bedeutung von Idealisierungen in ihrem Unterricht zu einem expliziten Lerngegenstand machen, ist weitgehend unklar. Mithilfe leitfadengestützter Interviews (n = vier Physiklehrkräfte) wurden unter Zuhilfenahme der qualitativen Inhaltsanalyse Kategorien identifiziert, die die unterrichtspraktische Auseinandersetzung mit Idealisierungen beschreiben.

Es zeigt sich, dass Idealisierungen bei den interviewten Lehrkräften meist einen impliziten Lerngegenstand darstellen. Darüber hinaus wurden sehr individuelle Herangehensweisen durch die Lehrkräfte berichtet, u.a. plastische Veranschaulichung von Idealisierungen, tagesaktuelle Bezüge sowie die historische Darstellung naturwissenschaftlicher Erkenntnisgewinnung.

Gleichzeitig wurden einige Bedenken zur Auseinandersetzung mit Idealisierungen thematisiert. Die Sorge, Irritationen zu stiften oder die Annahme, Schüler*innen seien nicht in der Lage mit Idealisierungen adäquat umzugehen, führen stellenweise dazu, dass diese von Lehrkräften bewusst verborgen werden.

Diese und weitere Ergebnisse der Interview-Studie werden auf dem Poster präsentiert.

DD 7.3 Mon 11:30 P

On Phase Transitions in the Early Universe - HANS-OTTO CARMESIN^{1,2,3} and •PHILIPP SCHÖNEBERG¹ – ¹Gymnasium Athenaeum, Harsefelder Straße 40, 21680 Stade — 2 Studienseminar Stade, Bahnhofstr. 5, 21682 Stade -³Universität Bremen, Fachbereich 1, Postfach 330440, 28334 Bremen

We summarize a covariant quantized field equation for the elements of spacetime. Using that equation and the droplet model, we derive phase transitions in the early universe. Moreover, we compare with three other derivations of these phase transitions using

(1) the van der Waals concept

(2) a connectivity of space concept,

(3) a bose gas.

We report about corresponding lessons and experience in a research club.

See e. g. Carmesin, H.-O. (March 2021): Quanta of Spacetime Explain Observations, Dark Energy, Gravitation and Nonlocality. Berlin: Verlag Dr. Köster, Carmesin, H.-O. (2019): Die Grundschwingungen des Universums - The Cosmic Unification - With 8 Fundamental Solutions based on G, c and h - With Answers to 42 Frequently Asked Questions. Berlin: Verlag Dr. Köster.

DD 7.4 Mon 11:30 P

An Easliy Comprehensible Analysis of the Anthropogenic Climate Change – HANS-OTTO CARMESIN^{1,2,3} and •JANNIS VON BARGEN^{1,4} – ¹Gymnasium Athenaeum, Harsefelder Straße 40, 21680 Stade – ²Studienseminar Stade, Bahnhofstr. 5, 21682 Stade — ³Universität Bremen, Fachbereich 1, Postfach 330440, 28334 Bremen — ⁴Brecht-Schule, Norderstraße 163, 20097 Hamburg In the following article we would like to point out the problem of the drastic increase in temperature due to climate change, and we present it in a comprehensible way. Anthropogenic climate change is a change in global temperature caused solely by human activity.

The starting point for our calculations is the worldwide average emission of carbon dioxide, which is 4.8 tons per person per year. To solve the problems of prediction, forecast, responsibility and traceability we use the Stefan-Boltzmann law, a power law describing the absorption, descriptions of other greenhouse gases and corresponding calculations, models, spreadsheets as well as diagrams.

DD 7.5 Mon 11:30 P

An Observation Station for Geomagnetism and Magnetic Storms - A **Project in a Research Club** — •FLORIAN VON BARGEN^{1,4} and HANS-OTTO CARMESIN^{1,2,3} — ¹Gymnasium Athenaeum, Harsefelder Straße 40, 21680 Stade - ²Studienseminar Stade, Bahnhofstr. 5, 21682 Stade — ³Universität Bremen, Fachbereich 1, Postfach 330440, 28334 Bremen — ⁴Brecht-Schule, Norderstraße 163, 20097 Hamburg

The project deals with the measurement and prediction of solar storms. Solar storms are the resultant phe- nomena of a solar flare. Violent solar storms can cause power outages, paralyze satellites and air traffic. This would lead to unpleasant restrictions in our daily life, and cause high costs.

The project deals with the measurement of these storms by recording magnetic fields with a fluxgate sensor. Based on this, a forecast can then be made regarding frequency and intensity. One sensor is already stationary. It delivers data permanently online in 80cm soil depth. The measurement results of the station are close to those of the metrology institute of the Federal Republic of Germany (PTB) in terms of measurement accuracy or measurement fluctuations. We report about results and experiences.

DD 8: Postersession 1: Bildung für nachhaltige Entwicklung

Time: Monday 11:30-12:30

DD 8.1 Mon 11:30 P

Bestimmung der Reliabilität eines Klimawandel-Konzepttests •Marko Jedamski¹, Carina Wöhlke¹, Rainer Wackermann¹, Thomas Schubatzky², Claudia Haagen-Schützenhöfer², Hannes Kasimir LINDEMANN¹ und KAI CARDINAL¹ — ¹Ruhr-Universität Bochum, AG Didaktik der Physik, Deutschland – ²Karl-Franzens-Universität Graz, Institut für Physik, Fachbereich Physikdidaktik, Österreich

Um den Wissensstand und die Vorstellungen deutschsprachiger Schülerinnen und Schüler zu naturwissenschaftlichen Grundlagen des Klimawandels feststellen zu können, werden geeignete Testinstrumente benötigt. Für die Entwicklung eines solchen, validen Klimawandel-Konzepttests wurden zentrale fachliche Inhalte identifiziert, offene Fragen entwickelt und schließlich Distraktoren und Attraktoren aus über 30 Interviews abgeleitet. Anschließend wurde der Konzepttest bei insgesamt rund 200 Schüler:innen, Studierenden und Expert:innen pilotiert. Als Teil der Pilotierung werden im Beitrag die Ergebnisse der Prä-Post-Erhebung einer Studierendengruppe vorgestellt, deren Ziel es war, die Reliabilität des Testinstruments zu untersuchen. Dabei wurde zur Schätzung der Retest-Reliabilität das Testverfahren an der gleichen Stichprobe zweimal durchgeführt und die Korrelation der Testwerte aus beiden Durchgängen berechnet. Berücksichtigt wurde hierbei, dass allen Testpersonen zwischen erstem und zweitem Testpunkt durch das Nennen von sogenannten Klimafakten in der Vorlesung Lerngelegenheiten angeboten wurden und sich somit systematische Veränderungen der wahren Personenwerte ergaben.

Location: P

DD 8.2 Mon 11:30 P

Closing the science-action gap: Kann ein Schülerlabor zur Verbesserung von Fachwissen und Klimawandel-Bewusstheit beitragen? - • JONATHAN GROT-HAUS, MARKUS ELSHOLZ und THOMAS TREFZGER — Lehrstuhl für Physik und ihre Didaktik, Julius-Maximilians-Universität Würzburg

Der Klimawandel und dessen Folgen sind ein komplexes Thema für Schüler:innen, die früher oder später selbst den Wandel der Gesellschaft gestalten. Von Interesse ist daher, inwiefern sich Fachwissen und Einstellung der Schüler:innen zum Thema durch gezielte Intervention beeinflussen lassen. An der JMU soll ein zweitägiges, fächerübergreifendes Schülerlabor Labs4Future am M!ND-Center entwickelt werden. Zielgruppe ist die 9. Jahrgangsstufe aller Schularten. Das Labor gliedert sich in den Besuch der Wissenschaftsausstellung TouchScience, zwei je vierstündigen Labordurchführungen und eine Vorstellung von Ausbildungs- und Berufsperspektiven, mit denen die Schüler:innen selbst Teil des Wandels sein können. In den Durchführungen sollen Ursachen und Wirkmechanismen des anthropogenen Treibhauseffekts und die Klimamodellierung, sowie Folgen und Wirksamkeit von Klimaschutzmaßnahmen thematisiert werden. Ein Ziel der Untersuchung ist es, herauszufinden, ob sich bezüglich der (Teil-)Merkmale von Fachwissen und Einstellung bestimmte markante Gruppen unter den Schüler:innen identifizieren lassen. Das Angebot soll dann entsprechend angepasst werden, sodass diese Teilgruppen jeweils in ihren Schwächen adressiert werden können. Auf dem Poster werden das Schülerlabor, die Ideen für die Materialien und das Forschungskonzept, vorgestellt.

DD 9: Postersession 1: Geschichte der Physik / Nature of Science

Time: Monday 11:30-12:30

Location: P

DD 9.1 Mon 11:30 P

Vorstellungen von Schüler_innen zur Tätigkeit von Forschenden in der Physik — •MORITZ KRIEGEL und VERENA SPATZ — Technische Universität Darmstadt

Schüler_innen haben ein stark verkürztes Bild von den Arbeitsweisen und Tätigkeiten von Naturwissenschaftler_innen, was zu naiven, stereotypischen Vorstellungen über dieses Berufsfeld führt (Wentorf et al., 2015). Besonders die theoretische Physik spielt bei den Vorstellungen der Lernenden über Physik eine stark untergeordnete Rolle und wurde in den Modellen der empirischen Forschung bisher eher vernachlässigt (Heine & Pospiech, 2015). Diese unvollständigen Vorstellungen über den physikalischen Forschungsalltag können zu unreflektierten Entscheidungen hinsichtlich der späteren Berufswahl führen. Die volitionalen

DD 10: Postersession 1: Lehr- und Lernforschung

Time: Monday 11:30-12:30

DD 10.1 Mon 11:30 P

Physikbezogene Mindsets in der gymnasialen Oberstufe — •LAURA GOLDHORN¹, THOMAS WILHELM¹ und VERENA SPATZ² — ¹Goethe-Universität Frankfurt, Deutschland — ²TU Darmstadt, Deutschland

Als Überzeugung zur Intelligenz wirkt sich das Mindset darauf aus, wie Schüler*innen mit (herausfordernden) Lernsituationen umgehen: während ein Fixed Mindset bei Herausforderungen schnell zum Aufgeben führen kann, lassen sich Schüler*innen mit einem Growth Mindset von schwierigen Aufgaben und Lernsituationen nicht abschrecken. Diese Reaktion ist unabhängig vom jeweiligen Könnens- und Wissensstand der Schüler*innen.

Ob Schüler*innen ein Fixed oder Growth Mindset haben, wird in den meisten Studien über einen allgemeinen Fragebogen bestimmt: wer Aussagen wie *Intelligenz bleibt gleich, auch wenn man viel lernt und übt* ablehnt, hat ein Growth Mindset. In einer Erhebung mit Oberstufenschüler*innen in Hessen (N = 760) wurde zusätzlich zur Überzeugung zur allgemeinen Intelligenz nach den fachbezogenen Überzeugungen zu Begabung und Anstrengung in Physik gefragt, um das Mindset bezogen auf Physik zu untersuchen. Die Ergebnisse dieser Mindset-Erhebung ermöglichen eine erste Untersuchung der fachbezogenen Mindset-Charakteristika, unter anderem den Zusammenhang zwischen physikbezogenen Mindset und der Wahl von Grund- oder Leistungskurs, sowie genderbezogene Unterschiede.

DD 10.2 Mon 11:30 P

Der didaktische Nutzen von Feynman-Diagrammen – •MERTEN DAHLKEMPER^{1,2}, PASCAL KLEIN², ANDREAS MÜLLER³, SASCHA SCHMELING¹ und JEFF WIENER¹ – ¹CERN – ²Universität Göttingen – ³Universität Genf Feynman-Diagramme sind als Repräsentationsform aus der Teilchenphysik nicht wegzudenken. Auch dank der Popularisierung durch Feynman selbst haben diese Diagramme eine große Bekanntheit weit über die wissenschaftliche Community hinaus erlangt. Innerhalb der Physikdidaktik sind in den letzten Jahren Debatten darüber entstanden, ob und wenn ja wie diese spezielle Repräsentationsform zur Vermittlung teilchenphysikalischer Inhalte in der Schule genutzt werden kann. In diesem Projekt wird deshalb mittels Design-Based Research erforscht, welche teilchenphysikalischen Inhalte geeignet sind, um diese mit Feynman-Diagrammen für Oberstufenschüler:innen verständlich zu machen und die eine Einführung dieser Repräsentationsform didaktisch rechtfertigen würden. Das Ziel der Forschung ist zweigeteilt: Einerseits entstehen Lehr-Lernmaterialien, hier im Bereich Teilchenphysik mit dem Ziel des Einsatzes in einem Massive Open Online Course (MOOC). Andererseits trägt die Forschung zu Erkenntnissen über Lehr-Lern-Prozesse, hier die Wahrnehmung von Repräsentationen in der Physik bei.

In diesem Poster werden Ergebnisse von Expert:inneninterviews vorgestellt, die Hinweise darauf geben, welche Themen grundsätzlich geeignet sind, um diese anhand von Feynman-Diagrammen zu erklären sowie ein Ausblick auf die weitere Forschung gegeben.

DD 10.3 Mon 11:30 P

Automatische Textanalyse von Schülerantworten auf offene Physikfragen — •FABIAN KIESER und PETER WULFF — Pädagogische Hochschule Heidelberg Durch Natural Language Processing (NLP), einem Teilbereich der Künstlichen Intelligenz Forschung, ist es Computern möglich, natürliche Sprache zu analysieren. Zahlreiche, aus dem Alltag bekannte Anwendungen, wie den Spamfilter der E-Mail Programme, Übersetzungssoftware, Chatbots oder persönliche Sprachassistenten auf verschiedenen technischen Geräten, greifen auf Methoden des NLP zurück, um natürliche Sprache automatisch zu verarbeiten. Um Laufbahnentscheidungen für oder gegen die Naturwissenschaften hängen dabei zudem auch vom Interesse und den Selbstwirksamkeitserwartungen der Lernenden in diesem Bereich ab (Taskinen, 2010).

Daher werden in einem Forschungsprojekt zunächst die Tätigkeiten von Physiker_innen in einem Sonderforschungsbereich der Kern- und Astrophysik mittels Interview- und Fragebogenstudie differenziert erfasst, um auf dieser Grundlage das RIASEC+N Modell für diesen Bereich zu spezifizieren. Mit dieser theoretischen Fundierung soll anschließend eine Projektwoche konzipiert werden, welche die unzureichenden Vorstellungen der Schüler_innen hierzu adressiert und ein ganzheitliches Bild des Spektrums an Tätigkeiten vermittelt. Auf dem Poster werden das Forschungsdesign sowie erste Ergebnisse der Befragung von Forschenden des SFB vorgestellt.

Location: P

eine maschinelle Analyse möglich zu machen, werden Dokumente typischerweise in einen Vektorraum abgebildet, sodass diese geclustert oder klassifiziert werden können. Diese Verfahren bieten auch für die Physikdidaktik besondere Potentiale. Der vorliegende Beitrag prüft, welche Möglichkeiten ausgewählte Methoden der maschinellen Textanalyse bieten, um Schülerantworten auf offene Physikfragen zu analysieren. Insbesondere wird überprüft, inwieweit die automatische Identifikation und Klassifikation mit Erkenntnissen der Physikdidaktik erklärbar sind. Auf dem Poster werden die Umsetzung der einzelnen Methoden, sowie deren Ergebnisse vorgestellt. Darüber hinaus werden Möglichkeiten für den Einsatz im schulischen Kontext diskutiert.

DD 10.4 Mon 11:30 P

Einfluss biologischer und technischer Kontexte auf das situationale Interesse: Ein empirischer Vergleich — •JOHANNES LEWING, PASCAL KLEIN und SUSANNE Schneider — Universität Göttingen

Zur Förderung des Interesses an physikalischen Inhalten wird zunehmend auf kontextorientierten Unterricht gesetzt. Durch den Bezug zur Lebenswelt soll die persönliche Relevanz fachlicher Inhalte für Schülerinnen und Schüler erfahrbar gemacht werden. Der positive Effekt kontextorientierten Lernens auf affektive Schülermerkmale konnte im Vergleich zu traditionellen Unterrichtsansätzen mehrfach bestätigt werden. Weniger erforscht ist die Frage, inwiefern sich die spezifische Wahl eines Kontexts auf das Interesse beim Lernen auswirkt und welche Kontext- bzw. Personenmerkmale dies beeinflussen. Zur Beantwortung dieser Frage, wurde eine Studie mit N = 315 Schülerinnen und Schülern durchgeführt, in der Lernaufgaben zum Basiskonzept Energie in technischen und biologischen Kontexten bearbeitet werden. Es zeigt sich, dass die Einbettung der physikalischen Inhalte in einen Themenbereich, für den bereits individuelles Interesse vorliegt, besonders die emotionale Komponente des situationalen Interesses fördert. Aber auch die für die stabile Interessenentwicklung entscheidende wertbezogene Komponente kann durch den passenden Kontext positiv beeinflusst werden. Im Beitrag werden Implikationen für die Schulpraxis und Forschungsdesiderate diskutiert.

DD 10.5 Mon 11:30 P

Zusammenhänge zwischen dem Blickverhalten und der Antwortsicherheit beim Lösen von Aufgaben zum Graphenverständnis — •HANNA BLUMEN-THAL, SASCHA SCHROEDER und PASCAL KLEIN — Universität Göttingen, Deutschland

Die Antwortsicherheit beim Lösen physikalischer Aufgaben ist ein wichtiger Indikator zur Ermittlung des Lernstands und zur Überprüfung des Verständnisses von Konzepten. So können inkorrekte Antworten, die mit einer hohen Antwortsicherheit gegeben wurden, auf Lernschwierigkeiten oder Missverständnisse hindeuten. Bislang wird die Antwortsicherheit nach Bearbeitung der Aufgabe abgefragt. Erste Eye-Tracking-Studien zeigen, dass Zusammenhänge zwischen der Antwortsicherheit und dem Blickverhalten während des Lösens der Aufgaben bestehen. In diesem Beitrag sollen diese Zusammenhänge anhand von Aufgaben zum Verständnis von Graphen überprüft werden. Es werden die Augenbewegungen von Versuchspersonen beim Lösen entsprechender Aufgaben und die gegebene Antwortsicherheit in einem gemischten linearen Modell in Beziehung gesetzt. Dabei wird zwischen verschiedenen Aufgabentypen unterschieden, welche sich aus der unterschiedlichen Verwendung von Graphen und Text in Frage und Antwort ergeben. Ziel der Untersuchung ist die Bestimmung von Merkmalen des Blickverhaltens, die zur Prädiktion der Antwortsicherheit genutzt werden können.

DD 10.6 Mon 11:30 P

Analyse des Kompetenzaufbaus zur Variablenkontrollstrategie mithilfe von Sankey-Diagrammen — Tobias Winkens, •Simon Goertz und Heidrun Heinke — RWTH Aachen University

Die Variablenkontrollstrategie (VKS) ist eine wichtige experimentelle Kompetenz im naturwissenschaftlichen Unterricht, die Schüler:innen in verschiedenen Lernsettings jedoch häufig falsch anwenden. Im Rahmen einer Design-Based Research Studie im Prä-Post-Design von Goertz wurde unter anderem evaluiert, wie die Kompetenzen der Schüler:innen hinsichtlich der VKS im Schulunterricht verbessert werden können. Mit einem Testinstrument zur Kompetenzerfassung der VKS wurden entsprechende Daten aufgenommen. Als Intervention kamen zwei verschiedene Lernzirkel mit je 5 Stationen (= Modulen) zum Einsatz. Der eine Lernzirkel besteht aus Modulen zu verschiedenen experimentellen Kompetenzen, wobei eines den Schwerpunkt auf die VKS legt, während beim anderen Lernzirkel alle Module die Schwerpunktkompetenz VKS behandeln. Eine detaillierte Auswertung der erhobenen Daten kann durch die Nutzung von Sankey-Diagrammen realisiert werden. Mit ihrer Hilfe kann das Antwortverhalten der Schüler:innen beim Testinstrument im Prä-Post-Vergleich insbesondere im Hinblick auf die beiden Lernzirkeltypen visualisiert und analysiert werden. Die daraus resultierenden Erkenntnisse und Ergebnisse werden auf dem Poster vorgestellt.

DD 10.7 Mon 11:30 P

Learning gains with the Newton's Third Law Open Source Tutorial in Austrian high schools — Iva N Sampaio-Kronister and •Michael M Hull — University of Vienna

I researched the use of the Newton's Third Law Open Source Tutorial in Austrian high schools in order to see if students achieve a better conceptual understanding of the physics concepts, compared to traditional instruction. The research was carried out in nine classes from three different schools (a total of 240 students). Pre-post testing was done with a "Force-Test" that included the Force Concept Inventory's Third Law dimension. All classes had already had their lessons in mechanics by the time of the pre-test. Therefore, the pre-tests results presented a good picture of what students had learned with traditional instruction. Between pre- and post-tests students had their normal classes with traditional instruction, which did not include mechanics, and only one 50-minute intervention with the Open Source Tutorial on Newton's third law. Subsequently they had the post-test, which showed what they learned with the tutorial. The results' analysis shows an evident gain on conceptual understanding of Newton's third law's concepts (g-factor=0,45). The survey also indicated that these concepts actually made sense to the students: many of them had reconciled their intuitive ideas with the correct scientific concepts.

DD 10.8 Mon 11:30 P

The Public's Knowledge on Radioactivity: What effect does the passing of time after graduation and the type of school attended make? — EVA HOLZINGER and •MICHAEL M HULL — University of Vienna

Physics is an unpopular topic in casual conversation today. This may be due to the fact that many adults do not remember any physics that they supposedly learned at school. This paper will address this hypothesis by seeing if time spent since graduating affects adults' understanding of radioactivity, and if school attended makes a difference in this retention. I created an online questionnaire composed of demographic questions and questions to probe understanding and misconceptions about radioactivity. I then collected data with this questionnaire from N = 386 individuals with Austrian school-leaving qualifications. I performed a three-way ANOVA and found that there is a difference in knowledge about radioactivity between recent school leavers and non-recent school leavers, with recent school leavers performing better. Nevertheless, even recent school graduates exhibited the typical misconceptions (they conflated irradiation and contamination, for example), with school attended making no significant difference.

DD 10.9 Mon 11:30 P

Schaltpläne als Repräsentation in der Elektrizitätslehre - eine Schulbuchanalyse — •Stefanie Peter und Olaf Krey — Universität Augsburg

Die Vermittlung grundlegender Konzepte der Elektrizitätslehre stellt eine große Herausforderung dar, wie sich auch daran erkennen lässt, dass auch nachdem die Elektrizitätslehre im Unterricht behandelt wurde, bei Lernenden eine Vielzahl von fachlich unangemessenen Vorstellungen festgestellt werden können. Eine grundlegende Repräsentationsform, die beim Lehren und Lernen der Elektrizitätslehre eine zentrale Rolle spielt, stellen Schaltpläne dar. Der Umgang mit Schaltplänen durch Lernende und Lehrende ist bisher wenig erforscht und auch die Schaltpläne selbst in ihrer Funktion als externale Repräsentation haben bisher wenig Aufmerksamkeit erfahren, was in Anbetracht ihrer Verwendung in Experimentieranleitungen, Sachtexten, Aufgabenstellungen etc. überrascht. Diesem Defizit widmet sich unser Forschungsvorhaben. In einem ersten Schritt werden dabei Schaltpläne als Repräsentationen gegen andere Repräsentationen abgegrenzt und die Verwendung von Schaltplänen in Schulbüchern systematisch erfasst. Aufbauend auf diesen Vorarbeiten soll dann der Umgang mit Schaltplänen insbesondere durch die Lernenden in den Blick genommen werden.

DD 10.10 Mon 11:30 P

Investigating differences in how teachers facilitate the Classbook "The Radiation Around Us" — MARKUS WINTERSTELLER, MAXIMILIAN JEIDLER, and •MICHAEL M HULL — Universität Wien, Wien, Österreich

HEC (Hypothesis Experiment Class) is a teaching method Kiyonobu Itakura first introduced. Similar to a gameshow, students in HEC are asked a series of problems to choose answers a, b or c from. After a discussion phase, the correct answer is revealed. In repeating this procedure, the students are constantly learning more about a specific topic. I am the first teacher to conduct the HEC lesson "The Radiation Around Us" (TRAU) at an Austrian school. My impression is that TRAU has much potential and I am focusing my MS thesis on evaluating this potential. According to Itakura, HEC has three goals: growth in conceptual understanding, enjoyment of the lesson, and reproducibility of the lesson across classrooms. In my thesis, I investigate the last of those three points; namely, I am investigating the effectiveness of TRAU with different teachers with different levels of experience in using HEC. To validate the HEC claims regarding TRAU, all 3 points need to be addressed. For that reason, Maximilian Jeidler is investigating the first two points in his MS thesis.

DD 10.11 Mon 11:30 P

The Influence of "The Radiation Around Us" on Student Conceptual Understanding and Interest — MAXIMILIAN JEIDLER, MARKUS WINTERSTELLER, and •MICHAEL M HULL — Universität Wien, Wien, Österreich

My MS thesis focuses on a special method of teaching called "Hypothesis-Experiment Class" (HEC), developed in Japan by Kiyonobu Itakura. In this method, the teacher is provided with a "Classbook" with which to conduct the lesson. In HEC, students are asked a question and choose between several possible answers. The number of votes for each answer choice is recorded on the blackboard. Individual learners are asked to explain why they chose one of the answer options. Afterwards, all learners have the opportunity to change their choice. Finally, an experiment decides which answer was the correct one. This process is repeated in a cycle with new questions and experiments. HEC has three goals (Itakura, 2019, p. 19-23): 1. make sure each and every student gains the ability to use the central theory or concept. 2. structure the class so that most students report that they like both science and these science classes. 3. make all necessary preparations so that any teacher sufficiently passionate about education, not just special veteran teachers, will be able to teach this type of class. In my work, I will focus on the first and second goals for the Classbook "The Radiation Around Us" (TRAU), which deals with where ionizing radiation can be found in what amounts in our everyday life. The third goal of HEC is discussed in the accompanying poster by Markus Wintersteller.

DD 10.12 Mon 11:30 P

Can focusing on the understanding sub-skill lead to enhanced mastery of the control of variable strategy? — STEPHANIE STEINDL and •MICHAEL M HULL — Universität Wien, Wien, Österreich

Since the mid-20th century, education researchers across disciplines have investigated how to better teach experimental competencies. The control-of-variablestrategy (CVS) is a fundamental part of experimental knowledge acquisition, as it allows the researcher to draw valid conclusions about cause-effect-relationships in experiments where only one variable is changed. Researchers have postulated CVS as being comprised of sub-skills, with the sub-skill "understanding the indeterminacy of confounded experiments (UN)" being the most difficult for students to master. Schwichow et al. (2020) proposed that direct instruction in the UN sub-skill might result in mastery of the UN-skill before or at the same time as the remaining sub-skills. In this presentation, I will discuss my investigation of this proposal, where I measure the effect of different interventions on the overall CVS-skills.

Schwichow, M., Osterhaus, C., & Edelsbrunner, P. A. (2020). The relation between the control-of-variables strategy and content knowledge in physics in secondary school. Contemporary Educational Psychology, 63, 101923. https://doi.org/10.1016/j.cedpsych.2020.101923

DD 11: Postersession 1: Lehreraus- und -weiterbildung

Time: Monday 11:30-12:30

DD 11.1 Mon 11:30 P

Reflexion von Physikunterricht - ein Online Assessment mit Feedback — •ANNA WEISSBACH und CHRISTOPH KULGEMEYER — AG Didaktik, Universität Paderborn, Paderborn, Deutschland

Die Unterrichtsreflexion spielt für Lehrkräfte eine zentrale Rolle. Sie dient der eigenen Professionalisierung und der Weiterentwicklung des Unterrichts. Die Förderung der Reflexionsfähigkeit ist daher auch schon in der ersten Ausbildungsphase wichtig. Ausgehend von einem bestehenden Performanztest zur Reflexion von Physikunterricht wird ein geschlossenes Instrument entwickelt, in welchem Proband*innen inhaltlich zusammengehörige Unterrichtsausschnitte (Videovignetten) aus dem Unterricht eines fiktiven Mitpraktikanten im Sinne einer Fremdreflexion beurteilen. Das Instrument wird mit (teil)automatisiertem Assessment Feedback sowie daran anschließenden Förderempfehlungen versehen. Studierende sollen das Instrument in Eigenregie zur fundierten Selbsteinschätzung dieser Fähigkeit durchführen, Dozierende Rückmeldungen zu ihrer Studierendengruppe erhalten können.

Ziel des Projekts ist die Validierung von Testinstrument und Rückmeldeformaten, sodass sichergestellt wird, dass sie zur Kompetenzentwicklung beitragen. Dazu wird im Sinne des Argument-Based-Approach von Kane (1992) ein Interpretations-Nutzungs-Argument formuliert und u.a. in Bezug auf die Angemessenheit der Testaufgaben, Testscores und resultierenden Konsequenzen evaluiert.

Vorgestellt werden das entwickelte Material (Testinstrument, Rückmeldeformat und Fördermaterial) sowie Ergebnisse einer Think-Aloud-Studie zur Evaluation des Testinstruments.

DD 11.2 Mon 11:30 P Quereinstiegsmasterstudiengänge verstetigen und ausbauen? Befunde und Implikationen aus der Evaluation des Quereinstiegsmasters an der Freien Universität Berlin — •NOVID GHASSEMI und VOLKHARD NORDMEIER — Freie Universität Berlin

Der Mangel an Lehrkräften für unterschiedliche Fächer und Schulformen dauert im Land Berlin weiter an. Demensprechend sieht der neue Koalitionsvertrag von SPD, Grünen und der Linken eine Verstetigung des sog. *Sonderprogramms Beste (Lehrkräfte-)Bildung* und einen Ausbau der Quereinstiegsmasterstudiengänge (Q-Master) vor (spd.berlin 2021). Als Argument für diese Pläne werden positive Erfahrungen genannt. Damit gemeint sind wohl auch vorläufige Evaluationsergebnisse des Q-Masters für die Grundschule (Lucksnat et al. 2021) und für Integrierte Sekundarschulen und Gymnasien (Ghassemi & Nordmeier 2021) im Land Berlin. Nicht ausreichend geklärt ist allerding, welche individuellen Motive und Erwartungen der Q-Masterstudierenden die positiven Befunde erklären könnten und welche zusätzlichen Implikationen und Limitationen sich hieraus ergeben. Um diese Fragen besser beantworten zu können, werden die Q-Masteranden im Fach Physik an der Freien Universität (FU) Berlin nicht allein mittels quantitativer Instrumente untersucht, sondern auch zu ihren Berufsbiografien, ihren Erwartungen an das Studium und dessen subjektiv erlebten Nutzen interviewt. Der Beitrag diskutiert anhand aktueller Ergebnisse der quantitativen und qualitativen Begleitforschung des Q-Masters im Fach Physik an der FU Berlin die Grenzen und Potenziale einer Verstetigung und Ausweitung der Q-Masterstudiengänge.

DD 11.3 Mon 11:30 P

Reflexionsprozesse im Lehr-Lern-Labor — •JENS DAMKÖHLER, THOMAS TREF-ZGER und MARKUS ELSHOLZ — Lehrstuhl für Physik und ihre Didaktik, Julius-Maximilians-Universität Würzburg

Praxisveranstaltungen wird in der Lehramtsausbildung eine bedeutende Rolle zugeschrieben. Eine mögliche Ausgestaltung dieser Praxisphase sind z.B. Lehr-Lern-Labore (LLL) mit iterativen Phasen, bei denen Studierende jeweils Kleingruppen von Schülerinnen und Schülern in zeitlichem Abstand an einem außerschulischen Lernort (M!ND-Center) unterrichten und zwischen den einzelnen Erprobungen Überarbeitungs- und Reflexionsphasen stattfinden. Bisherige Forschungsergebnisse legen nahe, dass ein Lehr-Lern-Labor mit intensiven Phasen der Reflexion die Professionelle Unterrichtswahrnehmung und das Selbstkonzept positiv begünstigen kann. In dem hier vorgestellten Vorhaben sollen die professionellen Reflexionsprozesse von Lehrkräften im LLL-Seminar mit qualitativen und quantitativen Methoden untersucht werden. Insbesondere wird der Frage nachgegangen werden, ob mit Hilfe systematischer Schülerfeedbacks zusätzlich Reflexionsprozesse initiiert werden können. Begleitend ist geplant, das Professionelle Selbstkonzept zu erheben.

DD 11.4 Mon 11:30 P

Digitalität im mathematisch-naturwissenschaftlichen Fachunterricht: Entwicklung und Beforschung einer Masterlehrveranstaltung für die Lehramtsausbildung — •Angelika Mandl, Claudia Haagen-Schützenhöfer, Phil-IPP Spitzer und Thomas Schubatzky — Universität Graz, Österreich Im Rahmen des Projekts ProDigiTrans wird an der Universität Graz im Paradigma des Design-Based Research eine Masterlehrveranstaltung für Lehramtsstudierende mathematisch-naturwissenschaftlicher Unterrichtsfächer entwickelt und beforscht.

Das Ziel ist es, zur Professionalisierung angehender Lehrkräfte beizutragen, sodass diese ihren Fachunterricht digital transformiert umsetzen können. Im Rahmen der Lehrveranstaltung sollen einerseits technisch-pädagogische Kompetenzen, ausgewählt auf Basis einer Curricula-Analyse und von Lehrendenund Studierenden-Befragungen, vermittelt und andererseits ein Verständnis für die Digitalität der Gesellschaft gefördert werden.

Das Poster zeigt die Ergebnisse dieser durchgeführten Curricula-Analyse und Befragungen sowie die daraus resultierenden an die Studierenden adressierten Kompetenzziele. Außerdem werden die Einführung in die digitale Messwerterfassung mit Arduino-Mikrocontrollern und Sensoren als ein inhaltlicher Schwerpunkt der Lehrveranstaltung und die Ergebnisse der dazu durchgeführten Akzeptanzbefragungen vorgestellt. Das auf speziellen Design-Kriterien beruhende Design der Lehrveranstaltung wird präsentiert und es werden Einblicke in die Beforschung des Masterlehrveranstaltungsformates gegeben.

DD 11.5 Mon 11:30 P

Essenzielle Features der Frankfurt/Grazer Optikkonzeption – •MARKUS OB-CZOVSKY, THOMAS SCHUBATZKY und CLAUDIA HAAGEN-SCHÜTZENHÖFER — University of Graz

Eine zentrale Aufgabe physikdidaktischer Forschung ist das Entwickeln innovativer Unterrichtskonzeptionen. Unterrichtsmaterialien bieten eine Möglichkeit diese Innovationen in die Schulpraxis zu bringen. Der Erfolg dieses Transfers ist mitunter abhängig davon, wie Unterrichtskonzeptionen im Unterricht implementiert werden. Wenn Lehrkräfte ihren Unterricht nach einer Unterrichtskonzeption planen, verwenden sie dafür meist beiliegende Unterrichtsmaterialien. Es zeigt sich jedoch, dass Lehrkräfte Unterrichtsmaterialien oft nur bruchstückhaft in ihrem Unterricht einsetzen (Breuer 2021), womit eventuell wesentliche Aspekte der Unterrichtskonzeption verloren gehen. Um Lehrkräfte dabei zu unterstützen, die fachdidaktischen Überlegungen und Leitideen einer Unterrichtskonzeption - welche möglicherweise nicht direkt aus den Unterrichtsmaterialien ersichtlich sind - im Unterricht umzusetzen, wird das Konzept der Essenziellen Features (EF) als fachdidaktische Charakteristika einer Unterrichtskonzeption vorgeschlagen. Eine explizite Kommunikation dieser EF an Lehrkräfte soll diese dabei unterstützen, ihren Unterricht individuell und adressatengerecht zu gestalten und dennoch die fachdidaktischen Grundideen einer Unterrichtskonzeption berücksichtigen zu können. Als Startpunkt wurden exemplarisch die EF der Frankfurt/Grazer Optikkonzeption (Haagen 2016) durch unterschiedliche Expert:innen mithilfe der Unterrichtsmaterialien und zusätzlicher Publikationen ermittelt und werden am Poster vorgestellt.

DD 11.6 Mon 11:30 P

Vignettenstudie zur Perspektive von Physiklehrkräften auf die Lernergebnissicherung im Physikunterricht — •LUCAS CARLOS TELEVANTOS UBEDA, JO-HANNES FRANK LHOTZKY, MARGARETE IMHOF und KLAUS WENDT — Johannes Gutenberg-Universität Mainz

Die Sicherung von Lernergebnissen ist die zentrale Gelingensbedingung für den erfolgreichen Abschluss von Lehr- und Lernprozessen. Trotz dieser zentralen Rolle für das unterrichtliche Handeln liefert die fachdidaktische Literatur kein einheitliches Verständnis darüber, was unter Lernergebnissicherung zu verstehen ist und welche Kriterien an die unterrichtspraktische Umsetzung anzulegen sind. Um die in der Unterrichtspraxis herrschenden Vorstellungen zur Lernergebnissicherung im Physikunterricht zu erfassen, fand für den vorliegenden Beitrag im Rahmen einer Masterarbeit an der JGU Mainz eine Erhebung unter Physiklehrkräften aus dem gesamten deutschsprachigen Raum mittels einer Online-Survey statt. Dazu wurden Unterrichtsvignetten entwickelt, die jeweils den aus der Theorie abgeleiteten Aspekten der Lernergebnissicherung entsprechen und von den Befragten bewertet wurden. Bei der Stichprobengröße der Untersuchung von N=27 fiel in der quantitativen Auswertung besonders die große Streuung in den Bewertungen der Physiklehrkräfte auf. Die Untersuchung liefert somit einen empirischen Beleg für die Verwirrung um den Begriff der Lernergebnissicherung. Das Poster fasst die Ergebnisse zusammen und diskutiert Implikationen für die Lehramts- Aus- und Weiterbildung.

DD 11.7 Mon 11:30 P

CERN Online-Fortbildungsprogramme für Lehrpersonen - **lessons learned** - **•**JEFF WIENER und SASCHA SCHMELING - CERN, Genf, Schweiz

Bereits seit 1998 werden am CERN Fortbildungsprogramme für MINT Lehrpersonen durchgeführt. In den vergangenen Jahren kamen so pro Jahr bis zu 1000 Lehrpersonen aus über 60 verschiedenen Ländern nach Genf, um im Rahmen von 'professional development programmes' ihr fachliches und fachdidaktisches Wissen zur Teilchenphysik aufzufrischen. Durch die COVID-19 Pande-

Location: P

mie konnten diese Programme allerdings seit Frühling 2020 nicht mehr vor Ort stattfinden und wurden stattdessen durch Online-Fortbildungsprogramme ersetzt. Diese wurden an die jeweiligen Anforderungen angepasst und erreichten so im vergangenen Jahr mehr als 2300 Lehrpersonen aus 80 verschiedenen Ländern. Die Bandbreite reicht von kurzen Halbtagesprogrammen hin zu wöchentlichen und sogar mehrmonatigen Programmen, in denen der Austausch der Lehrpersonen untereinander explizit ermöglicht und gefördert wird. Eine essentielle Komponente der Online-Programme stellen Arbeitsaufgaben dar, die von den Lehrpersonen während und nach dem Programm erledigt werden. Diese dienen einerseits der Verdichtung des jeweiligen Programminhalts, stellen andererseits aber auch eine reichhaltige Datenquelle für die Weiterentwicklung der Programme dar. Im Beitrag werden die Designprinzipien und Charakteristika der Programme vorgestellt, sowie relevante fachdidaktische Aspekte diskutiert, die aus der qualitativen Inhaltsanalyse der Arbeitsaufgaben gewonnen wurden.

DD 11.8 Mon 11:30 P Einsatz digitaler Werkzeuge im Physikunterricht - eine Interviewstudie — •DANIEL WALPERT und RITA WODZINSKI — Universität Kassel

Die integrative Vermittlung digitaler Kompetenzen im Regelunterricht setzt bei (angehenden) Lehrkräften technologiebezogene Kompetenzen voraus, welche beispielsweise im TPACK-Modell beschrieben werden. Darüber hinaus ist eine reflektierte Einstellung der Lehrkräfte gegenüber dem Einsatz digitaler Werkzeuge im Unterricht eine wichtige Gelingensbedingung zur Vermittlung digitaler Kompetenzen. Deshalb sollte bereits früh in der (Physik-)Lehramtsausbildung der Erwerb und die vertiefte Auseinandersetzung mit technologiebezogenen Inhalten erfolgen. Ziel des Forschungsvorhabens ist die Entwicklung und Evaluation von Lernarrangements, die eine Förderung technologiebezogener Kompetenzen (TK, TPK, TCK, TPACK) bei angehenden Physik-Lehrkräften erzielen sollen. Im Rahmen der Begleitforschung wird der Einfluss der Teilnahme auf die Einstellungen und die Ausprägung technologiebezogener Wissensfacetten der Studierenden untersucht. Die Erfassung der Einstellungen der Studierenden erfolgt mithilfe teilstrukturierter Interviews im Prä-Post-Design und wird mittels induktiver qualitativer Inhaltsanalyse ausgewertet. Unter den Einstellungen der Studierenden werden folgende Teilfacetten zusammengefasst: Selbstwirksamkeitserwartung, Motivation und wahrgenommene Relevanz zur Vermittlung digitaler Kompetenzen. Es sollen weiterhin Begründungsmuster der Studierenden zum Einsatz digitaler Werkzeuge im Unterricht offengelegt werden.

DD 11.9 Mon 11:30 P

Perceived Agency of In-Service Physics Teachers in Japan and Austria — •MICHAEL M HULL¹ and HARUKO UEMATSU² — ¹Universität Wien, Wien, Österreich — ²Tokyo Gakugei University, Tokyo, Japan

Dissemination of reformed curriculum requires teachers to feel that they have the freedom to implement the curriculum in the classroom. Even instructors who are trained in research-based instruction and are convinced of its value might fail to implement the curriculum in the classroom if, for example, they feel like doing so would jeopardize their ability to cover the contents required by the national standards. We created the "Perceived Agency Survey" to assess teacher views about such issues and administered it to physics teachers in Austria (where teachers are given considerable freedom in their teaching by the national standards) and Japan (which has national standards that are regarded as more demanding of teachers). In this presentation, I will show which items of the survey indicate differences in views between the two groups of teachers, and I will discuss recent interviews with teachers about how Corona has affected their perceived agency.

DD 12: Postersession 1: Neue Konzepte

Time: Monday 11:30-12:30

DD 12.1 Mon 11:30 P

Schülerlabor to go - MINT-Angebote im Freizeitbereich — •MARIA HINKEL-MANN und HEIDRUN HEINKE — RWTH Aachen University, I. Physikalisches Institut IA

Die Städteregion Aachen bietet Schüler:innen verschiedener Jahrgangsstufen bereits eine Vielzahl unterschiedlicher MINT-Angebote, doch haben sich besonders im Nachmittags- und Freizeitbereich weitere Förderbedarfe aufgezeigt. Deshalb soll an der RWTH Aachen University ein Programm entstehen, bei welchem niederschwellige MINT-Kurse im Nachmittagsbereich angeboten werden. Auf Grundlage der vielfältigen Schülerlabor-Angebote in Bereichen wie Physik, Chemie, Biologie oder Informatik sollen vierwöchige Kurse mit jeweils 90minütigen Einheiten entstehen, welche nachmittags zunächst in Schulen umgesetzt werden. Die kurze Laufzeit eines Kurses, die vielfältigen Themen und die Umsetzung an den Schulen sollen das Angebot für die Schüler:innen möglichst leicht zugänglich und für die Schulen flexibel nach Bedarf und Interesse buchbar machen. Die Zielgruppe stellen Jugendliche im Alter von 10 - 16 Jahren dar, wobei ein besonderer Fokus auf die Förderung von jungen Mädchen gelegt wird. Durch den Einsatz im non-formalen Bildungsbereich und in der Freizeit soll besonders die Motivation zur Auseinandersetzung mit MINT-Themen im Alltag gefördert sowie aktiv gegen den akuten Fachkräftemangel vorgegangen werden, der in nahezu allen MINT-Branchen vorzufinden ist. Langfristig sollen die Erfahrungen mit dem Programm auch außerhalb der Städteregion Aachen nutzbar werden und in den Erfahrungsaustausch zwischen MINT-Clustern in Deutschland einfließen.

DD 12.2 Mon 11:30 P

Entwicklung eines Unterrichtsansatzes zum Thema Modellierung und numerische Simulation als naturwissenschaftliche Arbeitsweisen der Erkenntnisgewinnung — •JOHANNA RÄTZ, JAN HEYSEL, INGA WOESTE, VERA MUNZ und FRANK BERTOLDI — Universität Bonn

Modellierung spielt eine zentrale Rolle in den Naturwissenschaften und ist damit auch von hoher Relevanz für gesellschaftspolitische Entscheidungen, z.B. bei der Vorhersage der Entwicklung des Erdklimas oder von Pandemien. Um Schüler*innen durch den Unterricht in ihrer Orientierung, Urteilsfähigkeit und kritischen Teilhabe an einer durch Naturwissenschaften geprägten Gesellschaft zu stärken, ist es unabdingbar die Rolle von Modellen für Erkenntnisprozesse zu thematisieren und explizit zu reflektieren.

Hierzu wurde ein Unterrichtsansatz entwickelt, in dem die Schüler*innen eine Wurfbewegung in einer Sportart ihrer Wahl zunächst modellieren und selber in einer authentischen Programmierumgebung (Jupyter Notebook für Python) numerisch simulieren. Anschließend wird das Ergebnis der Simulation mit einem Real-Experiment verglichen und der Vergleich diskutiert.

Auf diesem Poster wird der Arbeitsstand nach der ersten Erprobung im Rahmen des Projekts "EduChallenge: Perspektiven auf Naturwissenschaften" dargestellt. Location: P

DD 12.3 Mon 11:30 P

Ein dynamisches Elektronenmodell für den Elektrikunterricht — •FABIAN BEIL, MICHAEL THEES, SEBASTIAN KAPP und JOCHEN KUHN — Technische Universität Kaiserslautern, Kaiserslautern, Deutschland

Um im schulischen Elektrikunterricht Präkonzepten der Lernenden zu begegnen, nutzen Lehrende oft Unterrichtskonzepte auf Basis einer Analogie wie z.B. das Wasserkreislaufmodell. Kürzlich wurde von J.P. Burde und T. Wilhelm eine potentialbasierte Analogie entwickelt, das Elektronengasmodell. Diese zeigt eine höhere Wirksamkeit gegen Präkonzepte als traditionelle Unterrichtsansätze. Allerdings erklären analoge Ansätze oft nicht alle Aspekte des Zielbereichs und müssen durch weitere Modelle ergänzt werden. In unserer Arbeit stellen wir ein fachnahes, analogiefreies Modell vor, das kurze transiente Zustände nutzt, um daraus das Potential im Stromkreis abzuleiten. Diese dynamischen Vorgänge auf der Grundlage von Kräften zwischen geladenen Teilchen werden dabei so weit vereinfacht, dass das Modell von Schüler*innen der Mittelstufe genutzt werden kann. Es bleibt dabei anschlussfähig an nachfolgende Themenbereiche, z.B. an die elektromagnetische Induktion, und muss zu keinem Zeitpunkt völlig verworfen werden. In diesem Beitrag wird eine Unterrichtssequenz mit dem dynamischen Elektronenmodell vorgestellt, die als Basis für zukünftige empirische Untersuchungen dienen soll.

DD 12.4 Mon 11:30 P

Entwicklung eines Unterrichtsansatzes zum Thema "Peer-Review-Verfahren als wissenschaftliche Qualitätssicherung" — •INGA WOESTE, JAN HEYSEL, JO-HANNA RÄTZ, VERA MUNZ und FRANK BERTOLDI — Universität Bonn

Das Peer-Review-Verfahren ist eine in der Wissenschaft etablierte Praxis zur Begutachtung von wissenschaftlichen Leistungen und Akteuren durch Fachkolleg*innen. Als Mechanismus der wissenschaftlichen Qualitätssicherung und Kernelement in der Selbststeuerung von Wissenschaft spielt es unter anderem im Publikationswesen, in Berufungsverfahren und bei Verteilungen von Forschungsressourcen eine wichtige Rolle. Im schulischen Kontext wird Peer-Feedback bereits als Methode eingesetzt, bei der sich die Schüler*innen gegenseitig Rückmeldungen zu ihren Arbeitsprodukten geben. Ein Bezug zum Peer-Review-Verfahren als wissenschaftliche Praxis wird hierbei, soweit uns bekannt, jedoch nicht hergestellt. Daher stellt sich die Frage, wie das Peer-Review-Verfahren als Unterrichtsmethode gestaltet und angewendet werden kann, um bei den Schüler*innen ein Verständnis des Konzepts und der Zielsetzung dieser wissenschaftlichen Praxis zu fördern. Im Rahmen des ersten Entwicklungszyklus des Design-Based Research Projekts "EduChallenge: Perspektiven auf Naturwissenschaften" wurde dazu ein Unterrichtsansatz entwickelt und erprobt, bei dem sich die Schüler*innen Hintergründe zum Peer-Review-Verfahren über eine digitale Lernumgebung erarbeiten, in Arbeitsgruppen einen wissenschaftlichen Artikel über ihr Forschungsprojekt verfassen und sich dann dazu im Rahmen eines Double-Blind-Verfahrens Feedback geben.

DD 12.5 Mon 11:30 P

Entwicklung eines Unterrichtsansatzes zur soziologischen Perspektive auf Nature of Science. Der Schulgarten als Analogie zur Wissenschaft. — •VERA MUNZ, JAN HEYSEL, INGA WOESTE, JOHANNA RÄTZ und FRANK BERTOLDI — Universität Bonn

Ein bedeutendes Bildungsziel naturwissenschaftlichen Unterrichts ist es, ein angemessenes Verständnis der Nature of Science (NOS) zu entwickeln. Dies soll den Lernenden ermöglichen, mündig an unserer modernen Informationsgesellschaft teilzuhaben. Wissenschaft als Funktionssystem innerhalb unserer Gesellschaft zu verstehen, ermöglicht es, aus dem gesellschaftlichen Kontext heraus, in dem Wissenschaft eingebettet ist, strukturelle und methodische Praktiken von Wissenschaft zu verstehen. Eine Betrachtung des sozialen Kontextes von Wissenschaft steht daher nicht im Widerspruch zu methodologischen Überlegungen (wie der hypothetisch-deduktiven Methodik), sondern ergänzt diese.

Um Lernenden ein Bild der Naturwissenschaften zugänglich zu machen, welches aus soziologischen Überlegungen entsteht, wurde ein Unterrichtsansatz entwickelt, in welchem ein Schulgarten als Analogie zur Wissenschaft genutzt wird. Die Lernenden planen hierfür ihren eigenen Schulgarten, um sich mit einem möglichen Aufbau eines Funktionssystems vertraut zu machen. Daraufhin wird der selbst entwickelte Schulgarten als Analogie für das System Wissenschaft betrachtet und diskutiert. Dieser Ansatz wurde an zwei Schulen in Nordrhein-Westfalen an Kursen der Einführungsphase erprobt und evaluiert.

DD 12.6 Mon 11:30 P

EduChallenge: Perspektiven auf Naturwissenschaften. Entwicklung eines innovativen Unterrichtskonzepts zum Bereich Nature of Science. – •Jan Heysel¹, Johanna Rätz¹, Inga Woeste¹, Vera Munz¹, Janina Beigel² und Frank Bertoldi¹ – ¹Universität Bonn – ²Universität Heidelberg

Wir sehen Handlungsbedarf aus der Problematik, dass Schülervorstellungen zur *Nature of Science* (NOS) oft sehr undifferenziert sind, dass konkrete Inhalte und Ziele von Unterricht im Bereich NOS unklar und kontrovers bleiben, und dass bislang wenige Konzepte zur expliziten Förderung von Kompetenzen in diesem Bereich existieren.

Auf der Basis der Lerntheorie und der Pädagogik des *Deeper Learning* entwickeln wir den allgemeinen Ansatz einer *EduChallenge* als motivierendes und explizit kompetenzorientiertes Unterrichtsformat. Außerdem entwickeln wir ein Konzept der *Perspektiven auf Naturwissenschaften* als kompetenzorientierten, schulpraktischen Ansatz, der neben methodologischen Aspekten auch historische, soziologische und philosophische Perspektiven umfasst. Unser aktuelles Projekt der *EduChallenge: Perspektiven auf Naturwissenschaften* kombiniert beide allgemeinen Ansätze zu einem konkreten Anwendungsbeispiel. Wir haben kürzlich den ersten Entwicklungszyklus unseres Design-Based Research Projekts mit einer Erprobung an zwei Schulen abgeschlossen und stellen mit diesem Poster den aktuellen Arbeitsstand dar.

DD 12.7 Mon 11:30 P

Didaktische Rekonstruktion der Strahlentherapie — •АХЕL-ТHILO РКОКОР und RONNY NAWRODT — Universität Stuttgart, 5. Physikalisches Institut - Abt. Physik und ihre Didaktik Pfaffenwaldring 57, 70569 Stuttgart

Die Anwendung radioaktiver Materialien zur Behandlung von Krebserkrankungen ist einschließlich typischer Nebenwirkungen landläufig bekannt. Die genaue Wirkungsweise der dabei verwendeten Isotope, die dafür notwendige Verarbeitung des Isotops und die allgemeine Wirkung ionisierender Strahlung auf den menschlichen Körper auf mikroskopischer Ebene sind hingegen nicht allgemein bekannt. Hier bietet sich die Möglichkeit Radioaktivität in Hinblick auf typische Alltagsvorstellungen neu zu denken und die Begriffe "radioaktive Materie" und "ionisierende Strahlung" sachgerecht einzuordnen.

DD 12.8 Mon 11:30 P

Weiterentwicklung eines interaktiven Steckbretts zum Aufbau elektrischer Schaltungen — • MESUT IBRAHIM TASTEKIN, SIMON GOERTZ, SEBASTIAN NELL und HEIDRUN HEINKE — RWTH Aachen University, Deutschland

Um Lehrkräfte bei der Planung und Durchführung von kompetenzorientiertem Unterricht zu unterstützen, werden an der RWTH Aachen University im Rahmen der Plattform *FLexKom* Unterrichtsmaterialien zum Fördern und Lernen experimenteller Kompetenzen entwickelt. Dabei werden Module angeboten, die Teilkompetenzen in den Bereichen Planung, Durchführung und Auswertung von Experimenten fördern sollen. Dieser Beitrag beschreibt die Evaluation und Weiterentwicklung eines interaktiven Steckbretts, das eine direkte Rückmeldung durch einen Mikrocontroller über die Korrektheit einer gesteckten elektrischen Schaltung geben kann. Der Einsatz des Steckbretts fand in physikalischen Praktikumsversuchen von Studierenden unterschiedlicher Studiengänge statt. Die Weiterentwicklung erfolgt mit dem Ziel der detaillierten Nachverfolgung des Steckprozesses der Schaltung. Dies soll die Analyse möglicher experimenteller Schwierigkeiten, typischer Vorgehensweisen und Schülervorstellungen ermöglichen. Auf dem Poster werden das evaluierte Steckbrett sowie die Umsetzung der Weiterentwicklung des Steckbretts beschrieben.

DD 13: Postersession 1: Physikdidaktik und Inklusion

Time: Monday 11:30-12:30

DD 13.1 Mon 11:30 P Zur Rolle von naturwissenschaftlicher Bildung in Straßenschulen — •MATTHIAS FISCHER und MANUELA WELZEL-BREUER — Pädagogische Hochschule Heidelberg, Heidelberg, Deutschland

Im Jahr 2015 wurden die Ziele für nachhaltige Entwicklung von den Vereinten Nationen verabschiedet, zu denen auch die Ziele 'No Poverty' und 'Quality Education' gehören. Das letztgenannte Bildungsziel beinhaltet für die Vereinten Nationen neben dem gleichberechtigten Zugang aller Bevölkerungsgruppen zu den verschiedenen Bildungsebenen auch ein sichergestellter sowie hochwertiger Sekundarschulabschluss aller. Diese beiden Unterziele gelten insbesondere auch für die ungefähr 37.000 in Deutschland lebenden Straßenjugendlichen. Dabei sind Straßenkarrieren und erfolgreiche Schulkarrieren für viele Jugendlichen nur schwer zu realisieren. So hat beispielsweise Hoch (2016) in einer Studie festgestellt, dass die Mehrheit der Straßenjugendlichen im Gegensatz zu den durchschnittlich erreichten Schulabschlüssen in Deutschland entweder gar keinen Schulabschluss oder nur einen Hauptschulabschluss besitzt. Gründe hierfür sind diverse Barrieren und Probleme, denen sich Straßenjugendliche während des Schulbesuchs sowie bei der Wiederaufnahme des Bildungsweges nach erfolgtem Schulabbruch gegenübersehen. In Deutschland gibt es mittlerweile mehrere Straßenschulen, die solchen Jugendlichen unter Berücksichtigung ihrer Lebenslagen das Nachholen eines hochwertigen Sekundarschulabschlusses ermöglichen. In unserer Studie wird untersucht, inwiefern naturwissenschaftliche und im Besonderen physikalische Bildung eine Rolle in den Straßenschulen spielt.

DD 13.2 Mon 11:30 P

MasterClasses — •AZADEH GHANBARI², STINA SCHEER¹, RAINER MÜLLER² und GUNNER FRIEGE¹ — ¹Leibniz Universität Hannover | Institut für Didaktik der Mathematik und Physik | Physikdidaktik — ²Technische Universität Braunschweig | IFDN | Physik und Physikdidaktik

MasterClasses sind halbtägige Workshops zu aktuellen Themen der Quantenmetrologie für Schülerinnen und Schüler der gymnasialen Oberstufe (Einführungs und Qualifikationsphase). Die Lernenden werden auf Grundlage ihrer im Physik und Mathematikunterricht erworbenen Fähigkeiten an Themen der aktuellen Forschung herangeführt. Dabei können sie nicht nur experimentieren und neue Bereiche der Physik kennen lernen, sondern bekommen durch Kontakt mit Forschenden auch einen Einblick in den Wissenschaftsbetrieb. Die MasterClasses sind Teil des Exzellenzclusters QuantumFrontiers.

DD 13.3 Mon 11:30 P

Location: P

Gestaltung inklusiver Lernsettings am Beispiel Getriebe — •LARISSA FÜHNER und Alexander Pusch — Institut für Didaktik der Physik, WWU Münster Im Zuge der Inklusion an Schulen steht die Physikdidaktik vor der aktuellen Herausforderung den regulären Physikunterricht für die Bedürfnisse von Schülerinnen und Schülern mit verschiedenen Beeinträchtigungen des Lernens zu adaptieren und weiterzuentwickeln. Dabei können grundlegende Handlungsempfehlungen aus der sonderpädagogischen Praxisarbeit helfen, die Lernsituationen konkret auszugestalten. Auf diesem Poster stellen wir Teile unseres Seminarkonzepts zu inklusivem Physikunterricht vor. Dort zeigen wir zunächst typische Probleme von lernbeeinträchtigten Schülerinnen und Schülern bei der (experimentellen) Erarbeitung physikalischer Sachverhalte auf. Grundlegende, aus der Förderpädagogik etablierte, Handlungsempfehlungen werden anschließend zur Adaption und Umsetzung physikalischer Inhalte angewandt. Darauf aufbauend gestalten wir gemeinsam mit den Studierenden eine Unterrichtseinheit als Best-Practice-Beispiel, welche mit kooperierenden Förderschulen verschiedener Förderschwerpunkte durchgeführt wird. Diese praktische Anwendung der Handlungsempfehlungen auf schulisches Experimentieren soll den Studierenden dabei helfen, ein erstes Verständnis dafür zu entwickeln, wie eine bessere Teilhabe Lernender mit sonderpädagogischer Unterstützung im regulären Physikunterricht gelingen kann.

DD 14: Postersession 1: Sprache und Physikunterricht

Time: Monday 11:30-12:30

DD 14.1 Mon 11:30 P

Sprachsensibilität im studentischen Erklärvideo zum Fach Physik. Eine explorative Studie — •LENA SCHENK¹, NIKLAS REICHEL² und RAINER MÜLLER¹ — ¹TU Braunschweig, IfDN, Physikdidaktik — ²TU Braunschweig, IfG, Sprachdidaktik

Im Beitrag werden professionsbezogene Kompetenzen zu sprachsensiblem Fachunterricht diskutiert, auf die von Lehramtsstudierenden der Physik bei der inhaltlichen Konzeption und Umsetzung von sprachsensiblen Erklärvideos zurückgegriffen wird. Im Rahmen eines kooperativen Projektseminars mit Studierenden der Physik und Germanistik konnte bereits festgestellt werden, dass der Transfer theoretischen Wissens zu sprachsensiblem Fachunterricht eine große Herausforderung darstellt. So gelingt zwar in den meisten Fällen die didaktische

DD 15: Communicating Physics and its History (joint session GP/DD)

Time: Monday 13:30-15:00

Invited TalkDD 15.1Mon 13:30GP-H7The development of The Lorentz Lab: bringing the scientific history of TeylersMuseum to life with working replicas — •TRIENKE VAN DER SPEK — TeylersMuseum, Haarlem, The Netherlands

In 2017 Teylers Museum opened The Lorentz Lab, a new extension to the permanent presentation of the museum. The Lorentz Lab is specifically dedicated to revive the institute's scientific past as a research institute and laboratory. The scientific instrument collection has long been a silent witness of this past, as are the old laboratory buildings. The Lorentz Lab brings the unique origins of this heritage into life in a carefully recreated historical setting of Teylers' Physics Laboratory.

Here working replica's, theatrical support and a tailor made educational program allow different groups of visitors to participate in science directly and engage with the activities and scientists that shaped Teylers Museum and its collections. The program The Lorentz Formula is dedicated to the general museum visitor, and Einstein was here is a high-level physics program, designed to provoke questions that fit in with educational requirements and with links to the school curriculum.

This lecture will discuss the development of the Lorentz Lab and the central role of working replicas therein. It will address its challenges, original goals, experiences and insights - both from the general public and the educational programs' point of view.

Invited Talk DD 15.2 Mon 14:00 GP-H7 Physics in Information Comics •HEIKE ELISABETH JUENGST FHWS Würzburg <td

Information comics are comics designed primarily for knowledge transfer. They can be funny or serious, long or short, and can be found all over the world. And,

of course, they can cover any topic.

However, information comics about physics are not very common. "The Physics of Superheroes" by Kakalios is justly popular but does not deal with information comics at all. Chemistry, on the other hand, is a popular topic for information comics.

Reduktion fachwissenschaftlicher Inhalte, eine sprachsensible Aufbereitung, wie

sie im Rahmen der fachdidaktischen Anforderung an angehende Lehrkräfte wünschenswert wären, erfolgt jedoch nur bedingt. Besonderes Augenmerk der

Autor:innen liegt hierbei auf der oft vorzufindenden fachlichen Entfernung zum

Thema Sprache bei den angehenden Physiklehrkräften, die im Studium sonst

kaum mit linguistischen Inhalten in Berührung kommen. Der Posterbeitrag

skizziert Ergebnisse einer explorativen Studie, in der untersucht wird, (a) auf

welche Strukturdiagnose- und -aufbereitungsmechanismen Lehramtsstudieren-

de der Fachdisziplin Physik bei der Aufbereitung eines fachwissenschaftlichen

Themas in einem sprachsensiblen Erklärvideo zurückgreifen und (b) wie die

sprachsensible Aufbereitung in einem Erklärvideo methodisch realisiert wird.

The reason for this imbalance is a mystery and cannot be solved in this presentation. Information comics presenting topics from physics will be shown and the audience will be encouraged to produce their own information comics for teaching purposes (and for fun).

Invited TalkDD 15.3Mon 14:30GP-H7Jenseits gewohnter Pfade: Ausstellungen neu denken – •CHRISTIAN SICHAU
– experimenta gGmbH, Heilbronn

Lange Zeit galten Science Center als Ausstellungsorte, in denen über die Darstellung natürlicher Phänomene Begeisterung für Naturwissenschaft und Technik geweckt werden sollte. Obwohl hinsichtlich der Besucherzahlen recht erfolgreich, wurde ihr Ansatz immer wieder als unzureichend kritisiert. Science Center würden kein adäquates Verständnis der Natur der Naturwissenschaften vermitteln, ihnen mangele es an kritischer Reflektion und an einer ernsthafter Auseinandersetzung mit der Wissenschaft und ihrer Geschichte. Bei dieser gelegentlich scharf und polemisch vorgetragenen Kritik wird häufig übersehen, dass es seit vielen Jahren und auf internationaler Ebene zahlreiche neue Ansätze gibt. So wird immer mehr und intensiver in der Science Center-Szene diskutiert, wie das öffentliche Verständnis für und über Wissenschaft vermittelt und gestärkt werden kann. Hierbei kann gerade die Stärke der Science Center - das Erreichen einer sehr breiten Öffentlichkeit mit niedrigschwelligen Angeboten - ein wichtiger Pluspunkt werden. Am Beispiel der experimenta soll aufgezeigt werden, dass hier neue - und vielleicht sehr wirksame - Narrative entstehen können.

DD 16: Neue / digitale Medien - Schule

Time: Monday 15:30-16:50

DD 16.1 Mon 15:30 DD-H8

Scientifically Speaking: Kollaboratives Lernen digital unterstützen — •THOMAS SEAN WEATHERBY und THOMAS WILHELM — Didaktik der Physik, Goethe-Universität Frankfurt am Main, Max-von-Laue-Straße 1, 60439 Frankfurt am Main

Die individuelle und flüchtige Natur des Schülergesprächs macht es schwierig, dieses in einer ganzen Klasse bei allen erfolgreich zu strukturieren. Da die Lehrkraft nicht bei jeder einzelnen Diskussion anwesend sein kann, um die Lernenden zu den naturwissenschaftlichen Vorstellungen zu führen, wird in diesem Vortrag eine Web-App vorgestellt, die die Lehrkräfte dabei unter-stützen soll, Schülergespräche zu einem einfachen und gewinnbringenden Teil des naturwissenschaftlichen Unterrichts zu machen. Aufbauend auf Forschungsergebnissen in der Literatur zu kollaborativem und digitalem Lernen werden die Ideen hinter dem Design und der Verwendung der Software vorgestellt. Die dazugehörige Studie, die gerade in mehreren Schularten in Großbritannien läuft, bettet die Web-App in einen erprobten Lehrplan zu einfachen Stromkreisen ein und vergleicht den Lernerfolg mit und ohne Benutzung der Software.

DD 16.2 Mon 15:50 DD-H8 Interaktive Augmentierung klassischer Lehrfilme — •Jürgen Kirstein und Volkhard Nordmeier — Freie Universität Berlin, Didaktik der Physik Lehrfilme, insbesondere solche mit experimentellem Bezug, sind im Physikunterricht ein Mittel der Anschauung. Studien zur Erklärqualität haben gezeigt, das die passive Rezeption eines Lehrfilms zum Verstehen nicht ausreicht. Verstehen erfordert vielmehr Lernaktivitäten, mit denen Lernende die Inhalte des Films selbstständig verarbeiten können. Zur Augmentierung verwenden wir originales Filmmaterial und bearbeiten es so, dass Lernende die im Film dargestellten (experimentellen) Handlungen oder die Auswertung von Messergebnissen interaktiv nachvollziehen können. Eingebettet in eine an die Dramaturgie des Films angelehnte digitale Lernumgebung, bietet das Zusammenspiel aus erklärenden und inaktiven Elementen neue didaktische Gestaltungsmöglichkeiten. Wir haben dieses Konzept anhand des Lehrfilms "Time dilation, an experiment with mu-mesons" modellhaft umgesetzt, der 1963 im Rahmen des US-Bildungsprogramms des Physical Science Study Committee (PSSC) entstand. Der Film dokumentiert den Aufbau und die Durchführung dieses klassischen Experiments zum Nachweis der Zeitdilatation im Detail.

DD 16.3 Mon 16:10 DD-H8

Location: DD-H8

Die Überzeugungskraft interaktiver Experimentiervideos — •LION CORNELI-US GLATZ, ROGER ERB und ALBERT TEICHREW — Goethe-Universität Frankfurt am Main

Die Vermittlung des Teilchenmodells als Teil des Basiskonzepts "Materie" hat

Location: P

Location: GP-H7

eine bedeutende Stellung im Physikunterricht. In einer Studie an der Goethe-Universität Frankfurt wird untersucht, welche Experimente für die Einführung des Teilchenmodells besser geeignet sind, mit besonderem Augenmerk auf ihre Überzeugungskraft. Mit Überzeugungskraft ist die Fähigkeit der Experimente gemeint, einen Wechsel von fachlich falschen zu adäquaten Vorstellungen herbeizuführen.

Um vor diesem Hintergrund geeignete Experimente reliabel vergleichen zu können, wurde eine Auswahl an Experimenten zum Teilchenmodell als interaktive Videos gestaltet. Ergebnisse der Pilotierung (n=47) mit vier Experimenten weisen darauf hin, dass die Experimente überzeugender wirken, deren Inhalte kognitiv weniger fordernd sind. Nun werden erste Ergebnisse einer Folgeerhebung vorgestellt, in der eine größere Anzahl an Experimenten in mehreren Klassen eingesetzt wurden. Es wird untersucht, wie sich die Vorstellungen der Schüler*innen zum Aufbau der Materie in Abhängigkeit des Vorwissens und der Eigenschaften der Experimente verhalten. Außerdem werden Aussagen über die Überzeugungskraft der Experimentiervideos getroffen.

DD 16.4 Mon 16:30 DD-H8 Quantitative Analysen zum Einsatz schüler- und schuleigener Smartphones im Physikunterricht — •DANIEL LAUMANN, MALTE UBBEN, SUSANNE HEINI-CKE und STEFAN HEUSLER — Westfälische Wilhelms-Universität Münster Monday

Smartphones besitzen im Physikunterricht das Potential eine Schnittstelle zwischen klassischen experimentellen und innovativen digitalen Medien darzustellen. Um vorhandene didaktische Konzepte zur Vermittlung physikalischer Inhalte in die Praxis zu transferieren, ist es notwendig, Gelingensbedingungen zu identifizieren. Dabei erscheint die Frage relevant, inwiefern Lernen im Physikunterricht durch schüler- oder schuleigene Smartphones beeinflusst wird. Das BMBF-Projekt smart for science untersucht beide Nutzungsszenarien durch einen quasi-experimentellen Gruppenvergleich für die Nutzung schüler- oder schuleigener Smartphones im Rahmen von drei jeweils mehrstündigen, am Regelunterricht der Fächer Physik, Mathematik und Chemie orientierten Workshops zum Thema "Elektromobilität". Die vorläufigen Ergebnisse im Fach Physik zeigen, inwiefern für den Lernprozess relevante Variablen (Fachwissen, kognitive Belastung, Interesse etc.) durch die verschiedenen Nutzungsszenarien beeinflusst werden. Weiterführende Ergebnisse berichten auch über mögliche positive Einflüsse auf das Lernen bei zeitlich eingeschränkter Nutzung schülereigener Smartphones.

DD 17: Neue / digitale Medien - AR

Time: Monday 15:30-16:50

DD 17.1 Mon 15:30 DD-H9

Technische Entwicklung eines Augmented-Reality-Experiments zu polarisationsverschränkten Photonenpaaren — •Adrian Abazi¹, Paul Schlummer², Jonas Lauströer³, Rasmus Borkamp³, Reinhard Schulz-Schaeffer³, Stefan Heusler², Daniel Laumann², Wolfram Pernice¹ und Carsten Schuck¹ — ¹Center for Nanotechnology, WWU Münster — ²Institut für Didaktik der Physik, WWU Münster — ³Department Design, HAW Hamburg

Viele didaktische Ansätze haben Schwierigkeiten, aufgrund von komplexen Aufbauten und weiter Separation zwischen abstrakten Modellen und dem Akt des Experimentierens, Quanten-Phänomene effektiv zu vermitteln. Wie eine Synthese von Experiment und Modell auf technischer Ebene umgesetzt werden kann, präsentieren wir beispielhaft an einem Experiment zu polarisationsverschränkten Photonenpaaren gestützt durch eine Augmented-Reality-Umgebung. In der AR-Umgebung werden dem Experimentierenden, in Echtzeit interaktive holografische Elemente, in Form von Visualisierungen von Messwerten und Modellen, welche die abstrakten Prozesse zugänglicher machen sollen, im Sichtfeld einer AR-Brille eingeblendet. Der entwickelte experimentelle Aufbau, eine nicht-linear optisch basierende Photonenpaarquelle und simplifizierter Bell-Messungsapparat, besteht nur aus kommerziell erhältlichen Komponenten, um eine simple Transferierbarkeit auf andere Standorte zu gewährleisten. Die nutzerfreundliche Engine Unity, mit der die AR-Umgebung entwickelt wurde und in Python geschriebene modulare Schnittstellen, ermöglichen eine hohe Modifizierbarkeit der Umgebung für weitere Experimente und Aufbauten.

DD 17.2 Mon 15:50 DD-H9

PUMA : Magnetlabor * AR-Applikationen für den Einsatz in Lernstationen im Lehr-Lern-Labor — •HAGEN SCHWANKE und THOMAS TREFZGER — Universität Würzburg

Experimente stehen im naturwissenschaftlichen Unterricht nach wie vor im Zentrum des Unterrichtsgeschehens. Durch den Digitalpakt Schule und die Weiterentwicklungen im informationstechnischen Bereich ergänzen inzwischen kostengünstige digitale Medien und Werkzeuge das Experiment im Unterricht. Die Sekundarstufe I bietet zum Thema der Elektrizitätslehre viele Experimente zur Anwendung einer augmentierten Lernungebung. Die in dem Projekt PhysikUnterricht Mit Augmentierung (PUMA) entwickelte Applikation PUMA: Magnetlabor soll hauptsächlich die Modelle der magnetischen Felder sichtbar machen.

In diesem Vortrag wird zunächst die Frage geklärt, warum sich Augmented Reality (AR) zum Thema Magnetismus anbietet. Daraufhin wird die Studie und deren Verlauf dargestellt. Dabei werden beispielhaft einzelne Stationen einer Lernumgebung vorgestellt, welche auf Grundlage eines Schülerexperimentiersatzes für ein Lehr-Lern-Labor an der Universität Würzburg konzipiert wurden. Eine Herausforderung stellt u.a. die richtige Gestaltung von Aufgaben zur förderlichen Anwendung von AR dar. Location: DD-H9

DD 17.3 Mon 16:10 DD-H9

PUMA : Spannungslabor - Eine AR-Applikation für den Einsatz in der E-Lehre der Sek I — •FLORIAN FRANK, CHRISTOPH STOLZENBERGER und THO-MAS TREFZGER — Julius-Maximilians-Universität Würzburg, Lehrstuhl für Physik und ihre Didaktik

Mit Hilfe von Augmented-Reality-Apps können virtuelle Objekte und Texte in Echtzeit in die reale Welt (z.B. auch bei physikalischen Experimenten) eingefügt werden. Unter dem Namen PUMA (PhysikUnterricht Mit Augmentierung) werden AR-Applikationen für den Einsatz in der schulischen Physiklehre entwickelt. Die AR-App PUMA : Spannungslabor erweitert Experimente zu einfachen Stromkreisen um AR-Darstellungen elektrischer Ströme und Potentiale Durch die Beobachtung dieser sonst nicht wahrnehmbaren Größen können Schüler*innen selbstständig erforschen, warum und auf welche Weise Elektronen in geschlossenen Stromkreisen fließen und welche Gesetzmäßigkeiten bei Reihen- oder Parallelschaltungen gelten. Die App bietet Möglichkeiten für die Bildung und Überprüfung eigener qualitativer Aussagen zum Stromkreis und erlaubt durch Einblendung von Kenn- und Messwerten zusätzlich die Gewinnung halbquantitativer Erkenntnisse. In Lupen-Ansichten verschiedener elektrischer Bauteile können außerdem die Interaktionen der Leitungselektronen mit den Atomrümpfen nach dem Drude-Modell beobachtet werden und ein qualitatives Verständnis für den elektrischen Widerstand gebildet werden.

In diesem Vortrag wird die Applikation PUMA : Spannungslabor mit ihren Funktionalitäten vorgestellt und die entwicklungsbegleitende Forschung zusammengefasst beschrieben.

DD 17.4 Mon 16:30 DD-H9

Lehrkräftefortbildung zu Augmented Reality-Experimenten im Physikunterricht — •Mareike Freese¹, Albert Teichrew¹, Jan Winkelmann², Roger Erb¹, Mark Ullrich¹ und Michael Tremmel¹ — ¹Goethe-Universität Frankfurt am Main — ²PH Schwäbisch Gmünd

Physikalische Modelle sind als Teil der naturwissenschaftlichen Erkenntnisgewinnung in den Bildungsstandards verankert. Studien zeigen jedoch, dass selbst Lehrkräfte häufig Schwierigkeiten mit Modellen haben, da sie als abstrakt gelten und daher weniger häufig explizit im Unterricht behandelt werden. Das digitale Werkzeug Augmented Reality (AR) ermöglicht, reale Experimente im Physikunterricht in Echtzeit um virtuelle Modellierungen zu erweitern, was die zugrunde liegenden Modelle besser zugänglich macht.

In unserem QLB-geförderten Projekt werden Lehrkräfte in Anlehnung an das TPACK-Modell in fünf Sitzungen innerhalb eines halben Jahres in einer Fortbildung darin geschult, AR-Experimente im eigenen Unterricht einzusetzen (TPACK), indem sie lernen, das Softwarepaket GeoGebra zu bedienen (TK), Modelle zu physikalischen Phänomenen zu konstruieren (TCK) und didaktische Materialien zu den Modellen zu entwerfen (TPK). In einer Abschlusssitzung wird die Implementation gemeinsam reflektiert.

Der Kompetenzzuwachs wird in einem Pre-Post-Design qualitativ und quantitativ evaluiert. Zudem werden semistrukturierte Interviews in einem Followup geführt. In dem Beitrag werden die Ergebnisse der Studie vorgestellt.

DD 18: Lehr-Lernforschung - Schülervorstellungen fachbezogen

Time: Monday 15:30-16:50

DD 18.1 Mon 15:30 DD-H10

Rekonstruktion von Begriffsnetzen aus Essays von Schüler*innen zum Basiskonzept Energie — •DENNIS DIETZ, KARL PFAFF und CLAUS BOLTE — Freie Universität Berlin

Sowohl aus bildungspolitischer als auch aus lerntheoretischer Perspektive wird der Vernetzung von Begriffen im Zuge von Lernprozessen eine besondere Rolle beigemessen. In den letzten Jahren wurden verschiedene Verfahren zur Vernetzungsanalyse vorgeschlagen, mit deren Hilfe Begriffsnetze aus Interviews, Schülergesprächen oder Schulbüchern rekonstruiert und mittels gängiger Netzwerkparameter in Bezug auf ihre jeweilige Vernetzung analysiert werden können. Die Untersuchungen zeigen bedeutsame Zusammenhänge zwischen der Netzwerkkohärenz der jeweils rekonstruierten Begriffsnetze und dem Lernerfolg der Schüler*innen auf (Kubsch et al., 2019; Podschuweit & Bernholt, 2020). Folgerichtig schlagen Podschuweit und Bernholt (2020, S. 14) vor, Vernetzungsanalysen dieser Art zur Untersuchung von Schülertexten zu nutzen. In unserem Beitrag stellen wir ein dafür eigens entwickeltes Analyseverfahren vor, das die systematische Rekonstruktion von Begriffsnetzen aus Schüler*innen-Essays mit Hilfe unseres Modells zur Analyse der Vernetzung von Begriffselementen in Essays ermöglicht (Dietz & Bolte, 2021). Auf der Grundlage einer theoriegeleiteten Identifikation naturwissenschaftlich relevanter Begriffe und auf der Basis theoretischer Anleihen aus der germanistischen Linguistik haben wir Begriffsnetze zum Basiskonzept Energie aus Essays von 50 Schüler*innen der 9. Jahrgangsstufe rekonstruiert. Im Rahmen unseres Vortrags stellen wir das Verfahren sowie prototypische Begriffsnetze zur Diskussion.

DD 18.2 Mon 15:50 DD-H10

Assessing Austrian high school students' understanding of basic wave optics phenomena using the Conceptual Survey on Wave Optics — •KAROLINA MATEJAK CVENIC¹, MAJA PLANINIC¹, ANA SUSAC², LANA IVANJEK³, KATARINA JELICIC¹, and MARTIN HOPF⁴ — ¹Department of Physics, Faculty of Science, University of Zagreb, Croatia — ²Department of Applied Physics, Faculty of Electrical Engineering and Computing, University of Zagreb, Croatia — ³Faculty of Physics, Physics Education Research, Technische Universität Dresden, Germany — ⁴University of Vienna, Austrian Educational Competence Centre Physics, Austria

The Conceptual Survey on Wave Optics (CSWO) is a new diagnostic instrument aimed at the assessment of high school students' understanding of some basic phenomena of wave optics. The CSWO consists of 26 multiple choice items that include questions about double-slit and optical grating interference, single-slit diffraction, and polarization of light. The final version of the CSWO was administered to 170 students of several Viennese high schools, who took the test in an online form after finishing regular school instructions on wave optics. The survey was conducted during the summer term of school year 2020/21. The Rasch analysis of the results of Austrian students on this survey will be presented and discussed, together with some implications for high school instruction on wave optics. The results suggest that wave optics is a rather difficult topic for high school students, and it seems that recognizing patterns and explaining the basic wave optics phenomena are especially difficult tasks for students.

DD 18.3 Mon 16:10 DD-H10

Entwicklung und Evaluation eines Optikprojekts zu Schülervorstellungen in der Primarstufe — •LINDA SEUFERT, WOLFGANG LUTZ und THOMAS TREFZGER — Julius-Maximilians-Universität Würzburg

Bereits in der Grundschule bringen die Schülerinnen und Schüler Präkonzepte zu physikalischen Phänomenen aus der Optik mit in den Unterricht, beispielsweise zum Sehvorgang oder dem Wahrnehmen von Farben. Um ihnen möglichst frühzeitig ein anschlussfähiges Konzept an die Hand zu geben, sollten physikalische Inhalte zur Optik bereits im Sachunterricht der Primarstufe behandelt werden. Die Entwicklung eines konzeptionellen Verständnisses stellt dabei ein zentrales Ziel des Unterrichts dar und kann durch verschiedene Experimente und Visualisierungen unterstützt werden. Vor diesem Hintergrund wurde für die vierte Jahrgangsstufe eine Unterrichtskonzeption zum Sehvorgang sowie zum Thema Licht & Farben entwickelt und in einer Intervention erprobt. Mit Hilfe von Interviews wurden die Vorstellungen der Schülerinnen und Schüler zu dieser Thematik sowohl vor als auch nach der Intervention erhoben. Im Beitrag werden die Unterrichtskonzeption, die zugrundeliegenden didaktischen Überlegungen und die verwendeten Materialien vorgestellt sowie Einblicke in die Erkenntnisse aus den Interviews gegeben.

DD 18.4 Mon 16:30 DD-H10 Student understanding of half-life and background radiation — •Michael M Hull, Eva Holzinger, Maximilian Jeidler, and Markus Wintersteller — Universität Wien, Wien, Österreich

We have been studying the conceptual understanding of high school students about radioactivity, particularly regarding background radiation and half-life. We have seen that it is difficult for learners to grapple with the idea that random behavior of individual atoms can give rise to predictable patterns in the collective, and many students have said both on the Fission as a Random Occurrence Survey (FAROS) and in interviews that, if you are looking at an individual atom, half of the atom will have fissioned after one half-life. Our findings have indicated, however, that this idea (of individual atoms fissioning in a predictably continuous manner) is often not a robust and intact mental structure; rather, in other contexts, the same students correctly discuss fission as being instantaneous and unpredictable. Approaches to teaching radioactivity that take this fluidity of student reasoning into account are desired. We created and validated an expanded version of FAROS and are using the survey to assess "The Radiation Around Us", curriculum that does exactly that.

DD 19: neue Konzepte - Physikunterricht

Time: Monday 15:30–16:50

DD 19.1 Mon 15:30 DD-H11

Analyse dynamischer Unterrichtsinteraktionen mit Hilfe von Advanced State Space Grids — •NIKLAS LITZENBERGER und ANDREAS PYSIK — Johannes Gutenberg-Universität, Mainz, Deutschland

Unterricht zeichnet sich durch komplexe Interaktionen aus. Dies stellt die Unterrichtsforschung vor eine methodische Herausforderung, wenn diese Dynamik erfasst werden soll. Häufig werden bislang Methoden eingesetzt, die auf globale Einschätzungen des Unterrichts abzielen, wobei Aussagen über dynamische Prozesse ausbleiben. Für die dynamische Erforschung von Unterrichtsaktionen bieten die von Hollenstein (2013) entwickelten State Space Grids (SSG) eine geeignete Grundlage. SSG ermöglichen es, den zeitlichen Verlauf zweier Indikatoren abzubilden. Das Ziel der hier vorgestellten Studie ist es, zusätzlich den Zusammenhang zweier Indikatoren numerisch zu erfassen zu können. Dazu werden SSG durch mathematisch fundierte Begriffe zu Advanced State Space Grids (ASSG) erweitert, mit Standardmethoden der Unterrichtsforschung verglichen und so ihre Praktikabilität geprüft. Dazu gehören unter anderem neue Zusammenhangsstärken zwischen den erfassten Indikatoren, Chi-Quadrat-Tests, Parameter zu Streudichten und Abweichungen, sowie ein Flussdiagramm für die Interpretation der Parameter. Kritische Werte der Parameter wurden durch zwei Millionen Simulationen von Kombinationen von Indikatoren festgestellt. Erste Ergebnisse zeigen, dass ASSGs Zusammenhänge zwischen Unterrichtsmerkmalen beschreiben können, die mit Globalbewertungen nicht untersuchbar sind. Entsprechende Implikationen sowie Vor- und Nachteile der ASSG-Methode werden in diesem Beitrag diskutiert.

Location: DD-H11

DD 19.2 Mon 15:50 DD-H11

Das kontextstrukturierte Unterrichtskonzept EPo-EKo — •Jan-Philipp Burde¹, Liza Dopatka², Verena Spatz², Martin Hopf³, Thomas Wilhelm⁴, Thomas Schubatzky⁵, Claudia Haagen-Schützenhöfer⁵, Lana Ivanjek⁶ und Benedikt Gottschlich¹ – ¹Universität Tübingen – ²TU Darmstadt – ³Universität Wien – ⁴Universität Frankfurt – ⁵Universität Graz – ⁶TU Dresden

Ein Ziel des binationalen Projektes "Elektrizitätslehre mit Potenzial - Elektrizitätslehre mit Kontexten" (EPo-EKo) besteht darin, den Elektrizitätslehreunterricht lernwirksamer und interessanter zu gestalten. Spätestens seit den KMK-Beschlüssen wird ein stärkerer Einbezug von Kontexten im Physikunterricht angestrebt. Vor dem Hintergrund, dass bisher jedoch kein kontextbasiertes und empirisch evaluiertes Unterrichtskonzept zu einfachen Stromkreisen existiert, wurde im Rahmen des EPo-EKo-Projektes ein kontextstrukturiertes Unterrichtskonzept auf Basis der Sachstruktur des bewährten, aber rein fachsystematisch ausgerichteten Unterrichtskonzepts "Eine Einführung in die Elektrizitätslehre mit Potenzial" (kurz "EPo-Konzept") entwickelt. Entsprechend der Erkenntnisse der Interessensforschung werden die fachlichen Inhalte des neuen Unterrichtskonzepts "Eine Einführung in die Elektrizitätslehre mit Potenzial und Kontexten" (kurz "EPo-EKo-Konzept") u. a. in biologische und medizinische Kontexte eingebettet und entlang dieser erarbeitet. Im Vortrag werden die Grundideen dieses kontextstrukturierten Unterrichtskonzepts vorgestellt, das auf Basis des Feedbacks von Lernenden und Lehrkräften weiterentwickelt wurde.

DD 19.3 Mon 16:10 DD-H11 Kreisbewegungen erklären mit abzählbaren radialen Stößen - •BRUNO HARTMANN — Humboldt Universität, Berlin

Wir entwickeln einen neuen dynamischen Zugang für das Unterrichten der Kreisbewegung. Nach dem Trägheitsprinzip bewegen sich Körper ohne äußere Beeinflussung gleichförmig und geradeaus. Mit einem Stoß kann die Bewegung zur Seite abgelenkt werden. Durch fortgesetztes Ablenken kann ein vollständiger Kreis entstehen. Mit Hilfe von Standardstößen, die alle die gleiche Stärke und eine wohldefinierte Richtung haben, konstruieren die Schüler Kreisbewegungen mit unterschiedlichen Radien, Bahngeschwindigkeiten und Massen. Durch Abzählen der anschaulichen Standardstöße werden alle Einflussfaktoren auf die Radialkraft quantifiziert. Die angegebenen Unterrichtsbeispiele wurden im gymnasialen Physikunterricht erprobt.

DD 19.4 Mon 16:30 DD-H11 Bilder der Kugellinse — • THOMAS QUICK und JOHANNES GREBE-ELLIS — Bergische Universität Wuppertal

phänomene, die sich als Kontexte der geometrischen Optik eignen. Aus der Perspektive des Physikunterrichts sind alltägliche Linsenerscheinungen allerdings problematisch, denn die beobachtbaren Phänomene erfüllen zum größten Teil gerade nicht die paraxiale Approximation, die man üblicherweise zur Behandlung von dünnen Linsen und der Bildformation für die möglichst fehlerfreie Abbildung in technischen Werkzeugen heranzieht. Eine Reihe von Erscheinungen wie z.B. Kaustiken, Bildverzerrungen oder Mehrfachabbildungen, die für Schülerinnen und Schüler motivierend und interessant sein könnten, liegen außerhalb des Beschreibungsrahmens der paraxialen Näherung und werden üblicherweise als Bildfehler angesprochen. Am Beispiel der Kugellinse zeigen wir, wie sich die Bedingungen der Bildentstehung im optischen Gesamtraum explorativ so erkunden lassen, dass einerseits auch komplexe Linsenphänomene abseits der optischen Achse einer systematischen Untersuchung zugänglich sind und sich andererseits die einschränkenden Annahmen der paraxialen Näherung schließlich als Spezialfall ergeben. Als geeignetes Untersuchungsobjekt erweist sich dabei die virtuelle Kaustik, die das Transformationsverhalten optischer Bilder regelt.

Regentropfen, Trinkgläser oder durchsichtige Vasen zeigen vielfältige Linsen-

DD 20: außerschulisches Lernen - Konzepte

Time: Monday 15:30-16:50

DD 20.1 Mon 15:30 DD-H12

Mädchenförderung in der Physik — • CHRISTIANE RICHTER und MICHAEL KOмокек — Universität Oldenburg Der Frauenanteil am IfP der Universität Oldenburg (UOL) liegt durchschnitt-

lich bei 25%. Um den Frauenanteil auf allen Oualifikationsebenen zu erhöhen, wurden zwei markante Stellen charakterisiert, an denen es sinnvoll ist anzusetzen. Bereits in der Schule muss es gelingen, mehr Schülerinnen für das Fach Physik zu begeistern. Cimpian et al. (2020) bringen einen weiteren interessanten Aspekt ein. Wer Mädchen für MINT-Fächer interessieren will, der sollte sich an die durchschnittlich Begabten richten, denn bei Letzteren sei das Frauendefizit am größten. Daraus lässt sich schließen, dass sich Mädchen mit niedrigem Fähigkeitsselbstkonzept ein MINT-Studium nicht zutrauen. Die UOL trägt durch verschiedenste Maßnahmen dazu bei, das Fähigkeitsselbstkonzept von Mädchen hinsichtlich der MINT-Disziplinen weiterzuentwickeln oder grundlegend aufzubauen und diese so für Physik zu begeistern. In einem dieser Maßnahmen werden Mädchen in einem Dreischritt - Vorbesprechung und Zielfestsetzung, Experimentiertag, Nachbesprechung und Beratung - dazu angeregt, sich mit einem physikalischen Problem auseinanderzusetzen. Begleitet durch zwei Bachelor-Arbeiten wird diese Maßnahme in Kooperation mit dem Projekt AHOI-MINT im WiSe 21/22 erprobt und evaluiert. Im Vortrag werden das Konzept, dessen Erprobung und Evaluation vorgestellt. Keywords: Mädchenförderung, Problemlösen, Fähigkeitsselbstkonzept

DD 20.2 Mon 15:50 DD-H12

Physikalische Bildung in komplementär vernetzen non-formalen Lernangeboten — •JONAS TISCHER, CHRISTIN SAJONS UND MICHAEL KOMOREK — UNIversität Oldenburg

Außerschulische Lernorte verfügen über das wertvolle Potenzial, komplexe Themen interdisziplinär zu beleuchten. Die physikalische Perspektive ist dabei notwendig, damit Kinder und Jugendliche naturwissenschaftlich-technische Zusammenhänge nachvollziehen und auch erklären können. In einer von der niedersächsischen BINGO-Umweltstiftung geförderten "Komplementären Projektwoche" werden vorhandene Angebote verschiedener non-formaler Lernorte (Museum, Schülerlabor, Umweltbildungszentrum etc.) zum Thema 'Herausforderung Leben im Klimawandel' ergänzend und auch im Kontrast zueinander kombiniert. An der Projektwoche haben fünf 6. Schulklassen an vier Tagen je einen der Lernorte besucht und am fünften Tag ihre Erfahrungen reflektiert. Der physikalische Fokus wurde an den Lernorten unterschiedlich gesetzt. Die Projektwoche wurde mittels Beobachtungen und Interviews mit Schüler:innen und Lehrpersonen (Schule, Lernorte) empirisch begleitet. Transkripte und Arbeitsergebnisse der Schüler:innen wurden kategorienbasiert ausgewertet und aufeinander bezogen. Im Vortrag wird berichtet, inwiefern die Schüler:innen angebotene Perspektiven rekonstruieren und notwendiges physikalisches und weiteres Wissen nutzen konnten, um ein komplexes Bild von den Herausforderungen des Klimawandels aufzubauen.

DD 20.3 Mon 16:10 DD-H12 SmartMatters4You - Intelligente Materie in Workshops und Hands-On-Laborbesuchen — Cornelia Denz¹, Christian Klein-Bösing¹, Christina $\label{eq:Kriegel} \begin{array}{l} {\rm Kriegel}^2, \bullet {\rm Barbara} \ {\rm Leibrock}^1 \ {\rm und} \ {\rm Julian} \ {\rm Repke}^1 - {}^1 {\rm MExLab} \ {\rm Physik}, {\rm Universit\"at} \ {\rm M} \\ {\rm unster} - {}^2 {\rm Center} \ {\rm for} \ {\rm Soft} \ {\rm Nanoscience}, \ {\rm Universit\mathstrutat} \ {\rm M} \\ {\rm unster} \end{array}$

Das Projekt "SmartMatters4You" ist eine Kooperation zwischen dem außerschulischen Lernort MExLab Physik an der Westfälischen Wilhelms-Universität Münster und dem Sonderforschungsbereich SFB1459 "Intelligente Materie: Von responsiven zu adaptiven Nanosystemen" und richtet sich an Schülerinnen der Oberstufe und Studentinnen der Chemie, Physik, Biologie und Informatik bis zum zweiten Semester.

Über die Spanne von drei Jahren werden drei Gruppen von je 50 Teilnehmerinnen nach einer ersten Auseinandersetzung mit Intelligenter Materie in der Bewerbung mehrere Workshops zum Themenkomplex absolvieren und die Chance bekommen, Einblicke in die Laborarbeit des Sonderforschungsbereichs und in Firmen im Münsterland zu nehmen. Das Projekt vermittelt dabei Wissen, experimentelle Erfahrungen und Berufseinblicke in den beteiligten Fachbereichen, bietet darüber hinaus aber zusätzlich eine Austauschmöglichkeit zwischen Schülerinnen der Oberstufe und Studentinnen der Naturwissenschaften. Im Vortrag werden die Projektplanung und Ausgestaltung für die erste Gruppe vorgestellt.

DD 20.4 Mon 16:30 DD-H12

Entwicklung von Hands On-Exponaten zur Mikrogravitation - •DAVID BORGELT¹, JESSICA OERTEL² und CORNELIA DENZ² – ¹Institut für Kernphysik, Wilhelm-Klemm-Str. 9, 48149 Münster — ²Institut für Angewandte Physik, Corrensstr. 2 48149 Münster

Im Rahmen des sechswöchigen Open-Air Wissenschaftsfestivals, dem Q.UNI Camp der WWU Münster mit der Zielgruppe von Kindern und Jugendlichen, wurden Hands On-Exponate zur Mikrogravitation entwickelt. Für dieses Vorhaben ist ein mobiler Fallturm mit einer Höhe von 4 m und ein Trampolinexponat für außerschulische Lernorte konstruiert worden. An diesen Exponaten wurden Freifallexperimente mit ausstellungsbezogenen Kriterien entworfen und es wurde untersucht, inwiefern die in der Entwicklung angestrebten didaktischen Aspekte erreicht werden konnten. Dazu ist eine explorative Studie anhand von Interviews mit Kindern durchgeführt worden. Im Vortrag wird ein Einblick in die Exponatsentwicklung und Ergebnisse der Studie gegeben sowie Möglichkeiten zur Weiterentwicklung als Workshop an außerschulischen Lernorten vorgestellt.

DD 21: Hauptvortrag 2

Time: Tuesday 11:00-11:45

Invited Talk

DD 21.1 Tue 11:00 DD-H8 Warum Lehrbuchdarstellungen der Physikgeschichte so schlecht sind - und was wir daraus lernen können — •OLIVER PASSON — Bergische Universität Wuppertal

Der Umgang von Lehrbüchern (auf Schul- und Hochschulniveau) mit der Geschichte der Physik ist bekanntlich erschreckend schlampig. Zahllose Ungenauigkeiten und Fehler tradieren sich durch die Auflagen hindurch und allem Anschein nach ist den Autorinnen und Autoren die historische Genauigkeit ihrer Darstellungen herzlich egal. Aber warum ist dies so? Mein Vortrag erläutert den Zusammenhang dieser Tatsache zum Verständnis von Wissenschaft, das den jeweiligen Darstellungen zugrunde liegt. Vereinfacht ausgedrückt: Räumt man den Naturwissenschaften einen weitgehenden epistemischen Vorrang ein

206

(etwa weil man die wissenschaftliche Methode für hinreichend selbstkorrigierend hält), dann ist die Geschichte der Naturwissenschaften bloß eine Summe kontingenter (und damit letztlich irrelevanter) Faktoren. Ich diskutiere damit zusammenhängende Fragen anhand von Fallbeispielen (unter anderem aus der Quantenphysik) und entwickle allgemeine Konsequenzen für die Physikdidaktik.

DD 22: Lehreraus- und Lehrerfortbildung - neue Ansätze

Time: Tuesday 12:00-13:00

DD 22.1 Tue 12:00 DD-H8

The Development of personal and professional attributes of students in a physics content media course — •KATHLEEN FALCONER, STEFAN HOFF-MANN, and ANDRÉ BRESGES — Institut für Physikdidaktik, Universität zu Köln, Deutschland

The Medienpraktikum I and Medienpraktikum II were courses developed to improve the media competences of the HRSeG (Haupt-, Real, Sekundar- Gesamtschule), SoPäd (Sonderpädagogik) , and GymGe (Gymnasium Gesamtschule) physics education students. The courses focus on the use and production of media, including videos, simulations, animations, iBooks, etc. Using a Problem Based Learning (PBL) approach, the students work on creation, implementation and embedding of a media product for a particular context. While the personal and professional attributes of successful students in physics have been studied, the same can*t be said for physics students who create physics content media for use in the classroom. We will discuss the changes in the personal and professional attributes of students in the Medienpraktikum courses. The students* reflective writings and pre/post online surveys were analyzed using grounded theory. In a preliminary analysis, the students* view of self-management and self-motivation, especially in regards to peer/group learning seems to develop through their experiences in the course. There is mixed evidence for change in the students* worldview viability of the creation of media for classrooms. Their view of the usage and utility of media seems to be changeable as well.

DD 22.2 Tue 12:20 DD-H8

Entfaltung der Rollen der Mentor*innen und Tutor*innen im Learningby-Teaching-Konzept der Physiklehrerausbildung — •STEFAN HOFFMANN, KATHLEEN FALCONER und ANDRÉ BRESGES — Institut für Physikdidaktik, Universität zu Köln, Deutschland

In der Ausbildung von Physiklehrer*innen setzt man in Köln in den typischen Erstsemesterkursen in Experimentalphysik bereits seit über 10 Jahren auf tutorielle Konzepte, in denen Physikstudierende direkt zu Studienbeginn erste Lehrerfahrung in der Leitung von kleinen Lerngruppen sammeln können. Traditionelle Vorlesungen wurden durch aktivierendere Formate und ein verschiedene Schulformen und -fächer integrierendes Lehr-Lern-Konzept ersetzt: Learning by Teaching. Die Forschung zur Entwicklung und Verfeinerung des Konzepts verwendete das Modell des *Lehrers als Reflexionspraktiker* der Aktionsforschung (vgl. Kemmis & McTaggert 1988). Dabei waren Lehrende gleichzeitig in der Rolle von Forschern, die Ihre eigene Lehrpraxis erforschten. Der Fokus lag zunächst auf der Rolle der Lehrenden und auf Umfang und Ablauf der studentischen Aktivitäten. In der weiteren Entwicklung fand eine deutliche Entfaltung der Handlungsfähigkeit der beteiligten Rollen der Mentor*innen und Tutor*innen statt. Der Vortrag konzentriert sich auf die stetige Weiterentwicklung mittels Methoden der Partizipatorischen und Emanzipatorischen Aktionsforschung (vgl. Eilks 2018). Besonders deutlich werden die Entwicklungen durch Betrachtung der einzelnen am Learning by Teaching beteiligten Rollen (Tutor*innen, Tutees, Mentor*innen, Lehrende) und deren Evolution im Rahmen des Action Research Personality Continuum (Gibbs 2016).

DD 22.3 Tue 12:40 DD-H8 Einstellungen von Lehramtsstudierenden zu naturwissenschaftlichintegriertem Unterricht — •JANA BIEDENBACH und VERENA SPATZ — Technische Universität Darmstadt

Ein Fach oder drei Fächer? Naturwissenschaften oder Biologie, Chemie und Physik? Insbesondere im Kontext der Sekundarstufe I wird dies immer wieder diskutiert.

Lehramtsstudierende studieren in der Regel maximal zwei der drei Fächer. Somit fehlt in den meisten Studiengängen eine ganzheitliche Herangehensweise an die Naturwissenschaften, sodass den zukünftigen Lehrkräften häufig der Zugang zu naturwissenschaftlich-integrierten Unterrichtsansätzen im Studium verborgen bleibt.

Der Vernetzungsbereich an der TU Darmstadt hat das generelle Ziel, Lehramtsstudierende über die Grenzen ihrer Fächer hinaus gezielt zusammenzubringen, um interdisziplinäre Zugänge zum Unterricht zu thematisieren. Für diesen Vernetzungsbereich wurde das Modul 'Erkenntnisgewinnung in den Naturwissenschaften' von Grund auf neu konzipiert.

Im Rahmen einer Seminarsitzung werden die Vor- und Nachteile des naturwissenschaftlich-integrierten Unterrichts diskutiert. Im Anschluss an dieses Thema fertigen die Studierenden eine persönliche Stellungnahme an, anhand derer ausgewertet wird, welche Argumente für die Lehramtsstudierenden bedeutsam sind. Der Vortrag stellt das Ergebnis dieser Auswertung vor.

DD 23: Neue / digitale Medien -VR

Time: Tuesday 12:00-13:00

DD 23.1 Tue 12:00 DD-H9

VRE Physik im digitalen Labor — •JOHANNES LHOTZKY, WILLIAM LINDLAHR und KLAUS WENDT — Johannes Gutenberg-Universität, Institut für Physik Der Physikunterricht lebt von seinen Experimenten und von der aktiven Auseinandersetzung mit physikalischen Untersuchungsgegenständen und Phänomenen. Wegen hoher Gefahrenquellen, die bspw. von radioaktiven Stoffen, Lasern oder Hochspannungseinsatz ausgehen, gibt es Versuche, die (bisher) nicht oder nur erschwert im Unterricht durchgeführt werden können. Eine Möglichkeit, dennoch eigenständig experimentieren zu können, bieten die an der Johannes Gutenberg-Universität entwickelten Virtual-Reality-Experimente (VRE). Diese bilden in einer digitalen, aber authentischen Welt reale Physik ab. Innerhalb der VRE können Experimente mit den dazugehörigen Apparaturen und Geräten benutzt und bedient werden. So sieht der Funktionsumfang der Software u.a. vor, Messbereiche an Multimetern zu variieren oder Spannungen/ Ströme, Timer und Zählvorrichtungen zu betätigen - gleichermaßen wie wir es auch am Realexperiment durchführen würden. Durch die plattformübergreifende Kompatibilität der Software ist es neben Smartphones, Tablets und Computern ebenso möglich, die virtuellen Versuche auf interaktiven Bildschirmen wie Whiteboards zu betreiben. Im Vortrag werden konkrete Realisierungen vom CsBa-Generator bis zur Röntgenspektroskopie präsentiert.

DD 23.2 Tue 12:20 DD-H9 Das Rastertunnelmikroskop - Konzeption eines Virtual-Reality-Experiments (VRE) — •JAN SIMON, AARON REITH, JOHANNES LHOTZKY UND KLAUS WENDT — Johannes Gutenberg-Universität, Institut für Physik

Virtual-Reality-Experimente (VRE) stellen realistische 3D-Simulationen naturwissenschaftlicher Versuche dar. Dabei sollen vor allem Experimente umgesetzt werden, deren Anschaffung für die Schule zu teuer ist oder von denen zu hohe Gefahrenpotentiale ausgehen. VRE ermöglichen durch ihre digitale Verfügbarkeit Partizipation, wo diese sonst nicht oder nur eingeschränkt möglich wäre. Das Rastertunnelmikroskop (RTM) kann aufgrund des Tunneleffekts die Oberflächenstruktur von leitenden Materialien, wie z.B. Graphit und Gold, auf atomarer Ebene visualisieren. Um die aufgenommenen Kristallstrukturen interpretieren zu können, muss das Orbitalmodell verwendet werden. Somit bietet das RTM eine einzigartige Möglichkeit, sowohl die physikalischen Effekte als auch die Modelle zu thematisieren. Das RTM wird in der Schule üblicherweise nur theoretisch behandelt. Gründe dafür sind vor allem der hohe Anschaffungspreis, zeitaufwendige Durchführung und die Fehleranfälligkeit aufgrund von äußeren Einflüssen. Somit eröffnet die Realisierung des RTM als virtuelles Versuchslabor eine Experimentierumgebung, in der sich Lernende eigenständig mit der Thematik beschäftigen und die Welt der Atome entdecken können. Im Vortrag wird der Weg der Entwicklung des RTM als VRE von den phys. Grundlagen und Modellierungen bis hin zur praktischen Umsetzung dargestellt.

Der Physikunterricht lebt von seinen Experimenten und von der aktiven Auseinandersetzung mit physikalischen Untersuchungsgegenständen und Phänomenen. Wegen hoher Gefahrenquellen, die von bspw. radioaktiven Stoffen, La-

Location: DD-H8

sern oder Hochspannung ausgehen, gibt es Versuche, die (bisher) nicht oder erschwert im Unterricht durchgeführt werden können.

Eine Möglichkeit, dennoch experimentieren zu können, bieten die an der Johannes Gutenberg-Universität entwickelten Virtual-Reality-Experimente (VRE). Diese bilden in einer digitalen, aber authentischen Welt reale Physik ab. Innerhalb der VRE können Experimente mit allen dazugehörigen Apparaturen und Geräten benutzt und bedient werden. Damit werden die Potenziale von "Touch-Medien" im Unterricht genutzt, um Schülerinnen und Schülern neue

DD 24: Physikdidaktik und Inklusion - Experimentieren

tiert

Time: Tuesday 12:00-13:00

DD 24.1 Tue 12:00 DD-H10

Akzeptanz und Nutzung von UDL-basierten Unterstützungsangeboten zum **Experimentieren** – •FRANZISKA KLAUTKE¹ und HEIKE THEYSSEN² ¹Universität Duisburg-Essen, Didaktik der Physik – ²Universität Duisburg-Essen, Didaktik der Physik

Das Universal Design for Learning (UDL) bietet einen Rahmen für die Materialentwicklung und Unterrichtsplanung, um die Diversität aller Lernenden im Fachunterricht zu berücksichtigen und mögliche Lernbarrieren zu reduzieren. Durch das Bereitstellen von Wahl- und Unterstützungsangeboten sollen verschiedene Zugänge ermöglicht werden. Jedoch ist unklar, unter welchen Voraussetzungen Schüler:innen insbesondere bei experimentellen Aufgabenstellungen in der Lage sind, die Wahlmöglichkeiten selbständig zur Unterstützung des eigenen Lernprozesses zu nutzen. Um dieser Frage nachzugehen, wird eine nach UDL-Prinzipien gestaltete Lerngelegenheit mit Experimenten und einem digitalen "Forscherheft" entwickelt und erprobt. Lerngegenstand sind dabei die Planung und der funktionsfähige Aufbau von Experimenten, einschließlich der Berücksichtigung der Variablenkontrollstrategie. Erste Ergebnisse der Pilotierung der Lerngelegenheit werden im Rahmen dieses Vortrags vorgestellt.

DD 24.2 Tue 12:20 DD-H10 Eine Lehrkräftefortbildung zum inklusiven Experimentieren - •LAURA Sührig¹, Katja Hartig¹, Roger Erb¹, Holger Horz¹, Albert Teichrew¹, MARK ULLRICH¹ und JAN WINKELMANN² — ¹Goethe-Universität Frankfurt — ²Pädagogische Hochschule Schwäbisch Gmünd

Fachunterricht nicht mehr an einer fiktiven Homogenität der Schülerschaft auszurichten, ist inzwischen nicht nur noch eine politische Forderung. Unterricht so zu gestalten, dass er allen Schüler*innen einer heterogenen Lerngruppe gerecht wird, ist jedoch eine anspruchsvolle Aufgabe. Im Physikunterricht ist das Experimentieren ein wesentlicher Zugang, um Wissen oder Arbeitsweisen zu vermitteln, und muss daher auch im inklusiven Unterricht eine zentrale Rolle spielen. Im Projekt *FINEX* versuchen wir deswegen, durch eine Lehrkräftefortbildung zum inklusiven Experimentieren einen Beitrag zu leisten.

Möglichkeiten des Experimentierens zu eröffnen.

der Kultusministerkonferenz.

In der Fortbildung lernen Lehrkräfte ein Konzept zur Gestaltung von Schülerexperimentierphasen kennen, welches individuelle Lernwege von Schüler*innen mit einem gemeinsamen (fachlichen) Ziel ermöglicht. Das Konzept bietet die Basis für eine exemplarische Unterrichtseinheit, die die Teilnehmenden in ihren Klassen durchführen, sowie für die Entwicklung eigenen Unterrichts innerhalb der Fortbildung. Nach einer Vorstudie mit Studierenden wurde die Fortbildung mit Lehrkräften durchgeführt. Begleitend wird dabei erhoben, ob sich die Einstellungen der Lehrenden zum inklusiven Experimentieren durch die Teilnahme ändern. In diesem Vortrag werden das Konzept und die Ergebnisse der Evaluation der Fortbildung vorgestellt.

Das Konzept konnte sich in der Zwischenzeit auch im Fernunterricht der

Im Vortrag wird das Konzept der VRE vorgestellt und im Überblick präsen-

Corona-Pandemie bewähren. Gleichzeitig bieten VRE einen fachspezifischen

Ansatz zur Erfüllung der Ziele der Strategie zur "Bildung in der digitalen Welt"

DD 24.3 Tue 12:40 DD-H10 Adaption einer Experimentierreihe zum Thema *Elektrische Stromkreise* für Kinder und Jugendliche mit Sehbehinderungen - •MANUELA WELZEL-BREUER¹, DANIEL MACISAAC², KATHLEEN FALCONER³ und PAMELA DETROIS⁴ — ¹Pädagogische Hochschule Heidelberg, Deutschland — ²State University of New York (SUNY) Buffalo State College, USA - ³Institut für Physikdidaktik, Universität zu Köln, Deutschland — ⁴Staatliche Schule für Blinde und Sehbehinderte in Ilvesheim, Deutschland

Das Experimentieren sollte (auch und gerade) im naturwissenschaftlichen Unterricht mit blinden und sehbehinderten Menschen eine zentrale Rolle spielen, da hierbei die Prinzipien des handelnden Lernens und der Einsatz vieler Sinne gefördert werden. Das schülerorientierte Durchführen von Experimenten stärkt ihre Handlungs-, Methoden- sowie Sachkompetenz und nimmt damit wichtige blinden- und sehbehindertenpädagogische Lernziele in den Fokus. Es ermöglicht außerdem die Berücksichtigung individueller Lernprozesse (vgl. www.vbs.eu/de). Im hier vorgestellten Projekt wurde ein Teil einer ursprünglich für Straßenkinder entwickelten Experimentierreihe für die Bedürfnisse sehbeeinträchtigter Schüler:innen adaptiert und in der Praxis erprobt. Im Vortrag werden Beobachtungen und Erkenntnisse aus der ersten Erprobungsphase vorgestellt.

DD 25: Hochschuldidaktik - Studieneingangsphase Studierendenperspektive

Time: Tuesday 12:00-13:00

DD 25.1 Tue 12:00 DD-H11

Welche Rolle spielt das Mindset beim Studienabbruch im MINT-Studium? -Eine fächerübergreifende Erhebung im ersten Semester. — •MALTE DIEDE-RICH UND VERENA SPATZ — TU Darmstadt

Viele Studierende aus dem MINT-Bereich stehen in ihrem Studium zum ersten Mal vor einer großen Herausforderung, welche zum Teil zu Zweifeln an der eigenen Begabung führen kann. So verzeichnen Dresel und Grassinger (2013) über das erste Semester einen dramatischen Motivationseinbruch bei einem erheblichen Anteil der Studierenden. Dies könnte mit dem akademischen Selbstkonzept und den impliziten Theorien (Mindsets nach Dweck) zur Intelligenz und zur Begabung im Fach zusammenhängen - allerdings ist die Befundlage hierzu heterogen. In der vorliegenden Studie soll daher der Zusammenhang von akademischem Selbstkonzept und impliziten Theorien (Mindsets nach Dweck) mit Studienabbruchs- und -wechselintentionen untersucht werden. Über eine fachspezifische Messung (Rehberg 2020) und die Berücksichtigung von weiteren Skalen aus dem Mindset Meaning System (Yeager & Dweck 2020) sollen die komplexen Auswirkungen in der Studieneingangsphase tiefer verstanden werden. Hierfür wurde im WS 21/22 an der TU Darmstadt eine Fragebogenerhebung mit 805 Studierenden aus dem ersten Semester der MINT-Studiengänge durchgeführt. Im Vortrag werden das Design der Erhebung und ausgewählte Ergebnisse vorgestellt.

Location: DD-H11

DD 25.2 Tue 12:20 DD-H11 Belastungstrajektorie in der Studieneingangsphase Physik — • SIMON Z. LAHme, Larissa Hahn, Ronja Langendorf, Jasper O. Cirkel, Susanne Schnei-DER und PASCAL KLEIN — Universität Göttingen, Deutschland

Bisherige Forschungsergebnisse zeigen, dass die Studieneingangsphase Physik für Studierende mit großen Herausforderungen fachlicher, überfachlicher, administrativer sowie persönlicher Natur verbunden ist. Diese können zu individuellen Belastungsempfindungen und bei unzureichender Regulation zum Studienabbruch führen. Bisherige Studien betonen die Bedeutung detaillierter Kenntnisse über die Quellen der wahrgenommenen Belastung sowie ihre Intensität im zeitlichen Verlauf der Studieneingangsphase. Sie dienen als empirische Grundlage für die Identifikation von Faktoren und Zeitintervallen, um vonseiten der Universität geeignete Innovationen und Gegenmaßnahmen zu diskutieren und umzusetzen. Zu diesem Zweck wurde an der Universität Göttingen wöchentlich ein Kurzfragebogen eingesetzt, der durch standardisierte Fragen die subjektiv wahrgenommene Belastungsintensität, den geschätzten Workload sowie individuelle Belastungsquellen von Physikstudierenden im ersten Studienjahr ab Beginn des Vorkurses erfasst. Neben einer Beschreibung der zeitlichen Belastungsentwicklung der gesamten Kohorte können so auch individuelle Belastungstrajektorien analysiert sowie auffällige Teilgruppen identifiziert werden.

Im Vortrag werden die Belastungstrajektorien des ersten Studiensemesters sowie die Ergebnisse einer induktiven Kategorienbildung zu den studentischen Belastungsquellen präsentiert und diskutiert.

DD 25.3 Tue 12:40 DD-H11 Studentische Wahrnehmungen zur Studieneingangsphase in Physik-Studiengängen — •CHRISTINA LÜDERS und HEIDRUN HEINKE — RWTH Aachen University

Nach Klemm (2020) wird es in den nächsten Jahren zu einem dramatischen Anstieg des MINT-Lehrkräftemangels kommen. Studierende von MINT-Lehramtsstudiengängen müssen daher gezielt dabei unterstützt werden ihr Studium erfolgreich abzuschließen. An der RWTH Aachen werden die Studierenden im Lehramt Physik daher von Studienbeginn an in besonderem Maße unterstützt. Außerdem bietet die RWTH Aachen seit dem WS 2020/21 ergänzend

DD 26: Quantenphysik - Konzepte

Time: Tuesday 12:00-13:00

DD 26.1 Tue 12:00 DD-H12

How Excitations of the Vacuum form Mass — •HANS-OTTO CARMESIN — Gymnasium Athenaeum, Harsefelder Straße 40, 21680 Stade — Studienseminar Stade, Bahnhofstr. 5, 21682 Stade — Universität Bremen, Fachbereich 1, Postfach 330440, 28334 Bremen

The Higgs mechanism describes the formation of mass by an excitation of vacuum. However, that Higgs mechanism describes the vacuum in terms of a scalar field or potential $V(\Phi)$ consisting of a Φ^2 - term and a Φ^4 - term, but without characterizing the nature of the variables Φ or V. In particular, Higgs mechanism does not predict the mass m_H of the Higgs boson.

I summarize my covariant quantized field equation for the elements of spacetime. Using that equation, I present elementary objects underlying the physical vacuum as well as the variables Φ and V. In particular, and as a test, I derive the mass m_H in precise accordance with observation. Hereby, I do not use any fit, instead I derive all results from the universal constants G, c and h. I report about corresponding lessons and my experience in a research club.

See e. g. Carmesin, H.-O. (March 2021): Quanta of Spacetime Explain Observations, Dark Energy, Gravitation and Nonlocality. Berlin: Verlag Dr. Köster, Carmesin, H.-O. (August 2021): Cosmological and Elementary Particles Explained by Quantum Gravity. Berlin: Verlag Dr. Köster, (Carmesin, H.-O. (October 2021): The Elementary Charge Explained by Quantum Gravity. Berlin: Verlag Dr. Köster.

DD 26.2 Tue 12:20 DD-H12

How do scientists explain basic concepts in quantum physics? — •STINA SCHEER and GUNNAR FRIEGE — Institut für Didaktik der Mathematik und Physik, Leibniz Universität Hannover zum klassischen sechssemestrigen Bachelorstudiengang Physik einen Reformstudiengang Physik Plus mit acht Semestern Regelstudienzeit an. Hier können in den ersten Semestern zusätzliche Veranstaltungen angeboten werden, die den Übergang von der Schule zur Hochschule erleichtern. Seit dem WS 2019/20 wurden an der RWTH Aachen Studierende des ersten Studienjahres verschiedener Physik-Studiengänge in Fragebogen- und Interviewerhebungen zu Eingangsvoraussetzungen, Studienwahlmotiven und Erwartungen sowie zu ihren Wahrnehmungen in der Studieneingangsphase befragt. In dem Vortrag werden zusammenfassende Ergebnisse zu diesen Erhebungen für die Studiengänge Physik, Physik Plus und Lehramt Physik vorgestellt.

Location: DD-H12

Quantum physics is often perceived as complicated, unintuitive and hard to explain. So how do experts in quantum physics explain its basic concepts, such as the uncertainty principle? In an ongoing study we asked young scientists doing research in quantum metrology to explain such concepts to an interested first year university student. Here we report our initial findings on content structure and explanatory elements used within these explanations. The design and evaluation are based on known concepts of explaining and expert-novice communication.

DD 26.3 Tue 12:40 DD-H12

Topologische Modelle für Ununterscheidbarkeit und Verschränkung – • STEFAN HEUSLER – WWU Münster

Durch die steigende Bedeutung von Quantentechnologien ergibt sich die Notwendigkeit, geeignete Modellierungen auch für komplexere Themen wie dem der Verschränkung für die Schule zu entwickeln und zu evaluieren.

In diesem Beitrag stellen wir zunächst unsere Verallgemeinerung eines topologischen Ansatzes von P. Dirac zur Beschreibung von Spin j=1/2 Zuständen ("Dirak-Gürtel") auf beliebige Spinzustände vor. Dieser Ansatz erlaubt es, Ununterscheidbarkeit und Verschränkung haptisch begreifbar zu machen.

Wir diskutieren kritisch die Tauglichkeit des Modells aus fachlicher und fachdidaktischer Sicht und vergleichen mit anderen Ansätzen zur Vermittlung des Konzepts der Verschränkung.

DD 27: außerschulisches Lernen - Metaperspektive

Time: Tuesday 16:15-16:55

DD 27.1 Tue 16:15 DD-H8

Bedeutung des non-formalen Lernens für die MINT-Bildung: Interviewstudie mit Stakeholdern — •KAI BLIESMER und MICHAEL KOMOREK — Carl von Ossietzky Universität Oldenburg

Die Vielfalt non-formaler MINT-Bildungsangebote belegt die Ausdifferenzierung der außerschulischen Bildungslandschaft, die es interessierten Menschen ermöglicht, sich im Sinne des free-choice learning (Falk & Dierking 1998) freiwillig mit MINT-Inhalten zu beschäftigen. Eine prosperierende außerschulische Bildungslandschaft ist dabei in besonderem Maße von der Unterstützung einflussreicher Personen (Stakeholder) aus Politik, Bildungswissenschaft, Fachdidaktik, Wirtschaft und Wissenschaftsjournalismus abhängig, die sich finanziell und/oder ideell für das non-formale Lernen einsetzen und es so mitgestalten. Doch woher stammt dieses Engagement und welche Bedeutung weisen Stakeholder non-formaler MINT-Bildung zu? Der Vortrag präsentiert eine Studie, in der Stakeholder in einem leitfadengestützten, teilstandardisierten Interview nach ihrer Einschätzung befragt werden, wie non-formale Angebote zur MINT-Bildung insgesamt beitragen; welche Rolle sie in Zukunft spielen werden; und welchen Kriterien erfolgreiche non-formale Lernangebote genügen müssen. Von Interesse ist auch, wie die Befragten das Image non-formaler MINT-Angebote in der Allgemeinbevölkerung einschätzen und was sie vorschlagen, um diese zu stärken und für Erwachsene attraktiver zu gestalten. Die Studienergebnisse sollen helfen, non-formale Lernangebote im MINT-Cluster AHOI_MINT (www.ahoimint.de) weiterzuentwickeln und Generalisierungen hinsichtlich non-formaler MINT-Bildung zu formulieren.

Location: DD-H8

DD 27.2 Tue 16:35 DD-H8

BMBF-MINT-Cluster AHOI_MINT: Erhebung der von Familien geäußerten Bedarfe an non-formalen MINT-Angeboten — •MICHAEL KOMOREK, IMKE AHRENHOLTZ, JONATHAN NABER, CHRISTIN SAJONS und KAI BLIESMER — Universität Oldenburg

Im vom BMBF geförderten MINT-Cluster AHOI_MINT (www.ahoi-mint.de) werden non-formale, außerschulische MINT-Bildungsangebote für 10-16jährige aufgeschlossen, vernetzt und beworben. In dieser Studie wird untersucht, welches Interesse und welcher Bedarf bei Kindern, Jugendlichen und Eltern an MINT-Lernangeboten besteht und von außerschulischen MINT-Lernorten oder von Schulen im AG-Bereich gedeckt werden könnten. In problemzentrierten qualitativen Interviews mit biografischen und episodischen Elementen berichten 10-16jährige bzw. ihre Eltern über Erfahrungen mit non-formalen Bildungsangeboten, über ihre Wünsche nach Inhalten und Formaten und darüber, unter welchen Bedingungen sie vorhandene oder neue Angebote im Sinne von 'free choice learning' nutzen oder nutzen würden. Herausgefordernd war es, eine Stichprobe optimalen Kontrasts zu ziehen und das Interview an die Voraussetzungen der Teilgruppen anzupassen. Im Vortrag werden Ergebnisse vorgestellt, mit denen die Angebote im Cluster verbessert werden können und die Generalisierungen hinsichtlich non-formaler MINT-Bildung erlauben. Zu erkennen sind z. B. die Wünsche nach differenzierter methodischer Öffnung von MINT-Angeboten und der adressatengerechten Integration von Problemlöseaufgaben.

DD 28: Geschichte der Physik und NoS

Time: Tuesday 16:15-16:55

DD 28.1 Tue 16:15 DD-H9

Wiederentdeckte astronomische Modelle aus dem frühen 19. Jahrhundert und ihre didaktischen Möglichkeiten — •OLAF KRETZER — Schul- und Volkssternwarte Suhl, Hoheloh 1, 98527 Suhl

Astronomie in der Schule - diese heute oft gestellte Forderung ist keine Erfindung des 20. Jahrhunderts. Die Vermittlung astronomischen Wissens und Inhalte lassen sich beispielsweise bis ins beginnende 17. Jahrhundert zurück verfolgen. Zumeist waren Schüler höherer Schulen das Zielklientel der Vermittlung. In der ersten Hälfte des 19. Jahrhunderts hielt in einigen Regionen die Astronomie an den Volksschulen Einzug. Stellvertretend soll hier Leben und Werk eines Pädagogen untersucht werden, der mit seinen Werken und vor allem mit seinen schulastronomischen Modellen Pionierarbeit geleistet hat. Am Beispiel von zwei seiner Modelle sollen die Möglichkeiten der Modelle im heutigen Schulunterricht aufgezeigt werden.

DD 28.2 Tue 16:35 DD-H9

Zombies im Labor: Das Beispiel des mechanischen Wärmeäquivalents – •PETER HEERING – Europa-Universität Flensburg

versitären Praktika, aber auch im Physikunterricht vornehmlich der Oberstufe wieder. Sei dies die Coulombsche Drehwaage, der Franck-Hertz Versuch oder aber der Millikansche Öltröpfchenversuch; bereits die Benennung verweist auf einen (oftmals nicht weiter thematisierten) historischen Kontext. Es stellt sich aber zum einen die Frage, welche Relation die entsprechenden Versuchsaufbauten zu den historischen Apparaturen haben und zum anderen, was die Praxis mit diesen Aufbauten eigentlich mit Experimentieren zu tun hat. Anhand des Vergleichs der Arbeiten Joules zur Bestimmung des mechanischen Wärmeäquivalents, historischer Lehrversionen zu diesem Experiment und aktueller Angebote von Lehrmittelfirmen werde ich diskutieren, welche Probleme gerade im Hinblick auf die Entwicklung eines angemessenen Verständnisses im Bereich Nature of Science durch derartige Versuche geschaffen werden können.

Eine Reihe von Experimenten aus der Geschichte der Physik findet sich in uni-

DD 29: Lehr-Lernforschung - Lernermerkmale

Time: Tuesday 16:15-16:55

Time: Tuesday 16:15-16:55

Universität zu Köln, 50923 Köln

DD 29.1 Tue 16:15 DD-H10

DD 30.1 Tue 16:15 DD-H11

Empathisches oder systematisches Denken im Physikunterricht? Testentwicklung für Lernende der Sekundarstufe I — •Julia Welberg, Daniel Laumann und Susanne Heinicke — Institut für Didaktik der Physik, WWU Münster

Es ist hinreichend untersucht, dass Physikunterricht eher Jungen und weniger Mädchen interessiert. Im Schuljahr 2020/21 machten Mädchen jedoch 1/5 der Lernenden in Physikleistungskursen in NRW aus - physikinteressierte Mädchen gibt es also doch?!

Das biologische Geschlecht scheint demnach nicht das einzige Merkmal zu sein, welches zum Interesse an Physik betrachtet werden sollte. Studien von Zeyer und Kollegen haben gezeigt, dass Konstrukte der "Empathising-Systemising Theory" besser die Motivation Naturwissenschaften zu lernen beschreiben als das Geschlecht. Bisherige Erhebungen fanden allerdings ausschließlich ab der Sekundarstufe II statt. Von Interesse sind jedoch auch Zusammenhänge von empathischem und systematischem Denken in der Sekundarstufe I, da hier besonders im Bereich der Mittelstufe das Interesse an Physik(unterricht) stark sinkt. Für die Erfassung der Konstrukte lag jedoch bis jetzt kein geeignetes Erhebungsinstrument vor.

Im Beitrag werden die Entwicklungsschritte zu einer geeigneten Kurzskala zur Erfassung des empathischen und systematischen Denkens von Schülerinnen und

Junge Lehrende stehen beim Übergang von der ersten zur zweiten Phase vor

bedeutenden Herausforderungen wie dem Experimentieren mit Lernenden in

einer ungewohnten Umgebung, die Anpassung fachlicher Inhalte an das Niveau der Zielgruppe und, besonders ungewohnt, die Entwicklung einer emotionalen

Professionalität. Das Arbeitsfeld Schule ist gekennzeichnet von widersprüchli-

chen Handlungsanforderungen: Erziehen, Disziplinieren, Sanktionieren und Be-

werten auf der einen Seite, Beziehungslernen und emotionale Unterstützung auf

der anderen Seite, welche ohnehin schwierig zu vereinbaren sind. Hinzu kommt,

dass Physik Lehrende auch noch Vorbilder in wissenschaftlicher Neutralität und objektiver Distanz sein sollten. Für den Berufsanfänger sind dies zum Teil unver-

einbare Anforderungen, die zu einem Praxisschock erhöhtem Bewertungsdruck

und Schaffenskrisen führen können. Der Aufbau einer emotionalen Professiona-

lität ist daher aus unserer Sicht ein wichtiger Faktor zur Entwicklung der Resi-

lienz bei Physik Lehrenden. Im Beitrag werden wird die Definition emotionaler

Professionalität in geeigneten Quellen betrachtet und Ansätze zur längsschnitt-

lichen Untersuchung der Entwicklung emotionale Professionalität diskutiert.

Schülern der Sekundarstufe I vorgestellt und erste Ergebnisse in Bezug auf das Fachinteresse Physik gezeigt.

DD 29.2 Tue 16:35 DD-H10

Location: DD-H10

Growth Mindset (Cultures), Uni-Schul-Kooperationen und innovative Schulkonzepte – •LARS MÖHRING, JANNIK HENZE und ANDRÉ BRESGES — Inst. f. Phydid, Universität zu Köln

Die Kultur im Klassenraum aber auch der ganzen Schule hat einen großen Einfluss auf das Mindset von Lernenden und beinflusst deren Lernstrategien und -bilder stark. Menschen mit einem Fixed Mindset sehen Fähigkeiten als prädeterminiert an, Menschen mit einem Growth Mindset verstehen Fähigkeiten und Potential als veränderlich und entwickelbar. Ein Growth Mindset fördert Resilienz gegenüber Rückschlägen, hilft aus Fehlern zu lernen und minimiert Rückstände von vernachlässigten Lerngruppen. Ein Growth Mindset förderliches Klima im Klassenraum umfasst viele lernförderliche und inklusive Aspekte.

Die Inklusive Universitätsschule Köln weist ein modernes Schulkonzept auf und das Intelligenz- und physikbezogene-Mindset der Lernenden konnte im Zusammenhang mit dem Projekt "Zukunft gestalten mit Mensch und Technik" untersucht werden. Einflüsse des Projektes sowie unterschiedlicher digitaler Lernwerkzeuge und ein Vergleich zu "klassischen" Schulen werden im Vortrag präsentiert.

DD 30: Lehreraus- und Lehrerfortbildung - neue Ansätze

Location: DD-H11

DD 30.2 Tue 16:35 DD-H11

Emotionale Professionalität: Ein wichtiger Faktor zur Entwicklung von der Resilienz bei Physiklehrer*innen? – •BENJAMIN NIEHS¹, MICHEL NOETHLICHS² und ANDRÉ BRESGES³ – ¹Europaschule Bornheim, 52223 Bornheim – ²ZfSL Leverkusen, 51379 Leverkusen – ³Institut für Physikdidaktik,

Die Sicherung von Lernergebnissen ist ein zentraler Teil von Lehr-Lernprozessen im Physikunterricht und bezweckt eine nachhaltige Integration des Lernzugewinns in das bereits bestehende Wissensnetz von Schüler:innen. Bisher findet die Planung und Gestaltung von Lernergebnissicherungen im Physikunterricht jedoch nur sehr wenig Beachtung in der fachdidaktischen Literatur. Zudem gibt es kein einheitliches Verständnis darüber, was unter der Sicherung von Lernergebnissen verstanden wird und welche Bedeutung dieser in der konkreten Unterrichtsgestaltung sowie in der Ausbildung zukünftiger Lehrkräfte zugeschrieben wird. Da Physikfachleiter:innen die Vorstellungen und Unterrichtsplanungen von angehenden Physiklehrkräften im Referendariat maßgeblich prägen, wurde im Rahmen einer Masterarbeit an der Johannes Gutenberg-Universität Mainz die Bedeutung der Lernergebnissicherung in der zweiten Ausbildungsphase Physik untersucht. Dazu wurden vier leitfadengestützte, qualitative Expert:inneninterviews mit rheinland-pfälzischen Physikfachleiter:innen durchgeführt und mit MAXQDA kategorienbasiert ausgewertet. Im Vortrag werden die Ergebnisse der durchgeführten Interviewstudie zur Bedeutung der Lernergebnissicherung im Physikunterricht vorgestellt.

DD 31: Praktika und neue Praktikumsversuche

Time: Tuesday 16:15-16:55

DD 31.1 Tue 16:15 DD-H12

Warum ist die Gerade krumm? Messdaten interpretieren lernen am Beispiel des Stefan-Boltzmann-Gesetzes — •MICHAEL DAAM¹, FABIENNE MÜLLER¹, ANTJE BERGMANN¹, CARSTEN ROCKSTUHL¹ und RONNY NAWRODT² — ¹Institut für Theoretische Festkörperphysik, Karlsruher Institut für Technologie — ²Physik und ihre Didaktik, Universität Stuttgart

Ein typischer Versuch im physikalischen Anfängerpraktikum vieler Universitäten ist die Verifizierung des Stefan-Boltzmann-Gesetzes mit einer Thermosäule und einem näherungsweise schwarzen Strahler. In unserem Aufbau, mit einem elektrisch geheizten Leslie-Würfel, passen die Messwerte nur auf den ersten Blick zur Theorie. Wird die gemessene Strahlungsleistung über der vierten Potenz der Temperatur aufgetragen, ergibt sich anstatt der vorhergesagten Geraden eine leicht gekrümmte Kurve. Obwohl diese Abweichung offenbar nicht statistischer Natur ist, gehen Studierende in der Regel darüber hinweg und sehen in ihren Messwerten eine Bestätigung des Stefan-Boltzmann-Gesetzes.

Wir stellen eine Umsetzung des Experiments vor, die den Studierenden ihren unkritischen Umgang mit den Messwerten und dem zugrunde liegenden Modell vor Augen führt, Methoden für die sorgsame Interpretation von Messergebnissen bereitstellt und eine Gelegenheit bietet, diese anzuwenden. Im Laufe des Versuchs sollen die Studierenden die Diskrepanz zwischen der Vorhersage und der Messung einsehen, mit selbst ausgedachten Modifikationen des Messaufbaus beseitigen und schließlich Daten aufnehmen, die im Einklang mit dem Stefan-Boltzmann-Gesetz stehen.

Location: DD-H12

Tuesday

DD 31.2 Tue 16:35 DD-H12

Optische Kohärenztomographie im Praktikum — KAI PIEPER¹, ANTJE BERGMANN¹, MARIAN CHRISTNER¹, CARSTEN ROCKSTUHL¹ und •JENS KÜCHENMEISTER² — ¹Institut für Theoretische Festkörperphysik, KIT — ²Thorlabs GmbH

Die optische Kohärenztomographie ist ein bildgebendes Verfahren, bei dem Grenzflächen im Volumen einer Probe vermessen werden und so ein dreidimensionales Bild der betrachteten Struktur ergeben. Sie findet seit der Entwicklung in den frühen 1990er Jahren beispielsweise in der Augenheilkunde ihre Anwendung und bietet dabei die Möglichkeit, nichtinvasiv Aufnahmen der Netzhaut durchzuführen. Im Wesentlichen wird im hier vorgestellten optischen Kohärenztomographen ein Weißlicht-Michelson-Interferometer mit einem Mikroskop kombiniert. Einer der Spiegel im Interferometer wird durch eine Probe ersetzt, welche reflektierende Grenzflächen im Inneren besitzt. Durch die Beleuchtung mit weißem Licht kommt es zur Interferenz, wenn eine der Grenzflächen der Probe ungefähr den gleichen Abstand zum Strahlteiler einnimmt wie der Referenzspiegel. Wird die Probe in axialer Richtung bewegt, können diese Aufnahmen mit einem Mikroskop vergrößert und schließlich mit einer Kamera aufgenommen werden. Eine geeignete Nachverarbeitung ermöglicht es, aus den aufgenommenen Messdaten ein dreidimensionales Bild der Probe mit ihren Tiefeninformationen zu rekonstruieren. Mit dem hier vorgestellten Aufbau kann dank der Beschränkung auf die essenziellen Komponenten das Funktionsprinzip der optischen Kohärenztomographie anschaulich und leicht zugänglich in einem Praktikumsversuch für Studierende dargestellt werden.

DD 32: Postersession 2: Astronomie

Time: Tuesday 17:00-18:00

DD 32.1 Tue 17:00 P

AR-Lineale: Astronomie und Planeten im Klassenzimmer — •ALEXANDER ARSHEWIZKIJ, ALEXANDER PUSCH und MALTE UBBEN — Wilhelm-Klemm-Straße 10

Die astronomische Bildung kommt mit allerlei bekannten Tücken daher. Eine bekannte Problematik sind fehlende Stützpunktvorstellungen in Bezug auf Größenordnungen verschiedener Körper im Sonnensystem. Dieser Beitrag stellt eine einfache AR-Anwendung vor, mit der SuS sich eigene Größenverständnisse aufbauen können, indem sie "AR-Lineale" nutzen, um Vergleiche mit Alltagsgegenständen durchzuführen. Dabei werden nicht nur die acht Planeten thematisiert, auch Zwergplaneten und ihre Monde werden mit ihren Abständen und Größen durch die Applikation vermittelt.

DD 32.2 Tue 17:00 P

Lernschwierigkeiten im Umgang mit dem Hertzsprung-Russell-Diagramm — •RONJA LANGENDORF, SUSANNE SCHNEIDER und PASCAL KLEIN — Universität Göttingen

Das Hertzsprung-Russell-Diagramm (HRD) gilt als wichtiges Werkzeug der

DD 33: Postersession 2: Außerschulisches Lernen

Time: Tuesday 17:00-18:00

DD 33.1 Tue 17:00 P

Science Gateway: Zukünftige SchülerInnenlabore am CERN — •JULIA WOI-THE, PATRICK THILL und SASCHA SCHMELING — CERN, Genf, Schweiz Am Teilchennhysik-Labor CERN in Genf Schweiz wird derzeit ein neu-

Am Teilchenphysik-Labor CERN in Genf, Schweiz, wird derzeit ein neuer Science Center Komplex gebaut (*CERN Science Gateway* sciencegateway.cern), welcher auch zwei SchülerInnenlabore beherbergen wird. In diesen Laboren werden Lernende ab 5 Jahren ihre wissenschaftliche Neugierde erkunden und lernen wissenschaftlich zu forschen. Im Rahmen des forschendentdeckenden Lernens interagieren Lernende aus aller Welt direkt mit Mitgliedern der wissenschaftlichen Gemeinschaft des CERN um einen Einblick in die Forschung, die Arbeitsweise und die Technologien des weltweit größten Teilchenphysik-Labors zu erhalten. Die geplanten Themengebiete umfassen dabei verschiedenste Aspekte, wie zum Beispiel die Detektion, Beschleunigung und Manipulation von Teilchen, Robotik, (Quantum) Computing, technische Herausforderungen der Hochenergiephysik, oder medizinische Anwendungen der Teilchenphysik. In diesem Beitrag werden die Konzeption der Labore und einige der geplanten Lernaktivitäten vorgestellt. Wir freuen uns insbesondere über Astrophysik und ist daher ein zentraler Bestandteil in universitären Astrophysikveranstaltungen. Es unterstützt Lernende dabei, die physikalischen Zusammenhänge zwischen den Eigenschaften von Sternen zu verstehen und den Prozess der Sternentwicklung zu beschreiben. Die Extraktion von Informationen z.B. zur Leuchtkraft ist dabei Voraussetzung für einen zielführenden Umgang mit dem HRD. Praxiserfahrungen deuten jedoch an, dass dies seitens der Studierenden häufig mit Lernschwierigkeiten verbunden ist. Diese sind vermutlich auf die visuelle Komplexität des HRD zurückzuführen, die eine hohe Aufmerksamkeit auf und sorgfältige Auseinandersetzung mit relevanten Diagrammelementen erfordert. Bisher mangelt es jedoch an empirischen Untersuchungen zu den konkreten Ursachen dieser Schwierigkeiten. Im Beitrag werden daher ausgewählte Ergebnisse einer Studie vorgestellt, in der 35 Physikstudierende 14 offene Aufgaben zum HRD bearbeiteten. Eye-Tracking-Daten ermöglichen eine Analyse der visuellen Aufmerksamkeit während der Aufgabenbearbeitung. In Kombination mit retrospektiven Interviews können mögliche Lernschwierigkeiten aufgedeckt werden. Hiermit geht das langfristige Ziel einher, Physikstudierenden das HRD durch zielgerichtete Instruktionen in einer digitalen Lehr-Lern-Umgebung besser zugänglich zu machen.

Location: P

über Ideen, die wir zusammen mit anderen Schülerlaboren entwickeln könnten (sciencegateway.labs@cern.ch).

DD 33.2 Tue 17:00 P

MINT-Begeisterung wecken, fördern und halten - zwei Beispiele für außerschulische Lernangebote — •Anne Geese, DINA AL KHARABSHEH und RAINER MÜLLER — TU Braunschweig, Institut für Fachdidaktik der Naturwissenschaften

Außerschulische MINT-Aktivitäten sind aktuell wenig populär. Sie können aber für mehr Bildungsgerechtigkeit sorgen, indem sie sich an benachteiligte Gruppen wenden. Dies sind z.B. Mädchen und junge Frauen, deren Weg in die MINT-Fächer durch gesellschaftliche Rollenzuschreibungen erschwert wird, aber auch Kinder aus bildungsfernen Regionen, die außerhalb der Schule keine Berührungspunkte mit naturwissenschaftlichen Themen haben.

Wir stellen zwei Beispiele für außerschulische Lernangebote vor, die genau diese beiden Zielgruppen ansprechen: Der Forschungsclub changING richtet sich an Mädchen und junge Frauen und bietet seit 2019 in zweiwöchentlichen Gruppentreffen Einblicke in die Luftfahrt der Zukunft. Angekoppelt an den Exzellenzcluster zur nachhaltigen Luftfahrt SE*A möchte er jungen Frauen den

211

Location: P

Weg in ein ingenieurwissenschaftliches Studium ebnen. Die MINT Liga startet im Sommer 2022 und ist ein vom BMBF geförderter MINT-Cluster. An Ankerpunkten wie Jugendzentren in sozial benachteiligten Quartieren finden regelmäßig MINT-Angebote statt, wobei die verschiedenen Ankerpunkte in einer Liga gegeneinander antreten und am Ende einen MINT-Meister küren.

Wir stellen diese beiden MINT-Angebote vor und betrachten ihre Gelingensbedingungen.

DD 33.3 Tue 17:00 P

Fast lichtschnell durch die Stadt — •STEPHAN PREISS — Universität Hildesheim, Hildesheim, Germany

In Einsteins Spezieller Relativitätstheorie treten die ungewöhnlichen Effekte der Zeitdilatation und Längenkontraktion auf. Diese meist durch Gedankenexperimente und Diagramme behandelten Erscheinungen stellen eine große kognitive Herausforderung für Lernende, die sich zum ersten Mal mit dieser Thematik auseinandersetzen, dar. Immerhin müssen die alltäglichen Konzepte von Länge und Zeitdauer völlig neu beurteilt werden.

In diesem Beitrag stellen wir eine Unterrichtseinheit vor, die unter Benutzung interaktiver Computersimulationen einen "experimentellen" Zugang zu den Effekten der Speziellen Relativitätstheorie bietet. Diese Einheit, wurde in zahlreichen Veranstaltungen des Schülerlabors *Raumzeitwerkstatt* mit Schülergruppen der Klassenstufen 10 bis 13 erfolgreich eingesetzt.

Um die anfängliche Frage, wie die Umgebung bei einer Bewegung mit nahezu Lichtgeschwindigkeit aussieht, zu beantworten, stellen wir eine virtuelle Welt mit drastisch reduzierter Lichtgeschwindigkeit (c = 1 m/s), in der sich der Nutzer frei bewegen kann, zur Verfügung. In diesem Labor können die Lernenden dann durch Beobachtung schnell bewegter Objekte und eigener Flüge MessunDD 33.4 Tue 17:00 P

Tuesday

 $\begin{array}{rcl} \textbf{MILeNa} & & \textbf{MINT-Lehrkräfte-Nachwuchsförderung} & weitergedacht & \\ \bullet \textbf{Christina} & Lüders^1, \ \textbf{Carsten} & Kaus^1, \ \textbf{Christian} & Salinga^1, \ \textbf{Rebecca} \\ \textbf{Grandrath}^2, \ \textbf{Sebastian} & HÜmbert-Schnurr^2, \ \textbf{Franziska} & Klautke^3, \ \textbf{Sebastian} & Keller^3, \ \textbf{Heike} & Theyssen^3, \ \textbf{Amélie} & Tessartz^4, \ \textbf{Jan} & Heysel^4, \\ \textbf{Ulrich} & Blum^4 & und \ \textbf{Heidrun} & Heinke^1 & {}^1\text{RWTH} & Aachen \ University \\ & {}^2\text{Bergische} & Universität \ Wuppertal & {}^3\text{Universität} \ Duisburg-Essen & {}^4\text{Rheinische Friedrich-Wilhelms-Universität} \ \textbf{Bonn} \end{array}$

MILeNa ist ein Programm zur MINT-Lehrkräfte-Nachwuchsförderung, welches seit 2013 existiert und dem sich dramatisch verschärfenden MINT-Lehrkräftemangel entgegenwirken soll. Interessierte Oberstufenschüler:innen werden in der Phase der Berufsentscheidung langfristig begleitet und erhalten tiefe Einblicke in das Berufsfeld einer MINT-Lehrkraft mit dem Ziel eine fundierte Studienentscheidung zu unterstützen. Seit 2017 wurde das Programm in einer modifizierten Variante zunächst von der RWTH Aachen angeboten und seit 2019 auf weitere Hochschulen ausgeweitet. Dabei werden die Programmbestandteile in einem Präsenz-, Hybrid- oder Onlineformat umgesetzt. Zur Grundstruktur gehören neben einer Auftakt- und Abschlussveranstaltung ein Basis-Workshop, welcher Grundlagen zur Unterrichtsgestaltung und -durchführung vermittelt. In anschließenden schulischen Angeboten erproben sich die Schüler:innen in der Lehrendenrolle. Eine Beteiligung vieler Hochschulen ermöglicht neben wachsenden Teilnehmer:innenzahlen außerdem vielfältige (digitale) Zusatzangebote zu unterrichtlichen Themenstellungen.

DD 34: Postersession 2: Hochschuldidaktik

Time: Tuesday 17:00-18:00

DD 34.1 Tue 17:00 P

Vektorielle Feldkonzepte verstehen durch Zeichnen? Erste Wirksamkeitsuntersuchungen — •LARISSA HAHN und PASCAL KLEIN — Universität Göttingen, Deutschland

Um vektorielle Feldkonzepte wie die Divergenz in physikalischen Kontexten anzuwenden, ist ein konzeptionelles Verständnis notwendig. Bisherige Forschungsergebnisse zeigen jedoch studentische Schwierigkeiten im Umgang mit Vektorfeldkonzepten. Im Einklang mit lerntheoretischen Erkenntnissen fordern Fachdidaktiker:innen daher die Entwicklung zielgerichteter Lehr-/Lernmaterialien, die den Einsatz und die Koordination multipler Repräsentationen fokussieren und unterstützen. In diesem Zusammenhang hat in den letzten Jahren die Technik des Zeichnens zunehmend an Aufmerksamkeit gewonnen. Theoretischen Überlegungen und früheren Studien zufolge hat Zeichnen das Potential, eine tiefe Auseinandersetzung mit dem Lerninhalt durch eine mentale Entlastung zu unterstützen und so den Wissensaufbau, insbesondere komplexer Konzepte, zu fördern. Vor diesem Hintergrund wurden Lehr-/Lernmaterialien entwickelt, die einen visuellen Zugang zu vektoriellen Feldkonzepten anhand von Vektorfelddiagrammen ermöglichen und unterstützende Zeichenaktivitäten, z.B. bei der Interpretation von Richtungsableitungen, integrieren. Die Analyse der kognitiven Prozesse im Zuge der Bearbeitung des Lehr-/Lernmaterials und beim zugehörigen Problemlösen wird hierbei durch Eye-Tracking unterstützt. Neben einem Überblick zum aktuellen Projektstand bezüglich bisheriger Materialentwicklungen und -wirksamkeitsuntersuchungen werden außerdem zukünftige Ziele und Ideen vorgestellt.

DD 34.2 Tue 17:00 P

Lehramtsspezifischer Professionsbezug in Fachveranstaltungen •BENEDIKT GOTTSCHLICH und JAN-PHILIPP BURDE — Universität Tübingen Bereits im Jahr 2006 forderte die DPG die stärkere Berücksichtigung professionsbezogener Anforderungen an zukünftige Physiklehrkräfte im Rahmen ihrer fachlichen universitären Ausbildung. Eine Analyse der Modulhandbücher der Physik-Lehramtsstudiengänge in Deutschland zeigt, dass Fachvorlesungen in höheren Semestern tatsächlich häufig separat für Lehramtsstudierende angeboten werden. Im Gegensatz dazu belegen Lehramtsstudierende jedoch üblicherweise einen Großteil der Vorlesungen zur Klassischen Physik gemeinsam mit Fachstudierenden. Es stellt sich daher die Frage, auf welche Art und Weise auch im Rahmen dieser Veranstaltungen ein lehramtsbezogener Professionsbezug hergestellt wird. Eine engere Verzahnung fachlicher und didaktischer Aspekte bereits in der Studieneingangsphase erscheint nicht nur vor dem Hintergrund der beschriebenen Forderungen der DPG wünschenswert, sondern kann zudem einen Beitrag zur Verringerung von Frustration und zur Senkung der Abbruchquote in der ersten Phase des Physik-Lehramtsstudiums leisten. Im Beitrag werden die Ergebnisse einer aktuellen Umfrage unter Standorten in Deutschland vorgestellt, die ein gymnasiales Lehramtsstudium anbieten. Ziel war es hierbei, systematisch zu erheben, in welcher Form und in welchem Umfang ergänzende professionsbezogene Maßnahmen für Lehramtsstudierende im Rahmen gemeinsamer Fachveranstaltungen angeboten werden.

DD 34.3 Tue 17:00 P

Location: P

Prüfungen im Physikstudium: Aktuelle Hochschulpraxis bildungswissenschaftlich betrachtet — •M. LÄNGLE¹, D. KERN-MICHLER², S. BRACKERTZ³, S. PENGER³, A. SICH³, L. LEHMANN⁴ und C. KRONBERGER⁵ — ¹Universität Wien, Wien, Österreich — ²ZaPF e.V., Frankfurt, Deutschland — ³UZK, Köln, Deutschland — ⁴TU Dresden, Dresden, Deutschland — ⁵TU Wien, Wien, Österreich

Die Corona-Pandemie hat dazu geführt, dass viele Prüfungsformate adaptiert werden mussten oder ganz an ihre Grenzen gestoßen sind. Gleichzeitig wurden zahlreiche neue Prüfungsformate ausprobiert.

Auch wegen der Reduzierung der persönlichen Kontakte haben Prüfungen die Kultur des Studiums besonders geprägt.

All dies geschah typischerweise nur selten mit systematischem Bezug zu Forschung und bildungspolitischen Debatten in der allgemeinen Pädagogik. Letztere ist einerseits sehr fundiert, andererseits nimmt sie nur sehr selten spezifisch auf das Physikstudium Bezug.

Angesichts dessen werden alte und neue Prüfungsformate verschiedener Physikstudiengänge mit dem aktuellen Stand der pädagogischen Debatte konfrontiert. Besonders relevante Aspekte sind:

Prüfungen als

- Selektionselement von Lebenschancen
- Qualifizierungsnachweis
- kulturell prägendes Element des Unialltags und Stressfaktor
- Strukturierungs- und Feedbackinstrument

DD 34.4 Tue 17:00 P

Die WiMINT-AGs Mathematik und Physik — BRITTA SCHÜTTER-KERNDL¹, •MANUELA BOIN¹, BERND ODER², ACHIM BOGER³ und KARIN LUNDE¹ — ¹Technische Hochschule Ulm — ²Hochschule Aalen — ³Gewerbliche Schule Schwäbisch Gmünd

Die Arbeitsgruppe cosh (Cooperation Schule-Hochschule) setzt sich für eine intensive Zusammenarbeit zwischen Schulen und Hochschulen in Baden-Württemberg ein. LehrerInnen erarbeiten gemeinsam mit ProfessorInnen Möglichkeiten, SchülerInnen besser auf ein Hochschulstudium vorzubereiten (siehe www.cosh-bw.de).

Viele Erstsemester im WiMINT (Wirtschaft, Mathematik, Informatik, Naturwissenschaft und Technik)-Bereich haben zu Studienbeginn fachliche Probleme. Ein Ziel von cosh ist es deshalb, Angebote für Studieninteressierte zu entwickeln, die diese Anfangsschwierigkeiten mindern sollen.

WiMINT-AGs sind studentische Tutorien, die in Kooperation zwischen einer Hochschule und einer Schule stattfinden. Didaktisch geschulte Studierende Das Verbundprojekt cosh wurde im Rahmen des Fonds Erfolgreich Studieren in Baden-Württemberg (FESt-BW, Förderlinie 4 *Eignung und Auswahl*) gefördert. Die Erarbeitung eines Konzeptes für die WiMINT-AG Physik, die Erstellung der Materialien und die Erprobung in einem Pilotprojekt waren ein Schwerpunkt der Förderung.

DD 34.5 Tue 17:00 P Auswirkungen der Lehr-Lernüberzeugungen studentischer Tutor*innen — •ROBIN DEXHEIMER-REUTER — Didaktik der Physik, Technische Universität Darmstadt

Von studentischen Tutor*innen geleitete Übungen stellen an vielen Universitäten einen wichtigen Baustein der Lehre dar, welcher empirisch jedoch noch wenig untersucht ist. Insbesondere zum Einfluss der Lehr-Lernüberzeugungen von Tutor*innen erscheinen weitere Erhebungen lohnenswert, da sich diese auf das Handeln der Tutor*innen und hierüber auch auf den Erfolg der Studierenden auswirken können (Mediationsmodell). In der Informatik erwiesen sich die Überzeugungen von Tutor*innen als prädiktiv für die Wahrnehmung ihrer Kompetenz durch die Studierenden (Glathe, 2017). Im Bereich der universitären Physikübungen steht eine ähnliche Überprüfung zum Einfluss der Überzeugungen von Tutor*innen noch aus. Daher ist es Ziel des geplanten Projektes, das Mediationsmodell in diesem Kontext zu untersuchen. In einer Vorstudie wird erhoben, welche Methoden in der Tutor*innenqualifikation eine besondere Rolle spielen. Diese werden in der anschließenden Hauptstudie besonders berücksichtigt. Dabei wird die Qualität der Übungsleitung in Anlehnung an die COACTIV-Studie (Kunter, 2011) durch Befragung der Studierenden zur wahrgenommenen Gruppenführung, kognitiven Aktivierung und konstruktiven Unterstützung operationalisiert. Auf Seite der Studierenden wird neben dem Veranstaltungserfolg auch ihre Zufriedenheit mit der Übung erfasst. Auf dem Poster werden die Forschungsfragen und das entsprechende Design der ab 2022 geplanten Studie vorgestellt.

DD 34.6 Tue 17:00 P

Paderborner Studieneingangsphase Physik - Entwicklung von Unterstützungsmaterialien für die Theoretische Physik — •NILAB ABBAS, ANNA B. BAU-ER und PETER REINHOLD — Universität Paderborn, Deutschland

Physik Studiengänge weisen eine hohe Abbruchquote vor allem zu Beginn des Studiums auf. Als häufigste Ursache für den Abbruch werden inhaltliche Anforderungen genannt, die vor allem beim Bearbeiten von Aufgaben in Übungszetteln und beim Bestehen von Klausuren besonders herausfordernd wirken. Studien zeigen ebenfalls, dass die hierzu nötigen fachspezifischen Problemlösefähigkeiten nicht auf einem ausreichenden Niveau entwickelt werden. Als Reaktion hierauf wird an der Universität Paderborn eine abgestimmte Studieneingangsphase aus einem Guss gestaltet. Ein Fokus stellt die Theoretische Physik und der mit ihr einhergehende hohe Grad an Mathematisierung dar. Es werden im Rahmen eines Design-Based Research Ansatzes passgenaue Unterstützungsmaßnahmen für die Veranstaltung Theoretische Physik entwickelt und hinsichtlich der Gelingensbedingungen mit Fokus auf die Lernwirksamkeit evaluiert. Zusammen mit den Akteuren werden typische Herausforderungen analysiert. Daraus werden digitale Maßnahmen zur Vermittlung von Problemlösefähigkeiten (Erklärvideos, Worked-Out-Examples etc.) entwickelt und als Selbstlernmaterialien zur Verfügung gestellt. Die Lernwirksamkeit der Maßnahmen wird multiperspektivisch (Videografie, Fragebögen, Fachwissenstest) untersucht. Der Beitrag stellt das Untersuchungsdesign und die identifizierten Herausforderungen der Studierenden vor.

DD 34.7 Tue 17:00 P

Paderborner Studieneingangsphase Physik - Gestaltung einer Studieneingangsphase aus einem Guss — •ANNA B. BAUER und PETER REINHOLD — Didaktik der Physik, Universität Paderborn, Deutschland

Die Studieneingangsphase Physik stellt für die Studienanfänger einen komplexen Lernprozess mit vielfältigen Anforderungen auf fachlicher, Metakognitionsund Sozialisations-Ebene dar, der sie im Rahmen der akademischen Identitätsbildung mit einbezieht und prägt. Ziel des Projektes Paderborner Studieneingangsphase Physik ist die evidenzbasierte Gestaltung eines strukturierten Studieneinstiegs und einer in sich kohärent abgestimmten Studieneingangsphase aus einem Guss. Die Maßnahmen werden multiperspektivisch mit Hilfe bestehender Instrumente evaluiert. Die Implementation eines neuen Übungsformats (Präsenzübungen) in den Fachvorlesungen, die Integration einer zentralen Ansprechperson sowie die systematische Unterstützung der Studierenden im Bereich des selbstregulierten Lernens zeigen positive Effekte in einer erhöhten Teilnahmequote sowie Zufriedenheit der Studierenden mit den Veranstaltungen, einem aktiveren Arbeitsverhalten sowie einer höheren Bestehensquote der Klausur im ersten Semester. In dem Beitrag werden die Gelingensbedingungen bzw. Strukturen für eine wirksame Zusammenarbeit von Fachdidaktik und Fachwissenschaft am Beispiel der Überarbeitung Studieneingangsphase im Rahmen einer community of practice (Aktionsforschung), von der Wirksamkeit der bisherigen Implementierung sowie von den nächsten Schritten berichtet.

DD 34.8 Tue 17:00 P

Das Projekt: Flexible Zusammenführung der Lehrkonzepte und Lehrmaterialien für den Experimental-Physik-Unterricht an der Universität Kiel — •IRINA SCHNEIDER — Institut für Experimentelle und Angewandte Physik Christian-Albrechts-Universität zu Kiel

Der Experimental-Physik-Unterricht ist manchmal nicht interessant für Studierende und das behindert auch das Verständnis von den Grundlagen des Faches. Besonders betroffen sind hier die Nebenfachstudenten, die oft wenig Vorlesungsund Praktika-Stunden haben. Ziel des Projektes ist, den Unterricht verständlicher und abwechslungsreicher zu machen und die Studierenden zu motivieren. Dafür werden die unterschiedlichen Lehrkonzepte, Lehrmaterialien usw. zusammengeführt zu einer flexiblen Sammlung, die die Lehrenden nutzen und auch weiterentwickeln können. Auch der Erfahrungsaustausch zwischen den Lehrenden ist in dem Rahmen des Projektes hier wichtig. Das Projekt können auch die Studierenden mit den Dozierenden gemeinsam weiterentwickeln.

DD 34.9 Tue 17:00 P

Einsatz von multiplen Repräsentationsformen zur qualitativen Beschreibung realer Phänomene der Fluiddynamik — •CHRISTIAN RABE, VINCENT DREWS, LARISSA HAHN und PASCAL KLEIN — Universität Göttingen, Deutschland

Bei der qualitativen Beschreibung realer Phänomene in der Fluiddynamik zeigen sich Schwierigkeiten für viele Lernende. In zwei getrennten Studien wurden das studentische Verständnis der Kontinuitätsgleichung in Flüssigkeitsströmungen sowie der aerodynamische Magnus-Effekt untersucht. Beiden Themen ist die Verwendung multipler Repräsentationsformen (Formeln, Stromlinien, Vektorfelder) gemein, die zum Lernen konstruiert und aufeinander bezogen werden müssen. In der Physikdidaktik ist bekannt, dass sich die Verwendung von multiplen Repräsentationsformen in vielen Fällen als lernförderlich erweisen kann; eine kohärente Übersetzung zwischen realem Phänomen und Repräsentationsform allerdings auch Schwierigkeiten bereitet. Im Stil des Design-Based Research wurden Lehr-/Lernmaterialien entwickelt, die einen multi-repräsentationalen Zugang zu den Themen ermöglichen und schon frühzeitig vektorielle Feldkonzepte adressieren. Durch Akzeptanzbefragungen (N >10) und den Einsatz von Testinventaren (N > 100) konnten Lerngelegenheiten und -schwierigkeiten identifiziert werden, die der Weiterentwicklung der Materialien dienen. Die Fluiddynamik erwies sich dabei als ein äußerst reichhaltiges Feld für physikdidaktische Forschungsarbeiten mit hoher Anschlussfähigkeit an die Elektrodynamik.

DD 34.10 Tue 17:00 H

Ausbildung in drei Dimensionen: Theorie, Praxis, Forschung — •JOHANNES LHOTZKY¹, NADINE BASTON², KLAUS WENDT¹ und MARIUS HARRING² — ¹Johannes Gutenberg-Universität Mainz, Institut für Physik — ²Johannes Gutenberg-Universität Mainz, Institut für Erziehungswissenschaften Im Projekt "Lehr-Lern-Forschungslabore als Orte vertieften Lernens: Das Main-

zer Modell kooperativer Lehrerbildung" der BMBF Qualitätsoffensive Lehrerbildung partizipieren Studierende des gymnasialen Lehramtsstudiengangs im Teilprojekt Physik an Lehr-Lern-Labor-Seminaren. Diese sind an bildungswissenschaftliche Veranstaltungen im Bachelor- und Mastertsudiengang gekoppelt. In der fachdidaktischen Lehrveranstaltung treffen drei zentrale und auch empirisch geforderte Aspekte der universitären Lehramtsausbildung an einem Ort aufeinander: Theorieinput, Praxisphase und Forschungsperspektive. Die Veranstaltung umfasst inhaltlich praxisrelevante Schwerpunkte der Unterrichtsplanung, deren Umsetzung als auch zentrale theoretische und methodische Schwerpunkte: kognitive Aktivierung, vertieftes Lernen sowie indikatorbasierte Videoanalyse. Mithilfe der Videoanalyse gelingt eine enge Verzahnung von Forschung und Theorie mit konkretem Praxisbezug. Zur Unterstützung des Analyseprozesses wird eine im Projektkontext entwickelte, interaktive Lehr-Lern-Plattform eingesetzt. Der kooperativ gestaltete Beitrag präsentiert und diskutiert die gemeinsame theoretische Verortung, das Seminarkonzept, konkrete Studierendenergebnisse, sowie die Resultate der begleitenden summativen Evaluation.

DD 34.11 Tue 17:00 P

Empirische Überprüfung der Wirksamkeit eines Strategietrainings im Rahmen eines Physikmoduls — •Katja Plicht¹, Hendrik Härtig² und Alexandra Dorschu¹ — ¹Hochschule Ruhr West, Mülheim an der Ruhr — ²Universität Duisburg-Essen

Der Übergang von der Schule zur Hochschule stellt sich besonders in den Naturund Ingenieurwissenschaften als problematisch dar und führt noch immer häufig zum Studienabbruch (Heublein, 2018). Dabei zeigt sich, dass ein notwendiges tieferes Verständnis physikalischer Konzepte und die Befähigung zum selbständigen Problemlösen oftmals nicht ausreichend entwickelt ist (Schecker & Klieme, 2001).

Der Experten-Novizen-Vergleich zeigt darüber hinaus, dass das erfolgreiche Problemlösen im Bereich der Physik einerseits das Verständnis von Tiefenstrukturen und andererseits das Erlernen von Problemschemata und Heuristiken beinhaltet (Friege, 2003). Daraus ergibt sich ein Bedarf nach einer systematischen Entwicklung von Problemlösestrategien entgegen dem oftmals verbreiteten plug-and-chuck-Verfahren, bei dem eine zur Zielgröße passende Formel gesucht und weiterverwendet wird (Redish, 2006).

Mit einem Fokus auf diese beiden Aspekte wurde ein Strategietraining zur Förderung der Problemlösekompetenz entwickelt. Dieses wurde im Rahmen der Physikübung im Studiengang Maschinenbau eingesetzt und im Kontrollgruppendesign evaluiert. Es werden das Konzept sowie erste Pilotierungsergebnisse präsentiert.

DD 34.12 Tue 17:00 P Quantenphysik und Astronomie - mehr als nur bunte Farben? — •TOBIAS REINSCH¹, LUKAS MACZEWSKY², HOLGER CARTARIUS³ und RONNY NAWRODT¹ — ¹Physik und ihre Didaktik, Universität Stuttgart, Pfaffenwaldring 57, 70569 Stuttgart — ²Experimentelle Festkörperphysik, Universität Rostock, Albert-Einstein-Str. 23, 18059 Rostock — ³AG Fachdidaktik der Physik und Astronomie, Friedrich-Schiller-Universität Jena, 07743 Jena

DD 35: Postersession 2: Neue / digitale Medien

Time: Tuesday 17:00-18:00

DD 35.1 Tue 17:00 P

Unterstützung eines qualitativen Verständnisses der elektromagnetischen Induktion durch Augmented Reality — •ROLAND BERGER¹ und PHILIPP LENSING² — ¹Universität Osnabrück — ²Hochschule Osnabrück Ein grundlegendes Verständnis der elektromagnetischen Induktion erfordert bei Lernenden die Einsicht, dass die zeitliche Änderung des Magnetfeldes in einer Leiterschleife eine elektrische Spannung induziert. Um dieses Grundprinzip möglichst deutlich zu machen, haben Erfmann und Berger ein qualitatives Unterrichtskonzept vorgeschlagen, welches auf dem Abzählen von Feldlinien basiert. Nach diesem Konzept wird die zeitliche Änderung des Magnetfeldes in einer Leiterschleife mithilfe eines Filmstreifens veranschaulicht. In dem Filmstreifen sind Bild für Bild Feldlinien eingezeichnet, sodass die zeitliche Änderung der Zahl der Feldlinien unmittelbar zu erkennen ist. Um dieses Feldlinienkonzept der elektromagnetischen Induktion zu unterstützen, haben wir eine App entwickelt, welche die Stärke des Magnetfeldes mit dem Sensor des verwendeten Tablets misst. Dem Prinzip von Augmented Reality folgend wird einer Induktionsspule innerhalb einer Helmholtzspule eine entsprechende Zahl virtueller Feldlinien überlagert. Mithilfe dieser App können somit zentrale Fehlvorstellungen unmittelbar adressiert werden. Beispielsweise ist entgegen der Erwartung vieler Schülerinnen und Schüler die induzierte Spannung in dem Moment maximal, in dem das Magnetfeld Null ist.

DD 35.2 Tue 17:00 P

Für alles eine App. Ein Buch mit Ideen für Physik mit dem Smartphone — •THOMAS WILHELM¹ und JOCHEN KUHN² — ¹Institut für Didaktik der Physik, Goethe-Universität Frankfurt — ²Physics Education Research Group, Department of Physics, University of Kaiserslautern

Welche Apps eignen sich, um Smartphones oder Tablets für physikalische Untersuchungen und Betrachtungen zu nutzen? Diese Frage stellen sich physikalisch Interessierte genauso wie Lehrkräfte. Auf dem Poster wird das Praxisbuch "Für alles eine App. Ideen für Physik mit dem Smartphone" vorgestellt. Es gibt einen schnellen und umfassenden Überblick über geeignete Apps und stellt anhand vieler Praxisbeispiele dar, wie man Smartphone und Tablet physikalisch nutzen kann.

Das Buch geht auf unterschiedlichste Anwendungen ein: Sie reichen von vorgefertigten Simulationen über physikalische Spiele bis hin zu Augmented Reality-Anwendungen. Zudem werden Apps vorgestellt, mit denen Messdaten mit den internen Sensoren oder externen Zusatzgeräten erfasst, von einer Datenbank abgerufen oder durch die Verwendung der Foto- und Videokamera gewonnen werden. In jedem einzelnen Abschnitt wird eine andere App kurz und überblicksweise vorgestellt und deren Verwendbarkeit für physikalische Untersuchungen in Schule und/oder Hochschule und/oder zur eigenen Unterhaltung an einem Beispiel erläutert. Zunächst gibt es eine App-Kurzbeschreibung, dann folgt die Beschreibung eines physikalischen Anwendungsbeispiels. Die Abschnitte sind nach klassischen Themenbereichen der Physik geordnet.

DD 35.3 Tue 17:00 P

Eine AR-Erweiterung des EPo-Konzepts zu einfachen Stromkreisen — •SASKIA RAUBER¹, JAN-PHILIPP BURDE¹, THOMAS WILHELM², MARTIN HOPF³, LIZA DOPATKA⁴, VERENA SPATZ⁴, THOMAS SCHUBATZKY⁵, CLAUDIA HAAGEN-SCHÜTZENHÖFER⁵ und LANA IVANJEK⁶ — ¹Universität Tübingen — ²Universität Frankfurt — ³Universität Wien — ⁴TU Darmstadt — ⁵Universität Graz — ⁶TU Dresden

Anders als der Name es vermuten lässt, stellen einfache Stromkreise für viele Schülerinnen und Schüler (SuS) sowie Lehrkräfte eine der größten Herausforderungen des Physikunterrichts in der Sek I dar. Insbesondere entwickeln SuS oftmals kein eigenständiges Spannungskonzept und verstehen nicht, dass die Spannung eine Differenzgröße darstellt. Der Einsatz von digitalen Medien wie bspw. Augmented Reality (AR) könnte dazu beitragen, die Motivation und das konDie Beobachtung des Sternenhimmels gehört nicht nur zu den ältesten Wissenschaften, sondern bietet noch immer einen interessanten Einstieg in moderne Fragestellungen der Physik für Schüler*innen und Studierende. So bietet die Analyse von Stern- und Sonnenspektren einen motivierenden und vielfältigen Einstieg in die Quantenphysik. In diesem Beitrag wird eine unterrichtstaugliche experimentelle Herangehensweise auf Basis einfacher spektroskopischer Messungen der Linien im Sonnenspektrum vorgestellt. Quantenphysikalische Grundkonzepte wie Spin, Wellen- und ferner Teilchencharakter lassen sich damit direkt experimentnah aus den Spektren ableiten. Zusätzlich können diskrete Übergänge als Lösungen der Schrödingergleichung hergeleitet werden. Aus der Beobachtung weit entfernter Prozesse in unserem Universum lassen sich also Rückschlüsse auf die mikroskopischen Prozesse quantenmechanischer Natur auf Atomebene schließen und umgekehrt.

zeptionelle Verständnis der SuS zu fördern. Vor diesem Hintergrund wurden zu einer Reihe von Übungsaufgaben der Unterrichtskonzeption "Eine Einführung in die Elektrizitätslehre mit Potential" diverse AR-Modelle von Stromkreisen erstellt. Diese lassen sich unkompliziert mit Hilfe von mobilen Endgeräten über QR-Codes aufrufen und erscheinen anschließend über den gedruckten Schaltplänen. Da in den AR-Modellen die Spannungsverhältnisse in den Stromkreisen mittels Farbkodierung visualisiert werden, können diese den SuS u.a. als Musterlösungen dienen. Auf dem Poster werden verschiedene AR-Modelle exemplarisch vorgestellt und ihr didaktischer Mehrwert für den Physikunterricht diskutiert.

DD 35.4 Tue 17:00 P Vorgehensweise von Schüler*innen bei der Nutzung von Software zur mathematischen Modellbildung und Videoanalyse — • JANNIS WEBER und THOMAS WILHELM — Institut für Didaktik der Physik, Goethe-Universität Frankfurt Mathematische Modellbildung und Videoanalyse sind zwei unterschiedliche Ansätze für das Erlernen und Vertiefen der Newton*schen Dynamik in der gymnasialen Oberstufe, die die Gemeinsamkeit haben, dass sie den Nutzer/die Nutzerin von der nötigen Mathematik entlasten und es damit ermöglichen sollen, reale und komplexe Bewegungen zu modellieren bzw. zu analysieren und Reibungskräfte bewusst zu thematisieren. Als Teil einer Gesamtstudie zum Einsatz von mathematischer Modellbildung und Videoanalyse wurden Bildschirmvideos mit Tonaufnahmen von Schülerzweiergruppen während der Arbeit mit der Software aufgenommen. Auf dem Poster werden neben den aufgetretenen Schwierigkeiten die Arbeitsweisen der Proband*innen bei der Nutzung der entsprechenden Software vorgestellt. Es wird zudem ein Vergleich zwischen erfolgreichen und weniger erfolgreichen Proband*innen gezogen, um Merkmale zu identifizieren, die die Wahrscheinlichkeit eines erfolgreichen Umgangs mit der Software und damit einem großen Lernzuwachs erhöhen. Aus diesen Beobachtungen werden Empfehlungen für eine Unterrichtsgestaltung bei Nutzung der entsprechenden Software abgeleitet.

DD 35.5 Tue 17:00 P

Vergleich von Videoanalyse-Apps auf Tablets — •VINIT SURI und THOMAS WILHELM — Institut für Didaktik der Physik, Goethe-Universität Frankfurt

Die Videoanalyse von Bewegungen ist im Physikunterricht bereits weit verbreitet. Für Schüler*innen ist es selbstverständlich, Videoaufnahmen zu nutzen, und sie verfügen über die Möglichkeit, jederzeit und überall digitale Videoclips aufzuzeichnen. So bietet sich im Physikunterricht der Einsatz von Videoanalyse an, um die Alltagswelt der Schüler*innen mit dem Mechanikunterricht zu verbinden. Besonders einfach ist die Videoanalyse auf mobilen Endgeräten. Für diese sogenannte mobile Videoanalyse gibt es bereits einige Videoanalyse-Apps für unterschiedliche Betriebssysteme, die sich in ihrer Bedienung und ihren Möglichkeiten zum Teil erheblich unterschieden, für die es aber bisher keinen systematischen Vergleich gab.

Fünf derzeit auf dem Markt verfügbare Videoanalyse-Apps für Tablets wurden systematisch getestet, kriteriengeleitet verglichen und hinsichtlich des Einsatzes im Physikunterricht bewertet, sodass den Lehrkräften die Entscheidungsfindung für eine geeignete App erleichtert wird. Das Poster stellt die fünf Apps vor und gibt einen Einblick in die Vergleichsergebnisse.

DD 35.6 Tue 17:00 P

3D-Druck-Spektrometer als Unterrichtsprojekt — •RIKE HÄUSSLER¹, ANTJE BERGMANN² und GÜNTER QUAST¹ — ¹Institut für Experimentelle Teilchenphysik, Karlsruher Institut für Technologie — ²Institut für Theoretische Festkörperphysik, Karlsruher Institut für Technologie

Das Spektroskop ist ein beliebtes Werkzeug, um Lernenden in niedrigen Klassenstufen die Aufteilung von weißem Licht in die Spektralfarben zu erläutern. Dafür gibt es verschiedene Projekte, die Spektroskope selber, z.B. mit Pappe,

Location: P

zu bauen. Für die Oberstufe kann das Spektrometer im Unterricht genutzt werden, um unterschiedliche Lichtquellen zu charakterisieren und die dazugehörigen Spektren auszumessen. Das in diesem Beitrag präsentierte Projekt beinhaltet das Konstruieren eines Spektrometers in einer CAD-Software. So kann es mit Hilfe der 3D-Druck Technologie realisiert werden. Des Weiteren werden mit dem Spektrometer und einer frei verfügbaren Software verschiedene Lichtquellen quantitativ ausgewertet. Hierfür dient eine handelsübliche Smartphone-Kamera als Objektiv und Sensor. Dadurch wird ein klassischer Schulversuch mit digitalen Medien und mit Hilfe von modernen Fertigungsverfahren neu aufbereitet und ist auch für den fächerübergreifenden Unterricht geeignet.

DD 35.7 Tue 17:00 P

Vergleich von Videoanalyseprogrammen für den Einsatz im Mechanik-Unterricht der Sekundarstufe I — FLORIAN BRÄUER, PETER RIEGER und •ANDREAS KAPS — Universität Leipzig, Fakultät für Physik und Geowissenschaften, Bereich Didaktik der Physik, Prager Straße 36, 04317 Leipzig

Zu zeitgemäßem modernen Messen im Physikunterricht gehört die Videoanalyse. Moderne Videoanalyseprogramme weisen dabei eine hohe Komplexität auf. Untersucht wurde, ob diese bei der Verwendung im Mechanik Unterricht der Sekundarstufe I lernhemmend sein können. Vorgestellt werden die Ergebnisse einer Kleingruppenuntersuchung (N = 16) mit Lernenden der neunten Klassenstufe zum Einsatz des Verfahrens der Videoanalyse. Auf Grundlage eines theoriebasierten Vergleichs bezüglich der Handhabbarkeit im Unterricht zwischen den drei gängigsten Analyseprogrammen wurde die Anwendung Tracker ausgewählt. Trotz des anspruchsvollen Programms konnte keine lernhemmende kognitive Belastung durch die Komplexität festgestellt werden. Zwischen der Abneigung gegenüber Routinehandlungen (klassische Messverfahren mit Lineal und Stoppuhr) und dem Interesse an modernen Messverfahren, wie Videoanalyse, konnte ein hoher positiver Zusammenhang mit r=0,62 gefunden werden. Basierend auf diesen Erkenntnissen wurden Implikationen für einen möglichen Einsatz im Physikunterricht der Sekundarstufe I abgeleitet.

DD 35.8 Tue 17:00 P

Physik-Erklärvideos: Einstellungen (angehender) Physiklehrkräfte — •LOTTE HAHN und THORID RABE — Martin-Luther-Universität Halle-Wittenberg, Didaktik der Physik

Die Erklärvideonutzung erlebte in den vergangenen Jahren einen deutlichen Anstieg. Allerdings zeigen erste Analysen, dass Erklärvideos zum Teil erhebliche fachliche und fachdidaktische Mängel an Erklärqualität aufweisen, die nachhaltigen Lernprozessen sogar entgegenwirken können (Krey & Rabe, 2021). Vor dem Hintergrund, dass Erklärvideos zunehmend prägenden Einfluss auf das Bild von Physik und Physiklernen haben werden, wird diesem Befund in einem Promotionsprojekt mit einer Analyse ausgewählter Erklärvideos weiter nachgegangen.

Außerdem werden Perspektiven und Einstellungen (zukünftiger) Physiklehrkräfte bezüglich Erklärvideos mittels Leitfadeninterviews erhoben. Die Herausarbeitung (expliziter) Einstellungen zu Physik-Erklärvideos erfolgt mittels qualitativer Inhaltsanalyse nach Mayring (2010). Ziel ist es weiterhin, Personen hinsichtlich ihrer Einstellungen und weiterer Merkmale zu typologisieren.

Im Poster werden das Vorgehen der Erklärvideoanalyse und das Forschungsdesign der qualitativen Erhebung zur Diskussion gestellt.

DD 35.9 Tue 17:00 P

Kontrastierend und vergleichend die Qualität von Erklärvideos beurteilen lernen - Methodisches Vorgehen — •DEBORAH MILWA¹ und RITA WODZINSKI² — ¹Universität Kassel — ²Universität Kassel

Aufgrund einer steigenden Popularität von Erklärvideos finden diese als Medium zunehmend Einzug in den Grundschulunterricht (Dorgerloh & Wolf, 2020). Um passende Erklärvideos für den Unterricht auszuwählen, benötigen Lehrkräfte ein tieferes Verständnis bezüglich der Qualität von Erklärvideos (Kulgemeyer, 2018). Folglich setzen sich Sachunterrichtsstudierende in einem Seminar mit Qualitätskriterien in Anlehnung an Kulgemeyer (2018) und Lipowsky & Pätzold (2020) auseinander und wenden diese auf Erklärvideos an. Dabei wird auf die Methode des Kontrastierens und Vergleichens zurückgegriffen, da empirische Ergebnisse zeigen, dass sich das Identifizieren von Gemeinsamkeiten und Unterschieden positiv auf das Lernen auswirkt (u.a. Alfieri, Nokes-Malach & Schunn, 2013). Im Sinne der Methode wird untersucht, wie sich die Verwendung von Videos unterschiedlicher Qualität auf die Kenntnis von Qualitätskriterien und ihre Anwendung bei der Analyse von Erklärvideos auswirkt. Um den Wissenszuwachs der Studierenden zu erfassen, beurteilen sie vor und nach dem Vergleich unterschiedlicher Videos ein vorgegebenes Erklärvideo. Das Poster stellt das methodische Vorgehen und das Pre-Post-Design der Studie vor. Zudem wird die aus wissenschaftlicher Literatur abgeleitete Wirkweise des Kontrastierens und Vergleichens auf das Anwenden der Qualitätskriterien auf Erklärvideos visuell dargelegt.

DD 35.10 Tue 17:00 P

Erstellung von interaktiven digitalen Experimenten für das Physiklernen — •NELSON FINKELMEYER, PETER RIEGER und HELENA FRANKE — Universität Leipzig, Fakultät für Physik und Geowissenschaften, Bereich Didaktik der Physik, Prager Straße 36, 04317 Leipzig Die Corona Pandemie hat über Maßnahmen der Kontaktbeschränkung häufig auch den Kontakt der Lernenden mit Experimenten reduziert. Naturwissenschaftlicher Unterricht ohne mit Experimenten in Berührung zu kommen bietet nicht genug Raum zur Anschauung und Kompetenzförderung der Lernenden. Um dieses Problem zu adressieren, wurde im Rahmen einer Staatsexamensarbeit für den schulischen Kontext eine eigene Interpretation interaktiver digitaler (Bildschirm-)Experimente entworfen und umgesetzt:

Als Motivation für die Lernenden wird im Vorfeld eine Challenge formuliert, die experimentell gelöst werden soll. Viele einzelne Videosequenzen wurden zu einer Pfadstruktur verknüpft und bieten verschiedene experimentelle Wege, die in Erfolg oder Misserfolg enden können. Die Lernenden können aktiv über eine Reihe von Auswahlmöglichkeiten entscheiden wie weiter experimentiert werden soll, um die Challenge zu lösen. Durch diese mediale Gestaltung ist es den Lernenden möglich, sich mit dem Experiment im Sinne des forschenden Lernens entdeckend auseinanderzusetzen.

Die Konzeption wird an Hand von einem Experiment zum Prinzip des Archimedes vorgestellt und Möglichkeiten und Herausforderungen zur Erstellung dieser Art des digitalen Experiments erläutert.

DD 35.11 Tue 17:00 P

PUMA: Web-AR-Techniken als Ergänzung des Physikunterrichts — •STEFAN KRAUS und THOMAS TREFZGER — Lehrstuhl für Physik und ihre Didaktik, Julius-Maximilians-Universität Würzburg

Schülerexperimente mit starken Lasern, radioaktiven Präparaten und extremen optischen Dichten? PUMA (**P**hysik**U**nterricht **Mit A**ugmentierung) stellt interessante Möglichkeiten zur Verfügung, unsere Welt anhand digitaler Hilfsmittel zu erweitern und zu verstehen. Zum einen als Unterstützung von Realexperimenten, zum anderen für Heimexperimente mit minimalem Materialaufwand. AR-Anwendungen sind meist mit der Installation einer eigenen App und daraus resultierenden Hürden für die Schülerinnen und Schüler verbunden. Web-AR-Anwendungen hingegen öffnen sich direkt im Browser des Geräts. Dieser ist auf Smartphones wie Tablet-PCs vorhanden und macht die App zudem unabhängig vom Betriebssystem der Nutzerinnen und Nutzer. Mit Blick auf den Physikunterricht soll hier zunächst beleuchtet werden, inwieweit Web-AR-Techniken mit den Features von nativen Apps (maßgeschneidert für iOS oder Android) mithalten können und welche Vorteile sich für den praktischen Einsatz ergeben. Dazu werden exemplarisch Anwendungen aus der geometrischen Optik präsentiert, die zum Ausprobieren einladen und weitere Perspektiven aufzeigen.

DD 35.12 Tue 17:00 P

Interactive application for visualizing 3D and 2+1D spacetime sector models in GR — •VASSILIOS MARAKIS — Institut für Physik, Universität Hildesheim, Universitätsplatz 1, 31131 Hildesheim

The movement in spacetime and its curvature are concepts of general relativity, which are not easy to grasp for beginners. An approach to visualizing curved spacetime is the introduction of sector models, which avoids the introduction of the mathematical necessities for a user. The sector model divides the coordinate space into blocks with euclidean geometry, where elemental mathematical knowledge is sufficient to understand the resulting visuals and interpret them qualitatively in the sense of general relativity. The developed application uses different metrics and scenarios like Schwarzschild or space of constant curvature to create these sectors and lets the user construct geodesics in spacetime or calculate curvatures on specific points in space.

DD 35.13 Tue 17:00 P

Einsatzmöglichkeiten der Satelliten-Box von phyphox für die schulische und universitäre Lehre — •Leo Bodewig¹, Dominik Dorsel², Dustin Kirwald¹, Sebastian Staacks², Christoph Stampfer² und Heidrun Heinke¹ — ¹RWTH Aachen University, I. Physikalisches Institut IA — ²RWTH Aachen University, II. Physikalisches Institut A

phyphox ist eine an der RWTH Aachen entwickelte App, die das Auslesen der internen Sensoren von Smartphones bzw. Tablets ermöglicht und so das physikalische Experiment in die Hände des Nutzers gibt. Um die Bandbreite an möglichen Experimenten zu erweitern, sind externe Sensorboxen entwickelt worden, die sich über Bluetooth mit der App koppeln und auslesen lassen. Diese nutzen verschiedene Sensoren und lassen sich somit zu unterschiedlichen thematischen Schwerpunkten variabel einsetzen. Für die Implementation in die schulische und universitäre Lehrpraxis müssen konkrete Experimente konzipiert und getestet werden und hierzu Arbeitsmaterialien ausgearbeitet und erprobt werden. Auf dem Poster werden Experimente mit der sogenannten "Satelliten-Box" vorgestellt. Durch ihre zylinderförmige Form eignet sie sich hervorragend für Rollexperimente auf einer schiefen Ebene. Mithilfe des eingebauten Gyroskops können Geschwindigkeit und zurückgelegte Strecke über die phyphox-Benutzeroberfläche angezeigt werden. Des Weiteren können mit Hilfe eines hochauflösenden Drucksensors Höhenunterschiede durch Differenzen im Luftdruck ermittelt werden. Dadurch lassen sich Experimente zum freien Fall realisieren oder die barometrische Höhenformel experimentell untersuchen.

Time: Tuesday 17:00-18:00

DD 36.1 Tue 17:00 P DigiPhysLab: Digital Physics Laboratory Work for Distance Learning -•SIMON Z. LAHME¹, PEKKA PIRINEN², BRUNO TOMRLIN³, ANTTI LEHTINEN², ANA SUŠAC³, ANDREAS MÜLLER⁴, and PASCAL KLEIN¹ — ¹U Göttingen, Germany — ²U Jyväskylä, Finland — ³U Zagreb, Croatia — ⁴U Geneva, Switzerland Integral part of studying physics are lab courses in which students learn how to link theory with practice, acquiring experimental and problem-solving competencies. Extensive research has shown that these learning goals are difficult to achieve, with multiple causes: (i) instructional (cf. cookbook-styled instructions), (ii) learner-related (e.g. low motivation) and (iii) content-related (i.e. inherent difficulties by e.g. multiple representations). Due to the Covid-19 pandemic, the challenges of effective lab courses increased as those needed to be implemented in distance rapidly. Thus, the EU-project DigiPhysLab pursues the development of physics lab tasks suitable for distance learning and tackles the general challenges by a competence-centred didactic concept. Based on a literature review, a framework for designing digital lab tasks is created. Building on it, 15 experiments on several topics of basic physics lectures are developed and evaluated. They enable engaging and authentic lab work both in distance and oncampus learning settings. For that, the wide availability of modern digital media is used as these technologies allow everyday data collection (e.g. with smartphone sensors) or immersive simulation of real experiments (e.g. with virtual reality). On the poster, we present the project conceptualiza- tion and current findings regarding the framework and experiments.

DD 36.2 Tue 17:00 P

Schallausbreitung in Festkörpern - ein Schülerversuch auf dem Prüfstand -•Sebastian Felleisen, Antje Bergmann und Carsten Rockstuhl — Institut für Theoretische Festkörperphysik, Karlsruher Institut für Technologie Das selbstständige Experimentieren stellt für Schülerinnen und Schüler (SuS) einen wichtigen didaktischen Zugang zu physikalischen Themen und Inhalten dar. Um den SuS diese Art des Lernens zu ermöglichen und zu vereinfachen, greifen Lehrerinnen und Lehrer vermehrt auf Versuch-Sets zurück, welche von Lehrmittelfirmen entwickelt werden. Im Rahmen der vorgestellten Bachelorarbeit wurde ein solches Versuchs-Set zur Bestimmung von Schallgeschwindigkeit in Festkörpern systematisch getestet und auf Verständlichkeit und physikalische Korrektheit untersucht. Die Anleitung des Herstellers wurden in Bezug auf Durchführungsvorschläge möglicher Experimente, der Versuchsauswertung, sowie Annahmen in der dort präsentierten Theorie kritisch hinterfragt. In unserem Beitrag zeigen wir Unstimmigkeiten in dieser Anleitung auf, was uns motivierte, die Theorie umfassend zu erarbeiten und mit Messwerten, die mit dem Set aufgenommen wurden, abzugleichen. Wir zeigen (a), dass der Versuch hervorragend geeignet ist, um physikalisches Verständnis für ein so elementares Phänomen wie die Schallausbreitung zu entwickeln aber auch (b), dass es essenziell ist, dies korrekt darzustellen, um die fachlichen Lehrinhalte richtig zu vermitteln.

DD 36.3 Tue 17:00 P

Optische Datenübertragung mit LEDs und Lasern im Schülerlabor — •Marcel Lauterwasser 1 , Antje Bergmann 2 und Günter Quast 1 — ¹Institut für Experimentelle Teilchenphysik, Karlsruher Institut für Technologie – ²Institut f
ür Theoretische Festk
örperphysik, Karlsruher Institut f
ür Technologie

DD 37: Postersession 2: Präsentation von Experimenten

Time: Tuesday 17:00-18:00

DD 37.1 Tue 17:00 P

Kostengünstige Simulation der Röntgen- und Elektronenbeugung mit Hilfe von optischen Gittern - • HUBERTUS GIEFERS - Humboldt Gymnasium, Bad Pyrmont, Deutschland

Beugungsversuche mit Röntgen- oder Elektronenstrahlen sind im schulischen Unterricht oft auf wenige Substanzen wie NaCl oder Graphit beschränkt. Eine Alternative zur Darstellung von Laue- und Debye-Scherrer-Aufnahmen stellen die Beugungsmuster nach Koppelmann dar, wobei auf dem Lehrmittelmarkt solche Beugungsgitter kaum erhältlich sind. In diesem Beitrag werden neu entwickelte Beugungsgitter für solche optischen Analogieversuche sowie die didaktische Hinführung vorgestellt. Die 2D-Transmissionsgitter zeigen Beugungsmuster ähnlich denen echter Materialsysteme und sie können mit Hilfe eines Lasers und des Transmissionsgitters im Diaformat kostengünstig und schnell gezeigt werden. Die neuen Transmissionsgitter sind so aufgebaut, dass die grobe Struktur für das menschliche Auge sichtbar auf dem Dia erkennbar ist, die Mikrostruktur für den Beugungsversuch allerdings erst mit dem MiOptische Datenübertragung ist durch die hohe Übertragungsrate höchst vielversprechend und wird in unserem Alltag durch Fernbedienungen, Glasfasern und neue Projekte wie Starlink immer präsenter. Gerade deshalb ist es auch relevant, diese Technologie Schülerinnen und Schülern näher zu bringen. In diesem Beitrag werden zwei Versuche zur Optischen Datenübertragung vorgestellt, die im Physik-Schülerlabor des KIT eingesetzt werden sollen. Im ersten Versuch wird ein analoges Musiksignal über eine einzelne, in der Helligkeit modulierte LED übertragen. Empfangen wird das Signal in einigen Metern Entfernung von einer Photodiode. Diese wandelt das optische Signal in ein elektronisches um, welches von einem Lautsprecher wiedergegeben werden kann. Dieser Versuch ist sehr kostengünstig und einfach zu realisieren und auch als Unterrichtsprojekt geeignet. Im zweiten Versuch werden zwei unterschiedlich farbige Laser mit jeweils einem Musiksignal moduliert. Die Laser sind durch ein optisches Gitter in dieselbe Glasfaser eingekoppelt (Frequenzmultiplexing). Nach dem Austritt aus der Glasfaser werden beide Lichtstrahlen durch einen dichroitischen Spiegel getrennt und von unterschiedlichen Lautsprechern wiedergegeben.

DD 36.4 Tue 17:00 P

Umgang mit digitalen Messsystemen: Lernhilfen bei der Fehlersuche -•CHRISTOPHER JOERGENS und CORNELIA GELLER - Universität Duisburg-Essen

Durch ihren Mehrwert für experimentelle Fragestellungen sind digitale Messsysteme aus den Experimentierpraktika der universitären Ausbildung nicht mehr wegzudenken, entsprechende Basiskompetenzen für Lehramtsstudierende sind bereits diskutiert worden. Wie genau der Umgang mit Messsystemen, also beispielsweise eine geeignete Messparameterwahl, als Teil der experimentellen Kompetenz systematisch gefördert werden kann, ist bisher aber wenig untersucht worden.

Mit dem Ansatz, den experimentellen Kompetenzerwerb mit mehr Varianz in den Aufgabenstellungen zu unterstützen, setzen wir in einem Praktikum auch Fehlersuchaufgaben ein, in denen die Studierenden über das Finden und Beheben von typischen Fehlern verschiedene Wissensarten (wie z.B. Gerätewissen und inhaltliches Wissen) verknüpfen können. An zwei Experimentieraufbauten mit digitalen Messsystemen wurde dabei untersucht, inwieweit Studierende das Angebot gestufter Lernhilfen wahrnehmen. Das Prinzip der Aufgaben sowie die Ergebnisse der Erprobung werden auf dem Poster vorgestellt und Implikationen für die experimentelle Ausbildung diskutiert.

DD 36.5 Tue 17:00 P

Location: P

Ziele eines Demonstrationspraktikums für Physik-Lehramtsstudierende — •KATHARINA STÜTZ und RONNY NAWRODT — Physik und ihre Didaktik, Universität Stuttgart, 70569 Stuttgart

Das Präsentieren von Demonstrationsexperimenten oder das Durchführen von Schülerexperimenten sind zentrale Bausteine des Physikunterrichts. In der universitären Ausbildung erfolgt die Vermittlung der notwendigen Fähigkeiten klassischerweise in diversen Praktika. In diesem Beitrag soll ein Überblick über die konkreten Ziele eines solchen Praktikums gegeben werden. Dazu werden die Zielvorstellungen von 17 Studierenden aus drei Semestern den Zielen aus Theorie und Praxis gegenübergestellt und diskutiert. Für einen standortübergreifenden Überblick über die universitäre Praxis wurden die Modulpläne aller 48 Universitäten und Hochschulen, an denen in Deutschland Physik für das Gymnasiale Lehramt studiert werden kann, zusammengefasst.

kroskop/Diaprojektor. Eine qualitative Auswertung der Beugungsmuster kann mit der Bragg-Gleichung erfolgen, da die auftretenden Beugungswinkel klein sind. Im Folgenden eine Auswahl an Beugungsgittern: verschiedene 2D-Bravais-Gitter; einkristalline, pulverförmige und amorphe Substanzen; Graphitpulver; Legierungen/intermetallische Verbindungen; isotrope/texturierte Substanzen; inkommensurable Strukturen; große Moleküle; Quasikristalle; Temperatureinflüsse. Diese neu entwickelten Transmissionsgitter sind selbstverständlich auch für die Lehre im Hochschulbereich interessant.

DD 37.2 Tue 17:00 P Exploration wichtiger ästhetischer Qualitäten der Wissenschaftsillustration am Beispiel von MR- AR- und Web3D-Applikationen zur Präsentation von Experimenten in der Quantenphysik — •JONAS LAUSTRÖER¹, REIN-HARD SCHULZ-SCHAEFFER¹, JOCHEN STUHRMANN¹, RASMUS BORKAMP¹, ADRI-

an Abazi², Carsten Schuck², Wolfram H. P. Pernice², Stefan Heusler²,

PAUL SCHLUMMER² und DANIEL LAUMANN² — ¹HAW Hamburg, Hochschu-

Location: P
le für angewandte Wissenschaften, Informative Illustration und Wissenschaftsillustration, Germany $-^2 \rm WWU$ Münster, Center for Nanotechnology, WWU Münster Institut für Didaktik der Physik, Germany

Um visuelle Wissenschaftskommunikation mit interpretierenden Bildern evaluieren zu können, wurde ein Katalog an visuellen ästhetischen Variablen erarbeitet. Visuellen Kriterien zur Evaluation ästhetischer Qualitäten in dynamischen, dreidimensionalen Gestaltungen lassen sich am besten visuell entwickeln und

DD 38: Postersession 2: Quantenphysik

Time: Tuesday 17:00-18:00

DD 38.1 Tue 17:00 P

Quantenteleportation und Verschränkung im Science Center mit erweiterter Realität: Projekt Holodeck:Q – •FRANZISKA GREINERT¹, OLIVER BODENSIER¹, DOMINIK ESSING² und GOWTHAM MUTHUSAMY¹ – ¹TU Braunschweig, Institut für Fachdidaktik der Naturwissenschaften, Germany – ²phaeno gGmbH, Wolfsburg, Germany

Quantentechnologien, die Verfahren für die physikalisch abhörsichere Kommunikation nutzen, gewinnen zunehmend an Bedeutung und Bekanntheit. Damit gehen jedoch auch vielfach Skepsis und Verunsicherung mit einher. Daher sind Hemmungen abbauen und Interesse wecken die Hauptziele der Outreach-Initiative Quantum aktiv des BMBF und auch des Projektes Holodeck:Q.

Im Wolfsburger Science Center phaeno entsteht im Rahmen des Projektes ein zweiteiliges Exponat. Im ersten Teil wird spielerisch in erweiterter Realität (AR) das Quantenteleportationsprotokoll eingeführt, bevor im zweiten, vertiefenden Teil Verschränkung erlebbar wird. Für letzteres wird eine Einzelphotonenquelle von qutools mit zwei Polarisatoren verwendet, sodass die Verschränkung in der Polarisationsrichtung von Photonenpaaren beobachtbar wird. Vorgestellt werden das didaktische Konzept, der aktuelle Umsetzungststand und die weiteren Pläne.

DD 38.2 Tue 17:00 P Explanation of Quantum Physics by Gravity and Relativity: A Possible Course — •HANS-OTTO CARMESIN — Gymnasium Athenaeum, Harsefelder Straße 40, 21680 Stade — Universität Bremen, Fachbereich 1, Postfach 330440, 28334 Bremen — Studienseminar Stade, Bahnhofstraße 5, 21682 Stade evaluieren. Im Gegensatz zum statischen 2D-Bild sind Qualitäten wie Komposition, Farbharmonie, Licht, Textur, Aufmerksamkeitssteuerung etc. in interaktiven 3D-Applikationen von dynamischen Einflüssen abhängig. Interaktive Elemente erweitern die Gestaltungsräume und müssen in die visuelle Konzeption einbezogen werden. Das Plakat stellt dynamische, ästhetische Variablen vor, die zur Verbesserung der Lesbarkeit des User Interface sowie zur Steigerung der Usability und der User Experience entwickelt wurden. Diese Variablen werden zur Datenerhebung in qualitativen Interviews adressiert.

Location: P

Since Planck discovered quantization in 1900, the nature of quanta was a mystery. That problem has now been resolved [1]. For it, I derived the postulates of quantum physics from the equivalence principles, gravity and relativity, whereby I analyzed the vacuum.

Using that derivation, I explain many quantities and properties of quantum physics in a precise manner. Examples are the nature of nonlocality, the physical quantity corresponding to the wave function ψ , the mathematical transformation describing the particle wave duality and the origin of the dynamics inherent to the Schrödinger equation. Moreover, I propose and derive the generalized Schrödinger equation. Furthermore, I identify the physical basis of the Planck constant *h*.

Altogether, quantum physics has now been derived, explained and extended in a direct and transparent manner on the basis of space, time and gravity. As an additional test, I derive the density parameter Ω_{Λ} of the vacuum by using the wave function ψ . The result is in precise accordance with observation, whereby I do not apply any fit. I propose a concept for a course in quantum physics, based on space, time and gravity.

[1] Carmesin, H.-O. (February 2022): Explanation of Quantum Physics by Gravity and Relativity. Berlin: Verlag Dr. Köster.

DD 39: Postersession 2: Sonstige

Time: Tuesday 17:00-18:00

DD 39.1 Tue 17:00 P

Physikdidaktik - Quo vadis? — •JOHANNES GREBE-ELLIS¹, SUSANNE HEINICKE², MICOL ALEMANI³, MARTIN HOPF⁴, HEIKO KRABBE⁵, DANIEL LAUMANN², HORST SCHECKER⁶, ERICH STARAUSCHEK⁷, HEIKE THEYSSEN⁸, THOMAS WILHELM⁹ und RITA WODZINSKI¹⁰ — ¹Universität Wuppertal — ²Universität Münster — ³Universität Potsdam — ⁴Universität Wien — ⁵Universität Bochum — ⁶Universität Bremen — ⁷PH Ludwigsburg — ⁸Universität Duisburg-Essen — ⁹Universität Frankfurt — ¹⁰Universität Kassel

Die Initiative "Physikdidaktik - Quo vadis?" widmet sich der Frage, wohin wir uns als Physikdidaktik zukünftig entwickeln möchten. Die Planungen für eine Tagung, bei der sich Professor*innen der Physikdidaktik in Klausur intensiv austauschen, laufen bereits seit 2019. Coronabedingt musste diese Tagung mehrfach verschoben werden. Um zumindest einen ersten Schritt gehen zu können, fand am 7. und 8. Oktober 2021 ein erster Teil der Tagung online statt. In vier "Schlaglichtern" wurden exemplarisch zentrale Forschungsgebiete der Physikdidaktik, ihre bisherige Entwicklung und der aktuelle Forschungsstand vorgestellt. In drei "Reflexionen" wurden übergreifende Themen wie Rahmenbedingungen und Methoden diskutiert. Der Online-Tagung soll eine Präsenztagung im Frühsommer 2022 folgen, die sich aufbauend auf einer Rückschau und Standortbestimmung vermehrt den Perspektiven physikdidaktischer Forschung widmet. Auf dem Poster werden das gemeinsame Anliegen der beiden "Quo vadis?"-Tagungen und ausgewählte Ergebnisse der Online-Tagung vorgestellt.

DD 39.2 Tue 17:00 P

Pythagoras und Euklid in höheren Dimensionen — •MARTIN ERIK HORN — iu - International University of Applied Sciences, Campus Berlin — ISM - International School of Management, Campus Berlin

Die Physik spielt sich in einer mindestens dreidimensionalen Welt ab, in der Richtungsbeziehungen eine wesentliche Rolle bei der Beschreibung physikalischer Phänomene spielen. Es ist deshalb sinnvoll im Rahmen physikalisch motivierter mathematischer Ansätze (wie z.B. der Geometrischen Algebra von Grassmann, Clifford und Hestenes) elementare geometrische Beziehungen nicht nur durch skalare, sondern durch richtungsbezogene Größen (also Vektoren, Bivektoren, Trivektoren, etc.) auszudrücken.

Am Beispiel der Satzgruppen von Pythagoras und von de Gua de Malves wird gezeigt, wie solche k-vektoriellen Beschreibungen gelingen und höherdimensionale Analogien zu den Höhen- und Kathetensätzen von Euklid gefunden werden können.

Und da die in der Physik betrachteten Größen auch nicht immer nur senkrecht zueinander stehen, ist es erfreulich, dass diese höherdimensionalen Verallgemeinerungen auch für nicht-rechtwinklige geometrische Objekte (in Analogie zu https://eldorado.tu-dortmund.de/bitstream/2003/39479/1/BzMu2020_HORN-id235.pdf) in eleganter Art und Weise beschrieben werden können.

DD 39.3 Tue 17:00 P

Location: P

Subjektive Relevanz von Physik im Alltag — •JASMINA $MUJAGIC^1$ und CLAU-DIA HAAGEN-SCHÜTZENHÖFER² — ¹Fritz Strobl Schulzentrum, Spittal an der Drau, Österreich — 2 Universität Graz, Institut für Physik, Graz, Österreich Das Unterrichtsfach Physik hat kein besonders gutes Image. Schülerinnen und Schüler verlieren im Laufe ihrer Schulzeit zunehmend das Interesse am Unterrichtsfach Physik, wie zahlreiche Studien zeigen. Im Gegensatz dazu boomen Wissenschaftsformate etwa im Fernsehen. Untersuchungen dazu, welchen Nutzen Schülerinnen und Schüler, die keinen beruflichen Weg in eine technischnaturwissenschaftliche Richtung einschlagen, vom erworbenen Wissen aus dem Physikunterricht auch nach der Schulzeit ziehen, gibt es bisher kaum. Es ist wenig darüber bekannt, wie Erwachsene die Relevanz ihres Schulwissens im Bereich Physik für ihren Alltag einschätzen. Die vorgestellte Qualifizierungsarbeit setzte an dieser Fragestellung an und untersuchte, ob Menschen, insbesondere auch jene ohne beruflichen Hintergrund im technisch-naturwissenschaftlichen Bereich, ihrer individuellen Einschätzung nach von ihrem Schulwissen in ihrem Lebensalltag profitieren, bzw. dieses für Entscheidungen heranziehen. Dazu wurden 72 Erwachsene mittels Fragebogen befragt, inhaltlich wurde auf die Bereiche Anfangselektrizitätslehre und Erneuerbare Energie fokussiert.

DD 40: Mitgliederversammlung des Fachverbands Didaktik der Physik

Time: Tuesday 18:00-19:30

Mitgliederversammlung

DD 41: Neue / digitale Medien - Experimente

Time: Wednesday 10:45–11:45

DD 41.1 Wed 10:45 DD-H8 **Mit Arduino und Spielzeugeisenbahn zur Relativitätstheorie** — •Jörg SCHNEIDER und HOLGER CARTARIUS — AG Fachdidaktik der Physik und Astronomie, Friedrich-Schiller-Universität Jena, 07743 Jena

Die Relativitätstheorie ist als Teil des Physikunterrichts der gymnasialen Oberstufe fest in den Bildungsplänen verankert. Leider mangelt es aber an konkreten Experimenten, mit denen sich relativistische Prinzipien und Effekte veranschaulichen und erklären lassen.

Um dieser Problematik entgegenzuwirken und die Experimentiermöglichkeiten zur Relativitätstheorie im Schulunterricht und in unserem Schülerlabor zu erweitern, wurden auf Grundlage von Arduino Analogversuche entwickelt, welche Simulationen mit konkreten, greifbaren Experimenten vereinen. In Verbindung mit einer Spielzeugeisenbahn können so beispielsweise die Zeitdilatation und die Längenkontraktion untersucht werden.

Im Rahmen des Vortrages soll eine Übersicht über die Anwendungsmöglichkeiten des Aufbaus gegeben werden sowie exemplarisch für einen Versuch zugehöriges Lehr- und Lernmaterial vorgestellt werden.

 $DD \ 41.2 \ Wed \ 11:05 \ DD-H8$ phyphox: Erste Testungen der externen Sensorboxen mit erweiterten Experimentiermöglichkeiten — •DUSTIN KIRWALD¹, DOMINIK DORSEL¹, LEO BODEWIG¹, SEBASTIAN STAACKS², CHRISTOPH STAMPFER² und HEIDRUN HEINKE¹ — ¹RWTH Aachen University, I. Physikalisches Institut IA — ²RWTH Aachen University, II. Physikalisches Institut A

Bis heute sind bereits viele faszinierende wie didaktisch gewinnbringende Experimente mit Hilfe der internen Sensoren von Smartphones entwickelt worden. Die kostenlose und quelloffene App phyphox greift auf die fest verbauten, internen Sensoren des Smartphones zu und stellt die Messdaten sowie deren Auswertung live dar. Damit die Lernenden weitere Inhalte über ihre eigenen mobilen Endgeräte experimentell erschließen können, müssen die bereits zugänglichen Messgrößen aus der internen Sensorik des Gerätes um weitere Größen erweitert werden. Dazu nutzt phyphox eine Bluetooth Low Energy Schnittstelle, welche es externer Messtechnik ermöglicht, Daten aufzunehmen und an phyphox zu übermitteln. Es sind vier Sensorboxen entwickelt worden, die eine quantitative Messung und Auswertung von Experimenten unter anderem aus der Elektrizitätsund Wärmelehre sowie der Mechanik in der phyphox-App ermöglichen. Diese Sensorboxen durchlaufen aktuell in Kooperation mit zwölf Partnerschulen und in den physikalischen Praktika der Universität erste Anwendungstests. Im Vortrag werden sowohl das Konzept zu den Sensorboxen als auch erste Rückmeldungen seitens der Schulen und Ergebnisse aus Testläufen in den Praktika vorgestellt.

DD 41.3 Wed 11:25 DD-H8

Location: DD-H9

Die PhyxBox – ein interdisziplinäres Lehrmittel für diverse Unterrichtsformen — ROBERT SCHNEEWEISS, SARA OGRISSEK, MICHAEL BECKSTEIN, •SIMEON VÖLKEL und AXEL ENDERS — Experimentalphysik XI, Universität Bayreuth, 95440 Bayreuth

Unsere PhyxBox stellt ein neu entwickeltes Lehrmittelkonzept dar, um interdisziplinäre Lerninhalte mit Schülerversuchen zu vermitteln. Konkret werden hier an der Grenzfläche zwischen Informatik und Physik angesiedelte PhyxBoxen vorgestellt.

Die PhyxBox zeichnet sich durch eine explizite Trennung in physikzentrierte und informatikzentrierte Aufgabenstellungen aus. Diese sind speziell aufeinander abgestimmt, so dass physikorientierte und informatikorientierte Teams parallel arbeiten und ihre Ergebnisse schließlich zu einem gemeinsamen Aufbau kombinieren können. Die PhyxBox-Hälften sind dabei so konzipiert, dass sie auch für sich genommen als einzelne Unterrichtseinheit direkt einsetzbar sind.

Dieses Konzept wird am Beispiel der PhyxBox zu Thermometern welche auf dem thermoelektrischen Effekt basieren illustriert. Eine aufbauende PhyxBox zu dessen Umkehrung, dem Peltiereffekt, ermöglicht durch die Kombination mit Arduino-basierter Datenerfassung und Auswertung den Bau eines Taupunkt-Hygrometers. Damit skalieren die PhyxBoxen von fachübergreifendem Unterricht über Begabtenförderung bis zu universitären Praktikumsversuchen. Durch die spezielle Zusammenstellung der Materialien sind die PhyxBoxen nicht nur für den Präsenzunterricht geeignet, sondern auch in der Distanzlehre oder für selbstreguliertes Lernen einsetzbar.

DD 42: Lehr- und Lernforschung - Repräsentatonsformen

Time: Wednesday 10:45-11:45

DD 42.1 Wed 10:45 DD-H9

Wie man multiple externe Repräsentationen gewinnbringend im Physikunterricht einsetzt: Resultate zweier Interventionsstudien — • Andreas Lichtenberger¹, Lennart Schalk² und Tommi Kokkonen³ — ¹ETH Zürich, Schweiz — ²PH Schwyz, Schweiz — ³Universität Helsinki, Finnland Im naturwissenschaftlichen Unterricht werden meist verschiedene externe Repräsentationen wie Manipulative (z.B. Experimente), Visualisierungen (z.B. Diagramme) und mathematische Formeln eingesetzt. Zahlreiche Studien belegen die positive Förderung des Lernens durch den Einsatz von multiplen im Vergleich zu einzelnen Repräsentationen. Weniger erforscht ist bislang, wie verschiedene Repräsentationen möglichst effektiv sequenziert und kombiniert werden können. Eine Methode, die als besonders förderlich für den Mathematikund Naturwissenschaftsunterricht vorgeschlagen wird, ist das Concreteness Fading. Dabei beginnt der Unterricht mit einer konkreten Repräsentation (z.B. einem Experiment) und geht schrittweise zu abstrakteren Repräsentationen (z.B. Formeln) über. In zwei experimentellen Studien am Gymnasium zum anspruchsvollen Thema der elektromagnetischen Induktion haben wir die Effektivität von Concreteness Fading mit der umgekehrten Sequenz, Concreteness Introduction (N = 70), und mit einem integrierten Ansatz, bei dem alle Repräsentationen gleichzeitig eingeführt werden (N = 115), verglichen. Es hat sich gezeigt, dass sich bezüglich des konzeptuellen Verständnisses keine Unterschiede zwischen den drei Ansätzen ergeben. Allerdings gibt es Hinweise darauf, dass der integrierte Ansatz die Repräsentationskompetenz stärkt.

DD 42.2 Wed 11:05 DD-H9 Räumliches Denken bei Studierenden bei der Lösung mathematischphysikalischer Aufgaben zu Bewegung und Veränderung — •MARION ZÖG-GELER — Universität Salzburg, Österreich Zur Lösung mathematisch-physikalischer Aufgaben fließt das räumliche Denken in unterschiedlicher Ausprägung und auf verschiedene Weise ein. Anhand einer qualitativen Studie wird untersucht, wie Studierende räumliche Denkprozesse bei der Bearbeitung ausgewählter Aufgaben aus den Bereichen Mathematik, Physik, Technik und Astronomie einsetzen. Ziel der Untersuchung ist es, eine Vielzahl an räumlichen Denkschritten zu sammeln. Die Datenerhebung dieser Studie an der Universität Salzburg erfolgt mittels individueller Bearbeitung der Aufgaben, einer paarweisen Besprechung des Lösungsweges und eines leitfadengestützten Interviews. Die Aufgaben beziehen sich auf Inhalte zu Bewegung und Veränderung als wesentliche Elemente des räumlichen Denkens. Eine eingehende Sachanalyse fachlicher Inhalte der STEM-Fächer zeigt nämlich, dass Bewegung als fachliches Element und als Element des räumlichen Denkens eine zentrale Rolle spielt: als Vorstellung einer realen Bewegung, als verändernder Vorgang in einem Lösungsprozess sowie als Erkennen von Bewegbarkeit von Teilen in einem System. Die Auswertung, die auf einer qualitativen Inhaltsanalyse beruht und auf die Aufstellung von Hypothesen zielt, zeigt spezielle Prozesse des räumlichen Denkens, wie die Fokussierung auf relevante Merkmale der Bewegung, aber auch die Verbindung mit weiteren mentalen Fähigkeiten, wie unter anderem mit logischem Denken.

DD 42.3 Wed 11:25 DD-H9 Förderung schriftlicher Erklärungen im Physikunterricht — •Неіко Ккавве, Сакіла Wöhlke und Lukas Elflein — Physikdidaktik, Ruhr-Universität Bochum

Sprachliche Handlungen wie das Erklären sind ein fester Bestandteil der Bildungsstandards für Physik (Tajmel, 2011). Wissenschaftliche Erklärungen haben eine spezifische Struktur, die auch sprachlich differenziert ist (Osborne & Patterson, 2010). In der Studie wird die Umsetzung von Erklärungen im Phy-

Location: DD-MV

Location: DD-H8

Location: DD-H10

Location: DD-H11

sikunterricht in fachlich-konzeptioneller und sprachlicher Hinsicht gefördert. In einer Untersuchung mit 5 Gesamtschulklassen erstellten die Schülerinnen und Schüler Erklärvideos, in denen sie die Bewegung eines startenden Raumschiffs anhand der auftretenden Kräfte erklären sollten. Sie mussten zuvor Drehbuchtexte verfassen und diese je nach Gruppe mit rein fachlich-konzeptionellen, rein sprachlichen oder fachlich-konzeptionellen und sprachlich kombinierten Gerüsten zur Strukturierung der Erklärungstexte überarbeiten. Die Analysen mittels Kodiermanual, t-Test und Kruskal-Wallis-Test zeigen, dass fachlichfachlich-konzeptionelle und sprachlich kombinierte Gerüste am ehesten beim Verfassen einer Erklärung helfen.

DD 43: BNE - Lernendenperspektive

Time: Wednesday 10:45-11:45

DD 43.1 Wed 10:45 DD-H10

Mit Interesse ermächtigt: die Welt verstehen und mit Hilfe der Physik han**deln** — •John-Luke Ingleson¹, André Bresges² und Alexander Strahl³ ¹Wöhlerschule, 60320 Frankfurt – ²Universität Köln – ³Universität Salzburg

Klimawandel, Pandemien, ein sich beschleunigender gesellschaftlicher Wandel, neue Technologien und eine höhere Dichte an Informationen verändern die Art und Weise, wie Lernende Ihre Umwelt wahrnehmen. Auf der Basis dreier Studien aus Salzburg und der IPN Interessenstudie soll anhand einer Befragung in Deutschland ein aktueller Überblick über das Interesse von jungen Menschen in den Naturwissenschaften ermöglicht werden. Hier werden insbesondere Bereiche, die sich mit den Problemstellungen der Umwelt, der Gesellschaft und den SDG beschäftigen, betrachtet. Bei den bisher vorliegenden Studien hat sich herausgestellt, dass sich unterschiedliche Interessengruppen von Lernenden herauskristallisiert haben. Vor allem bei den zwei größten dieser Interessensgruppen deutet sich das Potential an, über umwelt- und gesellschaftsrelevante Themen zu motivieren und damit die physikalische Breitenbildung zu stärken. Es ist geplant gemeinsam mit Lernenden in einem nächsten Schritt durch Aufbau und Testen von Hypothesen zu prüfen, ob die obenen genannten Themen bei den Interessentypen auf Akzeptanz stoßen. Das Ziel ist es, darauf aufbauend in Lerngruppen mit Hilfe agiler Arbeitsweisen Interventionen zu entwickeln, die kraft der Physik Lösungswege für Zukunftsthemen aufzeigen.

DD 43.2 Wed 11:05 DD-H10

Bildung für nachhaltige Entwicklung, 4K und digitale Kreativitätswerkzeuge - •JANNIK HENZE, LARS MÖHRING und ANDRÉ BRESGES — Institut für Physikdidaktik, Universität zu Köln, Köln, Deutschland

Bildung für nachhaltige Entwicklung beinhaltet neben der Ebene der Information auch die Ebene des Kompetenzerwerbs. Im mathematischnaturwissenschaftlichen Unterricht können und sollen die 4K-Kompetenzen Kooperation, Kreativität, Kritisches Denken und Problemlösen sowie Kom-

munikation durch die Unterrichtseinheiten vertieft werden. Moderne Bildung jedoch verbindet dies gleichzeitig mit einer digitalen Komponente. In dem Projekt *Zukunft gestalten mit Mensch und Technik* wurde jene Kombination mithilfe digitaler Kreativitätswerkzeuge berücksichtigt, erprobt und die sich daraus ergebende Auswirkung der Lernwerkzeuge hinsichtlich der 4K evaluiert. Auch die Auswirkung digitaler Kreativitätswerkzeuge auf das Growth Mindset und die Bereitschaft solche Tools im eigenen Unterricht einzusetzen wurden untersucht.

Pilotierungsergebnisse eines Klimawandel-Konzepttests — •THOMAS Schubatzky¹, Rainer Wackermann², Carina Wöhlke², Claudia Haagen-Schützenhöfer¹, Hannes Kasimir Lindemann², Kai Cardinal² und Marko Jedamski² — ¹Universität Graz, Graz, Österreich — ²Ruhr-Universität Bochum, Bochum, Deutschland

Der aktuelle Klimawandel stellt unsere Gesellschaft vor große Herausforderungen. Der Klimawandel ist somit besonders für junge Menschen, deren Leben künftig von Klimawandelauswirkungen geprägt ist, zentral. Ein grundlegendes Verständnis des Klimawandels kann Personen insbesondere dabei helfen, falsche oder widersprüchliche Darstellungen besser einzuschätzen, um so am gesellschaftlichen Diskurs zum Thema Klimawandel teilhaben zu können. Die Erfassung von Schülervorstellungen ist außerdem zentral für die Entwicklung von Lernangeboten, etwa im Sinne der didaktischen Rekonstruktion. Um Aussagen über das Verständnis von zentralen fachlichen Inhalten zum Klimawandel zuverlässig treffen zu können, braucht es geeignete Testinstrumente. Für die Entwicklung eines derartigen Klimawandel-Konzepttests wurden zentrale fachliche Inhalte identifiziert, offene Fragen entwickelt und schließlich Distraktoren und Attraktoren aus über 30 Interviews abgeleitet. Anschließend wurde der Konzepttest bei insgesamt rund 200 Schüler:innen, Studierenden und Expert:innen pilotiert. Im Beitrag werden die Ergebnisse der (quantitativen) Pilotierung und abgeleitete Überarbeitungsschritte vorgestellt. Zudem wird ein Ausblick auf die geplante Hauptstudie gegeben.

DD 44: Hochschuldidakitk - neue Konzepte

Time: Wednesday 10:45-11:45

DD 44.1 Wed 10:45 DD-H11

Hochschuldidaktischer Vergleich von Experimentierhausaufgaben und klasischen Übungsaufgaben — • ANDREAS KAPS und FRANK STALLMACH — Universität Leipzig, Fakultät für Physik und Geowissenschaften, Bereich Didaktik der Physik, Prager Straße 36, 04317 Leipzig

Im Rahmen einer quasi experimentellen Feldstudie im Zweikohortendesign wurden smartphonebasierte Experimentierhausaufgaben mit klassischen Übungsaufgaben verglichen. Es wurde der Einfluss auf die Motivation und das Interesse mit validierten Testinstrumenten untersucht. Außerdem wurde die Lernwirksamkeit über Pre- und Posttest quantifiziert. Das Studiendesign und die Aufgaben werden detailliert vorgestellt. Die Ergebnisse des Vergleichs werden abschließend diskutiert und Implikationen für die Lehre abgeleitet.

DD 44.2 Wed 11:05 DD-H11

Strukturanalyse von Physik Studiengängen — •DANIELA KERN-MICHLER¹, Stefan Brackertz², Sophie Penger², Manuel Längle³, Christoph Kronberger⁴, Annemarie Sich², Lisa Lehmann⁵, Wanda Witte⁶ und Amr EL MINIAWY⁷ — 1 ZaPF e.V., Frankfurt, Deutschland — 2 Uni zu Köln, Köln, Deutschland — ³Uni Wien, Wien, Österreich — ⁴TU Wien, Wien Österreich — ⁵TU Dresden, Dresden, Deutschland — ⁶Uni Rostock, Rostock, Deutschland — ⁷Humbolt Uni, Berlin, Deutschland

Bei der Frage wie Studiengänge flexibel studierbar gestaltet werden können, spielen vor allem die Struktur also die Module und ihre Zusammenhänge sowie Voraussetzungen eine große Rolle. Einem Vorschlag eines vorherigen Beitrags im Studienrefom-Forum folgend wurde ein Onlinetool erstellt, welches aus der Eingabe von Modulen und Voraussetzungen "Strukturformeln" oder "Explosionszeichnungen" erstellt [1].

In diesem Vortrag werden bisherige Ergebnisse des Projektes zur Erhebung

und zum Vergleich von Physikstudiengängen vorgestellt und diskutiert. Neben der Frage nach gemeinsamen Strukturmustern und Unterschieden stellt sich auch die Frage nach dem Einfluss der Darstellungsart.

[1] http://studiengang-diagramm.de/

DD 44.3 Wed 11:25 DD-H11 **Peer Instruction in der Theoretischen Physik** — • PHILIPP SCHEIGER^{1,2}, RONNY NAWRODT² und HOLGER CARTARIUS¹ — ¹Fachdidaktik der Physik und Astronomie, Friedrich-Schiller-Universität Jena, 07743 Jena — ²Physik und ihre Didaktik, Universität Stutt- gart, 70569 Stuttgart

Die Peer Instruction ist ein effektives Konzept, aktivierende Lehre praktisch umzusetzen. Durch gezielte Verständnis- oder Konzeptfragen mit Multiple-Choice-Antworten können Lernende animiert werden, das eigene Verständnis an konkreten Beispielen noch in der Vorlesung zu testen. In der Peerdiskussion lernen sie eigene Argumente zu formulieren und zu überprüfen. Seit diese Methode von Eric Mazur vorgestellt wurde, fand sie Einzug in unzählige Lehrveranstaltungen.

Im Bereich der Theoretischen Physik, mit meist vom Formalismus geprägten Vorlesungen, gibt es allerdings noch kaum Umsetzungen oder konkret ausgearbeitete Beispiele, die im Sinne der Peer Instruction eingesetzt werden können. Wir möchten in diesem Vortrag Beispiele vorstellen, wo und wie Konzeptfragen nach den fünf Grundregeln von Mazur (Fragen sollten: -sich auf ein Konzept konzentrieren; -nicht durch Anwendung von Formeln zu lösen sein; -attraktive Multiple-Choice-Distraktoren anbieten; eindeutig formuliert sein; nicht zu leicht oder zu schwer sein.) in der Theoretischen Physik umgesetzt werden können. Darüber hinaus möchten wir aber ebenfalls aufzeigen, wie das Feld der Peer Instruction neben den Konzeptfragen erweitert werden kann, um das Wechselspiel Physik-Mathematik zu trainieren und so die Formalismen physikalischer Theorien nahbarer zu machen.

DD 43.3 Wed 11:25 DD-H10

DD 45: Astronomie

Location: DD-H12

Time: Wednesday 10:45-11:45

DD 45.1 Wed 10:45 DD-H12

Fern, ferner, am fernsten - warum es sinnvoll ist, auch im Mathematikunterricht über Exoplaneten und Kosmologie zu sprechen — •ELEEN HAMMER und HOLGER CARTARIUS — AG Fachdidaktik der Physik und Astronomie, Friedrich-Schiller-Universität Jena, 07743 Jena

In lediglich 3 deutschen Bundesländern wird Astronomie als eigenständiges Fach in der Sekundarstufe I unterrichtet, obwohl Studien belegen, dass das Interesse der Schüler am Fach hoch ist und Astronomie einen wichtigen Beitrag zum Allgemeinbildungsauftrag der Schule leistet. Da Astronomie stark interdisziplinär ist, wird die enge Verknüpfung zur Mathematik ausgenutzt, um den Schülern im Mathematikunterricht astronomische Inhalte näher zu bringen. In diesem Vortrag werden konkrete Sachaufgaben für den modernen, kompetenzorientierten Mathematikunterricht vorgestellt, die den Schülern der Sekundarstufe I Fakten und Gesetzmäßigkeiten der Themenbereiche Exoplaneten und Kosmologie vermitteln.

DD 45.2 Wed 11:05 DD-H12

Stellarium Gornergrat - ein fergesteuertes astronomisches Observatorium für Bildungszwecke — •Stéphane Gschwind¹, Sascha Hohmann³, Andre-AS MUELLER¹ und TIMM RIESEN² — ¹IUFE - Institut Universitaire de Formation des Enseignants, Université de Genève, Schweiz — 2 CSH - Center for Space and Habitability, Universität Bern, Schweiz – ³IPN - Leibniz Institute for Science and Mathematics Education, Kiel, Deutschland

Das Stellarium Gornergrat gegenüber vom Matterhorn in der Schweiz ist eines

Time: Wednesday 12:00-13:00

DD 46.1 Wed 12:00 DD-H8

Bildung für Nachhaltige Entwicklung und Physikunterricht: Wie geht das am besten zusammen? — • André Bresges, Carina Schatz, Lars Möhring, Sa-SCHA THEROLF, JANNIK HENZE, FLORIAN GENZ UND CRISTAL SCHULT — Institut für Physikdidaktik, Universität zu Köln, 50923 Köln

Die globalen Ziele für nachhaltige Entwicklung der Agenda 2030 richten sich an Regierungen, Zivilgesellschaft, Privatwirtschaft und auch an die Wissenschaft. Die meisten dieser Ziele lassen sich ohne grundsätzliche mathematischnaturwissenschaftliche Kenntnisse weder durchdringen noch erreichen. Es ist jedoch die Komplexität von Nachhaltigkeitszielen wie Klimaschutz, nachhaltigen Städte und Gemeinden oder bezahlbarer und sauberer Energie, die dem Physikunterricht zu schaffen macht. Von der Berechnungsformel des Anhalteweges bis zum Ziel der sicheren und nachhaltigen Mobillität ist es unterrichtlich ein weiter Weg. Die Auseinandersetzung mit den komplexen und zum Teil nicht reversiblen Wirkungskreisen des Klimawandels kann bei Schüler*innen schnell das Gefühl von Verzweifelung bis hin zur erlernten Hilflosigkeit erzeugen. Aus Sicht von Klimaaktivisten wie der *Fridays for Future* Bewegung informiert Schule über den Status Quo und verstärkt ihn damit, da in den Schüler*innen keine *Kompetenz zur Veränderung* gefördert wird. Aufbauend auf dem 5E Modell von Bybee et al. erproben wir in der Inklusiven Universitätsschule der Universität zu Köln in verschiedenen Projekten, wie sich der Physikunterricht den Globalen Nachhaltigkeitsszielen annähern und die Kompetenz zur Veränderung des Status Quo in Schüler*innen anlegen kann.

DD 46.2 Wed 12:20 DD-H8

Climate Escape - Entkommen aus der Klimakatastrophe?! — •TIMO GRAFFE, JOHANNES LHOTZKY, FILIP SIRRENBERG, UWE OBERLACK UNd KLAUS WENDT — Institut für Physik, JGU Mainz

Das "Climate Escape" ist ein interaktiv gestaltetes Schülerlabor in Form eines Escape Games, welches im NaTLab Physik der Universität Mainz durchlebt werden kann. Dafür wurden sechs Stationen zum Thema "Klimawandel und Nachhaltigkeit" mit Fokus auf den naturwissenschaftlichen Grundlagen des Klimawandels für die Mittelstufe konzipiert. Neben einer physikalischen Schwerpunktsetzung werden im Labor auch immer wieder Bezüge zu chemischen, geographischen und biologischen Inhalten hergestellt. Dabei werden die Theder leistungsfähigsten roobotischen Observatorien für didaktische und "citizen science"-Zwecke in Europa. Über ein Webportal können SchülerInnen und LehrInnen von der Primarstufe bis zur gymnasialen Oberstufe verschiedene astronomische Beobachtungsaktivitäten durchführen. Die inhaltliche Bandbreite reicht von erdnahen Sonnensystem bis zu entfernten Galaxien.

Zu jeder Aktivität stehen Erläuterungen, Arbeitsblätter sowie Beobachtungsaufträge zur Verfügung, die von den Lernenden selbstständig online gebucht und daraufhin vom Teleskop eigenständig aufgenommen werden können. Diese sind wenig später auf dem Portal abrufbar und können ausgewertet werden.

Wir geben einen Überblick über neue Entwicklungen und Beobachtungsaktivitäten mit dem Stellarium Gornergrat, z.B. zu Exoplaneten oder zum Mond.

DD 45.3 Wed 11:25 DD-H12 Entwicklung eines Konzepttests zur Astronomie - Erste Ergebnisse •PHILIPP BITZENBAUER¹ und MALTE UBBEN² — ¹Staudtstr. 7 / B2, 91058 Erlangen — ²Wilhelm-Klemm-Str. 10, 48149 Münster

Sterne und das Weltall sind bei Lernenden seit jeher hoch im Kurs, wie die Ergebnisse von Interessensstudien zeigen. Empirische Instrumente zur Erhebung von Vorstellungen oder Konzeptverständnissen Lernender zu Konzepten der Astronomie sind jedoch rar. Im vorgestellten Projekt wird die Entwicklung und Pilotierung von solchen Instrumenten für den deutschen Sprachraum angestrebt. Wir berichten in diesem Beitrag die Ergebnisse der Pilotierung eines ersten Konzepttests zu Sternen, der auf international bereits etablierten Concept Inventories basiert.

DD 46: BNE - Konzepte

Location: DD-H8

men den Schüler:innen anhand von Modellversuchen im Escape Room-Design nähergebracht. Die Idee dahinter ist, den Bildungsinhalt mit dem Lösen von Rätseln und Öffnen von Schlössern zu verbinden. Die "Escape"-Umgebung ist durch gezielte Gamification, progressive Arbeitsmaterialgestaltung und Methoden der Binnendifferenzierung darauf ausgelegt, von Lernenden in eigenständiger Kleingruppenarbeit bewältigt werden zu können. Der Escape Room-Ansatz bietet so die Möglichkeit, in Form von Educational Gaming das Lernen zu einem Erlebnis zu machen und dabei die Schüler:innen intrinsisch zu motivieren. Der Vortrag im Rahmen der DPG-Tagung zielt darauf ab, den Zuhörer:innen einen Einblick in die Gestaltung des "Climate Escapes" und der Versuche zu geben sowie eine rezeptartige Anleitung für die Konzeption eines eigenen Escape Rooms aufzuzeigen.

DD 46.3 Wed 12:40 DD-H8 Bildung für nachhaltige Entwicklung im Naturwissenschaftsunterricht der IGS: Die Bedeutung von physikalischem Wissen im BNE-orientierten Unterricht — •Christin Sajons, Simon Hermanns, Mitja Evers und Michael Koмокек — Carl von Ossietzky Universität, Oldenburg, Deutschland

Bildung findet unter den Bedingungen einer sich schnell wandelnden Gesellschaft statt, was die Bedeutung einer Bildung für eine nachhaltige Entwicklung (BNE) immer stärker in den Fokus rückt. BNE soll Kinder und Jugendliche auch dazu befähigen, ein angemessenes und positives Bild von Wissenschaft zu entwickeln, um komplexe persönliche und gesellschaftliche Herausforderungen und Krisen (Pandemien, Klimawandel) anzunehmen und damit verbundene Chancen zu nutzen. Dennoch findet BNE bisher meist nur in Randformaten (AGs, Projektwochen) und selten direkt im Fachunterricht statt. Eine kooperierende Gesamtschule hat sich deshalb als Ziel gesetzt, das schuleigene Curriculum von den Global Goals der Vereinten Nationen aus neu zu strukturieren und den Fachunterricht daran auszurichten. Hierzu ist ein fachbezogener Unterricht nötig, in dem disziplinäre und übergreifende Kompetenzen (vgl. cross-cutting key competencies, UNESCO) erreicht werden können. Im Vortrag werden das Verhältnis zwischen physikalischen und übergreifenden Kompetenzen diskutiert und konkrete Entwicklungsergebnisse des Kooperationsprojektes präsentiert. Es wird anhand zweier exemplarischer Unterrichtsansätze demonstriert, wie der Anchored-Instruction-Ansatz genutzt wird, um einen BNE-orientierten Fachunterricht umzusetzen.

DD 47: Lehreraus- und -weiterbildung - Lehrkonzepte

Time: Wednesday 12:00-13:00

DD 47.1 Wed 12:00 DD-H9

Das Storytelling im Physikunterricht - konkrete Ausblicke zur Untersuchung der Wirksamkeit des Storytelling als Planungsansatz für Praxissemester Studierende — •MICHEL NOETHLICHS¹, BENJAMIN NIEHS², ALEXANDER STRAHL³ und ANDRÉ BRESGES⁴ — ¹ZfSL Leverkusen, 51379 Leverkusen — ²Europaschule Bornheim, 52223 Bornheim — ³Universität Salzburg, 5020 Salzburg — ⁴Institut für Physikdidaktik, Universität zu Köln, 50923 Köln

Die Lehrerausbildung in Deutschland unterteilt sich in das eher theoretische Studium und dem anschließenden Referendariat, in dem es eher darum geht, Inhalte in die Praxis umzusetzen. Dabei kommt der Planung von Physikunterricht eine sehr große Rolle zu und stellt für die Studierenden eine große Kompetenz da, die es gilt im Lauf der Ausbildung zu entwickeln. In diesem Forschungsfeld soll daher das Storytelling als Planungsansatz anhand der Umsetzung durch die Praxissemester Studierenden untersucht werden. Erste vielversprechende Ergebnisse zur Wirkung des Storytellings bei Lehrpersonen wurden bereits von der Arbeitsgruppe um Prof. Heering an der Uni Flensburg untersucht. Daran angelehnt soll ein etwas abweichender Ansatz des Storytelling als Planungsprämisse für Physikunterricht durch Praxissemester Studierende untersucht werden. Ein wesentliches Merkmal soll dabei die aktive Teilnahme der Lernenden in die Story sein - die Lernenden werden somit selbst Teil einer Geschichte und müssen Probleme oder Erkenntnisse mit Hilfe physikalischer Erkenntnisse lösen. Inwiefern sich dieser Planungsansatz auf die Planungskompetenz der Studierenden auswirkt, bleibt zu untersuchen.

DD 47.2 Wed 12:20 DD-H9

Beeinflusst eine professionsbezogene Lehre der Physik im Studium das schulpraktische Handeln? — •TILMANN STEINMETZ und ERICH STARAUSCHEK — PSE Stuttgart-Ludwigsburg

Der Professionsbezug gilt als wichtiger Bestandteil der Physiklehramtsausbildung. Es gibt verschiedene Ansätze diesen im Rahmen der Physikfachveranstaltungen an Universitäten zu erhöhen. Damit ist u.a. das Ziel verbunden, Studierende gezielter auf Anforderungen der Schulpraxis vorzubereiten. Unsere Studie untersucht, ob und wie sich ein Professionsbezug auf das Handeln angehender Physiklehrkräfte im Physikunterricht der Schulpraxis auswirkt. Eine vergleichende qualitative Interviewstudie mit Physik-Lehramtsstudierenden im Schulpraxissemester umfasst Gruppen: (1) Studierende mit und (2) Studierende ohne Professionsbezug in der physikalischen Fachausbildung des Grundstudiums. Bei den Studierenden der ersten Gruppe wurde der Professionsbezug durch das Konzept des kumulativen Lehrens im Lehramtsstudium Physik realisiert. Eine Typenbildung anhand einer qualitativen Inhaltsanalyse ergibt, dass Studierende der ersten Gruppe eher schülerorientiert unterrichten, während Studierende der zweiten Gruppe eher inhaltsorientiert unterrichten. Ein expliziter Professionsbezug könnte somit eine Sozialisation als Lehrkraft während des Fachstudiums

DD 47.3 Wed 12:40 DD-H9 Weniger rechnen, mehr sprechen: Einblicke in einen neuen Lehramtsstudiengang — •Martin Dickmann, Cornelia Geller und Hendrik Härtig — Universität Duisburg-Essen, Essen, Deutschland

Ausgehend von aktuellen Befunden der Unterrichts- und Professionalisierungsforschung wird an der Universität Duisburg-Essen ein spezifischer Bachelorstudiengang für das Sek I-Lehramt an nicht gymnasialen Schulformen neu konzipiert. Ziele der Neukonzeption sind der Erwerb von konzeptuellem Verständnis in unterrichtsrelevanten Inhalts- und Handlungsfeldern bei geringerem Mathematisierungsgrad sowie die Unterstützung der Studienmotivation durch Kompetenzerleben und Kohärenzwahrnehmung. Dazu werden sowohl inhaltlich als auch organisatorisch neue Strukturen erprobt, beispielsweise werden klassische Veranstaltungsformate durch kognitiv aktivierende Lehr- und Lernformate ersetzt. Im Vortrag werden die Grundideen der Neukonzeption vorgestellt und an konkreten Beispielen diskutiert.

DD 48: Lehr- und Lernforschung - Methodik

Time: Wednesday 12:00–13:00

DD 48.1 Wed 12:00 DD-H10

Was ist eigentlich tet.folio? — •SEBASTIAN HAASE¹, MARKUS ELSHOLZ², WOLF-GANG LUTZ² und THOMAS TREFZGER² — ¹Fachbereich Erziehungswissenschaft und Psychologie, AB Schulpädagogik/Schulentwicklungsforschung, Freie Universität Berlin — ²Lehrstuhl für Physik und ihre Didaktik, Julius-Maximilians-Universität Würzburg

tet.folio ist als offene Entwicklungs- und Lehr-Lern-Plattform aus dem Projekt *Technology Enhanced Textbook (TET)* an der FU Berlin entstanden. Das zugrundeliegende didaktisch-technologische Prinzip fordert, dass sich digital angereicherter Unterricht nicht von den momentan verfügbaren technischen Lösungen einschränken lassen darf. Didaktische Unterrichtskonzeptionen werden vielmehr als Anforderungen an eine technologische Umsetzung verstanden und Lehrende als kreative Autor:innen digitaler Lernumgebungen gesehen. Im Sinne des SAMR-Modells können sie interaktive und kollaborative digitale Lerninhalte erstellen, die sich nahtlos in das individuelle didaktische Konzept einfügen. Im Vergleich zu einer menügesteuerten Learning Management Umgebung orientiert sich tet.folio an der Buch-Metapher und fokussiert auf die Freiheit der Gestaltungsmöglichkeiten. Im Beitrag zeigen wir exemplarisch verschiedene Anwendungsszenarien und aktuelle Entwicklungsrichtungen der tet.folio-Umgebung, wie z.B. real-time Kollaboration und Erhebungsmöglichkeiten für die Lehr-Lern-Forschung.

Computerbasiertes Testen mit tet.folio — •WOLFGANG LUTZ¹, SEBASTI-AN HAASE², MARKUS ELSHOLZ¹ und THOMAS TREFZGER¹ — ¹Lehrstuhl für Physik und ihre Didaktik, Julius-Maximilians-Universität Würzburg — ²Fachbereich Erziehungswissenschaft und Psychologie, AB Schulpädagogik/Schulentwicklungsforschung, Freie Universität Berlin

In der empirischen Forschung werden zur Erfassung von Personenmerkmalen häufig Fragebögen oder Multiple Choice Tests im Papierformat eingesetzt. Die erhobenen Daten werden anschließend zur weiteren Auswertung codiert und digitalisiert. Computerbasierte Testungen bieten demgegenüber viele Vorteile z.B. im Bereich der Organisation, Informationsverarbeitung und automatischer Auswerteverfahren. Im Beitrag wird am Beispiel eines Forschungsprojekts zum Einsatz der Flipped Classroom Methode im Physikunterricht der Sek I gezeigt, wie sich mit der Plattform tet.folio online eine pseudonymisierte Testungebung gestalten lässt. Durch die Entwicklung eines so genannten *Action-Loggers* können nicht nur Ergebnisse, sondern alle von den Nutzern getätigte Aktionen mit einem Zeitstempel erfasst werden. Daraus ergeben sich neue Möglichkeiten zur Aufbereitung der erhobenen Daten aber auch zur Erhebung des Nutzungsverhaltens digitaler Unterrichtsmaterialien, beispielsweise von Lernvideos.

DD 48.3 Wed 12:40 DD-H10

Location: DD-H10

Concept Inventory Development - New lessons learned with the Flight Physics Concept Inventory (FliP-CoIn) — •FLORIAN GENZ^{1,3}, ANDRÉ BRESGES², KATHLEEN FALCONER², and LARS MÖHRING^{2,3} — ¹ZuS - Science Labs, University of Cologne / GERMANY — ²Universität zu Köln — ³ComeIN - Communities of Practice NRW

Concept Inventories surged in physics (e.g.: physport.org) and leaked fast into many other educational disciplines. However, Concept Inventories differ widely in the way they are developed and validated. We present particular important methods and techniques for Concept Inventory development as well as critical turning points by the practical example of the Flight Physics Concept Inventory (FliP-CoIn, https://zus.uni-koeln.de/flip-coin.html).

We shed a practical light on the methodological steps: Literature review and the decision about the underlying test theory, expert opinions, student concepts, think-aloud interviews, focus groups, data collection, data transformation, and analysis.

This will be a humbling talk with focus on avoidable mistakes...inspired by life experience, hence, past mistakes.

keywords: Research-based assessment instrument, conceptual test

Location: DD-H9

Location: DD-H11

DD 49: Lehreraus- und -weiterbildung - digitale Medien

Time: Wednesday 12:00-13:00

DD 49.1 Wed 12:00 DD-H11

Digitale Medien im Physikunterricht: Entwicklung eines Seminarkonzepts — •DAVID WEILER¹, JAN-PHILIPP BURDE¹, ANDREAS LACHNER¹, RIKE GROSSE-HEILMANN², JOSEF RIESE² und THOMAS SCHUBATZKY³ — ¹Universität Tübingen, Tübingen, Deutschland — ²RWTH Aachen, Aachen, Deutschland — ³Universität Graz, Graz, Österreich

Ein zentrales Ziel des Verbundprojekts DiKoLeP (Digitale Kompetenzen von Lehramtsstudierenden im Fach Physik) ist es, ein Lehrkonzept zu entwickeln, das die Kompetenzen zum fachdidaktisch sinnvollen Einsatz von digitalen Medien im Physikunterricht bei Lehramtsstudierenden fördert. Das für Graz und Tübingen entwickelte Seminar ist dabei zweigeteilt: In einer theoretischen Phase werden die Grundlagen des Einsatzes digitaler Medien inkl. entsprechender empirischer Befunde erarbeitet, während Studierende in der anschließenden praktischen Phase den Einsatz digitaler Medien im Rahmen exemplarischer Unterrichtssequenzen planen, umsetzen und reflektieren. Nach der Entwicklung des Seminars auf Basis von Literaturrecherche und Bedarfserhebung mit Studierenden wurde dieses im Sommersemester 2021 pilotiert. Dabei wurden der Kompetenzzuwachs der Studierenden und qualitative Rückmeldungen zu einzelnen Seminarsitzungen erhoben. Zudem werden aktuell retrospektive leitfadengestützte Interviews mit den Studierenden sowie Expertenbefragungen durchgeführt, um das Seminar zu optimieren. Im Vortrag werden das bisherige Seminarkonzept sowie die aus der Pilotierung und den Interviews gewonnenen Erkenntnisse vorgestellt und ein Ausblick auf das Re-Design gegeben.

DD 49.2 Wed 12:20 DD-H11

DiKoLeP: Digitale Kompetenzen von Lehramtsstudierenden im Fach Physik – •THOMAS SCHUBATZKY¹, JAN-PHILIPP BURDE², RIKE GROSSE-HEILMANN³, JOSEF RIESE³ und DAVID WEILER² – ¹Universität Graz, Graz, Österreich – ²Universität Tübingen, Tübingen, Deutschland – ³RWTH Aachen, Aachen, Deutschland

Digitale Medien spielen eine immer größer werdende Rolle im physikalischen Fachunterricht. Für eine lernförderliche Integration digitaler Medien braucht es aber dahingehend professionalisierte Lehrkräfte. Fachspezifische digitale Kompetenzen sollten dementsprechend auch in der fachdidaktischen Lehrerbildung gezielt adressiert werden. Im Verbundprojekt DiKoLeP der RWTH Aachen, der Universität Graz und der Universität Tübingen wird daher ein übergeordnetes Lehrkonzept mit standortspezifischen Ausprägungen entwickelt, implementiert und evaluiert. Durch dieses Lehrkonzept sollen fachspezifische, digitale Kompetenzen von Lehramtsstudierenden der Physik gefördert werden. Die Evaluation erfolgt einerseits im Hinblick auf den Erwerb physikdidaktischen Wissens zum Einsatz digitaler Medien im Physikunterricht, wozu ein entsprechender Leistungstest entwickelt wurde. Darüber hinaus werden die Entwicklung der Motivation und Überzeugungen und die Rolle des physikdidaktischen Wissens für die Entwicklung der Motivation zum Einsatz digitaler Medien untersucht. Insgesamt sollen so Hypothesen für lernförderliche, standortunabhängig einsetzbare Lerngelegenheiten in der Physik-Lehramtsausbildung abgeleitet werden. Vorgestellt werden die grundlegenden Ideen des Lehrkonzepts, das Studiendesign und der entwickelte Test.

DD 49.3 Wed 12:40 DD-H11 Digitalisierung beginnt im Kopf - Eine Akzeptanzstudie im Lehr-Lern-Labor Physik – •JOHANNES LHOTZKY und KLAUS WENDT – Johannes Gutenberg-Universität, Institut für Physik

Physiklehrer:innen sind technisch versiert, motiviert und nutzen dennoch digitale Medien nur sehr eingeschränkt. Diverse Untersuchungen haben Ressentiments gegen den Medieneinsatz auch bei Physiklehrkräften konstatiert. Um diesem Umstand zu begegnen, wurde an der JGU Mainz im Rahmen der vom BMBF geförderten Qualitätsoffensive Lehrerbildung ein speziell ausgerichtetes Lehr-Lern-Labor (LLL) entwickelt. Innerhalb dieser Lehrveranstaltungen werden die Studierenden zu einer vertieften Beschäftigung mit digitalen Medien mit dem Fokus auf Physikunterrichtsspezifika geführt und zur praktischen Auseinandersetzung mit diversen Medien von Calliope, Arduino zur programmierbaren Flugdrohne motiviert. Hierbei steht zudem die aktive Gestaltung von Medien/Medieninhalten im Vordergrund, die über eine Anwendungen fertiger Inhalte hinausgeht. Innerhalb des LLL entwickeln die Studierenden sowohl "klassische" als auch "digital-angereicherte" Experimentierumgebungen, die an Schüler:innen praktisch erprobt werden. Mithilfe eines Mixed-Methode-Ansatzes, bestehend aus Fragebogen und Gruppendiskussionen, wird in einem Prä-Post-Design die Wirkung des Seminarbesuchs identifiziert und evaluiert, speziell hinsichtlich Akzeptanz und Relevanzempfinden der Studierenden gegenüber dem Einsatz digitaler Medien im Physikunterricht. Im Vortrag wird das LLL skizziert, Forschungsergebnisse präsentiert und Implikationen für die Ausund Weiterbildung abgeleitet.

DD 50: Praktika und neue Praktikumsversuche

Time: Wednesday 12:00-13:00

DD 50.1 Wed 12:00 DD-H12

Experimente zur Förderung von Kompetenzen zur digitalen Messwerterfassung — •GREGOR BENZ, DANIEL ISELE, OLGA WALTER und TOBIAS LUDWIG — Pädagogische Hochschule Karlsruhe, Institut für Physik und Technische Bildung, Karlsruhe, Deutschland

Obwohl digitale Messwerterfassungssysteme (dMS) seit Jahrzehnten existieren, werden entsprechende Systeme im Physikunterricht nur selten eingesetzt (Wenzel & Wilhelm, 2017). Grund hierfür sind u.a. geringe Kompetenzen der Physiklehrenden im Umgang mit dMS. Daher sollen im Rahmen des QLB-Digitalprojekts InDiKo entsprechende Kompetenzen von Physiklehramtsstudierenden gefördert werden. Hierzu haben wir in Vorarbeiten zunächst 15 relevante Kompetenzen identifiziert (Benz et al., 2021). Vor diesem Hintergrund präsentieren wir nun drei speziell entwickelte Experimentiersettings, die zum Ziel haben, Kompetenzen zu a) der Funktionsweise von Sensoren, b) der Samplerate und c) der Auflösung zu fördern. In einer Interventionsstudie mit pre-post-Testdesign konnte anschließend Evidenz dafür gesammelt werden, dass entsprechende Kompetenzen zum Umgang mit dMS durch das Bearbeiten der Lernumgebungen erworben werden können (d=.51). Die entwickelten Experimentiersettings können nun dazu verwendet werden, angehende Physiklehrkräfte im Umgang mit dMS zu schulen.

DD 50.2 Wed 12:20 DD-H12

Naturwissenschaftliche Denk- und Arbeitsweisen in (physikalischen) Praktika – •JULIA ORTMANN¹, ANDREAS VORHOLZER² und NICOLE GRAULICH³ – ¹Institut für Didaktik der Physik, JLU Gießen – ²Didaktik der Physik, TU München – ³Institut für Didaktik der Chemie, JLU Gießen

Im Physikstudium sollen Studierende neben Fachinhalten auch naturwissenschaftliche Denk- und Arbeitsweisen (NDAW; z. B. Planung von Experimenten, Auswerten von Daten) erlernen. Zur Förderung von NDAW eignen sich insbesondere die physikalischen Praktika, da Studierende dort in der Regel umfassende Gelegenheiten zum selbstständigen Experimentieren erhalten. Es stellt sich jedoch die Frage, inwiefern NDAW in Praktika gezielt thematisiert werden und welche Relevanz NDAW im Vergleich zu anderen Lernzielen zugeschrieben wird. Um dieser Frage nachzugehen, wurden mit einem Online-Fragebogen N = 86 Praktikumsbetreuende (davon 27 Physik) und N = 399 Studierende naturwissenschaftlicher Studiengänge (davon 99 Physik) zu den Zielen von Praktika und der Rolle von NDAW befragt. Die Ergebnisse zeigen u. a., dass NDAW aus Sicht der Betreuenden eine große Rolle spielen, die Relevanz und die Maßnahmen zur Förderung sich aber zwischen einzelner NDAW z. T. deutlich unterscheiden. Aus Sicht der Betreuenden der Physik ist insbesondere das Auswerten und Reflektieren von Daten von großer Relevanz. Zudem zeigt sich, dass die Rolle von NDAW in Praktika von Lehrenden und Studierenden in der Physik sehr ähnlich wahrgenommen wird. Diese und weiter Ergebnisse sollen im Vortrag diskutiert werden.

DD 50.3 Wed 12:40 DD-H12

Experimentierpraktika: Lernangebote mit System oder System mit Lernangeboten? — •CORNELIA GELLER, MARTIN DICKMANN und HEIKE THEYSSEN — Universität Duisburg-Essen

Experimentierpraktika stellen in der universitären Ausbildung wesentliche Bausteine dar, die Physik-Studierende sowohl zu einer vertieften Auseinandersetzung mit Fachinhalten anregen als auch praktisch-methodische Grundlagen wie den Umgang mit Geräten oder Messunsicherheiten legen sollen. Sie bewegen sich damit in dem Spannungsfeld, fachinhaltlichen Kompetenzerwerb mit anderen Lernwegen innerhalb der Praktika in Einklang zu bringen, die eine systematische Entwicklung fachmethodischer Kompetenzen erlauben.

Da in der Lehramtsausbildung noch die fachdidaktische Perspektive hinzukommt, sind für diese Herausforderungen standortspezifische Lösungen entstanden, die sich nicht einfach mit Blick auf die Kohärenz der fachinhaltlichen, fachmethodischen und fachdidaktischen Lernangebote vergleichen lassen. Da-

Location: DD-H12

her soll im Vortrag eine Systematisierung von Praktikumskonzeptionen zur Diskussion gestellt werden, in die wir verschiedene Praktika des Standorts Essen einordnen. Vertieft werden soll die Diskussion am Beispiel eines Praktikums, in dem wir die Strukturierung der Lernangebote der studentischen Wahrnehmung (Ergebnisse der Evaluation) gegenüberstellen.

DD 51: Workshop: Konsequenzen aus drei Jahren Studienreformforschung

Time: Wednesday 16:00-17:30

Das Studienreform-Forum hat seit 2018 konkrete Studienreformen zusammengetragen und dokumentiert sowie vielerorts aufkommende Entwicklungsfragen bildungsphilosophisch reflektiert. In diesem Workshop werden Erkenntnisse zu rei Jahren Studienreformforschung Location: DD-H8

bereits (breit) diskutierten Fragen vorgestellt, um daraus ein Zwischenfazit zu ziehen. Ein Fokus soll darauf gelegt werden, Konsequenzen aus den Ergebnissen zu ziehen. Dies betrifft einerseits die Studienreformarbeit vor Ort, andererseits aber auch allgemeinere Empfehlungen für Akkreditierungsprozesse etc.

Geplante Kernthemen: - Entlinearisierung und Flexibilisierung von Studiengängen; Schaffung von (Wieder-)Einstiegsmöglichkeiten - Abschaffung von Klausurversuchsbeschränkungen - Systematische Feedbackschleifen als Bestandteil der Lehre - Relevanz und Förderung informeller Kommunikation

History of Physics Division Fachverband Geschichte der Physik (GP)

Peter Heering Abteilung für Physik und ihre Didaktik und Geschichte Europa Universität Flensburg Peter.Heering@uni-flensburg.de Christian Forstner Ernst-Haeckel-Haus - Wissenschaftsgeschichte Friedrich Schiller University Jena christian.forstner@uni-jena.de

19th Symposium of the History of Physics Division "Communicating Physics in History Communicating Physics through its History"

Writing papers, gauging instruments, teaching at school or chatting over coffee: these are just a few ways in which physical knowledge is communicated, and many more exist today or have existed in the past. In the last decades research in the history, sociology and didactics of the sciences has investigated the communication of scientific knowledge, underscoring the relevance for the development of existing fields, the emergence of new ones and the shaping of scientists' identities and communities. These reflections also apply to the past and present of the physical sciences and the contributions to the conference are invited to reflect all aspects of this diversity.

We understand here "communicating physics" in a broad sense, encompassing formal and informal communications among physicists, teaching at university and school level and public outreach. Beyond that, it also includes the transfer, appropriation and assimilation of knowledge between different cultures, such as interdisciplinary cooperation between physicists and scientists from other disciplines or knowledge transfer in colonial contexts or within processes of globalization.

Furthermore we want to address also another perspective of communication in physics: Communicating Physics through its History. We want to ask how the history of physics is used in the communication of physics today: The communication of physics through its history in formal education as well as through mass media, museums, archives and at public sites.

Overview of Invited Talks and Sessions

(Lecture hall GP-H7)

Plenary Talk by Allison Marsh

PV IV	Wed	9:00- 9:45	Audimax	Teaching with Objects and Teaching with Video: The Challenges of Informal Educa-
				tion in Physics — •Allison Marsh

Invited Talks

GP 1.1	Mon	13:30-14:00	GP-H7	The development of The Lorentz Lab: bringing the scientific history of Teylers Museum
				to life with working replicas — • Trienke van der Spek
GP 1.2	Mon	14:00-14:30	GP-H7	Physics in Information Comics — •Heike Elisabeth Juengst
GP 1.3	Mon	14:30-15:00	GP-H7	Jenseits gewohnter Pfade: Ausstellungen neu denken — •Christian Sichau

Invited Talks of the joint symposium The Nature of Science (SYNS)

See SYNS for the full program of the symposium.

SYNS 1.1	Tue	14:00-14:30	Audimax	The Role of Nature of Science Education for Science Media Literacy — •DIETMAR HÖTTECKE
SYNS 1.2	Tue	14:30-15:00	Audimax	What kinds of identities are deemed in/our of place in physics? — $\bullet Lucy$ AVRAAMI-
SYNS 1.3	Tue	15:00-15:30	Audimax	DOU Some thoughts on the status of theoretical physics — •DANIEL HARLOW

Sessions

GP 1.1-1.3	Mon	13:30-15:00	GP-H7	Communicating Physics and its History (joint session GP/DD)
GP 2.1-2.3	Mon	15:30-16:30	GP-H7	History and Teaching
GP 3.1-3.2	Mon	17:00-17:40	GP-H7	Physics and Media
GP 4	Mon	18:00-19:00	GP-MV	Meeting of Early Career Scholars
GP 5.1-5.5	Tue	10:30-12:30	GP-H7	Physics and the Museum
GP 6.1-6.3	Tue	16:15-17:15	GP-H7	Physicists as Popularisators
GP 7	Tue	17:45-19:45	GP-MV	Annual General Meeting
GP 8.1-8.3	Wed	10:30-11:30	GP-H7	Physics and Instruments
GP 9.1-9.4	Wed	15:40-17:00	GP-H7	History of Physics
GP 10.1-10.1	Wed	18:00-18:20	GP-H7	Physics and Culture

Annual General Meeting of the History of Physics Division

Tuesday, March 22, 2022 17:45–19:45 GP-MV

Come Together for Early Career Scholars

Monday, March 21, 2022 18:00–19:00 GP-MV Ansprechpartnerin: Julia Bloemer (j.bloemer@deutsches-museum.de)

Sessions

Invited and Contributed Talks –

GP 1: Communicating Physics and its History (joint session GP/DD)

Time: Monday 13:30-15:00

Invited Talk

GP 1.1 Mon 13:30 GP-H7 The development of The Lorentz Lab: bringing the scientific history of Teylers Museum to life with working replicas — •TRIENKE VAN DER SPEK — Teylers Museum, Haarlem, The Netherlands

In 2017 Teylers Museum opened The Lorentz Lab, a new extension to the permanent presentation of the museum. The Lorentz Lab is specifically dedicated to revive the institute's scientific past as a research institute and laboratory. The scientific instrument collection has long been a silent witness of this past, as are the old laboratory buildings. The Lorentz Lab brings the unique origins of this heritage into life in a carefully recreated historical setting of Teylers' Physics Laboratory.

Here working replica's, theatrical support and a tailor made educational program allow different groups of visitors to participate in science directly and engage with the activities and scientists that shaped Teylers Museum and its collections. The program The Lorentz Formula is dedicated to the general museum visitor, and Einstein was here is a high-level physics program, designed to provoke questions that fit in with educational requirements and with links to the school curriculum.

This lecture will discuss the development of the Lorentz Lab and the central role of working replicas therein. It will address its challenges, original goals, experiences and insights - both from the general public and the educational programs' point of view.

GP 1.2 Mon 14:00 GP-H7 Invited Talk Physics in Information Comics — •Heike Elisabeth Juengst — FHWS Würzburg

Information comics are comics designed primarily for knowledge transfer. They can be funny or serious, long or short, and can be found all over the world. And,

GP 2: History and Teaching

Time: Monday 15:30-16:30

GP 2.1 Mon 15:30 GP-H7 Transformations: On the relation between research experiments and teach-

ing demonstrations — •PETER HEERING — Europa-Universität Flensburg Some experiments from the history of physics were so relevant that they were not only included in textbooks, but also found their way into (university) physics education as teaching demonstrations. The instruments developed for this purpose can be found in a number of museum collections, whereby frequently hardly any differentiation is made between teaching devices and research instruments. However, a somewhat closer analysis of the instruments and the practice associated with them makes it clear that significant differences can be found here. In the context of this paper, I am going to discuss some aspects of such demonstration devices that prove relevant precisely in distinguishing them from the corresponding experimental instruments and the practices associated with them.

GP 2.2 Mon 15:50 GP-H7

Bringing some light into the dark and some darkness into light: Young's double-slit experiment (1807) — • MICHELLE MERCIER — Europa-Universität Flensburg

In 1807, Thomas Young published the description of an experiment that is nowadays canonized as Young's double-slit experiment. Today, the basic principle of the experiment is well-known, however, the experiment performed by Young is not. And if one goes back to the initial description, his text is difficult to understand and leaves several questions unanswered both in respect to the details of the apparatus he used and the exact observations in his experiment.

As part of my PhD project, Young's double-slit experiment is analyzed by using the replication method. In this talk, I will describe the experiences made (experimentally) and focus in particular on the difficulties in observing as well of course, they can cover any topic.

However, information comics about physics are not very common. "The Physics of Superheroes" by Kakalios is justly popular but does not deal with information comics at all. Chemistry, on the other hand, is a popular topic for information comics.

The reason for this imbalance is a mystery and cannot be solved in this presentation. Information comics presenting topics from physics will be shown and the audience will be encouraged to produce their own information comics for teaching purposes (and for fun).

Invited Talk GP 1.3 Mon 14:30 GP-H7 Jenseits gewohnter Pfade: Ausstellungen neu denken — •CHRISTIAN SICHAU - experimenta gGmbH, Heilbronn

Lange Zeit galten Science Center als Ausstellungsorte, in denen über die Darstellung natürlicher Phänomene Begeisterung für Naturwissenschaft und Technik geweckt werden sollte. Obwohl hinsichtlich der Besucherzahlen recht erfolgreich, wurde ihr Ansatz immer wieder als unzureichend kritisiert. Science Center würden kein adäquates Verständnis der Natur der Naturwissenschaften vermitteln, ihnen mangele es an kritischer Reflektion und an einer ernsthafter Auseinandersetzung mit der Wissenschaft und ihrer Geschichte. Bei dieser gelegentlich scharf und polemisch vorgetragenen Kritik wird häufig übersehen, dass es seit vielen Jahren und auf internationaler Ebene zahlreiche neue Ansätze gibt. So wird immer mehr und intensiver in der Science Center-Szene diskutiert, wie das öffentliche Verständnis für und über Wissenschaft vermittelt und gestärkt werden kann. Hierbei kann gerade die Stärke der Science Center - das Erreichen einer sehr breiten Öffentlichkeit mit niedrigschwelligen Angeboten - ein wichtiger Pluspunkt werden. Am Beispiel der experimenta soll aufgezeigt werden, dass hier neue - und vielleicht sehr wirksame - Narrative entstehen können.

as documenting and communicating what is seen. The meaning of the experiences made for the understanding of Young's description of the experiment will be discussed in conclusion.

GP 2.3 Mon 16:10 GP-H7

Location: GP-H7

Garavito s work in Colombia on the theory of light aberration: some didactic reflections for the teaching of physics — •LISBETH ALVARADO-GUZMAN^{1,2}, ISABEL MALAQUIAS², and ROBERTO NARDI¹ — ¹São Paulo State University (Unesp), School of Sciences, Bauru, Brazil - ²University of Aveiro (UA), Dep. Physics, CIDTFF, Aveiro, Portugal

This article aims to recognize some elements about the scientific research developed in Colombia at the beginning of the 20th century and the communication with the international scientific community of the time. Julio Garavito Armero s (1865-1920) original article on the theory of light aberration (1912) is taken as a paradigmatic example and with didactic interest. He was the first professor of mathematics (Arboleda, 2021), graduated as an engineer, also director of the Astronomical Observatory in Colombia (1893-1919). From the original article, the problem Garavito addresses was the one proposed by the astronomer David Gill (1843-1914) in 1896: if one may consider as exact the generally accepted theory of aberration. Garavito concluded that the annual aberration constant was correct and considered incorrect Huygens interpretation of the superposition of waves effect, using the analogy of water waves (Lleras, 1915). To conclude, the problem addressed by Garavito can have a special interest in the teaching of Physics and Astronomy context from: 1) the analysis of the phenomenon of aberration of light as an argument in favor of Earth s motion; 2) questions on the nature of science connected with the communication by and between researchers.

Location: GP-H7

GP 3: Physics and Media

Time: Monday 17:00-17:40

GP 3.1 Mon 17:00 GP-H7

From the Preprint Information Exchange to "Computopia:" the development of preprint culture in High Energy Physics — •ARIANNA BORRELLI uphana University Lueneburg, Universitaetsallee 1, 21335 Lueneburg

The term "preprint" has recently become known also among the general public as indicating research papers which are circulated in digital form, but have not (yet) gone through the peer review process. In particular, debates in the biomedical sciences on whether and how far such results should play a role in the management of the pandemics have found large resonance in the media. A look back at the history of printed and digital preprints shows that such discussions are not new and often found a different outcome in the biomedical field, where preprints were long discouraged if not forbidden, and in physics, where preprints became an increasingly important means of academic communication already in the late 20th century.

High-energy physicists were path-breaking in this development and it has been suggested that they traditionally had a more open approach to the diffusion of knowledge than life scientists. Is this really the case? I will offer a brief overview of the growth of preprint culture in physics from the 1960s and its eventual move to the digital sphere, which physicist David Mermin in 1991 hailed as "Computopia." Comparing these developments with those in other academic fields, I will ask about the historical factors which may help explain the differences and similarities.

Monday

GP 3.2 Mon 17:20 GP-H7

Poggendorffs Briefe an seinen Verleger Barth: Eine neue Quelle zu den Annalen der Physik und Chemie — • MICHAEL BARTH — Nordstraße 7 31249 Hohenhameln

J.C.Poggendorff war von 1824 bis 1876 Herausgeber der Annalen der Physik und Chemie, der mit Abstand bedeutendsten Fachpublikation im 19. Jahrhundert im deutschen Sprachraum. Die 147 Briefe, die er im Zeitraum von 1824 - 1859 an J.A.Barth, Verleger der Annalen, schickte, wurden mir vor längerer Zeit überraschend zugänglich gemacht. Für den Zeitraum 1830 bis 1845 habe ich davon 51 Briefe transkribiert und in meiner Dissertation publiziert. Sie dienten mir vor allem zur möglichst exakten Datierung der Annalen, um die Rezeption von Faradays Arbeiten im deutschen Sprachraum genau zu verfolgen.

Diese bislang nach meinem Wissen nicht publizierte Quelle hat aber noch erheblich mehr Potential. Sie liefert für mehrere Dekaden detaillierten Einblick in den alltäglichen Ablauf von Poggendorffs Herausgebertätigkeit, den konkreten Druck.- und Publikationsprozess, Finanzfragen, technische Probleme, Einschätzungen von Kollegen, Animositäten, privaten Erlebnissen, Reisen usw. usw., dies auch für Poggendorffs bekanntes Biographisch-literarisches Handwörterbuch. Alles Informationen, die interessant für die unterschiedlichste Nutzung in der wissenschaftshistorischen Forschung sein können.

Im Vortrag werde ich diese Quelle vorstellen und ihr Potential diskutieren. Dabei werde ich auch kurz auf sehr interessante Passagen der Korrespondenz zwischen Poggendorff und Schönbein eingehen, die in den bisherigen Publikationen fortgelassen wurden.

GP 4: Meeting of Early Career Scholars

Time: Monday 18:00-19:00

Virtual Come Together

GP 5: Physics and the Museum

Time: Tuesday 10:30-12:30

GP 5.1 Tue 10:30 GP-H7

Physik erleben durch "Ein naturkundliches Spiel-System". Hugo Kükelhaus und die deutsche (Vor-)Geschichte des Science Centers — •ARNE SCHIRRMAснек — Humboldt-Universität zu Berlin

Mehr und mehr haben die Science Center die Vermittlung von Naturwissenschaften und damit insbesondere von physikalischen Phänomenen übernommen, häufig noch bevor Schule oder Museum dies tun. Während die Vermittlung über historische Arefakte in Museen eine europäische Errungenschaft war und etwa in Paris, München oder London Modellcharakter bekam, entstand die heutige Form des Science Centers Ende der 1960er Jahre in Nordamerika. In meinem Vortrag möchte ich mit Hugo Kükelhaus einen unkonventionellen deutschen Weg zur Vermittlung von physikalischen Phänomenen vorstellen, die dieser seit einem Besuch im Deutschen Museum im Juli 1965 entwickelte und mit Unterstützung von Otto Hahn und der Max-Planck-Gesellschaft realisierte, u.a. für die Weltausstellung 1967 in Montreal. Das Projekt eines "Phänodrom" scheiterte indes. Lagen die Probleme bei der deutschen Entwicklung zu einem Science Center in Kükelhaus' Biographie oder in dem Wandel der deutschen Wissenschaftskultur?

GP 5.2 Tue 10:50 GP-H7 Von Fehlergrenzen und Straßenbahnen: Die Experimente von Alice Golsen

zum Strahlungsdruck des Lichts - • JOHANNES-GEERT HAGMANN - Deutsches Museum, München, Deutschland

Laserkühlung, Optischen Pinzetten, Sonnensegel: für eine Vielzahl von physikalischen Konzepten und Anwendungen kommt der durch die elektromagnetische Strahlung vermittelte Druck zum Tragen. Die Geschichte der Entdeckung und Messung des Strahlungsdrucks wird gelegentlich auf die Arbeiten von Pjotr Nikolajewitsch Lebedew (1886-1912), Ernest Fox Nichols (1869-1924) und Gordon Ferrie Hull (1870-1956) reduziert. Der vorliegende Beitrag ordnet die Arbeiten der Physikerin Alice Golsen (1889-1940), die in Frankfurt gemeinsam mit Walther Gerlach (1889-1979) forschte, in diese Untersuchungen ein. Die damit verbundene Rekonstruktion der Biografie Golsens ist Teil der Vorbereitungen zur geplanten Ausstellung *Licht und Materie* im Deutschen Museum.

Location: GP-MV

Location: GP-H7

GP 5.3 Tue 11:10 GP-H7

Elsa Garmires kohärente Kunst: Eine interaktive Laserinstallation als di $daktisches \ Ausstellungselement - \bullet \texttt{Eckhard Wallis} - \texttt{Deutsches Museum},$ München, Deutschland

Das Ausstellungsprojekt "Licht und Materie" am Deutschen Museum verfolgt das Ziel, die Quantenphysik des Lichts auch Besucherinnen und Besuchern zugänglich zu machen, die keine ausgeprägte Affinität zur Physik mitbringen. Ein Beispiel für einen historischen Zugang zur Thematik sind die Arbeiten der Laserpionierin Elsa Garmire (geb. 1939). Neben ihrer physikalischen Arbeit am Caltech war Garmire ab den späten 1960er Jahren auch an den ersten experimentellen Kunstprojekten mit Laserlicht beteiligt. Inspiriert von Garmires Projekten ist für die Ausstellung eine interaktive Laser-Installation geplant, die die Besucherinnen und Besuchern zu physikalischen und historischen Erkundungen einladen soll. In diesem Beitrag gehe ich der Frage nach, welche Lerneffekte wir durch die historische Perspektive erwarten.

Coffee Break

GP 5.4 Tue 11:50 GP-H7

Label: Fallacy. Communicating Nature of Science in a Museum Exhibit -•JULIA BLOEMER — Deutsches Museum, Munich, Germany

Physics often seems to be incomprehensible and divorced from reality. Highspecialized instruments rarely help to bridge the gap in science communication. The interferometer built by Georg Joos in 1930 is one example. It is a person-high steal construct with four arms and a complex mirror construction inside. Meant to measure the hypothetic aether and to answer the question about a medium for light propagation, it is the end of a long line of different interferometer experiments since the 1880s with larger and larger instruments. Finally, special relativity theory replaced any need for a luminiferous aether and the hypothesis was abandoned. Why should anybody care about an old and gigantic historical instrument connected with an outdated theory? In the past, the Joos-interferometer served as a museum object to transport several different messages: to explain the theory of relativity and its history or to emphasize the high-standard of the German optical industry in the 1930s. This talk presents a different perspective. Especially instruments can tell stories about nature of science aspects, about the way physicists ask and answer questions. In this case, which role does fallacy play? How can repetition stabilize knowledge? In times of science skepticism, this is as important as never before.

GP 5.5 Tue 12:10 GP-H7 Mont Blanc, the laboratory of 18th century Geneva scientists — •Stéphane FISCHER — Musée d'histoire des sciences de Genève, 128, rue de Lausanne, 1202 Genève

At the end of the 18th century, through a combination of political, social and cultural circumstances, Geneva became the scientific capital of the Alps. In less than fifty years, several of the city's scientists embarked on a scientific exploration of the Savoyard Pre-Alps and the Mont Blanc Massif, 60 km away. The mountain became a veritable open-air laboratory where various measurements were taken: altitude, air or water boiling temperatures, topographical surveys, purity of the atmosphere, composition of the gases in the atmosphere at altitude,

GP 6: Physicists as Popularisators

Time: Tuesday 16:15-17:15

GP 6.1 Tue 16:15 GP-H7

The role of aesthetics in science communication: History and histories in Werner Heisenberg's popular writings — •ELENA SCHAA — Trinity College Dublin, Dublin, Ireland

In 1969, Werner Heisenberg published his memoir Der Teil und das Ganze. Gespräche im Umkreis der Atomphysik. Driven by the motivation to empower the lay audience to engage in the "philosophical, ethical, and political discussions" arising from modern physics, he sets out to share his recollections of the development of modern physics.

Departing from a close examination of Heisenberg's bestseller, the paper discusses how Heisenberg used history and historization to communicate physics beyond the scientific community while taking on a double role as a spectator and a narrator of the history of physics. The paper analysis first, the different ways history becomes relevant Heisenberg's narratives. Secondly, focusing on the form of his popular writings, the paper highlights the interference of science and religion, specifically the Romantic aesthetics of immediate experience of nature and knowledge production. Ultimately, I situate Heisenberg's communication of physics in the wider context of the German Bildungsbürgertum together with the masculine ideal of academic authority within which his popularisation of modern physics becomes effective.

GP 6.2 Tue 16:35 GP-H7 Writing History as a Way to Teach Physics: Edmund T. Whittaker's two editions of the A History of the Theories of Ether and Electricity, 1910 and 1951-

3. — •JAUME NAVARRO — University of the Basque Country, Spain At the end of his career as a physicist and mathematician, Edmund T. Whittaker (1873-1956) decided to prepare a second, much enlarged version of his by then classic 1910 book A History of the Theories of Ether and Electricity. While the first edition had received nothing but praise, the second, two-volume edition had mixed reviews, especially due to its treatment of Einstein's role in the development of special relativity. In a recent paper I have analysed the reasons behind his moderate portrayal of Einstein in the history of twentieth century physics etc.This quest culminated in the summer of 1788 with the ascent of Mont Blanc by the Genevan naturalist Horace-Bénédict de Saussure.

The Museum of the History of Science is devoting its next temporary exhibition to this scientific epic. The heart of this exhibition is made up of instruments from the collections, in this case the instruments from the Saussure collection kept at the Museum. The exhibition is an ideal medium for tracing a storyline, a narrative that allows the context in which these instruments were invented and manufactured at the time. A great deal of attention is paid to the functioning of these instruments. Alongside the old barometers, two modern replicas of mercury barometers taken this summer to heights of more than 3000m for altitude measurements will be presented. Various interactive experiments offer visitors the opportunity to learn the basics of barometric levelling, surveying and slope inclination in a playful way.

Location: GP-H7

(Navarro 2021). Yet, much has to be studied about the role both editions played in the teaching of physics to generations of physicists, historians of physics and physics aficionados. In this paper I intend to track the implicit agendas of Whittaker in the two editions of his book and the ways they were received. I shall argue that A History of the Theories of Ether and Electricity had an historiographical vision on how to use history to teach the contents of physics, to shape a discipline at a time of profound transformations and to portray physics as a collective and non-teleological enterprise. The highly unexplored correspondence with his son, the also mathematician John Whittaker, helps us give this more complex picture of the A History of the Theories of Ether and Electricity.

GP 6.3 Tue 16:55 GP-H7 Spacetime as popularised by Arthur S. Eddington - •FLORIAN LAGUENS -IPC-Facultés Libres, 70 avenue Denfert-Rochereau, 75014 Paris, France Arthur S. Eddington (1882-1944) certainly was the world's most famous astronomer during the interwar period. For thirty years he was the director of Cambridge Observatory and a Fellow of Trinity College. He also plunged into philosophy while discovering Einstein*s general relativity in 1916. From then on, he developed some personal thoughts about physics, its methods and its limits. Along with widely acclaimed scientific treatises, Eddington published some controversial books such as The Nature of the Physical World (1928), Relativity Theory of Protons and Electrons (1936) and The Philosophy of Physical Science (1939). In particular, The Nature of the Physical World is still considered a masterpiece regarding the popularisation of general relativity theory and quantum mechanics. Indeed, spacetime is discussed at length in several chapters. This paper intends to highlight, thanks to key passages of Eddington's works, his very conception of popularisation. It then allows to exemplify its role in helping both students and colleagues coping with the relativity major conceptual changes. Finally, Eddington's attitude towards popularisation reveals is way of considering the relationship between physics and mathematics. All in all, as he replies to some critics, his aim is to "convey exact thought in inexact language" (New Pathways in Science, 1935).

GP 7: Annual General Meeting

Time: Tuesday 17:45–19:45 Annual General Meeting

GP 8: Physics and Instruments

Time: Wednesday 10:30-11:30

GP 8.1 Wed 10:30 GP-H7

The Communication of Object-Bound Knowledge in Networks of Lead-Users — •CHRISTIAN FORSTNER — Ernst-Haeckel-Haus, Friedrich_Schiller-University, Jena, Germany

Research technologies circulate between different fields of society and are continuously adapted to new contexts in a process of dis- and re-embedding. But how is the tacit knowledge that is tied to their practical use made explicit? Does it reveal itself to the user? What strategies were taken by the instrument makers to make tacit knowledge accessible to the user? In my talk I will discuss a historical example of communicating physical practices. Therefore I analyze the optical measurement instruments of the Carl Zeiss Company, in particular analytical interferometers, to show how a network of lead users was created and how practical knowledge circulated in this network. Location: GP-MV

Location: GP-H7

GP 8.2 Wed 10:50 GP-H7

Analysis of Portable Quadrants from Different Cultures — •ENES TEPE — Europa-Universität Flensburg, Auf dem Campus 1, 24943 Flensburg

Portable quadrants (*rub^cal-dā²iras*) are instruments that were used for purposes such as astronomical observations, timekeeping, navigation, surveying, maritime, ballistics, and mathematical calculations for more than a Millennium. In a previous study (my MA thesis), the role of this class of instruments in classical astronomy was analyzed within the context of a comparison between the Islamic World and the Western Europe. For this purpose, information of surviving instruments that can be found through the online catalogues of museums and auctions were examined. Therefore, it has been shown that the astronomical portable quadrant traditions in Mamluk, Maghreb, Iran, Ottoman Syria, Ottoman, Continental Europe, Italy and England can be distinguished and studied in detail. In my current project, two of the most well-established portable quadrants, one from the Islamic World and the other from the West, are chosen in order to analyze the respective practice with the replication method. One of them is a quadrant of almucantars (*rub^cal-muqantarāt*) for Damascus by Zayn al-Dīn [Shams al-Dīn] Abū 'Abd Allāh Muhammad ibn Ahmad ibn 'Abd al-Rahīm al-Mizzī (d. 1349). The other one is a Sutton-type large quadrant by Henricus Sutton Londini (d. 1665) for London. In this talk, the prominence of these two instruments will be discussed and their general features of them will be shared.

GP 8.3 Wed 11:10 GP-H7

Binocular or stereoscopic telemeters? Two countries, two concepts -ANDREAS JUNK — Europa-Universität Flensburg

GP 9: History of Physics

Time: Wednesday 15:40-17:00

GP 9.1 Wed 15:40 GP-H7

M. I. Kaganow und die Elektronentheorie der Metalle — •Peter Bussemer¹ und VLADIMIR RZHEVSKIJ 2 – 1 Cooperative University Gera-Eisenach -²Lomonosov-University Moscow

Moisej I. Kaganow (1921-2019) war einer der letzten Repräsentanten der sowjetischen Schule der Festkörpertheorie, begründet in Charkow um 1930 von Lew Landau und dort bis 1970 fortgeführt von Ilja M. Lifschitz (1917-1982), danach in Moskau. Unter dessen Leitung war Kaganow wesentlich an modernen Entwicklungen beteiligt: Quantenmechanik von Elektronen mit beliebigem Dispersionsgesetz, Topologie der Fermi-Flächen (Fermiologie) mit der Möglichkeit offener Elektronenbahnen, Oszillationen thermodynamischer Größen im Magnetfeld, HF-Eigenschaften-veröffentlicht in Deutsch 1975 in "Elektronentheorie der Metalle" mit Lifschitz und Asbel. Seine Untersuchungen zu den Lifschitzschen topologischen Phasenübergängen 2 1/2-ter Ordnung sind ein früher Vorgriff auf die aktuellen Anwendungen der Topologie bei Phasenübergängen. Kaganow war ein brillanter Hochschullehrer: Von 1970-1994 bildete er an der Moskauer Universität mehrere Studentengenerationen aus. Ebenso ein engagierter Wissenschaftsvermittler, verfasste er zahlreiche populäre Einführungen in die Quantenund Festkörperphysik mit Übersetzungen in Deutsch, Englisch und Polnisch. Enge Kontakte pflegte er zur DDR und Polen: Dr.h.c. TU Wroclaw 1988. Nach 1994 lebte er als jüdischer Emigrant in den USA.

GP 9.2 Wed 16:00 GP-H7

John Herschel, not only an astronomer, in the footsteps of father and aunt -•HARALD GROPP — VIGN,Heidelberg, Germany

150 years ago, on May 11, 1871 John Herschel died in Hawkhurst, Kent. He is mainly known as an astronomer but also contributed to mathematics, chemistry, philosophy of science and related fields. Together with the astronomical couple, his aunt Caroline (1750-1848) and his father William (1738-1822), John(1792-1871) covers a period of more than 130 years of astronomical research. Caroline is probably the least known of them. She not only assisted her male colleague (brother) William, but was paid a salary for her work.

In this paper the cooperation of 3 astronomers will be discussed, between William and Caroline directly, but also independently, between William and John by John following in his father*s footsteps. Letters between John in England (and South Africa) and his aunt Caroline back in Germany played a main role in the last period. They were edited by John*s wife, *Mrs. John Herschel* (1810-1884) and tell interesting details about the *collective work*

GP 9.3 Wed 16:20 GP-H7

"Reducing physics to pure mathematics": Johann I Bernoulli on Brook Taylor's taut string problem — •IULIA MIHAI — Ghent University (Belgium)

The development of rangefinders in Europe towards the end of the 19th century seems to have been a competition between the British company Barr&Stroud and their German counterparts Carl Zeiss Jena. These companies used different approaches for their so-called telemeters whilst the instruments looked very much alike.

Barr & Stroud preferred a binocular approach for a coincidence rangefinder, Zeiss in turn designed a stereoscopic rangefinder. Whilst the preliminary test results indicated, that the stereoscopic approach would produce better results, Barr & Stroud insisted, that these results could only be produced by users, who did not have certain physiological handicaps. In my paper, I want to line out the (dis)advantages of the instruments as well as the respective motivation and conditions for their construction.

Location: GP-H7

The vibrating string problem is outstanding in the history of eighteenth-century mathematical physics. This paper focuses on the conception of the continuous string that was dominant for three decades before the use of partial differential equations at the end of the 1740s. Scholars have emphasized the novel mathematical techniques in Johann I Bernoulli's reappraisal (1732) of Brook Taylor's initial investigation (1713), but the standard view has it that the conception of the string with which they work remains unchanged. By contrast, this paper argues that how the string is conceived evolves due to the conceptual changes brought about by Bernoulli's scientific practice. Whereas Taylor approaches the string (also) by drawing analogies with other mechanical objects on the basis of shared (physical and geometrical) properties, Bernoulli steers clear of physical analogies in investigating the string's properties. Moreover, Bernoulli's extensive use of algebraic symbolism enables innovative notational interventions which result in a more robust handling of both physical and geometrical quantities; this goes beyond the fact that Bernoulli uses the differential calculus and Taylor the fluxional calculus. Ultimately, it is Bernoulli's methodology of "reducing physics to pure mathematics" which is behind the evolving conception of the string.

GP 9.4 Wed 16:40 GP-H7

Location: GP-H7

Black Hole Imaging and Framings of the Observer - • Emilie Skulberg -Institute of Physics, University of Amsterdam

Based on the study of an extensive collection of visual representations of black holes from 1970 to the present, I trace the history of visual and textual framings of the observer in the context of black hole imaging. I argue that the framing of the observer changed significantly in this period. Some peer-reviewed papers containing early visual representations of the immediate surroundings of black holes had brief references to telescopic observation. More commonly, such images were framed as part of thought experiments of what an observer (appearing in thought experiments as a hypothetical human being) would see or photograph if equipped with a camera directed at a black hole. The observer here became a point of view in a visual sense as images showed what an observer "saw" or "photographed". Towards the new millennium, visualizations from simulations began to form part of arguments that observing the shadow of a black hole would in fact be possible. Rather than an observer with a camera in thought experiments, visualizations showed how a specific celestial object believed to be a black hole might look if observed from Earth using Very Long Baseline Interferometry. This method, in which data from observations by multiple telescopes placed far apart are combined, was what would later enable the Event Horizon Telescope Collaboration to produce images such as the first observation of the shadow of a black hole. At the same time, virtual reality now offers the immersive and embodied experience of seemingly being an observer approaching a black hole.

GP 10: Physics and Culture

Time: Wednesday 18:00-18:20

GP 10.1 Wed 18:00 GP-H7 European and American Research Traditions in the 20th Century -•ALEXANDER UNZICKER — Pestalozzi-Gymnasium München While the European research tradition, in the spirit of natural philosophy, was

focused on the fundamental laws of nature and pursued the question of 'what holds the world together at its innermost folds, in the US, a technologically oriented culture dominated, with a desire to realize large, visionary projects, such as the atomic bomb and landing on the moon. It is argued that this change in scientific culture is still visible in contemporary physics.

Working Group on Physics, Modern IT and Artificial Intelligence Arbeitskreis Physik, moderne Informationstechnologie und Künstliche Intelligenz (AKPIK)

Tim Ruhe TU Dortmund Otto Hahn-Straße 4a 44227 Dortmund tim.ruhe@tu-dortmund.de

Overview of Invited Talks and Sessions

(Lecture hall AKPIK-H13)

Sessions

AKPIK 1.1–1.9	Mon	16:15-18:30	AKPIK-H13	Data Integration & Processing
AKPIK 2.1-2.9	Wed	16:15-18:30	AKPIK-H13	Data Analytics & Machine Learning
AKPIK 3	Wed	19:00-21:00	AKPIK-MV	Mitgliederversammlung AKPIK
AKPIK 4.1-4.9	Thu	16:15-18:30	AKPIK-H13	Deep Learning

Annual General Meeting of the Working Group on Physics, Modern IT and Artificial Intelligence

Wednesday 19:00-21:00 AKPIK-MV

- 1. Bericht des Vorstandes
- 2. Entlastung des aktuellen Vorstandes
- 3. Wahl des neuen Vorstandes
- 4. Die Zukunft in unserer Hand! (Anregungen und Wünsche)

Sessions

– Talks –

AKPIK 1: Data Integration & Processing

Time: Monday 16:15-18:30

AKPIK 1.1 Mon 16:15 AKPIK-H13

The PUNCH4NFDI Consortium in the NFDI - status, first results and outlook — •THOMAS SCHÖRNER for the PUNCH4NFDI-Collaboration — Deutsches Elektronen-Synchrotron DESY, Notkestr. 85, 22607 Hamburg

With the "Nationale Forschungsdateninfrastruktur" (NFDI, national research data infrastructure), a massive effort is undertaken in Germany to provide a coherent research data management, to make research data sustainably utilisable and to implement the FAIR data principles. PUNCH4NFDI is the consortium of particle, astro- and astroparticle, as well as hadron and nuclear physics within the NFDI. It aims for a FAIR future of the data management of its community and at harnessing its massive experience not least in "big data" and "open data" for the benefit of "PUNCH" sciences (Particles, Universe, NuClei and Hadrons) as well as for physics in general and the entire NFDI. In this presentation, we will introduce the work programme of PUNCH4NFDI, its connection to every-day work in the physical sciences and beyond, and in particular the idea of digital research products and the PUNCH science data platform.

AKPIK 1.2 Mon 16:30 AKPIK-H13

Community Initiative for a VHE Open Data Format — •MAXIMILIAN NÖTHE¹ and LARS MOHRMANN² — ¹Astroparticle Physics WG Elsässer, TU Dortmund University — ²Max-Planck-Institut für Kernphysik, Heidelberg

The operation of the next-generation gamma-ray telescopes as observatories, the wish of currently operating instruments to archive and publish their data in an accessible format, and enabling multi-instrument analyses are strong reasons for developing an open, software independent format for gamma-ray data.

A first attempt of a common specification has been developed by members of different Imaging Atmospheric Cherenkov Telescopes (IACT) within the "Data formats for gamma-ray astronomy" initiative. The current version defines formats for high-level gamma-ray data, including event lists of candidate photons and instrument response functions, serialized as FITS files.

Open-source software for gamma-ray analyses, including gammapy and ctools, have recently developed support for this format and, as a result, a series of publications relying on standardized datasets and software have been issued.

Currently, an effort to formalize the endeavor is underway, creating a Coordination Committee formed from representatives of the participating instruments to steer the future development of the specification.

In this talk, current developments and future plans will be presented, including the already implemented extension to ground-based wide-field experiments and possible extension to other messengers.

AKPIK 1.3 Mon 16:45 AKPIK-H13

From sample management to workflow integration: Semantic research data management with CaosDB — •DANIEL HORNUNG¹, FLORIAN SPRECKELSEN¹, and JOHANNES FREITAG² — ¹IndiScale Gmbh, Göttingen — ²Alfred Wegener Institute, Helmholtz Centre for Polar and Marine Research, Bremerhaven Organizing data from a diversity of sources, from acquisition to publication, can be a tough challenge. We present research data management implementations using the flexible open-source toolkit CaosDB at the Alfred Wegener Institute. CaosDB is used in a diversity of fields such as turbulence physics, legal research, animal behavior and glaciology. CaosDB links research data, makes it findable and retrievable, and keeps data consistent, even if the data model changes.

In the presented example, CaosDB keeps track of ice core samples and to whom samples are loaned for analyses. It made possible additional features such as: A revision system to track all changes to the data and the sample state at the time of analysis. Automated gathering of information for the publication in FAIR-DO meta-data repositories, e.g. Pangaea. Tools for storing, displaying and querying geospatial information and graphical summaries of all analyses performed on each ice core. Automatic data extraction and refinement into data records in CaosDB to minimize manual users interaction. A state machine which guarantees certain workflows, simplifies development and can be extended to trigger additional actions upon transitions.

We demonstrate how CaosDB simplifies semantic data in science and enables advanced data processing and understanding.

AKPIK 1.4 Mon 17:00 AKPIK-H13

CaosDB – a scientific research data management toolkit — •DANIEL HORNUNG¹, FLORIAN SPRECKELSEN¹, and JOHANNES FREITAG² — ¹IndiScale Gmbh, Göttingen — ²Alfred Wegener Institute, Helmholtz Centre for Polar and Marine Research, Bremerhaven

Processing interconnected, multi-modal data poses a challenge in many fields, especially when the data model, i.e. the way how data is organized, changes over time or when its structure is poorly documented. The open-source software **CaosDB** is a toolkit for research data management which was originally developed at the Max Planck Institute for Dynamics and Self-Organization (Göttingen) because existing software could not fulfill the needs of the scientists.

We present examples where CaosDB helped make data FAIR (Findable, Accessible, Interoperable, Retrievable) and how it can simplify the workflows for researchers: Automated data collection and integration, export to data repositories, API libraries for third-party programs, integrated revisioning and workflow state machines. If the data model needs to change, existing data can remain asis and future search queries will return matching results containing "old" and "new" data. We demonstrate how raw and processed data, analysis settings and results, and even labnotebooks and publications can be linked against each other, to improve long-term usability of data and reproducibility of results.

AKPIK 1.5 Mon 17:15 AKPIK-H13

Data curation in astroparticle physics data centers on example of KCDC and GRADLCI — •VICTORIA TOKAREVA¹, ANDREAS HAUNGS¹, DORIS WOCHELE¹, JÜRGEN WOCHELE¹, FRANK POLGART¹, ALEXANDER KRYUKOV², MINH-DUC NGUYEN², ANDREY MIKHAILOV³, and ALEXEY SHIGAROV³ — ¹Karlsruhe Institute of Technology, IAP, 76021 Karlsruhe, Germany — ²Moscow State University, SINP, Moscow 119991, Russia — ³Matrosov Institute for System Dynamics and Control Theory, Irkutsk 664033, Russia

The KASCADE Cosmic Ray Data Center (KCDC), introduced in 2013, is a multifunctional public data center for high-energy astroparticle physics. Its distinctive features include use of open standards and technologies, providing materials and service both for professional scientists and a broad outreach audience and furnishing open access to scientific data. The GRADLCI (German-Russian Astroparticle Data Life cycle Initiative), which spawned from KCDC in 2018, proposed an alternative approach to metadata management and utilized the optimized models and algorithms for processing requests. Today, the work on organizing flexible cross-collaboration data sharing is going on in various areas of science within the framework of the EOSC project and others, such as PUNCH4NFDI. A big share of this work includes collection and analysis of the data curation practices in order to reach a more abstract and complex understanding of the challenges of data curation committing into new advanced solutions ready for further extension. In this report the use cases of the KCDC and GRADLCI data centers will be considered.

AKPIK 1.6 Mon 17:30 AKPIK-H13

Optimizing Computer Vision for Radiosource Detection – •JANIS SOWA and KEVIN SCHMIDT — Astroparticle Physics AG Elsässer, TU Dortmund University, Germany

Earthbound radio astronomy utilizes interferometric arrays to achieve the highest possible resolution by combining the measurements of multiple telescopes. The resolution then depends on the distance between telescopes as opposed to the diameter of a single dish. Modern improvements in computing performance and telescope design are allowing radio astronomers to collect increasing amounts of data. In sky surveys, information about hundreds of thousands of astronomical sources are obtained. On this scale, a manual analysis is a timeconsuming task. Deep Learning-based source detection thus naturally comes to mind as a candidate for identifying these individual objects. In a previous work, a Convolutional Neural Network architecture was shown to be faster but less accurate in comparison to the state-of-the-art source detection tool PyBDSF, when tested on simulated data. This talk will showcase how the existing model can be further improved and fine-tuned for application on real data.

AKPIK 1.7 Mon 17:45 AKPIK-H13

Evaluation of deep learning accelerators for the usage in the cosmic ray simulation CORSIKA~8 — •DOMINIK BAACK and JEAN-MARCO ALAMEDDINE for the CORSIKA 8-Collaboration — Astroparticle Physics, WG Elsässer, TU Dortmund University, D-44227 Dortmund, Germany

The proliferation of neural networks has led to the acquisition of specialized hardware to accelerate training and application at an increasing number of scientific sites.

To take advantage of this growth, we investigated the extent to which this hardware can be used to accelerate the complex simulation of cosmic particle showers and which parts of the simulation benefit most.

Location: AKPIK-H13

A number of examples based on CORSIKA~8 are presented to illustrate advantages, disadvantages, and limitations in the choice of methods. In particular, the widely used Nvidia accelerators that was used very successfully for ray tracing of optical photons (e.g. Cherenkov light) will be discussed.

AKPIK 1.8 Mon 18:00 AKPIK-H13

Structured Sparsity for CNNs on Reconfigurable Hardware - •HENDRIK BORRAS, GÜNTHER SCHINDLER, and HOLGER FRÖNING — Institute of Computer Engineering; Heidelberg University; Heidelberg (Germany)

While Convolutional Neural Networks (CNNs) are gaining crucial importance for various applications, including modern analysis and trigger systems, their memory and compute requirements are increasing steadily and the requirements of many CNNs impose serious challenges for achieving high inference throughput and low latency on edge devices, situated close to an experiment. To improve the performance of CNNs on such resource-constrained devices, model compression through quantization and pruning has been proposed and evaluated as a possible solution in the past. Field Programmable Gate Arrays (FP-GAs) are a prime example of low-power devices and suitable for a pervasive deployment. Here, FINN is one of the most widely used frameworks for deploying highly quantized CNN models on edge devices. In this work, we extend FINN for pruning by introducing two methods for column pruning, enabling further compression of CNN-based models. The two techniques vary in their granularity and implementation complexity. The coarse-grain method only prunes blocks of columns, while the fine-grained method is able to prune single columns. Both approaches are then evaluated on the CIFAR10 image classification task. We demonstrate significant throughput improvements of on average

AKPIK 1.9 Mon 18:15 AKPIK-H13

IEA-GAN: Intra-Event Aware GAN for the Fast Simulation of PXD Background at Belle II — •HOSEIN HASHEMI, NIKOLAI HARTMANN, THOMAS KUHR, and MARTIN RITTER — Faculty of Physics, Ludwig Maximilians University of Munich, Germany

The pixel vertex detector (PXD) is the newest and the most sensitive sub-detector at the Belle~II. Data from the PXD and other sensors allow us to reconstruct particle tracks and decay vertices. The effect of background processes on track reconstruction is simulated by adding measured or simulated background hit patterns to the hits produced by simulated signal particles that originate from the processes of interest. This model requires a large set of statistically independent PXD background noise samples to avoid a systematic bias of reconstructed tracks. However, the fine-grained PXD data requires a substantial amount of storage. As an efficient way of producing background information for fast simulation, we introduce the idea of an on-demand PXD background generator with Intra-Event Aware GAN (IEA-GAN), conditioned over the number of PXD sensors in order to produce sensor-dependent PXD images by approximating the concept of an "event" in the detector as these PXD images share both semantic and statistical features that makes it extremely hard for even the State of the Art GANs to mimic these exact properties. As a result, we developed the IEA-GAN model which captures these dependencies by imposing relational inductive bias over the batch dimension.

AKPIK 2: Data Analytics & Machine Learning

Time: Wednesday 16:15-18:30

AKPIK 2.1 Wed 16:15 AKPIK-H13

Interpolation of Instrument Response Functions for the Cherenkov Telescope Array — •RUNE MICHAEL DOMINIK and MAXIMILIAN NÖTHE for the CTA Consortium — Astroparticle Physics WG Elsässer, TU Dortmund University, Germany

In very-high-energy gamma-ray astronomy, the Instrument Response Function (IRF) relates the observed and reconstructed properties to the original properties of the primary particles. The IRFs are usually factored into multiple components, namely the Effective Area, the Energy Dispersion and the Point Spread Function that are needed for the proper reconstruction of spectral and spacial information. These quantities are derived from Monte Carlo Simulations but depend on observation conditions like telescope pointing direction or atmospheric transparency. Producing a complete IRF for every observation taken is a time consuming task and not feasible on the short timescales needed to release e.g. an alert for a transient event. In consequence, IRFs are typically produced at fixed combinations of observation conditions. To derive the optimal IRFs for a given observation, interpolation techniques are investigated. This talk will summarize interpolation strategies that are being tested for the Cherenkov Telescope Array IRFs.

AKPIK 2.2 Wed 16:30 AKPIK-H13

Investigating the Potential Application of Neural Networks for Data Denoising at the Einstein Telescope — •DAVID BERTRAM, MARKUS BACHLECHNER, and ACHIM STAHL — III. Physikalisches Institut B, RWTH Aachen

The Einstein Telescope is a proposed third-generation gravitational wave detector aiming to improve the sensitivity over the whole frequency band compared to the previous generation. For this purpose solely hardware improvements could turn out to be insufficient and novel data processing techniques are crucial. A promising idea for the latter is the implementation of neural networks that can operate on potential irregularly structured additional inputs like seismic sensors. This talk investigates the potential of such techniques in terms of data denoising.

AKPIK 2.3 Wed 16:45 AKPIK-H13

Anomaly detection for Belle II PXD cluster data — •STEPHANIE KÄS, JENS SÖREN LANGE, JOHANNES BILK, and TIMO SCHELLHAAS — Justus-Liebig-Universität Gießen

The Belle II pixeldetector (PXD) has a trigger rate of up to 30 kHz for 8 M pixels. Highly ionizing particles such as antideuterons, pions with small transverse momenta <100 MeV ("slow pions"), magnetic monopoles or stable tetraquarks generate characteristic clusters in the PXD. A large fraction of those does not reach outer detectors and therefore does not generate reconstructable tracks. We study their identification based exclusively on the PXD data, by means of anomaly detection algorithms. In this presentation, we show results from multivariate statistics analysis and tree-based multiclassifiers. In a first step, principal component analysis, linear discriminant analysis, t-distributed stochastic neighbor embedding and random forests are used for each anomaly to investigate the separability of signal and background. In a second step, a multiclassifier system shall be used. The design and output of this approach will be presented, including results on accuracy and sensitivity as well as a comparison to other methods such as Convolutional Neural Networks and Support Vector Machines.

AKPIK 2.4 Wed 17:00 AKPIK-H13

Location: AKPIK-H13

Fast simulation of the HGCAL using generative models — soham bhattacharya¹, samuel bein², engin eren¹, frank gaede¹, gregor kasieczka², •william korcari², dirk kruecker¹, peter mckeown¹, and moritz scham¹ — ¹DESY — ²Universität Hamburg

Accurate simulation of the interaction of particles with the detector materials is of utmost importance for the success of modern particle physics. Software libraries like GEANT4 are tools that already allow the modeling of physical processes inside detectors with high precision. The downside of this method is its computational cost in terms of time. Recent developments in generative machine learning models seem to provide a promising alternative for faster and accurate simulations to accelerate this process. For the challenges of the High Luminosity phase of the LHC, CMS will deploy the High Granularity Calorimeter (HGCal), an imaging calorimeter for the endcap region with a high cell density, and irregular geometry. In this talk, we will show the taken steps in the development of a GraphGAN for the simulation of particle showers in the HGCal and the first achieved results.

AKPIK 2.5 Wed 17:15 AKPIK-H13

Ephemeral Learning - Augmenting Triggers with online-trained normalizing flows — •SASCHA DIEFENBACHER — Institut für Experimentalphysik, Universität Hamburg, Germany

The high collision rates at the Large Hadron Collider (LHC) make it impossible to store every single observed interaction. For this reason, only a small subset that passes so-called triggers - which select potentially interesting events are saved while the remainder is discarded. This makes it difficult to perform searches in regions that are usually ignored by trigger setups, for example at low energies. However a sufficiently efficient data compression method could help these searches by storing information about more events than can be saved offline. We investigate the use of a generative machine learning model (specifically a normalizing flow) for the purpose of this compression. The model is trained to learn the underlying data structure of collisions events in an online setting, meaning we can never have a repeated look at past data. After the training the underlying distribution encoded into the network parameters can be analyzed and, for example, probed for anomalies. We initially demonstrate this method for a simple bump hunt, showing that the online trained flow model can recover sensitivity compared to a classical trigger setup. We then extend this demonstration to more complex examples using the LHC Olympics Anomaly Detection Challenge dataset.

AKPIK 2.6 Wed 17:30 AKPIK-H13

Simulation of High-Granularity Calorimeter Showers for the ILD Using Normalizing Flows — •IMAHN SHEKHZADEH — Universität Hamburg, Hamburg, Deutschland

The large computational cost of Monte Carlo simulations together with recent advances in deep learning motivate using deep generative models to speed up simulations. This talk explores the use of normalizing flows (NFs) for highgranularity calorimeter simulations, such as the ones planned for the International Large Detector (ILD). We show that NFs are able to generate high-fidelity showers of simulated photons in the electromagnetic calorimeter of the ILD. Strictly monotonic rational quadratic spline flows are used to enhance the fidelity in comparison to the generally used affine-linear transformations. Finally, we compare the generative performance of the NFs to other state-of-the-art generative network architectures

AKPIK 2.7 Wed 17:45 AKPIK-H13

Identifying Slow Pions using Support Vector Machines — •Тімо Schellhaas¹, Jens Sören Lange², and Stephanie Käs³ — ¹II. Physikalisches Institut, JLU Gießen, Germany — ²II. Physikalisches Institut, JLU Gießen, Germany — ³II. Physikalisches Institut, JLU Gießen, Germany

Finding new physics beyond the standard model is of highest interest. Pions with a low transversal momentum (slow pions) are linked to interesting decay scenarios and are therefore studied at the Belle II experiment. However, it is dificult to detect slow pions due to their low momenta: a large amount of them does only reach the Belle II pixeldetector (PXD), but not the outer detectors (e.g. the drift chamber). In order to improve the detection rate it is suggested to use a machine learning model. One possible model is the support vector machine (SVM) algorithm. Therefore a simulated data set is used to train the SVM model with different parameters, including a modified kernel, with the goal of reaching better results than other models.

AKPIK 2.8 Wed 18:00 AKPIK-H13 Deep Learning Accelerated Maximum Likelihood Reconstruction of IACT Events — •NOAH BIEDERBECK and MAXIMILIAN NÖTHE FOR THE CTA CON-SORTIUM — Astroparticle Physics WG Elsässer, TU Dortmund University, Germany The Cherenkov Telescope Array will be the next generation ground-based gamma-ray observatory, consisting of tens of Imaging Atmospheric Cherenkov Telescopes (IACTs) at two sites once its construction is finished.

In this talk we present a deep learning accelerated maximum likelihood reconstruction of gamma-ray events. A generative neural network predicts IACT camera images from a set of physical event parameters. These generated images are then compared to Monte Carlo simulated event images using a Poissonian likelihood loss in order to reconstruct the event properties, e.g. the energy of the primary particle and its direction.

First results on simulated single-telescope events will be presented and extensions to predictions of array events will be outlined.

AKPIK 2.9 Wed 18:15 AKPIK-H13 Adding Errors to the Quantum Circuit Model — •Tom Weber¹, Matthias RIEBISCH¹, KERSTIN BORRAS^{2,4}, KARL JANSEN³, and DIRK KRÜCKER² — ¹Universität Hamburg, Hamburg, Germany — ²Deutsches Elektronen-Synchrotron DESY, Hamburg, Germany — ³Deutsches Elektronen-Synchrotron DESY, Zeuthen, Germany — ⁴RWTH Aachen University, Aachen, Germany The full potential of quantum computers cannot yet be realised because existing quantum hardware is still error-prone. It is essential to understand the impact of these errors on calculations to counteract them with methods like quantum error mitigation. Models can provide this understanding of the complexity of quantum noise. In addition, they can be a tool for communication between different quantum computing stakeholders who do not necessarily have an education in physics. While the quantum circuit model is commonly used to model gate-based quantum computation, errors are modelled mathematically by quantum operations on density operators. However, the quantum circuit model is restricted to the description of error-free processes. On the other hand, mathematical models are difficult to understand without a background in theoretical physics. Therefore, we present a way to couple both models, combining the comprehensibility of the quantum circuit model with the mathematical models' ability to represent quantum noise accurately.

AKPIK 3: Mitgliederversammlung AKPIK

Time: Wednesday 19:00–21:00 AKPIK Mitgliederversammlung.

AKPIK 4: Deep Learning

Time: Thursday 16:15-18:30

AKPIK 4.1 Thu 16:15 AKPIK-H13 Using Graph Neural Networks for improving Cosmic-Ray Composition

Analysis at IceCube Observatory — •PARAS KOUNDAL for the IceCube-Collaboration — Institute for Astroparticle Physics, KIT Karlsruhe, Germany Graph Neural Networks (GNNs) is one of the most emerging and promising research topics in the field of deep-learning. Described using nodes and edges, graphs allow us to efficiently represent relational data and learn hidden representations of input data to obtain better model-prediction accuracy. The success of GNNs is mainly attributed to their unique ability to represent complex input data in its most natural representation. GNNs have hence accelerated and extended the pattern learning, inference drawing of standard deep-learning architectures. This has also made it possible for faster and more precise analysis in astroparticle physics, enabling new insights from massive volumes of input data. IceCube Neutrino Observatory, a multi-component detector concealed deep under the South Pole ice provides a suitable test-case to implement such methods.

The talk will discuss the GNN-based methods for improving cosmic-ray composition understanding in the transition region from Galactic to extragalactic sources, at IceCube Observatory. The implementation benefits by using full signal-footprint information, in addition to reconstructed cosmic-ray air shower parameters. The talk will also explain improvement to individual GNN based model by ensemble methods. The implementation will reduce the time and computing cost for performing cosmic-ray composition analysis while boosting sensitivity.

AKPIK 4.2 Thu 16:30 AKPIK-H13

Amplifying Calorimeter Simulations with Deep Neural Networks — •SEBASTIAN GUIDO BIERINGER¹, ANJA BUTTER², SASCHA DIEFENBACHER¹, EN-GIN EREN³, FRANK GAEDE³, DANIEL HUNDSHAUSEN¹, GREGOR KASIECZKA¹, BENJAMIN NACHMAN⁴, TILMAN PLEHN², and MATHIAS TRABS⁵ — ¹Institut

Location: AKPIK-H13

Location: AKPIK-MV

für Experimentalphysik, Universität Hamburg, Germany — ²Institut für Theoretische Physik, Universität Heidelberg, Germany — ³Deutsches Elektronen-Synchrotron, Hamburg, Germany — ⁴Physics Division, Lawrence Berkeley National Laboratory, Berkeley, CA, USA — ⁵Fachbereich Mathematik, Universität Hamburg, Germany

Speeding up detector simulation over the computationally expensive Monte Carlo tools is a key effort for upcoming studies at the LHC and future colliders. Machine-learned generative surrogate models show great potential to accelerate such and other simulations. However, estimating the relation between the statistics of the training data and the generated distribution of the model is essential to determine the gains and use-cases of these methods.

We present a detailed study of this relation for the concrete physics example of photon showers in a highly granular calorimeter. For established metrics on calorimeter images, the amplification properties of a VAE-GAN model are examined in terms of an approximation to the Jenson-Shannon-divergence between generated data, training data and a high-statistics batch.

AKPIK 4.3 Thu 16:45 AKPIK-H13

Deep Learning-based Imaging in Radio Interferometry — •FELIX GEYER and KEVIN SCHMIDT — Astroparticle Physics AG Elsässer, TU Dortmund University, Germany

Radio interferometry is used to monitor and observe distant astronomical sources and objects with high resolution. Especially Very Long Baseline Interferometry allows achieving the highest resolutions by combining the data of multiple telescopes. This results in an effective diameter corresponding to the greatest distance between two telescopes. The taken data consists of visibilities, which depend on the baselines between the telescopes. Because the distribution of these baselines is sparse, the sample of visibilities is incomplete. After transforming this sample to spatial space, this so-called "dirty image" is inadequate

for physical analyses. Thus, the image undergoes an elongated and mostly manually performed cleaning process in order to remove background artifacts and to restore the original source distribution. We developed a new and fast approach to reconstruct missing data reasonably using Neural Networks. This talk will present the current state of the radionets framework and an outlook on upcoming projects. One focus will be on the simulation improvements using the RIME formalism.

AKPIK 4.4 Thu 17:00 AKPIK-H13

Binary Black Hole Parameter Reconstruction using Deep Neural Networks — •MARKUS BACHLECHNER, DAVID BERTRAM, and ACHIM STAHL — III. Physikalisches Institut B, RWTH Aachen

The proposed Einstein Telescope, as the first of the third generation of gravitational wave detectors, is expected to be an order of magnitude more sensitive compared to current interferometers like LIGO or Virgo. On the one hand the higher sensitivity increases the observable volume. On other hand the frequency range is broadened, which in return bares the potential to extend the observable time of binary coalescences from seconds to hours. These long observable times make it possible to send multi-messenger alerts before the end of the coalescences. For this it is essential to apply a fast real-time analysis handling event detection, classification, and reconstruction. In this talk an approach for the parameter reconstruction of binary black holes using deep neural networks is presented.

AKPIK 4.5 Thu 17:15 AKPIK-H13

A Recurrent Neural Network for Radio Imaging — •STEFAN FRÖSE and KEVIN SCHMIDT — Astroparticle Physics WG Elsässer, TU Dortmund University, Germany

In radio astronomy, an array of correlated antennas, called a radio interferometer, is used to produce high-resolution images of the sky. The measurements take place in the complex Fourier space due to the pairwise correlation of antennas. Therefore, the amount of information to receive from such an array is restricted by the number of antennas. The resulting spatial dirty map of these measurements will be cleaned using a Neural Network. The architecture for this network is based on a Recurrent Neural Network (RNN). RNNs can be used to extract information from sequential data, like text or speech. In the context of inverse problems the RNN can be derived directly from a *maximum a posteriori* approach. Furthermore the iterative behaviour of the network can be exploited to construct a CLEAN-like network to reconstruct a map of the sky. This results in the so-called RIM architecture published by Patrick Putzky & Max Welling (arxiv:1706.04008). The Neural Network is able to clean given dirty maps for simulated radio images and also shows convergence for the EHT dataset of M87.

AKPIK 4.6 Thu 17:30 AKPIK-H13

Measurement of the Mass Composition using the Surface Detector of the Pierre Auger Observatory and Deep Learning — MARTIN ERDMANN, •JONAS GLOMBITZA, and NIKLAS LANGNER for the Pierre Auger-Collaboration — III. Physics Institute A, RWTH Aachen

Measuring the mass composition of ultra-high energy cosmic rays (UHECRs) constitutes one of the biggest challenges in astroparticle physics. Nowadays, the most precise measurements can be obtained from measurements of the depth of maximum of air showers, $X_{\rm max}$, with the use of Fluorescence Detectors (FD), which can be operated only during clear and moonless nights.

With the advent of deep learning, it is now possible for the first time to perform an event-by-event reconstruction of X_{max} using the Surface Detector (SD) of the Pierre Auger Observatory. Therefore, previously recorded data can be analyzed for information on X_{max} , and thus the cosmic-ray composition. Since the SD features a duty cycle of nearly 100%, the gain in statistics is a factor of 15 for energies above $10^{19.5}$ eV compared to the FD.

This contribution introduces the neural network specifically designed for the SD of the Pierre Auger Observatory. We evaluate its performance using three different hadronic interaction models and verify its functionality using Auger hybrid measurements. Finally, we quantify the expected systematic uncertainties and determine the UHECR mass composition using the first two moments of the X_{max} distributions up to the highest energies.

AKPIK 4.7 Thu 17:45 AKPIK-H13 Graph Neural Networks for Low Energy Neutrino Reconstruction at IceCube — •RASMUS ØRSØE — Trøjborggade 6, 3.sal, 1757 Copenhagen

A presentation on the application of graph neural networks for low energy neutrino reconstruction of IceCube events. Comparisons with current methods will be shown. Brief introduction to graph neural networks and motivation is included.

 $AKPIK 4.8 \quad Thu \ 18:00 \quad AKPIK-H13$ Event-by-event estimation of high-level observables with data taken by the Surface Detector of the Pierre Auger Observatory using deep neural networks — •Steffen Hahn¹, Markus Roth¹, Darko Veberic¹, David Schmidt¹, Ralph Engel¹, and Brian Wundheiler² — ¹KIT, IAP, Germany — ²UNSAM, ITEDA, Argentina

Probing physics beyond the scales of human-made accelerators with cosmic rays requires accurate estimation of high-level observables, such as the energy of the primary particle or the maximum of the shower depth. Measurements of the shower cascade, however, consist mainly of various, hard-to-interpret time signals which potentially contain non-trivial correlations. Deep neural networks are a convenient way to tackle such a problem in a general way.

The shower footprint measured by the surface detector of the Pierre Auger Observatory provides us with time slices of the ground signal of a shower cascade. This gives us an ideal test bed to determine the quality of network based reconstruction methods compared to that of regular analysis methods. However, a caveat of this approach is that the networks must be trained on Monte-Carlo simulations. Since present hadronic interaction models for energies beyond 10 EeV are extrapolations there are discrepancies between simulations and real data for which we have to correct for.

Here, we present a multi-purpose architecture and correction-method to predict high-level observables on measured data as well as physics results.

AKPIK 4.9 Thu 18:15 AKPIK-H13

Reconstruction of primary particle energy from data taken by the Surface Detector of the Pierre Auger Observatory using deep neural networks — RALF ENGEL, MARKUS ROTH, DARKO VEBERIC, DAVID SCHMIDT, STEFFEN HAHN, and •FIONA ELLWANGER for the Pierre Auger-Collaboration — Karlsruhe Institute of Technology (IAP), Karlsruhe, Germany

To probe physics beyond the scales of human-made accelerators with cosmic rays demands an accurate knowledge of their energy. Indirect, ground-based experiments reconstruct this primary particle energy from measurements of the emitted fluorescence light or the time-signal of the shower footprint. Using fluorescence detectors, one is able to estimate former with good accuracy. These, however, exhibit a rather low duty cycle.

At the Pierre Auger Observatory the shower footprint is measured by a regular triangular grid of water-Cherenkov detectors. Since the shower development is a very intricate process the time signals of the detectors are fairly complex. Additionally, the sheer amount of data makes it non-trivial to find hidden patterns in their spatial and temporal distributions. Neural networks provide a straightforward way of tackling such a problem doing a data-driven analysis.

With large simulation data sets we are able to train more complex networks. Systematic differences between simulations and measured data require special attention to possible biases, which are quantified. In this work, we present a neural network architecture that gives an estimate on the energy for real data.

Working Group on Philosophy of Physics Arbeitsgruppe Philosophie der Physik (AGPhil)

Dennis Lehmkuhl Institut für Philosophie Rheinische Friedrich-Wilhelms-Universität Bonn 53113 Bonn dennis.lehmkuhl@uni-bonn.de Radin Dardashti Interdisciplinary Center for Science and Technology Studies Bergische Universität Wuppertal 42119 Wuppertal dardashti@uni-wuppertal

Overview of Invited Talks and Sessions

(Lecture hall AGPhil-H14)

Invited Talks

AGPhil 2.1	Mon	16:15-17:00	AGPhil-H14	To G or not to G: J. H. Poynting and the gravitational constant in the 19th century — •ISOBEL FALCONER
AGPhil 4.1	Tue	14:00-14:45	AGPhil-H14	Hypothetical Waveforms and Unmodeled or Pipeline Searches in Gravita- tional Wave Astronomy ALYDIA DATTON
AGPhil 5.1	Tue	16:15-17:00	AGPhil-H14	Portrait of a Black Hole: Objectivity and the Imaging of M87* by the Event
AGPhil 5.2	Tue	17:00-17:45	AGPhil-H14	Horizon Telescope — •Peter Galison When is a black hole spacetime "as large as it can be"? — •Juliusz Do-
AGPhil 6.1	Wed	14:00-14:45	AGPhil-H14	BOSZEWSKI Spacetime Conventionalism Revised: Tidal Forces and Weyl Curvature —
AGPhil 7.1	Wed	16:15-17:00	AGPhil-H14	•Karim Thébault, Ufuk Tasdan On an inferential role of spacetime in particle physics — •Tushar Menon

Sessions

Mon	11:00-13:00	AGPhil-H14	Symmetry and Geometry
Mon	16:15-18:00	AGPhil-H14	History and Philosophy of Gravity
Tue	11:30-13:00	AGPhil-H14	Black Holes I
Tue	14:00-15:45	AGPhil-H14	Gravitational and Electromagnetic Waves
Tue	16:15-18:15	AGPhil-H14	Black Holes II
Wed	14:00-16:15	AGPhil-H14	Foundations of Gravity
Wed	16:15-18:30	AGPhil-H14	Symmetries and Principles
Wed	18:30-19:30	AGPhil-MV	Annual Meeting of the AGPhil
Thu	11:00-13:00	AGPhil-H14	Quantum Mechanics I
Thu	14:00-16:00	AGPhil-H14	Quantum Mechanics II
Thu	16:15-18:45	AGPhil-H14	Time and Temperature
Fri	11:00-13:00	AGPhil-H14	Processes, Events and Time
	Mon Mon Tue Tue Wed Wed Thu Thu Thu Fri	Mon11:00-13:00Mon16:15-18:00Tue11:30-13:00Tue14:00-15:45Tue16:15-18:15Wed14:00-16:15Wed16:15-18:30Wed18:30-19:30Thu11:00-13:00Thu14:00-16:00Thu16:15-18:45Fri11:00-13:00	Mon11:00-13:00AGPhil-H14Mon16:15-18:00AGPhil-H14Tue11:30-13:00AGPhil-H14Tue14:00-15:45AGPhil-H14Tue16:15-18:15AGPhil-H14Wed14:00-16:15AGPhil-H14Wed16:15-18:30AGPhil-H14Wed18:30-19:30AGPhil-H14Thu11:00-13:00AGPhil-H14Thu16:15-18:45AGPhil-H14Fri11:00-13:00AGPhil-H14

Annual General Meeting of the Working Group on Philosophy of Physics

Wednesday 18:30-19:30 AGPhil-MV

- Report
- Plans for 2022/23
- Miscellaneous

Location: AGPhil-H14

Sessions

- Invited and Contributed Talks -

AGPhil 1: Symmetry and Geometry

Time: Monday 11:00-13:00

AGPhil 1.1 Mon 11:00 AGPhil-H14

A Proposal for a Metaphysics of Self-Subsisting Structures — •ANTONIO VASSALLO¹ and PEDRO NARANJO^{1,2} — ¹Faculty of Administration and Social Sciences, Warsaw University of Technology, Plac Politechniki 1, 00-661 Warsaw, Poland — ²Faculty of Philosophy, University of Warsaw, Krakowskie Przedmieście 3, 00-047 Warsaw, Poland

We present a new metaphysical framework for physics that is conceptually clear, ontologically parsimonious, and empirically adequate. This framework relies on the notion of self-subsisting structure, that is, a set of fundamental physical elements whose individuation and behavior are described in purely relational terms, without any need for a background spacetime. Although the specification of the fundamental elements of the ontology depends on the particular physical domain considered – and is thus susceptible to scientific progress–, the structural features of the framework are preserved through theory change. The kinematics and dynamics of these self-subsisting structures are technically implemented using the theoretical framework of Pure Shape Dynamics, which provides a completely relational physical description of a system in terms of the intrinsic geometry of a suitably defined Riemannian space, called shape space.

AGPhil 1.2 Mon 11:30 AGPhil-H14

Arguments from scientific practice in the debate about the physical equivalence of symmetry-related models — •JOANNA LUC — Jagiellonian University, Kraków, Poland

In the recent philosophical literature, several counterexamples to the interpretative principle that symmetry-related models are physically equivalent have been suggested (Belot 2013, Belot 2018, Fletcher 2020). Arguments based on these counterexamples can be understood as arguments from scientific practice of roughly the following form: because in scientific practice such-and-such symmetry-related models are treated as representing distinct physical situations, these models indeed represent distinct physical situations. I will argue that if we are exclusively interested in models understood as representing entire possible worlds (not their subsystems), arguments from scientific practice should involve some additional assumptions to guarantee that they are relevant for models understood in this way. However, none of the examples presented in the literature satisfy all these additional assumptions, which leads to the conclusion that arguments from scientific practice based on these examples do not undermine the interpretative principle that different symmetry-related models represent the same possible world. An important ingredient of my argumentation is the distinction between implicit and explicit modes of representing in physics; symmetry-related models understood as representing subsystems are in some contexts physically inequivalent only because they represent implicitly some physical object (associated with a reference frame).

AGPhil 1.3 Mon 12:00 AGPhil-H14 A new view of the history of electromagnetic theory. An alternative formulation to Maxwell — •ANTONINO DRAGO — via Benvenuti 3, 56011 Calci, Italy The exceptional role played by electromagnetic theory within the history of classical physics is stressed and characterized. The notion of incommensurability between different approaches explains why this case-study constitutes a hard subject for the historians of physics and hence why in the past they were content to consider as the decisive event of this history the birth of Maxwell equations. The usual historical account on the completion of electromagnetic theory is contested; electromagnetic theory has to be considered a completed theory not before the requirement of a symmetric explanation of electromagnetic induction between moving bodies was fulfilled. Actually in a retrospective view from the introduction of Lorentz*s group some scholars have suggested new foundations of electromagnetism. Among these new foundations I recognize in a recent one (Diener et al. 2013) a substantial anticipation of an alternative formulation to Maxwell-Hertz-Lorentz one. I improve it in a formal way according to an interpretation of the foundations of the electromagnetism as constituted by the choices on two basic dichotomies: one about two kinds of mathematics, and another about two kinds of logic.

AGPhil 1.4 Mon 12:30 AGPhil-H14 The Spatially of the Universe in Einstein*s paper *Geometry and Experience* — •TAIMARA PASSERO — University of São Paulo, São Paulo, Brazil.

The aim of this talk is to present and discuss the role of Euclidean geometry in Einstein*s argument concerning the spatially of the Universe. Albert Einstein analyzes this topic in the paper *Geometry and Experience*, given as a public address on January 27, 1921 at the Prussian Academy of Sciences. In the first part of his paper, Einstein distinguishes between *purely axiomatic geometry* and *practical geometry*. In the second part, Einstein discusses whether the Universe is spatially finite or not. He presents a beautiful argument to illustrate the theory of a finite Universe by means of a mental picture using his notion of practical geometry. To obtain this, Einstein goes from the thinking and visualization offered by Euclidean geometry to acquire a mental picture of the spherical geometry. This process leads him to conclude that *the human faculty of visualization is by no means bound to capitulate to non-Euclidean geometry*.

AGPhil 2: History and Philosophy of Gravity

Time: Monday 16:15-18:00

Invited TalkAGPhil 2.1Mon 16:15AGPhil-H14To G or not to G: J. H. Poynting and the gravitational constant in the 19th

century — •ISOBEL FALCONER — University of St Andrews, UK The increasing precision of gravitational measurement is sometimes given as a reason for the acceptance of the gravitational constant, G, in the late 19th century. However, as late as the 1890s, John Henry Poynting, the doyen of British workers on gravitation, persistently refused to cast measurement of the gravitational

ers on gravitation, persistently refused to cast measurement of the gravitational constant as his experimental aim; he preferred to present it as measurement of the mean density of the Earth. Despite his detailed analysis, in his Adams Prize Essay of 1894, of the improvements in experimental method that were enabling ever more precise measurement, he similarly interpreted all previous measurements as of the mean density of the earth. His reservations about G alert us to the mathematical, physical and metaphysical interpretative work involved in the shift that had occurred during the previous 100 years, from expressing the laws of physics as ratio equations to expressing them as functional relationships between algebraic symbols that denoted the numerical values of physical quantities.

This talk will encompass gravitational work in Britain, France, and Germany, in exploring the introduction of G into physics and some of the questions raised by Poynting's reservations about G as a useful physical construct.

Location: AGPhil-H14

AGPhil 2.2 Mon 17:00 AGPhil-H14

The Renaissance of General Relativity in the 1960s — •DENNIS LEHMKUHL — Lichtenberg Group for History and Philosophy of Physics, Institute of Philosophy, University of Bonn

This talk will focus on the development of new mathematical methods during the 1960s that allowed for new ways of understanding the solution space of the Einstein equations, and subsequently for new avenues to work on cosmology. The focus will be on the classification schemes for vacuum solutions developed by Petrov, Penrose, and Pirani, as well as the global methods developed during the work on the singularity theorems by Penrose and Hawking. Building on this, the talk will outline how both the singularity theorems themselves and the new methods developed in proving them have influenced subsequent work on cosmology.

AGPhil 2.3 Mon 17:30 AGPhil-H14 Holistic Eliminative Reasoning for Astronomy and Astrophysics — •SHANNON SYLVIE ABELSON — Indiana University Bloomington, IN, United States

I argue that a promising epistemology for astronomy and astrophysics (A&A) involves a certain kind of eliminative reasoning. Unlike the traditional conceptions of such reasoning that propose to eliminate rival theories or models based upon quality of evidence, I build upon work by Paul Horwich (1982), Patrick Forber (2011), and Elisabeth Lloyd (2013; 2015) to argue that it is particular model as-

Working Group on Philosophy of Physics (AGPhil)

sumptions (variables, parameters, etc.) that are weighed and eliminated. Rather than a veridical comparison between theory predictions and individual observational results, holistic eliminative reasoning has a web-like structure. Elimination is the end result of a multi-step reasoning program that holistically evaluates the introduction of a proposed assumption into the state space of previously accepted evidence. In particular and where possible, model assumptions should cohere with our well-confirmed pictures of dynamical processes and the mechanisms that underlie them. Holistic elimination then becomes a project of capturing dynamical accuracy. These ideas have been explored in the context of biology and genetics (see Lloyd, Lewontin, and Feldman (2008), Forber (2011) and Ratti (2015)), but have not been extended to A&A. I outline how this epistemic framework can be applied to competing dynamical pictures of the mechanisms and conditions underlying the evolutionary histories of black holes, including gas accretion, intermediate mass black hole mergers, and direct-collapse black hole models.

AGPhil 3: Black Holes I

Time: Tuesday 11:30-13:00

AGPhil 3.1 Tue 11:30 AGPhil-H14

No membrane at the black hole horizon? — •MARCO SANCHIONI — Via Timoteo Viti 10, Department of Pure and Applied Sciences, University of Urbino, Italy

Since the discovery of Hawking radiation (Hawking, 1976), it has been ac- accepted among physicists, and later on also by philosophers (Wallace, 2018), that black holes are thermodynamic objects in the total sense. To have a statistical mechanical underpinning of black hole thermodynamics, as is the case for thermodynamics of ordinary things, it has been argued that a quantum membrane should be posited at the black hole horizon. This paper is an inquiry on the status of the quantum membrane paradigm in light of recent theoretical results on black hole physics obtained within the research program of semiclassical gravity (Penington et al., 2019; Almheiri et al., 2020) and ultimately grounded on the ER=EPR proposal (Maldacena and Susskind, 2013). However, we do not discuss the problematic aspect of such a research program, which would be a project on its own, and our result is thus conditional to its validity. In particular, the paper starts an investigation on the picture of black holes that underlies these new calculations. The main result of this paper is that, within the central assumption on the validity of semiclassical gravity, the quantum membrane paradigm should be abandoned.

AGPhil 3.2 Tue 12:00 AGPhil-H14

Stellar gravitational collapse, singularity formation and theory breakdown – •KIRIL MALTSEV – Heidelberg Institute for Theoretical Studies / University of Heidelberg

A critical examination of the main physical arguments against the prediction of gravitational singularity formation in stellar core collapse is given, restricted in

Location: AGPhil-H14 scope to a historically oriented survey of the decades spanning in between the Schwarzschild 1916 solution and the Penrose 1965 theorem. We first review the 3 definitions (missing point(s), infinite curvature, and geodesic incompleteness) of what a singularity is, and argue that its prediction is problematic for GR, indicating breakdown of Lorentzian geometry, only insofar as infinite curvature is concerned. In contrast, geodesic incompleteness is its innovating hallmark, which is not meaningfully available in Newtonian gravity formulations (infinite density, and infinite gravitational force) of what a gravitational singularity is. The Oppenheimer-Snyder 1939 solution derives the formation of locally infinite curvature and of incomplete geodesics, while Penrose's 1965 theorem concerns the formation of incomplete (null) geodesics only. We assess as the most robust curvature pathology formation counter-argument Markov's derivation of an upper bound on the quadratic curvature invariant from a ratio of natural constants, in connection with Wheeler's conjecture that the Planck scale is ultimate. Finally, we recall Landau's objection to fermionic infinite density point mass formation, which still provides strong reasons to believe that by the least an intermediate state towards the final fate of gravitational collapse must be a bosonic configuration. AGPhil 3.3 Tue 12:30 AGPhil-H14

Alice meets Bob! or: The association of infinity and finiteness within the Schwarzschild metric — •René FRIEDRICH — Strasbourg

The Schwarzschild metric is the basic description of a gravitational field, but it is more than that: It provides us with some hints about the way how the universe is working. One main feature of the Schwarzschild metric is the association of finite and infinite time structures, and it includes even proposals for the solution of the so-called "information paradox" of black holes and the supposed "breakdown of general relativity" near singularities.

AGPhil 4: Gravitational and Electromagnetic Waves

Time: Tuesday 14:00-15:45

Invited TalkAGPhil 4.1Tue 14:00AGPhil-H14Hypothetical Waveforms and Unmodeled or Pipeline Searches in Gravita-
tional Wave Astronomy — •LYDIA PATTON — Virginia Tech, Blacksburg, Vir-
ginia, USA

The multiple theoretical and instrumental advances in gravitational wave astronomy (GWA) have allowed for the construction of an increasingly flexible platform for discovery. This paper will investigate novel research methods being constructed on the ground in GWA. It will begin by evaluating the comparison of and contrast between two methods of analysis: the construction of waveforms that incorporate hypothetical parameters for new searches (EOB and novel extended EOB methods), and the use of unmodeled and pipeline-based searches of existing data. In both cases, varying hypothetical assumptions or models allows for more flexible, broad search methods. The question then is how to move from the broader a priori models to the detection of an event. We will examine several recent papers to reconstruct how these broader methods can be used to support novel detection, and examining how search and detection methods work together in this context.

AGPhil 4.2 Tue 14:45 AGPhil-H14

What Gravitational Waves Really Teach Us about Energy — •SAMUEL FLETCHER — University of Minnesota, Twin Cities, Minneapolis, USA

Gravitational wave solutions to the Einstein field equation of general relativity are commonly regarded as examples proving how gravity in general relativity transmits energy from a source body to a distant body. The famous 1955 Feynman sticky bead thought experiment illustrates the reality of this phenomenon by imagining two beads generating heat in a rod on which they slide with friction, due to their changing proper distance in the presence of the waves. I argue that while this lesson is not entirely wrong, it is much too simplistic. It does not reconLocation: AGPhil-H14

cile its conclusion with the fact that conservation of local energy-momentum, in the sense that appears in the field equation, prevents energy transmission across a vacuum. Thus "energy transmission" must employ a different concept of energy, raising the possibility of pluralism with regard to the energy concept. Another (compatible) possibility is that gravitational waves, rather than transmitting energy, facilitate the transformation between different types or stores of energy locally. Key to these possibilities is analysis of the Weyl tensor. Time permitting, I discuss these possibilities' implications for a re-evaluation of the scope of Mach's Principle, the idea that the distribution of matter determines the geometry of spacetime.

AGPhil 4.3 Tue 15:15 AGPhil-H14 Absorbing the Arrow of Electromagnetic Radiation - MARIO HUBERT and •CHARLES SEBENS — California Institute of Technology, Pasadena, CA, USA We argue that the asymmetry between diverging and converging electromagnetic waves is just one of many asymmetries in observed phenomena that can be explained by a past hypothesis and statistical postulate (together assigning probabilities to different states of matter and field in the early universe). The arrow of electromagnetic radiation is thus absorbed into a broader account of temporal asymmetries in nature. We give an accessible introduction to the problem of explaining the arrow of radiation and compare our preferred strategy for explaining the arrow to three alternatives: (i) modifying the laws of electromagnetism by adding a radiation condition requiring that electromagnetic fields always be attributable to past sources, (ii) removing electromagnetic fields and having particles interact directly with one another through retarded action-at-a-distance, (iii) adopting the Wheeler-Feynman approach and having particles interact directly through half-retarded half-advanced action-at-a-distance. In addition to the asymmetry between diverging and converging waves, we also consider the related asymmetry of radiation reaction.

AGPhil 5: Black Holes II

Time: Tuesday 16:15-18:15

Invited TalkAGPhil 5.1Tue 16:15AGPhil-H14Portrait of a Black Hole: Objectivity and the Imaging of M87* by the EventHorizon Telescope• PETER GALISONBlack Hole Initiative, Harvard University, Cambridge, Mass. United States

In thousands of atlases depicting the working objects of scientific inquiry-from skeletons, clouds, and plants, to crystals, elementary particles, and stars, physicians and scientists across many domains worked out what counted as scientific objectivity. This long-term history, with its various takes on what a reliable image should be, converged in the yearslong struggle of the Event Horizon Telescope (EHT) to produce a picture of a black hole robust enough to make public. As a member of the imaging group, I was part of this effort–offering an occasion for the direct interaction of philosophy and physics as we in the collaboration thought through the different forms of images in consideration: ideal images, mechanically objective images, and expert judgment images. On April 10, 2019, the team released the first image of a black hole, an image viewed within a very few days by more than a billion people. This is a talk about how the EHT team of some 200 scientists came to assess as objective the glowing, crescent-like ring around the supermassive black hole M87*.

Invited TalkAGPhil 5.2Tue 17:00AGPhil-H14When is a black hole spacetime "as large as it can be"?- JULIUSZ DO-BOSZEWSKI — University of Bonn (Lichtenberg Group for History and Philoso-
phy of Physics) — Black Hole Initiative, Harvard University

AGPhil 6: Foundations of Gravity

Time: Wednesday 14:00-16:15

Invited Talk AGPhil 6.1 Wed 14:00 AGPhil-H14 Spacetime Conventionalism Revised: Tidal Forces and Weyl Curvature — •KARIM THÉBAULT and UFUK TASDAN — University of Bristol

Our goal in this paper is to better understand the physical interpretation of tidal forces and Weyl curvature in general relativity by considering novel articulations of thesis of 'spacetime conventionality'. We will first consider a specific rendition of the conventionality thesis in the context of the debates regarding the status of energy conservation and the effects of tidal forces. This will then, in turn, motivate a discussion of the two most physically important forms of curvature - Ricci and Weyl - which can be isolated in general relativity, focusing upon the extent to which such formal distinction may be employed to articulate an entirely non-conventional analysis of the causal origin of tidal forces. We next consider the idea that the Ricci vs. Weyl curvature distinction can be further deployed to anchor a conventionalism-proof distinction between 'pure geometric' Weyl curvature and 'matter-energy-coupled' Ricci curvature. To foreshadow our main conclusion, what we find is that the complex of couplings between Ricci curvature and stress-energy, via the Einstein equation, and Weyl and Ricci curvature, via the Bianchi identity, leads us away from such attractively clean distinctions. Finally, we will outline some open questions and possible lines of future work as an envoi.

AGPhil 6.2 Wed 14:45 AGPhil-H14

Perturbing the hole argument — •JOHN DOUGHERTY — Munich Center for Mathematical Philosophy, LMU Munich

The recent literature on the hole argument has seen a reappraisal of its mathematical aspects. According to this reappraisal, as Halvorson and Manchak succinctly put it, there are two mathematical claims that might be thought to underwrite the hole argument, and neither in fact does. The claim that there are isomorphic but distinct Lorentzian manifolds is trivial, and the claim that there is a diffeomorphism that spoils the determinism is false. In this paper I argue that at least one version of the hole argument is underwritten by a third mathematical claim: that the configuration space of general relativity is "natural", which is to say that it depends functorially on the base manifold. This claim is nontrivial in the sense that it is not true in many theories, such as those containing spinor fields. But it is true in a tensorial theory like general relativity. And it underwrites the version of the hole argument that analogizes general covariance to the "gauge" nature of general relativity as it is used in perturbative contexts such as calculations concerning gravitational radiation and semiclassical effects.

AGPhil 6.3 Wed 15:15 AGPhil-H14

A Case for Further Inquiry into Spin and Gravity — \bullet ZACHARY HALL — Stanford University

Location: AGPhil-H14

Multiple conditions have been proposed in the literature aiming at capturing the idea that a general relativistic spacetime is "as large as it can be". I will consider some of them in the context of particular black hole spacetimes, including standard solutions, regular black holes, and fully evaporating black holes. The emerging landscape is not just subtle but also surprising. Interesting connections arise between these issues and certain versions of the cosmic censorship conjecture. Philosophical consequences involve a notion of a time machine and impact the viability of metaphysical principles such as the principle of sufficient reason.

AGPhil 5.3 Tue 17:45 AGPhil-H14 **A Role for the 'Fauxrizon' in the Semiclassical Limit of a Fuzzball** — •MIKE D. SCHNEIDER — University of Illinois at Chicago, Chicago IL, USA

Recent remarks by Huggett and Matsubara ("Lost Horizon? - Modeling Black Holes in String Theory", 2021) indicate that a 'fauxrizon' (portmanteau of 'faux horizon'), such as is relevant to understanding astrophysical black holes according to the fuzzball proposal within string theory (and perhaps in firewall proposals, more generally), might ultimately solve the familiar black hole evaporation paradox. I clarify, with general upshots for quantum gravity research, some of what this suggestion would amount to: namely, identification of intertheoretic constraints on global spacetime structure in semiclassical models of fuzzballs.

Location: AGPhil-H14

I present an undiscussed instance of the tension between the backgrounddependent formalism of quantum theory and the background-independence of classical general relativity. Notably, the issue is subject to empirical testing, for which reason it also holds interest for those who eschew backgroundindependent methods or interpretations in gravitational theory. The issue is that the representations of spin-states in quantum theory depend prima facie on an embedding of those states in a flat background geometry. This raises the question of whether we should continue using a background geometry in representing spin-states in a world with gravitation. The empirical questions are apparent with knowledge of how experimentalists align (a) preparing and measuring devices of spin-states undergoing no non-gravitationally induced precession and (b) measuring devices in multiple wings of experiments on spin-entangled states. The aligning procedure is operational, meaning that the question of how the aligned measurement axes should be represented in the spacetime has been so-far uninvestigated. While some may be inclined to think that they should be represented with the Christoffel symbols and path information of the system, it is not clear that this is the only acceptable solution a priori.

It is often said that Hawking radiation just is a kind of Unruh radiation. In this work, we clarify the ways in which Hawking radiation can and cannot be seen as a kind of Unruh radiation. Hawking radiation is analogous to Unruh radiation in that the Schwarzschild metric near the horizon is isomorphic to the Rindler metric, which allows us to employ the derivation of Unruh radiation to obtain Hawking radiation. But the isomorphism is restricted to the near-horizon region. This observation leads to the way in which Hawking radiation is not a kind of Unruh radiation: the analogy between them is not due to the equivalence principle. One might think that because observers near - but outside of the horizon of a black hole are equivalent, via the equivalence principle, to an accelerating observer in empty space, Hawking radiation observed by a hovering observer outside a black hole just is the kind of Unruh radiation that an accelerating observer in empty space would see. We argue that this is an incorrect way of thinking of Hawking radiation. Indeed, this would imply that hovering observers outside gravitating bodies that are not black holes - such as stars and planets - would also observe Unruh/Hawking radiation, and this is not the case. Throughout we emphasize the ways in which Hawking and Unruh radiation can be seen as varieties of geometric radiation, i.e., radiation generated by the structure of a metric containing horizons.

Tuesday

AGPhil 7: Symmetries and Principles

Time: Wednesday 16:15-18:30

Invited TalkAGPhil 7.1Wed 16:15AGPhil-H14On an inferential role of spacetime in particle physics — •TUSHAR MENON —Faculty of Philosophy, University of Cambridge, Sidgwick Avenue, CambridgeCB3 9DA

Here is a plausible claim from particle physics: the states in a gauge multiplet correspond to (possibly distinct) configurations of the same type of particle. Take, for example, the spin-up and spin-down states (with respect to some axis) of an electron in an SU(2) multiplet. But surely, one might worry, not all such formal unifications count. Consider an electron-neutrino doublet, which is also an SU(2) multiplet. It seems less straightforward (or correct!) to consider these two be two states of the same type of particle. But why?

Consider what we might call the Redhead-Weingard thesis: two or more states in a multiplet of the gauge group of a quantum field theory are ontologically unified if they transform into each under the action of a spacetime transformation. The Redhead-Weingard thesis seems to generate the intuitively correct verdict in a number of cases, including the two SU(2) cases presented above. In ordinary relativistic QFTs, it works because the question of what structure counts as spatiotemporal is settled pretheoretically. But this fact conceals a contingent fact that is tacitly assumed across much theorising about spacetime: that spacetime plays the same inferential role regardless of the theoretical framework within which it is employed. The primary goal of this talk is to demonstrate this contingency by discussing how these roles come apart in supersymmetric quantum field theories.

AGPhil 7.2 Wed 17:00 AGPhil-H14

Cassirer and Weyl on the Constitutive Structure of Physical Theory — •NOAH STEMEROFF — University of Bonn

Though representative of divergent philosophical and intellectual traditions, both Ernst Cassirer and Hermann Weyl held that a given mathematical framework must always serve as a necessary presupposition of scientific thought (within a broadly Kantian position). Neither thought that this framework was fixed, a priori, as it was for Kant. However, in allowing for the revision of the constitutive framework of scientific thought, both were forced to face the spectre of a pervasive relativism. In response, each suggested that the relativist abolition of the standard of objectivity does not entail the abolition of the difference in value and performance of various scientific theories. On this view, scientific theories do not stand apart in their relation to the world, to be judged solely on their own merits, but rather as part of a progressive series. In this context, Cassirer and Weyl both highlighted the fundamental role that group theory played as a constitutive feature of our understanding of objectivity through the progress of science. In this paper, I will examine the differing views of Cassirer and Weyl concerning the constitutive role of group theory in physical enquiry, and what lessons we Location: AGPhil-H14

can draw from this history concerning modern debates on the methodology of physics.

AGPhil 7.3 Wed 17:30 AGPhil-H14

Naturalness and the Heuristic Role of Scientific Principles — •ENNO FIS-CHER — Bergische Universität Wuppertal, Interdisziplinäres Zentrum für Wissenschafts- und Technikforschung

The naturalness principle roughly demands that a theory should not involve independent parameters that are finely tuned. This principle was employed heavily over the last 40 years by theoretical physicists as a guideline for developing theories of beyond the Standard Model physics (BSM). However, since experiments at the Large Hadron Collider (LHC) have not found conclusive signs for new physics, the significance of naturalness arguments has been questioned and it has been suggested that high-energy physics has reached the "dawn of the postnaturalness era."

I argue that an explanation of the current shift in attitude towards naturalness can be given if we acknowledge that the naturalness principle has experienced epistemic support through the theories it has inspired. I argue that the potential coherence between major BSM proposals and the naturalness principle led to an increasing degree of credibility of the principle. The absence of new physics at the LHC has undermined the potential coherence and has led to the principle's current loss of significance. On the basis of this account I assess the heuristic role of naturalness as a guiding principle in high-energy physics and draw some tentative conclusions about the role of principles in the context of scientific progress.

AGPhil 7.4 Wed 18:00 AGPhil-H14

A neo-Kantian approach to the epistemology of the LHC flavour anomalies — •ALEX SEUTHE — Technische Universität Dortmund

Large scale experiments at the LHC, like the LHCb experiment, seek to answer questions about the fundamental structure of matter and the nature of the cosmos. Since the discovery of the Higgs boson and correspondingly of all predicted particles of the Standard Model, the field is faced with an open horizon for gaining knowledge. In flavour physics, various anomalies in $b \rightarrow s\ell^+\ell^-$ decays have been attracting attention in recent years. Although the single measurements are not yet statistically significant for a discovery, the overall picture might hint at possible extensions or modifications of the Standard Model. So far, only little epistemological reflection on the scientific process related to these anomalies has been presented. In my talk, I suggest a first attempt utilizing Ernst Cassirer's concept of science as a series process towards the limit of reality. Here, the experimental anomalies and theoretical explanatory models, including specific models or model-independent effective theories, stand in an alternating series of cognitions, oriented towards the ideal of reality as a regulative principle.

AGPhil 8: Annual Meeting of the AGPhil

Time: Wednesday 18:30–19:30 Annual Meeting of the AGPhil

AGPhil 9: Quantum Mechanics I

Time: Thursday 11:00-13:00

AGPhil 9.1 Thu 11:00 AGPhil-H14

A Heuristic Route to Nonlinear Quantum Mechanics — •ALIREZA JAMALI — 3rd Floor - Block No. 6 - Akbari Alley - After Dardasht Intersection - Janbazane Sharghi - Tehran - Iran

It is known since Madelung that the Schrödinger equation can be thought of as governing the evolution of an incompressible fluid, but the current theory fails to mathematically express this incompressibility in terms of the wavefunction without facing problem. In this paper after showing that the current definition of quantum-mechanical momentum as a linear operator is neither the most general nor a necessary result of the de Broglie hypothesis, a new definition is proposed that can yield both a meaningful mathematical condition for the incompressibility of the Madelung fluid, and nonlinear generalizations of Schrödinger and Klein-Gordon equations. The derived equations satisfy all conditions that are expected from a proper generalization: simplification to their linear counterparts by a well-defined dynamical condition; Galilean and Lorentz invariance (respectively); and signifying only rays in the Hilbert space. Location: AGPhil-MV

Location: AGPhil-H14

AGPhil 9.2 Thu 11:30 AGPhil-H14

Evidence for Interactive Common Causes. Resuming the Cartwright-Hausman-Woodward Debate — •PAUL M. Näger — University of Münster, Germany

The causal Markov condition (CMC), which is a central principle of causal modelling, requires that conditional on a common cause the correlation between its effects vanishes (the common cause screens off the correlation). Since Salmon (1978) presented the first counterexamples, joined by van Fraassen (1980, 1982) and Cartwright (1988 and many more), there is a debate about whether there are also common causes that fail to screen off (interactive common causes, ICCs), violating the CMC. Since indeterminism is a necessary requirement, the most serious candidates for ICCs refer to quantum phenomena. In her seminal debate with Hausman and Woodward, Cartwright early on focussed on unfortunate non-quantum examples (chemical factory). Especially, Hausman and Woodward's redescriptions of quantum cases saving the CMC remain unchallenged. This paper takes up this lose end of the discussion and aims to resolve the debate in favour of Cartwright's position. It systematically considers redescriptions of ICC structures, including those by Hausman and Woodward, and explains why

239

these are inappropriate, when quantum mechanics (in a dynamic collapse interpretation) is true. It first shows that all cases of purported quantum ICCs are cases of entanglement and then, using the tools of causal modelling, it provides an analysis of the quantum mechanical formalism for the case that the collapse of entangled systems is best described as a causal model with an ICC.

AGPhil 9.3 Thu 12:00 AGPhil-H14 Aristotelian Grounding for GRW's Flash Ontology — •RYAN MILLER — University of Geneva, Switzerland

The flash (i.e., event) ontology for the GRW objective-collapse formulation of quantum mechanics (Goldstein et al., 2012) has become popular for maintaining both a primitive ontology in 4D spacetime (Allori et al., 2014; Allori, 2015; Tumulka, 2017) and serious Lorentz invariance (Tumulka, 2009; Petrat & Tumulka, 2014a, 2014b; Tumulka, 2021). Valia Allori's (Allori et al., 2008; Allori, 2016) straightforward reading of this ontology suggests that the flashes are fundamental, grounding both the other elements of the theory and our everyday macro-scale ontology. This view has come under pressure on both points, however: Tim Maudlin (1997, 2010, 2011, 2019) argues that the GRW wavefunction cannot be wholly grounded in the flashes, while Elizabeth Miller (forthcoming) argues that flashes are an inadequate ground for everyday macro-scale ontology.

I suggest resolving these difficulties with the GRW flash ontology by grounding the flashes in entangled macro-objects. On this Aristotelian proposal, macroobjects like Schrodinger's cat maintain the entangled wavefunction that governs their micro-scale powers, realized in flash events. Because entangled particle families flash together (Maudlin, 2011), the density of micro-events will support macro-observations without the GRW parameters departing from observed values (Feldmann & Tumulka, 2012). Neo-Aristotelian grounding is thus attractive for GRW's flash ontology.

AGPhil 9.4 Thu 12:30 AGPhil-H14

Does the weak trace show the past of a quantum particle in an unperturbed system? — •JONTE R HANCE¹, JOHN RARITY¹, and JAMES LADYMAN² — ¹Quantum Engineering Technology Laboratories, Department of Electrical and Electronic Engineering, University of Bristol, Woodland Road, Bristol, BS8 1US, UK — ²Department of Philosophy, University of Bristol, Cotham House, Bristol, BS6 6JL, UK

We investigate the weak trace method for determining the path of a quantum particle in an unperturbed system. Specifically, looking at nested interferometer experiments, when internal interferometers are tuned to destructive interference, we show that the weak trace method gives misleading results. This is because the methods used experimentally to obtain the weak value of the position operator necessarily perturb the system, hence, in some cases the assumption that weak coupling being equivalent to no coupling is incorrect. Further, even if we assume there is no disturbance, there is no reason to associate the weak value of the spatial projection operator with the classical idea of 'particle presence', especially if it has features which go against the classical ideas associated with a particle being present (i.e. a particle having a single, continuous path). Experiments performed that are claimed to support the interpretation simply show the effects of this coupling acting as measurement, rather than tapping into the underlying reality of what happens in a quantum system when no-one is looking.

AGPhil 10: Quantum Mechanics II

Time: Thursday 14:00-16:00

AGPhil 10.1 Thu 14:00 AGPhil-H14

How to distinguish between indistinguishable particles — •MICHAEL TE VRUGT — Institut für Theoretische Physik, Center for Soft Nanoscience, Philosophisches Seminar, Westfälische Wilhelms-Universität Münster, 48149 Münster, Germany

A long and intense debate in philosophy is concerned with the question whether there can be haecceistic differences between possible worlds, that is, nonqualitative differences that only arise from different de re representations. According to haecceitism, it can give rise to a different situation if the positions of two qualitatively identical particles are exchanged, while according to anti-haecceitism, this is not the case. It has been suggested that classical statistical mechanics might provide evidence for one of these positions. However, most philosophers of physics argue that it does not. In this work [1], I show that order-preserving dynamics, a novel method from statistical mechanics developed for the description of nonergodic systems, changes this situation: It is intrinsically haecceistic and makes different experimental predictions than non-haecceistic alternatives. Thereby, it provides an empirical argument for the existence of modality de re.

[1] M. te Vrugt, British Journal for the Philosophy of Science (forthcoming), https://doi.org/10.1086/718495

AGPhil 10.2 Thu 14:30 AGPhil-H14 Who's afraid of retrocausation? A retrocausal explanation of Bell-type corre-

lations — • MATTHIAS ACKERMANN — Leibniz University Hanover Bell's theorem is commonly understood to have demonstrated that the observed statistics in quantum experiments rule out a 'locally causal' explanation. However, almost always the temporal aspect of 'local causality' seems to be implicitly assumed, rather than explicitly defined. Recent work by Wharton and Argaman (2020) does just that and with it offers a retrocausal framework that accounts for the correlations at the cost of an explicit relaxation of the usually implicit arrowof-time-thus, the argument goes, operationally saving Bell-compatible locality. This work assesses their proposal based on the central aspects of causal modelling (Pearl, 2009) and an influential no-go theorem by Wood and Spekkens (2015). Taking seriously the relaxation of the standard past-to-future description of physical systems, one can defend causal fine-tuning from being deemed 'unnatural' (Wood and Spekkens, 2015) or 'unsatisfactory' (Allen et al., 2017). Although Wharton and Argaman's (2020) retrocausal model indeed does fall victim to fine-tuning, this is due to an assumed underlying symmetry. The main finding is that taking these underlying symmetry considerations seriously lets one reasonably entertain the possibility that causes and signals do not necessarily co-occur. It is concluded that the framework of classical causal modelling is too restrictive of a framework to be home to and therefore to capture the notion of retrocausality.

Location: AGPhil-H14

AGPhil 10.3 Thu 15:00 AGPhil-H14

Configuration Space Realism and Fundamentality – •GABRIELLE KERBEL and NINA EMERY² — ¹University of Michigan — ²Mount Holyoke College The central question of quantum ontology is: what does the wavefunction represent? According to configuration space realism, the wavefunction represents a field (the 'wavefunction field') in a high-dimensional space (what we call 'cfspace'). According to the standard version of configuration space realism, which we call configuration space fundamentalism, the wavefunction field and cf-space are fundamental. We present a novel version of configuration space realism, called configuration space non-fundamentalism, according to which the wavefunction field and cf-space are non-fundamental. Instead, the wavefunction field and cf-space depend on three-dimensional space and the entities therein. We argue that configuration space non-fundamentalism should be taken at least as seriously as configuration space fundamentalism. Along the way we show how choosing between these different versions of configuration space realism will encourage metaphysicians and philosophers of physics alike to confront significant questions about the structure of grounding relations, the importance of locality and separability, and the nature of supervenience and scientific explanation

AGPhil 10.4 Thu 15:30 AGPhil-H14 Change and Time in Quantum Mechanics — •BRITTANY GENTRY — Utah State University, Logan, USA

While it is apparent that leading physical theories such as Relativity Theory and standard interpretations of Quantum Mechanics do not posit a real, or fundamental, time, the search for real time persists. One reason for continuing to posit real time is the concern that time is necessary to change. Examples of this concern as well as confusing claims that may lead others to that concern abound in philosophy of physics, even from physicists who agree that real time is unnecessary to physical theories. To address that concern, this paper argues that one way to separate time and change is to understand time as a construct that we use to slice up 4-dimensional Hilbert space into 3-dimensional space for the purpose of further distinguishing differences in the basic stuff occupying Hilbert space-namely, particles. On such a view, changes are the differences in positions that we observe in the stuff of Hilbert space and time is a construct that we sometimes place on this space to articulate these differences-and this conception of our QM models allows us to conceive of changes in a way that is independent of time. Time is a helpful feature of the model that we apply at certain levels, but not essential to the existence of the changes that we study. It leaves unaddressed the question of whether changes are real or apparent. However, this explanation makes progress in tidying up concerns regarding time in QM by removing the confusions surrounding the relationship between time and change.

AGPhil 11: Time and Temperature

Time: Thursday 16:15-18:45

AGPhil 11.1 Thu 16:15 AGPhil-H14

Taking seriously the problem of time of quantum gravity — •ALVARO MOZOTA FRAUCA — Autonomous University of Barcelona

In this paper I raise a worry about the most extended resolutions of the problem of time of canonical quantizations of general relativity. The reason for this is that these resolutions are based on analogies with deparametrizable models for which the problem can be solved, while I argue in this paper that there are good reasons for doubting about these resolutions when the theory is not deparametrizable, which is the case of general relativity. I introduce an example of a non-deparametrizable model, a double harmonic oscillator system expressed by its Jacobi action, and argue that the problem of time for this model is not solvable, in the sense that its canonical quantization doesn't lead to the quantum theory of two harmonic oscillators and the standard resolutions of the problem of time don't work for this case. I argue that as general relativity is strongly analogous to this model, one should take seriously the view that the canonical quantization of general relativity doesn't lead to a meaningful quantum theory. Finally, I comment that this has an impact on the foundations of different approaches to quantum gravity.

AGPhil 11.2 Thu 16:45 AGPhil-H14

quantum gravity and time's arrow: why primitivism should leave the floor to (local) reductionism — •LUCA GASPARINETTI — Venice, Italy

According to some primitivist approaches about the debate on time's arrow, spacetime is characterized by an intrinsic and global anisotropy of time, i.e., the temporal direction is a primitive and no further analyzable feature of the universe's geometry (Earman 1974 and Maudlin 2007). However, in several approaches to quantum gravity (e.g., causal set theory, loop quantum gravity, string theory), most philosophers of physics, e.g., Huggett (2021), Le Bihan (2021), Wüthrich (2018), state that spacetime disappears at the fundamental level and emerges in some sense from a non-spatiotemporal structure. Thus, the following question arises: given the disappearance of spacetime from the fundamental structure, what are the consequences for the primitivist approach about time's arrow?

In this paper, I argue that primitivism about time's arrow is seriously challenged by what quantum gravity theorists state about spacetime. More specifically, since spacetime is emergent, the direction of time, if it exists, reduces on a more fundamental asymmetry. It follows that if time's arrow is not primitive, the primitivist approach is false in the context of a theory of quantum gravity. Hence, I conclude that quantum gravity theorists have at their disposal only (local) reductionism, i.e., time's arrow is an extrinsic and, local or global, anisotropy of time.

AGPhil 11.3 Thu 17:15 AGPhil-H14

On the Status of Temperature and Thermodynamics in Relativity — •EUGENE Y. S. CHUA — University of California San Diego, La Jolla, CA

The project to understand black holes thermodynamically (i.e. black hole thermodynamics) was motivated by how stationary black holes can be characterized by laws analogous to the laws of classical thermodynamics. Taking this analogy seriously as evidence that black holes are thermodynamical seems to require that thermodynamics be relevant in the large-scale relativistic regime, viz. that there is a relativistic thermodynamics to speak of. However, an unresolved debate from the 1960s over the (lack of a) canonical Lorentz transformation for a central thermodynamic concept - temperature (and heat) - undermines this very assumption by asking whether thermodynamics could be relativized at all. By examining this debate, I argue that temperature, like absolute simultaneity, is not relativistic. We can readily judge simultaneity within a frame, just as co-moving observers can readily discern a system's temperature. However, the debate suggests there is no fact of the matter about the temperature of a moving object, just as there is no absolute sense that two objects moving relative to one another are simultaneous with each other. This pushes back against the idea that classical thermodynamics should be extended into the relativistic regime. The upshot for black hole thermodynamics: the thermodynamical analogy should not be taken too seriously.

AGPhil 11.4 Thu 17:45 AGPhil-H14 **The physical reality of a directed time** — •GRIT KALIES — HTW University of Applied Sciences, Dresden, Germany

Irreversibility has occupied philosophers and physicists for centuries. While quantum mechanics and special and general relativity interpret processes as reversible, thermodynamics describes every macroscopic process as irreversible. This divergence is called "Paradox of Time" [1].

In the 19th century, Max Planck was searching for a genuine irreversible microscopic process and refused to accept Ludwig Boltzmann's purely statistical interpretation of the second law of thermodynamics, which does not describe irreversibility at the quantum level [2]. Later, Boltzmann's interpretation was accepted.

Recent studies [3-7] show that Boltzmann and Clausius could not yet formulate the second law comprehensively due to the limited data available. As a result, physics was founded on symmetry principles. And yet: it exists, the irreversible process at the quantum level. The second law of thermodynamics can be further developed and understood as a fundamental law of nature, i.e. time symmetry is excluded.

[1] I. Prigogine, I. Stengers: Das Paradox der Zeit, Piper, München, Zürich, 1993; [2] L. Boltzmann, Sitzungsber. kaiserl. Akad. Wiss. Wien 66 (1872) 275-370; [3] G. Kalies: Vom Energieinhalt ruhender Körper, De Gruyter, Berlin, 2019; [4] G. Kalies, Z. Phys. Chem. 234 (2020) 1567-1602; [5] G. Kalies, Z. Phys. Chem. 235 (2021) 849-874; [6] G. Kalies: Back to the roots: The concepts of force and energy, Z. Phys. Chem. (2021) 1-53, DOI: 10.1515/zpch-2021-3122; [7] G. Kalies: On the unification of mechanics and thermodynamics, submitted (2021).

AGPhil 11.5 Thu 18:15 AGPhil-H14

Breaking Symmetry in Scientific Explanation — •BENJAMIN FALTESEK — Texas A&M University, College Station, TX, USA

The causal asymmetry problem plagues argument-form accounts of scientific explanation such as Kitcher*s unificationism. Such accounts require explanations to be sound and have some additional property; for Kitcher, the additional property is unifyingness: theory A is more unifying than theory B iff A explains more phenomena than B using as many or fewer ultimate facts and argument forms than B.

The causal asymmetry problem is that such accounts cannot distinguish good from bad explanations when there is an equation among the premises. An argument pattern that explains the length of a building*s shadow from the building*s height, for instance, can equally well explain the building*s height from the length of its shadow. Each explanation is equally sound and unifying, but the latter goes against causal dependence.

I propose a solution without relying on causal intuitions. For any explanation with an equation premise, the equation has a term E such that if E takes the value 0, the system at issue in the explanation cannot exist. This is not true of the other terms C. E represents the effect of the system, the phenomenon to be explained. I provide a schema for constructing explanatory argument forms that avoids the causal asymmetry problem by conditionalizing equations on the Cs.

AGPhil 12: Processes, Events and Time

Time: Friday 11:00-13:00

AGPhil 12.1 Fri 11:00 AGPhil-H14 Die Erfindung der Zeit — •HELMUT HILLE — Fritz-Haber-Straße 34, 74081

Heilbronn Die Zeit ist nicht nur Physikern und Philosophen ein Rätsel, das sie in immer neuen Anläufen zu entschlüsseln versuchen. Richtig ist, sie als Dimension zu bezeichnen, nämlich die des (zeitlichen) Nacheinanders, neben den 3 räumlichen Dimensionen des Neben-, Über- und Hintereinanders. Das räumliche Erleben geschieht dadurch, dass das Gehirn die 2-dimensionalen Bilder der Wahrnehmung * es gibt keine anderen! * so überlagert, dass ein räumlicher Eindruck entsteht. Gleiches geschieht mit den gerichteten Schallwahrnehmungen der beiden Ohren beim Stereohören. So wie ferner das Gehirn das Farbensehen zur besseren Unterscheidung von Objekten erfunden hat, wo gar keine Farben sind, so hat es auch das Zeiterleben erfunden, obwohl alle Dinge nur in der Gegenwart existieren, die zeitlos ist. Jeder Moment ist so gegenwärtig wie jeder andere, Die Rolle des Beobachters in allen Wahrnehmungen kann also gar nicht überschätzt werden. Trotzdem wird sie so wenig verstanden, weil das Gehirn das so will. Es möchte ungestört arbeiten können, weshalb es sich bedeckt hält. Es kann hier mit Hilfe der Neurophilosophie* gezeigt werden, wie es zum Zeiterleben kommt, das für unser Menschsein unverzichtbar ist. *Verbindung von Ergebnissen der Hirnforschung mit philosophischen Fragen

Location: AGPhil-H14

Location: AGPhil-H14

AGPhil 12.2 Fri 11:30 AGPhil-H14 Alfred North Whitehead und die Philosophiegeschichte — •Christian Tho-Mas Kohl — Daumstr.105, 13599 Berlin

Jede Philosophie bezieht ihre Farbe von der geheimen Lichtquelle eines Vorstellungshintergrunds, der niemals ausdrücklich in ihren Gedankenketten auftaucht

Vorstellungshintergrund. Ein Vorstellungshintergrund kann auch aus Vorurteilen und oberflächlichen Klischees bestehen oder aus einer Mischung von allem. In der Mathematikgeschichte ist manchmal von Axiomen oder von Grundsätzen die Rede. Als Physikhistoriker wende ich mich vor allem der Geschichte der Physik zu, meinem eigenen Hintergrund. Innerhalb einer physikalischen Theorie ist eine These ein Satz, der bestätigt werden soll.

AGPhil 12.3 Fri 12:00 AGPhil-H14

Events, structures and processes — •HANS JÜRGEN PIRNER — Institut für Theoretische Physik, Heidelberg

What are events? To answer this question, the talk analyzes physical and mental events. A singular event like the heat catastrophe of 2003 becomes meaningful when one relates it to the evolution of the climate. Structures emerge when events repeat themselves or when the experimenter makes them repeatable. In this way, physicists discovered the standard model of elementary particles, the cosmic microwave radiation and condensates of ultracold atoms. In general, events are parts of processes i.e. chains of events as will be shown for the birth of galaxies. In the university, researchers in separate faculties investigate physics and philosophy. This talk attempts to bridge this gap. By using the concept of event

one understands reality from both perspectives. The author discusses Whitehead's philosophy and gives examples of how to structure events. Phenomena in the brain or in artificial intelligence show the interaction of nature and mind. Paradoxical results in decisions and in cognition interpret the connection we are looking for. If events are fundamental, a deep insight into the interaction of nature and mind opens up.

AGPhil 12.4 Fri 12:30 AGPhil-H14

Heisenberg*s loop of knowledge and a mathematical model of the *thing in itself*: Circles theory — •MOHAMMED SANDUK — Department of Chemical and Process Engineering, University of Surrey, Guildford, GU2 7XH, UK

In philosophy of microscopic physics, Heinsberg introduced two concepts. The first concept is for the *nature in itself* and *nature as appears*. The observation can be regarded as a transformation from *nature in itself* to *nature as appears*. In the second one, Heinsberg, unlike Kant, opened a possibility to have a mathematical model for the *thing in itself*. This process may be a type of another transformation. It is a transformation from observable nature to nature in itself. These two concepts may form a loop of knowledge in microscopic nature. In an attempt to explain the complex harmonic oscillator (microscopic thing as appears), this loop has been adopted. This attempt led to develop a theory *circles theory*. This theory is not in quantum mechanics. The theory shows the process of the two transformations. The wave function has a form of complex harmonic oscillator. In spite of the fact that this theory is not in quantum mechanics, the results of transformations show a good similarity to relativistic quantum mechanics.

Abazi, Adrian DD 3.3, •DD 17.1,
DD 37.2
Abeln Andreas T 77 6
Abelson, Shannon SylvieAGPhil 2.3
Aberham, Vito•EP 7.4
Abicht, Nils Julius•T 6.8
Abouabid, Hamza T 85.3
Abraham, Velda Azalia•ST 2.2
Abreu, Samuel I 2.2 Abuladza Mariam
Ackermann Anke
Ackermann, Karl-Heinz T 104.3
Ackermann, Markus T 18.8, T 18.9,
T 46.2
Ackermann, Matthias•AGPhil 10.2
Ackermann, Nicola
Adan Danver Perez T 76 7
Adibekvan, Vardan
Aehle, Stefan•DD 6.3
Afsharipour, Armindokht•T 84.1,
T 84.4
Agarwal, Jessica
Ageev, Serger
Ahlburg Patrick T 95 7
Ahrenholtz, ImkeDD 27.2
AKMETE, ATAKAN TUĞBERK•T 3.8
Akolkar, Nilima•T 6.5
Al Kharabsheh, Dina DD 33.2
Alameddine, Jean-Marco EP 11.8,
•I 51.0, I 83.8, AKPIK I./
Albrecht Johannes T 4 3 T 43 6
T 57.1, T 57.2, T 57.3, T 57.5, T 57.7,
T 58.4, T 68.3, T 68.4
Albrecht, Maximilian
Albrecht, Steffen •T 32.6, T 87.1,
I 87.5 Aldava Maria T 25.4
Alemani Micol DD 39 1
Alemanno, Giulia EP 3.1
Alkakhi, Wael•T 12.3
Allison, HayleyEP 5.7
ALPS-Kollaboration T 38.2, T 38.3,
1 38.4
Alverade Curmen Liebeth CD 2.2
Alvarado-Guzman, Lisbeth•GP 2.3
Alvarado-Guzman, Lisbeth •GP 2.3 Alvarez Fernandez, Adrian •T 69.3 Alves Junior, Antonio Augusto
Alvarado-Guzman, Lisbeth•GP 2.3 Alvarez Fernandez, Adrian•T 69.3 Alves Junior, Antonio Augusto •T 47.8, T 72.5
Alvarado-Guzman, Lisbeth•GP 2.3 Alvarez Fernandez, Adrian•T 69.3 Alves Junior, Antonio Augusto •T 47.8, T 72.5 Ambrus, Victor EugenT 2.9
Alvarado-Guzman, Lisbeth•GP 2.3 Alvarez Fernandez, Adrian•T 69.3 Alves Junior, Antonio Augusto •T 47.8, T 72.5 Ambrus, Victor EugenT 2.9 Amendola, Luca
Alvarado-Guzman, Lisbeth•GP 2.3 Alvarez Fernandez, Adrian•T 69.3 Alves Junior, Antonio Augusto •T 47.8, T 72.5 Ambrus, Victor EugenT 2.9 Amendola, LucaGR 6.2 Ames, Christoph•T 15.1, •T 91.3,
Alvarado-Guzman, Lisbeth•GP 2.3 Alvarez Fernandez, Adrian•T 69.3 Alves Junior, Antonio Augusto •T 47.8, T 72.5 Ambrus, Victor EugenT 2.9 Amendola, LucaGR 6.2 Ames, Christoph•T 15.1, •T 91.3, T 91.4 Amin Neerai•T 101 9
Alvarado-Guzman, Lisbeth•GP 2.3 Alvarez Fernandez, Adrian•T 69.3 Alves Junior, Antonio Augusto •T 47.8, T 72.5 Ambrus, Victor EugenT 2.9 Amendola, LucaGR 6.2 Ames, Christoph•T 15.1, •T 91.3, T 91.4 Amin, Neeraj•T 101.9 Ammerman-Yebra, JuanT 72.7
Alvarado-Guzman, Lisbeth•GP 2.3 Alvarez Fernandez, Adrian•T 69.3 Alves Junior, Antonio Augusto •T 47.8, T 72.5 Ambrus, Victor EugenT 2.9 Amendola, LucaGR 6.2 Ames, Christoph•T 15.1, •T 91.3, T 91.4 Amin, Neeraj•T 101.9 Ammerman-Yebra, JuanT 72.7 Amorsos, Simone•T 1.2, T 88.2
Alvarado-Guzman, Lisbeth•GP 2.3 Alvarez Fernandez, Adrian•T 69.3 Alves Junior, Antonio Augusto •T 47.8, T 72.5 Ambrus, Victor Eugen T 2.9 Amendola, LucaGR 6.2 Ames, Christoph•T 15.1, •T 91.3, T 91.4 Amin, Neeraj•T 101.9 Ammerman-Yebra, JuanT 72.7 Amoroso, Simone•T 1.2, T 88.2 Anagnostatou, VassiaST 2.1
Alvarado-Guzman, Lisbeth•GP 2.3 Alvarez Fernandez, Adrian•T 69.3 Alves Junior, Antonio Augusto •T 47.8, T 72.5 Ambrus, Victor Eugen T 2.9 Amendola, LucaGR 6.2 Ames, Christoph•T 15.1, •T 91.3, T 91.4 Amin, Neeraj•T 101.9 Ammerman-Yebra, JuanT 72.7 Amoroso, Simone•T 1.2, T 88.2 Anagnostatou, VassiaT 9.6
Alvarado-Guzman, Lisbeth•GP 2.3 Alvarez Fernandez, Adrian•T 69.3 Alves Junior, Antonio Augusto •T 47.8, T 72.5 Ambrus, Victor Eugen T 2.9 Amendola, LucaGR 6.2 Ames, Christoph•T 15.1, •T 91.3, T 91.4 Amin, Neeraj•T 15.1, •T 91.3, T 91.4 Amin, Neeraj•T 101.9 Ammerman-Yebra, JuanT 72.7 Amoroso, Simone•T 12, T 88.2 Anagnostatou, VassiaT 31.6, T 92.5 Anglelscu, AndreiT 31.6, T 92.5
Alvarado-Guzman, Lisbeth•GP 2.3 Alvarez Fernandez, Adrian•T 69.3 Alves Junior, Antonio Augusto • T 47.8, T 72.5 Ambrus, Victor EugenT 2.9 Amendola, LucaGR 6.2 Ames, Christoph•T 15.1, •T 91.3, T 91.4 Amin, Neeraj•T 15.1, •T 91.3, T 91.4 Amin, Neeraj•T 12, T 88.2 Anagnostatou, VassiaT 9.6 Angelescu, AndreiT 9.6 Angelescu, AndreiT 9.4 ANNIF.KollaborationT 9.4
Alvarado-Guzman, Lisbeth•GP 2.3 Alvarez Fernandez, Adrian•T 69.3 Alves Junior, Antonio Augusto •T 47.8, T 72.5 Ambrus, Victor EugenT 2.9 Amendola, LucaGR 6.2 Ames, Christoph•T 15.1, •T 91.3, T 91.4 Amin, Neeraj•T 15.1, •T 91.3, T 91.4 Amin, Neeraj•T 12, T 88.2 Anagnostatou, VassiaT 9.6 Angelescu, AndreiT 9.6 Anglescu, AndreiT 9.4, T 59.5 Ansorae, HenrikT 9.4, T 59.5
Alvarado-Guzman, Lisbeth•GP 2.3 Alvarez Fernandez, Adrian•T 69.3 Alves Junior, Antonio Augusto •T 47.8, T 72.5 Ambrus, Victor Eugen7 2.9 Amendola, LucaGR 6.2 Ames, Christoph•T 15.1, •T 91.3, T 91.4 Amin, Neeraj•T 101.9 Ammerman-Yebra, JuanT 72.7 Amoroso, Simone•T 1.2, T 88.2 Anagnostatou, VassiaST 2.1 Androsov, KonstantinT 9.6 Angelescu, AndreiT 31.6, T 92.5 Angloher, GodeT 104.6 ANNIE-KollaborationT 59.4, T 59.5 Ansorge, HenrikT 04.3 ANTARES-KM3NET-ERLANGEN-
Alvarado-Guzman, Lisbeth•GP 2.3 Alvarez Fernandez, Adrian•T 69.3 Alves Junior, Antonio Augusto •T 47.8, T 72.5 Ambrus, Victor Eugen7 2.9 Amendola, LucaGR 6.2 Ames, Christoph•T 15.1, •T 91.3, T 91.4 Amin, Neeraj•T 101.9 Ammerman-Yebra, JuanT 72.7 Amoroso, Simone•T 1.2, T 88.2 Anagnostatou, VassiaST 2.1 Androsov, Konstantin
Alvarado-Guzman, Lisbeth•GP 2.3 Alvarez Fernandez, Adrian•T 69.3 Alves Junior, Antonio Augusto •T 47.8, T 72.5 Ambrus, Victor EugenT 2.9 Amendola, LucaGR 6.2 Ames, Christoph•T 15.1, •T 91.3, T 91.4 Amin, Neeraj•T 101.9 Ammerman-Yebra, JuanT 72.7 Amoroso, Simone•T 1.2, T 88.2 Anagnostatou, VassiaST 2.1 Androsov, KonstantinT 9.6 Angelescu, AndreiT 31.6, T 92.5 Angloher, GodeT 104.6 ANNIE-KollaborationT 59.4, T 59.5 Ansorge, HenrikT 104.3 ANTARES-KM3NET-ERLANGEN- Kollaboration T 25.6, T 71.8, T 99.1, T 99.2, T 103.1, T 103.2, T 103.3,
Alvarado-Guzman, Lisbeth•GP 2.3 Alvarez Fernandez, Adrian•T 69.3 Alves Junior, Antonio Augusto •T 47.8, T 72.5 Ambrus, Victor Eugen7 2.9 Amendola, LucaGR 6.2 Ames, Christoph•T 15.1, •T 91.3, T 91.4 Amin, Neeraj•T 101.9 Ammerman-Yebra, JuanT 72.7 Amoroso, Simone•T 1.2, T 88.2 Anagnostatou, VassiaST 2.1 Androsov, KonstantinT 9.6 Angelescu, AndreiT 31.6, T 92.5 Angloher, GodeT 31.6, T 92.5 Angloher, GodeT 104.3 ANTARES-KM3NET-ERLANGEN- Kollaboration T 25.6, T 71.8, T 99.1, T 99.2, T 103.1, T 103.2, T 103.3, T 107.2, T 107.3
Alvarado-Guzman, Lisbeth•GP 2.3 Alvarez Fernandez, Adrian•T 69.3 Alves Junior, Antonio Augusto •T 47.8, T 72.5 Ambrus, Victor Eugen7 2.9 Amendola, LucaGR 6.2 Ames, Christoph•T 15.1, •T 91.3, T 91.4 Amin, Neeraj•T 15.1, •T 91.3, T 91.4 Amin, Neeraj•T 101.9 Ammerman-Yebra, JuanT 72.7 Amoroso, Simone•T 1.2, T 88.2 Anagnostatou, VassiaST 2.1 Androsov, KonstantinT 9.6 Angelescu, AndreiT 31.6, T 92.5 Angloher, GodeT 104.6 ANNIE-KollaborationT 59.4, T 59.5 Ansorge, HenrikT 104.3 ANTARES-KM3NET-ERLANGEN- Kollaboration T 25.6, T 71.8, T 99.1, T 99.2, T 103.1, T 103.2, T 103.3, T 107.2, T 107.3 Anton, Gisela
Alvarado-Guzman, Lisbeth•GP 2.3 Alvarez Fernandez, Adrian•T 69.3 Alves Junior, Antonio Augusto •T 47.8, T 72.5 Ambrus, Victor Eugen7 2.9 Amendola, LucaGR 6.2 Ames, Christoph•T 15.1, •T 91.3, T 91.4 Amin, Neeraj•T 101.9 Ammerman-Yebra, JuanT 72.7 Amoroso, Simone•T 1.2, T 88.2 Anagnostatou, VassiaST 2.1 Androsov, KonstantinT 9.6 Angelescu, AndreiT 31.6, T 92.5 Angloher, GodeT 104.3 ANTARES-KM3NET-ERLANGEN- Kollaboration T 25.6, T 71.8, T 99.1, T 99.2, T 103.1, T 103.2, T 103.3, T 107.2, T 107.3 Anton, Gisela
Alvarado-Guzman, Lisbeth•GP 2.3 Alvarez Fernandez, Adrian•T 69.3 Alves Junior, Antonio Augusto •T 47.8, T 72.5 Ambrus, Victor Eugen7 Amendola, LucaGR 6.2 Ames, Christoph•T 15.1, •T 91.3, T 91.4 Amin, Neeraj•T 15.1, •T 91.3, T 91.4 Amin, Neeraj•T 12, T 88.2 Anagnostatou, Vassia•T 12, T 88.2 Anagnostatou, VassiaT 9.6 Angelescu, AndreiT 9.6 Angloher, GodeT 104.3 ANNIE-KollaborationT 59.5 Ansorge, HenrikT 104.3 ANTARES-KM3NET-ERLANGEN- Kollaboration T 25.6, T 71.8, T 99.1, T 99.2, T 103.1, T 103.2, T 103.3, T 107.2, T 107.3 Anton, Gisela
Alvarado-Guzman, Lisbeth•GP 2.3 Alvarez Fernandez, Adrian•T 69.3 Alves Junior, Antonio Augusto •T 47.8, T 72.5 Ambrus, Victor Eugen7 2.9 Amendola, LucaGR 6.2 Ames, Christoph•T 15.1, •T 91.3, T 91.4 Amin, Neeraj•T 15.1, •T 91.3, T 91.4 Amin, Neeraj•T 12, T 88.2 Anagnostatou, Vassia•T 12, T 88.2 Anagnostatou, VassiaT 9.6 Angelescu, AndreiT 9.6 Angelescu, AndreiT 9.6 Angloher, GodeT 04.3 ANTIAES-KM3NET-ERLANGEN- Kollaboration T 25.6, T 71.8, T 99.1, T 99.2, T 103.1, T 103.2, T 103.3, T 107.2, T 107.3 Anton, GiselaT 12.9 Antonio Fernandez Pretel, Jose T 39.2 Anuar, AfiqT 33.1, T 36.1, •T 54.3,
Alvarado-Guzman, Lisbeth•GP 2.3 Alvarez Fernandez, Adrian•T 69.3 Alvarez Fernandez, Adrian•T 69.3 Alves Junior, Antonio Augusto • T 47.8, T 72.5 Ambrus, Victor Eugen T 2.9 Amendola, Luca
Alvarado-Guzman, Lisbeth•GP 2.3 Alvarez Fernandez, Adrian•T 69.3 Alves Junior, Antonio Augusto •T 47.8, T 72.5 Ambrus, Victor Eugen7 2.9 Amendola, LucaGR 6.2 Ames, Christoph•T 15.1, •T 91.3, T 91.4 Amin, Neeraj•T 15.1, •T 91.3, T 91.4 Amin, Neeraj•T 15.1, •T 91.3, T 91.4 Amin, Neeraj•T 12, T 88.2 Anagnostatou, VassiaT 12, T 88.2 Anagnostatou, VassiaT 9.6 Angelescu, AndreiT 9.6 Angelescu, AndreiT 9.6 Angelescu, AndreiT 9.6 Angloher, GodeT 9.6 Ansorge, Henrik
Alvarado-Guzman, Lisbeth•GP 2.3 Alvarez Fernandez, Adrian•T 69.3 Alves Junior, Antonio Augusto •T 47.8, T 72.5 Ambrus, Victor Eugen7 2.9 Amendola, Luca
Alvarado-Guzman, Lisbeth•GP 2.3 Alvarez Fernandez, Adrian•T 69.3 Alves Junior, Antonio Augusto •T 47.8, T 72.5 Ambrus, Victor Eugen7 2.9 Amendola, Luca
Alvarado-Guzman, Lisbeth•GP 2.3 Alvarez Fernandez, Adrian•T 69.3 Alves Junior, Antonio Augusto •T 47.8, T 72.5 Ambrus, Victor Eugen7 2.9 Amendola, LucaGR 6.2 Ames, Christoph•T 15.1, •T 91.3, T 91.4 Amin, Neeraj•T 15.1, •T 91.3, T 91.4 Amin, Neeraj•T 101.9 Ammerman-Yebra, JuanT 72.7 Amoroso, Simone•T 1.2, T 88.2 Anagnostatou, VassiaST 2.1 Androsov, Konstantin
Alvarado-Guzman, Lisbeth•GP 2.3 Alvarez Fernandez, Adrian•T 69.3 Alves Junior, Antonio Augusto •T 47.8, T 72.5 Ambrus, Victor Eugen7 2.9 Amendola, LucaGR 6.2 Ames, Christoph•T 15.1, •T 91.3, T 91.4 Amin, Neeraj•T 15.1, •T 91.3, T 91.4 Amin, Neeraj•T 101.9 Ammerman-Yebra, JuanT 72.7 Amoroso, Simone•T 1.2, T 88.2 Anagnostatou, VassiaST 2.1 Androsov, Konstantin
Alvarado-Guzman, Lisbeth GP 2.3 Alvarez Fernandez, Adrian
Alvarado-Guzman, Lisbeth +GP 2.3 Alvarez Fernandez, Adrian +T 69.3 Alves Junior, Antonio Augusto • T 47.8, T 72.5 Ambrus, Victor Eugen +T 2.9 Amendola, Luca GR 6.2 Arnes, Christoph +T 15.1, •T 91.3, T 91.4 Amin, Neeraj +T 15.1, •T 91.3, T 91.4 Amin, Neeraj +T 101.9 Ammerman-Yebra, Juan +T 2.7 Amoroso, Simone •T 1.2, T 88.2 Anagnostatou, Vassia ST 2.1 Androsov, Konstantin T 9.6 Angelescu, Andrei T 31.6, T 92.5 Angloher, Gode T 104.3 ANTARES-KM3NET-ERLANGEN- Kollaboration T 59.4, T 59.5 Ansorge, Henrik T 104.3 ANTARES-KM3NET-ERLANGEN- Kollaboration T 25.6, T 71.8, T 99.1, T 99.2, T 103.1, T 103.2, T 103.3, T 107.2, T 107.3 Anton, Gisela T 51.8 Antonello, Massimiliano T 12.9 Antonio Fernandez Pretel, Jose T 39.2 Anuar, Afiq T 33.1, T 36.1, •T 54.3, T 76.7 Apel, Frederike EP 7.6 Appel, Simon T 20.8 Appelt, Christian +T 37.1, T 43.4 Araujo, Gabriela R T 78.5 Araya, Ignacio AGPhil 6.4 Arbet-Engels, Axel EP 11.2, T 83.2 Archilli, Flavio 4.4 Argyropoulos, Spyridon T 62.4
Alvarado-Guzman, Lisbeth•GP 2.3 Alvarez Fernandez, Adrian•T 69.3 Alverez Fernandez, Adrian•T 69.3 Alves Junior, Antonio Augusto • T 47.8, T 72.5 Ambrus, Victor Eugen T 2.9 Amendola, LucaGR 6.2 Ames, Christoph•T 15.1, •T 91.3, T 91.4 Amin, Neeraj•T 15.1, •T 91.3, T 91.4 Amin, Neeraj•T 101.9 Ammerman-Yebra, JuanT 72.7 Amoroso, Simone•T 1.2, T 88.2 Anagnostatou, VassiaST 2.1 Androsov, KonstantinT 9.6 Angelescu, AndreiT 104.3 ANTARES-KM3NET-ERLANGEN- KollaborationT 59.4, T 59.5 Ansorge, HenrikT 104.3 ANTARES-KM3NET-ERLANGEN- Kollaboration T 25.6, T 71.8, T 99.1, T 99.2, T 103.1, T 103.2, T 103.3, T 107.2, T 107.3 Anton, GiselaT 12.9 Antonio Fernandez Pretel, Jose T 39.2 Anuar, AfiqT 33.1, T 36.1, •T 54.3, T 76.7 Apel, FrederikeT 20.8 Appelt, Christian
Alvarado-Guzman, Lisbeth
Alvarado-Guzman, Lisbeth
Alvarado-Guzman, Lisbeth
Alvarado-Guzman, Lisbeth GP 2.3 Alvarez Fernandez, Adrian
Alvarado-Guzman, Lisbeth•GP 2.3 Alvarez Fernandez, Adrian•T 69.3 Alves Junior, Antonio Augusto •T 47.8, T 72.5 Ambrus, Victor Eugen
Alvarado-Guzman, Lisbeth•GP 2.3 Alvarez Fernandez, Adrian•T 69.3 Alves Junior, Antonio Augusto •T 47.8, T 72.5 Ambrus, Victor Eugen
Alvarado-Guzman, Lisbeth
Alvarado-Guzman, Lisbeth •GP 2.3 Alvarez Fernandez, Adrian •T 69.3 Alves Junior, Antonio Augusto • T 47.8, T 72.5 Ambrus, Victor Eugen T 2.9 Amendola, Luca GR 6.2 Arnes, Christoph •T 15.1, •T 91.3, T 91.4 Amin, Neeraj •T 15.1, •T 91.3, T 91.4 Amin, Neeraj •T 101.9 Ammerman-Yebra, Juan T 72.7 Amoroso, Simone •T 1.2, T 88.2 Anagnostatou, Vassia ST 2.1 Androsov, Konstantin T 9.6 Angelescu, Andrei T 31.6, T 92.5 Angloher, Gode T 104.3 ANTARES-KM3NET-ERLANGEN- Kollaboration T 59.4, T 59.5 Ansorge, Henrik T 104.3 ANTARES-KM3NET-ERLANGEN- Kollaboration T 25.6, T 71.8, T 99.1, T 99.2, T 103.1, T 103.2, T 103.3, T 107.2, T 107.3 Anton, Gisela T 51.8 Antonello, Massimiliano T 12.9 Antonio Fernandez Pretel, Jose T 39.2 Anuar, Afiq

Bäcker, Claus Maximilian •ST 1.6, ST 6.2 Badawi, Ramsey D. ST 9.2 Bao Ta, Duc T 92.1 Barnes, Micah ST 3.4 Barth, Arnulf•T 102.4, T 102.5, T 102.6 T 102.6 T 94.4 Basan, Alexander •T 34.2, T 34.3, T 35.2, T 35.3

 Bechtle, Philip
 T. 5.1, T 8.2, T 37.4, T 63.7, T 64.8, T 92.6, T 107.6

 Becker, Anna
 T 65.1

 Becker, Jürgen
 T 43.8

 Becker, Maik
 T 57.1

 Becker, T, 73.9, R T 01.1, T 101.2
 Beckstein, Michael

 Behrends, Carina
 DD 41.3

 Behrends, Carina
 DD 12.6

 Bein, Fabian
 DD 12.6

 Beil, Fabian
 DD 12.6

 Beil, Fabian
 DD 12.3

 Bein, Samuel
 T 9.2, T 9.3, T 9.7, T 53.9, AKPIK 2.4

 Beirer, Joshua
 T 70.9

 Beisegel, Florian
 T 61.4

 Belenchia, Alessio
 GR 6.4

 Bell, Guido
 T 2.5

 BellageT 79.5, T 63.3, T 3.4, T 30.6, T 84.2, T 30.7, T 95.1, T 3.5, T 95.3, T 95.7, Bender, Stefan •EP 5.4 Benecke, Anna T 66.9
 Bento, Antonio
 T 104.6

 Benz, Gregor
 •DD 50.1

 Berdalovic, Ivan
 T 40.8

 Berg, Eric
 ST 9.2

 Berger, Bernd
 ... T 65.5, T 65.6, T 65.7,
 T 94.4 Berger, Marten•T 31.4 Berger, Tilmann DD 5.2 Bergermann, Dominik T 105.10 Bergert, Roman ST 9.1

 Augusthy, Amala
 •T 12.8

 Austrup, Volker
 T 10.2, •T 37.3

 Averbeck, Janis
 •T 106.7

 Avraamidou, Lucy
 •SYNS 1.2

 Awal, Awal
 \$T 3.5

•AKPIK 4.4

Berggren, Mikael •T 9.9, T 63.6 Bergmann, Antje ... DD 31.1, DD 31.2, DD 35.6, DD 36.2, DD 36.3
 DD 35.6, DD 36.2, DD 36.3

 Berker, Daniel
 ST 6.5

 Berlea, Vlad Dumitru
 •T 13.7

 Bernabò, Lia Marta
 •EP 2.6

 Bernhard, Peter
 T 41.1, T 59.2, T 97.5

 Bernloghoff, Daniel
 T 96.6

 Bernlochner, Florian
 T 3.1, T 3.3, T 3.4, T 3.5, T 16.5, T 16.6, T 30.6, T 20.0, 5 - 1 4.1
 T 30.9, T 52.1, T 64.7, T 79.6, T 84.1, T 84.2, T 84.4, T 84.6, T 84.7, T 84.8, T 95 3 T 95 7 Bespin, Christian •T 40.8, T 40.9, T 952 AKPIK 2.4 Biedenbach, Jana•DD 22.3 •AKPIK 4.2 Biermann, Lisa •T 92.3 Bifani, Simone T 4.4 Bilk, Johannes •T 70.2, AKPIK 2.3 Bilokin, Sviatoslav •T 30.2 Bilow, Uta •T 24.1, T 24.2, T 24.3 Bindi, Marcello T 12.6, T 14.6 Birk, Joschka •T 26.4 Birkenfeld, Thilo •T 99.3 Bismark, Kathrin •T 19.5 Bister, Teresa T 107.4 Bitzenbauer, Philipp •DD 45.3 Blake, Matthew •T 85.2, •T 85.9 Blake, Thomas T 56.8 Blank, Jonah •T 57.6

 Blake, Matthew
 1103.2, 1103.3

 Blake, Jonah
 156.8

 Blank, Michael
 145.7

 Blasi, Simone
 131.6

 Biesmer, Kai
 0D 27.1, DD 27.2

 Block, Fabian
 1721.3, T 21.9, T 49.6

 Bloemer, Julia
 6P 5.4

 Blum, David
 T 20.3, T 20.9, T 48.6, T 48.8, T 102.3

 Blumenthal, Hanna
 0D 10.5

 Blumenthal, Julian
 15.6 7 69.3

 Bock-Röllar, Gianluca
 6P 4.2

 Böckh, Tobias
 16.6

 Bock-Müller, Klaus
 DD 6.2

 Bodensiek, Oliver
 ... DD 3.2, DD 38.1

 Bodensiek, Oliver
 ... DD 3.5, 13, DD 41.2

 Boehauve, Julian
 15.4, D 41.2

 Boehauve, Julian
 17.4, S 10.4

 Boebe, Artur
 194.3

 Boelhauve, Julian
 • T 58.4

 Boger, Achim
 DD 34.4

 Bohlen, Simon
 T 44.5, T 44.6

 Böhler, Michael
 ... T 8.3, T 15.2, T 15.5

 Böhles, Manuel
 ... T 48.10, • T 102.2

 Böhn, Burkhard
 ... • T 11.7

 Böhm, Burkhard
 •T 11.7

 Böhme, Lukas
 •EP 2.1

 Bohnensack, Jan-Niklas
 •T 10.2

 Boin, Manuela
 .DD 5.2, •DD 34.4

 Boisson, Catherine
 EP 7.5

 Bolles, Janina
 T 72.2, •T 72.3

 Bollig, Julian
 •T 69.6

 Bolte, Claus
 .DD 18.1

 Bolz, Arthur
 T 95.5, T 95.6

 Bonet, Hannes
 .T 23.1

 Boncore, Domenico
 .MP 2.6, •T 2.1

 Bontempo, Federico
 •T 100.3

Borexino-Kollaboration T 103.8, T 80.1, EP 11.9 AKPIK 2.9 T 44 6 Borysova, Maryna Böser, Sebastian T 23.5, T 23.6, T 49.2, T 49.3 T 102.7 Bostelmann, Henning • MP 7.1 Brandes, Jürgen•GR 10.7 Brandl, Carlos Andres ST 2.4, •ST 3.2
 Brandl, Carlos Andres
 ST 2.4, •ST 3.2

 Brandt, Gerhard
 T 14.7

 Bräuer, Florian
 DD 35.7

 Braun, Daniel
 GR 6.4

 Breer, Nils
 •T 96.7, T 96.8

 Brehmer, Johann
 T 90.3

 Breisch, Marc
 T 20.3, T 20.9, T 48.6, T 48.8, •T 59.4, T 102.3

 Breitmoser, Dominik
 •T 105.3

 Brenner, Anja
 T 77.7

 Breschi, Matteo
 GR 9.4

 Bresges, André
 DD 22.1, DD 22.2, DD 29.2, DD 30.1, DD 43.1, DD 43.2, •DD 46.1, DD 47.1, DD 48.3

 Bretz, Thomas
 GR 14.2, T 72.9
 •0D 46.1, DJ 47.1, DJ 48.3 Bretz, Thomas GR 14.2, T 72.9 Breuer, Doris EP 3.1 Breuhaus, Mischa •T 101.5 Brignoli, Alessia T 68.2 Brinkmann, Kai-Thomas ST 9.1, Brommer, Sebastian . •T 25.1, T 36.4, T 36.5 Brostean-Kaiser, Jannes •T 52.4
 T 62.0, 1 90.7, 1 92.1

 Büsken, Max
 •T 51.1

 Bußmann, Lara
 •ST 3.3

 Buson, Sara
 •T 29.4

 Buss, Thorsten
 •T 93.4

 Bussemer, Peter
 •GP 9.1

 Pursony Peter
 T 10.9

Bütikofer Polf ED 5.8
Butter, AnjaT 2.6, T 2.7, •T 2.8, T 45 3 •T 55 1 T 90 3 AKPIK 4 2
Cacciapaglia, Giacomo
Caicedo, Ivan
Calefice, Lukas T 57.1, •T 57.7, T 68.4
T 41.4, T 41.5, T 41.6
Campello, Fabio•T 31.5 Campos Fauth, AndersonT 51.4
Canonica, Lucia •T 104.4, T 104.6 Cao, Lu T 84.7
Capel, Angela•MP 1.3 Caracas, Ioana AlexandraT 71.1,
•T 71.2 Cardinal, Kai •DD 5.3, DD 8.1, DD 43.3
Cardini, Andrea T 9.6 Carmesin, Hans-Otto•GR 10.6,
DD 7.3, DD 7.4, DD 7.5, •DD 26.1, •DD 38.2
Carminati, Marco T 21.1, T 21.8 Carone Ludmila
Cartarius, Holger DD 6.3, DD 34.12,
DD 41.1, DD 44.3, DD 45.1 Cartwright Casey MP 2.3
Carus, Leon
Casado Moran, David
Castro-Ruiz, Esteban MP 5.1
Çekmecelioğlu, Yusuf Can •T 58.5
Cerruti, Matteo EP 11.2, T 83.2
Cervato, Beatrice
Ceslik, Harald 194.3 Cetin, Nadine
Chakraborty, Sagnik MP 5.4
Chataigner, Leonardo•SYMD 1.1 Chaurasia, Swami Vivekanandii
GR 13.2
Chauvin, Antoine T 50.4, •T 50.5 Chop Zo
Cheng, Chun•T 94.5
Cherry, Simon RST 9.2
Chiesa, Mauro 1 90.2 Chobola. Tomas
Choi, Kyungeon 1 60.3, 1 60.4
Choi, Kyungeon 1 60.3, 1 60.4 Christner, MarianDD 31.2 Chruściński Dariusz MP 5 4
Choi, Kyungeon 1 60.3, 1 60.4 Christner, MarianDD 31.2 Chruściński, Dariusz MP 5.4 Chua, Eugene Y. S •AGPhil 11.3
Choi, Kyungeon 1 60.3, 1 60.4 Christner, Marian DD 31.2 Chruściński, Dariusz MP 5.4 Chua, Eugene Y. S •AGPhil 11.3 Chwalek, Thorsten T 7.5, T 87.4, T 87.5
Choi, Kyungeon 160.3, 160.4 Christner, Marian DD 31.2 Chruściński, Dariusz MP 5.4 Chua, Eugene Y. S. •AGPhil 11.3 Chwalek, Thorsten T.7.5, T 87.4, T 87.5 Cichon, Dominick •T 78.2 Cidala Ludia EP 12
Choi, Kyungeon 1 60.3, 1 60.4 Christner, Marian DD 31.2 Chruściński, Dariusz MP 5.4 Chua, Eugene Y. S. AGPhil 11.3 Chwalek, Thorsten T 7.5, T 87.4, T 87.5 Cichon, Dominick T 78.2 Cidale, Lydia EP 1.3 Cieri, Davide
Choi, Kyüngeon 1 60.3, 1 60.4 Christner, Marian DD 31.2 Chruściński, Dariusz MP 5.4 Chua, Eugene Y. S AGPhil 11.3 Chwalek, Thorsten T 7.5, T 87.4, T 87.5 Cichon, Dominick T 78.2 Cidale, Lydia EP 1.3 Cieri, Davide T 14.5, •T 43.2, T 67.3 Ciezarek, Gregory Max T 43.6
Choi, Kyüngeon 160.3, 160.4 Chruściński, Dariusz DD 31.2 Chruściński, Dariusz MP 5.4 Chua, Eugene Y. S. •AGPhil 11.3 Chwalek, Thorsten T 7.5, T 87.4, T 87.5 Cichon, Dominick •T 78.2 Cidale, Lydia EP 1.3 Cieri, Davide •T 14.5, •T 43.2, T 67.3 Cizearek, Gregory Max T 43.6 Cirkel, Jasper O. DD 25.2 Claessens, Christine •SYMD 1.2
Choi, Kyüngeon 1 60.3, 1 60.4 Christner, Marian DD 31.2 Chruściński, Dariusz MP 5.4 Chua, Eugene Y. S •AGPhil 11.3 Chwalek, Thorsten T 7.5, T 87.4, T 87.5 Cichon, Dominick T 78.2 Cidale, Lydia T 78.2, T 77.3 Ciezarek, Gregory Max T 43.2, T 67.3 Ciezarek, Gregory Max T 43.2, T 67.3 Ciezarek, Gregory Max T 43.6 Cirkel, Jasper O DD 25.2 Claessens, Christine
Choi, Kyungeon 160.3, 160.4 Christner, Marian DD 31.2 Chruściński, Dariusz MP 5.4 Chua, Eugene Y. S. AGPhil 11.3 Chwalek, Thorsten T 7.5, T 87.4, T 87.5 Cichon, Dominick T 78.2 Cidale, Lydia EP 1.3 Cieri, Davide •T 14.5, •T 43.2, T 67.3 Ciezarek, Gregory Max T 43.6 Cirkel, Jasper O. DD 25.2 Classen, Christine •SYMD 1.2 Classen, Lew T 106.5, T 106.6, T 106.7, T 106.8 Clavijo Columbié, José Manuel
Choi, Kyüngeon 1 60.3, 1 60.4 Christner, Marian DD 31.2 Chruściński, Dariusz MP 5.4 Chua, Eugene Y. S AGPhil 11.3 Chwalek, Thorsten T 7.5, T 87.4, T 87.5 Cichon, Dominick T 78.2 Cidale, Lydia EP 1.3 Cieri, Davide•T 14.5, •T 43.2, T 67.3 Ciezarek, Gregory Max T 43.6 Cirkel, Jasper O DD 25.2 Claessens, Christine•SYMD 1.2 Claessen, Lew T 106.5, T 106.6, T 106.7, T 106.8 Clavijo Columbié, José Manuel •T 89.3 Clercy, Joshua
Choi, Kyüngeon 1 60.3, 1 60.4 Christner, Marian DD 31.2 Chruściński, Dariusz MP 5.4 Chua, Eugene Y. S •AGPhil 11.3 Chwalek, Thorsten T 7.5, T 87.4, T 87.5 Cichon, Dominick T 78.2 Cidale, Lydia EP 1.3 Cieri, Davide T 14.5, •T 43.2, T 67.3 Ciezarek, Gregory Max T 43.6 Cirkel, Jasper O DD 25.2 Claessens, Christine •SYMD 1.2 Classen, Lew T 106.5, T 106.6, T 106.7, T 106.8 Clavijo Columbié, José Manuel •T 89.3 Clercx, Joshua T 92.4 Climescu, Matei T 41.7
Choi, Kyüngeon 1 60.3, 1 60.4 Christner, Marian
Choi, Kyüngeon 1 60.3, 1 60.4 Christner, Marian
Choi, Kyüngeon 1 60.3, 1 60.4 Christner, Marian
Choi, Kyüngeon 1 60.3, 1 60.4 Christner, Marian
Choi, Kyüngeon 1 60.3, 1 60.4 Christner, Marian
Choi, Kyüngeon 1 60.3, 1 60.4 Christner, Marian
Choi, Kyüngeon 1 60.3, 1 60.4 Christner, Marian
Choi, Kyüngeon 1 60.3, 1 60.4 Christner, Marian
Choi, Kyüngeon 1 60.3, 1 60.4 Christner, Marian
Choi, Kyüngeon 1 60.3, 1 60.4 Christner, Marian
Choi, Kyüngeon 1 60.3, 1 60.4 Christner, Marian
Choi, Kyüngeon 1 60.3, 1 60.4 Christner, Marian
Choi, Kyüngeon 1 60.3, 1 60.4 Christner, Marian
Choi, Kyüngeon 1 60.3, 1 60.4 Christner, Marian
Choi, Kyüngeon 1 60.3, 1 60.4 Christner, Marian
Choi, Kyüngeon 1 60.3, 1 60.4 Christner, Marian
Choi, Kyüngeon
Choi, Kyüngeon 1 60.3, 1 60.4 Christner, Marian

Crivellin, AndreasT 56.1 CRPropa-Kollaboration T 47.5 Csakli, Simon T 20.8, T 99.5 Csizmadia, Szilárd EP 2.6 CTA-Kollaboration T 17.8, T 17.1, AKPIK 2.1, T 17.4 Cuenca-García, Jose•T 50.3 Dall'Occo, Elena . T 4.3, T 68.3, T 68.4 T 50.6 de la Hamette, Anne-Catherine MP 5.1 de Leo, Ksenia •T 10.4, T 10.8, T 62.6, T 62.7

 de Oña-Wilhelmi, Emma
 T 45.1

 de Silva, Malinda
 • T 41.3

 De Vellis, Fabio
 • T 57.3

 Debowski, Karol
 • T 49.8

 Deckert, Rosanna
 • T 75.4

 Dedu, Vlad
 • T 4.4

 Defranchis, Matteo
 • T 7.1, T 7.5

 Defranchis, Matteo M.
 • SYMD 1.3

 Deka, Pranab
 EP 6.7

 Del Rio, Fernado
 • T 43.3

 Del Rio Viera, Manuel Alejandro
 • T 40.1

 •T 40.1 • 1 40.1 Delampady, Srijan T 20.3, T 20.9, T 48.6, T 48.8, T 102.3 Deliot, Frederic T 60.3, T 60.4 Dembinski, Hans • EP 10.4, •T 3.7, T 58.4, •T 81.4 Demchenko, Ivan T 24.4, T 52.2 Denzia Imbef Delampady, Srijan Dennis, Imhof DD 4.2, DD 20.3, DD 20.4 Derugin, Evelin ST 7.2, ST 7.3, ST 7.4 Desch, Klaus T 5.1, T 8.2, T 14.2, T 37.4, T 38.5, T 39.8, T 63.7, T 64.8, T 68.1, T 69.7, T 97.1, T 97.6, T 107.6 •T 31.8 Dickmann, Martin . •DD 47.3, DD 50.3 AKPIK 4.2 T 94 1 T 3.5, T 12.4, T 13.5, T 16.5, T 16.7, T 3.0, T 3.0, G, T 30.9, T 40.5, T 40.8, T 40.9, T 52.1, T 61.4, T 62.2, T 62.3, T 84.1, T 84.2, T 84.4, T 84.6, T 84.7, T 84.8, T 89.2, T 95.1, T 95.2, T 95.3, T 95.7 Dinkel, Martin • T 46.8 Dirr, Gunther • MP 5.3

Dittmer, Markus T 106.5, T 106.6, T 106.7
Dittrich, CarstenT 20.8, T 99.5 Do, Mia Giang •EP 4.3, EP 4.4, EP 4.5, EP 4.6
Doboszewski, Juliusz •AGPhil 5.2 Dodonova, Alena T 26.3, T 89.2, •T 90.6
Doerenkamp, Maike•T 50.4, T 50.5 Domann, Osvaldo•EP 7.11 Domcke, Valerie•EP 8.2.•T 80.2
Dominik, Rune Michael •AKPIK 2.1 Dopatka, Liza DD 19.2, DD 35.3 Dorband, Moritz MP 4.1
Dörflinger, DavidT 20.8, T 48.7, T 48.9, T 48.10, T 99.5 Döring, Lydia
Dornbusch, Sven T 77.4 Dörner, Julien •T 47.5, T 73.3, •T 101.1, T 101.2
Dorrer, Holger T 102.6 Dorschu, Alexandra DD 34.11 Dorsel, Dominik DD 35.13, DD 41.2
Dort, Katharina 1 /0.2 Dostal, Petr
drago, antonino •AGPhil 1.3 Drews, Vincent DD 34.9 Dreyer, Sascha •T 63.1 Dreyler, Sascha •T 63.1
Dickeck, Günter T 05.0, T 05.1, T 94.4 Duckeck, Günter T 15.1 Duda, Dominik T 64.5, T 64.6, •T 82.1 Dudar Bohdan •T 44.1 T 44.2
Duehrssen, Michael T 70.9 Duenkel, Fabian T 101.6 Duhr, Claude T 1.1
Dujmovic, Hrvoje•T 100.7 Düllmann, ChristophT 102.6 Düren, MichaelT 3.9
Durrer, Ruth GR 6.3 Dutta, Juhi T 36.8 Dyar, Melinda Darby EP 3.1
Dybalski, Wojciech •MP 7.2 Dziwok, Christian T 65.1, T 65.2, T 65.3, •T 65.4, T 94.6
Eberlein, Nathalie •1 30.4 Eberwein, Gregor
•T 48.8, T 102.3 Ecker, Patrick
Edzards for the KATRIN collaboration, Frank
Edzards Frank T 23.3 T 75.2
Edzards, Frank

Engelke, Frederic	
Englert, Christoph 1 85.2	-
Enss, Christian 1 //.0, 1 102.0)
Erb Androop T104 E	;
Erb, Anureas))
EID, ROYEL . DD 10.3, DD 17.4, DD 24.2	-
T 6 9 T 25 4 T 24 1 T 24 5 T 59 2	,
T 60 2 T 70 1 T 70 7 T 70 8 T 107 8	
Erdmann Martin T 61 5 T 61 6	
T 61 7 T 90 4 T 90 5 T 107 4	,
Τ 107 5 ΔΚΡΙΚ 4 6	
Erdmenger Johanna MP 4 1 MP 4 2	,
Fren Engin •T 53 5 AKPIK 2 4	
AKPIK 4.2	'
Erhart, Andreas	
T 78.7. T 104.5. T 104.7	ʻ
Erhart, Max)
Ernst. Andrea	ŧ
Erpenbeck, Hannah . •T 72.9. T 103.4	Ċ.
T 106.1. T 106.2. T 106.3	í
Espe, Clemens EP 4.3, EP 4.4, EP 4.5	5
Espy, Patrick EP 5.4	ł
Essing, DominikDD 38.1	l
Eurin, Guillaume T 78.2	2
Evers, Mitja DD 46.3	3
Fabbri, Francesco Maria • GR 13.2	2
Fabian, Sven•T 31.7, T 31.8	3
Fackeldey, Peter •T 61.5, T 61.6, T 61.7	1
Fahn, Max Joseph . •GR 12.3, •MP 9.3	3
Fahrendholz, Ulrike T 20.8, •T 48.7	,
T 48.9, T 48.10, T 99.5	
Falaki, Jaouad El T 85.3	3
Falcke, Heino GR 1.6	j
Falconer, Isobel•AGPhil 2.1	l
Falconer, Kathleen •DD 22.1, DD 22.2	,
DD 24.3, DD 48.3	
Faltermann, Nils . T 7.5, T 87.4, T 87.5	5
Faltesek, Benjamin •AGPhil 11.5	5
Farkas, Ralf	5
Fattorini, Alicia•T 101.3	3
Febres Cordero, Fernando T 2.2	2
Feigl. Nora •T 52.3. T 52.4. •T 106.4	ŧ
Feindt, Finn	,
Feindt, Michael T 3.1	1
Feld. Lutz T 33.2. T 65.1. T 65.2	
T 65.3, T 65.4, T 94.6	
Feldmann, Marco EP 4.3, EP 4.4	,
Feldmann, Marco EP 4.3, EP 4.4 EP 4.5	,
Feldmann, Marco EP 4.3, EP 4.4 EP 4.5 Feldmann, Thorsten	;
Feldmann, MarcoEP 4.3, EP 4.4,EP 4.5Feldmann, ThorstenFeldmann, Sebastian	}
Feldmann, Marco EP 4.3, EP 4.4, EP 4.5 Feldmann, Thorsten	, , ,
Feldmann, Marco EP 4.3, EP 4.4, EP 4.5 Feldmann, Thorsten	, ,
Feldmann, Marco EP 4.3, EP 4.4, EP 4.5 Feldmann, Thorsten T 56.3 Felleisen, Sebastian •DD 36.2 Ferber, Torben T 15.7, T 63.1, T 63.3, T 63.4, T 63.5, T 76.9, T 79.2, T 79.3, T 79.4, T 79.5, T 84.3	, 32,
Feldmann, Marco EP 4.3, EP 4.4, EP 4.5 Feldmann, Thorsten •DD 36.2 Ferber, Torben •DD 36.2 Ferber, Torben •DD 36.2 T 63.4, T 63.5, T 76.9, T 79.2, T 79.3, T 79.4, T 79.5, T 84.3 Ferencz, Lars •T 88.3	, 32, 3
Feldmann, Marco EP 4.3, EP 4.4, EP 4.5 Feldmann, Thorsten T 56.3 Felleisen, Sebastian DD 36.2 Ferber, Torben T 15.7, T 63.1, T 63.3, T 63.4, T 63.5, T 76.9, T 79.2, T 79.3, T 79.4, T 79.5, T 84.3 Ferencz, Lars	, <u>32</u> , <u>33</u>
Feldmann, Marco EP 4.3, EP 4.4, EP 4.5 Feldmann, Thorsten T 56.3 Felleisen, Sebastian DD 36.2 Ferber, Torben T 15.7, T 63.1, T 63.3, T 63.4, T 63.5, T 76.9, T 79.2, T 79.3, T 79.4, T 79.5, T 84.3 Ferencz, Lars T 88.3 Fernández, Marcos ST 9.3 Fernández Pretel, José Antonio	, 32, 33
Feldmann, Marco EP 4.3, EP 4.4, EP 4.5 Feldmann, Thorsten T 56.3 Felleisen, Sebastian DD 36.2 Ferber, Torben T 15.7, T 63.1, T 63.3, T 63.4, T 63.5, T 76.9, T 79.2, T 79.3, T 79.4, T 79.5, T 84.3 Ferencz, Lars	, <u>32</u> , <u>33</u> , .
Feldmann, Marco	, 3 <u>2</u> , 3 <u>3</u> , 4
Feldmann, Marco	, <u>82</u> , <u>88</u> , <u>44</u> ,
Feldmann, Marco	, 3 <u>2</u> , 3 <u>3</u> , 4 <u>4</u> ,
Feldmann, Marco	, <u>82</u> , <u>81</u> , <u>81}, 81}, <u>81</u>, <u>81</u>,</u>
Feldmann, Marco	32, 33, 443, 443, 443, 443, 443, 443, 44
Feldmann, Marco	32, 33 ++3,5+2
Feldmann, Marco EP 4.3, EP 4.4, EP 4.5 Feldmann, Thorsten T 56.3 Felleisen, Sebastian DD 36.2 Ferber, Torben T 15.7, T 63.1, T 63.3, T 63.4, T 63.5, T 76.9, T 79.2, T 79.3, T 79.4, T 79.5, T 84.3 Ferencz, Lars T 88.3 Fernández, Marcos ST 9.3 Fernández Pretel, José Antonio •T 39.1 Ferrari, Anna •T 96.4 Ferrari, Anna •T 96.4 Ferreira, Pedro M. T 85.3 Ferreiro, Nahuel T 104.6 Fertl, Martin T 49.1, T 49.4, 4 Fesik, Liudmila •GR 5.2 Fichtner, Horst •EP 12.3, MP 4.4, 4	, <u>82</u> , <u>88</u> , <u>448</u> , <u>448}, 448, <u>448}, 448</u>, <u>448}, 448, <u>448}, 488}, 488}, 488}, 488}, 488}, 488}, 488}, 2888}, 2888}, 2888}, 2888}, 2</u></u></u>
Feldmann, Marco EP 4.3, EP 4.4, EP 4.5 Feldmann, Torsten T 56.3 Felleisen, Sebastian DD 36.2 Ferber, Torben T 15.7, T 63.1, T 63.3, T 63.4, T 63.5, T 76.9, T 79.2, T 79.3, T 79.4, T 79.5, T 84.3 Ferencz, Lars T 88.3 Fernández, Marcos ST 9.3 Fernández, Marcos ST 9.3 Fernández Pretel, José Antonio •T 39.1 Ferrari, Anna T 96.4 Ferrari, Anna T 89.3 Ferreira, Pedro M. T 85.3 Ferreiro, Nahuel T 104.6 Fertl, Martin -G 85.2 Fesik, Liudmila -G 85.2 Fichtner, Horst EP 12.3, MP 4.4 T 101.1 T 104.6	32, 33 H-13542, -
Feldmann, Marco EP 4.3, EP 4.4, EP 4.5 Feldmann, Torsten T 56.3 Felleisen, Sebastian •DD 36.2 Ferber, Torben T 15.7, T 63.1, T 63.3, T 63.4, T 63.5, T 76.9, T 79.2, T 79.3, T 79.4, T 79.5, T 84.3 Fernerz, Lars •T 88.3 Fernández, Marcos ST 9.3 Fernández, Marcos ST 9.3 Fernández Pretel, José Antonio •T 39.1 Ferrari, Arnaud T 89.4 Ferreira, Pedro M. T 85.3 Ferreiro, Nahuel T 104.6 Ferti, Martin •C 9.1, 23, MP 4.4, T 101.1 Fichtner, Horst •EP 12.3, MP 4.4, T 101.1 Figueira, Juan Manuel T 47.1	
Feldmann, Marco	
Feldmann, Marco	
Feldmann, Marco EP 4.3, EP 4.4, EP 4.5 Feldmann, Torsten T 56.3 Felleisen, Sebastian DD 36.2 Ferber, Torben T 15.7, T 63.1, T 63.3, T 63.4, T 63.5, T 76.9, T 79.2, T 79.3, T 79.4, T 79.5, T 84.3 Ferencz, Lars T 88.3 Fernández, Marcos ST 9.3 Fernández, Marcos ST 9.3 Ferrari, Anna T 96.4 Ferreiro, Nahuel T 104.6 Fertl, Martin	
Feldmann, Marco EP 4.3, EP 4.4, EP 4.5 Feldmann, Torsten T 56.3 Felleisen, Sebastian DD 36.2 Ferber, Torben T 15.7, T 63.1, T 63.3, T 63.4, T 63.5, T 76.9, T 79.2, T 79.3, T 79.4, T 79.5, T 84.3 Fernerz, Lars T 88.3 Fernández, Marcos ST 9.3 Fernández, Marcos ST 9.3 Ferrari, Anna T 96.4 Ferrari, Arnaud T 89.4 Ferreira, Pedro M. T 85.3 Ferreiro, Nahuel T 104.6 Ferti, Martin 49.1, T 49.4 Fersik, Liudmila 6R 5.2 Fichtner, Horst EP 12.3, MP 4.4, T 101.1 Figueira, Juan Manuel T 47.1 Finazzi, Lucas T 66.3 Fink, Christopher T 23.1, T 21.8 Fink, David T 21.1, T 21.8 Fink, Christopher T 45.3, T 93.3, T 93.4	
Feldmann, Marco EP 4.3, EP 4.4, EP 4.5 Feldmann, Torsten T 56.3 Felleisen, Sebastian •DD 36.2 Ferber, Torben T 15.7, T 63.1, T 63.3, T 63.4, T 63.5, T 76.9, T 79.2, T 79.3, T 79.4, T 79.5, T 84.3 Ferner, Lars •T 88.3 Fernández, Marcos ST 9.3 Fernández, Marcos ST 9.3 Ferrari, Anna •T 96.4 Ferrari, Anna •T 96.4 Ferrari, Anna •T 96.4 Ferrerai, Pedro M. T 85.3 Ferreiro, Nahuel T 104.6 Fertl, Martin T 49.1, T 49.4 Feisk, Liudmila •GR 5.2 Fichtner, Horst •EP 12.3, MP 4.4 T 101.1 Figueira, Juan Manuel T 47.1 Finazzi, Lucas •T 66.3 Fink, Christopher T 23.4 Fink, Christopher T 21.1, T 21.8 Finke, Thorben •T 45.3, T 9.3, T 9.	
Feldmann, Marco EP 4.3, EP 4.4, EP 4.5 Feldmann, Torsten T 56.3 Felleisen, Sebastian DD 36.2 Ferber, Torben T 15.7, T 63.1, T 63.3, T 63.4, T 63.5, T 76.9, T 79.2, T 79.3, T 79.4, T 79.5, T 84.3 Fernardez, Marcos ST 9.3 Fernández, Marcos ST 9.3 Fernández Pretel, José Antonio •T 39.1 Ferrari, Anna •T 96.4 Ferrari, Anna •T 96.4 Ferrari, Anna •T 96.4 Ferreiro, Nahuel T 104.6 Ferreiro, Nahuel T 104.6 Ferteiro, Nahuel T 49.1, T 49.4 Feik, Liudmila •GR 5.2 Fichtner, Horst •EP 12.3, MP 4.4, T 101.1 Finazzi, Lucas •T 66.3 Fink, Christopher T 23.4 Fink, Oxid T 21.1, T 21.8 Finke, Thorben •T 45.3, T 93.3, T 93.3, T 93.4 Finke, Thorben •T 45.3, T 9.3, T 9.4, T 21.8 Finke, Thorben •T 45.3, T 9.3, T 9.3, T 51.4 Finke, Thorben •T 45.3, T 9.3, T 9.3, T 51.4 Finke, Thorben •T 45.3, T 9.3, T 9.3, T 51.4 Finke, Thorben •T 45.3, T 9.3, T 9.3, T 51.4	
Feldmann, Marco EP 4.3, EP 4.4, EP 4.5 Feldmann, Torsten T 56.3 Felleisen, Sebastian DD 36.2 Ferber, Torben T 15.7, T 63.1, T 63.3, T 63.4, T 63.5, T 76.9, T 79.2, T 79.3, T 79.4, T 79.5, T 84.3 Ferner, Lars - T 88.3 Fernández, Marcos ST 9.3 Ferrari, Anna - T 96.4 Ferreiro, Nahuel T 104.6 Fertl, Martin - 49.4 Fersik, Ludmila - GR 5.2 Fichtner, Horst - EP 12.3, MP 4.4 T 101.1 Figueira, Juan Manuel T 47.1 Finazzi, Lucas - T 66.3 Fink, Christopher T 23.4 Finke, Thorben • T 45.3, T 93.3, T 93.4 Finke, Thorben • T 45.3, T 93.4 Finke, Thorben - T 45.3, T 93.4 Finke, Torben • T 45.3, T 93.4 Finke, Torben - T 45.3, T 93.4 Finke, Benjami	
Feldmann, Marco EP 4.3, EP 4.4, EP 4.5 Feldmann, Torsten T 56.3 Felleisen, Sebastian DD 36.2 Ferber, Torben T 15.7, T 63.1, T 63.3, T 63.4, T 63.5, T 76.9, T 79.2, T 79.3, T 79.4, T 79.5, T 84.3 Ferencz, Lars T 88.3 Fernández, Marcos ST 9.3 Fernández, Marcos ST 9.3 Ferrari, Anna T 96.4 Ferrari, Anna T 87.3 Ferrari, Anna T 87.3 Ferrari, Anna T 87.4 Fichtner, Horst T 66.3 Fink	
Feldmann, Marco EP 4.3, EP 4.4, EP 4.5 Feldmann, Torsten T 56.3 Felleisen, Sebastian \bullet D 36.2 Ferber, Torben T 15.7, T 63.1, T 63.3, T 63.4, T 63.5, T 76.9, T 79.2, T 79.3, T 79.4, T 79.5, T 84.3 Ferencz, Lars T 88.3 Fernández, Marcos ST 9.3 Fernández, Marcos ST 9.3 Ferrari, Arnau T 96.4 Ferrari, Anna T 96.4 Ferrari, Arnaud T 89.4 Ferreira, Pedro M. T 85.3 Ferreiro, Nahuel T 104.6 Fert, Martin T 49.1, T 49.4 Feisk, Liudmila •GR 5.2 Fichtner, Horst •EP 12.3, MP 4.4, T 101.1 Figueira, Juan Manuel T 47.1 Fink, Christopher T 23.4 Fink, Christopher T 21.1, T 21.8 Finke, Thorben •T 45.3, T 93.3, T 93.4 Finkelmeyer, Nelson •DD 35.10 Fiorini, Carlo T 21.1, T 21.8 Fischer, Benjamin T 61.5, T 61.6 •T 61.7 Fischer, Enno	
Feldmann, Marco EP 4.3, EP 4.4, EP 4.5 Feldmann, Thorsten T 56.3 Felleisen, Sebastian DD 36.2 Ferber, Torben T 15.7, T 63.1, T 63.3, T 63.4, T 63.5, T 76.9, T 79.2, T 79.3, T 79.4, T 79.5, T 84.3 Ferner, Lars - T 88.3 Fernández, Marcos ST 9.3 Fernández Pretel, José Antonio - T 39.1 Ferrari, Anna - T 96.4 Ferreira, Pedro M. T 85.3 Ferreiro, Nahuel - T 104.6 Fert, Martin - T 49.1, T 49.4 Feit, Liudmila - GR 5.2 Fichtner, Horst - EP 12.3, MP 4.4, T 101.1 Figueira, Juan Manuel - T 47.1 Finazzi, Lucas - T 66.3 Fink, Christopher T 23.4 Fink, David - T 21.1, T 21.8 Finke, Thorben - T 45.3, T 93.3, T 93.4 Finkelmeyer, Nelson - DD 35.10 Fiorini, Carlo - T 21.1, T 21.8 Fischer, Benjamin - T 61.5, T 61.6 - T 61.7 Fischer, Enno - AGPhil 7.3<	
Feldmann, Marco EP 4.3, EP 4.4, EP 4.5 Feldmann, Torsten T 56.3 Felleisen, Sebastian DD 36.2 Ferber, Torben T 15.7, T 63.1, T 63.3, T 63.4, T 63.5, T 76.9, T 79.2, T 79.3, T 79.4, T 79.5, T 84.3 Fernardez, Marcos ST 9.3 Fernández, Marcos ST 9.3 Fernández, Marcos ST 9.3 Ferrari, Anna T 96.4 Ferreiro, Nahuel T 104.6 Fertl, Martin T 49.1, T 49.4 Fersik, Ludmila GR 5.2 Fichtner, Horst EP 12.3, MP 4.4 T 101.1 Figueira, Juan Manuel T 47.1 Finazzi, Lucas T 66.3 Fink, Christopher T 23.4 Finke, Thorben T 45.3, T 93.3, T 93.4 Finke, Thorben T 45.3, T 93.4 Finke, Benjamin T 61.5, T 61.6 •T 61.7 Fischer, Enno AGPhil 7.3 Fischer, Enno AGPhil 7.3 Fischer, Enno	
Feldmann, Marco EP 4.3, EP 4.4, EP 4.5 Feldmann, Thorsten T 56.3 Felleisen, Sebastian DD 36.2 Ferber, Torben T 15.7, T 63.1, T 63.3, T 63.4, T 63.5, T 76.9, T 79.2, T 79.3, T 79.4, T 79.5, T 84.3 Ferencz, Lars T 88.3 Fernández, Marcos ST 9.3 Fernández, Marcos ST 9.3 Ferrari, Anna T 96.4 Ferrerio, Nahuel T 104.6 Fertl, Martin T 49.1, T 49.4 Fersi, Ludmila GR 52.2 Fichtner, Horst EP 12.3, MP 4.4 T 101.1 Figueira, Juan Manuel T 47.1 Finazzi, Lucas T 66.3 Fink, Christopher T 21.3, T 93.3, T 93.4 Finkke, Thorben •T 45.3, T 93.3, T 93.4 Finkke, Thorben •T 45.3, T 93.3, T 93.4 Finkke, Benjamin 61.5, T 61.6 •T 61.7 Fischer, Benjamin 61.5, T 61.6 •T 61.7 Fischer, Felix •T 93.4 Fischer, Felix	
Feldmann, Marco EP 4.3, EP 4.4, EP 4.5 Feldmann, Torsten T 56.3 Felleisen, Sebastian DD 36.2 Ferber, Torben T 15.7, T 63.1, T 63.3, T 63.4, T 63.5, T 76.9, T 79.2, T 79.3, T 79.4, T 79.5, T 84.3 Ferencz, Lars T 88.3 Ferendez, Marcos ST 9.3 Fernández, Marcos ST 9.3 Fernández, Marcos ST 9.3 Ferrari, Anna T 96.4 Ferrari, Arnaud T 89.4 Ferrari, Anna T 96.4 Ferrari, Arnaud T 89.4 Ferreiro, Nahuel T 104.6 Fert, Martin 49.1, T 49.4 Ferreiro, Nahuel T 104.6 Fetrik, Martin -G 8.5.2 Fichtner, Horst -EP 12.3, MP 4.4 T 101.1 Figueira, Juan Manuel T 47.1 Finazzi, Lucas -T 66.3 Finke, Thorben -T 45.3, T 93.3, T 93.4 Finke, Thorben -T 45.3, T 93.3, T 93.4 Finke, Thorben -T 45.3, T 93.3, T 93.4 Finke, Thorben -T 61.5, T 61.6 •T 61.7 Fischer, Benjamin -T 61.5, T 61.6 •T 61.7 Fischer, Enno	
Feldmann, Marco EP 4.3, EP 4.4, EP 4.5 Feldmann, Thorsten T 56.3 Felleisen, Sebastian DD 36.2 Ferber, Torben T 15.7, T 63.1, T 63.3, T 63.4, T 63.5, T 76.9, T 79.2, T 79.3, T 79.4, T 79.5, T 84.3 Fernact, Lars - T 88.3 Fernandez, Marcos ST 9.3 Fernández Pretel, José Antonio - T 39.1 Ferrari, Anna - T 96.4 Ferreira, Pedro M. T 85.3 Ferreiro, Nahuel - T 104.6 Fertl, Martin - T 49.1, T 49.4 Feitl, Liudmila - GR 5.2 Fichther, Horst - EP 12.3, MP 4.4, T 101.1 Figueira, Juan Manuel - T 47.1 Finazzi, Lucas - T 66.3 Fink, Christopher - T 23.4 Fink, Christopher - T 23.4 Fink, Christopher - T 23.4 Finkelmeyer, Nelson - DD 35.10 Fiorini, Carlo - T 21.1, T 21.8 Fischer, Benjamin	
Feldmann, Marco EP 4.3, EP 4.4, EP 4.5 Feldmann, Torsten T 56.3 Felleisen, Sebastian DD 36.2 Ferber, Torben T 15.7, T 63.1, T 63.3, T 63.4, T 63.5, T 76.9, T 79.2, T 79.3, T 79.4, T 79.5, T 84.3 Fernearc, Lars T 88.3 Fernández, Marcos ST 9.3 Ferrari, Anna T 96.4 Ferrari, Anna T 96.4 Ferrari, Anna T 96.4 Ferrerio, Nahuel T 104.6 Ferrerio, Nahuel T 104.6 Ferreiro, Nahuel T 49.1, T 49.4 Feisk, Ludmila GR 5.2 Fichtner, Horst EP 12.3, MP 4.4 T 101.1 Finazzi, Lucas T 66.3 Fink, Christopher T 23.4 Finke, Thorben •T 45.3, T 93.3, T 93.4 Finke, Thorben •T 45.3, T 93.4 Finke, Fels •D 35.10 Fiorini, Carlo T 21.1, T 21.8 Fischer, Enno ·AGPhil 7.3 Fischer, Enno ·AGPhil 7.3 Fisc	
Feldmann, Marco EP 4.3, EP 4.4, EP 4.5 Feldmann, Torsten T 56.3 Felleisen, Sebastian DD 36.2 Ferber, Torben T 15.7, T 63.1, T 63.3, T 63.4, T 63.5, T 76.9, T 79.2, T 79.3, T 79.4, T 79.5, T 84.3 Ferner, Lars T 88.3 Fernández, Marcos ST 9.3 Fernández, Marcos ST 9.3 Ferrari, Anna T 96.4 Ferrerio, Nahuel T 104.6 Fertl, Martin T 49.1, T 49.4 Fersi, Ludmila GR 52.2 Fichtner, Horst EP 12.3, MP 4.4 T 101.1 Figueira, Juan Manuel T 47.1 Finazzi, Lucas T 66.3 Fink, Christopher T 23.4 Finke, Thorben T 45.3, T 93.3, T 93.4 Finke, Thorben T 45.3, T 93.4 Finke, Benjamin T 61.5, T 61.6 • T 61.7 Fischer, Benjamin T 61.5, T 61.6 • T 61.7	
Feldmann, Marco EP 4.3, EP 4.4, EP 4.5 Feldmann, Torsten T 56.3 Felleisen, Sebastian DD 36.2 Ferber, Torben T 15.7, T 63.1, T 63.3, T 63.4, T 63.5, T 76.9, T 79.2, T 79.3, T 79.4, T 79.5, T 84.3 Ferencz, Lars T 88.3 Fernández, Marcos ST 9.3 Fernández, Marcos ST 9.3 Ferrari, Anna T 96.4 Ferra	
Feldmann, Marco EP 4.3, EP 4.4, EP 4.5 Feldmann, Torsten T 56.3 Felleisen, Sebastian DD 36.2 Ferber, Torben T 15.7, T 63.1, T 63.3, T 63.4, T 63.5, T 76.9, T 79.2, T 79.3, T 79.4, T 79.5, T 84.3 Ferencz, Lars T 88.3 Fernández, Marcos ST 9.3 Fernández, Marcos ST 9.3 Fernández, Marcos ST 9.3 Ferrari, Anna T 96.4 Ferreiro, Nahuel T 104.6 Fertl, Martin - 49.1, T 49.4 Ferreiro, Nahuel T 104.6 Fichtner, Horst - EP 12.3, MP 4.4 T 101.1 Figueira, Juan Manuel T 47.1 Finazzi, Lucas - T 66.3 Finke, Thorben - T 45.3, T 93.3, T 93.4 Finke, David T 21.1, T 21.8 Finke, Thorben - T 45.3, T 93.3, F 94.4 Finke, Thorben - T 61.5, T 61.6 Fiorini, Carlo - D 13.1 Fischer, Benjamin </td <td></td>	
Feldmann, Marco EP 4.3, EP 4.4, EP 4.5 Feldmann, Thorsten T 56.3 Felleisen, Sebastian DD 36.2 Ferber, Torben T 15.7, T 63.1, T 63.3, T 63.4, T 63.5, T 76.9, T 79.2, T 79.3, T 79.4, T 79.5, T 84.3 Ferner, Lars T 88.3 Fernández, Marcos ST 9.3 Ferrari, Anna T 96.4 Ferreira, Pedro M. T 85.3 Ferreiro, Nahuel T 104.6 Ferti, Martin T 49.1, T 49.4 Feitl, Liudmila •GR 5.2 Fichtner, Horst •EP 12.3, MP 4.4, T 101.1 Finazzi, Lucas T 66.3 Fink, Christopher T 23.4 Fink, David T 21.1, T 21.8 Finke, Thorben •T 45.3, T 93.3, T 93.4 Finke, Droben •T 45.3, T 93.3, T 93.4 Finke, Broben •D 35.10 Fiorini, Carlo T 21.1, T 21.8 Fischer, Enno •A GPhil 7.3 Fischer, Kenpamin T 61.5, T 61	
Feldmann, Marco EP 4.3, EP 4.4, EP 4.5 Feldmann, Torsten T 56.3 Felleisen, Sebastian DD 36.2 Ferber, Torben T 15.7, T 63.1, T 63.3, T 63.4, T 63.5, T 76.9, T 79.2, T 79.3, T 79.4, T 79.5, T 84.3 Fernardez, Marcos ST 9.3 Fernández, Marcos ST 9.3 Fernández, Marcos ST 9.3 Ferrari, Anna T 96.4 Ferreiro, Nahuel T 104.6 Ferreiro, Nahuel T 104.6 Fertl, Martin T 49.1, T 49.4 Fersiz, Lucas F 66.3 Fink, Christopher T 23.4 Fink, Christopher T 23.4 Fink, Christopher T 23.4 Finke, Thorben •T 45.3, T 93.3, T 93.4 Finke, Thorben •T 45.3, T 93.3, T 93.4 Finke, Thorben •T 45.3, T 93.4 Finke, Thorben •T 45.3, T 93.4 Finke, Tenno •A 64Phil 7.3 Fischer, Enno •A 64Phil 7.3 Fischer, Enno •A 64Phil 7.3 </td <td></td>	
Feldmann, Marco EP 4.3, EP 4.4, EP 4.5 Feldmann, Torsten T 56.3 Felleisen, Sebastian DD 36.2 Ferber, Torben T 15.7, T 63.1, T 63.3, T 63.4, T 63.5, T 76.9, T 79.2, T 79.3, T 79.4, T 79.5, T 84.3 Ferner, Lars T 88.3 Fernández, Marcos ST 9.3 Ferrari, Anna T 96.4 Ferreiro, Nahuel T 104.6 Fertl, Martin T 49.1, T 49.4 Fersiz, Ludmila GR 5.2 Fichtner, Horst EP 12.3, MP 4.4 T 101.1 Figueira, Juan Manuel T 47.1 Finazzi, Lucas T 66.3 Fink, Christopher T 23.4 Finke, Thorben T 45.3, T 93.3, T 93.4 Finke, Thorben T 45.3, T 93.4 Fischer, Benjamin T 61.5, T 61.6 • T 61.7 Fischer, Felix T 98.4 Fischer, Felix	
Feldmann, Marco EP 4.3, EP 4.4, EP 4.5 Feldmann, Thorsten T 56.3 Felleisen, Sebastian DD 36.2 Ferber, Torben T 15.7, T 63.1, T 63.3, T 63.4, T 63.5, T 76.9, T 79.2, T 79.3, T 79.4, T 79.5, T 84.3 Ferencz, Lars T 88.3 Fernández, Marcos ST 9.3 Fernández, Marcos ST 9.3 Ferrari, Anna T 96.4 Ferreiro, Nahuel T 104.6 Fertl, Martin T 49.1, T 49.4 Fersi, Ludmila GR 52.2 Fichtner, Horst EP 12.3, MP 4.4 T 101.1 Figueira, Juan Manuel T 47.1 Finazzi, Lucas T 66.3 Fink, Christopher T 23.3, T 93.4 Finke, Thorben I 45.3, T 93.3, P3.4 Fischer, Benjamin T 61.5, T 61.6 If 61.7 Fischer, Benjamin T 61.5, T 61.6 Fischer, Feli	
Feldmann, Marco EP 4.3, EP 4.4, EP 4.5 Feldmann, Torsten T 56.3 Felleisen, Sebastian DD 36.2 Ferber, Torben T 15.7, T 63.1, T 63.3, T 63.4, T 63.5, T 76.9, T 79.2, T 79.3, T 79.4, T 79.5, T 84.3 Ferencz, Lars T 88.3 Fernández, Marcos ST 9.3 Fernández, Marcos ST 9.3 Fernández, Marcos ST 9.3 Ferrari, Anna T 96.4 Ferreiro, Nahuel T 104.6 Fert, Martin	
Feldmann, Marco EP 4.3, EP 4.4, EP 4.5 Feldmann, Torsten T 56.3 Felleisen, Sebastian DD 36.2 Ferber, Torben T 15.7, T 63.1, T 63.3, T 63.4, T 63.5, T 76.9, T 79.2, T 79.3, T 79.4, T 79.5, T 84.3 Ferencz, Lars T 88.3 Fernández, Marcos ST 9.3 Ferrari, Anna T 96.4 Ferrari, Anna T 96.4 Ferrari, Anna T 96.4 Ferrari, Anna T 96.4 Ferreriro, Nahuel T 104.6 Ferreriro, Nahuel T 49.1, T 49.4 Feitl, Martin T 49.1, T 49.4 Figueira, Juan Manuel T 47.1 Finazzi, Lucas T 66.3 Fink, Christopher T 23.4 Finke, Thorben T 45.3, T 93.3, T 93.4 Finke, Thorben T 45.3, T 93.3, T 93.4 Finkelmeyer, Nelson • DD 35.10 Foirni, Carlo T 21.1, T 21.8 Fischer, Enno • AGPhil 7.3 Fischer, Enno • AGPhil 7.3 Fischer, Katthias • DD 2.1, • DD 3.1 Fischer, Max T 15.4 Fischer, Max T 15.3 Fischer, Max T 51.	
Feldmann, Marco EP 4.3, EP 4.4, EP 4.5 Feldmann, Torsten T 56.3 Felleisen, Sebastian DD 36.2 Ferber, Torben T 15.7, T 63.1, T 63.3, T 63.4, T 63.5, T 76.9, T 79.2, T 79.3, T 79.4, T 79.5, T 84.3 Fernardez, Marcos ST 9.3 Fernández, Marcos ST 9.3 Fernández, Marcos ST 9.3 Ferrari, Anna T 96.4 Ferreiro, Nahuel T 104.6 Fertl, Martin T 49.1, T 49.4 Fersik, Ludmila GR 5.2 Fichtner, Horst EP 12.3, MP 4.4 T 101.1 Figueira, Juan Manuel T 47.1 Finazzi, Lucas T 66.3 Fink, Christopher T 23.4 Finke, Thorben •T 45.3, T 93.3, T 93.4 Finke, Thorben •T 45.3, T 93.3, T 93.4 Finke, Thorben •T 45.3, T 93.4 Finke, Thorben •T 45.3, T 93.4 Finke, Thorben •T 45.3, T 93.4 Finke, Tenno •A 64.1, T 21.8 Fi	

Engelbrecht, Eugene EP 12.3

Froonz Markua	T 39.4
Frahm, Mathis	•T 61.8
Francke, Gero	. EP 4.3, EP 4.4, EP 4.5
Frank, Florian . Frank Torsten	•DD 17.3 T 104 3
Franke, Helena	DD 35.10
Franken, Julia	DD 5.3
Frauendiener, Jö	rg GR 2.2
Frediani, Nikolas	T 26.1
Freese, Mareike Freienhofer Mai	•DD 1/.4 T 106 1 T 106 2
T 106.3	
Freitag, Johanne	s AKPIK 1.3,
Fretwurst, Eckha	rt
Frey, Ariane	. T 84.9, T 95.1, T 95.4,
Frevermuth Olive	er T 64 8
Friedrich, René	.•MP 11.3, •AGPhil 3.3
Friedrich, Sarah	
Friege, Gunner	
Friend, Pia	EP 4.2
Fries, Pascal	MP 4.2
Fritzsche, Nick	T 41.8, •T 69.1
Fröb, Markus B.	•GR 3.2, •MP 4.7
Froch, Alexander	· •T 26.2, T 26.4
Fröse, Stefan	•AKPIK 4.5
Frost, Torben C.	•GR 1.2
Fuchs, Dominik	Т 104.6
Fühner, Larissa	•DD 13.3
Führing, Quentin	T 26.6, •T 60.8
Fujan, Kim Fuks Benjamin	
Funcke, Lena .	T 79.8, T 79.9
Funk, Stefan	
•T 46.7. T 72.4	. 1 18.8, 1 46.1, 1 46.3, . T 103.4
Fuss, Alexander	T 77.8
Gabici, Stefano	EP 6.6
gaede, frank	
Gagneur, Sophie	•T 97.2
Galison, Peter	•AGPhil 5.1
	1 • GR 9.3, GR 9.4
Ganster, Erik	•T 18.8, T 46.1, T 46.2,
T 46.3, T 46.7,	•T 18.8, T 46.1, T 46.2, T 72.4
T 46.3, T 46.7, Garai, Abhijit Garbers, Christol	. •T 18.8, T 46.1, T 46.2, T 72.4
T 46.3, T 46.7, Garai, Abhijit Garbers, Christo Garbiso Amano,	•T 18.8, T 46.1, T 46.2, T 72.4 T 104.6 phT 7.4 Markus MP 2.3
Ganster, Erik . T 46.3, T 46.7, Garai, Abhijit . Garbers, Christol Garbiso Amano, Gardikiotis, Anto Cartage Pobogog	. •T 18.8, T 46.1, T 46.2, T 72.4
Ganster, Erik T 46.3, T 46.7, Garai, Abhijit Garbers, Christoj Garbiso Amano, Gardikiotis, Anto Gartner, Rebecca Garutti, Erika	•T 18.8, T 46.1, T 46.2, T 72.4
Garaister, Erik T 46.3, T 46.7, Garai, Abhijit Garbiso Amano, Gardikiotis, Anto Gartner, Rebecca Garutti, Erika T 66.2, T 66.4,	•T 18.8, T 46.1, T 46.2, T 72.4
Ganster, Erik T 46.3, T 46.7, Garai, Abhijit Garbiso Amano, Gardikiotis, Anto Gartner, Rebecca Garutti, Erika T 66.2, T 66.4, gasparinetti, lucc Gasnarvan Sart	•T 18.8, T 46.1, T 46.2, T 72.4 T 104.6 phT 7.4 MarkusMP 2.3 nios•T 105.6 aT 65.8 T 12.8, T 12.9, T 66.1, T 66.5, T 66.8, •T 98.1 a•AGPhil 11.2 is FP 112 T 83.2
Garaister, Erik T 46.3, T 46.7, Garai, Abhijit Garbiso Amano, Gardikiotis, Anto Gartner, Rebecca Garutti, Erika T 66.2, T 66.4, gasparinetti, luca Gasparyan, Sarg Gastaldo, Loredd	•T 18.8, T 46.1, T 46.2, T 72.4 T 104.6 phT 7.4 MarkusMP 2.3 niosT 105.6 aT 63.8 T 12.8, T 12.9, T 66.1, T 66.5, T 66.8, •T 98.1 a
Ganster, Erik T 46.3, T 46.7, Garai, Abhijit Garbiso Amano, Gardikiotis, Anto Gartner, Rebeccc Garutti, Erika T 66.2, T 66.4, gasparinetti, luca Gasparyan, Sarg Gastaldo, Loreda T 102.6	•T 18.8, T 46.1, T 46.2, T 72.4
Ganster, Erik T 46.3, T 46.7, Garai, Abhijit Garbiso Amano, Gardikiotis, Anto Gartner, Rebeccc Garutti, Erika T 66.2, T 66.4, gasparinetti, luca Gasparyan, Sarg Gastaldo, Loreda T 102.6 Gastine, Thomas Gauda Kevin	•T 18.8, T 46.1, T 46.2, T 72.4
Ganster, Erik T 46.3, T 46.7, Garai, Abhijit Garbiso Amano, Gardikiotis, Anto Gartner, Rebecca Garutti, Erika T 66.2, T 66.4, gasparinetti, luca Gasparyan, Sarg Gastaldo, Loreda T 102.6 Gastine, Thomas Gauda, Kevin GeDet-Kollabora	•T 18.8, T 46.1, T 46.2, T 72.4
Ganster, Erik T 46.3, T 46.7, Garai, Abhijit Garbers, Christoj Garbiso Amano, Gardikiotis, Anto Gartner, Rebeccz Garutti, Erika T 66.2, T 66.4, gasparinetti, luca Gasparyan, Sarg Gastaldo, Loreda T 102.6 Gastine, Thomas Gauda, Kevin GeDet-Kollabora Geese, Anne	•T 18.8, T 46.1, T 46.2, T 72.4 T 72.4 Markus T 104.6 ph T 7.4 Markus MP 2.3 nios •T 105.6 a T 63.8 T 12.9, T 12.9, T 66.1, T 66.5, T 66.8, •T 98.1 a • AGPhil 11.2 is EP 11.2, T 83.2 ana T 77.6, T 102.5, C EP 3.4 T 21.4, T 21.5 tion T 13.1, T 13.2 • D 33.2 • T 21.4
Ganster, Erik T 46.3, T 46.7, Garai, Abhijit Garbiso Amano, Gardikiotis, Anto Gartner, Rebeccz Garutti, Erika T 66.2, T 66.4, gasparinetti, luca Gasparyan, Sarg Gastaldo, Loreda T 102.6 Gastine, Thomas Gauda, Kevin GeDet-Kollabora Geese, Anne Geesmann, Henr Gehlert, Janette	•T 18.8, T 46.1, T 46.2, T 72.4 T 72.4 Markus T 7, 4 Markus MP 2.3 nios •T 105.6 a T 63.8 T 12.8, T 12.9, T 66.1, T 66.5, T 66.8, •T 98.1 a AGPhil 11.2 is EP 11.2, T 83.2 ana T 77.6, T 102.5, s EP 3.4 T 21.4, T 21.5 tion T 13.1, T 13.2
Ganster, Erik T 46.3, T 46.7, Garai, Abhijit Garbiso Amano, Gardikiotis, Anto Gartner, Rebecca Garutti, Erika T 66.2, T 66.4, gasparinetti, luca Gasparyan, Sarg Gastaldo, Loreda T 102.6 Gastine, Thomas Gauda, Kevin GeDet-Kollabora Geese, Anne Geesmann, Henr Gehlert, Janette Geib, Karl-Heinz	•T 18.8, T 46.1, T 46.2, T 72.4 T 72.4 Markus T 104.6 ph T 7.4 Markus MP 2.3 nios •T 105.6 a T 105.6 a T 63.8 T 12.8, T 12.9, T 66.1, T 66.5, T 66.8, •T 98.1 a •AGPhil 11.2 is EP 11.2, T 83.2 ana T 77.6, T 102.5, S EP 3.4 T 21.4, T 21.5 tion T 13.1, T 13.2
Ganster, Erik T 46.3, T 46.7, Garai, Abhijit Garbiso Amano, Gardikiotis, Anto Gartner, Rebecco Garutti, Erika T 66.2, T 66.4, gasparinetti, luca Gasparyan, Sarg Gastaldo, Loreda T 102.6 Gastine, Thomas Gauda, Kevin GeDet-Kollabora Geese, Anne Geesemann, Henr Geigenberger, Ka Caibeurt, Manté	•T 18.8, T 46.1, T 46.2, T 72.4 T 104.6 ph
Ganster, Erik T 46.3, T 46.7, Garai, Abhijit Garbiso Amano, Gardikiotis, Anto Gartner, Rebecca Garutti, Erika T 66.2, T 66.4, gasparinetti, luca Gasparyan, Sarg Gastaldo, Loreda T 102.6 Gastine, Thomas Gauda, Kevin GeDet-Kollabora Geese, Anne Geese, Anne Geesemann, Henr Gehlert, Janette Geib, Karl-Heinz Geigenberger, Ka Geilhaupt, Manfr Geißhercht, Nin	•T 18.8, T 46.1, T 46.2, T 72.4 T 104.6 ph
Ganster, Erik T 46.3, T 46.7, Garai, Abhijit Garbers, Christoj Garbiso Amano, Gardikiotis, Anto Gartner, Rebecca Garutti, Erika T 66.2, T 66.4, gasparinetti, luca Gasparyan, Sarg Gastaldo, Loreda T 102.6 Gastine, Thomas Gauda, Kevin GeDet-Kollabora Geese, Anne Geese, Anne Geese, Anne Geigenberger, Ka Geilhaupt, Manfr Geißelbrecht, Niu Geiser, Achim	•T 18.8, T 46.1, T 46.2, T 72.4 T 104.6 ph
Ganster, Erik T 46.3, T 46.7, Garai, Abhijit Garbers, Christoj Garbiso Amano, Gardikiotis, Anto Gartner, Rebecca Garutti, Erika T 66.2, T 66.4, gasparinetti, luca Gasparyan, Sarg Gastaldo, Loreda T 102.6 Gastine, Thomas Gauda, Kevin GeDet-Kollabora Geese, Anne Geese, Anne Geese, Anne Geigenberger, Ka Geilhaupt, Manfr Geißelbrecht, Niu Geiser, Achim Geller, Cornelia	•T 18.8, T 46.1, T 46.2, T 72.4 T 72.4 T 104.6 ph
Ganster, Erik T 46.3, T 46.7, Garai, Abhijit Garbers, Christoj Garbiso Amano, Gardikiotis, Anto Gartner, Rebecca Garutti, Erika T 66.2, T 66.4, gasparinetti, luca Gasparyan, Sarg Gastaldo, Loreda T 102.6 Gastine, Thomas Gauda, Kevin GeDet-Kollabora Geese, Anne Geese, Anne Geese, Anne Geigenberger, Ka Geigenberger, Ka Geilhaupt, Manfr Geißelbrecht, Niù Geiser, Achim Geller, Cornelia •DD 50.3 Gemmecke, Just	.•T 18.8, T 46.1, T 46.2, T 72.4 T 72.4 Markus T 104.6 ph T 7.4 Markus MP 2.3 nios •T 105.6 a T 63.8, •T 105.6 a T 63.8, •T 98.1 a AGPhil 11.2 is EP 11.2, T 83.2 ana T 77.6, T 102.5, s EP 3.4 T 21.4, T 21.5 tion T 13.1, T 13.2 DD 33.2 ning ST 2.1 T 41.1, T 59.2, T 97.5 ed T 86.4 cole T 16.1, T 58.1 DD 36.4, DD 47.3, tine ST 8.2, ST 8.3,
Ganster, Erik . T 46.3, T 46.7, Garai, Abhijit . Garbiso Amano, Gardikiotis, Anto Garthiso Amano, Gardikiotis, Anto Gartner, Rebecca Garutti, Erika . T 66.2, T 66.4, gasparinetti, luca Gasparyan, Sarg Gastaldo, Loreda T 102.6 Gastine, Thomas Gauda, Kevin . GeDet-Kollabora Geese, Anne . Geesmann, Henr Gehlert, Janette Geib, Karl-Heinz Geigelbrecht, Nic Geißelbrecht, Nic Geiser, Achim Geller, Cornelia •DD 50.3 Gemmecke, Just ST 8.4	.•T 18.8, T 46.1, T 46.2, T 72.4
Ganster, Erik . T 46.3, T 46.7, Garai, Abhijit . Garbiso Amano, Gardikiotis, Anto Gartner, Rebecca Garutti, Erika . T 66.2, T 66.4, gasparinetti, lucz Gasparyan, Sarg Gastaldo, Loreda T 102.6 Gastine, Thomas Gauda, Kevin . GeDet-Kollabora Geese, Anne . Geesmann, Henr Gehlert, Janette Geib, Karl-Heinz Geigelbrecht, Nic Geilen, Cornelia •DD 50.3 Gemmecke, Just ST 8.4 Gentry, Brittany Genz, Florian	•T 18.8, T 46.1, T 46.2, T 72.4 T 72.4 T 104.6 ph T 7.4 Markus MP 2.3 nios •T 105.6 a
Ganster, Erik T 46.3, T 46.7, Garai, Abhijit Garbiso Amano, Gardikiotis, Anto Gartner, Rebecca Garutti, Erika T 66.2, T 66.4, gasparinetti, luca Gasparyan, Sarg Gastaldo, Loreda T 102.6 Gastine, Thomas Gauda, Kevin GeDet-Kollabora Geese, Anne Geesmann, Henr Gehlert, Janette Geib, Karl-Heinz Geigelbrecht, Nic Geilhaupt, Manfr Geißelbrecht, Nic Geilhaupt, Manfr Geißelbrecht, Nic Geilhaupt, Achim Geller, Cornelia •DD 50.3 Gemmecke, Just ST 8.4 Gentry, Brittany Geoffray, Emma	•T 18.8, T 46.1, T 46.2, T 72.4 T 72.4 T 104.6 ph T 7.4 Markus MP 2.3 nios •T 105.6 a T 12.8, T 12.9, T 66.1, T 162.8, T 12.9, T 66.1, T 66.5, T 66.8, •T 98.1 a •AGPhil 11.2 is EP 11.2, T 83.2 ana T 77.6, T 102.5, S EP 3.4
Ganster, Erik . T 46.3, T 46.7, Garai, Abhijit . Garbiso Amano, Gardikiotis, Anto Gartner, Rebecca Garutti, Erika . T 66.2, T 66.4, gasparinetti, luca Gasparyan, Sarg Gastaldo, Loreda T 102.6 Gastine, Thomas Gauda, Kevin . GeDet-Kollabora Geese, Anne . Geesmann, Henr Gehlert, Janette Geißelbrecht, Nic Geigelberger, Ka Geilhaupt, Manfr Geißelbrecht, Nic Geigelberger, Ka Geilhaupt, Manfr Geißelbrecht, Nic Geiser, Achim Geller, Cornelia •DD 50.3 Gemmecke, Just ST 8.4 Gentry, Brittany George, William Carber, Wilc	•T 18.8, T 46.1, T 46.2, T 72.4 T 72.4 T 104.6 ph T 7.4 Markus MP 2.3 nios •T 105.6 a T 12.9, T 66.1, T 162.8, T 12.9, T 66.1, T 66.5, T 66.8, •T 98.1 a •AGPhil 11.2 is EP 11.2, T 83.2 ana T 77.6, T 102.5, S EP 3.4
Ganster, Erik . T 46.3, T 46.7, Garai, Abhijit . Garbiso Amano, Gardikiotis, Anto Gartner, Rebecca Garutti, Erika . T 66.2, T 66.4, gasparinetti, luca Gasparyan, Sarg Gastaldo, Loreda T 102.6 Gastine, Thomas Gauda, Kevin . GeDet-Kollabora Geese, Anne . Geesmann, Henr Gehlert, Janette Geib, Karl-Heinz Geigelberger, Ka Geilhaupt, Manfr Geißelbrecht, Nic Geiser, Achim Geller, Cornelia •DD 50.3 Gemmecke, Just ST 8.4 Gentry, Brittany George, William Gerber, Nils GERDA-Kollabora	•T 18.8, T 46.1, T 46.2, T 72.4 T 72.4 T 104.6 ph T 7.4 Markus MP 2.3 nios •T 105.6 a T 12.9, T 66.1, T 12.8, T 12.9, T 66.1, T 66.5, T 66.8, •T 98.1 a •AGPhil 11.2 is EP 11.2, T 83.2 ana T 77.6, T 102.5, S EP 3.4
Ganster, Erik T 46.3, T 46.7, Garai, Abhijit Garbiso Amano, Gardikiotis, Anto Gartner, Rebecca Garutti, Erika T 66.2, T 66.4, gasparinetti, luca Gasparyan, Sarg Gastaldo, Loreda T 102.6 Gastine, Thomas Gauda, Kevin GeDet-Kollabora Geese, Anne Geese, Anne Geese, Anne Gelert, Janette Geißelbrecht, Nic Geißelbrecht, Nic Gernecke, Just ST 8.4 Gentry, Brittany George, William Geroge, William Geroge, William	•T 18.8, T 46.1, T 46.2, T 72.4 T 72.4 Markus T 104.6 ph T 7.4 Markus MP 2.3 nios •T 105.6 a T 12.9, T 66.1, T 66.5, T 66.8, •T 98.1 a •AGPhil 11.2 is EP 11.2, T 83.2 ana T 77.6, T 102.5, S EP 3.4 T 21.4, T 21.5 S EP 3.4 DD 33.2 T 41.1, T 59.2, T 97.5 S EP 4.5 S EP 3.4 DD 36.4, DD 47.3, tine •ST 8.2, ST 8.3, AGPhil 10.4 DD 46.1, •DD 48.3 T 53.9 ation T 75.3, T 75.5, T 87.8
Ganster, Erik T 46.3, T 46.7, Garai, Abhijit Garbers, Christoj Garbiso Amano, Gardikiotis, Anto Gartner, Rebeccc Garutti, Erika T 66.2, T 66.4, gasparinetti, luca Gasparyan, Sarg Gastaldo, Loredd T 102.6 Gastine, Thomas Gauda, Kevin GeDet-Kollabora Geese, Anne Geese, Anne Geese, Anne Geese, Anne Geese, Anne Geilert, Janette Geib, Karl-Heinz Geißelbrecht, Nic Geißelbrecht, Nic Gerse, Achim George, William Geroge, William Geroge, William Geroge, Tim	•T 18.8, T 46.1, T 46.2, T 72.4 T 72.4 Markus T 104.6 ph T 7.4 Markus MP 2.3 nios •T 105.6 a T 3.8, T 12.8, T 12.9, T 66.1, T 66.5, T 66.8, •T 98.1 a •AGPhil 11.2 isEP 11.2, T 83.2 ana T 77.6, T 102.5, SEP 3.4 T 21.4, T 21.5 SEP 3.4 DD 33.2 ning ST 2.1 DD 33.2 ningEP 6.5 tedT 16.1, T 58.1 DD 36.4, DD 47.3, tineT 8.2, ST 8.3, AGPhil 10.4 T 53.9 ation
Ganster, Erik T 46.3, T 46.7, Garai, Abhijit Garbers, Christoj Garbiso Amano, Gardikiotis, Anto Gartner, Rebecca Garutti, Erika T 66.2, T 66.4, gasparinetti, luca Gasparyan, Sarg Gastaldo, Loredd T 102.6 Gastine, Thomas Gauda, Kevin GeDet-Kollabora Geese, Anne Geesmann, Henr Gehlert, Janette Geißelbrecht, Nic Geißelbrecht, Nic Gerser, Achim Geller, Cornelia •DD 50.3 Germecke, Just ST 8.4 Gentry, Brittany George, William Geroge, Nils GERDA-Kollabora T 75.6, T 75.7, Gershon, Tim Gesualdi, Flavia	•T 18.8, T 46.1, T 46.2, T 72.4 T 72.4 T 104.6 ph T 7.4 Markus MP 2.3 nios •T 105.6 a
Ganster, Erik T 46.3, T 46.7, Garai, Abhijit Garbiso Amano, Gardikiotis, Anto Gartner, Rebecca Garutti, Erika T 66.2, T 66.4, gasparinetti, luca Gasparyan, Sarg Gastaldo, Loreda T 102.6 Gastine, Thomas Gauda, Kevin GeDet-Kollabora Geese, Anne Geesmann, Henr Gehlert, Janette Geißelbrecht, Nic Geißelbrecht, Nic Germecke, Just ST 8.4 Gentry, Brittany Genz, Florian Geoffray, Emma George, William Gerber, Nils GERDA-Kollabora T 75.6, T 75.7, Gerken, Louis Gershon, Tim Gesualdi, Flavia	•T 18.8, T 46.1, T 46.2, T 72.4 T 72.4 T 104.6 ph T 7.4 Markus MP 2.3 nios •T 105.6 a
Ganster, Erik . T 46.3, T 46.7, Garai, Abhijit . Garbiso Amano, Gardikiotis, Anto Gartner, Rebecca Garutti, Erika . T 66.2, T 66.4, gasparinetti, luca Gasparyan, Sarg Gastaldo, Loreda T 102.6 Gastine, Thomas Gauda, Kevin . GeDet-Kollabora Geese, Anne . Geesmann, Henr Gehlert, Janette Geißelbrecht, Nic Geißelbrecht, Nic Gerta, Elorian . Geoffray, Emma George, William Gerber, Nils GERDA-Kollabora T 75.6, T 75.7, Gerken, Louis . Gershon, Tim Gesualdi, Flavia Geyer, Felix Geyik, Marvin E.	•T 18.8, T 46.1, T 46.2, T 72.4 T 72.4 T 104.6 ph T 7.4 Markus MP 2.3 nios •T 105.6 a
Ganster, Erik T 46.3, T 46.7, Garai, Abhijit Garbiso Amano, Gardikiotis, Anto Gartner, Rebecca Garutti, Erika T 66.2, T 66.4, gasparinetti, luca Gasparyan, Sarg Gastaldo, Loreda T 102.6 Gastine, Thomas Gauda, Kevin GeDet-Kollabora Geese, Anne Geesmann, Henr Gehlert, Janette Geiß, Karl-Heinz Geigenberger, Ka Geilhaupt, Manfr Geißelbrecht, Nic Geiser, Achim Geller, Cornelia •DD 50.3 Gemmecke, Jusi ST 8.4 Gentry, Brittany Genz, Florian Geoffray, Emma Geoffray, Emma Geoffray, Emma Geoffray, Emma George, William Gerber, Nils Gershon, Tim Gesualdi, Flavia Geyer, Felix Geyik, Marvin En Ghanbari, Azade	•T 18.8, T 46.1, T 46.2, T 72.4 T 72.4 Markus T 104.6 ph T 7.4 Markus MP 2.3 nios •T 105.6 a
Ganster, Erik T 46.3, T 46.7, Garai, Abhijit Garbiso Amano, Gardikiotis, Anto Gartner, Rebecca Garutti, Erika T 66.2, T 66.4, gasparinetti, luca Gasparyan, Sarg Gastaldo, Loreda T 102.6 Gastine, Thomas Gauda, Kevin GeDet-Kollabora Geese, Anne Geese, Anne Geesmann, Henr Gehert, Janette Geib, Karl-Heinz Geigenberger, Ka Geilhaupt, Manfr Geißelbrecht, Nic Geiser, Achim Geller, Cornelia •DD 50.3 Gemmecke, Just ST 8.4 Gentry, Brittany Genz, Florian Geoffray, Emma George, William Gerber, Nils Gershon, Tim Gesualdi, Flavia Geyer, Felix Geyik, Marvin En Ghanbari, Azade Ghassemi, Novic Charbaria	•T 18.8, T 46.1, T 46.2, T 72.4 T 72.4 T 104.6 ph T 7.4 Markus MP 2.3 nios •T 105.6 a T 105.6 a T 12.8, T 12.9, T 66.1, T 66.5, T 66.8, •T 98.1 a •AGPhil 11.2 is EP 11.2, T 83.2 ana T 77.6, T 102.5, S EP 3.4 T 21.4, T 21.5 c EP 3.4 T 41.1, T 59.2, T 97.5 c ed T 82.2, ST 8.3, C AGPhil 10.4 C AGPhil 10.4 C AGPhil 10.4 C AGPhil 7.4, T 60.5 T 52.2 C T 60.6 T 56.2 C T 51.7 C AKPIK 4.3 C T 60.6 T 56.2 C T 51.7 C AKPIK 4.3 C D 13.2 C AGP 11.2 C AGP 12.2 C

Ghosh Saranya Samik T 6/ 1
Ghoshdastidar Debardhya T 26.8
Giacinti. Gwenaël
Giakoustidis, Georgios •T 95.1,
T 95.3, T 95.7
Giebenhain, Kim Tabea •ST 9.1
Giefers, Hubertus
Giese, Albrecht•GR 10.1
Giese, H EP 5.2
Giese, Hanna
Glesel, Kristina •GR 12.1, GR 12.3,
•IVIP 9.1, IVIP 9.3, IVIP 10.2 Ciffold Manual T 15.2 T 15.4 T 15.6
T 15 7
Gilles. Geoffrey
Giraldi. Angela•T 35.4
Glas, Cynthia•EP 6.4
Glaser, Fabian
Glatz, Lion Cornelius •DD 16.3
Gligorov, Vladimir
Glombitza, Jonas T 107.5, •AKPIK 4.6
Glück, Ferenc
Gocke, Benedikt 16.2, 134.5, •170.1
Goertz Elorian T 216 T 217 T 219
T 92 5
Goertz, Simon •DD 10 6 DD 12 8
Goettel, Alexandre T 20.2, T 20.5
Goettlicher, Peter
Goetz, Stefanie
Goksu, Hazal•T 45.6
Goldenzweig, Pablo T 3.1, T 63.3,
T 63.4, T 63.5, T 76.9, T 79.2, T 79.3,
I /9.5, I 84.3
Golunom, Laura
Golmar Federico T 66 3
Gomes Gabriel T 26.5 •T 60.5
Gonzalez, Alejandra
Goriely, Stephane EP 1.4
Goswami, Abhishek•MP 12.3
Goth, Florian MP 4.2
Götschel, Sebastian T 53.8
Gottardo, Carlo T 60.5
Gottel, Alexandre•I 20.1, I 43.1,
1 48.4, 1 103.7, 1 103.9
Cottmann Artur T 15 3
Gottmann, Artur
Gottmann, Artur T 15.3 Gottowik, Marvin •T 19.7 Göttsche, Malte •T 102.8
Gottmann, Artur
Gottmann, Artur
Gottmann, Artur T 15.3 Gottowik, Marvin • T 19.7 Göttsche, Malte T 102.8 Gottschlich, Benedikt DD 19.2, •DD 34.2 Götz, Stefanie
Gottmann, Artur

Gruber, Markus I II.8, •I I4.2, I 39.8, T 68 1	
Grünauer. Gina T 20.3. T 20.9. T 48.6.	
T 48.8, T 102.3	
Grunwald, Cornelius	
Gschwind, Stephane•DD 45.2	
Gubernari. Nico	
Guffanti, Daniele T 59.5, T 102.2	
Gugel, Ralf T 90.7	
Gugiatti, Matteo	
Guglieimi, valentina	
Günther. Christoph T 98.7. •T 106.9	
Günther, Maximilian N • EP 2.7	
Gürbuz, Saime . T 39.8, •T 68.1, T 97.1	
Gurdasani, Simran•T 76.5	
Gussmann, MarcDD 5.2 Güth Andreas T 4.2	
Guth, Manuel	
Gutjahr, Pascal EP 11.8, •T 46.6,	
T 83.8	
Gutsche, Christian	
H P Pernice Wolfram DD 37 2	
Ha Minh, Martin	
Haack, ChristianT 46.9	
Haagen-Schützenhöfer, Claudia	
•DD 4.3, DD 8.1, DD 11.4, DD 11.5, DD 10.2, DD 35.3, DD 30.3, DD 43.3	
Haase, Sebastian •DD 48.1, DD 48.2	
Habbaba, Moritz•T 88.5	
Habedank, Martin•T 64.4	
Hackfeld, Jonas•T 17.4	
Hackmann, EvaGR 2.6, GR 2.8	
Hadef, Asma . •T 16.9. T 34.2. T 34.3.	
T 35.2, T 35.3, T 41.1, T 59.2	
Hagemann, Felix•T 13.1	
Hagenbuchner, MarkusST 3.4	
Hagmann, Jonannes-Geert •GP 5.2 Hagner Caren T 102 1	
Hahn. Jan Joachim	
Hahn, Larissa DD 25.2, •DD 34.1,	
DD 34.9	
Hahn Lotte •DD 35.8	- 1
Liehn Steffen AKDIK 4.0 AKDIK 4.0	
Hahn, Steffen AKPIK 4.8, AKPIK 4.9 Haide Isabel T 79.3	
Hahn, Steffen . •AKPIK 4.8, AKPIK 4.9 Haide, Isabel	
Hahn, Steffen . •AKPIK 4.8, AKPIK 4.9 Haide, Isabel	
Hahn, Steffen •AKPIK 4.8, AKPIK 4.9 Haide, Isabel •T 79.3 Hajheidari, Mohammadtaghi •T 12.9 Hajian, Kamal •GR 1.5 Hakenmüller, Janina T 23.1, •T 74.4	
Hahn, Steffen •AKPIK 4.8, AKPIK 4.9 Haide, Isabel •T 79.3 Hajheidari, Mohammadtaghi •T 12.9 Hajian, Kamal •GR 1.5 Hakenmüller, Janina T 23.1, •T 74.4 Haleem, Mahsana ······ T 88.4	
Hahn, Steffen •AKPIK 4.8, AKPIK 4.9 Haide, Isabel •T 79.3 Hajheidari, Mohammadtaghi •T 12.9 Hajian, Kamal •GR 1.5 Hakenmüller, Janina T 23.1, •T 74.4 Haleem, Mahsana T 88.4 Hall, Zachary •AGPhil 6.3 Haller, Johannes T 6.3 T 7 2 T 10.4	
Hahn, Steffen •AKPIK 4.8, AKPIK 4.9 Haide, Isabel •T 79.3 Hajheidari, Mohammadtaghi •T 12.9 Hajian, Kamal •GR 1.5 Hakenmüller, Janina T 23.1, •T 74.4 Haleem, Mahsana T 88.4 Hall, Zachary •AGPhil 6.3 Haller, Johannes T 6.3, T 7.2, T 10.4, T 10.5, T 10.8, T 61.8, T 62.6, T 62.7,	
Hahn, Steffen •AKPIK 4.8, AKPIK 4.9 Haide, Isabel •T 79.3 Hajheidari, Mohammadtaghi •T 12.9 Hajian, Kamal •GR 1.5 Hakenmüller, Janina T 23.1, •T 74.4 Haleem, Mahsana T 88.4 Hall, Zachary •AGPhil 6.3 Haller, Johannes T 6.3, T 7.2, T 10.4, T 10.5, T 10.8, T 61.8, T 62.6, T 62.7, T 69.5	
Hahn, Steffen •AKPIK 4.8, AKPIK 4.9 Haide, Isabel •T 79.3 Hajheidari, Mohammadtaghi •T 12.9 Hajian, Kamal •GR 1.5 Hakenmüller, Janina T 23.1, •T 74.4 Haleem, Mahsana T 88.4 Hall, Zachary •AGPhil 6.3 Haller, Johannes T 6.3, T 7.2, T 10.4, T 10.5, T 10.8, T 61.8, T 62.6, T 62.7, T 69.5 Hallford, John •T 44.4	
Hahn, Steffen •AKPIK 4.8, AKPIK 4.9 Haide, Isabel •T 79.3 Hajheidari, Mohammadtaghi •T 12.9 Hajian, Kamal •GR 1.5 Hakenmüller, Janina T 23.1, •T 74.4 Haleem, Mahsana T 88.4 Hall, Zachary •AGPhil 6.3 Haller, Johannes T 6.3, T 7.2, T 10.4, T 10.5, T 10.8, T 61.8, T 62.6, T 62.7, T 69.5 Hallford, John •T 44.4 Halve, Lasse •T 106.1, T 106.2, T 106.2, T 106.2, T 106.2	
Hahn, Steffen •AKPIK 4.8, AKPIK 4.9 Haide, Isabel •T 79.3 Hajheidari, Mohammadtaghi •T 12.9 Hajian, Kamal •GR 1.5 Hakenmüller, Janina T 23.1, •T 74.4 Haleem, Mahsana T 88.4 Hall, Zachary •AGPhil 6.3 Haller, Johannes T 6.3, T 7.2, T 10.4, T 10.5, T 10.8, T 61.8, T 62.6, T 62.7, T 69.5 Hallford, John •T 44.4 Halve, Lasse •T 106.1, T 106.2, T 106.3, T 106.3 Hamacher-Baumann, Philip T 39.5	
Hahn, Steffen •AKPIK 4.8, AKPIK 4.9 Haide, Isabel •T 79.3 Hajheidari, Mohammadtaghi •T 12.9 Hajian, Kamal •GR 1.5 Hakenmüller, Janina •T 23.1, •T 74.4 Haleem, Mahsana T 88.4 Hall, Zachary •AGPhil 6.3 Haller, Johannes •T 6.3, T 7.2, T 10.4, T 10.5, T 10.8, T 61.8, T 62.6, T 62.7, T 69.5 Hallford, John •T 44.4 Halve, Lasse •T 106.1, T 106.2, T 106.3, T 106.3, T 39.5, T 39.6, •T 39.7	
Hahn, Steffen •AKPIK 4.8, AKPIK 4.9 Haide, Isabel •T 79.3 Hajheidari, Mohammadtaghi •T 12.9 Hajian, Kamal •GR 1.5 Hakenmüller, Janina •T 23.1, •T 74.4 Haleem, Mahsana T 88.4 Hall, Zachary •AGPhil 6.3 Haller, Johannes •T 6.3, T 7.2, T 10.4, T 10.5, T 10.8, T 61.8, T 62.6, T 62.7, T 69.5 Hallford, John •T 44.4 Halve, Lasse •T 106.1, T 106.2, T 106.3, T 39.5, T 39.6, •T 39.7 Hamacher-Baumann, Philip T 39.5, T 39.7 Hamer, Matthias T 64.8, T 97.6	
Hahn, Steffen •AKPIK 4.8, AKPIK 4.9 Haide, Isabel •T 79.3 Hajheidari, Mohammadtaghi •T 12.9 Hajian, Kamal •GR 1.5 Hakenmüller, Janina •CR 1.5 Haleem, Mahsana •SR 3.5 Halle, Johannes •AGPhil 6.3 Haller, Johannes •T 6.3, T 7.2, T 10.4, T 10.5, T 10.8, T 61.8, T 62.6, T 62.7, T 69.5 Hallford, John •T 44.4 Halve, Lasse •T 106.1, T 106.2, T 106.3 Hamacher-Baumann, Philip T 39.5, T 39.6, •T 39.7 Hamer, Matthias ·T 64.8, T 97.6 Hamilton, Fiona ·T 04.5, •T 104.6	
Hahn, Steffen •AKPIK 4.8, AKPIK 4.9 Haide, Isabel •T 79.3 Hajheidari, Mohammadtaghi •T 12.9 Hajian, Kamal •GR 1.5 Hakenmüller, Janina •CR 1.5 Haleem, Mahsana •SR 1.5 Haller, Johannes •AGPhil 6.3 Haller, Johannes •T 6.3, T 7.2, T 10.4, T 10.5, T 10.8, T 61.8, T 62.6, T 62.7, T 69.5 Hallford, John •T 44.4 Halve, Lasse •T 106.1, T 106.2, T 106.3 Hamacher-Baumann, Philip T 39.5, T 39.6, •T 39.7 Hammann, Robert T 22.5, T 102.6 Hamilton, Fiona T 104.5, •T 104.6 Hammann, Robert T 22.5, T 102.5	
Hahn, Steffen •AKPIK 4.8, AKPIK 4.9 Haide, Isabel •T 79.3 Hajheidari, Mohammadtaghi •T 12.9 Hajian, Kamal •GR 1.5 Hakenmüller, Janina .T 23.1, •T 74.4 Haleem, Mahsana T 88.4 Hall, Zachary •AGPhil 6.3 Haller, Johannes .T 6.3, T 7.2, T 10.4, T 10.5, T 10.8, T 61.8, T 62.6, T 62.7, T 69.5 Hallford, John •T 44.4 Halve, Lasse •T 106.1, T 106.2, T 106.3 Hamacher-Baumann, Philip .T 39.5, T 39.6, •T 39.7 Hammann, Robert T 24.5, T 104.6 Hammann, Robert T 22.5, T 102.5 Hammerich, Jan	
Hahn, Steffen •AKPIK 4.8, AKPIK 4.9 Haide, Isabel •T 79.3 Hajheidari, Mohammadtaghi •T 12.9 Hajian, Kamal •GR 1.5 Hakenmüller, Janina T 23.1, •T 74.4 Haleem, Mahsana T 88.4 Hall, Zachary •AGPhil 6.3 Haller, Johannes T 6.3, T 7.2, T 10.4, T 10.5, T 10.8, T 61.8, T 62.6, T 62.7, T 69.5 Hallford, John •T 44.4 Halve, Lasse •T 106.1, T 106.2, T 106.3 Hamacher-Baumann, Philip T 39.5, T 39.6, •T 39.7 Hamer, Matthias T 64.8, T 97.6 Hamilton, Fiona T 104.5, •T 104.6 Hammann, Robert •T 22.5, T 102.5 Hammerich, Jan •T 13.4 Han Thanh, Linh •EP 6.8	
Hahn, Steffen •AKPIK 4.8, AKPIK 4.9 Haide, Isabel •T 79.3 Hajheidari, Mohammadtaghi •T 12.9 Hajian, Kamal •GR 1.5 Hakenmüller, Janina .T 23.1, •T 74.4 Haleem, Mahsana T 88.4 Hall, Zachary •AGPhil 6.3 Haller, Johannes T 6.3, T 7.2, T 10.4, T 10.5, T 10.8, T 61.8, T 62.6, T 62.7, T 69.5 Hallford, John •T 44.4 Halve, Lasse •T 106.1, T 106.2, T 106.3 Hamacher-Baumann, Philip T 39.5, T 39.6, •T 39.7 Hamer, Matthias T 64.8, T 97.6 Hamiton, Fiona T 104.5, •T 104.6 Hammer, Robert Hanmer, Eleen •D 45.1 Hammerich, Jan Han Thanh, Linh •EP 6.8 Han, Xiaoxue T 96.6	
Hahn, Steffen •AKPIK 4.8, AKPIK 4.9 Haide, Isabel •T 79.3 Hajheidari, Mohammadtaghi •T 12.9 Hajian, Kamal •GR 1.5 Hakenmüller, Janina T 23.1, •T 74.4 Haleem, Mahsana T 88.4 Hall, Zachary •AGPhil 6.3 Haller, Johannes T 6.3, T 7.2, T 10.4, T 10.5, T 10.8, T 61.8, T 62.6, T 62.7, T 69.5 Hallford, John •T 44.4 Halve, Lasse •T 106.1, T 106.2, T 106.3 T 39.5, Hamacher-Baumann, Philip T 39.5, T 39.6, •T 39.7 Hamer, Matthias	
Hahn, Steffen •AKPIK 4.8, AKPIK 4.9 Haide, Isabel •T 79.3 Hajheidari, Mohammadtaghi •T 12.9 Hajian, Kamal •GR 1.5 Hakenmüller, Janina T 23.1, •T 74.4 Haleem, Mahsana T 88.4 Hall, Zachary •AGPhil 6.3 Haller, Johannes T 6.3, T 7.2, T 10.4, T 10.5, T 10.8, T 61.8, T 62.6, T 62.7, T 69.5 Hallford, John •T 44.4 Halve, Lasse •T 106.1, T 106.2, T 106.3 T 39.5, Hamacher-Baumann, Philip T 39.5, T 39.6, •T 39.7 Hamer, Matthias T 64.8, T 97.6 Hammern, Fienen •DD 45.1 Hammerich, Jan •T 13.4 Han Thanh, Linh •EP 6.8 Han, Xiaoxue T 96.6 Hance, Jonte R •AGPhil 9.4 Hannen, Ines T 39.5	
Hahn, Steffen •AKPIK 4.8, AKPIK 4.9 Haide, Isabel •T 79.3 Hajheidari, Mohammadtaghi •T 12.9 Hajian, Kamal •GR 1.5 Hakenmüller, Janina T 23.1, •T 74.4 Haleem, Mahsana T 88.4 Hall, Zachary •AGPhil 6.3 Haller, Johannes T 63.7 7.2, T 10.4, T 10.5, T 10.8, T 61.8, T 62.6, T 62.7, T 69.5 Hallford, John •T 44.4 Halve, Lasse •T 106.1, T 106.2, T 106.3 T 39.5, Hamacher-Baumann, Philip T 39.5, T 39.6, •T 39.7 Hamer, Matthias •T 64.8, T 97.6 Hammerich, Jan •T 13.4 Hammerich, Jan •T 13.4 Han Thanh, Linh •EP 6.8 Han, Xiaoxue T 96.6 Hance, Jonte R •AGPhil 9.4 Hannen, Ines T 39.5, T 39.5, T 39.5 Hannen, Ker •AGPhil 9.4 Hannen, Ker •AGPhil 9.4	
Hahn, Steffen •AKPIK 4.8, AKPIK 4.9 Haide, Isabel •T 79.3 Hajheidari, Mohammadtaghi •T 12.9 Hajian, Kamal •GR 1.5 Hakenmüller, Janina T 23.1, •T 74.4 Haleem, Mahsana T 88.4 Hall, Zachary •AGPhil 6.3 Haller, Johannes T 6.3, T 7.2, T 10.4, T 10.5, T 10.8, T 61.8, T 62.6, T 62.7, T 69.5 Hallford, John •T 44.4 Halve, Lasse •T 106.1, T 106.2, T 106.3 T 39.7 Hamer, Matthias T 64.8, T 97.6 Hamilton, Fiona T 104.5, •T 104.6 Hammerich, Jan •T 13.4 Han Thanh, Linh •EP 6.8 Han, Xiaoxue T 96.6 Hance, Jonte R •AGPhil 9.4 Hannen, Ines T 39.5 Hannen, Volker T 21.4, T 21.5 Hannen, Volker T 39.5, T 39.5 Hannen, Volker T 22.5, T 102.5 Hammerich, Jan •T 13.4 Han Thanh, Linh •EP 6.8 Hannen, Sobert T 39.5 Hannen, Volker T 39.5 Hannen, Sobert	
Hahn, Steffen •AKPIK 4.8, AKPIK 4.9 Haide, Isabel •T 79.3 Hajheidari, Mohammadtaghi •T 12.9 Hajian, Kamal •GR 1.5 Hakenmüller, Janina T 23.1, •T 74.4 Haleem, Mahsana T 88.4 Hall, Zachary •AGPhil 6.3 Haller, Johannes T 6.3, T 7.2, T 10.4, T 10.5, T 10.8, T 61.8, T 62.6, T 62.7, T 69.5 Hallford, John •T 44.4 Halve, Lasse •T 106.1, T 106.2, T 106.3 T 39.7 Hamer, Matthias T 64.8, T 97.6 Hamilton, Fiona T 104.5, •T 104.6 Hammerich, Jan •T 13.4 Han Thanh, Linh •EP 6.8 Han, Xiaoxue T 96.6 Hance, Jonte R •AGPhil 9.4 Hannen, Ines T 39.5 T 39.5 T 39.5 Hannen, Volker T 21.4, T 21.5 Hannen, Volker T 21.4, T 21.5 Hannen, Volker T 50.5 Hansmann-Menzemer, Stephanie •PV VI, T 50.4, T 50.5	
Hahn, Steffen •AKPIK 4.8, AKPIK 4.9 Haide, Isabel •T 79.3 Hajheidari, Mohammadtaghi •T 12.9 Hajian, Kamal •GR 1.5 Hakenmüller, Janina T 23.1, •T 74.4 Haleem, Mahsana T 88.4 Hall, Zachary •AGPhil 6.3 Haller, Johannes T 6.3, T 7.2, T 10.4, T 10.5, T 10.8, T 61.8, T 62.6, T 62.7, T 69.5 Hallford, John •T 44.4 Have, Lasse •T 106.1, T 106.2, T 106.3 Hamacher-Baumann, Philip T 39.5, T 39.6, •T 39.7 Hamer, Matthias T 64.8, T 97.6 Hamilton, Fiona .T 104.5, •T 104.6 Hammer, Keleen •DD 45.1 Hammerich, Jan •T 13.4 Han, Xiaoxue T 96.6 Hance, Jonte R •AGPhil 9.4 Hannen, Ines T 39.5, Hannen, Volker T 21.4, T 21.5 Hansch, Walter ST 9.5, T 66.4 Hansmann-Menzemer, Stephanie •PV VI, T 50.4, T 50.5 Hardy Edma DD 5 2	
Hahn, Steffen •AKPIK 4.8, AKPIK 4.9 Haide, Isabel •T 79.3 Hajheidari, Mohammadtaghi •T 12.9 Hajian, Kamal •GR 1.5 Hakenmüller, Janina T 23.1, •T 74.4 Haleem, Mahsana T 88.4 Hall, Zachary •AGPhil 6.3 Haller, Johannes T 6.3, T 7.2, T 10.4, T 10.5, T 10.8, T 61.8, T 62.6, T 62.7, T 69.5 Hallford, John •T 44.4 Have, Lasse •T 106.1, T 106.2, T 106.3 T 39.7 Hamer, Matthias T 64.8, T 97.6 Hamilton, Fiona T 104.5, •T 104.6 Hammer, Keleen •DD 45.1 Hammerich, Jan •T 13.4 Han Thanh, Linh •EP 6.8 Hance, Jonte R •AGPhil 9.4 Hannen, Ines T 39.5, T 39.5 T 39.5 Hannen, Volker T 21.4, T 21.5 Hansch, Walter ST 9.5, T 66.4 Hansmann-Menzemer, Stephanie •PV VI, T 50.4, T 50.5 Harder, Sven T 53.8 Hardy, Edme D 5.2	
Hahn, Steffen •AKPIK 4.8, AKPIK 4.9 Haide, Isabel •T 79.3 Hajheidari, Mohammadtaghi •T 12.9 Hajian, Kamal •GR 1.5 Hakenmüller, Janina T 23.1, •T 74.4 Haleem, Mahsana T 88.4 Hall, Zachary •AGPhil 6.3 Haller, Johannes T 6.3, T 7.2, T 10.4, T 10.5, T 10.8, T 61.8, T 62.6, T 62.7, T 69.5 Hallford, John •T 44.4 Have, Lasse •T 106.1, T 106.2, T 106.3 Hamacher-Baumann, Philip T 39.5, T 39.6, •T 39.7 Hamer, Matthias T 64.8, T 97.6 Hamilton, Fiona T 104.5, •T 104.6 Hammer, Keleen •DD 45.1 Hammerich, Jan •T 13.4 Han, Xiaoxue T 96.6 Hance, Jonte R 4GPhil 9.4 Hannen, Ines T 39.5, T 66.4 Hannen, Volker T 21.4, T 21.5 Hansch, Walter ST 9.5, T 66.4 Hansch, Walter ST 9.5, T 66.4 Hander, Sven T 53.8 Hardy, Edme D 5.2 Harder, Sven T 53.8 Hardy, Edme	
Hahn, Steffen •AKPIK 4.8, AKPIK 4.9 Haide, Isabel •T 79.3 Hajheidari, Mohammadtaghi •T 12.9 Hajian, Kamal •GR 1.5 Hakenmüller, Janina T 23.1, •T 74.4 Haleem, Mahsana T 88.4 Hall, Zachary •AGPhil 6.3 Haller, Johannes T 6.3, T 7.2, T 10.4, T 10.5, T 10.8, T 61.8, T 62.6, T 62.7, T 69.5 Hallford, John •T 44.4 Have, Lasse •T 106.1, T 106.2, T 106.3 Hamacher-Baumann, Philip T 39.5, T 39.6, •T 39.7 Hamer, Matthias T 64.8, T 97.6 Hamilton, Fiona T 104.5, •T 104.6 Hammer, Keleen •DD 45.1 Hammerich, Jan •T 13.4 Han Thanh, Linh •EP 6.8 Hance, Jonte R 4GPhil 9.4 Hannen, Ines T 39.5, T 66.4 Hansch, Walter ST 9.5, T 66.4 Hansen, Welter ST 9.5, T 66.4 Hansch, Walter ST 9.5, T 66.4 Hander, Robert V. T 31.1 Harlow, Daniel •SVNS 1.3, •MP 2.1 Harlander, Robert V. T 31.1	
Hahn, Steffen •AKPIK 4.8, AKPIK 4.9 Haide, Isabel •T 79.3 Hajheidari, Mohammadtaghi •T 12.9 Hajian, Kamal •GR 1.5 Hakenmüller, Janina T 23.1, •T 74.4 Haleem, Mahsana T 88.4 Hall, Zachary •AGPhil 6.3 Haller, Johannes T 6.3, T 7.2, T 10.4, T 10.5, T 10.8, T 61.8, T 62.6, T 62.7, T 69.5 Hallford, John •T 44.4 Halve, Lasse •T 106.1, T 106.2, T 106.3 T 39.5, Hamacher-Baumann, Philip T 39.5, T 39.6, •T 39.7 Hamer, Matthias T 64.8, T 97.6 Hamilton, Fiona T 104.5, •T 104.5 Hammer, Keleen •D 45.1 Hammer, Kleen •D 45.1 Hammer, Leen •D 45.1 Hanc, Jonte R •AGPhil 9.4 Hannen, Ines T 39.5 Hance, Jonte R •AGPhil 9.4 Hannen, Nolker T 21.4, T 21.5 Hansch, Walter ST 9.5, T 66.4 Hansmann-Menzemer, Stephanie •PV VI, T 50.4, T 50.5 Harder, Sven T 53.8 Hardy, Edme <td></td>	
Hahn, Steffen •AKPIK 4.8, AKPIK 4.9 Haide, Isabel •T 79.3 Hajheidari, Mohammadtaghi •T 12.9 Hajian, Kamal •GR 1.5 Hakenmüller, Janina T 23.1, •T 74.4 Haleem, Mahsana T 88.4 Hall, Zachary •AGPhil 6.3 Haller, Johannes T 63.7 7.2, T 10.4, T 10.5, T 10.8, T 61.8, T 62.6, T 62.7, T 69.5 Hallford, John •T 44.4 Halve, Lasse •T 106.1, T 106.2, T 106.3 Hamacher-Baumann, Philip T 39.5, T 39.6, •T 39.7 Hamer, Matthias T 64.8, T 97.6 Hammer, Kelen •D 45.1 Hammer, Ioea Hann, Robert •T 22.5, T 102.5 Hammer, Keleen •D 45.1 Hann, Robert •T 24.4, T 21.5 Hanne, Ines T 39.5 Hanne, Ines T 39.5 Hanne, Nolker T 21.4, T 21.5 Hansch, Walter ST 9.5, T 66.4 Hansmann-Menzemer, Stephanie •PV VI, T 50.4, T 50.5 Harder, Sven T 53.8 Hardy, Edme D 5.2 Harlander, Robert V. T 31.1 <t< td=""><td></td></t<>	
Hahn, Steffen •AKPIK 4.8, AKPIK 4.9 Haide, Isabel •T 79.3 Hajheidari, Mohammadtaghi •T 12.9 Hajian, Kamal •GR 1.5 Hakenmüller, Janina T 23.1, •T 74.4 Haleem, Mahsana T 88.4 Hall, Zachary •AGPhil 6.3 Haller, Johannes T 63.7 7.2, T 10.4, T 10.5, T 10.8, T 61.8, T 62.6, T 62.7, T 69.5 Hallford, John •T 44.4 Halve, Lasse •T 106.1, T 106.2, T 106.3 Hamacher-Baumann, Philip T 39.5, T 39.6, •T 39.7 Hamer, Matthias T 64.8, T 97.6 Hamitton, Fiona T 104.5, •T 104.6 Hammer, Robert Hammer, Kauthias T 22.5, T 102.5 Hammer, Leen Han Thanh, Linh •EP 6.8 Han, Xiaoxue T 96.6 Hance, Jonte R •AGPhil 9.4 Hannen, Ines T 39.5 Hannen, Ines T 39.5 T 39.5 Hane, Volker T 21.4, T 21.5 Hansch, Walter ST 9.5, T 66.4 Hansmann-Menzemer, Stephanie •PV VI, T 50.4, T 50.5 Harder, Sven T 53.8 Hardy, Edme D 5.2 Harlander, Robert V. T 31.1	
Hahn, Steffen •AKPIK 4.8, AKPIK 4.9 Haide, Isabel •T 79.3 Hajheidari, Mohammadtaghi •T 12.9 Hajian, Kamal •GR 1.5 Hakenmüller, Janina T 23.1, •T 74.4 Haleem, Mahsana T 88.4 Hall, Zachary •AGPhil 6.3 Haller, Johannes T 6.3, T 7.2, T 10.4, T 10.5, T 10.8, T 61.8, T 62.6, T 62.7, T 69.5 Hallford, John •T 44.4 Halve, Lasse •T 106.1, T 106.2, T 106.3 T 104.5, •T 104.5, *T 104.6, Hamacher-Baumann, Philip T 39.5, T 39.6, •T 39.7 Hamer, Matthias T 64.8, T 97.6 Hamitton, Fiona T 104.5, •T 104.6 Hammer, Robert •T 22.5, T 102.5 Hammer, Kauthias T 64.8, T 97.6 Han Thanh, Linh •EP 6.8 Han, Xiaoxue T 96.6 Hance, Jonte R •AGPhil 9.4 Hannen, Ines T 39.5 Hande, Sven T 53.8 Harde, Sven T 53.8 Hardy, Edme D 5.2 Harlander, Robert V. T 31.1 Harlow, Daniel <td></td>	
Hahn, Steffen •AKPIK 4.8, AKPIK 4.9 Haide, Isabel •T 79.3 Hajheidari, Mohammadtaghi •T 12.9 Hajian, Kamal •GR 1.5 Hakenmüller, Janina T 23.1, •T 74.4 Haleem, Mahsana T 88.4 Hall, Zachary •AGPhil 6.3 Haller, Johannes T 6.3, T 7.2, T 10.4, T 10.5, T 10.8, T 61.8, T 62.6, T 62.7, T 69.5 Hallford, John •T 44.4 Halve, Lasse •T 106.1, T 106.2, T 106.3 T 106.3 Hamacher-Baumann, Philip T 39.5, T 39.6, •T 39.7 Hamer, Matthias T 64.8, T 97.6 Hamiton, Fiona T 104.5, •T 104.6 Hammer, Katthias T 64.8, T 97.6 Hammer, Katon T 104.5, •T 104.6 Hammer, Katon T 104.5, •T 104.6 Hammer, Nobert T 22.5, T 102.5 Hammer, Katon T 64.8, T 97.6 Hammer, Jan T 104.5, •T 104.6 Hanner, Nobert T 21.5, T 104.6 Hanner, Kaon T 13.4 Han Thanh, Linh EP 6.8 Han, Xiaoxue T 96.6 </td <td></td>	
Hahn, Steffen •AKPIK 4.8, AKPIK 4.9 Haide, Isabel •T 79.3 Hajheidari, Mohammadtaghi •T 12.9 Hajian, Kamal •GR 1.5 Hakenmüller, Janina T 23.1, •T 74.4 Haleem, Mahsana T 88.4 Hall, Zachary •AGPhil 6.3 Haller, Johannes T 6.3, T 7.2, T 10.4, T 10.5, T 10.8, T 61.8, T 62.6, T 62.7, T 69.5 Hallford, John •T 44.4 Halve, Lasse •T 106.1, T 106.2, T 106.3 Hamacher-Baumann, Philip T 39.5, T 39.6, •T 39.7 Hamer, Matthias T 104.5, •T 104.6 Hammer, Katthias T 104.5, •T 104.6 Hammer, Kathias Hamer, Matthias T 104.5, •T 104.6 Hammer, Leen Han Thanh, Linh •EP 6.8 Han, Xiaoxue T 96.6 Hance, Jonte R •AGPhil 9.4 Hannen, Ines T 39.5, T 39.5, T 66.4 Hansmann-Menzemer, Stephanie •PV VI, T 50.4, T 50.5 Harder, Sven T 53.8 Hardy, Edme DD 5.2 Harlander, Robert V. T 31.1 Harlow, Daniel •SYNS 1.3, •MP 2.1 Harring, Marius DD 34.10 Härtig, Hen	
Hahn, Steffen •AKPIK 4.8, AKPIK 4.9 Haide, Isabel •T 79.3 Hajheidari, Mohammadtaghi •T 12.9 Hajian, Kamal •GR 1.5 Hakenmüller, Janina T 23.1, •T 74.4 Haleem, Mahsana T 88.4 Hall, Zachary •AGPhil 6.3 Haller, Johannes T 6.3, T 7.2, T 10.4, T 10.5, T 10.8, T 61.8, T 62.6, T 62.7, T 69.5 Hallford, John •T 44.4 Halve, Lasse •T 106.1, T 106.2, T 106.3 T 39.5, Hamacher-Baumann, Philip T 39.5, T 39.6, •T 39.7 Hamer, Matthias T 104.5, •T 104.6 Hammerich, Jan •T 13.4 Han Thanh, Linh •EP 6.8 Han, Xiaoxue T 96.6 Hance, Jonte R •AGPhil 9.4 Hannen, Ines T 39.5, Hang, Volker T 21.4, T 21.5 Hansmann-Menzemer, Stephanie •PV VI, T 50.4, T 50.5 Harder, Sobert V. T 31.1 Harlow, Daniel •SYNS 13.4, MP 2.1 Harring, Marius DD 34.10 Hartig, Katja DD 24.2 Hartmann, Nikola	
Hahn, Steffen •AKPIK 4.8, AKPIK 4.9 Haide, Isabel •T 79.3 Hajheidari, Mohammadtaghi •T 12.9 Hajian, Kamal •GR 1.5 Hakenmüller, Janina T 23.1, •T 74.4 Haleem, Mahsana T 88.4 Hall, Zachary •AGPhil 6.3 Haller, Johannes T 6.3, T 7.2, T 10.4, T 10.5, T 10.8, T 61.8, T 62.6, T 62.7, T 69.5 Hallford, John •T 44.4 Halve, Lasse •T 106.1, T 106.2, T 106.3 T 104.5, •T 104.6, +T 104.6, Hamc, Matthias T 64.8, T 97.6 Hamilton, Fiona T 104.5, •T 104.6 Hammer, Keleen •DD 45.1 Hammerich, Jan •T 13.4 Han Thanh, Linh •EP 6.8 Han, Xiaoxue T 96.6 Hance, Jonte R •AGPhil 9.4 Hannen, Ines T 39.5, Hang, Kolker T 21.4, T 21.5 Hansmann-Menzemer, Stephanie •PV VI, T 50.4, T 50.5 Harder, Sobert V. T 31.1 Harlow, Daniel •SYNS 1.3, •MP 2.1 Harring, Marius DD 34.10 Härtig, Katja DD 24.2	
Hahn, Steffen •AKPIK 4.8, AKPIK 4.9 Haide, Isabel •T 79.3 Hajheidari, Mohammadtaghi •T 12.9 Hajian, Kamal •GR 1.5 Hakenmüller, Janina T 23.1, •T 74.4 Haleem, Mahsana T 88.4 Hall, Zachary •AGPhil 6.3 Haller, Johannes T 6.3, T 7.2, T 10.4, T 10.5, T 10.8, T 61.8, T 62.6, T 62.7, T 69.5 Hallford, John •T 44.4 Halve, Lasse •T 106.1, T 106.2, T 106.3 T 39.6, •T 39.7 Hamer, Matthias T 64.8, T 97.6 Hamilton, Fiona T 104.5, •T 104.6 Hammerich, Jan •T 13.4 Han Thanh, Linh •EP 6.8 Han, Xiaoxue T 96.6 Hance, Jonte R •AGPhil 9.4 Hannen, Ines T 39.5, Hannen, Ines T 39.5 Hannen, Volker T 21.4, T 21.5 Hansmann-Menzemer, Stephanie •PV VI, T 50.4, T 50.5 Harder, Sobert V. T 31.1 Harlow, Daniel •SYNS 1.3, •MP 2.1 Harting, Marius DD 34.10 Hartig, Katja DD 24.2	
Hahn, Steffen •AKPIK 4.8, AKPIK 4.9 Haide, Isabel •T 79.3 Hajheidari, Mohammadtaghi •T 12.9 Hajian, Kamal •GR 1.5 Hakenmüller, Janina T 23.1, •T 74.4 Haleem, Mahsana T 88.4 Hall, Zachary •AGPhil 6.3 Haller, Johannes T 6.3, T 7.2, T 10.4, T 10.5, T 10.8, T 61.8, T 62.6, T 62.7, T 69.5 Hallford, John •T 44.4 Halve, Lasse •T 106.1, T 106.2, T 106.3 T 39.6, •T 39.7 Hamer, Matthias T 64.8, T 97.6 Hamilton, Fiona T 104.5, •T 104.6 Hammerich, Jan •T 13.4 Han Thanh, Linh •EP 6.8 Han, Xiaoxue T 96.6 Hance, Jonte R •AGPhil 9.4 Hannen, Ines T 39.5, Hannen, Nes T 39.5 Hannen, Volker T 21.4, T 21.5 Harder, Sobert V. T 31.1 Harlow, Daniel •SYNS 1.3, •MP 2.1 Harls, Katja DD 24.2 Hartmann, Nikolai •T 25.7, T 53.1, T 53.8 Hardy, Edme D 5.2 Harlande	
Hahn, Steffen •AKPIK 4.8, AKPIK 4.9 Haide, Isabel •T 79.3 Hajheidari, Mohammadtaghi •T 12.9 Hajian, Kamal •GR 1.5 Hakenmüller, Janina T 23.1, •T 74.4 Haleem, Mahsana T 88.4 Hall, Zachary •AGPhil 6.3 Haller, Johannes T 63, T 7.2, T 10.4, T 10.5, T 10.8, T 61.8, T 62.6, T 62.7, T 69.5 Hallford, John •T 44.4 Halve, Lasse •T 106.1, T 106.2, T 106.3 Hamacher-Baumann, Philip T 39.5, T 39.6, •T 39.7 Hamer, Matthias T 64.8, T 97.6 Hamilton, Fiona T 104.5, •T 104.6 Hammer, Keleen •DD 45.1 Hammerich, Jan •T 13.4 Han Thanh, Linh •EP 6.8 Han, Xiaoxue T 96.6 Hance, Jonte R •AGPhil 9.4 Hannen, Ines T 39.5 Harlene, Kobert V. T 31.1 Harlow, Daniel •SYNS 1.3, •MP 2.1 Harlsonder, Robert V. T 31.1 Harlow, Daniel •SYNS 1.3, •MP 2.1 Harting, Marius DD 34.10 Härtig, Katja	

Gr

Hauser, Marc ... T 66.6, T 66.7, T 94.1
 Häußler, Rike
 • DD 35.6

 Haußmann, Manuel
 T 25.3

 Haveresch, Helge
 T 102.8

 Havrylenko, Vitalii
 T 11.2, T 11.3,
 •T 11.4 Havrylenko, Vitaliy T 11.1, T 11.6, T 97.7 HAWC's Eye-Kollaboration EP 7.1
 Hebberer, HIOMIAS
 GR 14.4, GR 14.5, GR 14.6, GR 14.6, T 14.3, T 37.2, T 64.1, T 64.2, T 64.3, T 67.4, T 67.9

 Heber, B.
 EP 5.2

 Heber, B.
 EP 5.3

 Heckmann, Lea
 EP 11.2, T 17.8, -T 18.2
 •T 83.2 Hedges, Michael •DD 28.2, •GP 2.1 Heibges, Tobias •T 71.3 Heidelbach. Alexander •T 63.5 Heimel, Theo • T 2.6, T 2.8 Heimeroth, Axel • T 79.6 Heimemann, Beate • T 36.2, T 79.8, T 79.9, T 87.3 Heinemeyer, Sven T 31.4, T 36.7, T 85.5, T 92.6 Heinen, Dirk ... EP 4.3, •EP 4.4, EP 4.5, EP 4.6, T 106.9 Heinicke, Susanne ... DD 16.4, DD 29.1, DD 39.1 DD 39.1 Heinke, Heidrun DD 10.6, DD 12.1, DD 12.8, DD 25.3, DD 33.4, DD 35.13, DD 41.2

 Heinrichs, Jan-Eric
 •T 64.8

 Heinz, Tobias
 •T 20.3, T 20.9, T 48.6, T 48.8, T 102.3

 Heinzel, Thomas
 •ST 2.2

 Heinzelmann, Feline
 •ST 8.1

 Hekker, Saskia
 •EP 1.1

 Helbert, Jörn
 •EP 3.1

 Helbig, Markus
 T 69.1

 Hellmann, Dominik
 •T 86.2, T 86.3

 Hellmich, Luis
 •T 36.9

 Hellrung, Jonas
 T 18.8, •T 46.1, T 46.3, T 46.7

 Hemperek, Tomasz
 ... T 12.4, T 13 5

 T 46.3, T 46.7 Hemperek, Tomasz T 12.4, T 13.5, T 13.6, T 14.2, T 40.5, T 40.8, T 40.9, T 95.2 Hempfling, Janine•T 74.5 Hengstler, DanielT 77.6, T 102.5, T 102.6 .. •T 74 5 Henkes, Florian•T 75.2 Henning Rehren, Karl MP 7.5 Herbst, Konstantin•EP 2.8, EP 3.2, EP 5.3
 Herbst, Matthew
 T 102.6

 Hermann, Leonie
 T 5.7

 Hermanns, Simon
 DD 46.3
 Hermannsgabner, Johanna ... EP 4.3, EP 4.6 EP 4.6 Herold, Adam•T 35.1, T 35.7 Herrero Gascón, Paula•T 43.7
 Herreno Gascoli, Falla
 T. 107.6

 Herrmann, Sven
 GR 2.8

 Herten, Gregor
 T 26.9, T 39.3

 Hertenberger, Ralf
 T 11.1, T 11.2, T 11.3, T 11.4, T 11.5, T 11.6, T 97.7
 Heuel, JohannesT 57.4 Heuermann, LarsT 48.1, •T 48.2, T 72.9 Heurs, Michèle •EP 8.3, •T 80.3 Heusler, Stefan DD 3.3, DD 16.4, DD 17.1, •DD 26.3, DD 37.2

Hils, Christopher•T 78.1
Hinkelmann, Maria
Hinterkeuser, Fiorian 1 12.4, •1 97.6 Hinton James Anthony T 101.5
Hinton, Jim
Hinze, Johannes•T 61.2
Hinzmann, Andreas T 10.1, T 32.6,
I 66.9, I 87.1 Hiropo Toko T 40.8 T 40.9
Hirschbühl Dominic T 6 2 T 6 9
T 34.4, T 60.4, •T 88.1
Hirzberger, Johann EP 12.7
Hmaid, Ramy•T 8.2
Hoche, Stefan 1 88.3 Hoek Matthias T 95.3
Hoepfner. Kerstin T 37.2. T 64.3.
Т 67.1, Т 67.8, Т 67.9
Hofer, Elisabeth DD 2.3
Höfer, Judith•T 36.2
Hoffmann Stefan DD 22.1 •DD 22.2
Höflich. Nina
Hofmann, Jan Niclas T 3.9
Höfner, Jessica T 34.2, •T 34.3
Hotsaess, Robin
Hohmann Sascha DD 45.2
Hölken. Patrick
Hollit, Sophie
Hollitt, Sophie . T 60.7, T 96.7, •T 96.8
Hollnagel, Annika•T 42.7
Holman Maro
Holzanfel Kilian T 23 4
Holzbock, Michael T 9.4, T 37.5
Holzinger, Eva DD 10.8, DD 18.4
Hönig, Cedric
Hood, Alan W EP 12.5
Hopt, Martin UU 18.2, UU 19.2,
Hörbe, Mario
Horn, Martin Erik
Horns, Dieter
Hornung, Daniel •AKPIK 1.3,
•AKPIK 1.4
Horz, Holger
Horzela Maximilian T 15 3 T 15 7
Horzela, Maximilian •T 15.3, T 15.7, T 32.7
Horzela, Maximilian •T 15.3, T 15.7, T 32.7 Hosseini, Sepideh•T 63.6
Horzela, Maximilian •T 15.3, T 15.7, T 32.7 Hosseini, Sepideh•T 63.6 Höttecke, Dietmar•SYNS 1.1
Horzela, Maximilian •T 15.3, T 15.7, T 32.7 Hosseini, Sepideh •T 63.6 Höttecke, Dietmar •SYNS 1.1 Hötting, Marius ST 1.2, •ST 4.3, ST 4.4 ST 4.5 ST 6.2
Horzela, Maximilian •T 15.3, T 15.7, T 32.7 Hosseini, Sepideh •T 63.6 Höttecke, Dietmar \$YNS 1.1 Hötting, Marius \$T 1.2, •ST 4.3, ST 4.4, ST 4.5, ST 6.2 Hötzsch Luisa •T 22 1
Horzela, Maximilian •T 15.3, T 15.7, T 32.7 Hosseini, Sepideh •T 63.6 Höttecke, Dietmar •SYNS 1.1 Hötting, Marius ST 1.2, •ST 4.3, ST 4.4, ST 4.5, ST 6.2 Hötzsch, Luisa •T 22.1 Hou, Wenjie •T 100.5
Horzela, Maximilian •T 15.3, T 15.7, T 32.7 Hosseini, Sepideh •T 63.6 Höttecke, Dietmar •SYNS 1.1 Hötting, Marius •ST 1.2, •ST 4.3, ST 4.4, ST 4.5, ST 6.2 Hötzsch, Luisa •T 22.1 Hou, Wenjie •T 100.5 Houdy, Thibaut T 23.3
Horzela, Maximilian •T 15.3, T 15.7, T 32.7 Hosseini, Sepideh •T 63.6 Höttecke, Dietmar •SYNS 1.1 Höttig, Marius •ST 1.2, •ST 4.3, ST 4.4, ST 4.5, ST 6.2 Hötzsch, Luisa •T 22.1 Hou, Wenjie •T 20.3 Houdy, Thibaut T 23.3 Hu, Jiangqiao •T 4.6
Horzela, Maximilian •T 15.3, T 15.7, T 32.7 Hosseini, Sepideh •T 63.6 Höttecke, Dietmar •SYNS 1.1 Hötting, Marius •ST 1.2, •ST 4.3, ST 4.4, ST 4.5, ST 6.2 Hötzsch, Luisa •T 22.1 Hou, Wenjie •T 20.3 Hu, Jiangqiao •T 4.6 Huber, Thomas •T 66.3
Horzela, Maximilian •T 15.3, T 15.7, T 32.7 Hosseini, Sepideh •T 63.6 Höttecke, Dietmar •SYNS 1.1 Hötting, Marius •ST 1.2, •ST 4.3, ST 4.4, ST 4.5, ST 6.2 Hötzsch, Luisa •T 22.1 Hou, Wenjie •T 100.5 Houdy, Thibaut T 23.3 Hu, Jiangqiao •T 4.6 Huber, Thomas T 66.3 Huber, Tobias T 56.6 Huber, Tobias • 56.6
Horzela, Maximilian •T 15.3, T 15.7, T 32.7 Hosseini, Sepideh •T 63.6 Höttecke, Dietmar •SYNS 1.1 Hötting, Marius •ST 1.2, •ST 4.3, ST 4.4, ST 4.5, ST 6.2 Hötzsch, Luisa •T 22.1 Hou, Wenjie •T 100.5 Houdy, Thibaut T 23.3 Hu, Jiangqiao •T 4.6 Huber, Thomas T 66.3 Huber, Tobias AGPhil 4.3 Huber, Tim
Horzela, Maximilian •T 15.3, T 15.7, T 32.7 Hosseini, Sepideh •T 63.6 Höttecke, Dietmar •SYNS 1.1 Hötting, Marius •ST 1.2, •ST 4.3, ST 4.4, ST 4.5, ST 6.2 Hötzsch, Luisa •T 22.1 Hou, Wenjie •T 100.5 Houdy, Thibaut T 26.3 Huber, Thomas •T 66.3 Huber, Tobias AGPhil 4.3 Huber, Tim AGPhil 4.3 Huber, Tim
Horzela, Maximilian • T 15.3, T 15.7, T 32.7 Hosseini, Sepideh • T 63.6 Höttecke, Dietmar • SYNS 1.1 Höttsch, Luisa • SYNS 1.2, • ST 4.3, ST 4.4, ST 4.5, ST 6.2 Hötzsch, Luisa • T 22.1 Hou, Wenjie • T 100.5 Houdy, Thibaut T 23.3 Huber, Thomas T 66.3 Huber, Tobias T 56.6, T 56.6 Huber, Tobias T 66.3 Huber, Tobias T 72.7, T 101.7 Hugg, Tim T 12.4, T 13.5, T 40.8, T 40.9, T 42.6, T 95.2, T 97.6
Horzela, Maximilian •T 15.3, T 15.7, T 32.7 Hosseini, Sepideh •T 63.6 Höttecke, Dietmar •SYNS 1.1 Hötting, Marius ST 1.2, •ST 4.3, ST 4.4, ST 4.5, ST 6.2 Hötzsch, Luisa •T 22.1 Hou, Wenjie •T 20.5 Houdy, Thibaut T 20.5 Houdy, Thibaut T 20.5 Huber, Thomas T 66.3 Huber, Thomas T 66.3 Huber, Tobias T 56.3, T 56.6 Hubert, Mario AGPhil 4.3 Huege, Tim T 72.7, T 101.7 Hügging, Fabian T 12.4, T 13.5, T 40.8, T 40.9, T 42.6, T 95.2, T 97.6 Hügli, Cédrine
Horzela, Maximilian •T 15.3, T 15.7, T 32.7 Hosseini, Sepideh •T 63.6 Höttecke, Dietmar •SYNS 1.1 Hötting, Marius ST 1.2, •ST 4.3, ST 4.4, ST 4.5, ST 6.2 Hötzsch, Luisa •T 22.1 Hou, Wenjie •T 100.5 Houdy, Thibaut T 23.3 Huber, Thomas •T 66.3 Huber, Thomas T 66.3 Huber, Tobias T 56.3, T 56.6 Hubert, Mario AGPhil 4.3 Huege, Tim T 72.7, T 101.7 Hüggling, Fabian T 12.4, T 13.5, T 40.8, T 40.9, T 42.6, T 95.2, T 97.6 Hügli, Cédrine T 9.8 Huhmann, Christian
Horzela, Maximilian T 15.3, T 15.7, T 32.7 Hosseini, Sepideh T 63.6 Höttecke, Dietmar SYNS 1.1 Höttig, Marius SYNS 1.1 Höttecke, Dietmar SYNS 1.1 Höttig, Marius ST 1.2, •ST 4.3, ST 4.4, ST 4.5, ST 6.2 Hötzsch, Luisa T 22.1 Hötzsch, Luisa T 20.5 Houdy, Thibaut T 23.3 Hu, Jiangqiao T 4.6 Huber, Thomas T 66.3 Huber, Tobias T 56.3, T 56.6 Huber, Tobias T 72.7, T 101.7 Hügging, Fabian T 12.4, T 13.5, T 40.8, T 40.9, T 42.6, T 95.2, T 97.6 Hügli, Cédrine T 9.8 Huhmann, Christian T 22.6 Hull, Michael M DD 10.7, DD 10.8, DD 10.17, .DD 10.12,
$\begin{array}{llllllllllllllllllllllllllllllllllll$
Horzela, Maximilian T 15.3, T 15.7, T 32.7 Hosseini, Sepideh T 63.6 Höttecke, Dietmar SYNS 1.1 Höttig, Marius ST 1.2, sT 4.3, ST 4.4, ST 4.5, ST 6.2 Hötzsch, Luisa T 22.1 Hötzsch, Luisa T 22.1 Höudy, Thibaut T 23.3 Hu, Jiangqiao T 66.3 Huber, Thomas T 56.3, T 56.6 Huber, Tobias T 56.3, T 56.6 Huber, Tobias T 72.7, T 101.7 Hügging, Fabian T 12.4, T 13.5, T 40.8, T 40.9, T 42.6, T 95.2, T 97.6 Hügli, Cédrine T 9.8 Huhmann, Christian T 22.6 Hul, Michael M DD 10.7, •DD 10.8, •DD 10.10, •DD 10.7, •DD 10.8, •DD 11.0, •DD 10.17, •DD 10.8, •DD 11.9, •DD 18.4
Horzela, Maximilian T 15.3, T 15.7, T 32.7 Hosseini, Sepideh T 63.6 Höttecke, Dietmar SYNS 1.1 Höttig, Marius ST 1.2, •ST 4.3, ST 4.4, ST 4.5, ST 6.2 Hötzsch, Luisa T 22.1 Hötzsch, Luisa T 20.3 Hu, Jiangqiao T 4.6 Huber, Thomas T 66.3 Huber, Tobias T 56.3, T 56.6 Huber, Tobias T 27, T 101.7 Hüging, Fabian T 12.4, T 13.5, T 40.8, T 40.9, T 42.6, T 95.2, T 97.6 Hügli, Cédrine T 9.8 Huhmann, Christian T 22.6 Hul, Michael M DD 10.7, •DD 10.8, •DD 10.12, •DD 10.12, •DD 10.11, •DD 10.12, •DD 11.9, •DD 18.4 Humair, Thibaud T 30.3, •T 54.2 Hümbert-Schnur, Sebastian D 33.4
Horzela, Maximilian T 15.3, T 15.7, T 32.7 Hosseini, Sepideh T 63.6 Höttecke, Dietmar SYNS 1.1 Höttsch, Luisa T 22.1 Hötzsch, Luisa T 20.5 Houdy, Thibaut T 23.3 Hu, Jiangqiao T 4.6 Huber, Thomas T 66.3 Huber, Tobias T 56.3, T 56.6 Huber, Thomas T 66.3 Huber, Tobias T 72.7, T 101.7 Hügging, Fabian T 12.4, T 13.5, T 40.8, T 40.9, T 42.6, T 95.2, T 97.6 Hügli, Cédrine T 93.6 Huhmann, Christian T 22.6 Hull, Michael M DD 10.7, •DD 10.18, •D 10.12, •D 10.11, •DD 10.12, •D 10.11, •DD 10.12, •D 10.11, •DD 10.12, •D 10.11, •DD 10.12, •D 10.11, •D 10.12, •D 10.11, •D 10.12, •D 10.11, •D 10.12, •D 11.9, •D 18.4 Humbert-Schnur, Sebastian DD 33.4 Hummer, Fabian T 41.2 Hummerich, Sander T 41.2
Horzela, Maximilian T 15.3, T 15.7, T 32.7 Hosseini, Sepideh T 63.6 Höttecke, Dietmar SYNS 1.1 Höttsch, Luisa T 22.1 Hou, Wenjie T 100.5 Houdy, Thibaut T 23.3 Hu, Jiangqiao T 4.6 Huber, Thomas T 66.3 Huber, Tobias T 56.3, T 56.6 Huber, Tobias T 56.3, T 56.6 Huber, Tobias T 56.3, T 56.6 Huber, Tobias T 72.7, T 101.7 Hügging, Fabian T 12.4, T 13.5, T 40.8, T 40.9, T 42.6, T 95.2, T 97.6 Hügli, Cédrine T 9.8 Huhmann, Christian T 22.6 Hul, Michael M DD 10.7, •DD 10.8, •DD 10.17, •DD 10.12, •DD 10.12, •DD 10.10, •DD 10.11, •DD 10.12, •DD 11.9, •DD 18.4 Humair, Thibaud T 30.3, •T 54.2 Hümbert-Schnurr, Sebastian DD 33.4 Hummer, Fabian T 2.8 Hundhausen, Daniel T 2.8 <
Horzela, Maximilian T 15.3, T 15.7, T 32.7 Hosseini, Sepideh T 63.6 Höttecke, Dietmar SYNS 1.1 Höttsch, Luisa T 22.1 Hou, Wenjie T 100.5 Houdy, Thibaut T 23.3 Huber, Thomas T 66.3 Huber, Tobias T 56.6 Huber, Tobias T 56.3, T 56.6 Huber, Mario AGPhil 4.3 Huege, Tim T 72.7, T 101.7 Hügging, Fabian T 12.4, T 13.5, T 40.8, T 40.9, T 42.6, T 95.2, T 97.6 Hügli, Cédrine T 9.4 Huhmann, Christian T 22.6 Hull, Michael M DD 10.7, •DD 10.8, •DD 10.1, •DD 10.12, •DD 10.12, •DD 10.19, •DD 11.8, •DD 10.19, •DD 18.4 Humair, Thibaud T 30.3, •T 54.2 Hümbert-Schnurr, Sebastian DD 33.4 Hummer, Fabian T 2.8 Hundhausen, Daniel T 2.7 Hundhausen, Daniel Christian T 62.6
Horzela, Maximilian •T 15.3, T 15.7, T 32.7 Hosseini, Sepideh •T 63.6 Höttecke, Dietmar •SYNS 1.1 Hötting, Marius ST 1.2, •ST 4.3, ST 4.4, ST 4.5, ST 6.2 Hötzsch, Luisa •T 22.1 Hou, Wenjie •T 100.5 Houdy, Thibaut T 23.3 Hu, Jiangqiao •T 4.6 Huber, Thomas T 66.3 Huber, Tobias T 56.3, T 56.6 Hubert, Mario AGPhil 4.3 Huege, Tim T 72.7, T 101.7 Hügging, Fabian T 12.4, T 13.5, T 40.8, T 40.9, T 42.6, T 95.2, T 97.6 Hügli, Cédrine T 92.4 Hul, Michael M •DD 10.7, •DD 10.8, •DD 10.10, •DD 10.11, •DD 10.12, •DD 11.9, •DD 18.4 Humari, Thibaud T 30.3, •T 54.2 Hümbert-Schnur, Sebastian DD 33.4 Hummer, Fabian T 2.8 Hundhausen, Daniel T 62.6
Horzela, Maximilian •T 15.3, T 15.7, T 32.7 Hosseini, Sepideh •T 63.6 Höttecke, Dietmar \$YNS 1.1 Hötting, Marius \$T 1.2, •ST 4.3, ST 4.4, ST 4.5, ST 6.2 Hötzsch, Luisa •T 22.1 Hou, Wenjie •T 20.5 Houdy, Thibaut T 20.5 Houdy, Thibaut T 20.5 Huber, Thomas T 66.3 Huber, Tobias T 66.3 Huber, Tobias T 66.3, T 56.6 Huber, Mario AGPhil 4.3 Hueg, Tim T 72.7, T 101.7 Hügging, Fabian T 12.4, T 13.5, T 40.8, T 40.9, T 42.6, T 95.2, T 97.6 Hügli, Cédrine T 9.8 Huhmann, Christian T 20.6 Hull, Michael M •DD 10.7, •DD 10.8, •DD 10.10, •DD 18.4 Hummerich, Sander T 2.8 Hundhausen, Daniel T 42.7 Hundhausen, Daniel T 62.7 Hundhausen, Daniel T 62.7 Hundhausen, Daniel T 62.7 Hundhausen, Daniel T 62.7 Hundhausen, Daniel T 64.7 Hundhausen, Daniel AKPIK 4.2 Hünnefeld, Mirco T 18.5, T 46.6
$\begin{array}{llllllllllllllllllllllllllllllllllll$
Horzela, Maximilian •T 15.3, T 15.7, T 32.7 Hosseini, Sepideh •SYNS 1.1 Hötting, Marius \$SYNS 1.1 Hötting, Marius \$ST 4.3, ST 4.4, ST 4.5, ST 6.2 Hötzsch, Luisa \$T 22.1 Houdy, Thibaut \$T 22.1 Huber, Thomas \$T 66.3 Huber, Tobias \$T 56.3, T 56.6 Hubert, Mario AGPhil 4.3 Huege, Tim \$T 72.7, T 101.7 Hügging, Fabian \$T 22.6, T 95.2, T 97.6 Hügli, Cédrine \$T 9.8 Huhmann, Christian \$T 22.6 Hull, Michael M \$D 10.7, •DD 10.18, •DD 10.10, •DD 10.11, •DD 10.12, •DD 11.9, •DD 18.4 Hummer, Fabian \$T 41.2 Hummerich, Sander \$T 4.2.8 Hundhausen, Daniel \$T 4.2.6 Hundhausen, Daniel \$T 4.2.6 Hundhausen, Daniel \$T 4.6.6 Husemann, Ulrich \$T 6.4, T 26.7, T 36.3, T 65.5, T 65.6, T 65.7, T 76.8, T 88.7, T 88.8, T 88.9, T 94.4
Horzela, Maximilian T 15.3, T 15.7, T 32.7 Hosseini, Sepideh SYNS 1.1 Hötting, Marius ST 4.3, ST 4.4, ST 4.5, ST 6.2 Hötzsch, Luisa T 22.1 Hou, Wenjie T 100.5 Houdy, Thibaut T 22.3 Hu, Jiangqiao T 4.6 Huber, Thomas T 66.3, T 56.6 Hubert, Mario AGPhil 4.3 Huege, Tim T 72.7, T 101.7 Hügging, Fabian T 12.4, T 13.5, T 40.8, T 40.9, T 42.6, T 95.2, T 97.6 Hügli, Cédrine T 98. Huhmann, Christian T 22.6 Hull, Michael M DD 10.12, •DD 10.10, •DD 10.11, •DD 10.12, •DD 11.9, •DD 18.4 Hummer, Fabian T 41.2 Hummerich, Sander T 2.8 Hundhausen, Daniel AKPIK 4.2 Hünhefeld, Mirco T 18.5, T 46.6 Husemann, Ulrich ST 6.4, T 26.7, T 36.3, T 65.5, T 65.6, T 65.7, T 76.8, T 88.7, T 88.7, T 88.8, T 88.9, T 94.4 Huseynov, Nazim T 35.1, T 35.6,
Horzela, Maximilian •T 15.3, T 15.7, T 32.7 Hosseini, Sepideh •SYNS 1.1 Hötting, Marius ST 1.2, •ST 4.3, ST 4.4, ST 4.5, ST 6.2 Hötzsch, Luisa •T 22.1 Hötzsch, Luisa •T 22.1 Hou, Wenjie •T 100.5 Houdy, Thibaut T 23.3 Hu, Jiangqiao •T 4.6 Huber, Thomas T 66.3 Huber, Tobias T 56.3, T 56.6 Hubert, Mario AGPhil 4.3 Huege, Tim T 72.7, T 101.7 Hügging, Fabian T 12.4, T 13.5, T 40.8, T 40.9, T 42.6, T 95.2, T 97.6 Hügli, Cédrine •T 21.4, T 13.5, T 40.8, T 40.9, T 42.6, T 95.2, T 97.6 Hügli, Cédrine T 20.3, •T 54.2 Hulhmann, Christian T 20.3, •T 54.2 Hümbert-Schnur, Sebastian DD 33.4 Hummer, Fabian T 41.2 Hümbert-Schnur, Sebastian DD 33.4 Hummer, Fabian T 2.8 Hundhausen, Daniel AKPIK 4.2 Hünhausen, Daniel
$\begin{array}{llllllllllllllllllllllllllllllllllll$
$\begin{array}{llllllllllllllllllllllllllllllllllll$
Horzela, Maximilian •T 15.3, T 15.7, T 32.7 Hosseini, Sepideh •T 63.6 Höttecke, Dietmar •SYNS 1.1 Hötting, Marius ST 1.2, •ST 4.3, ST 4.4, ST 4.5, ST 6.2 Hötzsch, Luisa •T 22.1 Höu, Wenjie •T 20.5 Houdy, Thibaut T 23.3 Hu, Jiangqiao •T 4.6 Huber, Thomas T 66.3 Huber, Tobias T 56.3, T 56.6 Hubert, Mario AGPhil 4.3 Huege, Tim T 72.7, T 101.7 Hügging, Fabian T 12.4, T 13.5, T 40.8, T 40.9, T 42.6, T 95.2, T 97.6 Hügli, Cédrine •T 9.8 Huhmann, Christian T 22.6 Hull, Michael M •DD 10.17, •DD 10.18, •DD 10.10, •DD 10.11, •DD 10.12, •DD 11.0, •DD 10.11, •DD 10.12, •DD 11.0, •DD 10.11, •DD 10.12, •DD 11.9, •DD 18.4 Hummerich, Sander T 2.8 Hundhausen, Daniel Christian •T 62.6 Hündhausen, Daniel Christian •T 62.6 Hündhausen, Daniel AKPIK 4.2 Hünnefeld, Mirco •T 18.5, T 46.6 Husemann, Ulrich ST 6.4, T 26.7, T 36.3, T 65.5, T 65.6, T 65.7, T 76.8, T 88.6, T 88.7, T 88.8, T 88.9, T 94.4 Huseynov, Nazim T 35.1, T 35.6, T 35.7 Hütten, Moritz
Horzela, Maximilian •T 15.3, T 15.7, T 32.7 Hosseini, Sepideh •T 63.6 Höttecke, Dietmar •SYNS 1.1 Hötting, Marius ST 1.2, •ST 4.3, ST 4.4, ST 4.5, ST 6.2 Hötzsch, Luisa •T 22.1 Höu, Wenjie •T 20.5 Houdy, Thibaut T 23.3 Hu, Jiangqiao •T 20.5 Houdy, Thibaut T 23.3 Hu, Jiangqiao •T 4.6 Huber, Thomas T 66.3 Huber, Tobias T 56.3, T 56.6 Hubert, Mario AGPhil 4.3 Huege, Tim T 72.7, T 101.7 Hügging, Fabian T 12.4, T 13.5, T 40.8, T 40.9, T 42.6, T 95.2, T 97.6 Hügli, Cédrine •T 9.8 Huhmann, Christian T 22.6 Hull, Michael M •DD 10.7, •DD 10.8, •DD 10.10, •DD 10.17, •DD 10.12, •DD 11.0, •DD 10.17, •DD 10.12, •DD 11.9, •DD 18.4 Hummer, Fabian •T 41.2 Hümbert-Schnurr, Sebastian DD 33.4 Hummer, Fabian T 42.8, T 48.6, Huusen, Daniel Christian •T 62.6 Hundhausen, Daniel AKPIK 4.2 Hünnefeld, Mirco
Horzela, Maximilian •T 15.3, T 15.7, T 32.7 Hosseini, Sepideh •T 63.6 Höttecke, Dietmar •SYNS 1.1 Hötting, Marius ST 1.2, •ST 4.3, ST 4.4, ST 4.5, ST 6.2 Hötzsch, Luisa •T 22.1 Höu, Wenjie •T 100.5 Houdy, Thibaut T 23.3 Hu, Jiangqiao •T 22.3 Huber, Thomas T 66.3 Huber, Thomas T 66.4 Huber, Thomas T 12.4, T 13.5, T 40.8, T 40.9, T 42.6, T 95.2, T 97.6 Hügli, Cédrine T 9.8 Huhmann, Christian T 22.6 Hull, Michael M •DD 10.7, •DD 10.8, •DD 10.10, •DD 10.11, •DD 10.12, •DD 11.9, •DD 18.4 Hummer, Fabian •T 42.4, T 43.3, •T 54.2 Hümbert-Schnurr, Sebastian DD 33.4 Hummer, Fabian •T 48.5, T 46.6 Husemann, Ulrich ST 6.4, T 26.7, Hundhausen, Daniel Christian •T 62.6 Hundshausen, Daniel AKPIK 4.2 Hünnefeld, Mirco T 18.5, T 46.6 Husemann, Ulrich ST 6.4, T 26.7, T 36.3, T 65.5, T 65.6, T 65.7, T 76.8, T 88.6, T 88.7, T 88.8, T 88.9, T 94.4 Huseynov, Nazim T 35.1, T 35.6, T 72.2, T 72.3, T 83.8 Iacobacci, Lorenzo •MP 2.4 Ibis, Philipp T 60.6, T 60.7 Ibragimov, Iskander T 12.7 Ibragimov, Iskander T 12.7
Horzela, Maximilian •T 15.3, T 15.7, T 32.7 Hosseini, Sepideh •T 63.6 Höttecke, Dietmar •SYNS 1.1 Hötting, Marius ST 1.2, •ST 4.3, ST 4.4, ST 4.5, ST 6.2 Hötzsch, Luisa •T 22.1 Höu, Wenjie •T 120.5 Houdy, Thibaut T 23.3 Hu, Jiangqiao •T 4.6 Huber, Thomas T 66.3 Huber, Tobias T 56.3, T 56.6 Hubert, Mario AGPhil 4.3 Huege, Tim T 12.4, T 13.5, T 40.8, T 40.9, T 42.6, T 95.2, T 97.6 Hügli, Cédrine •T 9.8 Huhmann, Christian T 22.6 Hull, Michael M •DD 10.7, •DD 10.8, •DD 10.10, •DD 10.11, •DD 10.12, •DD 11.9, •DD 18.4 Humair, Thibaud T 30.3, •T 54.2 Hümbert-Schnur, Sebastian DD 33.4 Hummer, Fabian •T 42.6, T 95.2, T 97.6 Hügli, Cédrine •T 30.3, •T 54.2 Hümbert-Schnur, Sebastian DD 33.4 Hummer, Fabian •T 41.2 Humhausen, Daniel Christian •T 62.6 Hundhausen, Daniel AKPIK 4.2 Hünnefeld, Mirco T 18.5, T 46.6 Husemann, Ulrich ST 6.4, T 26.7, T 36.3, T 65.5, T 65.6, T 65.7, T 76.8, T 88.6, T 88.7, T 88.8, T 88.9, T 94.4 Huseynov, Nazim T 35.1, T 35.6, T 72.2, T 72.3, T 83.8 Iacobacci, Lorenzo •MP 2.4 Ibis, Philipp T 60.6, T 60.7 Ibragimov, Iskander T 12.7 IceCube-Kollaboration T 18.5, T 46.1, T 98.7, T 46.7, T 93.9, T 106.2, T 72.3, T 72.3, T 73.5, T 12.7
Horzela, Maximilian •T 15.3, T 15.7, T 32.7 Hosseini, Sepideh •T 63.6 Höttecke, Dietmar \$YNS 1.1 Hötting, Marius \$T 1.2, •ST 4.3, ST 4.4, ST 4.5, ST 6.2 Hötzsch, Luisa T 22.1 Höu, Wenjie T 120, •T 100.5 Houdy, Thibaut T 23.3 Hu, Jiangqiao •T 4.6 Huber, Thomas T 66.3 Huber, Tobias T 66.3, T 56.6 Hubert, Mario AGPhil 4.3 Huege, Tim T 12.4, T 13.5, T 40.8, T 40.9, T 42.6, T 95.2, T 97.6 Hügli, Cédrine T 9.8 Huhmann, Christian T 22.6 Hull, Michael M •DD 10.7, •DD 10.8, •DD 10.10, •DD 10.11, •DD 10.12, •DD 11.9, •DD 18.4 Humair, Thibaud T 30.3, •T 54.2 Hümbert-Schnur, Sebastian DD 33.4 Hummer, Fabian T 42.6, T 95.7, T 66.8 Hundhausen, Daniel Christian •T 62.6 Hundhausen, Daniel Christian •T 62.6 Hundhausen, Daniel ST 6.4, T 26.7, T 36.3, T 65.5, T 65.6, T 65.7, T 76.8, T 88.6, T 88.7, T 88.8, T 88.9, T 94.4 Huseynov, Nazim T 35.1, T 35.6, T 35.7 Hütten, Moritz T 18.5, T 46.6 Husemann, Jensen T 12.7 IceCube-Kollaboration T 18.5, T 46.6, T 12.7, T 23.7, T 33.8 Iacobacci, Lorenzo •MP 2.4 Ibis, Philipp T 60.6, T 60.7 Ibragimov, Iskander T 12.7 IceCube-Kollaboration T 18.5, T 46.1, T 93.8, T 23.5, T 103.4, AKPIK 4.1,
Horzela, Maximilian •T 15.3, T 15.7, T 32.7 Hosseini, Sepideh •T 63.6 Höttecke, Dietmar \$YNS 1.1 Hötting, Marius \$T 1.2, •ST 4.3, ST 4.4, ST 4.5, ST 6.2 Hötzsch, Luisa T 22.1 Höu, Wenjie T 12.4, ST 4.3, Hu, Jiangqiao T 4.6 Huber, Thomas T 66.3 Huber, Tobias T 66.3 Huber, Tobias T 66.3, T 56.6 Hubert, Mario AGPhil 4.3 Hueg, Tim T 12.4, T 13.5, T 40.8, T 40.9, T 42.6, T 95.2, T 97.6 Hügli, Cédrine T 12.4, T 13.5, T 40.8, T 40.9, T 42.6, T 95.2, T 97.6 Hügli, Cédrine T 98 Huhmann, Christian T 22.6 Hull, Michael M •DD 10.7, •DD 10.8, •DD 10.10, •DD 10.11, •DD 10.12, •DD 11.9, •DD 18.4 Humair, Thibaud T 30.3, •T 54.2 Hümbert-Schnur, Sebastian DD 33.4 Hummer, Fabian T 48.4, T 42.6, T 95.7, T 66.7 Hundhausen, Daniel Christian •T 62.6 Hundshausen, Daniel Christian •T 62.7 Hundhausen, Daniel ST 6.4, T 26.7, T 36.3, T 65.5, T 65.6, T 65.7, T 76.8, T 38.6, T 88.7, T 88.8, T 88.9, T 94.4 Huseynov, Nazim T 35.1, T 35.6, T 35.7 Hütten, Moritz T 17.8 Hymon, Karolin EP 11.8, •T 46.5, T 72.2, T 72.3, T 83.8 Iacobacci, Lorenzo MP 2.4 Ibis, Philipp T 60.6, T 60.7 Ibragimov, Iskander T 12.7 IceCube-Kollaboration T 18.5, T 46.1, T 98.7, T 46.7, T 93.9, T 106.2, T 72.3, T 93.8, T 23.5, T 103.4, AKPIK 4.1, EP 11.1, T 100.1, T 106.5, T 46.2, T 100.1, T 105.5, T 46.2, T 100.
Horzela, Maximilian •T 15.3, T 15.7, T 32.7 Hosseini, Sepideh •T 63.6 Höttecke, Dietmar \$YNS 1.1 Hötting, Marius \$T 1.2, •ST 4.3, ST 4.4, ST 4.5, ST 6.2 Hötzsch, Luisa T 22.1 Höu, Wenjie T 12., •T 23.3 Hu, Jiangqiao T 4.6 Huber, Thomas T 66.3 Huber, Tobias T 66.3, T 56.6 Huber, Thomas T 12.4, T 13.5, T 40.8, T 40.9, T 42.6, T 95.2, T 97.6 Hügli, Cédrine T 12.4, T 13.5, T 40.8, T 40.9, T 42.6, T 95.2, T 97.6 Hügli, Cédrine T 98. Huhmann, Christian T 22.6 Hull, Michael M •DD 10.7, •DD 10.8, •DD 10.10, •DD 10.11, •DD 10.12, •DD 11.9, •DD 18.4 Humair, Thibaud T 30.3, •T 54.2 Hümbert-Schnur, Sebastian DD 33.4 Hummer, Fabian T 48.4, T 46.7 Hundhausen, Daniel Christian •T 62.6 Hundshausen, Daniel ST 6.4, T 26.7, T 36.3, T 65.5, T 65.6, T 65.7, T 76.8, T 38.6, T 88.7, T 88.8, T 88.9, T 94.4 Huseynov, Nazim T 35.1, T 35.6, T 35.7 Hütten, Moritz T 17.8 Hymon, Karolin EP 11.8, •T 46.5, T 72.2, T 72.3, T 83.8 Iacobacci, Lorenzo MP 2.4 Ibis, Philipp T 60.6, T 60.7 Ibragimov, Iskander T 12.7 IceCube-Kollaboration T 18.5, T 46.1, T 98.7, T 46.7, T 93.9, T 106.2, T 72.3, T 93.8, T 23.5, T 103.4, AKPIK 4.1, EP 11.1, T 100.1, T 106.5, T 46.2, T 100.3, T 46.4, T 106.3, T 106.4, T 46.5, T 106.7, T 106.4, T 46.5, T 106.7, T 106.9, T 100.6,
Horzela, Maximilian •T 15.3, T 15.7, T 32.7 Hosseini, Sepideh •T 63.6 Höttecke, Dietmar \$YNS 1.1 Hötting, Marius \$T 1.2, •ST 4.3, ST 4.4, ST 4.5, ST 6.2 Hötzsch, Luisa T 22.1 Höu, Wenjie T 12., •T 30.3 Hu, Jiangqiao T 20.3 Huber, Thomas T 66.3 Huber, Thomas T 66.3 Huber, Thomas T 66.3, T 56.6 Huber, Mario AGPhil 4.3 Huege, Tim T 12.4, T 13.5, T 40.8, T 40.9, T 42.6, T 95.2, T 97.6 Hügli, Cédrine T 12.4, T 13.5, T 40.8, T 40.9, T 42.6, T 95.2, T 97.6 Hügli, Cédrine T 9.8 Huhmann, Christian T 22.6 Hull, Michael M •DD 10.7, •DD 10.8, •DD 10.10, •DD 10.11, •DD 10.12, •DD 11.9, •DD 18.4 Humair, Thibaud T 30.3, •T 54.2 Hümbert-Schnur, Sebastian DD 33.4 Hummer, Fabian T 42.4, T 42.6 Hundhausen, Daniel AKPIK 4.2 Hünnefeld, Mirco T 18.5, T 46.6 Husemann, Ulrich ST 6.4, T 26.7, T 36.3, T 65.5, T 65.6, T 65.7, T 76.8, T 88.6, T 88.7, T 88.8, T 88.9, T 94.4 Huseynov, Nazim T 35.1, T 35.6, T 35.7 Hütten, Moritz T 17.8 Hymon, Karolin EP 11.8, •T 46.5, T 72.2, T 72.3, T 83.8 Iacobacci, Lorenzo MP 2.4 Ibis, Philipp T 60.6, T 60.7 Ibragimov, Iskander T 18.5, T 46.1, T 98.7, T 46.7, T 93.9, T 106.2, T 72.3, T 93.8, T 23.5, T 103.4, AKPIK 4.1, EP 11.1, T 100.1, T 106.5, T 46.2, T 100.3, T 46.4, T 106.3, T 106.4, T 46.5, T 106.7, T 106.9, T 100.6, T 100.8, T 18.2, T 72.2, T 103.5.
Horzela, Maximilian •T 15.3, T 15.7, T 32.7 Hosseini, Sepideh •T 63.6 Höttecke, Dietmar \$YNS 1.1 Hötting, Marius \$T 1.2, •ST 4.3, ST 4.4, ST 4.5, ST 6.2 Hötzsch, Luisa T 22.1 Höu, Wenjie T 12., •T 30.3 Hu, Jiangqiao T 22.3 Hu, Jiangqiao T 4.6 Huber, Thomas T 66.3 Huber, Tobias T 56.3, T 56.6 Hubert, Mario AGPhil 4.3 Hueg, Tim T 12.4, T 13.5, T 40.8, T 40.9, T 42.6, T 95.2, T 97.6 Hügli, Cédrine T 12.4, T 13.5, T 40.8, T 40.9, T 42.6, T 95.2, T 97.6 Hügli, Cédrine T 9.8 Huhmann, Christian T 22.6 Hull, Michael M •DD 10.7, •DD 10.8, •DD 10.10, •DD 10.11, •DD 10.12, •DD 11.9, •DD 18.4 Hummer, Fabian T 40.3, •T 54.2 Hümbert-Schnur, Sebastian DD 33.4 Hummer, Fabian T 48.4, T 40.2, Hundhausen, Daniel AKPIK 4.2 Hünnefeld, Mirco T 18.5, T 46.6 Husemann, Ulrich ST 6.4, T 26.7, T 36.3, T 65.5, T 65.6, T 65.7, T 76.8, T 88.6, T 88.7, T 88.8, T 88.9, T 94.4 Huseynov, Nazim T 35.1, T 35.6, T 35.7 Hütten, Moritz T 17.8 Hymon, Karolin EP 11.8, •T 46.5, T 72.2, T 72.3, T 83.8 Iacobacci, Lorenzo MP 2.4 Ibis, Philipp T 60.6, T 60.7 Ibragimov, Iskander T 18.5, T 46.1, T 98.7, T 46.7, T 93.9, T 106.2, T 72.3, T 93.8, T 23.5, T 103.4, AKPIK 4.1, EP 11.1, T 100.1, T 106.5, T 46.2, T 100.3, T 46.4, T 106.3, T 106.4, T 46.5, T 106.7, T 106.9, T 100.6, T 100.8, T 18.2, T 72.2, T 103.5, T 93.7, T 18.6, T 72.1, T 100.4, T 18.4,

Т 100.2, Т 18.1, Т 18.3, Т 23.6, Т 46.3,
T 100.9, T 106.1, EP 11.8, T 18.7,
EP 11.5, T 72.8
Idaszek, BerenikaT 107.5
Imnor, MargareteDD 11.6
Ingleson, John-Luke
Isele Daniel DD 50 1
Issever, Cigdem T 16.4, T 58.5, T 70.5,
T 70.6
Ita, Harald T 2.2
Ito, Hirotaka EP 1.4
Iuliano, Antonio
Ivanjek, LanaDD 18.2, DD 19.2,
Ivone Francesco
Izraelevitch. Federico
Jabusch, Henrik
Jackson, Paul
Jacobi, Daniel•T 30.6
Jacobs, Hanno•T 73.5
Jafari, Abideh 16.4, 133.5, 18/.4
Jage, Hendrik
Jagfeld Christoph T 11 1 T 11 2
•T 11.3. T 11.4. T 11.5. T 11.6. T 97.7
Jain, Prasham•T 87.3
Jaitly, Annanay•EP 7.9
Jäkel, GunnarT 6.9
Jakob, Johanna•T 22.2
Jakobs, Karl I 8.I, I 94.I
Jamali Alireza
Janardhan Shruthi •T 53.8
Janka. Thomas
Jansen, Hannah•T 13.6
Jansen, Karl •PV X, T 79.8, T 79.9,
AKPIK 2.9
Januschek, Friederike •PV IX
Jaroschewski, Ilja •EP 11.4, EP 11.6,
• 1 83.4, 1 83.6, 1 99.8 Jodamski Marka - DD 9.1 DD 42.2
leidler Maximilian DD 10 10
DD 10.11. DD 18.4
Jelicic, Katarina
Jentzen, Walter ST 1.6
Jeppe, LauridsT 76.7
Jeremy, Mah Zhee Kein EP 2.2
Jevtic Vukan •T 26 6 T 60 9
Jimeno, Doramas•T 68.2
Jimeno, Doramas
Jimeno, Doramas • T 68.2 Joergens, Christopher • DD 36.4 Jomhari, Nur Zulaiha • DD 36.4 Jones, Stephen P T 2.7 Joppe, Robert GR 14.3, • GR 14.4, GR 14.5, GR 14.6 Jörg, Florian T 22.7, T 78.2 Joshi, Vikas T 45.4, T 101.5 Jozsef, Erik • ST 9.6 Jueid, Adil
Jimeno, Doramas • T 68.2 Joergens, Christopher • DD 36.4 Jomhari, Nur Zulaiha T 58.1 Jones, Stephen P T 2.7 Joppe, RobertGR 14.3, •GR 14.4, GR 14.5, GR 14.6 Jörg, Florian T 22.7, T 78.2 Joshi, Vikas T 45.4, T 101.5 Jozsef, Erik •ST 9.6 Jueid, Adil •ST 9.6 Jueid, Adil •DD 15.2,
Jimeno, Doramas
Jimeno, Doramas T 68.2 Joergens, Christopher DD 36.4 Jonhari, Nur Zulaiha T 58.1 Jones, Stephen P T 2.7 Joppe, Robert GR 14.3, •GR 14.4, GR 14.5, GR 14.6 Jörg, Florian T 22.7, T 78.2 Joshi, Vikas T 45.4, T 101.5 Jozsef, Erik ST 9.6 Jueid, Adil T 31.3 Juengst, Heike Elisabeth •DD 15.2, •GP 1.2 Jung, Alexander T 105.10 Jung, Hannes T 32.4 Junk, Andreas •GP 8.3 Junkerkalefeld, Henrik •T 3.3, T 52.1 Junkerkalefeld, Henrik T 20.1, T 48.3, T 99.5, T 99.3, T 20.8 Juraebit Detrick T 65.1 T 65.2
Jimeno, Doramas T 68.2 Joergens, Christopher DD 36.4 Jomhari, Nur Zulaiha T 58.1 Jones, Stephen P T 2.7 Joppe, Robert GR 14.3, •GR 14.4, GR 14.5, GR 14.6 Jörg, Florian T 22.7, T 78.2 Joshi, Vikas T 45.4, T 101.5 Jozsef, Erik ST 9.6 Jueid, Adil T 31.3 Juengst, Heike Elisabeth DD 15.2, •GP 1.2 Jung, Alexander T 105.10 Jung, Hannes T 32.4 Junkerkalefeld, Henrik T 3.3, T 52.1 Junkermann, Thomas T 14.4 JUNO-Kollaboration T 20.1, T 48.3, T 99.5, T 99.3, T 20.8 Juraschitz, Patrick T 65.1, •T 65.3 Just Johannes T 09.8
Jimeno, Doramas
Jimeno, Doramas • T 68.2 Joergens, Christopher • DD 36.4 Jomhari, Nur Zulaiha T 58.1 Jones, Stephen P T 2.7 Joppe, Robert GR 14.3, •GR 14.4, GR 14.5, GR 14.6 Jörg, Florian T 22.7, T 78.2 Joshi, Vikas T 45.4, T 101.5 Jozsef, Erik • ST 9.6 Jueid, Adil T 31.3 Juengst, Heike Elisabeth • DD 15.2, • GP 1.2 Jung, Alexander • T 105.10 Junkerkalefeld, Henrik .•T 3.3, T 52.1 Junkerkalefeld, Henrik .•T 3.3, T 52.1 Junkermann, Thomas • T 14.4, JUNO-Kollaboration T 20.1, T 48.3, T 99.5, T 99.3, T 20.8 Juraschitz, Patrick T 65.1, •T 65.3 Just, Oliver • EP 1.4 Kääpä, Alex
Jimeno, Doramas T 68.2 Joergens, Christopher DD 36.4 Jomhari, Nur Zulaiha T 58.1 Jones, Stephen P T 2.7 Joppe, Robert GR 14.3, •GR 14.4, GR 14.5, GR 14.6 Jörg, Florian T 22.7, T 78.2 Joshi, Vikas T 45.4, T 101.5 Jozsef, Erik ST 9.6 Jueid, Adil T 31.3 Juengst, Heike Elisabeth DD 15.2, •GP 1.2 Jung, Alexander T 105.10 Jung, Hannes T 3.3, T 52.1 Junkermann, Thomas T 14.4 JUNO-Kollaboration T 20.1, T 48.3, T 99.5, T 99.3, T 20.8 Juraschitz, Patrick T 65.1, •T 65.3 Just, Johannes T 73.2 Kabel, Viktoria MP 5.1
Jimeno, Doramas T 68.2 Joergens, Christopher DD 36.4 Jomhari, Nur Zulaiha T 58.1 Jones, Stephen P T 2.7 Joppe, Robert GR 14.3, •GR 14.4, GR 14.5, GR 14.6 Jörg, Florian T 22.7, T 78.2 Joshi, Vikas T 45.4, T 101.5 Jozsef, Erik ST 9.6 Jueid, Adil T 31.3 Juengst, Heike Elisabeth DD 15.2, •GP 1.2 Jung, Alexander T 105.10 Jung, Hannes T 32.4 Junkerkalefeld, Henrik T 3.3, T 52.1 Junkermann, Thomas T 14.4 JUNO-Kollaboration T 20.1, T 48.3, T 99.5, T 99.3, T 20.8 Juraschitz, Patrick T 65.1, •T 65.3 Just, Johannes T 32.4 Just, Oliver F 91.4 Kääpä, Alex T 73.2 Kabel, Viktoria MP 5.1 Kahlhoefer, Felix T 86.6
Jimeno, Doramas T 68.2 Joergens, Christopher DD 36.4 Jonnari, Nur Zulaiha T 58.1 Jones, Stephen P T 2.7 Joppe, Robert GR 14.3, •GR 14.4, GR 14.5, GR 14.6 Jörg, Florian T 22.7, T 78.2 Joshi, Vikas T 45.4, T 101.5 Jozsef, Erik ST 9.6 Jueid, Adil T 31.3 Juengst, Heike Elisabeth •DD 15.2, •GP 1.2 Jung, Alexander T 105.10 Jung, Hannes GR 8.3 Junkerkalefeld, Henrik T 3.3, T 52.1 Junkermann, Thomas T 14.4 JUNO-Kollaboration T 20.1, T 48.3, T 99.5, T 99.3, T 20.8 Juraschitz, Patrick T 65.1, •T 65.3 Just, Johannes T 99.8 Just, Oliver F 9.14 Kääpä, Alex T 85.1 Kabilhoefer, Felix T 86.6 Kaiser, Benedict T 00.9, T 20.9,
Jimeno, Doramas T 68.2 Joergens, Christopher DD 36.4 Jomhari, Nur Zulaiha T 58.1 Jones, Stephen P T 2.7 Joppe, Robert GR 14.3, •GR 14.4, GR 14.5, GR 14.6 Jörg, Florian T 22.7, T 78.2 Joshi, Vikas T 45.4, T 101.5 Jozsef, Erik ST 9.6 Jueid, Adil T 45.4, T 101.5 Jozsef, Erik •ST 9.6 Jueid, Adil T 45.4, T 101.5 Joung, Heike Elisabeth •DD 15.2, •GP 1.2 Jung, Alexander T 105.10 Jung, Hannes T 32.4 Junk, Andreas •GP 8.3 Junkerkalefeld, Henrik T 33.7 52.1 Junkerkalefeld, Henrik T 65.1, •T 65.3 Just, Ohannes T 65.1, •T 65.3 Just, Oliver EP 1.4 Kääpä, Alex T 73.2 Kabel, Viktoria MP 5.1 Kahlhoefer, Felix T 86.6 Kaiser, Benedict T 20.3, T 20.9, T 48.6, T 48.8, •T 102.3
Jimeno, Doramas T 68.2 Joergens, Christopher DD 36.4 Jomhari, Nur Zulaiha T 58.1 Jones, Stephen P T 2.7 Joppe, Robert GR 14.3, •GR 14.4, GR 14.5, GR 14.6 Jörg, Florian T 22.7, T 78.2 Joshi, Vikas T 45.4, T 101.5 Jozsef, Erik ST 9.6 Jueid, Adil T 45.4, T 101.5 Jozsef, Frik ST 9.6 Jueid, Adil T 45.4, T 101.5 Jourgst, Heike Elisabeth DD 15.2, •GP 1.2 Jung, Alexander T 105.10 Jung, Hannes T 33.3, T 52.1 Junkerkalefeld, Henrik T 33.3, T 52.1 Junkerkalefeld, Henrik T 65.1, •T 65.3 Just, Johannes T 65.1, •T 65.3 Just, Johannes T 99.8 Just, Oliver EP 1.4 Kääpä, Alex T 65.3, T 73.2 Kabel, Viktoria MP 5.1 Kahlhoefer, Felix T 86.6 Kaiser, Benedict T 20.3, T 20.9, T 48.6, T 48.8, •T 102.3 Kalaitzidou, Ilia
Jimeno, Doramas • T 68.2 Joergens, Christopher • DD 36.4 Jomhari, Nur Zulaiha T 58.1 Jones, Stephen P 7 2.7 Joppe, Robert GR 14.3, • GR 14.4, GR 14.5, GR 14.6 Jörg, Florian T 22.7, T 78.2 Joshi, Vikas T 45.4, T 101.5 Jozsef, Erik • ST 9.6 Jueid, Adil T 45.4, T 101.5 Jouegst, Heike Elisabeth • DD 15.2, • GP 1.2 Jung, Alexander • T 105.10 Jung, Hannes T 32.4 Junkerkalefeld, Henrik • T 3.3, T 52.1 Junkerkalefeld, Henrik • T 3.3, T 52.1 Junkerkalefeld, Henrik • T 3.3, T 52.1 Junkerkalefeld, Henrik • T 65.1, • T 65.3 Just, Johannes • T 14.4 JUNO-Kollaboration T 20.1, T 48.3, T 99.5, T 99.3, T 20.8 Juraschitz, Patrick T 65.1, • T 65.3 Just, Johannes • T 99.8 Just, Oliver • EP 1.4 Kääpä, Alex T 73.2 Kabel, Viktoria MP 5.1 Kahlhoefer, Felix • MP 11.2, • AGPhil 11.4 Kallonen Kimmo T 95.
Jimeno, Doramas T 68.2 Joergens, Christopher DD 36.4 Jomhari, Nur Zulaiha T 58.1 Jones, Stephen P T 2.7 Joppe, Robert GR 14.3, •GR 14.4, GR 14.5, GR 14.6 Jörg, Florian T 22.7, T 78.2 Joshi, Vikas T 45.4, T 101.5 Jozsef, Erik •ST 9.6 Jueid, Adil T 45.4, T 101.5 Joueid, Adil T 45.4, T 101.5 Juengst, Heike Elisabeth •DD 15.2, •GP 1.2 Jung, Alexander •T 105.10 Jung, Hannes T 3.3, T 52.1 Junkerkalefeld, HenrikT 3.3, T 52.1 Junkerkalefeld, HenrikT 3.3, T 52.1 Junkerkalefeld, HenrikT 3.3, T 52.1 Junkerkalefeld, HenrikT 48.3, T 99.5, T 99.3, T 20.8 Juraschitz, Patrick T 65.1, •T 65.3 Just, Johannes T 99.8 Just, Oliver EP 1.4 Kääpä, Alex T 20.3, T 20.9, T 48.6, T 48.8, •T 102.3 Kalaitzidou, Ilia T 62.4 Kailonen, Kimmo T 18.6, T 14.2
Jimeno, Doramas T 68.2 Joergens, Christopher DD 36.4 Jomhari, Nur Zulaiha T 58.1 Jones, Stephen P T 2.7 Joppe, Robert GR 14.3, •GR 14.4, GR 14.5, GR 14.6 Jörg, Florian T 22.7, T 78.2 Joshi, Vikas T 45.4, T 101.5 Jozsef, Erik ST 9.6 Jueid, Adil T 31.3 Juengst, Heike Elisabeth DD 15.2, •GP 1.2 Jung, Alexander T 105.10 Jung, Hannes T 32.4 Junkerkalefeld, Henrik .•T 3.3, T 52.1 Junkerkalefeld, Henrik .•T 7.3, T 48.3, T 99.5, T 99.3, T 20.8 Juraschitz, Patrick T 65.1, •T 65.3 Just, Oliver EP 1.4 Kääpä, Alex T 73.2 Kabel, Viktoria MP 5.1 Kahlhoefer, Felix T 86.6 Kaiser, Benedict T 20.3, T 20.9, T 48.6, T 48.8, •T 102.3 Kalaitzidou, Ilia T 62.4 Kalies, Grit •MP 11.2, •AGPhil 11.4 Kallonen, Kimmo T 9.5 Kaminski, Jochen T 11.8, T 14.2, T 38.5, T 39.8, T 69.7, T 97.1
Jimeno, Doramas • T 68.2 Joergens, Christopher • DD 36.4 Jomhari, Nur Zulaiha • T 58.1 Jones, Stephen P • T 2.7 Joppe, Robert • GR 14.3, • GR 14.4, GR 14.5, GR 14.6 Jörg, Florian • T 22.7, T 78.2 Joshi, Vikas • T 45.4, T 101.5 Jozsef, Erik • ST 9.6 Jueid, Adil • ST 9.6 Jueid, Adil • ST 9.6 Jueid, Adil • ST 9.6 Jueid, Adil • T 31.3 Juengst, Heike Elisabeth • DD 15.2, • GP 1.2 Jung, Alexander • T 105.10 Jung, Hannes • T 32.4 Junkerkalefeld, Henrik .• T 3.3, T 52.1 Junkermann, Thomas • T 14.4 JUNO-Kollaboration T 20.1, T 48.3, T 99.5, T 99.3, T 20.8 Juraschitz, Patrick T 65.1, • T 65.3 Just, Oliver • EP 1.4 Kääpä, Alex • T 73.2 Kabel, Viktoria MP 5.1 Kahlhoefer, Felix T 86.6 Kaiser, Benedict T 20.3, T 20.9, T 48.6, T 48.8, • T 102.3 Kalaitzidou, Ilia • T 62.4 Kalies, Grit • MP 11.2, • AGPhil 11.4 Kallonen, Kimmo T 91.7 Kaminski, Jochen T 11.8, T 14.2, T 38.5, T 39.8, T 68.1, T 69.7, T 97.1 Kaminski, Matthias • MP 2.3
Jimeno, Doramas T 68.2 Joergens, Christopher DD 36.4 Jomhari, Nur Zulaiha T 58.1 Jones, Stephen P T 2.7 Joppe, Robert GR 14.3, •GR 14.4, GR 14.5, GR 14.6 Jörg, Florian T 22.7, T 78.2 Joshi, Vikas T 45.4, T 101.5 Jozsef, Erik ST 9.6 Jueid, Adil T 22.7, T 78.2 Jung, Adil T 45.4, T 101.5 Jouegst, Heike Elisabeth DD 15.2, •GP 1.2 Jung, Alexander T 105.10 Jung, Hannes GR 8.3 Junkerkalefeld, Henrik T 33.7, 52.1 Junkerkalefeld, Henrik T 65.1, •T 65.3 Just, Johannes T 65.1, •T 65.3 Just, Johannes T 99.8 Just, Oliver F 65.1, *T 65.3 Just, Johannes T 99.8 Just, Oliver T 86.6 Kaiser, Benedict T 20.3, T 20.9, T 48.6, T 48.8, •T 102.3 Kalalizidou, Ilia T 62.4 Kalies, Grit MP 11.2, •AGPhil 11.4, Kallonen, Kimmo T 11.8, T 14.2, T 38.5, T 39.8, T 68.1, T 69.7, T 97.1 Kaminski, Matthias MP 2.3 Kampert, Karl-Heinz T 71.1
Jimeno, Doramas T 68.2 Joergens, Christopher DD 36.4 Jomhari, Nur Zulaiha T 58.1 Jones, Stephen P T 2.7 Joppe, Robert GR 14.3, •GR 14.4, GR 14.5, GR 14.6 Jörg, Florian T 22.7, T 78.2 Joshi, Vikas T 45.4, T 101.5 Jozsef, Erik ST 9.6 Jueid, Adil T 45.4, T 101.5 Jozsef, Erik •ST 9.6 Jueid, Adil T 31.3 Juengst, Heike Elisabeth •DD 15.2, •GP 1.2 Jung, Alexander T 105.10 Jung, Hannes T 32.4 Junk, Andreas •GP 8.3 Junkerkalefeld, Henrik T 33.7 52.1 Junkerkalefeld, Henrik T 65.1, •T 65.3 Just, Johannes T 105.1, •T 65.3 Just, Johannes T 99.8 Just, Oliver EP 1.4 Kääpä, Alex T 73.2 Kabel, Viktoria T 86.6 Kaiser, Benedict T 20.3, T 20.9, T 48.6, T 48.8, •T 102.3 Kalaitzidou, Ilia •T 62.4 Kalies, Grit •MP 11.2, •AGPhil 11.4 Kallhoer, Kimmo T 9.5 Kaminski, Jochen T 11.8, T 14.2, T 38.5, T 39.8, T 68.1, T 69.7, T 97.1 Kaminski, Matthias •MP 2.3 Kampert, Karl-Heinz •T1.1 Kampmann, Philipp T 20.2, T 20.5,
Jimeno, Doramas T 68.2 Joergens, Christopher DD 36.4 Jomhari, Nur Zulaiha T 58.1 Jones, Stephen P T 2.7 Joppe, Robert GR 14.3, •GR 14.4, GR 14.5, GR 14.6 Jörg, Florian T 22.7, T 78.2 Joshi, Vikas T 45.4, T 101.5 Jozsef, Erik •ST 9.6 Jueid, Adil T 45.4, T 101.5 Jozsef, Erik •ST 9.6 Jueid, Adil T 45.4, T 101.5 Jozsef, Heike Elisabeth •DD 15.2, •GP 1.2 Jung, Alexander T 105.10 Jung, Hannes T 3.3, T 52.1 Junkerkalefeld, Henrik T 3.3, T 52.1 Junkerkalefeld, Henrik T 3.3, T 52.1 Junkerkalefeld, Henrik T 65.1, •T 65.3 Just, Johannes T 105.1, •T 65.3 Just, Johannes T 99.8 Just, Oliver EP 1.4 Kääpä, Alex T 73.2 Kabel, Viktoria MP 51.1 Kahlhoefer, Felix *MP 11.2, •AGPhil 11.4 Kalionen, Kimmo T 9.5 Kaminski, Jochen T 11.8, T 14.2, T 38.5, T 39.8, T 68.1, T 69.7, T 97.1 Kaminski, Matthias *MP 2.3 Kaper, Karl-Heinz T 71.1 Kampmann, Philipp T 20.2, T 20.5, T 43.1, T 48.4
Jimeno, Doramas T 68.2 Joergens, Christopher DD 36.4 Jomhari, Nur Zulaiha T 58.1 Jones, Stephen P T 2.7 Joppe, RobertGR 14.3, •GR 14.4, GR 14.5, GR 14.6 Jörg, Florian T 22.7, T 78.2 Joshi, Vikas T 45.4, T 101.5 Jozsef, Erik ST 9.6 Jueid, Adil T 31.3 Juengst, Heike Elisabeth DD 15.2, •GP 1.2 Jung, Alexander T 105.10 Jung, Hannes T 32.4 Junk, Andreas GP 8.3 Junkerkalefeld, Henrik T 3.3, T 52.1 Junkerkalefeld, Henrik T 3.3, T 52.1 Junkerkalefeld, Henrik T 65.1, •T 65.3 Just, Johannes T 20.1, T 48.3, T 99.5, T 99.3, T 20.8 Juraschitz, Patrick T 65.1, •T 65.3 Just, Johannes T 99.8 Just, Oliver EP 1.4 Kääpä, Alex T 20.3, T 20.9, T 48.6, T 48.8, •T 102.3 Kalaitzidou, Ilia 4P 11.2, •AGPhil 11.4 Kallonen, Kimmo T 9.5 Kaminski, Jochen T 11.8, T 14.2, T 38.5, T 39.8, T 68.1, T 69.7, T 97.1 Kaminski, Matthias MP 2.3 Kampert, Karl-Heinz T 100.4, T 100.5 Kanp Donghwa F 100.4, T 100.5 Kanp Donghwa F 100.4, T 100.5 Kanp Senetice P 12.2
Jimeno, Doramas T 68.2 Joergens, Christopher DD 36.4 Jomhari, Nur Zulaiha T 58.1 Jones, Stephen P T 2.7 Joppe, Robert GR 14.3, •GR 14.4, GR 14.5, GR 14.6 Jörg, Florian T 22.7, T 78.2 Joshi, Vikas T 45.4, T 101.5 Jozsef, Erik •ST 9.6 Jueid, Adil T 31.3 Juengst, Heike Elisabeth •DD 15.2, •GP 1.2 Jung, Alexander •T 105.10 Jung, Hannes T 32.4 Junkerkalefeld, Henrik •T 3.3, T 52.1 Junkerkalefeld, Henrik •T 3.3, T 52.1 Juraschitz, Patrick T 65.1, •T 65.3 Just, Johannes •T 99.8 Just, Oliver •EP 1.4 Kääpä, Alex •T 73.2 Kabel, Viktoria MP 5.1 Kahlhoefer, Felix •MP 11.2, •AGPhil 11.4 Kallonen, Kimmo T 20.3, T 20.9, T 48.6, T 48.8, •T 102.3 Kaaninski, Jochen T 11.8, T 14.2, T 38.5, T 39.8, T 68.1, T 69.7, T 97.1 Kaminski, Jochen T 11.8, T 14.2, T 38.5, T 39.8, T 68.1, T 69.7, T 97.1 Kaminski, Jochen T 11.8, T 14.2, T 38.5, T 39.8, T 68.1, T 69.7, T 97.1 Kaminski, Matthias •MP 2.3 Kampert, Karl-Heinz •T 71.1 Kampmann, Philipp T 20.2, T 20.5, T 43.1, T 48.4 Kang, Donghwa •T 100.4, T 100.5 Kapp, Sebastian DD 12.3 Kannes Alexander T 52.5 T 106.5
Jimeno, Doramas T 68.2 Joergens, Christopher DD 36.4 Jomhari, Nur Zulaiha T 58.1 Jones, Stephen P T 2.7 Joppe, Robert GR 14.3, •GR 14.4, GR 14.5, GR 14.6 Jörg, Florian T 22.7, T 78.2 Joshi, Vikas T 45.4, T 101.5 Jozsef, Erik ST 9.6 Jueid, Adil T 31.3 Juengst, Heike Elisabeth DD 15.2, •GP 1.2 Jung, Alexander T 105.10 Jung, Hannes T 32.4 Junkerkalefeld, Henrik .•T 3.3, T 52.1 Junkerkalefeld, Henrik .•T 3.3, T 52.1 Juraschitz, Patrick T 65.1, •T 65.3 Just, Johannes T 99.8 Just, Oliver EP 1.4 Kääpä, Alex T 73.2 Kabel, Viktoria MP 5.1 Kahlhoefer, Felix T 86.6 Kaiser, Benedict T 20.3, T 20.9, T 48.6, T 48.8, •T 102.3 Kalaitzidou, Ilia T 62.4 Kalies, Grit MP 11.2, •AGPhil 11.4 Kallonen, Kimmo T 11.8, T 14.2, T 38.5, T 39.8, T 68.1, T 69.7, T 97.1 Kaminski, Jochen T 118, T 14.2, T 38.5, T 39.8, T 68.1, T 69.7, T 97.1 Kampan, Nehilipp T 20.2, T 20.5, T 43.1, T 48.4 Kang, Donghwa •T 100.4, T 100.5 Kapp, Sebastian DD 12.3 Kappes, Alexander T 52.5, T 106.5, T 106.6, T 106.7, T 106.8
Jimeno, Doramas • T 68.2 Joergens, Christopher • DD 36.4 Jonnari, Nur Zulaiha T 58.1 Jones, Stephen P. T 2.7 Joppe, Robert GR 14.3, •GR 14.4, GR 14.5, GR 14.6 Jörg, Florian Jörg, Florian T 22.7, T 78.2 Joshi, Vikas T 45.4, T 101.5 Jozsef, Erik ST 9.6 Jueid, Adil T 31.3 Juengst, Heike Elisabeth •DD 15.2, •GP 1.2 Jung, Alexander •T 105.10 Jung, Hannes T 32.4 Junk, Andreas •GP 8.3 Junkerkalefeld, Henrik •T 33.7 52.1 Junkermann, Thomas •T 14.4 JUNO-Kollaboration T 20.1, T 48.3, Ju9 5, T 99.3, T 20.8 Juast, Johannes Just, Johannes •T 99.8 Just, Oliver •EP 14 Kääpä, Alex T 73.2 Kabel, Viktoria MP 5.1 Kalaitzidou, Ilia •T 62.4 Kääpä, Alex T 20.3, T 20.9, T 48.6, T 48.8, •T 102.3 Kalaitzidou, Ilia Kalitse, Grit •MP 11.2, •AGPhil 11.4
Jimeno, Doramas • T 68.2 Joergens, Christopher • DD 36.4 Jonnari, Nur Zulaiha T 58.1 Jones, Stephen P. T 2.7 Joppe, Robert GR 14.3, •GR 14.4, GR 14.5, GR 14.6 Jörg, Florian Jozsef, Erik T 22.7, T 78.2 Joshi, Vikas T 45.4, T 101.5 Jozsef, Erik ST 9.6 Jueid, Adil T 31.3 Juengst, Heike Elisabeth •DD 15.2, •GP 1.2 Jung, Alexander T 105.10 Jung, Hannes T 32.4 Junk, Andreas GP 8.3 Junkerkalefeld, Henrik •T 3.3, T 52.1 Junkerkalefeld, Henrik T 65.1, •T 65.3 Just, Johannes T 99.8 Just, Oliver EP 1.4 Kääpä, Alex T 73.2 Kabel, Viktoria MP 5.1 Kalaitzidou, Ilia T 62.4 Kalizzidou, Ilia T 62.4 Kaiser, Benedict T 20.3, T 20.9, T 48.6, T 48.8, •T 102.3 Kaalexi, Matthias Kalaitzidou, Ilia T 62.4 Kalaitzidou, Ilia T
Jimeno, Doramas T 68.2 Joergens, Christopher DD 36.4 Jomhari, Nur Zulaiha T 58.1 Jones, Stephen P T 2.7 Joppe, Robert GR 14.3, •GR 14.4, GR 14.5, GR 14.6 Jörg, Florian T 22.7, T 78.2 Joshi, Vikas T 45.4, T 101.5 Jozsef, Erik ST 9.6 Jueid, Adil T 45.4, T 101.5 Jozsef, Erik ST 9.6 Jueid, Adil T 45.4, T 101.5 Joung, Heike Elisabeth DD 15.2, •GP 1.2 Jung, Alexander T 105.10 Jung, Hannes T 32.4 Junk, Andreas GP 8.3 Junkerkalefeld, Henrik T 3.3, T 52.1 Junkerkalefeld, Henrik T 65.1, •T 65.3 Just, Johannes
Jimeno, Doramas T 68.2 Joergens, Christopher DD 36.4 Jomhari, Nur Zulaiha T 58.1 Jones, Stephen P T 2.7 Joppe, Robert GR 14.3, •GR 14.4, GR 14.5, GR 14.6 Jörg, Florian T 22.7, T 78.2 Joshi, Vikas T 45.4, T 101.5 Jozsef, Erik ST 9.6 Jueid, Adil T 45.4, T 101.5 Jozsef, Erik ST 9.6 Jueid, Adil T 31.3 Juengst, Heike Elisabeth DD 15.2, •GP 1.2 Jung, Alexander T 105.10 Jung, Hannes T 32.4 Junk, Andreas GP 8.3 Junkerkalefeld, Henrik T 3.3, T 52.1 Junkerkalefeld, Henrik T 3.3, T 52.1 Junkerkalefeld, Henrik T 65.1, •T 65.3 Just, Johannes T 105.1, •T 65.3 Just, Johannes T 99.8 Just, Oliver EP 1.4 Kääpä, Alex T 65.1, •T 65.1 Kahlhoefer, Felix T 86.6 Kaiser, Benedict T 20.3, T 20.9, T 48.6, T 48.8, •T 102.3 Kalaitzidou, Ilia •MP 11.2, •AGPhil 11.4 Kallonen, Kimmo T 9.5 Kaminski, Jochen T 11.8, T 14.2, T 38.5, T 39.8, T 68.1, T 69.7, T 97.1 Kaminski, Matthias MP 2.3 Kapp, Sebastian DD 12.3 Kappe, Alexander T 52.5, T 106.5, T 106.6, T 106.7, T 106.8 Kaps, Andreas DD 35.7, •DD 44.1 Karan, Kanioar •EP 11.8, •T 83.8 "Comparison" •KP 11.8, •T 83.8

	T 21.7
Karl, Martina	•T 18.1, T 18.4
Karsten, Florian	DD 5.2
Käs. Stephanie	70.2. •AKPIK 2.3.
AKPIK 2.7	
Kasemann Christoph	Т 77 4
Kasjeczka Gregor	T 37 8 T 37 0
	1 T 02 2
	Ω, 1 90.2, Ω
ANPIN 2.4, ANPIN 4	
Kasper, Jonas	
Kass, Hanno	•DD 5.2
KATRIN-Kollaboration	. T 21.3, T 21.4,
T 21.5, T 21.8, T 21.9	9, T 95.8, T 21.6,
T 49.6, T 49.7, T 49.	8, T 49.9, T 21.7,
T 98.6, T 21.2	
Katzy Judith T 8	3.3. T 89.3. T 92.6
Kaus Carsten	DD 33 4
Kazok Sandra Laura	9T16
Kazek, Saliula Laula	
	ILd 1 /4./,
•1 //.8, 1 /8./, 1 10	4.5, 1 104.0,
1 104.7	
Keaveney, James	
Kegel, Sophie	
Keijzer, Sietske	T 4.4
Keil, Felicitas	T 45.3
Keller. Sebastian	DD 33.4
Keller Stephan	Т 47 4
Kellermann Moritz	•T 104 3
Kempkens Tim	T 21 1
Kempster Jacob	
Kon Kroul	
Konnedy Dhili-	• I O.I ا•
Kennedy, Philip	
Kennedy, Philip David	•T 5.2
Kenzie, Matt	T 3.7
Kepkar, Ankur	EP 5.1
Kerbel, Gabrielle	•AGPhil 10.3
Kerner, Matthias	T 2.7
Kern-Michler, D.	DD 34.3
Kern-Michler, Daniela	•DD 44.2,
DD 51.1	
Khalil Mohammed	GR 9 2
Khodiamirian Alexand	ler T 56 9
Khoo Teng lian	Т 16 Л
Khoriauli Cia	т 47 0
Kiloliduli, Gid	I 07.2
Кіеск, тот	
kiefer, claus	GR 3.0
Kieseler, Jan	I /.5
Kieser, Fabian	0 10 2
	•DD 10.3
Kiesling, Christian .	T 30.5, T 43.8
Kiesling, Christian . Kilian, Schwarz	T 30.5, T 43.8 •EP 2.2
Kiesling, Christian . Kilian, Schwarz Kilian, Wolfgang	•DD 10.3 •T 30.5, T 43.8 •EP 2.2 •T 31.2
Kiesling, Christian Kilian, Schwarz Kilian, Wolfgang Kilminster, Ben	•DD 10.3 T 30.5, T 43.8 •EP 2.2 T 31.2 T 66.9
Kiesling, Christian Kilian, Schwarz Kilian, Wolfgang Kilminster, Ben Kim, Dohun	•DD 10.3 T 30.5, T 43.8 •EP 2.2 T 31.2 T 66.9 •T 94.7
Kiesling, Christian Kilian, Schwarz Kilian, Wolfgang Kilminster, Ben Kim, Dohun Kim, Ye Joon	• • • • • • • • • • • • • • • • • • •
Kiesling, Christian Kilian, Schwarz Kilian, Wolfgang Kilminster, Ben Kim, Dohun Kim, Ye Joon Kim, Yoonyoung	•DD 10.3 T 30.5, T 43.8 •EP 2.2 T 31.2 T 66.9 T 94.7
Kiesling, Christian Kilian, Schwarz Kilian, Wolfgang Kim, Inster, Ben Kim, Dohun Kim, Ye Joon Kim, Yoonyoung Kinast. Angelina	• DD 10.3 T 30.5, T 43.8 • EP 2.2 T 31.2 T 66.9 • T 94.7 • T 8.4 EP 3.7 • T 104.5, T 104.5
Kiesling, Christian Kilian, Schwarz Kilian, Wolfgang Kiminster, Ben Kim, Dohun Kim, Ye Joon Kim, Ye Joon Kim, Yoonyoung Kinast, Angelina T 104.7	•DD 10.3 •EP 2.2 T 31.2 T 66.9 •T 94.7 •T 8.4 EP 3.7 •T 104.5, T 104.6,
Kiesling, Christian Kilian, Schwarz Kilian, Wolfgang Kim, Dohun Kim, Dohun Kim, Ye Joon Kim, Yoonyoung Kinast, Angelina T 104.7 King Andreas	• DD 10.3 T 30.5, T 43.8
Kiesling, Christian Kilian, Schwarz Kilian, Wolfgang Kilminster, Ben Kim, Dohun Kim, Ye Joon Kinast, Angelina T 104.7 King, Andreas King Markus	• DD 10.3 T 30.5, T 43.8 • EP 2.2 T 31.2 T 66.9 • T 94.7 • T 84. EP 3.7 • T 104.5, T 104.6, GR 2.2 • GR 2.2
Kiesling, Christian Kilian, Schwarz Kilian, Wolfgang Kim, Dohun Kim, Dohun Kim, Ye Joon Kim, Ye Joon Kinast, Angelina T 104.7 King, Andreas King, Markus King, Markus	• DD 10.3 T 30.5, T 43.8 • EP 2.2 T 31.2 T 66.9 • T 94.7 • T 8.4 EP 3.7 • T 104.5, T 104.6, GR 2.2 • GR 2.2 • T 66.2
Kiesling, Christian Kilian, Schwarz Kilian, Wolfgang Kilminster, Ben Kim, Dohun Kim, Ye Joon Kinast, Angelina T 104.7 King, Andreas King, Markus King, Montague	•DD 10.3 •EP 2.2 T 30.5, T 43.8 •EP 2.2 T 31.2 T 66.9 •T 94.7 •T 8.4 EP 3.7 •T 104.5, T 104.6, GR 2.2 •GR 2.2 •GR 2.2 •GR 2.2
Kiesling, Christian Kilian, Schwarz Kilian, Wolfgang Kim, Dohun Kim, Ye Joon Kim, Ye Joon Kinast, Angelina T 104.7 King, Andreas King, Markus King, Montague King, Monty	•DD 10.3 T 30.5, T 43.8 EP 2.2 T 31.2 T 66.9 T 94.7 T 94.7 T 94.7 T 94.7 T 94.7 T 94.7 T 94.7 T 104.5, T 104.6, GR 2.2 GR 2.2 T 66.7 T 66.7 T 66.6
Kiesling, Christian Kilian, Schwarz Kilian, Wolfgang Kim, Dohun Kim, Ye Joon Kim, Ye Joon Kim, Yoonyoung Kim, Yoonyoung King, Andreas King, Andreas King, Markus King, Montague King, Pietro	• DD 10.3
Kiesling, Christian Kilian, Schwarz Kilian, Wolfgang Kim, Dohun Kim, Dohun Kim, Ye Joon Kim, Ye Joon Kinast, Angelina T 104.7 King, Andreas King, Markus King, Montague King, Montague King, Pietro Kinchoff, Andreas	•DD 10.3 •EP 2.2 T 30.5, T 43.8 •EP 2.2 T 31.2 •T 66.9 •T 94.7 •T 8.4 EP 3.7 •T 104.5, T 104.6, GR 2.2 •GR 2.2 •GR 2.2 •GR 2.2 •GR 2.2 •GR 2.2 •T 66.7 T 66.6 T 21.1, T 21.8 •T 12.6, •T 33.8 •T 12.6, •T 33.8
Kiesling, Christian Kilian, Schwarz Kilian, Wolfgang Kim, Nolfgang Kim, Dohun Kim, Ye Joon Kim, Ye Joon Kinast, Angelina T 104.7 King, Andreas King, Markus King, Montague King, Pietro Kirchhoff, Andreas	•DD 10.3 T 30.5, T 43.8 •EP 2.2 T 31.2 T 66.9 T 94.7 T 94.7 T 94.7 T 94.7 T 94.7 T 94.7 T 104.5, T 104.6, GR 2.2 GR 2.2
Kiesling, Christian Kilian, Schwarz Kilian, Wolfgang Kim, Dohun Kim, Ye Joon Kim, Ye Joon Kinast, Angelina T 104.7 King, Andreas King, Montague King, Montague King, Montague King, Montague King, Pietro Kirchhoff, Andreas Kirfel, Florian	•DD 10.3 T 30.5, T 43.8 •EP 2.2 T 31.2 T 66.9 T 94.7 •T 94.7 •T 94.7 •T 8.4 EP 3.7 •T 104.5, T 104.6, GR 2.2 GR 2.2 GR 2.2 GR 2.2
Kiesling, Christian Kilian, Schwarz Kilian, Wolfgang Kim, Volfgang Kim, Dohun Kim, Ye Joon Kim, Ye Joon Kim, Ye Joon Kim, Ye Joon King, Andreas King, Markus King, Montague King, Montague King, Pietro Kirchhoff, Andreas Kirfel, Florian Kirschenmann, Henni	• DD 10.3 T 30.5, T 43.8 EP 2.2 T 31.2 T 66.9 T 94.7 • T 94.7 • T 94.7 • T 104.5, T 104.6, GR 2.2 GR 2.2 GR 2.2 GR 2.2 T 66.7 T 66.6 T 21.1, T 21.8 •T 102.6, • T 33.8 •T 102.7 T 45.1 ngT 9.5
Kiesling, Christian Kilian, Schwarz Kilian, Wolfgang Kilminster, Ben Kim, Dohun Kim, Ye Joon Kim, Ye Joon Kinast, Angelina T 104.7 King, Andreas King, Markus King, Markus King, Montague King, Montague King, Pietro Kirchhoff, Andreas Kirfel, Florian Kirk, John Kirschenmann, Hennii Kirstein, Jürgen	•DD 10.3
Kiesling, Christian Kilian, Schwarz Kilian, Wolfgang Kim, Dohun Kim, Dohun Kim, Ye Joon Kim, Ye Joon Kinast, Angelina T 104.7 King, Andreas King, Markus King, Montague King, Pietro Kirchhoff, Andreas Kirfel, Florian Kirk, John Kirschenmann, Hennii Kirstein, Jürgen	•DD 10.3 T 30.5, T 43.8 •EP 2.2 T 31.2 T 66.9 •T 94.7 •T 94.7 •T 94.7 •T 104.5, T 104.6, GR 2.2 GR 2.2 GR 2.2 T 66.7 T 66.6 T 21.1, T 21.8 T 102.6, •T 33.8 T 107.7 T 45.1 ngT 9.5 D 16.2 D 35.13, •DD 41.2
Kiesling, Christian Kilian, Schwarz Kilian, Wolfgang Kim, Dohun Kim, Ye Joon Kim, Ye Joon Kim, Yoonyoung King, Yoonyoung T 104.7 King, Andreas King, Markus King, Montague King, Montague King, Pietro Kirchhoff, Andreas Kirfel, Florian Kirschenmann, Henni Kirstein, Jürgen Kirstein, Jürgen	•DD 10.3 T 30.5, T 43.8 EP 2.2 T 31.2 T 66.9 T 94.7 •T 94.7 •T 94.7 •T 94.7 •T 94.7 T 94.7 T 94.7 T 104.6, GR 2.2 GR 2.2 GR 2.2 GR 2.2 GR 2.2 GR 2.2 T 66.7 T 66.6 T 21.1, T 21.8 T 12.6, •T 33.8 T 107.7 T 45.1 ngT 9.5 D 16.2 D 35.13, •DD 41.2 EP 6.7
Kiesling, Christian Kilian, Schwarz Kilian, Wolfgang Kim, Dohun Kim, Dohun Kim, Ye Joon Kim, Ye Joon Kim, Ye Joon King, Andreas King, Andreas King, Markus King, Montague King, Montague King, Montague King, Pietro Kirchhoff, Andreas Kirfel, Florian Kirschenmann, Henni Kirstein, Jürgen Kirstein, Jurgen Kirstafan Kissenbeck, Andreas	•DD 10.3 T 30.5, T 43.8 EP 2.2 T 31.2 T 66.9 T 94.7 •T 94.7 •T 94.7 •T 94.7 •T 104.5, T 104.6, GR 2.2 GR 2.1 GR 2.2 GR 2.2
Kiesling, Christian Kilian, Schwarz Kilian, Wolfgang Kim, Dohun Kim, Dohun Kim, Ye Joon Kim, Ye Joon Kinast, Angelina T 104.7 King, Andreas King, Markus King, Markus King, Montague King, Montague King, Pietro Kirchhoff, Andreas Kirfel, Florian Kirschenmann, Hennii Kirstein, Jürgen Kirsud, Dustin Kirsud, Dustin Kissenbeck, Andreas Kissemann, Ralf	•DD 10.3 •EP 2.2 T 30.5, T 43.8 •EP 2.2 T 31.2 •F 94.7 •T 94.7 •G 8.2 •G 8.2.2 •G 7.2.2 •G 7.2.2 •G 7.2.2 •G 8.2.2 •G 8.2.2 •D 10.2.2 •D 10.2.2 •D 2.2.2 •D 2.
Kiesling, Christian Kilian, Schwarz Kilian, Wolfgang Kim, Dohun Kim, Dohun Kim, Ye Joon Kim, Ye Joon Kinast, Angelina T 104.7 King, Andreas King, Markus King, Markus King, Montague King, Pietro Kirchhoff, Andreas Kirfel, Florian Kirk, John Kirschenmann, Hennii Kirstein, Jürgen Kirwald, Dustin Kissenbeck, Andreas Kissenbeck, Andreas Kissenbeck, Andreas Kissenbeck, Andreas	•DD 10.3 T 30.5, T 43.8 EP 2.2 T 31.2 T 66.9 T 94.7 T 104.6, GR 2.2 GR 2.2 GR 2.2 T 66.7 T 66.7 T 66.6 T 211, T 21.8 T 12.6,T 33.8 T 107.7 T 45.1 ngT 9.5 D 16.2 D 16.2 D 16.2 D 16.2 D 16.7 D 16.7
Kiesling, Christian Kilian, Schwarz Kilian, Schwarz Kilminster, Ben Kim, Dohun Kim, Ye Joon Kim, Ye Joon Kim, Ye Joon Kim, Yoonyoung Ti04.7 King, Andreas King, Markus King, Montague King, Montague King, Montague King, Pietro Kirchhoff, Andreas Kirfel, Florian Kirschenmann, Henni Kirstein, Jürgen Kirstein, Jürgen Kirstein, Jürgen Kirstein, Jürgen Kirstein, Jürgen Kirstein, Andreas Kissenbeck, Andreas Kissenbeck, Andreas Kisamann, Ralf Kitagawa, Hussain Kivernyik, Oleh	•DD 10.3 T 30.5, T 43.8 EP 2.2 T 31.2 T 66.9 T 94.7 •T 94.7 •T 94.7 •T 94.7 T 94.7 T 94.7 T 94.7 T 94.7 GR 2.2 GR 2.2
Kiesling, Christian Kilian, Schwarz Kilian, Wolfgang Kilminster, Ben Kim, Dohun Kim, Ye Joon Kim, Ye Joon King, Angelina T 104.7 King, Andreas King, Markus King, Markus King, Montague King, Montague Kirchhoff, Andreas Kirfel, Florian Kirk, John Kirschenmann, Hennii Kirstein, Jürgen Kirstein, Jürgen Kirstefan Kissenbeck, Andreas Kissmann, Ralf Kissenin, Kielen Kirvernyik, Oleh Kiwit, Florian	•DD 10.3 •EP 2.2 T 30.5, T 43.8 •EP 2.2 T 31.2 T 66.9 •T 94.7 •T 8.4 EP 3.7 •T 104.5, T 104.6,
Kiesling, Christian Kilian, Schwarz Kilian, Wolfgang Kim, Nolfgang Kim, Dohun Kim, Ye Joon Kim, Ye Joon Kinast, Angelina T 104.7 King, Andreas King, Markus King, Markus King, Montague King, Montague Kirchhoff, Andreas Kirfel, Florian Kirschenmann, Henni Kirschenmann, Henni Kirstein, Jürgen Kirstein, Jürgen Kis Stefan Kissenbeck, Andreas Kistagawa, Hussain Kivernyik, Oleh Kiwet, Florian Kiwet, Florian Kikut, Florian Kiwet, Florian Kiwit, Florian Kiwit, Florian	•DD 10.3 •EP 2.2 T 30.5, T 43.8 •EP 2.2 T 31.2 T 66.9 •T 94.7 •T 8.4 •EP 3.7 •T 104.6, GR 2.2 •GR 2.2 •DR 10.2 •DD 10.2 •DP 2.2 •EP 6.7 •DD 2.2 •T 10.9 •T 5.1 •T 5.1 •T 64.5 ·T 13.3, T 66.1
Kiesling, Christian Kilian, Schwarz Kilian, Wolfgang Kim, Dohun Kim, Dohun Kim, Ye Joon Kim, Ye Joon Kinast, Angelina T 104.7 King, Andreas King, Markus King, Montague King, Montague King, Pietro Kirchhoff, Andreas Kirfel, Florian Kirschenmann, Henni Kirschenmann, H	•DD 10.3 T 30.5, T 43.8 EP 2.2 T 31.2 T 66.9 T 94.7 T 94.7 T 94.7 T 94.7 T 94.7 T 104.6, GR 2.2 GR 2.2 GR 2.2 GR 2.2 T 66.7 T 66.6 T 21.1, T 21.8 T 12.6, -T 33.8 T 10.7, T 45.1 ngT 9.5 DD 16.2 D 16.2 D 16.7 DD 16.2 DD 10.2 DD
Kiesling, Christian Kilian, Schwarz Kilian, Wolfgang Kim, Wolfgang Kim, Dohun Kim, Ye Joon Kim, Ye Joon Kim, Ye Joon Kim, Ye Joon King, Andreas T 104.7 King, Andreas King, Montague King, Montague King, Montague King, Montague King, Pietro Kirchhoff, Andreas Kirfel, Florian Kirschenmann, Henni Kirstein, Jürgen Kirstein, Jürgen Kirstein, Jürgen Kirstein, Jürgen Kirstein, Jürgen Kissenbeck, Andreas Kissenbeck, Andreas Kissenbeck, Andreas Kissenbeck, Andreas Kissenbeck, Andreas Kissenbeck, Andreas Kissenbeck, Andreas Kissenbeck, Andreas Kissmann, Ralf Kiwit, Florian Kiwit, Florian Kiwit, Florian Kiwit, Florian Kianner, Robert Klanner, Robert	•DD 10.3 T 30.5, T 43.8 EP 2.2 T 31.2 T 66.9 T 94.7 •T 94.7 •T 94.7 •T 94.7 T 94.7 T 94.7 T 104.6, GR 2.2 GR 2.
Kiesling, Christian Kilian, Schwarz Kilian, Wolfgang Kilminster, Ben Kim, Dohun Kim, Ye Joon Kim, Ye Joon Kinast, Angelina T 104.7 King, Andreas King, Markus King, Markus King, Montague King, Montague Kirchhoff, Andreas Kirfel, Florian Kirschenmann, Henni Kirschenmann, Henni Kirstein, Jürgen Kirschenmann, Henni Kirstein, Jürgen Kirsehenkan, Hussain Kissenbeck, Andreas Kissmann, Ralf Kitagawa, Hussain Kivernyik, Oleh Kiwit, Florian Kiwit, Florian Kikut, Florian Kitagawa, Hussain Kitagawa, Hussain Kivernyik, Oleh Kiwit, Florian Kiautke, Franziska Kleimann, Jens Kleim, Daniel	•DD 10.3 •EP 2.2 T 30.5, T 43.8 •EP 2.2 T 31.2 •F 94.7 •T 94.7 •T 94.7 •T 94.7 •T 94.7 •T 94.7 •T 94.7 •T 94.7 •F 94
Kiesling, Christian Kilian, Schwarz Kilian, Wolfgang Kilminster, Ben Kim, Dohun Kim, Ye Joon Kim, Yoonyoung Kinast, Angelina T 104.7 King, Andreas King, Markus King, Markus King, Montague King, Montague King, Montague King, Montague King, Pietro Kirchhoff, Andreas Kirfel, Florian Kirschenmann, Hennin Kirschenmann, Hennin Kirschenmann, Hennin Kirstein, Jürgen Kissenbeck, Andreas Kissenbeck, Andreas Kissenbeck, Andreas Kistagawa, Hussain Kivernyik, Oleh Kiwait, Florian Kivernyik, Oleh Kiwait, Florian Kiautke, Franziska Kleimann, Jens Klein, Daniel	•DD 10.3 T 30.5, T 43.8 EP 2.2 T 31.2 T 66.9 T 94.7 T 94.7 T 94.7 T 94.7 T 94.7 T 104.5, T 104.6, GR 2.2 GR 2.2 GR 2.2 GR 2.2 GR 2.2 T 66.7 T 66.7 T 66.7 T 66.7 T 66.7 T 107.7 T 45.1 ng T 9.5 DD 16.2 DD 22.2 EP 6.7 DD 2.2 EP 6.7 T 10.9 T 5.1 T 10.3, T 66.1 T 10.9 T 5.1 T 13.3, T 66.1 DD 24.1, DD 33.4 EP 12.3 T 38.1, T 38.7 T 13.5 T 65.1 EP 12.3 T 38.1, T 38.7 T 10.7 T 65.1 T 13.3, T 7 55.1 T 10.5 T 13.3, T 7 55.1 T 10.5 T 13.3, T 7 55.1 T 10.5 T 10
Kiesling, Christian Kilian, Schwarz Kilian, Schwarz Kilminster, Ben Kim, Dohun Kim, Ye Joon Kim, Ye Joon Kim, Ye Joon King, Andreas T 104.7 King, Andreas King, Monty King, Monty King, Monty King, Monty King, Pietro Kirchhoff, Andreas Kirfel, Florian Kirk, John Kirschenmann, Henni Kirstein, Jürgen Kirstein, Jürgen Kissenbeck, Andreas Kissmann, Ralf Kivernyik, Oleh Kiwit, Florian Klanner, Robert Klanner, Robert Klein, Daniel Klein, Katja T 654 T 94 6	•DD 10.3 T 30.5, T 43.8 EP 2.2 T 31.2 T 66.9 T 94.7 T 94.7 T 94.7 T 94.7 T 104.5, T 104.6, GR 2.2 GR 2.2 GR 2.2 GR 2.2 T 66.7 T 66.6 T 21.1, T 21.8 T 10.7, T 45.1 ngT 9.5 DD 16.2 D 16.2 D 16.2 EP 6.7 DD 2.2 EP 6.7 T 13.3, T 66.1 T 13.3, T 66.1 DD 24.1, DD 33.4 EP 12.3 T 38.1, T 38.7 5.1, T 65.2, T 65.3,
Kiesling, Christian Kilian, Schwarz Kilian, Wolfgang Kilian, Wolfgang Kim, Dohun Kim, Dohun Kim, Ye Joon Kim, Ye Joon King, Andreas King, Andreas King, Markus King, Markus King, Montague King, Montague Kirchhoff, Andreas Kirfel, Florian Kirschenmann, Henni Kirstein, Jürgen Kirschenmann, Henni Kirstein, Jurgen Kirstefan Kissenbeck, Andreas Kissmann, Ralf Kissenbeck, Andreas Kissmann, Ralf Kissenbeck, Andreas Kissmann, Ralf Kisur, Florian Kivernyik, Oleh Kiwit, Florian Kiwat, Florian Kivernyik, Oleh Kiwit, Florian Kianner, Robert Klautke, Franziska Kleim, Daniel Klein, Katja T 65.4, T 94.6	•DD 10.3 •EP 2.2 T 30.5, T 43.8 •EP 2.2 T 31.2 •F 9.7 •T 94.7 •T 94.7 •T 104.5, T 104.6, •GR 2.2 •GR 3.3 •DD 41.2 •DD 2.2 ·GR 6.7 •T 10.9 ·T 5.1 •DD 2.2 ···································
Kiesling, Christian Kilian, Schwarz Kilian, Wolfgang Kilminster, Ben Kim, Dohun Kim, Ye Joon Kim, Ye Joon Kinast, Angelina T 104.7 King, Andreas King, Markus King, Montague King, Montague Kirg, Montague Kirg, Horian Kirchhoff, Andreas Kirfel, Florian Kirschenmann, Henni Kirstein, Jürgen Kirstein, Jürgen Kirstein, Jürgen Kissenbeck, Andreas Kissmann, Ralf Kissenbeck, Andreas Kissmann, Ralf Kistagawa, Hussain Kivernyik, Oleh Kissenbeck, Andreas Kismann, Ralf Kitagawa, Hussain Kivernyik, Oleh Kistein, Jens Kikitagawa, Hussain Kivernyik, Oleh Kiautke, Franziska Kleimann, Jens Klein, Katja T 65.4, T 94.6 Klein, Lucas	•DD 10.3
Kiesling, Christian Kilian, Schwarz Killian, Schwarz Kilminster, Ben Kim, Dohun Kim, Ye Joon Kim, Ye Joon Kim, Yoonyoung Ti 104.7 King, Andreas King, Andreas King, Montague King, Montague King, Montague King, Montague King, Montague King, Pietro Kirchhoff, Andreas Kirfel, Florian Kirkschenmann, Henni Kirschenmann, Henni Kirschen, Jürgen Kirschenmann, Henni Kirstein, Jürgen Kissenbeck, Andreas Kissenbeck, Andreas Kissenbeck, Andreas Kissenbeck, Andreas Kissenbeck, Andreas Kissenbeck, Andreas Kissenbeck, Andreas Kissenbeck, Andreas Kissenbeck, Andreas Kissenbeck, Andreas Kismann, Ralf Kivernyik, Oleh Kiwit, Florian Klautke, Franziska Kleimann, Jens Klein, Katja T 65.4, T 94.6 Klein, Lucas	•DD 10.3
Kiesling, Christian Kilian, Schwarz Kilian, Wolfgang Kilminster, Ben Kim, Dohun Kim, Ye Joon Kim, Ye Joon Kim, Ye Joon King, Andreas T 104.7 King, Andreas King, Montague King, Montague King, Montague King, Montague King, Pietro Kirchhoff, Andreas Kirfel, Florian Kirschenmann, Henni Kirstein, Jürgen Kirstein, Jürgen Kirstein, Jürgen Kissenbeck, Andreas Kissenbeck, Andreas Kisternyik, Oleh Kiwit, Florian Kivernyik, Oleh Kiwit, Florian Kivernyik, Oleh Kiwit, Florian Kikein, Aussain Klein, Daniel Klein, Katja T 65.4, T 94.6 Klein, Pascal	•DD 10.3 •EP 2.2 T 30.5, T 43.8 •EP 2.2 T 31.2 •F 9.7 •T 94.7 •T 104.5, T 104.6, •GR 2.2 •GR 4.1 •DD 10.2 •T 88.4 DD 10.2, DD 10.4, D 32.2, DD 34.1,
Kiesling, Christian Kilian, Schwarz Kilian, Wolfgang Kilminster, Ben Kim, Dohun Kim, Ye Joon Kim, Ye Joon Kinast, Angelina T 104.7 King, Andreas King, Markus King, Montague King, Montague King, Montague King, Montague Kirchhoff, Andreas Kirfel, Florian Kirk, John Kirschenmann, Henni Kirstein, Jürgen Kirsden Andreas Kissenbeck, Andreas	•DD 10.3 •EP 2.2 T 30.5, T 43.8 •EP 2.2 T 31.2 •F 66.9 •T 94.7 •T 94.7 •T 94.7 •T 94.7 •T 94.7 •T 94.7 •T 94.7 •F 94.7 •F 8.4 EP 3.7 •T 104.5, T 104.6, GR 2.2 •GR 3.8 •T 10.7 •DD 16.2 ·DD 10.2 ·DD 24.1, DD 33.4 ·DD 10.2, DD 10.4, D 32.2, DD 34.1, ·
Kiesling, Christian Kilian, Schwarz Kilian, Wolfgang Kilminster, Ben Kim, Dohun Kim, Ye Joon Kim, Ye Joon Kinast, Angelina T 104.7 King, Andreas King, Markus King, Montague King, Montague Kirchhoff, Andreas Kirfel, Florian Kirschenmann, Henni Kirstein, Jürgen Kissenbeck, Andreas Kissembeck, Andreas Kissembeck, Andreas Kistagawa, Hussain Kivernyik, Oleh Kikautke, Franziska Kleimann, Jens Klein, Daniel T 65.4, T 94.6 Klein, Lucas Klein, Pascal DD 10.5, DD 25.2, D DD 34.9, DD 36.1	•DD 10.3 •EP 2.2 T 30.5, T 43.8 •EP 2.2 T 31.2 T 66.9 •T 94.7 •T 94.7 •GR 2.2 •GR 4.5 ·T 10.7 ·T 10.9 ·T 5.1 •DD 24.1, DD 33.4 EP 12.3 ·T 38.1, T 38.7 ·T 7.5, T 65.2, T 65.3, •T 88.4 DD 10.2, DD 10.4, D 32.2, DD 34.1, ·T 24.2,
Kiesling, Christian Kilian, Schwarz Kilian, Wolfgang Kilminster, Ben Kim, Dohun Kim, Ye Joon Kim, Ye Joon Kim, Ye Joon King, Andreas T 104.7 King, Andreas King, Montague King, Montague King, Montague King, Montague King, Pietro Kirchhoff, Andreas Kirchhoff, Andreas Kirchhoff, Andreas Kirschenmann, Henni Kirstein, Jürgen Kirstein, Jürgen Kirstein, Jürgen Kirstein, Jürgen Kirstein, Jürgen Kirstein, Jürgen Kirstein, Jürgen Kirstein, Jürgen Kirstein, Jürgen Kissenbeck, Andreas Kissenbeck, Andreas Kissenbeck, Andreas Kissenbeck, Andreas Kissenbeck, Andreas Kisennan, Alf Kitagawa, Hussain Kivernyik, Oleh Kiwit, Florian Kivernyik, Oleh Kiein, Lucas Klein, Lucas Klein, Pascal D 10.5, DD 25.2, D D D 34.9, DD 36.1	•DD 10.3 T 30.5, T 43.8 EP 2.2 T 31.2 T 66.9 T 94.7 T 94.7 T 94.7 T 94.7 T 104.5, T 104.6, GR 2.2 GR 2.2 GR 2.2 GR 2.2 GR 2.2 T 66.7 T 66.7 T 66.6 T 21.1, T 21.8 T 10.4, 5, T 104.6, T 10.7, T 45.1 ng T 9.5 DD 16.2 EP 6.7 DD 16.2 EP 6.7 T 10.3, T 66.1 T 10.3, T 66.1 T 10.3, T 65.2, T 65.3, T 88.4 DD 10.2, DD 10.4, D 32.2, DD 34.1, T 24.2,
Kiesling, Christian Kilian, Schwarz Kilian, Wolfgang Kilminster, Ben Kim, Dohun Kim, Ye Joon Kim, Ye Joon Kim, Ye Joon King, Andreas King, Andreas King, Markus King, Montague King, Montague King, Montague King, Pietro Kirchhoff, Andreas Kirfel, Florian Kirschenmann, Henni Kirstein, Jürgen Kirschenmann, Henni Kirstein, Jürgen Kirstefan Kissenbeck, Andreas Kissenbeck, Andreas Kissenbeck, Andreas Kissenbeck, Andreas Kissenbeck, Andreas Kissenbeck, Andreas Kissenbeck, Andreas Kissmann, Ralf Kiwald, Dustin Kivernyik, Oleh Kiwit, Florian Kivernyik, Oleh Kiwit, Florian Klein, Daniel Klein, Katja T 65.4, T 94.6 Klein, Lucas Klein, Pascal DD 10.5, DD 25.2, D D3.4.9, DD 36.1 Klein-Bösing, Christian DD 20.3 Kliem, Bernhard	•DD 10.3 •EP 2.2 T 30.5, T 43.8 •EP 2.2 T 31.2 •F 66.9 •T 94.7 •T 794.7 •T 104.5, T 104.6, •GR 2.2 •GR 4.5 •DD 10.2 •DD 2.2 •GR 4.5 •GR 2.2 •GR 4.5 •GR 4.5 •GR 2.2 •GR 4.5 •GR 4.5 •GR 2.3 •GR 4.5 •GR 2.3 •GR 4.5 •GR 4.5 •GR 4.5 •GR 2.3 •GR 4.5 •GR 4.5 •
Kiesling, Christian Kilian, Schwarz Kilian, Wolfgang Kilminster, Ben Kim, Dohun Kim, Ye Joon Kim, Ye Joon Kim, Ye Joon King, Andreas King, Angelina T 104.7 King, Andreas King, Montague King, Montague King, Montague Kirg, Montague Kirg, Horian Kirchhoff, Andreas Kirfel, Florian Kirschenmann, Henni Kirschenmann, Henni Kirstein, Jürgen Kirwald, Dustin Kirstein, Jürgen Kissenbeck, Andreas Kissmann, Ralf Kissenbeck, Andreas Kissmann, Ralf Kitagawa, Hussain Kivernyik, Oleh Kissenbeck, Andreas Kishen, Katja Klein, Katja Klein, Katja Klein, Katja Klein, Katja T 65.4, T 94.6 Klein, Lucas Klein, Pascal DD 10.5, DD 25.2, D DD 34.9, DD 36.1 Klein-Bösing, Christian DD 20.3 Kliem, Bernhard Kling, Felix	•DD 10.3 •EP 2.2 T 30.5, T 43.8 •EP 2.2 T 31.2 T 66.9 •T 94.7 •T 94.7 •GR 2.2 •GR 4.1 •DD 4.1 •DD 4.1 •D 3.4 ·G 4.5 ·G 5.3 ····································
Kiesling, Christian Kilian, Schwarz Killian, Schwarz Kilminster, Ben Kim, Yoonyoung Kim, Yoonyoung T 104.7 King, Andreas King, Montague King, Montague King, Montague King, Montague King, Montague King, Montague King, Montague King, Pietro Kirchhoff, Andreas Kirfel, Florian Kirk, John Kirschenmann, Henni Kirstein, Jürgen Kirstein, Jürgen Kissenbeck, Andreas Kissmann, Ralf Kisennyik, Oleh Kiwit, Florian Klanner, Robert Klanner, Robert Klein, Lucas Klein, Lucas Klein, Pascal DD 10.5, DD 25.2, D DD 34.9, DD 36.1 Klein-Bösing, Christian DD 20.3 Kliem, Bernhard Kling, Felix	•DD 10.3 T 30.5, T 43.8 EP 2.2 T 31.2 T 66.9 T 94.7 T 94.7 T 94.7 T 94.7 T 94.7 T 104.5, T 104.6, GR 2.2 GR 2.2 GR 2.2 GR 2.2 GR 2.2 GR 2.2 GR 2.2 T 104.5, T 104.6, T 104.5, T 104.6, T 104.5, T 104.6, T 105, T 104.6, T 107, T 45.1 ng T 95. DD 16.2 DD 16.2 EP 6.7 DD 2.2 EP 6.7 DD 2.2 EP 6.7 DD 2.2 EP 6.7 T 10.9 T 5.1 T 13.3, T 66.1 EP 12.3 T 38.1, T 38.7 T 88.4 DD 10.2, DD 10.4, D 32.2, DD 34.1, T 24.2,
Kiesling, Christian Kilian, Schwarz Kilian, Wolfgang Kim, Ponyoung Kim, Ye Joon Kim, Ye Joon Kim, Ye Joon Kim, Ye Joon Kim, Ye Joon King, Andreas King, Andreas King, Montague King, Montague King, Montague King, Montague Kirchhoff, Andreas Kirchhoff, Andreas Kirfel, Florian Kirk, John Kirschenmann, Henni Kirstein, Jürgen Kirstein, Jürgen Kirstein, Jürgen Kirstein, Jürgen Kirstefan Kissenbeck, Andreas Kissenbeck, Andreas Kis	•DD 10.3 •EP 2.2 T 30.5, T 43.8 •EP 2.2 T 31.2 T 66.9 •T 94.7 •T 8.4 EP 3.7 •T 104.5, T 104.6, GR 2.2 •GR 2.2 •DD 3.3 •DD 4.2 ·GR 2.2 •EP 6.7 •T 10.9 ·T 5.1 •EP 6.7 ·T 13.3, T 66.1 •DD 2.2, DD 3.4 · · · · · · · · · · · · ·
Kiesling, Christian Kilian, Schwarz Kilian, Wolfgang Kim, Nohun Kim, Dohun Kim, Ye Joon Kim, Ye Joon Kinast, Angelina T 104.7 King, Andreas King, Markus King, Montague King, Montague King, Montague King, Montague Kirchhoff, Andreas Kirfel, Florian Kirk, John Kirschenmann, Henni Kirstein, Jürgen Kirschenmann, Henni Kirstein, Jürgen Kirsden Andreas Kissenbeck, Andreas Kissenbeck, Andreas Kissmann, Ralf Kitagawa, Hussain Kivernyik, Oleh Kiwit, Florian Kivernyik, Oleh Kiwit, Florian Kikein, Daniel Klein, Katja T 65.4, T 94.6 Klein, Katja D 10.5, DD 25.2, D D 34.9, DD 36.1 Klein-Bösing, Christian DD 20.3 Kliem, Bernhard Klinkert, Maximilian	•DD 10.3 •EP 2.2 T 30.5, T 43.8 •EP 2.2 T 31.2 T 66.9 •T 94.7 •T 94.7 •GR 2.2 •GR 2.2 •DR 4.1 •D 10.2, DD 33.4 ·GR 4.5 ·GR 5.3, •T 88.4 DD 10.2, DD 10.4, D 32.2, DD 34.1, •T 24.2, •EP 12.11 •EP 10.3, •T 81.3 •T 78.6 ·T 78.6
Kiesling, Christian Kilian, Schwarz Kilian, Wolfgang Kilminster, Ben Kim, Dohun Kim, Ye Joon Kim, Ye Joon Kim, Ye Joon King, Andreas King, Angelina T 104.7 King, Andreas King, Montague King, Montague King, Montague Kirg, Montague Kirg, Montague Kirg, Montague Kirg, Montague Kirg, Montague Kirg, Montague Kirg, Montague Kirchhoff, Andreas Kirfel, Florian Kirschenmann, Henni Kirstein, Jürgen Kirsden Jürgen Kirsden Justin Kirstein, Jürgen Kissenbeck, Andreas Kissmann, Ralf Kissenbeck, Andreas Kissmann, Ralf Kitagawa, Hussain Kivernyik, Oleh Kisin, Katja Klein, Katja Klein, Katja T 65.4, T 94.6 Klein, Lucas Klein, Pascal D 10.5, DD 25.2, D DD 34.9, DD 36.1 Klein-Bösing, Christian DD 20.3 Kline, Bernhard Kling, Felix Klinkent, Maximilian Kluke, Markus	• DD 10.3 • EP 2.2 • T 30.5, T 43.8 • EP 2.2 • T 31.2 • T 66.9 • T 94.7 • T 94.7 • T 94.7 • T 104.5, T 104.6, • GR 2.2 • CR 6.7 • DD 16.2 • DD 2.2 • EP 6.7 • DD 2.2 • T 10.9 • T 5.1 • T 2.2 / I, T 25.2 / I, T 25.2
Kiesling, Christian Kilian, Schwarz Kilian, Schwarz Kilminster, Ben Kim, Dohun Kim, Ye Joon Kim, Ye Joon Kim, Ye Joon Kina, Angelina T 104.7 King, Andreas King, Montague King, Montague King, Montague King, Montague King, Montague King, Pietro Kirchhoff, Andreas Kirchhoff, Andreas Kirchhoff, Andreas Kirschenmann, Henni Kirstein, Jürgen Kirstein, Jürgen Kirstein, Jürgen Kirstein, Jürgen Kirstein, Jürgen Kirstein, Jürgen Kirstein, Jürgen Kirstein, Jürgen Kirstein, Jürgen Kissenbeck, Andreas Kissenbeck, Andreas Kissenbeck, Andreas Kissenbeck, Andreas Kissmann, Ralf Kitagawa, Hussain Kivernyik, Oleh Kiwit, Florian Kivernyik, Oleh Kiwit, Florian Kivernyik, Oleh Kiwit, Florian Kiein, Daniel Klein, Lucas Klein, Pascal D 10.5, DD 25.2, D D D 34.9, DD 36.1 Klein, Berinhard Kling, Felix Klingenmeyer, Hannal Klinkenberg, Ludwig Klinkert, Maximilian Klute, Markus	• DD 10.3 • EP 2.2 • T 30.5, T 43.8 • EP 2.2 • T 31.2 • T 66.9 • T 94.7 • T 8.4 • EP 3.7 • T 104.5, T 104.6, • GR 2.2 • GR 4.1, DD 3.3 • EP 12.3 • T 38.1, T 38.7 • T 24.2, • EP 12.11 • EP 10.3, • T 81.3 • T 24.2, • EP 12.11 • CP 10.3, • T 81.3 • T 78.6 • T 2.2 / II, T 25.7, T 25.2, 6 • T 6.2, T 94.4
Kiesling, Christian Kilian, Schwarz Kilian, Wolfgang Kilminster, Ben Kim, Dohun Kim, Ye Joon Kim, Ye Joon Kim, Ye Joon King, Andreas King, Andreas King, Markus King, Montague King, Montague King, Montague King, Montague King, Montague Kirchhoff, Andreas Kirfel, Florian Kirk, John Kirschenmann, Henni Kirstein, Jürgen Kirschenmann, Henni Kirstein, Jürgen Kirstefan Kirstefan Kissenbeck, Andreas Kissmann, Ralf Kissenbeck, Andreas Kissmann, Ralf Kissenbeck, Andreas Kissmann, Ralf Kissenbeck, Andreas Kissmann, Ralf Kistein, Dastin Kivernyik, Oleh Kiwit, Florian Kivernyik, Oleh Kiwit, Florian Kiein, Daniel Kiein, Katja T 65.4, T 94.6 Klein, Lucas Klein, Bernhard Kling, Felix Kling, Felix Kling, Felix Klingenmeyer, Hannal Klinkenberg, Ludwig Klinkert, Maximilian Klute, Markus PT 36.4, T 65.5, T 65.5	• DD 10.3 • EP 2.2 • T 30.5, T 43.8 • EP 2.2 • T 31.2 • T 66.9 • T 94.7 • T 94.7 • T 104.5, T 104.6, • GR 2.2 • GR 2.5 • GR 2
Kiesling, Christian Kilian, Schwarz Kilian, Wolfgang Kilminster, Ben Kim, Dohun Kim, Ye Joon Kim, Ye Joon Kim, Ye Joon King, Andreas T 104.7 King, Andreas King, Markus King, Montague King, Montague King, Montague Kirg, Montague Kirg, Pietro Kirchhoff, Andreas Kirfel, Florian Kirschenmann, Henni Kirschenmann, Henni Kirstein, Jürgen Kirwald, Dustin Kirstein, Jürgen Kirwald, Dustin Kirstefan Kissenbeck, Andreas Kissmann, Ralf Kitagawa, Hussain Kivernyik, Oleh Kiwit, Florian Kivernyik, Oleh Kiwit, Florian Kikein, Katja Klein, Katja Klein, Katja T 65.4, T 94.6 Klein, Lucas DD 10.5, DD 25.2, D DD 34.9, DD 36.1 Klein-Bösing, Christian DD 20.3 Kliem, Bernhard Klinkent, Maximilian Klute, Markus Varang, Ludwig Klinkert, Maximilian Klute, Markus Klein, Stefan Klute, Markus	• DD 10.3 • EP 2.2 • T 30.5, T 43.8 • EP 2.2 • T 31.2 • T 66.9 • T 94.7 • T 94.7 • T 104.5, T 104.6, • GR 2.2 • GR 2

ŀ

Kneip, Nina	T 102.6
Knodel, Oliver	T 96.4
Knoll, Alois	T 43.8
Knue Andrea T73T7	6 T 26 2
T 26.4, T 26.9, •T 27.3	.0, 1 20.2,
Kobel, Michael	4.1, T 24.2
Kobler, Michael GR 12	.3, MP 9.3
Kogler, Roman I 6.3, I /	/.2, 1 10.4,
Kohl Christian Thomas	GPhil 12 2
Köhler, Christoph	•T 49.6
Kohler, Frieder . T 20.3, T 20	.9, T 48.6,
T 48.8, T 102.3	T 00 1
Köhli Markus T398 T6	8 1 T 97 1
Kok. Karel	•DD 4.1
Koke, David	•T 52.5
Kokkonen, Tommi	DD 42.1
Kokkotas, Kostas GR 11	2, GR 11.4
Kolhey, Patrick	•EP 3.4
Kolk, Lars	1 35./
Köllenberger Leonard	T 21 6
Kołodziej Magdalena	ST 1 5
Komarov, Ihor	•T 69.5
Komm, Matthias	.5, •T 69.4
Komorek, Michael . DD 20.7	I, DD 20.2,
DD 27.1, •DD 27.2, DD 46.3	
Koneke, Karsten . 18.1, 16	1.2, 1 61.3
Konig, Philipp	5.1, 1 5.Z
Konstantinova Maria	•T 95 6
Kontaxis, Pantelis	
Kontrimas, Tomas	•T 18.4
Koole, Gabriël	T 2.4
Kooshk, Kaveh	•ST 6.5
Köper, Maximilian	T 96.2
Köppel, Marius	•T 14.9
Koppennoter, Roland .•1 65	0.5, I 65.6,
korcari william	
Korn Steffen •T 12	5 •T 33 4
Korosec, Michele	•T 75.5
Korsmeier, Michael	T 86.6
Kortner, Oliver . T 14.5, T 43	8.2, T 67.3,
T 67.4, T 67.5, T 67.6, T 67.	7
Kortnor Condro TOATA	
	3.2, T 61.1,
T 64.5, T 64.6	3.2, T 61.1,
T 64.5, T 64.6 Kösker, Gani T 65.5, T 65 T 94 4	3.2, T 61.1, 5.6, T 65.7,
Kortriel, Santra 19.4, 14. T 64.5, T 64.6 Kösker, Gani T 65.5, T 65 T 94.4 Kostyukhin. Vadim T 20	3.2, T 61.1, 5.6, T 65.7, 5.5. T 42.6
Kortner, Sandra 194, 14 T 64.5, T 64.6 Kösker, Gani T 65.5, T 65 T 94.4 Kostyukhin, Vadim T 20 Koumtzis, Argiris	3.2, T 61.1, 5.6, T 65.7, 5.5, T 42.6 . •EP 12.8
Kortner, Sandra 1.19.4, 14. T 64.5, T 64.6 Kösker, Gani Kösker, Gani T 65.5, T 65 T 94.4 Kostyukhin, Vadim Kountzis, Argiris T 20 Koundal, Paras T 72.1, •	3.2, T 61.1, 5.6, T 65.7, 5.5, T 42.6 . •EP 12.8 AKPIK 4.1
Kortner, Sandra 1.19.4, 14. T 64.5, T 64.6 Kösker, Gani Kösker, Gani T 65.5, T 65. T 94.4 Kostyukhin, Vadim Koutral, Paras T 72.1, 6 Kovac, Neven	3.2, T 61.1, 5.6, T 65.7, 5.5, T 42.6 •EP 12.8 AKPIK 4.1 5, •T 102.6
Kortner, Sandra 1.19.4, 14. T 64.5, T 64.6 Kösker, Gani Kösker, Gani T 65.5, T 65. T 94.4 Kostyukhin, Vadim Koundal, Paras T 72.1, e Kovac, Neven T 102.5 Kovad, Oksana	3.2, T 61.1, 5.6, T 65.7, 5.5, T 42.6 •EP 12.8 AKPIK 4.1 5, •T 102.6 T 35.6
Kortner, Sandra 1.19.4, 14. T 64.5, T 64.6 Kösker, Gani Kösker, Gani T 65.5, T 65 T 94.4 Kostyukhin, Vadim Kountzis, Argiris	3.2, T 61.1, 5.6, T 65.7, 5.5, T 42.6 • EP 12.8 AKPIK 4.1 5, •T 102.6 T 35.6 •T 96.5 •T 96.5
Kortner, Sandra 1.19.4, 1.4. T 64.5, T 64.6 Kösker, Gani Kösker, Gani T 65.5, T 65 T 94.4 Kostyukhin, Vadim Koundal, Paras • T 72.1, • Kovac, Neven T 102.5 Koval, Oksana	3.2, T 61.1, 5.6, T 65.7, 5.5, T 42.6 • EP 12.8 AKPIK 4.1 5, •T 102.6 T 35.6 •T 96.5 , •DD 42.3 • T 30.7
Kortner, Sandra 19,4, 14. T 64.5, T 64.6 Kösker, Gani T 65.5, T 65 T 94.4 Kostyukhin, Vadim T 20 Koundal, Paras • T 72.1, • Kovac, Neven T 102.5 Koval, Oksana Kozlinskiy, Alexandr Krabe, Heiko D 39.1 Kraetzschmar, Thomas Krämer, Michael	3.2, T 61.1, 5.6, T 65.7, 5.5, T 42.6 EP 12.8 AKPIK 4.1 5, •T 102.6 T 35.6 •T 96.5 •T 30.7 III. T 31.2.
Kortner, Sandra 1.19.4, 14. T 64.5, T 64.6 Kösker, Gani Kösker, Gani T 65.5, T 65 T 94.4 Kostyukhin, Vadim Koundzis, Argiris	3.2, T 61.1, 5.6, T 65.7, 5.5, T 42.6 • EP 12.8 AKPIK 4.1 5, • T 102.6 T 35.6 • T 96.5 , •DD 42.3 • T 30.7 III, T 31.2, 4
Kortner, Sandra 1.9.4, 1.4. T 64.5, T 64.6 Kösker, Gani T 65.5, T 65 T 94.4 Kostyukhin, Vadim T 20 Kountzis, Argiris Koundal, Paras T 72.1, • Kovac, Neven T 102.5 Koval, Oksana Koval, Oksana Kozlinskiy, Alexandr Krabbe, Heiko DD 39.1 Kraetzschmar, Thomas Krämer, Michael	3.2, T 61.1, 5.6, T 65.7, 5.5, T 42.6 • EP 12.8 AKPIK 4.1 5, •T 102.6 T 35.6 •T 96.5 , •DD 42.3 •T 30.7 III, T 31.2, 4 .6, •T 10.7
Kortner, Sandra 1.19.4, 14. T 64.5, T 64.6 Kösker, Gani T 65.5, T 65 T 94.4 Kostyukhin, Vadim T 20 Koundal, Paras T 72.1, 6 Kovac, Neven T 102.5 Kovad, Oksana T 102.5 Koval, Oksana Kozlinskiy, Alexandr Krabe, Heiko D 39.1 Kraetzschmar, Thomas FVA.3, T 86.6, T 93.3, T 93. Krämer, Peter T 10 Kramer, Tobias T 9	3.2, T 61.1, 5.6, T 65.7, 5.5, T 42.6 .•EP 12.8 AKPIK 4.1 5, •T 102.6 T 35.6 •T 96.5 , •DD 42.3 •T 30.7 III, T 31.2, 4 .6, •T 10.7 1.5, T 91.6
Kortner, Sandra 1.19.4, 14. T 64.5, T 64.6 Kösker, Gani T 65.5, T 65 T 94.4 Kostyukhin, Vadim T 20 Koundal, Paras T 72.1, 6 Kovac, Neven T 102.5 Kovac, Neven T 102.5 Koval, 0ksana Kozilinskiy, Alexandr Krabe, Heiko D 39.1 Kraetzschmar, Thomas Krämer, Michael PV 45.3, T 86.6, T 93.3, T 93. Krämer, Peter T 10 Kramer, Tobias T 9 Krasznahorkay, Attila Model Y 9	3.2, T 61.1, 5.6, T 65.7, 5.5, T 42.6 EP 12.8 AKPIK 4.1 5, •T 102.6 T 35.6 •T 96.5 •T 30.7 III, T 31.2, 4 .6, •T 10.7 1.5, T 91.6 T 16.7
Korner, Sandra 19,4,14. T 64.5, T 64.6 Kösker, Gani T 65.5, T 65 T 94.4 Kostyukhin, Vadim T 24 Koundal, Paras T 72.1, 4 Kovac, Neven T 102.5 Kovad, Oksana Krabbe, Heiko DD 39.1 Kraetzschmar, Thomas Krämer, Michael Krämer, Michael PV T 45.3, T 86.6, T 93.3, T 93. Krämer, Tobias T 9 Krasznahorkay, Attila Kraus, Johanna Wanda	3.2, T 61.1, 5.6, T 65.7, 5.5, T 42.6 5.5, T 42.6 4.6, T 102.6 5.5, T 102.6 5.5, T 102.6 5.5, T 96.5 5.5, D 42.3 5.5, T 10.7 111, T 31.2, 4 5.5, T 10.7 1.5, T 91.6 5.5, T 10.7 1.5, T 10
Korner, Sandra 19,4,14. T 64.5, T 64.6 Kösker, Gani T 65.5, T 65 T 94.4 Kostyukhin, Vadim T 24 Koundal, Paras T 72.1, 4 Kovac, Neven T 102.5 Kovad, Oksana Krabbe, Heiko DD 39.1 Kraetzschmar, Thomas Krämer, Michael PV T 45.3, T 86.6, T 93.3, T 93. Krämer, Peter T 10 Kramer, Tobias T 9 Krasznahorkay, Attila Kraus, Johanna Wanda Kraus, Michaela EP 1 Kraus, Stefan	3.2, T 61.1, 5.6, T 65.7, 5.5, T 42.6 5.5, T 42.6 5.5, T 102.6 T 35.6 T 35.6 T 96.5 , •DD 42.3 T 30.7 III, T 31.2, 4 T 10.7 1.5, T 91.6 T 16.7 T 5.4 .2, •EP 1.3 •DD 35.11
Korner, Sandra 19,4,14. T 64.5, T 64.6 Kösker, Gani T 65.5, T 65 T 94.4 Kostyukhin, Vadim T 24 Koundal, Paras T 72.1, 6 Kovac, Neven T 102.5 Koval, Oksana Kozlinskiy, Alexandr Krabbe, Heiko DD 39.1 Kraetzschmar, Thomas Krämer, Michael PV T 45.3, T 86.6, T 93.3, T 93. Krämer, Tobias T 9 Krasznahorkay, Attila Kraus, Johanna Wanda Kraus, Stefan Krause, Christopher	3.2, T 61.1, 5.6, T 65.7, 5.5, T 42.6 EP 12.8 AKPIK 4.1 AKPIK 4.1 5., -T 102.6 T 35.6 T 30.7 III, T 31.2, 4 T 10.7 II.5, T 91.6 T 16.7 F 5.4 .2, EP 1.3 •DD 35.11 •DD 35.11
Korner, Sandra 19,4, 14 T 64.5, T 64.6 Kösker, Gani T 65.5, T 65 T 94.4 Kostyukhin, Vadim T 24 Koundal, Paras - 7 72.1, 6 Kovac, Neven T 102.5 Koval, Oksana Kozlinskiy, Alexandr Krabbe, Heiko D 39.1 Kraetzschmar, Thomas Krämer, Michael PV T 45.3, T 86.6, T 93.3, T 93. Krämer, Peter T 10 Kraetzschorkay, Attila Kraus, Johanna Wanda Kraus, Stefan Kraus, Christopher Krause, Johannes	3.2, T 61.1, 5.6, T 65.7, 5.5, T 42.6 EP 12.8 AKPIK 4.1 5, -T 102.6 T 35.6 T 96.5 T 96.5 T 30.7 III, T 31.2, 4 T 16.7 T 5.4 .2, •EP 1.3 •DD 35.11 T 88.3
Korner, Sandra 19,4, 14. T 64.5, T 64.6 Kösker, Gani T 65.5, T 65 T 94.4 Kostyukhin, Vadim T 20 Koumtzis, Argiris Kovac, Neven T 102.4 Koval, Oksana Kozlinskiy, Alexandr Krabbe, Heiko D 39.1 Kraetzschmar, Thomas Krämer, Michael PV T 45.3, T 86.6, T 93.3, T 93. Krämer, Peter T 10 Kraetzschmar, Tobias Krämer, Tobias Kramer, Tobias Kraus, Johanna Wanda Kraus, Stefan Kraus, Stefan Krause, Christopher Krause, Johannes Krause, Patrick •T 75.1, T 75	3.2, T 61.1, 5.6, T 65.7, 5.5, T 42.6 EP 12.8 AKPIK 4.1 5, •T 102.6 T 35.6 •T 96.5 •T 96.5 •T 30.7 110, T 31.2, 4 T 16.7 •T 5.4 .2, •EP 1.3 •DD 35.11 •ST 1.1 •ST 1.1 •ST 1.1 •ST 4.4, T 78.4,
Kortner, Sandra 1.9.4, 1.4. T 64.5, T 64.6 Kösker, Gani T 65.5, T 65 T 94.4 Kostyukhin, Vadim T 20 Kountzis, Argiris Koutais, Argiris Koundal, Paras T 72.1, 4 Kovac, Neven T 102.5 Kovad, Oksana Kovad, Oksana Kozlinskiy, Alexandr Krabbe, Heiko DD 39.1 Kraetzschmar, Thomas Krämer, Michael PV T 45.3, T 86.6, T 93.3, T 93. Krämer, Tobias Kraus, Johanna Wanda Kraus, Johanna Wanda Krause, Christopher Krause, Abhanes Krause, Patrick •T 75.1, T 75 T 78.5	3.2, T 61.1, 5.6, T 65.7, 5.5, T 42.6 EP 12.8 AKPIK 4.1 5, •T 102.6 T 35.6 •T 96.5 •T 96.5 •T 96.5 •T 30.7 110, T 31.2, 4 T 16.7 •T 5.4 .2, •EP 1.3 •DD 35.11 •T 81.3 5.4, T 78.4, T 62.7
Korner, Sandra 19,4,14. T 64.5, T 64.6 Kösker, Gani T 65.5, T 65 T 94.4 Kostyukhin, Vadim T 24 Koundal, Paras T 72.1, Kovac, Neven T 102.5 Koval, Oksana Kozlinskiy, Alexandr Krabbe, Heiko DD 39.1 Kraetzschmar, Thomas Krämer, Michael PV T 45.3, T 86.6, T 93.3, T 93. Krämer, Peter T 10 Kramer, Tobias T 9 Krasznahorkay, Attila Kraus, Johanna Wanda Krause, Christopher Krause, Christopher Krause, Patrick •T 75.1, T 75 T 78.5 Kravchenko, Michail	3.2, T 61.1, 5.6, T 65.7, 5.5, T 42.6 5.5, T 42.6 5.5, T 42.6 5.5, T 42.6 5.5, T 42.6 5.5, T 102.6 5.5, T 102.6 5.5, T 96.5 5.5, T 90.5 5.5, T 91.6 5.5, T 91.6 5
Korner, Sandra 194, 14 T 64.5, T 64.6 Kösker, Gani T 65.5, T 65 T 94.4 Kostyukhin, Vadim T 24 Koundal, Paras T 72.1, 4 Kovac, Neven T 102.3 Koval, Oksana Kozlinskiy, Alexandr Krabbe, Heiko DD 39.1 Kraetzschmar, Thomas Krämer, Michael PV T 45.3, T 86.6, T 93.3, T 93. Krämer, Peter T 10 Kramer, Tobias T 9 Krasznahorkay, Attila Kraus, Johanna Wanda Krause, Christopher Krause, Christopher Krause, Patrick •T 75.1, T 75 T 78.5 Kravchenko, Michail Kreikemeverel orenzo. Dama	3.2, T 61.1, 5.6, T 65.7, 5.5, T 42.6 5.5, T 42.6 5.5, T 42.6 5.5, T 102.6 T 35.6 5.5, T 102.6 T 35.6 5.5, T 96.5 5.5, D 42.3 5.5, T 91.6 5.5, T 91.6
Korner, Sandra 19,4,14. T 64.5, T 64.6 Kösker, Gani T 65.5, T 65 T 94.4 Kostyukhin, Vadim T 24 Koundal, Paras T 72.1, 4 Kovac, Neven T 102.5 Kovad, Oksana Kozlinskiy, Alexandr Krabbe, Heiko DD 39.1 Kraetzschmar, Thomas Krämer, Michael Krämer, Michael PV T 45.3, T 86.6, T 93.3, T 93. Krämer, Peter T 0 Kraetz, Sohanna Wanda Kraus, Johanna Wanda Krause, Christopher Krause, Christopher Krause, Christopher Krause, Datrick •T 75.1, T 75 T 78.5 Kravchenko, Michail Kreikemeyer-Lorenzo, Dagma •T 105.7	3.2, T 61.1, 5.6, T 65.7, 5.5, T 42.6 5.5, T 42.6 5.5, T 42.6 5.5, T 102.6 5.5, T 102.6 5.5, T 102.6 5.5, T 102.6 5.5, T 102.6 5.5, T 102.6 5.5, T 102.7 111, T 31.2, 4 5.6, •T 10.7 1.5, T 91.6 5.7, T 10.7 1.5,
Korner, Sandra 194, 14 T 64.5, T 64.6 Kösker, Gani T 65.5, T 65 T 94.4 Kostyukhin, Vadim T 24 Koundal, Paras T 72.1, 4 Kovac, Neven T 102.5 Kovad, Oksana Krabbe, Heiko DD 39.1 Kraetzschmar, Thomas Krämer, MichaelPV T 45.3, T 86.6, T 93.3, T 93. Krämer, Peter T 10 Kramer, TobiasT 9 Krasznahorkay, Attila Krause, Johanna Wanda Krause, Johannes Krause, Christopher Krause, Dahannes Krause, Patrick •T 75.1, T 75 T 78.5 Kravchenko, Michail Kreikemeyer-Lorenzo, Dagma •T 105.7 Kremer, Jakub 1	3.2, T 61.1, 5.6, T 65.7, 5.5, T 42.6 5.5, T 42.6 5.5, T 42.6 5.5, T 102.6 T 35.6 T 96.5 T 96.5 T 30.7 III, T 31.2, 4 T 30.7 III, T 31.2, 4 T 10.7 1.5, T 91.6 T 16.7 T 5.4 T 5.4 T 88.3 T 88.3 T 88.3 T 68.7 T 2.8 r 5.1, T 5.2
Kortner, Sandra 1.9.4, 1.4. T 64.5, T 64.6 Kösker, Gani T 65.5, T 65 T 94.4 Kostyukhin, Vadim T 24 Koundal, Paras T 72.1, 6 Kovac, Neven T 102.5 Kovac, Neven T 102.5 Kovad, 0ksana T 24.0 Krabe, Heiko DD 39.1 Kraetzschmar, Thomas T 25.0 Krabe, Heiko DD 39.1 Kraetzschmar, Thomas T 102.5 Krater, Michael PV T 45.3, T 86.6, T 93.3, T 93.3 Krämer, Peter T 10 Kramer, Tobias T 9 Krasznahorkay, Attila Kraus, Johanna Wanda Krause, Johanna Wanda Krause, Christopher Krause, Patrick •T 75.1, T 75 Krause, Dahannes T 78.5 Kravchenko, Michail Krebs, Tobias T 105.7 Kreikemeyer-Lorenzo, Dagma •T 105.7 T 78.5 Kreekenger-Lorenzo, Dagma	3.2, T 61.1, 5.6, T 65.7, 5.5, T 42.6 EP 12.8 AKPIK 4.1 5.5, T 102.6 T 35.6 T 36.6 T 96.5 T 30.7 III, T 31.2, 4 T 10.7 II.5, T 91.6 T 10.7 II.5, T 91.6 T 16.7 F 5.4 T 5.4 T 88.3 5.4, T 78.4, T 68.7 T 2.8 5.1, T 5.2 1.9, T 62.9
Kortner, Sandra 1.19.4, 1.4. T 64.5, T 64.6 Kösker, Gani T 65.5, T 65 T 94.4 Kostyukhin, Vadim T 26 Koundal, Paras T 72.1, 6 Kovac, Neven T 102.5 Kovad, Oksana T 102.5 Kovad, Oksana T 102.5 Krabbe, Heiko DD 39.1 Kratzschmar, Thomas T 102.5 Krabbe, Heiko DD 39.1 Kratzschmar, Thomas T 102.5 Kratzschmar, Thomas T 102.5 Krämer, Petker T 102.5 Krämer, Peter T 102.5 Kraus, Johanna Wanda T 9 Krasznahorkay, Attila F 9 Krause, Johannes T 75.1, T 75 Krause, Patrick •T 75.1, T 75 T 78.5 Kravchenko, Michail Krebs, Tobias T 75.1, T 75 Kreikemeyer-Lorenzo, Dagma •T 105.7 Kreemer, Jakub T 4 Kretzer, Olaf T 4	3.2, T 61.1, 5.6, T 65.7, 5.5, T 42.6 EP 12.8 AKPIK 4.1 5 T 35.6 T 35.6 T 36.5 T 96.5 T 96.5 T 30.7 III, T 31.2, 4 T 10.7 1.5, T 91.6 T 10.7 1.5, T 91.6 T 5.4 .2, •EP 1.3 •D 35.11 T 88.3 5.4, T 78.4, T 68.7 T 2.8 r 5.1, T 5.2 1.9, T 62.9 D 28.1
Kortner, Sandra 1.1.9.4, 1.4. T 64.5, T 64.6 Kösker, Gani T 65.5, T 65 T 94.4 Kostyukhin, Vadim T 21 Kountzis, Argiris Kountzis, Argiris Koundal, Paras T 72.1, • Kovac, Neven T 102.5 Koval, Oksana Koval, Oksana Kozlinskiy, Alexandr Krabbe, Heiko DD 39.1 Kraetzschmar, Thomas Krämer, Michael PV T 45.3, T 86.6, T 93.3, T 93.3 Krämer, T 100.5 Kramer, Tobias T 9 Krasznahorkay, Attila	3.2, T 61.1, 5.6, T 65.7, 5.5, T 42.6 EP 12.8 AKPIK 4.1 5T 102.6 T 35.6 T 35.6 T 96.5 T 96.5 T 10.7 11, T 31.2, 4 T 10.7 1.5, T 91.6 T 16.7 T 5.4 .2, •EP 1.3 •D 35.11 T 88.3 5.4, T 78.4, T 68.7 T 2.8 r 5.1, T 5.2 1.9, T 62.9 D 28.1 DD 10.9
Kortner, Sandra 1.1.9.4, 1.4. T 64.5, T 64.6 Kösker, Gani T 65.5, T 65 T 94.4 Kostyukhin, Vadim T 24 Kostyukhin, Vadim T 24 Kostyukhin, Vadim Koumdal, Paras T 72.1, 4 Kovac, Neven T 102.5 Kovad, Oksana Kovad, Oksana Kozlinskiy, Alexandr Krabbe, Heiko Krabbe, Heiko DD 39.1 Kraetzschmar, Thomas Frü Krämer, Michael PV T 45.3, T 86.6, T 93.3, T 93. Krämer, Tobias Kraus, Johanna Wanda Krause, Johanna Wanda Kraus, Stefan	3.2, T 61.1, 5.6, T 65.7, 5.5, T 42.6 EP 12.8 AKPIK 4.1 5, -T 102.6 T 35.6 T 96.5 T 96.5 T 96.5 T 30.7 III, T 31.2, 4 T 16.7 T 16.7 T 5.4 .2, •EP 1.3 •D 35.11 T 88.3 5.4, T 78.4, T 68.7 T 68.7 T 2.8 F 5.1, T 5.2 1.9, T 62.9 DD 62.
Kortner, Sandra 1.9.4, 1.4. T 64.5, T 64.6 Kösker, Gani T 65.5, T 65 T 94.4 Kostyukhin, Vadim T 24 Koutdal, Paras T 72.1, 4 Koundal, Paras T 72.1, 4 Kovac, Neven T 102.5 Kovad, Oksana Kozlinskiy, Alexandr Krabbe, Heiko DD 39.1 Kraetzschmar, Thomas Krämer, Michael Krämer, Michael PV T 45.3, T 86.6, T 93.3, T 93. Krämer, Tobias Kraus, Johanna Wanda Krause, Johanna Wanda Kraus, Stefan T 75.1, T 75 Krause, Christopher Krause, Johannes Krause, Johannes Krause, T 105.7 Krewchenko, Michail T 75.1, T 75 T 78.5 Kravchenko, Michail Kreize, Tolas 4 Kretzer, Olaf 4 Kretzer, Olaf 4 Kriegel, Christina Kriegel, Moritz	3.2, T 61.1, 5.6, T 65.7, 5.5, T 42.6 5.5, T 42.6 5.5, T 42.6 5.5, T 42.6 5.5, T 42.6 5.7, T 02.6 5.1, T 30.7 1.5, T 91.6 5.1, T 5.2 1.9, T 62.9 5.1, T 5.2 1.9, T 62.9 5.1, T 5.2 1.9, T 62.9 5.1, D 28.1 5.1, D 10.9 5.1, D 28.1 5.1, D 20.3 5.1, D
Kortner, Sandra 1.9.4, 1.4. T 64.5, T 64.6 Kösker, Gani T 65.5, T 65 T 94.4 Kostyukhin, Vadim T 24 Koutdal, Paras T 72.1, 4 Kostage Koundal, Paras T 72.1, 4 Kovac, Neven T 102.3 Kovac, Neven T 102.3 Koval, 0ksana Kozlinskiy, Alexandr Krabbe, Heiko DD 39.1 Kraetzschmar, Thomas Krämer, Michael PV 45.3, T 93. Krämer, Michael PV T 45.3, T 86.6, T 93.3, T 93. Krämer, Tobias T 10 Kramer, Tobias T 10 Kramer, Tobias T 10 Kraus, Johanna Wanda Krause, Johannes Krause, Johannes Krause, Christopher Krause, Johannes T 75.1, T 75 T 78.5 Kravchenko, Michail Freise, Tobias 105.7 Kremer, Jakub T 105.7 Kremer, Jakub T 4 Kretzer, Olaf 74 Kriege, Tom T 4 Kretzer, Olaf Kriegel, Christina Kriegel, Moritz Kriegel, Christina	3.2, T 61.1, 6.6, T 65.7, 5.5, T 42.6 EP 12.8 AKPIK 4.1 5, •T 102.6 T 35.6 T 35.6 T 30.7 III, T 31.2, 4 T 30.7 III, T 31.2, 4 F 130.7 III, T 31.2, 4 F 191.6 T 16.7 T 5.4 •T 68.7 T 68.7 T 68.7 T 2.8 5.1, T 5.2 1.9, T 62.9 DD 28.1 DD 10.9 DD 20.3 •DD 9.1 •DD 9.1 •T 0.5 •T 0.5
Kortner, Sandra 1.9.4, 1.4. T 64.5, T 64.6 Kösker, Gani T 65.5, T 65 T 94.4 Kostyukhin, Vadim T 24 Koutdal, Paras T 72.1, 6 Koskar, Argiris Koundal, Paras T 72.1, 6 Kovac, Neven T 102.5 Kovac, Neven T 102.5 Koval, 0ksana Koval, 0ksana Koval, 0ksana Kozlinskiy, Alexandr Krabbe, Heiko DD 39.1 Kraetzschmar, Thomas Krämer, Michael PV Krabe, Heiko DD 39.1 Kraetzschmar, Thomas T 10 Kramer, Michael PV T 45.3, T 86.6, T 93.3, T 93. Krämer, Peter T 10 Kramer, Tobias T 9 Krasznahorkay, Attila Kraus, Johanna Wanda Krause, Johannes Krause, Johannes Krause, Johannes Krause, Johannes Krause, Johannes Krause, Johannes T 75.1, T 75 T 78.5 Kravchenko, Michail Kreise, Tobias T 105.7 Kremer, Jakub T 4 Kretzer, Olaf Kriegel, Christina T 4 Kriegel, Christina Kriegel, Mori	3.2, T 61.1, 6.6, T 65.7, 5.5, T 42.6 EP 12.8 AKPIK 4.1 5, •T 102.6 T 35.6 T 96.5 , •DD 42.3 •T 30.7 III, T 31.2, 4 •T 30.7 III, T 31.2, 4 •T 10.7 1.5, T 91.6 T 16.7 •T 5.4 .2, •EP 1.3 •ST 1.1 T 88.3 3.4, T 78.4, •T 68.7 T 2.8 r 5.1, T 5.2 1.9, T 62.9 DD 28.1 DD 28.1 DD 28.1 DD 20.3 •DD 9.1 •T 105.1 •T 15.5
Kortner, Sandra 1.9.4, 1.4. T 64.5, T 64.6 Kösker, Gani T 65.5, T 65 T 94.4 Kostyukhin, Vadim T 24 Koundal, Paras T 72.1, 6 Koskar, Argiris Koundal, Paras T 72.1, 6 Koskar, Argiris Kovac, Neven T 102.5 Kovad, 0ksana Kovad, 0ksana Kozlinskiy, Alexandr Krabe, Heiko DD 39.1 Kraetzschmar, Thomas T 70 Kräbe, Heiko DD 39.1 Kraetzschmar, Thomas T 102.5 Krämer, Michael PV T 45.3, T 86.6, T 93.3, T 93. Krämer, Peter Kraus, Johanna Wanda Kraus, Johanna Wanda Krause, Johannes T 9 Krause, Johannes Krause, Johannes Krause, Johannes Krause, Johannes Krause, Patrick •T 75.1, T 75 T 78.5 Kravchenko, Michail Kreibs, Tobias T 74 Kreibe, Tom T 4 Kreizer, Olaf T 9.4 Kriegel, Christina T 9.4, T 14 Kriegel, Christina T 9.4, T 14	3.2, T 61.1, 6.6, T 65.7, 5.5, T 42.6 EP 12.8 AKPIK 4.1 5, •T 102.6 T 35.6 T 35.6 T 36.7 T 30.7 III, T 31.2, 4 T 30.7 III, T 31.2, 4 T 10.7 1.5, T 91.6 T 16.7 T 5.4 •T 68.7 T 2.8 r 5.1, T 5.2 1.9, T 62.9 DD 28.1 DD 28.1 DD 28.1 DD 28.1 DD 28.1 DD 28.3 DD 28.3 0D 9.1 T 15.5 •T 15.5 •T 15.5 •T 15.5 •T 15.5 •T 15.5 •T 105.1 •T 15.5 •T 15.5 •T 37.5, •T 37.5,
Kortner, Sandra 1.9.4, 1.4. T 64.5, T 64.6 Kösker, Gani T 65.5, T 65 T 94.4 Kostyukhin, Vadim T 24 Kountzis, Argiris T 72.1, 6 Kovac, Neven T 102.5 Kovad, Oksana T 72.1, 6 Kovad, Oksana T 72.1, 6 Kozlinskiy, Alexandr Krabbe, Heiko Krabbe, Heiko DD 39.1 Kraetzschmar, Thomas T 102.5 Krabe, Heiko DD 39.1 Kraetzschmar, Thomas T 102.5 Kramer, Michael PV T 45.3, T 86.6, T 93.3, T 93.3 Krämer, Peter Kramer, Tobias T 9 Krasznahorkay, Attila Kraus, Johanna Wanda Krause, Johannes T 9 Krause, Johannes T 75.1, T 75 T 78.5 Kravchenko, Michail Krebs, Tobias T 75.1, T 75 T 78.5 Kravchenko, Michail Krebs, Tobias T 9.4, T 14 Kriege, Christina T 4 Kriegel, Moritz Kriegel, Moritz Krey, Olaf Stejan <tr< td=""><td>3.2, T 61.1, 5.6, T 65.7, 5.5, T 42.6 EP 12.8 AKPIK 4.1 5T 102.6 T 35.6 T 35.6 T 36.7 T 30.7 III, T 31.2, 4 T 10.7 III, T 31.2, 4 T 16.7 T 5.4 T 68.7 T 2.8 7 T 2.8 7 T 0.2 T 0.5 1.0, T 62.9 DD 20.3 DD 20.3 DD 20.3 T 10.5 T 15.5 JD 20.3 T 15.5 JD 20.3 T 15.5 JD 20.3 JD 20</td></tr<>	3.2, T 61.1, 5.6, T 65.7, 5.5, T 42.6 EP 12.8 AKPIK 4.1 5T 102.6 T 35.6 T 35.6 T 36.7 T 30.7 III, T 31.2, 4 T 10.7 III, T 31.2, 4 T 16.7 T 5.4 T 68.7 T 2.8 7 T 2.8 7 T 0.2 T 0.5 1.0, T 62.9 DD 20.3 DD 20.3 DD 20.3 T 10.5 T 15.5 JD 20.3 T 15.5 JD 20.3 T 15.5 JD 20.3 JD 20
Kortner, Sandra 1.9.4, 1.4. T 64.5, T 64.6 Kösker, Gani T 65.5, T 65 T 94.4 Kostyukhin, Vadim T 21 Koattyukhin, Vadim T 22 Kountzis, Argiris Kountzis, Argiris T 72.1, • Kovac, Neven T 102.5 Kovad, Oksana Kovac, Neven T 102.5 Koval, Oksana Kovac, Neven T 21, • Kovac, Neven T 102.5 Kovad, Oksana Kovac, Neven T 22, • Kozlinskiy, Alexandr Krabbe, Heiko D 39.1 Kräetzschmar, Thomas PV Krämer, Michael PV T 45.3, T 86.6, T 93.3, T 93.3 Krämer, Peter T 100 Kramer, Peter T 10 Kramer, Tobias T 9 Yrasznahorkay, Attila Krause, Johanna Wanda Krause, Christopher Krause, Johannes Krause, Christopher Krause, Christopher T 78.5 Kravchenko, Michail T 75.1, T 75 T 78.5 Kravchenko, Michail T 75.1, T 75 T 78.5 T 45.7 Kreikemeyer-Lorenzo, Dagma T	3.2, T 61.1, 5.6, T 65.7, 5.5, T 42.6 EP 12.8 AKPIK 4.1 5., -T 102.6 T 35.6 T 35.6 T 30.7 III, T 31.2, 4 T 10.7 III, T 31.2, 4 T 16.7 T 5.4 .2, eEP 1.3 •DD 35.11 T 88.3 5.4, T 78.4, T 68.7 T 2.8 5.1, T 5.2 1.9, T 62.9 D 20.3 DD 10.9 D 20.3 •T 0.5, 3, T 67.4, T 15.5 3, T 67.4, T 6.5 T 10.7 T 10.5 T 10.7 T 10.5 T 10.7 T 10.5 T 10.7 T 10.5 T 10.7 T 10.5 T 10.7 T 10.5 T 10.7 T 10.7 T 2.8 T 10.7 T 10.5 T 10.7 T 10.5 T 10.7 T 2.8 T 10.7 T 10.5 T 10.7 T 10.5 T 10.5 T 10.7 T 10.5 T 10.5 T 10.5 T 10.5 T 10.5 T 10.7 T 2.8 T 10.5 T 10.7 T 2.8 T 10.5 T 10.5
Kortner, Sandra 1.9.4, 1.4. T 64.5, T 64.6 Kösker, Gani T 65.5, T 65 T 94.4 Kostyukhin, Vadim T 24 Kostyukhin, Vadim T 24 Kostyukhin, Vadim T 24 Koundal, Paras T 72.1, 4 Kovac, Neven T 102.5 Kovad, Oksana Kovad, Oksana Kovad, Oksana Kovad, Oksana Kozlinskiy, Alexandr Krabbe, Heiko DD 39.1 Kräetzschmar, Thomas Krämer, Michael PV T 45.3, T 86.6, T 93.3, T 93. Krämer, Michael PV Krass, Johanna Wanda Krause, Johanna Wanda Krause, Johannes Krause, Christopher Krause, Stefan Krause, Christopher T 75.1, T 75 T 78.5 Kravchenko, Michail Krebs, Tobias T 4 Kreikemeyer-Lorenzo, Dagma T 105.7 T 44 Kreize, Olaf T 4 T 4 Kriege, Christoph T 4 T 4.7 Kreige, Christoph T 14 T 4.7 Kreige, Christoph T 4.7 T 4.7<	3.2, T 61.1, 5.6, T 65.7, 5.5, T 42.6 EP 12.8 AKPIK 4.1 5, •T 102.6 T 35.6 T 96.5 •T 30.7 III, T 31.2, 4 6, •T 10.7 1.5, T 91.6 T 16.7 •T 5.4 •ST 1.1 •ST 1.1 T 88.3 6.4, T 78.4, T 68.7 T 2.8 5.1, T 5.2 1.9, T 62.9 DD 28.1 DD 10.9 DD 28.1 DD 10.9 DD 28.3 DD 10.9 DD 6.2 DD 9.1 T 15.5 .5, T 37.5, 3, T 67.4, T 38.7 T 38.7
Kortner, Sandra 1.9.4, 1.4., 1.4., 1.4., 1.4., 1.6., 5.4., 6.4., 7.6., 7.6., 7.9., 4.4., 7.6., 7.6., 7.9., 4.4., 7.6., 7.9., 7.9., 7.6., 7.6., 7.9	3.2, T 61.1, 5.6, T 65.7, 5.5, T 42.6 5.5, T 42.6 5.5, T 42.6 5.5, T 42.6 5.5, T 42.6 5.5, T 30.7 10, T 30.7 11, T 31.2, 4 5, •T 10.7 1.5, T 91.6 5.1, T 5.2 1.9, T 62.9 •DD 28.1 •T 105.1 •T 15.5 5.5, T 37.5, 3, T 67.4, 1.1, •T 38.7 DD 44.2 0 10.3 0 10.9 0 24.2 0 10.9 0 10.9 0 10.9 0 28.1 0 10.9 0 20.3 0 0 20.3 0 0 0.1 0 0 0.1
Kortner, Sandra 1.1.9.4, 1.4. T 64.5, T 64.6 Kösker, Gani T 65.5, T 65 T 94.4 Kostyukhin, Vadim T 24 Koundal, Paras T 72.1, 4 Kovac, Neven T 102.5 Kovad, Oksana Kovac, Neven T 102.5 Kovad, Oksana Kovac, Neven T 102.5 Kovad, Oksana Kovac, Neven T 102.5 Kozalinskiy, Alexandr	3.2, T 61.1, 6.6, T 65.7, 5.5, T 42.6 EP 12.8 AKPIK 4.1 5, •T 102.6 T 35.6 T 35.6 T 36.7 T 30.7 III, T 31.2, 4 T 30.7 III, T 31.2, 4 T 16.7 T 16.7 T 16.7 T 16.7 T 16.7 T 16.7 T 88.3 3.4, T 78.4, T 68.7 T 68.7 T 68.7 T 2.8 5.1, T 5.2 1.9, T 62.9 DD 28.1 DD 0.9 DD 28.1 T 15.5 5.5, T 37.5, 3, T 67.4, T 38.7 DD 34.3 DD 34.3 DD 34.3 DD 34.3 DD 44.2 2 ST 16
Kortner, Sandra 1.1.9.4, 1.4. T 64.5, T 64.6 Kösker, Gani T 65.5, T 65 T 94.4 Kostyukhin, Vadim T 24 Koutdal, Paras T 72.1, 4 Kostage Koundal, Paras T 72.1, 4 Kovac, Neven T 102.3 Kovad, Oksana Kovad, Oksana Kovad, Oksana Kovad, Oksana Kozlinskiy, Alexandr Krämer, Michael PV T 45.3, T 86.6, T 93.3, T 93.3 Krämer, Michael PV T 45.3, T 86.6, T 93.3, T 93.3 Krämer, Peter T 10 Kramer, Tobias T 9 Krasznahorkay, Attila Krause, Johanna Wanda Krause, Johannes Krause, Johannes Krause, Johannes Krause, Johannes Krause, Johannes Krause, Christopher Krause, Johannes Krause, Johannes Krause, Johannes Krause, Johannes T 75.1, T 75. T 78.5 Kravchenko, Michail Kreise, Tobias Kreibe, Tobias Michaela T 75.1, T 75.1, T 75.5 Kreibe, Tobias Kreibe, T 05.7 Kreke, Tom T	3.2, T 61.1, 6.6, T 65.7, 5.5, T 42.6, EP 12.8 AKPIK 4.1 5, •T 102.6 T 35.6 T 96.5 •T 30.7 III, T 31.2, 4 •T 30.7 III, T 31.2, 4 •T 10.7 •T 16.7 •T 16.7 •T 5.4 .2, •EP 1.3 •ST 1.1 •ST 1.1 •ST 1.1 •ST 1.1 •T 68.7 T 2.8 F 5.1, T 5.2 1.9, T 62.9 •DD 28.1 •DD 28.1 •DD 28.1 •DD 28.3 •DD 28.3 •DD 28.3 •DD 28.3 •T 105.1 •T 15.5 5.5, T 37.5, 3, T 67.4, .1, •T 38.7 DD 34.3 DD 44.2 1.2, ST 1.6, .4, ST 1.4,
Kortner, Sandra 1.9.4, 1.4. T 64.5, T 64.6 Kösker, Gani T 65.5, T 65 T 94.4 Kostyukhin, Vadim T 24 Koutralis, Argiris	3.2, T 61.1, 6.6, T 65.7, 5.5, T 42.6, • EP 12.8 AKPIK 4.1 5, •T 102.6 T 35.6 T 35.6 T 36.7 T 30.7 III, T 31.2, 4 T 30.7 III, T 31.2, 4 T 10.7 1.5, T 91.6 T 16.7 F 5.4 •T 68.7 T 2.8 r 5.1, T 5.2 1.9, T 62.9 DD 28.1 •T 68.7 T 2.8 r 5.1, T 5.2 1.9, T 62.9 DD 28.1 •T 105.1 •T 105.1 •T 105.1 •T 15.5 5.5, T 37.5, 3, T 67.4, D 34.3 DD 44.2 .2, ST 1.6, 3, ST 7.4.
Kortner, Sandra 1.1.9.4, 1.4. T 64.5, T 64.6 Kösker, Gani T 65.5, T 65 T 94.4 Kostyukhin, Vadim T 24 Koundal, Paras T 72.1, 6 Koskar, Argiris Koundal, Paras T 72.1, 6 Kovac, Neven T 102.5 Kovad, Oksana T 72.1, 6 Kovac, Neven T 102.5 Krabbe, Heiko DD 39.1 Kraetzschmar, Thomas Krabbe, Heiko DD 39.1, 7 Kraetzschmar, Thomas Kramer, Michael PV T 45.3, T 86.6, T 93.3, T 93. Krämer, Peter T 10 Kramer, Tobias T 9 Krasus, Johanna Wanda Krause, Johannes Krause, Johannes	3.2, T 61.1, 6.6, T 65.7, 5.5, T 42.6, • EP 12.8 AKPIK 4.1 5, •T 102.6 • T 35.6 • T 96.5 • •D 42.3 • T 30.7 III, T 31.2, 4 •T 30.7 III, T 31.2, 4 •T 6.7 •T 5.4 •T 6.7 •T 6.8 T 2.8 r 5.1, T 5.2 1.9, T 62.9 • DD 28.1 •T 68.7 T 2.8 r 5.1, T 5.2 1.9, T 62.9 • DD 28.1 •T 105.1 •T 105.5 •T 105.5 •T 37.5, 3, T 67.4, 1.4, ST 4.5, •T 38.7 •D 34.3 0D 44.2 .2, ST 1.6, 4, ST 4.5, 3, ST 7.4, T 34.1,

Kropf, Annabel
Krucker, Dirk 19.5, 19.6, 153.4,
AKPIK 2.9
kruecker, dirk AKPIK 2.4
Krueger, Christian
Krüger, Hans T 12.4, T 13.5, T 40.8,
T 40.9, T 95.2, T 97.6
Krüger, Katja
Kryukov, AlexanderAKPIK I.5
Kübelbäck Elorian T 00 5
Küchenmeister lens
Kuesters Roman •T 62.5
Kugathasan Thanushan T 40.8
Kuhlbusch. Tim GR 14.3. GR 14.4.
GR 14.6
Kuhlen, Marco•T 73.6
Kühn, Eileen
Kuhn, Jochen DD 12.3, DD 35.2
Kühn, Stefan T 79.8, T 79.9
Kühn, Susanne
Kuhr, Thomas 1 3.2, •1 27.1, 1 30.2,
1 30.4, 1 53.1, 1 53.2, 1 58.0, AKDIK 1 0
Kujiners Kolia GR16
Kuiper Bolf •FP 16
Kulesza, Anna MP 2.6. T 2.1. T 31 3
Kulgemeyer, Christoph DD 11.1
Kulii, Yaroslav•T 58.6
Kullmann, Ina EP 1.4
Kumar, Naveen
Kumaran, Sindhujha •EP 11.9, •T 83.9,
T 103.7, T 103.9
Kunert, Nina
Kunkel, Manuel
Kunz, Lucas
Kuprash Oleg T 87.3
Kurt Atila T 97.5
Kuruoalu Cenaiz T 72 9
Kutzner. Viktor
Kybic, Jan
Kyriacou, Alex EP 4.2
L. Biermann, Peter EP 11.4, T 83.4
La Cagnina, Salvatore•T 58.3
Labe, Finn•T 10.5, T 69.5
Lachenmaier, Tobias . 1 20.3, 1 20.9,
1 48.6, 1 48.8, 1 102.3
Lachner, Andreas
Lachner, Andreas DD 49.1 Lacker, Heiko T 37.1, T 43.4, T 64.9, T 68.2, T 94.2 Lad, Neha
Lachner, Andreas

Lechner, Peter
Lee, Chang•T 105.9
Lee, Ming-Yan T 26.3, T 90.6, •T 91.1
Lee Seh Wok T93
Lee Wei-Chieh
LEGEND-Kollaboration T 75 1 T 75 2
Lenmann, L
Lenmann, Lisa
Lehmkuhl, Dennis •AGPhil 2.2
Lehtinen, Antti DD 36.1
Leibrock, Barbara•DD 20.3
Leitgeb, Clara Elisabeth T 58.5, T 70.5
Lenke, Ralf R
Lensing, Philipp
Lenz, Alexander
T 56 4 T 85 2
Lenz Tatiana T 61 / T 62 2 T 62 3
Leonnardt, Andreas
Leontsinis, Stefanos
Leppla-Weber, David•T 105.8
Lerch, Michael ST 3.4
Leschke, Hajo MP 4.6
Lessing, Nadja•T 103.2
Leszczvnska, Agnieszka T 100.2
Leuna, Edwin K ST 9.2
Leverington Blake ST 4 2 ST 6 3
T 06 6
Lewing, Jonannes•DD 10.4
Lewis, Peter 1 3.3, 1 3.5, 1 30.6,
Т 30.9, Т 84.1, Т 84.2, Т 84.4, Т 84.8
Leyva Pernia, Daina•T 36.6
LHCb MightyTracker
group-Kollaboration
Lhotzky, Johannes •DD 23.1, DD 23.2,
DD 23.3. •DD 34.10. DD 46.2.
•DD 49 3
L hotzky Johannes F DD 30.2
Lhotzky, Johannes Frank DD 11.6
Linutzky, Solidines Hank DD 11.0
LI, Bao-Fei GR 12.1, MP 9.1, MP 10.2
LI, Cheng•1 36.7, 1 85.5
Li, Peilian
Liao, Chuan•T 66.8
Liberati, Stefano GR 1.5, GR 11.5
Liberatore, Marianna•T 76.6
Lichtenberger, Andreas •DD 42.1
Lichtenberger, Andreas•DD 42.1 Lika. Florian
Lichtenberger, Andreas •DD 42.1 Lika, Florian
Lichtenberger, Andreas•DD 42.1 Lika, Florian
Lichtenberger, Andreas•DD 42.1 Lika, Florian
Lichtenberger, Andreas•DD 42.1 Lika, Florian
Lichtenberger, Andreas•DD 42.1 Lika, Florian
Lichtenberger, Andreas•DD 42.1 Lika, Florian
Lichtenberger, Andreas•DD 42.1 Lika, Florian
Lichtenberger, Andreas•DD 42.1 Lika, Florian
Lichtenberger, Andreas •DD 42.1 Lika, Florian
Lichtenberger, Andreas •DD 42.1 Lika, Florian
Lichtenberger, Andreas •DD 42.1 Lika, Florian
Lichtenberger, Andreas •DD 42.1 Lika, Florian
Lichtenberger, Andreas •DD 42.1 Lika, Florian
Lichtenberger, Andreas •DD 42.1 Lika, Florian
Lichtenberger, Andreas •DD 42.1 Lika, Florian T 36.7 Lincetto, Massimiliano •T 18.7 Lindemann, Hannes Kasimir DD 8.1, DD 43.3 Linden, Lars T 91.3, •T 91.4 Lindenau, Philipp T 24.2 Linder, Jakob W T 31.1 Lindlahr, William DD 23.1, •DD 23.3 Lindman, Alec •T 49.3 Lindner, Axel GR 14.1 Lindner, Manfred T 23.1 Linhoff, Lena T 17.5, T 17.6 Link, Marco •T 7.5 Lipinski, Martin T 65.1, T 65.2, T 65.3, T 65.4, T 94.6
Lichtenberger, Andreas •DD 42.1 Lika, Florian
Lichtenberger, Andreas •DD 42.1 Lika, Florian
Lichtenberger, Andreas •DD 42.1 Lika, Florian T 36.7 Lincetto, Massimiliano •T 18.7 Lindemann, Hannes Kasimir DD 8.1, DD 43.3 Linden, Lars T 91.3, •T 91.4 Lindenau, Philipp T 24.2 Linder, Jakob W T 31.1 Lindlahr, WilliamDD 23.1, •DD 23.3 Lindman, Alec •T 49.3 Lindner, Manfred T 49.3 Lindner, Manfred •T 49.3 Lindner, Manfred •T 7.5 Lipinski, Martin T 65.1, T 65.2, T 65.3, T 65.4, T 94.6 Lipka, Katerina SYMD 1.3, T 7.1, T 32.2, T 65.8, T 88.2 Lisovskiv, Vitalii T 57.1, T 57.2, T 57.3.
Lichtenberger, Andreas •DD 42.1 Lika, Florian
Lichtenberger, Andreas •DD 42.1 Lika, Florian T 36.7 Lincetto, Massimiliano •T 18.7 Lindemann, Hannes Kasimir DD 8.1, DD 43.3 Linden, Lars T 91.3, •T 91.4 Lindenau, Philipp T 24.2 Linder, Jakob W T 31.1 Lindlahr, William DD 23.1, •DD 23.3 Lindman, Alec •T 49.3 Lindner, Axel
$eq:linear_line$
Lichtenberger, Andreas •DD 42.1 Lika, Florian 736.7 Lincetto, Massimiliano •T 18.7 Lindemann, Hannes Kasimir DD 8.1, DD 43.3 Linden, Lars T 91.3, •T 91.4 Linden, Lars T 91.3, •T 91.4 Linden, Lars T 91.3, •T 91.4 Linden, Jakob W T 31.1 Lindlahr, William DD 23.1, •DD 23.3 Lindman, Alec T 49.3 Lindmar, Axel GR 14.1 Lindner, Manfred T 23.1 Linhoff, Lena T 17.5, T 17.6 Link, Marco •T 7.5 Lipinski, Martin T 65.1, T 65.2, T 65.3, T 65.4, T 94.6 Lipka, Katerina SYMD 1.3, T 7.1, T 32.2, T 65.8, T 88.2 Lisovskyi, Vitalii T 57.1, T 57.2, T 57.3, •T 57.5, T 57.7 List, Jenny T 44.1, T 44.2, T 44.3, T 44.5, T 44.6, T 63.6, T 90.8 Litvinenko, Yuri MP 4.4 Litzenberger, Niklas DD 19.1 Liu, Runxuan T 20.2, T 20.5, •T 43.1, T 48.4 Llorente Anaya, Arturo T 46.9 Loaiza, Victor GR 6.2 Lobanov, Andrei T 77.4, T 77.5
Lichtenberger, Andreas •DD 42.1 Lika, Florian
Lichtenberger, Andreas •DD 42.1 Lika, Florian T 36.7 Lincetto, Massimiliano •T 18.7 Lindemann, Hannes Kasimir DD 8.1, DD 43.3 Linden, Lars T 91.3, •T 91.4 Lindenau, Philipp 22.2 Linder, Jakob W. T 31.1 Lindlahr, William DD 23.1, •DD 23.3 Lindman, Alec •T 49.3 Lindner, Axel GR 14.1 Lindner, Axel GR 14.1 Linhoff, Lena T 17.5, T 17.6 Link, Marco •T 7.5 Lipinski, Martin T 65.1, T 65.2, T 65.3, T 65.4, T 94.6 Lipka, Katerina SYMD 1.3, T 7.1, T 32.2, T 65.8, T 88.2 Lisovskyi, Vitalii T 57.1, T 57.2, T 57.3, •T 57.5, T 57.7 List, Jenny T 44.1, T 44.2, T 44.3, T 44.5, T 44.6, T 63.6, T 90.8 Litvinenko, Yuri MP 4.4 Litzenberger, Niklas •DD 19.1 Liu, Runxuan T 20.2, T 20.5, •T 43.1, T 48.4 Llorente Anaya, Arturo T 46.9 Loaiza, Victor GR 6.2 Lobanov, Andrei T 77.4, T 77.5 Lobanov, Andrei T 77.4, T 77.5 Lobanov, Andrei T 70.4, T 10.5, T 10.8, T 61.8, T 69.5, T 69.6, T 69.9
Lichtenberger, Andreas •DD 42.1 Lika, Florian
Lichtenberger, Andreas •DD 42.1 Lika, Florian 736.7 Lincetto, Massimiliano •T 18.7 Lindemann, Hannes Kasimir DD 8.1, DD 43.3 Linden, Lars T 91.3, •T 91.4 Linden, Lars T 91.3, •T 91.4 Linden, Lars T 91.3, •T 91.4 Linden, Jakob W T 31.1 Lindlahr, William DD 23.1, •DD 23.3 Lindman, Alec T 49.3 Lindner, Axel GR 14.1 Lindner, Manfred T 25.1 Linhoff, Lena T 17.5, T 17.6 Link, Marco •T 7.5 Lipinski, Martin T 65.1, T 65.2, T 65.3, T 65.4, T 94.6 Lipka, Katerina SYMD 1.3, T 7.1, T 32.2, T 65.8, T 88.2 Lisovskyi, Vitalii T 57.1, T 57.2, T 57.3, •T 57.5, T 57.7 List, Jenny T 44.1, T 44.2, T 44.3, T 44.5, T 44.6, T 63.6, T 90.8 Litvinenko, Yuri MP 4.4 Litzenberger, Niklas DD 19.1 Liu, Runxuan T 20.2, T 20.5, •T 43.1, T 48.4 Llorente Anaya, Arturo T 46.9 Loaiza, Victor GR 6.2 Lobanov, Andrei T 77.4, T 77.5 Lobanov, Andrei T 77.4, T 77.5 Lobanov, Andrei T 77.4, T 77.5 Lobanov, Andrei T 70.3, T 61.8, T 69.5, T 69.6, T 69.9 Lohmann, Wolfgang T 8.6 Loiacono, Eleonora T 70.3 Lokhov, Alexey T 21.3, T 21.4, T 21.5 Lomax, Tony PV I ombardi Daniele T 20.4
Lichtenberger, Andreas •DD 42.1 Lika, Florian T 36.7 Lincetto, Massimiliano •T 18.7 Lindemann, Hannes Kasimir DD 8.1, DD 43.3 Linden, Lars T 91.3, •T 91.4 Lindenu, Philipp T 24.2 Linder, Jakob W T 31.1 Lindlahr, William DD 23.1, •DD 23.3 Lindman, Alec •T 49.3 Lindner, Axel
Lichtenberger, Andreas •DD 42.1 Lika, Florian T 36.7 Lincetto, Massimiliano •T 18.7 Lindemann, Hannes Kasimir DD 8.1, DD 43.3 Linden, Lars T 91.3, •T 91.4 Lindenau, Philipp 7 24.2 Linder, Jakob W T 31.1 Lindlahr, William DD 23.1, •DD 23.3 Lindman, Alec T 49.3 Lindner, Axel GR 14.1 Lindner, Axel GR 14.1 Lindner, Axel T 17.5, T 17.6 Link, Marco T 17.5, T 17.6 Lipka, Katerina SYMD 1.3, T 7.1, T 32.2, T 65.8, T 88.2 Lisovskyi, Vitalii T 57.1, T 57.2, T 57.3, •T 57.5, T 57.7 List, Jenny T 44.1, T 44.2, T 44.3, T 44.5, T 44.6, T 63.6, T 90.8 Litvinenko, Yuri MP 4.4 Litzenberger, Niklas •DD 19.1 Liu, Runxuan T 20.2, T 20.5, •T 43.1, T 48.4 Llorente Anaya, Arturo T 46.9 Loaiza, Victor GR 6.2 Lobanov, Andrei T 77.4, T 77.5 Lobanov, Andrei T 77.4, T 77.5 Lobanov, Andrei T 77.4, T 77.3 Lobanov, Andrei T 77.4, T 77.5 Lobanov, Andrei T 71.3, T 21.4, T 21.5 Lomax, Tony PV 1 Lombardi, Daniele T 2.4 Lommler, Jan T 45.8 Loo Kai T 42.1 T 49.5
Lichtenberger, Andreas •DD 42.1 Lika, Florian 736.7 Lincetto, Massimiliano •T 18.7 Lindemann, Hannes Kasimir DD 8.1, DD 43.3 Linden, Lars
Lichtenberger, Andreas •DD 42.1 Lika, Florian 736.7 Lincetto, Massimiliano •T 18.7 Lindemann, Hannes Kasimir DD 8.1, DD 43.3 Linden, Lars 791.3, •T 91.4 Linden, Lars 791.3, •T 91.4 Linden, Jakob W T 31.1 Linden, Jakob W T 31.1 Linden, William DD 23.1, •DD 23.3 Lindman, Alec 49.3 Lindner, Axel GR 14.1 Lindner, Manfred 717.5 Lipinski, Martin T 65.1, T 65.2, T 65.3, T 65.4, T 94.6 Lipka, Katerina SYMD 1.3, T 7.1, T 32.2, T 65.8, T 88.2 Lisovskyi, Vitalii T 57.1, T 57.2, T 57.3, •T 57.5, T 57.7 List, Jenny T 44.1, T 44.2, T 44.3, T 44.5, T 44.6, T 63.6, T 90.8 Litvinenko, Yuri MP 4.4 Litzenberger, Niklas •DD 19.1 Liu, Runxuan T 20.2, T 20.5, •T 43.1, T 48.4 Llorente Anaya, Arturo T 46.9 Loaiza, Victor GR 6.2 Lobanov, Andrei T 77.4, T 77.5 Lobanov, Andrei T 77.4, T 77.5 Lobanov, Andrei T 77.4, T 77.5 Lobanov, Andrei T 70.3 Lokhov, Alexey T 21.3, T 21.4, T 21.5 Lomax, Tony PVI Lombardi, Daniele T 2.4 Loo, Kai T 43.1, T 48.5 Loope Z Solis, Alvaro T 76.1 Lione Dubardi, Daniele T 76.1 Lione Z 5015, Alvaro T 76.1
Lichtenberger, Andreas •DD 42.1 Lika, Florian 736.7 Lincetto, Massimiliano •T 18.7 Lindemann, Hannes Kasimir DD 8.1, DD 43.3 Linden, Lars T 91.3, •T 91.4 Lindenau, Philipp 74.2 Linder, Jakob W T 31.1 Lindlahr, William DD 23.1, •DD 23.3 Lindman, Alec T 49.3 Lindner, Axel
Lichtenberger, Andreas •DD 42.1 Lika, Florian 736.7 Lincetto, Massimiliano •T 18.7 Lindemann, Hannes Kasimir DD 8.1, DD 43.3 Linden, Lars 791.3, •T 91.4 Lindenau, Philipp 724.2 Linder, Jakob W T 31.1 Lindlahr, William DD 23.1, •DD 23.3 Lindman, Alec T 49.3 Lindner, Axel GR 14.1 Lindner, Manfred 73.1 Linhoff, Lena T 17.5, T 17.6 Link, Marco T 7.5 Lipinski, Martin T 65.1, T 65.2, T 65.3, T 65.4, T 94.6 Lipka, Katerina SYMD 1.3, T 7.1, T 32.2, T 65.8, T 88.2 Lisovskyi, Vitalii T 57.1, T 57.2, T 57.3, •T 57.5, T 57.7 List, Jenny T 44.1, T 44.2, T 44.3, T 44.5, T 44.6, T 63.6, T 90.8 Litvinenko, Yuri MP 4.4 Litzenberger, Niklas DD 19.1 Liu, Runxuan T 20.2, T 20.5, •T 43.1, T 48.4 Llorente Anaya, Arturo T 46.9 Loaiza, Victor GR 6.2 Lobanov, Antur T 10.4, T 10.5, T 10.8, T 61.8, T 69.5, T 69.6, T 69.9 Lohmann, Wolfgang T 8.6 Loiacono, Eleonora •T 70.3 Lokhov, Alexey T 21.3, T 21.4, T 21.5 Lomax, Tony •PV1 Lombardi, Daniele T 2.4 Lommler, Jan •F 45.8 Loo, Kai •T 43.1, T 48.5 Lopez Solis, Alvaro •T 43.1, T 78. Lorkowski, Florian *T 32.1
Lichtenberger, Andreas •DD 42.1 Lika, Florian 736.7 Lincetto, Massimiliano •T 18.7 Lindemann, Hannes Kasimir DD 8.1, DD 43.3 Linden, Lars 791.3, •T 91.4 Linden, Lars 791.3, •T 91.4 Linden, Jakob W T 31.1 Lindlahr, William DD 23.1, •DD 23.3 Lindman, Alec •T 49.3 Lindner, Axel GR 14.1 Lindner, Axel GR 14.1 Lindner, Manfred 723.1 Linhoff, Lena T 17.5, T 17.6 Link, Marco •T 7.5 Lipinski, Martin T 65.1, T 65.2, T 65.3, T 65.4, T 94.6 Lipka, Katerina SYMD 1.3, T 7.1, T 32.2, T 65.8, T 88.2 Lisovskyi, Vitalii T 57.1, T 57.2, T 57.3, •T 57.5, T 57.7 List, Jenny T 44.1, T 44.2, T 44.3, T 44.5, T 44.6, T 63.6, T 90.8 Litvinenko, Yuri MP 4.4 Litzenberger, Niklas DD 19.1 Liu, Runxuan T 20.2, T 20.5, •T 43.1, T 48.4 Llorente Anaya, Arturo T 46.9 Loaiza, Victor GR 6.2 Lobanov, Andrei T 77.4, T 77.5 Lobanov, Andrei T 77.4, T 77.5 Lobanov, Andrei T 77.4, T 77.5 Lobanov, Alexey T 21.3, T 21.4, T 21.5 Lomax, Tony PVI Lombardi, Daniele •T 2.4 Loommler, Jan T 43.1, T 48.5 Loope Solis, Alvaro T 43.1, T 48.5 Lovkowski, Florian T 32.1 Louis, Daniel 65.2
Lichtenberger, Andreas •DD 42.1 Lika, Florian T 36.7 Lincetto, Massimiliano T 18.7 Lindemann, Hannes Kasimir DD 8.1, DD 43.3 Linden, Lars T 91.3, •T 91.4 Lindenu, Philipp T 24.2 Linder, Jakob W T 31.1 Lindlahr, William DD 23.1, •DD 23.3 Lindman, Alec T 49.3 Lindner, Axel R 14.1 Lindner, Axel R 14.1 Lindner, Axel R 14.1 Lindner, Axel R 17.5, T 17.6 Link, Marco T 7.5, T 17.6 Lipka, Katerina T 55.1, T 65.2, T 65.3, T 65.4, T 94.6 Lipka, Katerina SYMD 1.3, T 7.1, T 32.2, T 65.8, T 88.2 Lisovskyi, Vitalii T 57.1, T 57.2, T 57.3, •T 57.5, T 57.7 List, Jenny T 44.1, T 44.2, T 44.3, T 44.5, T 44.6, T 63.6, T 90.8 Litvinenko, Yuri MP 4.4 Litzenberger, Niklas DD 19.1 Liu, Runxuan T 20.2, T 20.5, •T 43.1, T 48.4 Llorente Anaya, Arturo T 46.9 Lobanov, Andrei T 77.4, T 77.5 Lobanov, Artur T 10.4, T 10.5, T 10.8, T 61.8, T 69.5, T 69.6, T 69.9 Lohmann, Wolfgang T 8.6 Loiacono, Eleonora T 7.3, Lokhov, Alexey T 21.3, T 21.4, T 21.5 Lomax, Tony PV 1 Lombardi, Daniele T 7.4, T 76.1 Lopez Coto, Ruben T 13.2 Lovis, Daniel T 65.2 Lovatti, Giullio T 65.2
Lichtenberger, Andreas •DD 42.1 Lika, Florian
Lichtenberger, Andreas •DD 42.1 Lika, Florian
Lichtenberger, Andreas •DD 42.1 Lika, Florian

Luc, Luchmann, Michel•T 25.3, T 25.3, T 31.2

Lück Thomas T	2 2 T 20 A
LUCK, Inomas	3.2, 1 30.4
Lüders. Christina . •DD 25.3	3. •DD 33.4
Ludhova Livia ED 9.1 T 20	2 7 20 5
Luunova, Livia •EP 8.1, 1 20	J.Z, I ZU.S,
T 43.1. T 48.4. •T 80.1. T 10	3.7.
T 102 0	,
1 103.9	
Ludwig Tobias	DD 50 1
Lunr, Armin ST 3	3.3, I 53.3
Lüledorf Sebastian	T 52 1
	I JZ.I
Lunde, Karin	DD 34.4
Lunharger Mishael T 6	01 T 60 7
Lupperger, Michael 1 6	8.1, I 09.7,
T 97.1	
Lunharger Michael	T 20 0
Lupperger, Micheal	1 39.8
Lutz Wolfgang DD 18	3 DD 48 1
DD 40.0	0,00 1011)
•DD 48.2	
Lyons Fairburst	•T 42 1
	•1 84./
Machado Paveras Allan	T 51 4
Wachaut Fayeras, Allan	
Machatschek, Moritz	T 49.7
Maalaaaa Danial	00040
	DD 24.3
Mackowiak, Patrick	T 60.9
Maazawalay Lukaa	DD 04 10
Maczewsky, Lukas	. DD 34.1Z
Madar Romain	T 34 7
Madan Malfaan	
Mader, worrgang	1 62.9
Madiarska Maria	FP 12 9
	EI 12.2
MADMAX-Kollaboration	138.6,
T 105 1 T 105 2 T 105 4 T	105 5
T 10E 6 T 10E 7 T 10E 6 T	105.0
i iud.o, i iud./, i iud.8, T	105.9
Madroñero Javier	GR 6 2
Mass Luly	
IVIAES, LUKAS	•EP 3.3
Mändefessel Sven T1	3 6 T 0/ 1
11111111111111111111111111111111111111	0.0, 1 94.1
Magee, Ryan	GR 9.5
Magerya Vitaly	ТОТ
	· · · · · I Z./
Maggio, Elisa	•GR 5.6
MAGIC Kollaboration T1	75 7 17 7
	7.0, I I/./,
T 101.3. T 17.6	
Major Stofan ST (4 T (5	5 .T 45 4
wale, sterall . 31 0.4, 1 03.	, • 1 0.0.0,
F 65.7, T 94.4	
Mäkelä Toni	•T 32 2
Malaina and David T 50	F T 00 0
Maksimovic, David 1 59	.5, •1 99.9
Malaguias, Isabel	GP 2.3
Malowski Frank	T 72 Q
Maltsev, Kiril	AGPhil 3.2
Malvetti Emanuel MP 5	2 • MP 5 3
	2, 1011 0.0
Malyshev, Dmitry 1 4	5.2, 1 51.8
Manai Chokri	•MP 4 6
Manconi, Silvia 14	5.3, 1 86.6
Mancuso Michele	T 104 6
Manual Annuality	DD 11 4
Mandl, Angelika	•DD 11.4
Mandl, Angelika	•DD 11.4 •MP 7.3
Mandl, Angelika Mandrysch, Jan	•DD 11.4 •MP 7.3
Mandl, Angelika Mandrysch, Jan Maneschg, Werner	•DD 11.4 •MP 7.3 T 23.1
Mandl, Angelika Mandrysch, Jan Maneschg, Werner Manfred, Geilhaupt	•DD 11.4 •MP 7.3 T 23.1 •GR 10.5
Mandl, Angelika Mandrysch, Jan Maneschg, Werner Manfred, Geilhaupt	•DD 11.4 •MP 7.3 T 23.1 •GR 10.5 2 •ST 8.3
Mandl, Angelika Mandrysch, Jan Maneschg, Werner Manfred, Geilhaupt Manke, Henning ST 8.	•DD 11.4 •MP 7.3 T 23.1 •GR 10.5 .2, •ST 8.3,
Mandl, Angelika Mandrysch, Jan Manschg, Werner Manfred, Geilhaupt Manke, Henning ST 8. ST 8.4	•DD 11.4 •MP 7.3 T 23.1 •GR 10.5 .2, •ST 8.3,
Mandl, Angelika Mandrysch, Jan Maneschg, Werner Manfred, Geilhaupt Manke, Henning ST 8. ST 8.4 Mann, Gottfried EP 11.	•DD 11.4 •MP 7.3 T 23.1 •GR 10.5 .2, •ST 8.3, 7, EP 12.4,
Mandl, Angelika Mandrysch, Jan Mansechg, Werner Manfred, Geilhaupt Manke, Henning ST 8.4 Mann, Gottfried T 83 7	•DD 11.4 •MP 7.3 T 23.1 •GR 10.5 .2, •ST 8.3, 7, EP 12.4,
Mandl, Angelika Mandrysch, Jan Maneschg, Werner Manfred, Geilhaupt Manke, Henning ST 8.4 Mann, Gottfried T 83.7 Manne J Therese	•DD 11.4 •MP 7.3 T 23.1 •GR 10.5 2, •ST 8.3, 7, EP 12.4,
Mandl, Angelika Mandrysch, Jan Mansechg, Werner Manfred, Geilhaupt Manke, Henning ST 8.4 Mann, Gottfried T 83.7 Mannel, Thomas	•DD 11.4 •MP 7.3 T 23.1 •GR 10.5 2, •ST 8.3, 7, EP 12.4, T 56.5
Mandl, Angelika Mandrysch, Jan Manfred, Werner Manfred, Geilhaupt Manke, Henning ST 8.4 Mann, Gottfried T 83.7 Mannel, Thomas Mans, Leon	•DD 11.4 •MP 7.3 T 23.1 •GR 10.5 2,•ST 8.3, 7, EP 12.4, T 56.5 T 39.5
Mandl, Angelika Mandrysch, Jan Maneschg, Werner Manfred, Geilhaupt Manke, Henning ST 8.4 Mann, Gottfried T 83.7 Mannel, Thomas Mans, Leon Mantegazzini Eederica	•DD 11.4 •MP 7.3 T 23.1 •GR 10.5 2, •ST 8.3, 7, EP 12.4, T 56.5 T 39.5 T 102 5
Mandl, Angelika Mandrysch, Jan Manfred, Geilhaupt Manfred, Geilhaupt Manke, Henning ST 8.4 Mann, Gottfried T 83.7 Mannel, Thomas Mans, Leon Mantegazzini, Federica	•DD 11.4 •MP 7.3 T 23.1 •GR 10.5 2, •ST 8.3, 7, EP 12.4, T 56.5 T 39.5 T 102.5,
Mandl, Angelika Mandrysch, Jan Manferd, Geilhaupt Manfred, Geilhaupt Manke, Henning ST 8.4 Mann, Gottfried T 83.7 Mannel, Thomas Mans, Leon Mantegazzini, Federica T 102.6	•DD 11.4 •MP 7.3 •GR 10.5 2, •ST 8.3, 7, EP 12.4, T 56.5 T 39.5 T 102.5,
Mandl, Angelika Mandrysch, Jan Mansechg, Werner Manfred, Geilhaupt Manke, Henning ST 8.4 Mann, Gottfried T 83.7 Mannel, Thomas Mans, Leon Mantegazzini, Federica T 102.6 Manthei, Alina	•DD 11.4 •MP 7.3 T 23.1 •GR 10.5 2, •ST 8.3, 7, EP 12.4, T 56.5 T 39.5 T 102.5, •T 3.5
Mandl, Angelika Mandrysch, Jan Manfred, Geilhaupt Manfred, Geilhaupt Manke, Henning ST 8.4 Mann, Gottfried T 83.7 Mannel, Thomas Mans, Leon Mantegazzini, Federica T 102.6 Manthei, Alina Marskie Vaesilioe	•DD 11.4 •MP 7.3 T 23.1 •GR 10.5 2, •ST 8.3, 7, EP 12.4, T 56.5 T 39.5 T 102.5, •T 3.5
Mandl, Angelika Mandrysch, Jan Mansechg, Werner Manfred, Geilhaupt Manke, Henning ST 8.4 Mann, Gottfried T 83.7 Mannel, Thomas Mans, Leon Mantegazzini, Federica T 102.6 Manthei, Alina Marakis, Vassilios	•DD 11.4 •MP 7.3 T 23.1 •GR 10.5 2, •ST 8.3, 7, EP 12.4, T 56.5 T 39.5 T 102.5, •T 3.5 •DD 35.12
Mandl, Angelika Mandrysch, Jan Mansechg, Werner Manfred, Geilhaupt Manke, Henning ST 8.4 Mann, Gottfried T 83.7 Mannel, Thomas Mans, Leon Mantegazzini, Federica T 102.6 Manthei, Alina Maria, Vassilios Marin Tobon, Cesar Augusto	•DD 11.4 •MP 7.3 T 23.1 •GR 10.5 2, •ST 8.3, 7, EP 12.4, T 56.5 T 39.5 T 102.5, •T 3.5 •DD 35.12 T 40.8
Mandl, Angelika Mandrysch, Jan Mansechg, Werner Manfred, Geilhaupt Manke, Henning ST 8.4 Mann, Gottfried T 83.7 Mannel, Thomas Mans, Leon Mantegazzini, Federica T 102.6 Manthei, Alina Marakis, Vassilios Maritor Max Mar T Marker Max Mar	•DD 11.4 •MP 7.3 T 23.1 •GR 10.5 2, •ST 8.3, 7, EP 12.4, T 56.5 T 39.5 T 102.5, •T 3.5 •DD 35.12 T 40.8 18 T 62 9
Mandl, Angelika Mandrysch, Jan Mansechg, Werner Manfred, Geilhaupt Manke, Henning ST 8.4 Mann, Gottfried T 83.7 Mannel, Thomas Mans, Leon Mantegazzini, Federica T 102.6 Manthei, Alina Marakis, Vassilios Marin Tobon, Cesar Augusto Märker, Max T 4	•DD 11.4 •MP 7.3 T 23.1 •GR 10.5 2, •ST 8.3, 7, EP 12.4, T 56.5 T 39.5 T 102.5, •T 3.5 •DD 35.12 T 40.8 1.8, T 62.9
Mandl, Angelika Mandrysch, Jan Mansechg, Werner Manfred, Geilhaupt Manke, Henning ST 8.4 Mann, Gottfried T 83.7 Mannel, Thomas Mans, Leon Mantegazzini, Federica T 102.6 Manthei, Alina Marakis, Vassilios Marin Tobon, Cesar Augusto Märker, Max Markhoos, Ahmed	•DD 11.4 •MP 7.3 T 23.1 •GR 10.5 2, •ST 8.3, 7, EP 12.4, T 56.5 T 39.5 T 102.5, •T 3.5 •DD 35.12 T 40.8 1.8, T 62.9 •T 61.3
Mandl, Angelika Mandrysch, Jan Mansechg, Werner Manfred, Geilhaupt Mankred, Geilhaupt Manke, Henning ST 8.4 Mann, Gottfried T 83.7 Mannel, Thomas Mans, Leon Mantegazzini, Federica T 102.6 Manthei, Alina Marakis, Vassilios Marin Tobon, Cesar Augusto Märker, Max Markhoos, Ahmed Marzaya, Hanna	•DD 11.4 •MP 7.3 T 23.1 •GR 10.5 2, •ST 8.3, 7, EP 12.4, T 56.5 T 39.5 T 102.5, •T 3.5 •DD 35.12 T 40.8 1.8, T 62.9 •T 61.3 •GR 14 2
Mandl, Angelika Mandrysch, Jan Mansechg, Werner Manfred, Geilhaupt Manfred, Geilhaupt Manke, Henning ST 8.4 Mann, Gottfried T 83.7 Mannel, Thomas Mans, Leon Mantegazzini, Federica T 102.6 Manthei, Alina Marakis, Vassilios Marin Tobon, Cesar Augusto Märker, Max Markhoos, Ahmed Marode In decerting	•DD 11.4 •MP 7.3 T 23.1 •GR 10.5 2, •ST 8.3, 7, EP 12.4, T 56.5 T 39.5 T 102.5, •T 3.5 •DD 35.12 •T 40.8 1.8, T 62.9 •T 61.3 •GR 14.2
Mandl, Angelika Mandrysch, Jan Mansechg, Werner Manfred, Geilhaupt Manke, Henning ST 8.4 Mann, Gottfried T 83.7 Mannel, Thomas Mans, Leon Mantegazzini, Federica T 102.6 Manthei, Alina Marakis, Vassilios Marin Tobon, Cesar Augusto Märker, Max Marzava, Hanna Marrodán Undagoitia, Teresa	•DD 11.4 •MP 7.3 T 23.1 •GR 10.5 2, •ST 8.3, 7, EP 12.4, T 56.5 T 39.5 T 102.5, •T 3.5 •DD 35.12 T 40.8 1.8, T 62.9 •GR 14.2 •GR 14.2
Mandl, Angelika Mandrysch, Jan Maneschg, Werner Manfred, Geilhaupt Manke, Henning ST 8.4 Mann, Gottfried T 83.7 Mannel, Thomas Mans, Leon Mantegazzini, Federica T 102.6 Manthei, Alina Marakis, Vassilios Marin Tobon, Cesar Augusto Märker, Max Markhoos, Ahmed Markhoos, Ahmed Marodán Undagoitia, Teresa Marsh, Allison	•DD 11.4 •MP 7.3 T 23.1 •GR 10.5 2, •ST 8.3, 7, EP 12.4, T 56.5 T 39.5 T 102.5, •T 3.5 •DD 35.12 T 40.8 1.8, T 62.9 •T 61.3 •GR 14.2 •PV IV
Mandl, Angelika Mandrysch, Jan Mansechg, Werner Manfred, Geilhaupt Manke, Henning ST 8.4 Mann, Gottfried T 83.7 Mannel, Thomas Mans, Leon Mantegazzini, Federica T 102.6 Manthei, Alina Marakis, Vassilios Marin Tobon, Cesar Augusto Märker, Max Markos, Ahmed Markoos, Ahmed Martodan Undagoitia, Teresa Marsh, Allison Martela	•DD 11.4 •MP 7.3 T 23.1 •GR 10.5 2, •ST 8.3, 7, EP 12.4, T 56.5 T 39.5 T 102.5, •T 3.5 •DD 35.12 T 40.8 1.8, T 62.9 •GR 14.2 •GR 14.2 T 78.2
Mandl, Angelika Mandrysch, Jan Maneschg, Werner Manfred, Geilhaupt Manke, Henning ST 8.4 Mann, Gottfried T 83.7 Mannel, Thomas Mans, Leon Mantegazzini, Federica T 102.6 Manthei, Alina Marakis, Vassilios Marin Tobon, Cesar Augusto Märker, Max Markhoos, Ahmed Markozava, Hanna Marrodán Undagoitia, Teresa Marsh, Allison Marsteller, Alexander	•DD 11.4 •MP 7.3 T 23.1 •GR 10.5 2, •ST 8.3, 7, EP 12.4, T 56.5 T 39.5 T 102.5, •T 3.5 •DD 35.12 T 40.8 1.8, T 62.9 •T 61.3 •GR 14.2 T 78.2 •PV IV T 49.6
Mandl, Angelika Mandrysch, Jan Mansechg, Werner Manfred, Geilhaupt Manke, Henning ST 8.4 Mann, Gottfried T 83.7 Mannel, Thomas Mans, Leon Mantegazzini, Federica T 102.6 Manthei, Alina Marakis, Vassilios Marin Tobon, Cesar Augusto Märker, Max Markhoos, Ahmed Martodan Undagoitia, Teresa Marsh, Allison Marsteller, Alexander Martens, Stephan Martens, Stephan Martens, Stephan Martens, Stephan Martens, Stephan Martens, Stephan Martens, Stephan Martens, Stephan Martens, Stephan Martens, Stephan	•DD 11.4 •MP 7.3 T 23.1 •GR 10.5 2, •ST 8.3, 7, EP 12.4, T 56.5 T 39.5 T 102.5, •T 3.5 •DD 35.12 •T 40.8 1.8, T 62.9 •GR 14.2 T 78.2 T 78.2 T 49.6 6.1, T 66.5
Mandl, Angelika Mandrysch, Jan Maneschg, Werner Manfred, Geilhaupt Manke, Henning ST 8.4 Mann, Gottfried T 83.7 Mannel, Thomas Mans, Leon Mantegazzini, Federica T 102.6 Manthei, Alina Marakis, Vassilios Marin Tobon, Cesar Augusto Märker, Max Markhoos, Ahmed Marozava, Hanna Marodah Undagoitia, Teresa Marsh, Allison Marsteller, Alexander Martin, Stephan Martin Jill	•DD 11.4 •MP 7.3 T 23.1 •GR 10.5 2, •ST 8.3, 7, EP 12.4, T 56.5 T 39.5 T 102.5, •T 3.5 •DD 35.12 •T 61.3 •GR 14.2 •GR 14.2 •GR 14.2 •PV IV T 49.6 6.1, T 66.5 T 2.6
Mandl, Angelika Mandrysch, Jan Mansechg, Werner Manfred, Geilhaupt Manke, Henning ST 8.4 Mann, Gottfried T 83.7 Mannel, Thomas Mans, Leon Mantegazzini, Federica T 102.6 Manthei, Alina Marakis, Vassilios Marin Tobon, Cesar Augusto Märker, Max Markoos, Ahmed Markoos, Ahmed Marodan Undagoitia, Teresa Marsh, Allison Marsteller, Alexander Martens, Stephan T 6 Martini, Till	•DD 11.4 •MP 7.3 T 23.1 •GR 10.5 2, •ST 8.3, 7, EP 12.4, T 56.5 T 39.5 T 102.5, •T 3.5 •DD 35.12 •T 40.8 1.8, T 62.9 •GR 14.2 •GR 14.2 •GR 14.2 •GR 14.2 T 78.2 •C 16.3 T 49.6 6.1, T 66.5 T 2.6
Mandl, Angelika Mandrysch, Jan Mansechg, Werner Manfred, Geilhaupt Manke, Henning ST 8.4 Mann, Gottfried T 83.7 Mannel, Thomas Mans, Leon Mantegazzini, Federica T 102.6 Manthei, Alina Marakis, Vassilios Marin Tobon, Cesar Augusto Märker, Max Marakis, Vassilios Marin Tobon, Cesar Augusto Märker, Max Marzava, Hanna Marzodan Undagoitia, Teresa Marsh, Allison Marsteller, Alexander Martens, Stephan T 6 Martini, Till Martinov, Tommy	•DD 11.4 •MP 7.3 T 23.1 •GR 10.5 2, •ST 8.3, 7, EP 12.4, T 56.5 T 39.5 T 102.5, •T 3.5 •DD 35.12 •T 61.3 •GR 14.2 •GR 14.2 •PV IV T 49.6 6.1, T 66.5 T 2.6 •T 95.5
Mandl, Angelika Mandrysch, Jan Mansechg, Werner Manfred, Geilhaupt Manke, Henning ST 8.4 Mann, Gottfried T 83.7 Mannel, Thomas Mans, Leon Mantegazzini, Federica T 102.6 Manthei, Alina Marakis, Vassilios Marin Tobon, Cesar Augusto Märker, Max Markos, Ahmed Markos, Ahmed Martozava, Hanna Marrodán Undagoitia, Teresa Marsh, Allison Marsteller, Alexander Martens, Stephan T 6 Martini, Till Martinov, Tommy Martius, Florin	•DD 11.4 •MP 7.3 T 23.1 •GR 10.5 2, •ST 8.3, 7, EP 12.4, T 56.5 T 39.5 T 102.5, •T 3.5 •DD 35.12 •T 40.8 1.8, T 62.9 •T 61.3 •GR 14.2 T 78.2 •PV IV T 49.6 6.1, T 66.5 T 29.5 T 25.1
Mandl, Angelika Mandrysch, Jan Mansechg, Werner Manfred, Geilhaupt Manfred, Geilhaupt Manke, Henning ST 8.4 Mann, Gottfried T 83.7 Mannel, Thomas Mans, Leon Mantegazzini, Federica T 102.6 Manthei, Alina Marakis, Vassilios Marin Tobon, Cesar Augusto Märker, Max Markhoos, Ahmed Marzava, Hanna Marzodan Undagoitia, Teresa Marsteller, Alexander Martens, Stephan Martinov, Tommy Martius, Florin Martinov, Hohann	•DD 11.4 •MP 7.3 T 23.1 •GR 10.5 2, •ST 8.3, 7, EP 12.4, T 56.5 T 39.5 T 102.5, •T 3.5 •DD 35.12 •T 40.8 1.8, T 62.9 •T 61.3 •GR 14.2 •FV IV •PV IV T 49.6 6.1, T 66.5 T 2.6 •T 95.5 T 52.1 •T 92.9
Mandl, Angelika Mandrysch, Jan Mansechg, Werner Manfred, Geilhaupt Manke, Henning ST 8.4 Mann, Gottfried T 83.7 Mannel, Thomas Mans, Leon Mantegazzini, Federica T 102.6 Manthei, Alina Marakis, Vassilios Marin Tobon, Cesar Augusto Märker, Max Marozava, Hanna Marozava, Hanna Marodán Undagoitia, Teresa Marsh, Allison Marsteller, Alexander Martens, Stephan Martinov, Tommy Martus, Florin Marty, Johann	•DD 11.4 •MP 7.3 T 23.1 •GR 10.5 2, •ST 8.3, 7, EP 12.4, T 56.5 T 39.5 T 102.5, •T 3.5 •DD 35.12 •T 61.3 •GR 14.2 •F 61.3 •GR 14.2 •PV IV T 49.6 6.1, T 66.5 T 2.6 •T 95.5 T 103.8
Mandl, Angelika Mandrysch, Jan Mansechg, Werner Manfred, Geilhaupt Manfred, Geilhaupt Manke, Henning ST 8.4 Mann, Gottfried T 83.7 Mannel, Thomas Mans, Leon Mantegazzini, Federica T 102.6 Manthei, Alina Marakis, Vassilios Marin Tobon, Cesar Augusto Märker, Max Markoos, Ahmed Markhoos, Ahmed Markoava, Hanna Marsh, Allison Marteller, Alexander Martens, Stephan Martinov, Tommy Martius, Florin Marty, Johann Marz, Martin Marz, Martin Martin .	•DD 11.4 •MP 7.3 T 23.1 •GR 10.5 2, •ST 8.3, 7, EP 12.4, T 56.5 T 39.5 T 102.5, •T 3.5 •DD 35.12 •T 40.8 1.8, T 62.9 •T 61.3 •GR 14.2 T 78.2 T 49.6 6.1, T 66.5 T 2.6 •T 95.5 T 52.1 •T 103.8 6.3, T 36.5
Mandl, Angelika Mandrysch, Jan Mansechg, Werner Manfred, Geilhaupt Manke, Henning ST 8.4 Mann, Gottfried T 83.7 Mannel, Thomas Mans, Leon Mantegazzini, Federica T 102.6 Manthei, Alina Marakis, Vassilios Marin Tobon, Cesar Augusto Märker, Max Markhoos, Ahmed Markhoos, Ahmed Markoda, Misson Marsteller, Alexander Marsteller, Alexander Martens, Stephan Martinov, Tommy Martinov, Tommy Martius, Florin Marsty, Johann Marz, Martin Marstillucia Masteller 16 P 7	•DD 11.4 •MP 7.3 T 23.1 •GR 10.5 2, •ST 8.3, 7, EP 12.4, T 56.5 T 39.5 T 102.5, •T 3.5 •DD 35.12 •T 40.8 1.8, T 62.9 •T 61.3 •PV IV T 49.6 6.1, T 66.5 T 2.6 •T 52.1 •T 103.8 6.3, T 36.5 1.2 T 34.3
Mandl, Angelika Mandrysch, Jan Mansechg, Werner Manfred, Geilhaupt Manke, Henning ST 8.4 Mann, Gottfried T 83.7 Mannel, Thomas Mans, Leon Mantegazzini, Federica T 102.6 Manthei, Alina Marakis, Vassilios Marin Tobon, Cesar Augusto Märker, Max Markoos, Ahmed Martodan Undagoitia, Teresa Marsh, Allison Martodan Undagoitia, Teresa Marsh, Allison Marteller, Alexander Martens, Stephan Martinov, Tommy Martinov, Tommy Martinov, Tommy Marty, Johann Marz, Martin Marton, Ta 16, 9, 73 Masetti, Lucia	•DD 11.4 •MP 7.3 T 23.1 •GR 10.5 2, •ST 8.3, 7, EP 12.4, T 56.5 T 39.5 T 102.5, •T 35.5 •DD 35.12 •T 40.8 1.8, T 62.9 •T 61.3 •GR 14.2 T 78.2 T 49.6 6.1, T 66.5 T 2.6 T 29.5 T 52.1 •T 103.8 6.3, T 36.5 4.2, T 34.3 2, T 74.2
Mandl, Angelika Mandrysch, Jan Maneschg, Werner Manfred, Geilhaupt Manke, Henning ST 8.4 Mann, Gottfried T 83.7 Mannel, Thomas Mans, Leon Mantejazzini, Federica T 102.6 Manthei, Alina Marakis, Vassilios Marin Tobon, Cesar Augusto Märker, Max Markhoos, Ahmed Marozava, Hanna Marrodán Undagoitia, Teresa Marsh, Allison Marsteller, Alexander Martens, Stephan Martin, Jill Martinov, Tommy Martius, Florin Martyn, Johann Mars, Martin Mars, T 19, T 3 Masteti, Lucia T 10, T 35, 2, T 35, 3, T 41.1, T 59.	•DD 11.4 •MP 7.3 T 23.1 •GR 10.5 2, •ST 8.3, 7, EP 12.4, T 56.5 T 39.5 T 102.5, •T 3.5 •DD 35.12 T 40.8 1.8, T 62.9 •T 61.3 •GR 14.2 T 49.6 6.1, T 66.5 T 2.6 •T 9.5 T 2.6 •T 9.5 T 103.8 6.3, T 36.5 1.2, T 34.3, 2, T 97.5
Mandl, Angelika Mandrysch, Jan Mansechg, Werner Manfred, Geilhaupt Manke, Henning ST 8.4 Mann, Gottfried T 83.7 Mannel, Thomas Mans, Leon Mantegazzini, Federica T 102.6 Manthei, Alina Marakis, Vassilios Marin Tobon, Cesar Augusto Märker, Max Markhoos, Ahmed Markoos, Ahmed Martodan Undagoitia, Teresa Marsh, Allison Marsteller, Alexander Martens, Stephan Martens, Stephan Martinov, Tommy Martun, Johann Marty, Johann Marz, Martin Masteti, Lucia T 16.9, T 35. Mastrolorenzo, Luca Mastoner	•DD 11.4 •MP 7.3 T 23.1 •GR 10.5 2, •ST 8.3, 7, EP 12.4, T 56.5 T 39.5 T 102.5, •T 3.5 •DD 35.12 •T 40.8 1.8, T 62.9 •T 61.3 •GR 14.2 •GR 14.2 •GR 14.2 T 49.6 6.1, T 66.5 T 2.6 T 2.5 T 103.8 6.3, T 36.5 1.2, T 34.3, 2, T 97.5 .3, •T 82.3,
Mandl, Angelika Mandrysch, Jan Mansechg, Werner Manfred, Geilhaupt Manke, Henning ST 8.4 Mann, Gottfried T 83.7 Mannel, Thomas Mans, Leon Mantegazzini, Federica T 102.6 Manthei, Alina Marakis, Vassilios Marin Tobon, Cesar Augusto Märker, Max Marakis, Vassilios Marin Tobon, Cesar Augusto Märker, Max Martakos, Ahmed Martozava, Hanna Martodán Undagoitia, Teresa Marsh, Allison Marteller, Alexander Martens, Stephan Martinov, Tommy Martius, Florin Martin, Till Martinov, Tommy Martius, Florin Mart, Martin Mart, Johann Marz, Martin Mart, Jas, 2, T 35.3, T 41.1, T 59. Mastrolorenzo, Luca T 89.1, T 89.2, T 90.6	•DD 11.4 •MP 7.3 T 23.1 •GR 10.5 2, •ST 8.3, 7, EP 12.4, T 56.5 T 39.5 T 102.5, •T 3.5 •DD 35.12 •T 61.3 •GR 14.2 •F 61.3 •GR 14.2 •PV IV T 49.6 6.1, T 66.5 T 2.6 •T 95.5 T 103.8 6.3, T 36.5 1.2, T 34.3, 2, T 97.5 3, •T 82.3,
Mandl, Angelika Mandrysch, Jan Mansechg, Werner Manfred, Geilhaupt Manke, Henning ST 8.4 Mann, Gottfried T 83.7 Mannel, Thomas Mans, Leon Mantejazzini, Federica T 102.6 Manthei, Alina Marakis, Vassilios Marin Tobon, Cesar Augusto Märker, Max Markos, Ahmed Markos, Ahmed Markos, Ahmed Markos, Ahmed Markos, Ahmed Marsh, Allison Marsteller, Alexander Martens, Stephan Martinov, Tommy Martinov, Tommy Martinov, Tommy Marty, Johann Marsh, Johann Marso, Jattin Marto, T 16.9, T Mastrolorenzo, Luca T 89.1, T 89.2, T 90.6	•DD 11.4 •MP 7.3 T 23.1 •GR 10.5 2, •ST 8.3, 7, EP 12.4, T 56.5 T 39.5 T 102.5, •T 3.5 •DD 35.12 •T 40.8 1.8, T 62.9 •T 61.3 •GR 14.2 •PV IV T 49.6 6.1, T 66.5 T 2.6 •T 95.5 T 52.1 •T 103.8 6.3, T 36.5 2, T 97.5 3, •T 82.3, •DD 19.2
Mandl, Angelika Mandrysch, Jan Mansechg, Werner Manfred, Geilhaupt Manke, Henning ST 8.4 Mann, Gottfried T 83.7 Mannel, Thomas Mans, Leon Mantegazzini, Federica T 102.6 Manthei, Alina Marakis, Vassilios Marin Tobon, Cesar Augusto Märker, Max Markhoos, Ahmed Martodan Undagoitia, Teresa Marsh, Allison Martodán Undagoitia, Teresa Marsh, Allison Marteller, Alexander Martens, Stephan Martinov, Tommy Martius, Florin Marty, Johann Marto, Johann Marto, Tar 53. Mastti, Lucia T 16.9, T 32 T 35.2, T 35.3, T 41.1, T 59. Mastrolorenzo, Luca T 89.1, T 89.2, T 90.6 Matejak Cvenic, Karolina	•DD 11.4 •MP 7.3 T 23.1 •GR 10.5 2, •ST 8.3, 7, EP 12.4, T 56.5 T 39.5 T 102.5, •T 3.5 •DD 35.12 T 40.8 1.8, T 62.9 •T 61.3 •GR 14.2 T 49.6 6.1, T 66.5 T 2.6 T 2.6 T 2.5 T 103.8 6.3, T 36.5 1.2, T 34.3, 2, T 97.5 3, •T 82.3, •D 18.2
Mandl, Angelika Mandrysch, Jan Mansechg, Werner Manfred, Geilhaupt Manke, Henning ST 8.4 Mann, Gottfried T 83.7 Mannel, Thomas Mans, Leon Mantejazzini, Federica T 102.6 Manthei, Alina Marakis, Vassilios Marin Tobon, Cesar Augusto Märker, Max Markoos, Ahmed Markoos, Ahmed Marcava, Hanna Marrodán Undagoitia, Teresa Marsh, Allison Marsteller, Alexander Martens, Stephan Martinov, Tommy Martinov, Tommy Martinov, Tonmy Martyn, Johann Marstolorenzo, Luca T 89.1, T 89.2, T 90.6 Matejak Cvenic, Karolina	•DD 11.4 •MP 7.3 T 23.1 •GR 10.5 2, •ST 8.3, 7, EP 12.4, T 56.5 T 39.5 T 102.5, •T 3.5 •DD 35.12 •T 40.8 1.8, T 62.9 •T 61.3 •GR 14.2 •PV IV T 49.6 6.1, T 66.5 T 2.6 •T 95.5 T 52.1 •T 103.8 6.3, T 36.5 4.2, T 34.3, 2, T 97.5 3, •T 82.3, •DD 18.2 MP 4.2
Mandl, Angelika Mandrysch, Jan Mansechg, Werner Manfred, Geilhaupt Manke, Henning ST 8.4 Mann, Gottfried T 83.7 Mannel, Thomas Mans, Leon Mantegazzini, Federica T 102.6 Manthei, Alina Marakis, Vassilios Marin Tobon, Cesar Augusto Märker, Max Markhoos, Ahmed Marozava, Hanna Marodán Undagoitia, Teresa Marsh, Allison Marsteller, Alexander Martens, Stephan Martinov, Tommy Martinov, Tommy Martinov, Tommy Martinov, Toshann Marto, Johann Marto, Johann Marto, Johann Marto, Johann Marto, Jasz, T 35.3, T 41.1, T 59. Mastrolorenzo, Luca T 89.1, T 89.2, T 90.6 Matthaiakakis, Ioannis Matthaiakakis, Ioannis	•DD 11.4 •MP 7.3 T 23.1 •GR 10.5 2, •ST 8.3, 7, EP 12.4, T 56.5 T 39.5 T 102.5, •T 3.5 •DD 35.12 •T 40.8 1.8, T 62.9 •T 40.8 1.8, T 62.9 •T 40.8 (GR 14.2 T 49.6 6.1, T 66.5 T 26.5 T 26.5 T 52.1 •T 103.8 6.3, T 36.5 4.2, T 34.3, 2, T 97.5 3, •T 82.3, MP 4.2 T 49.2
Mandl, Angelika Mandrysch, Jan Manrechg, Werner Manfred, Geilhaupt Manke, Henning ST 8.4 Mann, Gottfried T 83.7 Mannel, Thomas Mans, Leon Mantegazzini, Federica T 102.6 Manthei, Alina Marakis, Vassilios Marin Tobon, Cesar Augusto Märker, Max Markoos, Ahmed Marcava, Hanna Marrodán Undagoitia, Teresa Marsh, Allison Marsteller, Alexander Marsteller, Alexander Martin, Till Martino, Tommy Martius, Florin Martin, Johann Marz, Martin Marto, Jast,	•DD 11.4 •MP 7.3 T 23.1 T 23.1 •GR 10.5 2, •ST 8.3, 7, EP 12.4, T 56.5 T 39.5 T 102.5, •T 3.5 •DD 35.12 •T 40.8 1.8, T 62.9 •T 61.3 •GR 14.2 T 78.2 •PV IV T 49.6 6.1, T 66.5 T 2.6 •T 95.5 T 52.1 •T 103.8 6.3, T 36.5 6.3, T 36.5 3, •T 82.3, •DD 18.2 •MP 4.2 • T 49.2
Mandl, Angelika Mandrysch, Jan Mansechg, Werner Manfred, Geilhaupt Manke, Henning ST 8.4 Mann, Gottfried T 83.7 Mannel, Thomas Mans, Leon Mantegazzini, Federica T 102.6 Manthei, Alina Marakis, Vassilios Marin Tobon, Cesar Augusto Märker, Max Markoos, Ahmed Markoos, Ahmed Martoda Undagoitia, Teresa Marsh, Allison Marsteller, Alexander Martens, Stephan Martens, Stephan Martinov, Tommy Martus, Florin Martun, Johann Marto, Johann Marto, Ta 53., T 41.1, T 59. Masttolorenzo, Luca T 26. T 89.1, T 89.2, T 90.6 Matthia, Kloeft	•DD 11.4 •MP 7.3 T 23.1 •GR 10.5 2, •ST 8.3, 7, EP 12.4, T 56.5 T 39.5 T 102.5, •T 3.5 •DD 35.12 •T 40.8 1.8, T 62.9 •T 40.8 6.1, T 66.3 •GR 14.2 •GR 14.2 •F 19.5 T 2.6 T 2.6 6.1, T 66.5 T 2.5 T 2.5 T 103.8 6.3, T 36.5 4.2, T 34.3, 2, T 97.5 3, •T 82.3, •DT 18.2 •T 49.2
Mandl, Angelika Mandrysch, Jan Maneschg, Werner Manfred, Geilhaupt Manke, Henning ST 8.4 Mann, Gottfried T 83.7 Mannel, Thomas Mans, Leon Mantejazzini, Federica T 102.6 Manthei, Alina Marakis, Vassilios Marin Tobon, Cesar Augusto Märker, Max Markoos, Ahmed Marozava, Hanna Marrodán Undagoitia, Teresa Marsh, Allison Marsteller, Alexander Marst, Allison Marsteller, Alexander Martins, Stephan Martins, Florin Martins, Florin Martins, Florin Martin, Jill Martinov, Tommy Martius, Florin Marto, Jastrolorenzo, Luca T 26. T 89.1, T 89.2, T 90.6 Matthaia Akoft	•DD 11.4 •MP 7.3 T 23.1 •GR 10.5 2, •ST 8.3, 7, EP 12.4, T 56.5 T 39.5 T 102.5, •T 3.5 •DD 35.12 •T 40.8 1.8, T 62.9 •T 61.3 •GR 14.2 T 78.2 •PV IV T 49.6 6.1, T 66.5 T 2.6 •T 95.5 T 52.1 •T 103.8 6.3, T 36.5 1.2, T 34.3, 2, T 97.5 3, •T 82.3, •DD 18.2 •T 49.2 •T 49.2
Mandl, Angelika Mandrysch, Jan Mansechg, Werner Manfred, Geilhaupt Manfred, Geilhaupt Manke, Henning ST 8.4 Mann, Gottfried T 83.7 Mannel, Thomas Mans, Leon Mantejazzini, Federica T 102.6 Manthei, Alina Marakis, Vassilios Marin Tobon, Cesar Augusto Märker, Max Marakos, Ahmed Markos, Ahmed Martozava, Hanna Marrodán Undagoitia, Teresa Marsh, Allison Marteller, Alexander Martens, Stephan Marteller, Alexander Martini, Till Martinov, Tommy Martius, Florin Martino, Tommy Martis, Florin Marty, Johann Marz 16, 9, 734 T 35.2, T 35.3, T 41.1, T 59. Mastolorenzo, Luca T 26 Matthias, Hoeft Matthias, Hoeft Matthias, Hoeft	•DD 11.4 •MP 7.3 T 23.1 •GR 10.5 2, •ST 8.3, 7, EP 12.4, T 56.5 T 39.5 T 102.5, •T 35. •DD 35.12 •T 40.8 1.8, T 62.9 •T 61.3 •GR 14.2 •GR 14.2 •F 61.3 •GR 14.2 T 49.6 6.1, T 66.5 T 2.6 •T 95.5 T 52.1 •T 103.8 6.3, T 36.5 4.2, T 34.3, 3, •T 82.3, •DD 18.2 •MP 4.2 •T 6.3 •T 6.3 •T 6.3 •T 95.5 T 52.1 •T 103.8 6.3, T 36.5 4.2, T 34.3, 3, •T 82.3, •DD 18.2 •T 6.3 •T 6.3
Mandl, Angelika Mandrysch, Jan Maneschg, Werner Manfred, Geilhaupt Manke, Henning ST 8.4 Mann, Gottfried T 83.7 Mannel, Thomas Mans, Leon Mantejazzini, Federica T 102.6 Manthei, Alina Marakis, Vassilios Marin Tobon, Cesar Augusto Märker, Max Markoos, Ahmed Marcaya, Hanna Martodán Undagoitia, Teresa Marsh, Allison Marsteller, Alexander Martens, Stephan Martius, Florin Martius, Florin Martius, Florin Martius, Florin Marti, Jill Martinov, Tommy Martius, Florin Marto, Lucia Marto, Lucia Marto, Lucia Mastolorenzo, Luca T 26. Matthaiakakis, Ioannis Matthe, Christian Matthias, Hoeft Mattini, Alessandro	•DD 11.4 •MP 7.3 T 23.1 •GR 10.5 2, •ST 8.3, 7, EP 12.4, T 56.5 T 39.5 T 102.5, •T 3.5 •DD 35.12 •T 40.8 1.8, T 62.9 •T 61.3 •GR 14.2 T 40.8 1.8, T 62.9 •T 40.8 6.1, T 66.5 T 2.6 T 2.6 T 103.8 6.3, T 36.5 4.2, T 34.3, 2, T 97.5 3, •T 82.3, •DD 18.2 •MP 4.2 •T 6.3 EP 3.1
Mandl, Angelika Mandrysch, Jan Mansechg, Werner Manfred, Geilhaupt Manke, Henning ST 8.4 Mann, Gottfried T 83.7 Mannel, Thomas Mans, Leon Mantejazzini, Federica T 102.6 Manthei, Alina Marakis, Vassilios Marin Tobon, Cesar Augusto Märker, Max Markoos, Ahmed Markoos, Ahmed Marozava, Hanna Marrodán Undagoitia, Teresa Marsh, Alison Marsteller, Alexander Martens, Stephan Martens, Stephan Martini, Till Martinov, Tommy Martius, Florin Marty, Johann Marto, Jasz, T 45, 37 Masteti, Lucia T 16, 9, 73 T 35.2, T 35.3, T 41.1, T 59. Mastolorenzo, Luca T 26 Matthias, Hoeft Matthias, Hoeft Matthias, Hoeft Matthies, Christopher Maturi, Alessandro Maturi, Alessandro Maturi, Alessandro	•DD 11.4 •MP 7.3 T 23.1 •GR 10.5 2, •ST 8.3, 7, EP 12.4, T 56.5 T 39.5 T 102.5, •T 3.5 •DD 35.12 •T 40.8 1.8, T 62.9 •T 61.3 •GR 14.2 •F 61.3 •GR 14.2 T 78.2 •PV IV T 49.6 6.1, T 66.5 T 25.1 •T 103.8 6.3, T 36.5 4.2, T 34.3, 2, T 97.5 3, •T 82.3, MP 4.2 •T 49.2 •F 9.31 •EP 3.1 •EP 6.1
Mandl, Angelika Mandrysch, Jan Maneschg, Werner Manke, Henning ST 8.4 Mann, Gottfried T 83.7 Mannel, Thomas Mans, Leon Mantegazzini, Federica T 102.6 Manthei, Alina Marategazzini, Federica T 102.6 Manthei, Alina Marakis, Vassilios Marin Tobon, Cesar Augusto Märker, Max Marin Cohon, Cesar Augusto Märker, Max Markhoos, Ahmed Marozava, Hanna Marodán Undagoitia, Teresa Marsh, Allison Marsteller, Alexander Martens, Stephan Martius, Florin Martius, Florin Martius, Florin Martius, Florin Marti, Jill Martinov, Tommy Marti, Jucia Marto, Lucia T 16.9, T 3/ T 35.2, T 35.3, T 41.1, T 59. Mastrolorenzo, Luca T 26. Matejak Cvenic, Karolina Matthias, Hoeft Matthias, Hoeft Mattinik, Alesandro Maturilli, Alessandro Mäusle, Raquel Mausolf, Florian	•DD 11.4 •MP 7.3 T 23.1 •GR 10.5 2, •ST 8.3, 7, EP 12.4, T 56.5 T 39.5 T 102.5, •T 3.5 •DD 35.12 •T 102.5, •T 3.5 •DD 35.12 •T 40.8 1.8, T 62.9 •T 61.3 •GR 14.2 T 78.2 T 49.6 6.1, T 66.5 T 2.6 T 29.5 T 52.1 •T 103.8 6.3, T 36.5 4.2, T 34.3, 2, T 97.5 3, •T 82.3, •DD 18.2 •T 49.2 EP 3.1 EP 3.1 EP 6.1 •T 84.5
Mandl, Angelika Mandrysch, Jan Mansechg, Werner Manfred, Geilhaupt Manke, Henning ST 8.4 Mann, Gottfried T 83.7 Mannel, Thomas Mans, Leon Mantejazzini, Federica T 102.6 Manthei, Alina Marakis, Vassilios Marin Tobon, Cesar Augusto Märker, Max Markoos, Ahmed Markoos, Ahmed Marcava, Hanna Marrodán Undagoitia, Teresa Marsh, Allison Marsteller, Alexander Martens, Stephan Marteller, Alexander Martini, Till Martinov, Tommy Martinov, Tommy Martus, Florin Martyn, Johann Marz, Martin Martolorenzo, Luca T 36.2, T 35.3, T 41.1, T 59. Masteller, Christian Matthias, Hoeft Matthias, Hoeft Matthias, Hoeft Matthias, Hoeft Matthias, Raquel Mausolf, Florian Maver Matthias	•DD 11.4 •MP 7.3 T 23.1 •GR 10.5 2, •ST 8.3, 7, EP 12.4, T 56.5 T 39.5 T 102.5, •T 3.5 •DD 35.12 •T 40.8 1.8, T 62.9 •T 61.3 •GR 14.2 •F 61.3 •GR 14.2 •PV IV T 49.6 6.1, T 66.5 T 2.6 •T 95.5 T 52.1 •T 103.8 6.3, T 36.5 4.2, T 34.3, 2, T 97.5 3, •T 82.3, •DD 18.2 •T 49.2 •F 9.31 •EP 6.1 •EP 6.1 •EP 6.1 •EP 6.1
Mandl, Angelika Mandrysch, Jan Mansechg, Werner Manfred, Geilhaupt Manke, Henning ST 8.4 Mann, Gottfried T 83.7 Mannel, Thomas Mans, Leon Mantegazzini, Federica T 102.6 Manthei, Alina Marakis, Vassilios Marin Tobon, Cesar Augusto Märker, Max Markhoos, Ahmed Markoos, Ahmed Marozava, Hanna Marcodán Undagoitia, Teresa Marsh, Allison Martens, Stephan Martens, Stephan Martinov, Tommy Martius, Florin Martin, Johann Marto, Johann Marto, Johann Marto, Johann Marto, Johann Marto, Johann Mart, Mattin Martino, Taj 5.2, T 35.3, T 41.1, T 59. Mastrolorenzo, Luca T 89.1, T 89.2, T 90.6 Matthaiakakis, Ioannis Matthias, Hoeft Matthias, Hoeft Matthias, Christopher Maturilli, Alessandro Mausolf, Florian Mausolf, Florian Masu, Matthias Mastel, Lucia Matthias, Raquel Mausolf, Florian Mayer, Matthias Mayer, Matthias	•DD 11.4 •MP 7.3 T 23.1 •GR 10.5 2, •ST 8.3, 7, EP 12.4, T 56.5 T 39.5 T 102.5, •T 3.5 •DD 35.12 •T 40.8 1.8, T 62.9 •T 61.3 •GR 14.2 T 40.8 6.1, T 66.5 T 2.6 T 2.6 T 95.5 T 52.1 •T 103.8 6.3, T 36.5 4.2, T 34.3 2, T 97.5 3, •T 82.3, •DT 18.2 •P 3.1 •EP 3.1 •EP 3.1 •EP 6.1 •T 34.5 0.8, T 99.5
Mandl, Angelika Mandrysch, Jan Mansechg, Werner Manfred, Geilhaupt Manke, Henning ST 8.4 Mann, Gottfried T 83.7 Mannel, Thomas Mans, Leon Mantegazzini, Federica T 102.6 Manthei, Alina Marakis, Vassilios Marin Tobon, Cesar Augusto Märker, Max Markoos, Ahmed Markoos, Ahmed Markoos, Ahmed Marcava, Hanna Marrodán Undagoitia, Teresa Marsh, Allison Marsteller, Alexander Martens, Stephan Martinov, Tommy Martiuov, Tommy Martiuo, Florin Martol, Lucia T 16,9, T 3 Masetti, Lucia T 35.2, T 35.3, T 41.1, T 59. Mastholorenzo, Luca T 89.1, T 89.2, T 90.6 Mattinas, Hoeft Matthias, Hoeft Matthias, Hoeft Matthias, Hoeft Matthias, Hoeft Matthias, Hoeft Matthias, Hoeft Matthias, Hoeft Matthias, Hoeft Matthias, Christopher Maturi, Simon Mayer, Matthias Mayer, Matthias Mayer, Matthias Mayer, Matthias Mayer, Matthias Mayer, Simon	•DD 11.4 •MP 7.3 T 23.1 •GR 10.5 2, •ST 8.3, 7, EP 12.4, T 56.5 T 39.5 T 102.5, •T 3.5 •DD 35.12 •T 40.8 1.8, T 62.9 •T 61.3 •GR 14.2 •PV IV T 49.6 6.1, T 66.5 T 2.6 •T 95.5 T 52.1 •T 103.8 6.3, T 36.5 4.2, T 34.3, 2, T 97.5 3, •T 82.3, •DD 18.2 •PD 18.2 •PC 14.9 •F 9.5 F 9.3 •EP 3.1 •EP 6.1 •T 99.5 •MP 4.3
Mandl, Angelika Mandrysch, Jan Mansechg, Werner Manfred, Geilhaupt Manke, Henning ST 8.4 Mann, Gottfried T 83.7 Mannel, Thomas Mans, Leon Mantegazzini, Federica T 102.6 Manthei, Alina Marakis, Vassilios Marin Tobon, Cesar Augusto Märker, Max Markhoos, Ahmed Markoos, Ahmed Marozava, Hanna Marcodan Undagoitia, Teresa Marsh, Allison Martens, Stephan Martens, Stephan Martinov, Tommy Martius, Florin Martin, Johann Marz, Martin Martino, Tonmy Martius, Florin Marty, Johann Marz, Martin Mart, Jas2, T 35.3, T 41.1, T 59. Masteliorenzo, Luca T 35.2, T 35.3, T 41.1, T 59. Masthorenzo, Luca T 89.1, T 89.2, T 90.6 Matthaiakakis, Ioannis Matthias, Hoeft Matthias, Hoeft Matthias, Hoeft Matthias, Hoeft Matthias, Christopher Maturilli, Alessandro Mäusle, Raquel Mausolf, Florian Mayer, Simon	•DD 11.4 •MP 7.3 T 23.1 •GR 10.5 2, •ST 8.3, 7, EP 12.4, T 56.5 T 39.5 T 102.5, •T 3.5 •DD 35.12 •T 3.5 •DD 35.12 •T 40.8 1.8, T 62.9 •T 61.3 •GR 14.2 T 40.8 6.1, T 66.5 T 2.6 T 2.6 6.1, T 66.5 T 2.5 T 52.1 •T 103.8 6.3, T 36.5 4.2, T 34.3 2, T 97.5 3, •T 82.3, •DD 18.2 •P 31. •EP 3.1 •EP 6.1 • T 34.5 0.8, T 99.5 MP 4.2 • T 9.5 • T 9.5 • T 3.5 • MP 4.2 • EP 3.1 • EP 6.1 • T 34.5 0.8, T 99.5 MP 4.2 • MP 4.3
Mandl, Angelika Mandrysch, Jan Mansechg, Werner Manfred, Geilhaupt Manke, Henning ST 8.4 Mann, Gottfried T 83.7 Mannel, Thomas Mans, Leon Mantegazzini, Federica T 102.6 Manthei, Alina Marakis, Vassilios Marin Tobon, Cesar Augusto Märker, Max Markoos, Ahmed Markoos, Ahmed Marcozava, Hanna Marrodán Undagoitia, Teresa Marsh, Allison Marsteller, Alexander Marsh, Allison Marsteller, Alexander Martin, Till Martino, Tommy Martius, Florin Martin, Stephan Marz, Martin Martio, T 169, T 32 T 35.2, T 35.3, T 41.1, T 59. Masthela, Christian Matthias, Hoeft Matthias, Hoeft Matthias, Hoeft Matthias, Christopher Matthias, Hoeft Matthias, Florin Matthias, Hoeft Matthias, Christopher Matthias, Hoeft Matthias, Florin Matthias, Hoeft Matthias, Christopher Matthias, Florin Matthias, Stopher Matthias, Signon Mayer, Simon Mayer, Simon Mather T 200000000000000000000000000000000000	•DD 11.4 •MP 7.3 T 23.1 •GR 10.5 2, •ST 8.3, 7, EP 12.4, T 56.5 T 39.5 T 102.5, •T 35. •DD 35.12 •T 40.8 1.8, T 62.9 •T 61.3 •GR 14.2 T 78.2 T 79.5 T 78.2 T 79.5 T 78.2 T 79.5 T 79.5
Mandl, Angelika Mandrysch, Jan Mansechg, Werner Manfred, Geilhaupt Manke, Henning ST 8.4 Mann, Gottfried T 83.7 Mannel, Thomas Mans, Leon Mantegazzini, Federica T 102.6 Manthei, Alina Marakis, Vassilios Marin Tobon, Cesar Augusto Märker, Max Markoos, Ahmed Markoos, Ahmed Martoda, Undagoitia, Teresa Marsh, Allison Marsteller, Alexander Martens, Stephan Marteler, Alexander Martini, Till Martinov, Tommy Martius, Florin Martu, Johann Marz, Martin Marto, Ta 53., T 41.1, T 59. Mastelier, Kasander Matthias, Hoeft Matthias, Hoeft Matthias, Hoeft Matthias, Hoeft Matthias, Hoeft Matthias, Christopher Matusi, Florian Matthias, Hoeft Matthias, Hoeft Matthias, Hoeft Matthias, Florian Matthias, Hoeft Matthias, Florian Matthias, Hoeft Matthias, Florian Matthias, Christopher Matthias, Florian Matthias, Florian Matthias, Florian Matthias, Hoeft Matthias, Florian Matthias, Florian Mayer, Matthias Florian Mayer, Simon KcKeown, Peter Klachlan, Thomas	•DD 11.4 •MP 7.3 T 23.1 •GR 10.5 2, •ST 8.3, 7, EP 12.4, T 56.5 T 39.5 T 102.5, •T 3.5 •DD 35.12 •T 40.8 1.8, T 62.9 •T 40.8 1.8, T 62.9 •GR 14.2 •GR 14.2 •F 16.3 •GR 14.2 •GR 14.2 •F 49.2 T 49.6 6.1, T 66.5 T 2.6 •T 95.5 T 52.1 •T 103.8 6.3, T 36.5 4.2, T 34.3, 2, T 97.5 .3, •T 82.3, •DT 18.2 •F 49.2 •F 6.1 •EP 3.1 •EP 6.1 •T 34.5 0.8, T 99.5 MP 4.2 •T 6.6
Mandl, Angelika Mandrysch, Jan Mansechg, Werner Manfred, Geilhaupt Manfred, Geilhaupt Manr, Gottfried T 83.7 Mannel, Thomas Mans, Leon Mantegazzini, Federica T 102.6 Manthei, Alina Marakis, Vassilios Marin Tobon, Cesar Augusto Märker, Max Markoos, Ahmed Marozava, Hanna Marrodán Undagoitia, Teresa Marsh, Allison Marsteller, Alexander Marst, Allison Marsteller, Alexander Martens, Stephan Martin, Till Martinov, Tommy Martius, Florin Martio, Johann Marto, Johann Marto, Johann Marto, Johann Marto, Jorenzo, Luca T 35.2, T 35.3, T 41.1, T 59. Masterlorenzo, Luca T 89.1, T 89.2, T 90.6 Mattihas, Hoeft Matthias, Hoeft Matthias, Christopher Matthias, Aloft Matthe, Christian Matthe, Christian Matthe, Christian Matthias, Hoeft Matuili, Alessandro Matthias, Raquel Mausolf, Florian Mayer, Matthias T 2000 Mayer, Simon Marte, Klaus	•DD 11.4 •MP 7.3 T 23.1 •GR 10.5 2, •ST 8.3, 7, EP 12.4, T 56.5 T 39.5 T 102.5, •T 35. •DD 35.12 •T 40.8 1.8, T 62.9 •T 61.3 •GR 14.2 T 78.2 •PV IV T 49.6 6.1, T 66.5 T 2.6 •T 95.5 T 52.1 •T 103.8 6.3, T 36.5 1.2, T 34.3, 2, T 97.5 3, •T 82.3, •DD 18.2 •MP 4.3 AKPIK 2.4 •T 6.5 F 7 49.2 •F 9.5 •F 9.5 •F 6.5 •F 9.5 •F 149.2 •F 6.5 •F 9.5 •F 149.2 •F 7.5 •F 149.2 •F 7.5 •F 7.
Mandl, Angelika Mandrysch, Jan Mansechg, Werner Manfred, Geilhaupt Manke, Henning ST 8.4 Mann, Gottfried T 83.7 Mannel, Thomas Mans, Leon Mantejazzini, Federica T 102.6 Manthei, Alina Marakis, Vassilios Marin Tobon, Cesar Augusto Märker, Max Markos, Ahmed Markos, Ahmed Markos, Ahmed Marozava, Hanna Marrodán Undagoitia, Teresa Marsh, Allison Marteller, Alexander Martens, Stephan Marteller, Alexander Martinov, Tommy Martius, Florin Marty, Johann Marto, Johann Marto, Jastro, T 26, T 35.2, T 35.3, T 41.1, T 59. Mastrolorenzo, Luca T 26, Matthias, Hoeft Matthias, Hoeft Matthias, Hoeft Matthias, Hoeft Matthias, Christopher Matus, Florian Matthe, Christian Matthe, Christian Matthias, Hoeft Matthias, Hoeft Matthias, Hoeft Matthias, Hoeft Matthias, Christopher Maturilli, Alessandro Maty, Simon Maty, Simon Mayer, Matthias Mathe, Christian Maty, Simon Maty, Simon Mayer, Mathias Mathas, Kaus Mayer, Simon McKeown, Peter Mausolf, Florian Mayer, Simon McKeown, Peter	•DD 11.4 •MP 7.3 T 23.1 •GR 10.5 2, •ST 8.3, 7, EP 12.4, T 56.5 T 39.5 T 102.5, •T 3.5 •DD 35.12 •T 40.8 1.8, T 62.9 •T 61.3 •GR 14.2 •T 40.8 6.1, T 66.5 T 2.6 T 2.6 T 2.6 T 2.5 T 52.1 •T 103.8 6.3, T 36.5 4.2, T 34.3, 2, T 97.5 3, •T 82.3, •EP 6.1 •EP 6.1 •EP 6.1 •T 34.5 0.8, T 99.5 3. ~MP 4.3 AKPIK 2.4 •T 6.6 •T 6.6 •T 6.6 •T 6.1 •EP 6.1 •EP 6.1 •T 6.6 •T 6.6 •T 6.6 •T 6.6 •T 6.6 •T 6.6 •T 6.6
Mandl, Angelika Mandrysch, Jan Maneschg, Werner Manke, Henning ST 8.4 Mann, Gottfried T 83.7 Mannel, Thomas Mans, Leon Mantegazzini, Federica T 102.6 Manthei, Alina Marakis, Vassilios Marin Tobon, Cesar Augusto Marker, Max Markos, Ahmed Marozava, Hanna Marodán Undagoitia, Teresa Marsh, Allison Marsteller, Alexander Marsteller, Alexander Martens, Stephan Martins, Florin Martius, Florin Martius, Florin Martius, Florin Martio, Jastrolorenzo, Luca T 16.9, T 32 T 35.2, T 35.3, T 41.1, T 59. Mastrolorenzo, Luca T 26. Mattheia, Aloeft Matthias, Hoeft Matthias, Hoeft Matthias, Christopher Matthe, Christian Matthe, Christian Matthe, Christian Matthe, Christian Matthe, Christian Matthias, Hoeft Mausel, Raquel Mausolf, Florian Mayer, Matthias Mathas Masetti Alessandro Mithe, Christian Matthias, Moeft Mausolf, Florian Mayer, Matthias Mathas Mayer, Simon Mayer, Simon Macke, Klaus Mecke, Klaus	•DD 11.4 •MP 7.3 T 23.1 •GR 10.5 2, •ST 8.3, 7, EP 12.4, T 56.5 T 39.5 T 102.5, •T 3.5 •DD 35.12 •T 40.8 1.8, T 62.9 •T 40.8 1.8, T 62.9 •T 40.8 (T 49.6 6.1, T 66.5 T 2.6 T 95.5 T 52.1 •T 103.8 6.3, T 36.5 4.2, T 34.3 2, T 97.5 3, •T 82.3, EP 3.1 •EP 6.1 •T 43.8 ON H 4.3 AKENE 2.4 •T 6.6 •MP 10.1 •T 43.8 •T 43.8
Mandl, Angelika Mandrysch, Jan Mansechg, Werner Manfred, Geilhaupt Manke, Henning ST 8.4 Mann, Gottfried T 83.7 Mannel, Thomas Mans, Leon Mantegazzini, Federica T 102.6 Manthei, Alina Marakis, Vassilios Marin Tobon, Cesar Augusto Märker, Max Markoos, Ahmed Markoos, Ahmed Markoos, Ahmed Markoos, Ahmed Martodán Undagoitia, Teresa Marsh, Allison Marsteller, Alexander Martens, Stephan Martinov, Tommy Martiuov, Tommy Martiuo, Florin Martolorenzo, Luca T 35.2, T 35.3, T 41.1, T 59. Mastholorenzo, Luca T 89.1, T 89.2, T 90.6 Mattinia, Hoeft Matthias, Christopher Matthias, Hoeft Matthias, Matthias Matthias Matthias Matthias Matthias Matthias Matthias Mayer, Matthias Mayer, Matthias Meggendorfer, Felix Meggendorfer, Felix	•DD 11.4 •MP 7.3 T 23.1 •GR 10.5 2, •ST 8.3, 7, EP 12.4, T 56.5 T 39.5 T 102.5, •T 3.5 •DD 35.12 •T 40.8 1.8, T 62.9 •T 61.3 •GR 14.2 •F 61.3 •GR 14.2 •PV IV T 49.6 6.1, T 66.5 •T 95.5 T 52.1 •T 103.8 6.3, T 36.5 4.2, T 34.3, 2, T 97.5 3, •T 82.3, •DD 18.2 •PD 18.2 •F 9.31 •EP 6.1 •T 6.6 MP 4.2 •F 6.6 •MP 4.3 AKPIK 2.4 •T 6.6 •MP 10.1 •T 43.8 •T 30.5
Mandl, Angelika Mandrysch, Jan Mansechg, Werner Manfred, Geilhaupt Manfred, Geilhaupt Manke, Henning ST 8.4 Mann, Gottfried T 83.7 Mannel, Thomas Mans, Leon Mantejazzini, Federica T 102.6 Manthei, Alina Marakis, Vassilios Marin Tobon, Cesar Augusto Märker, Max Markoos, Ahmed Markoos, Ahmed Martoda, Undagoitia, Teresa Marsh, Allison Marsteller, Alexander Martens, Stephan Marteller, Alexander Martino, Tommy Martius, Florin Martin, Johann Mart, Johann Marz, Martin Mattino, Tas2, T 35.3, T 41.1, T 59. Mastelier, Karsolina Matthias, Hoeft Matthias, Hoeft Matthias, Hoeft Matthias, Hoeft Matthias, Hoeft Matthias, Hoeft Matthias, Hoeft Matthias, Hoeft Matthias, Florian Matthias, Hoeft Matthias, Hoeft Matthias, Hoeft Matthias, Hoeft Matthias, Florian Matthias, Hoeft Matthias, Hoeft Matthias, Hoeft Matthias, Hoeft Matthias, Hoeft Matthias, Florian Matthias, Hoeft Matthias, Florian Matthias, Hoeft Matthias, Hoeft Mausolf, Florian Mayer, Matthias T 20 Mayer, Simon Metka, Klaus Meggendorfter, Felix Meggendorfter, Felix Mehta, Ankita T 10	•DD 11.4 •MP 7.3 T 23.1 •GR 10.5 2, •ST 8.3, 7, EP 12.4, T 56.5 T 39.5 T 102.5, •T 3.5 •DD 35.12 •T 40.8 1.8, T 62.9 •T 40.8 1.8, T 62.9 •T 40.8 6.1, T 66.5 T 2.6 •T 95.5 T 2.2 •F 103.8 6.3, T 36.5 4.2, T 34.3, 2, T 97.5 .3, •T 82.3, •EP 3.1 •EP 3.1 •EP 6.1 •T 34.5 0.8, T 99.5 MP 4.2 •T 6.6 •T 6.6 •T 6.5 •T 34.5 0.8, T 99.5 MP 4.2 •T 6.3 •EP 3.1 •EP 6.1 •T 34.5 0.8, T 99.5 MP 4.3 AKPIK 2.4 •T 6.6 •T 30.5 0.1, •T 87.1
Mandl, Angelika Mandrysch, Jan Manschg, Werner Manfred, Geilhaupt Mantred, Geilhaupt Manke, Henning ST 8.4 Mann, Gottfried T 83.7 Mannel, Thomas Mans, Leon Mantegazzini, Federica T 102.6 Manthei, Alina Marakis, Vassilios Marin Tobon, Cesar Augusto Märker, Max Markoos, Akmed Marozava, Hanna Marrodán Undagoitia, Teresa Marsh, Allison Marsteller, Alexander Marsh, Allison Marsteller, Alexander Martodán Undagoitia, Teresa Marsh, Allison Marsteller, Alexander Martodán Undagoitia, Teresa Marsh, Allison Marsteller, Alexander Martodán Undagoitia, Teresa Marsh, Allison Marsteller, Alexander Martin, Till Martinov, Tommy Martius, Florin Martyn, Johann Marz, Martin Marto, Tonmy Martis, Stephan Marto, T 169, T 32 T 35.2, T 35.3, T 41.1, T 59. Mastrolorenzo, Luca T 89.1, T 89.2, T 90.6 Matejak Cvenic, Karolina Matthias, Hoeft Mauthias, Hoeft Mausolf, Florian Matthias, Hoeft Mausolf, Florian Mayer, Matthias Florian Mayer, Simon Mayer, Simon Mayer, Simon Mayer, Simon Mayer, Kalus Meggendorfer, Felix Meggendorfer, Felix Medgendorfer, Felix Meing Gervin	•DD 11.4 •MP 7.3 T 23.1 •GR 10.5 2, •ST 8.3, 7, EP 12.4, T 56.5 T 39.5 T 102.5, •T 35. •DD 35.12 •T 40.8 1.8, T 62.9 •T 61.3 •GR 14.2 T 78.2 T 77.5 T 78.2 T 78.2 T 78.2 T 78.2 T 79.5 T 78.2 T 78.2 T 77.5 T 78.2 T 78.2 T 78.2 T 79.5 T 78.2 T 78.2 T 77.5 T 78.2 T 78.2 T 78.2 T 78.2 T 77.5 T 78.2 T 77.5 T 78.2 T 77.5 T 78.2 T 77.5 T 78.2 T 79.5 T 78.2 T 79.5 T 78.2 T 77.5 T 77.5

T 77.7 Meinel, Stefan
Melo, Andrés • T 89.7 Meloni, Federico T 79.8, T 79.9 Melzer-Pellmann, Isabell T 9.5, T 9.6, T 60.1 T 9.5, T 9.6, Mender, Simone • T 40.4 Mendizabal, Mikel • T 32.4 Menen, Marco • T 92.4
T 60.1 Mender, Simone
Mendizabal, Mikel•T 32.4 Menen. Marco•T 92.6
Menon, Tushar•AGPhil 7.1
Mentzel, Florian ST 1.1, ST 3.3, •ST 3.4, ST 4.4, ST 7.2, ST 7.4, T 53.3, T 107.8
Mercier, Michelle
Mertens, SusanneT 21.7, T 23.3, T 75.2 Mertsch, Philipp EP 6.6, T 46.1,
T 72.4, T 73.5, T 73.6 Messner, Jonathan
Metzner, Felix
Meuser, Danilo
Meyer, René .•MP 1.2, MP 4.1, MP 4.2 Meyer zu Theenhausen, Hanno •T 69.2
Meyers, Mark •EP 7.1, T 72.9 Meyers, Zachary
Mihai, Iulia ••GP 9.3 Mihaylov, Deyan ••GP 9.2 Mikaylov, Deyan •••••GR 9.2
Miller, Ryan
Minh Phan, Vo-Hong T 73.5 Mitchell, Alison •EP 10.2, T 17.2, T 45.4, •T 81.2
Mitzel, Dominik Stefan•T 82.2 Mkrtchyan, Artak
Mocellin, Giovanni
Mödden, Antje T 60.6, T 60.7 Moder, Paul
Mohammed, Ashraf
Möhring, Lars •DD 29.2, DD 43.2, DD 46.1, DD 48.3
Molifinaliti, Lars AKPIK 1.2 Moll, MichaelST 9.3 Mondal, Buddhadeb•T 33.7
Mondal, Spadan T 26.1, •T 26.3, T 89.2, T 90.6
Mondragon, Elizabeth T 102.7 Monteil, Stéphane T 34.7 Montero, Raúl ST 9.3
MONUMENT-Kollaboration T 102.7 Moortgat-Pick, Gudrid MP 12.1, MP 12.2, T 9.1, T 31.4, T 36.7, T 36.8,
T 36.9, T 38.1, T 38.7, T 44.5, T 44.6, T 63.6 Moos, Lars-Thorben
Moralejo, Abelardo
Morejon, Leonel
Morgenthaler, Maurice•T 63.8 Morgenthaler, Maurice•T 63.8 Moser, Hans-Günther
Moskalets, Tetiana 1 62.4, 1 62.5 Mossel, Kyra T 23.6 Moureaux, Louis T 93.2
Moustakas, Konstantinos T 40.8
Moyano, Helena
Moyano, Helena 18.5 Mozota Frauca, Alvaro • AGPhil 11.1 Mrowietz, Malte • T 9.7 Mrozik, Konstantin T 106.1, T 106.2, T 106.3 • T 9.7

Mück, Alexander
Muddu Venkata Sai Saketh GR 5.4
Muddu, Venkata Gar Gaketh . • On 0.4
Mueller, Andreas
Mueller, Stefan E T 96.4
Mühlleitner, Margarete T 85.3, T 92.2.
т 02 3
1 92.3
Mujagic, Jasmina•DD 39.3
Mukherjee, Swagata•T 1.3, T 64.2,
T 64 3
Multherine Tiete ED 11 1 T 02 1
Mukherjee, fista •EP II.I, •I 83.I
Mulder, Mick
Mulders Martiin T 7 5
Müller Andrees DD 10.2 DD 26.1
Mullel, AlluleasDD 10.2, DD 30.1
Muller, Axel 1 20.3, 1 20.9, 1 48.6,
T 48.8, T 102.3
Müller Fahienne DD 311
Müller Jakob
Muller, Jakob
Müller, Jonas
Müller Julia •T 50 2
Müllor Martin T 14.9
Muller, Nils EP 3.1
Müller, Rainer DD 6.2, DD 13.2,
14 1 0 33 2
Muller, Sala•51 1.5
Müller, Saskia . ST 8.2, ST 8.3, ST 8.4
Müller, Thomas
Τ 65 7 Τ 87 4 Τ 87 5 Τ 04 4
Müller Welf Obsistister ED (1 ED (2
wuller, woll-Christian . EP 6.1, EP 6.2
Müller-Gosewisch, Jan-Ole T 65.6,
T 65.7. T 94.4
Multi-wavelength collaborators and the
water-wavelength collaborators and the
MAGIC and Fermi-LAT-Kollaboration
EP 11.2
Mundlens MP72
Munz, vera DD 12.2, DD 12.4,
•DD 12.5, DD 12.6
Muthukrishnan Siddharth
Muthusamy, GowthamDD 6.2,
DD 38.1
NA61/SHINE-Kollaboration T 101.9
Nach Dishard T10.0 T46.0
Naab, Richard 1 18.8, •1 40.2
Naaijkens, Pieter
Naber, Jonathan
Nachman Benjamin AKPIK / 2
Nachtmann, Utto 1 85.8
Nackenhorst, Olaf ST 3.4, T 34.5,
Nackenhorst, Olaf ST 3.4, T 34.5, T 70 7 T 107 8
Nackenhorst, Olaf ST 3.4, T 34.5, T 70.7, T 107.8
Nackenhorst, Olaf ST 3.4, T 34.5, T 70.7, T 107.8 Nadol, Lilli
Nackenhorst, Olaf ST 3.4, T 34.5, T 70.7, T 107.8 Nadol, Lilli
Nackenhorst, Olaf ST 3.4, T 34.5, T 70.7, T 107.8
Nackenhorst, Olaf ST 3.4, T 34.5, T 70.7, T 107.8 •EP 12.10 Nagar, Alessandro GR 9.4 Nagataki, Shigehiro EP 14.1 Näner Paul M •AGPhil 9.2
Nackenhorst, Olaf ST 3.4, T 34.5, T 70.7, T 107.8 Nadol, Lilli •EP 12.10 Nagar, Alessandro GR 9.4 Nagataki, Shigehiro P1.4 Näger, Paul M. •AGPhil 9.2 Democolde wildon T 47.6
Nackenhorst, Olaf ST 3.4, T 34.5, T 70.7, T 107.8 Nadol, Lilli
Nackenhorst, Olaf ST 3.4, T 34.5, T 70.7, T 107.8 Nadol, Lilli
Nackenhorst, Olaf ST 3.4, T 34.5, T 70.7, T 107.8 Nadol, Lilli •EP 12.10 Nagar, Alessandro GR 9.4 Nagataki, Shigehiro EP 14.1 Näger, Paul M. •AGPhil 9.2 namasaka, wilson •T 47.6 Nandi, Gerrit DD 5.2 Naranjo, Pedro AGPhil 1.1
Nackenhorst, Olaf ST 3.4, T 34.5, T 70.7, T 107.8 Nadol, Lilli
Nackenhorst, Olaf ST 3.4, T 34.5, T 70.7, T 107.8 Nadol, Lilli
Nackenhorst, Olaf ST 3.4, T 34.5, T 70.7, T 107.8 Nadol, Lilli •EP 12.10 Nagar, Alessandro GR 9.4 Nagataki, Shigehiro EP 14. Näger, Paul M. •AGPhil 9.2 namasaka, wilson •T 47.6 Naraijo, Pedro AGPhil 19.2 Naranjo, Pedro GP 5.2 Naranjo, Pedro GP 2.3 Nardo, Lorenzo ST 9.2
Nackenhorst, Olaf ST 3.4, T 34.5, T 70.7, T 107.8 Nadol, Lilli EP 12.10 Nagar, Alessandro GR 9.4 Nagataki, Shigehiro EP 1.4 Näger, Paul M. AGPhil 9.2 namasaka, wilson T 47.6 Nandi, Gerrit DD 5.2 Naranjo, Pedro GP 2.3 Nardo, Lorenzo ST 9.2 Narimani Charan, Abtin T 79.4
Nackenhorst, Olaf ST 3.4, T 34.5, T 70.7, T 107.8 Nadol, Lilli •EP 12.10 Nagar, Alessandro GR 9.4 Nagataki, Shigehiro EP 14.10 Näger, Paul M. •AGPhil 9.2 namasaka, wilson •T 47.6 Naranjo, Pedro AGPhil 1.1 Nardi, Roberto GP 2.3 Nardo, Lorenzo ST 9.2 Narimani Charan, Abtin •T 9.4
Nackenhorst, Olaf ST 3.4, T 34.5, T 70.7, T 107.8 Nadol, Lilli •EP 12.10 Nagar, Alessandro GR 9.4 Nagataki, Shigehiro EP 14.10 Näger, Paul M. •AGPhil 9.2 namasaka, wilson •T 47.6 Nardi, Gerrit DD 5.2 Naranjo, Pedro AGPhil 1.1 Nardi, Roberto GP 2.3 Nardo, Lorenzo ST 9.2 Narimani Charan, Abtin •T 79.4 Nasc Esfahani, Alina •T 51.5
Nackenhorst, Olaf ST 3.4, T 34.5, T 70.7, T 107.8 Nadol, Lilli •EP 12.10 Nagar, Alessandro GR 9.4 Nagataki, Shigehiro EP 1.4 Näger, Paul M. •AGPhil 9.2 namasaka, wilson •T 47.6 Naranjo, Pedro AGPhil 1.1 Nardi, Roberto GP 2.3 Nardo, Lorenzo ST 9.2 Narimani Charan, Abtin •T 79.4 Nasr Esfahani, Alina •T 51.5 Nass, Christian •T 62.3
Nackenhorst, Olaf ST 3.4, T 34.5, T 70.7, T 107.8 Nadol, Lilli •EP 12.10 Nagar, Alessandro GR 9.4 Nagataki, Shigehiro EP 14.10 Näger, Paul M. •AGPhil 9.2 namasaka, wilson •T 47.6 Naranjo, Pedro AGPhil 1.1 Nardi, Gerrit D 5.2 Naranjo, Pedro GP 2.3 Nardo, Lorenzo ST 9.2 Narimani Charan, Abtin •T 79.4 Nasz Esfahani, Alina •T 62.3 Nast, Natalie T 107.4
Nackenhorst, Olaf ST 3.4, T 34.5, T 70.7, T 107.8 Nadol, Lilli •EP 12.10 Nagar, Alessandro GR 9.4 Nagataki, Shigehiro EP 1.4 Näger, Paul M. •AGPhil 9.2 namasaka, wilson •T 47.6 Nandi, Gerrit DD 5.2 Naranjo, Pedro AGPhil 1.1 Nardi, Roberto GP 2.3 Nardo, Lorenzo ST 9.2 Narimani Charan, Abtin •T 79.4 Nasr Esfahani, Alina •T 51.5 Nass, Christian •T 62.3 Nast, Natalie T 107.4 Nattland, Phillipp T 33.2
Nackenhorst, Olaf ST 3.4, T 34.5, T 70.7, T 107.8 Nadol, Lilli •EP 12.10 Nagar, Alessandro GR 9.4 Nagataki, Shigehiro EP 14.10 Näger, Paul M. •AGPhil 9.2 namasaka, wilson •T 47.6 Nardi, Gerrit DD 5.2 Naranjo, Pedro AGPhil 1.1 Nardi, Roberto GP 2.3 Nardo, Lorenzo ST 9.2 Narimani Charan, Abtin •T 79.4 Nast, Stafani, Alina T 162.3 Nast, Natalie T 107.4 Nattland, Philipp T 33.2 Navarro, Laume •GP 6.2
Nackenhorst, Olaf ST 3.4, T 34.5, T 70.7, T 107.8 Nadol, Lilli •EP 12.10 Magar, Alessandro GR 9.4 Nagataki, Shigehiro EP 14.10 Näger, Paul M. •AGPhil 9.2 namasaka, wilson •T 47.6 Nardi, Gerrit DD 5.2 Naranjo, Pedro AGPhil 1.1 Nardi, Roberto GP 2.3 Nardo, Lorenzo ST 9.2 Narimani Charan, Abtin •T 79.4 Nasz Esfahani, Alina •T 51.5 Nast, Natalie T 107.4 Nattland, Philipp T 33.2 Navarro, Jaume • GP 6.2
Nackenhorst, Olaf ST 3.4, T 34.5, T 70.7, T 107.8 Nadol, Lilli EP 12.10 Nagar, Alessandro GR 9.4 Nagataki, Shigehiro EP 1.4 Näger, Paul M. AGPhil 9.2 namasaka, wilson T 47.6 Nardi, Gerrit DD 5.2 Naranjo, Pedro AGPhil 1.1 Nardi, Roberto GP 2.3 Nardo, Lorenzo ST 9.2 Narimani Charan, Abtin T 79.4 Nast Esfahani, Alina T 51.5 Nass, Christian T 62.3 Nattlae T 107.4 Nattlae T 107.4 Nattlae T 107.4 Nattlae GP 6.2 Navarro, Jaume GP 6.2 Navarto, Ronny DD 5.2, DD 12.7,
Nackenhorst, Olaf ST 3.4, T 34.5, T 70.7, T 107.8 Nadol, Lilli
Nackenhorst, Olaf ST 3.4, T 34.5, T 70.7, T 107.8 Nadol, Lilli EP 12.10 Nagar, Alessandro GR 9.4 Nagataki, Shigehiro EP 14.10 Näger, Paul M. GPhil 9.2 namasaka, wilson T 47.6 Nandi, Gerrit DD 5.2 Naranjo, Pedro GPhil 1.1 Nardi, Roberto GP 2.3 Nardo, Lorenzo ST 9.2 Narimani Charan, Abtin T 79.4 Nast, Schristian T 62.3 Nast, Natalie T 107.4 Nattland, Philipp T 33.2 Nawrodt, Ronny D 5.2, DD 12.7, DD 31.1, DD 34.12, DD 36.5, DD 44.3 T 64.4
Nackenhorst, Olaf ST 3.4, T 34.5, T 70.7, T 107.8 Nadol, Lilli •EP 12.10 Nagar, Alessandro GR 9.4 Nagataki, Shigehiro EP 14.10 Nagar, Alessandro GR 9.4 Nagataki, Shigehiro EP 1.4 Näger, Paul M. •AGPhil 9.2 namasaka, wilson •T 47.6 Nandi, Gerrit DD 5.2 Nardo, Lorenzo ST 9.2 Narimani Charan, Abtin •T 79.4 Nasr Esfahani, Alina •T 51.5 Nass, Christian •T 62.3 Nattland, Philipp T 33.2 Navarro, Jaume •GP 6.2 Nawrot, Ronny DD 5.2, DD 12.7, DD 31.1, DD 34.12, DD 36.5, DD 44.3 Nayaz, Abdullah Necker, Jannis •T 16.4
Nackenhorst, Olaf ST 3.4, T 34.5, T 70.7, T 107.8 Nadol, Lilli •EP 12.10 Nagar, Alessandro GR 9.4 Nagataki, Shigehiro EP 14.10 Näger, Paul M. •AGPhil 9.2 namasaka, wilson •T 47.6 Naranjo, Pedro AGPhil 9.2 Naranjo, Pedro AGPhil 1.1 Nardi, Gerrit D 5.2 Naranjo, Pedro GP 2.3 Nardo, Lorenzo ST 9.2 Narimani Charan, Abtin •T 79.4 Nast, Schristian •T 62.3 Nast, Natalie T 107.4 Natland, Philipp T 33.2 Navarro, Jaume •GP 6.2 Nawrodt, Ronny DD 5.2, DD 12.7, DD 31.1, DD 34.12, DD 36.5, DD 44.3 Nayaz, Abdullah Nayaz, Abdullah •T 16.4 Necker, Jannis •T 18.2
Nackenhorst, Olaf ST 3.4, T 34.5, T 70.7, T 107.8 Nadol, Lilli •EP 12.10 Nagar, Alessandro GR 9.4 Nagataki, Shigehiro EP 1.4 Näger, Paul M. •AGPhil 9.2 namasaka, wilson •T 47.6 Nardi, Gerrit DD 5.2 Naranjo, Pedro AGPhil 1.1 Nardi, Roberto GP 2.3 Nardo, Lorenzo ST 9.2 Narimani Charan, Abtin •T 79.4 Nast Esfahani, Alina •T 51.5 Nass, Christian •T 62.3 Nattland, Philipp T 33.2 Navarot, Jaume •GP 6.2 Navarot, Ronny DD 5.2, DD 12.7, D 31.1, DD 34.12, DD 36.5, DD 44.3 Nayaz, Abdullah Nayaz, Abdullah •T 16.4 Necker, Jannis •T 18.2 Nedora, Vsevolod •GR 11.3
Nackenhorst, Olaf ST 3.4, T 34.5, T 70.7, T 107.8 Nadol, Lilli •EP 12.10 Nagar, Alessandro GR 9.4 Nagataki, Shigehiro EP 14.10 Näger, Paul M. •AGPhil 9.2 namasaka, wilson •T 47.6 Nandi, Gerrit DD 5.2 Naranjo, Pedro AGPhil 1.1 Nardi, Roberto GP 2.3 Nardo, Lorenzo ST 9.2 Narimani Charan, Abtin •T 79.4 Nast Esfahani, Alina •T 51.5 Nass, Christian •T 62.3 Nattland, Philipp T 33.2 Navarto, Jaume •GP 6.2 Nawrodt, Ronny DD 5.2, DD 12.7, DD 31.1, DD 34.12, DD 36.5, DD 44.3 Nayaz, Abdullah Necker, Jannis •T 18.2 Nedora, Vsevolod •GR 11.3 Negar, Farjoud Masouleh EP 2.2
Nackenhorst, Olaf ST 3.4, T 34.5, T 70.7, T 107.8 Nadol, Lilli •EP 12.10 Nagar, Alessandro GR 9.4 Nagataki, Shigehiro EP 1.4 Näger, Paul M. •AGPhil 9.2 namasaka, wilson •T 47.6 Nandi, Gerrit DD 5.2 Naranjo, Pedro AGPhil 1.1 Nardi, Roberto GP 2.3 Nardo, Lorenzo ST 9.2 Narimani Charan, Abtin •T 79.4 Nasr Esfahani, Alina •T 51.5 Nass, Christian •T 62.3 Nast, Natalie T 107.4 Nattland, Philipp T 33.2 Navarro, Jaume •GP 6.2 Navarot, Ronny DD 5.2, DD 12.7, D 31.1, DD 34.12, DD 36.5, DD 44.3 Nayaz, Abdullah Necker, Jannis •T 18.2 Nedora, Vsevolod •GR 11.3 Negar, Farjoud Masouleh EP 2.2
Nackenhorst, Olaf ST 3.4, T 34.5, T 70.7, T 107.8 Nadol, Lilli •EP 12.10 Nagar, Alessandro GR 9.4 Nagataki, Shigehiro EP 1.4 Näger, Paul M. •AGPhil 9.2 namasaka, wilson •T 47.6 Nardi, Gerrit DD 5.2 Naranjo, Pedro AGPhil 1.1 Nardi, Roberto GP 2.3 Nardo, Lorenzo ST 9.2 Narimani Charan, Abtin •T 79.4 Nasr Esfahani, Alina •T 51.5 Nass, Christian •T 62.3 Nattland, Philipp T 33.2 Navarro, Jaume •GP 6.2 Nawrot, Ronny DD 5.2, DD 12.7, DD 31.1, DD 34.12, DD 36.5, DD 44.3 Nayaz, Abdullah Necker, Jannis •T 16.4 Necker, Jannis •T 18.2 Nedora, Vsevolod •GR 11.3 Negar, Farjoud Masouleh EP 2.2 Nell, Sebastian DD 12.8
Nackenhorst, Olaf ST 3.4, T 34.5, T 70.7, T 107.8 Nadol, Lilli •EP 12.10 Nagar, Alessandro GR 9.4 Nagataki, Shigehiro EP 14.10 Näger, Paul M. •AGPhil 9.2 namasaka, wilson •T 47.6 Naranjo, Pedro AGPhil 1.1 Nardi, Gerrit DD 5.2 Naranjo, Pedro GP 2.3 Nardo, Lorenzo ST 9.2 Narimani Charan, Abtin •T 79.4 Nast Esfahani, Alina •T 612.3 Nast, Natalie T 107.4 Navarro, Jaume •GP 6.2 Navarot, Ronny DD 5.2, DD 12.7, DD 31.1, DD 34.12, DD 36.5, DD 44.3 Nayaz, Abdullah Nedar, Vsevolod •GR 11.3 Nedara, Vsevolod •GR 11.3 Nedar, Vsevolod •GR 11.3 Negar, Farjoud Masouleh EP 2.2 Nell, Sebastian DD 12.8 Nemeth-Csoka, Marton •T 30.5
Nackenhorst, Olaf ST 3.4, T 34.5, T 70.7, T 107.8 Nadol, Lilli •EP 12.10 Nagar, Alessandro GR 9.4 Nagataki, Shigehiro EP 1.4 Näger, Paul M. •AGPhil 9.2 namasaka, wilson •T 47.6 Nardi, Gerrit DD 5.2 Naranjo, Pedro AGPhil 1.1 Nardi, Roberto GP 2.3 Nardo, Lorenzo ST 9.2 Narimani Charan, Abtin •T 79.4 Nast Esfahani, Alina •T 51.5 Nass, Christian •T 62.3 Nattl, Natalie T 107.4 Nattland, Philipp T 33.2 Navarot, Jaume •GP 6.2 Nawarodt, Ronny DD 5.2, DD 12.7, DD 31.1, DD 34.12, DD 36.5, DD 44.3 Nayaz, Abdullah Nayaz, Abdullah •T 16.4 Necker, Jannis •T 18.2 Nedora, Vsevolod •GR 11.3 Negar, Farjoud Masouleh EP 2.2 Nell, Sebastian DD 12.8 Nemeth-Csoka, Marton •T 30.5 Netzwerk Teilchenwelt-Kollaboration
Nackenhorst, Olaf ST 3.4, T 34.5, T 70.7, T 107.8 Nadol, Lilli •EP 12.10 Nagar, Alessandro GR 9.4 Nagataki, Shigehiro EP 1.4 Näger, Paul M. •AGPhil 9.2 namasaka, wilson •T 47.6 Nardi, Gerrit DD 5.2 Nardio, Lorenzo ST 9.2 Naranjo, Pedro AGPhil 1.1 Nardi, Roberto GP 2.3 Nardo, Lorenzo ST 9.2 Narimani Charan, Abtin •T 79.4 Nasr Esfahani, Alina •T 51.5 Nass, Christian •T 62.3 Navarto, Jaume •GP 6.2 Nawrodt, Ronny DD 5.2, DD 12.7, DD 31.1, DD 34.12, DD 36.5, DD 44.3 Nayaz, Abdullah •T 16.4 Necker, Jannis •T 18.2 Nedra, Vsevolod •GR 11.3 Negar, Farjoud Masouleh EP 2.2 Nell, Sebastian DD 12.8 Nemeth-Csoka, Marton •T 30.5 Netzwerk Teilchenwelt-Kollaboration T 24.3, T 24.1, T 52.1, T 52.4, T 24.2,
Nackenhorst, Olaf ST 3.4, T 34.5, T 70.7, T 107.8 Nadol, Lilli EP 12.10 Nagar, Alessandro GR 9.4 Nagataki, Shigehiro EP 1.4 Näger, Paul M. AGPhil 9.2 namasaka, wilson T 47.6 Nandi, Gerrit DD 5.2 Naranjo, Pedro AGPhil 1.1 Nardi, Roberto GP 2.3 Nardo, Lorenzo ST 9.2 Narimani Charan, Abtin T 79.4 Nast Esfahani, Alina T 51.5 Nass, Christian T 62.3 Navarro, Jaume GP 6.2 Navaro, Jaume GP 6.2 Navaro, Jaume GP 12.7 D 31.1, DD 34.12, DD 36.5, DD 44.3 Nayaz, Abdullah Negar, Farjoud Masouleh EP 2.2 Negar, Farjoud Masouleh EP 2.2 Negar, Farjoud Masouleh EP 2.2 Netwerk Teilchen
Nackenhorst, Olaf ST 3.4, T 34.5, T 70.7, T 107.8 Nadol, Lilli EP 12.10 Nagar, Alessandro GR 9.4 Nagataki, Shigehiro EP 1.4 Näger, Paul M. AGPhil 9.2 namasaka, wilson T 47.6 Naranjo, Pedro AGPhil 9.2 Naranjo, Pedro AGPhil 1.1 Nardi, Roberto GP 2.3 Nardo, Lorenzo ST 9.2 Narimani Charan, Abtin T 79.4 Nasr Esfahani, Alina T 51.5 Nass, Christian T 62.3 Nattland, Philipp T 33.2 Navaro, Jaume GP 6.2 Nawrodt, Ronny DD 5.2, DD 12.7, DD 31.1, DD 34.12, DD 36.5, DD 44.3 Nayaz, Abdullah T 16.4 Necker, Jannis T 16.4 Necker, Jannis T 18.2 Nedora, Vsevolod GR 11.3 Negar, Farjoud Masouleh P 2.2 Nell, Sebastian DD 12.8 Nemeth-Csoka, Marton T 30.5 Netzwerk Teilchenwelt-Kollaboration T 24.3, T 24.1,
Nackenhorst, Olaf ST 3.4, T 34.5, T 70.7, T 107.8 Nadol, Lilli •EP 12.10 Nagar, Alessandro GR 9.4 Nagataki, Shigehiro EP 14.10 Näger, Paul M. •AGPhil 9.2 namasaka, wilson •T 47.6 Nardi, Gerrit DD 5.2 Naranjo, Pedro AGPhil 1.1 Nardi, Roberto GP 2.3 Nardo, Lorenzo ST 9.2 Narimani Charan, Abtin •T 79.4 Nast Esfahani, Alina •T 51.5 Nass, Christian •T 62.3 Nattland, Philipp T 33.2 Navarro, Jaume •GP 6.2 Nawrodt, Ronny DD 5.2, DD 12.7, DD 31.1, DD 34.12, DD 36.5, DD 44.3 Nayaz, Abdullah •T 16.4 Necker, Jannis •T 18.2 Nedra, Vsevolod •GR 11.3 Negar, Farjoud Masouleh EP 2.2 Nell, Sebastian DD 12.8 Netwerk Teilchenwelt-Kollaboration T 30.5 Netzwerk Teilchenwelt-Kollaboration T 24.3, T 24.1, T 52.1, T 52.4, T 24.2, T 52.3 Neuberger, Moritz •T 75.3 Neu
Nackenhorst, Olaf ST 3.4, T 34.5, T 70.7, T 107.8 Nadol, Lilli EP 12.10 Nagar, Alessandro GR 9.4 Nagataki, Shigehiro EP 1.4 Näger, Paul M. AGPhil 9.2 namasaka, wilson T 47.6 Nardi, Gerrit DD 5.2 Naranjo, Pedro AGPhil 1.1 Nardi, Roberto GP 2.3 Nardo, Lorenzo ST 9.2 Narimani Charan, Abtin T 79.4 Nasr Esfahani, Alina T 51.5 Nass, Christian T 62.3 Nast, Natalie T 107.4 Nattland, Philipp T 32.2 Navarro, Jaume GP 6.2 Navarodt, Ronny DD 5.2, DD 12.7, DD 31.1, DD 34.12, DD 36.5, DD 44.3 Nayaz, Abdullah T 18.2 Nedora, Vsevolod GR 11.3 Negar, Farjoud Masouleh EP 2.2 Nell, Sebastian D 12.5, Netzwerk Teilchenwelt-Kollaboration T 24.3, T 24.1, T 52.1, T 52.4, T 24.2, T 52.3 T 53.3 Neuberger, Moritz T 75.3 Neuendorff
Nackenhorst, Olaf ST 3.4, T 34.5, T 70.7, T 107.8 Nadol, Lilli EP 12.10 Nagar, Alessandro GR 9.4 Nagataki, Shigehiro EP 1.4 Näger, Paul M. AGPhil 9.2 namasaka, wilson T 47.6 Nandi, Gerrit DD 5.2 Naranjo, Pedro AGPhil 1.1 Nardi, Roberto GP 2.3 Nardo, Lorenzo ST 9.2 Narimani Charan, Abtin T 79.4 Nasr Esfahani, Alina T 51.5 Nass, Christian T 62.3 Nattland, Philipp T 33.2 Navaro, Jaume GP 6.2 Nawrott, Ronny DD 5.2, DD 12.7, DD 31.1, DD 34.12, DD 36.5, DD 44.3 Nayaz, Abdullah T 16.4 Necker, Jannis T 18.2 Nedora, Vsevolod GR 11.3 Negar, Farjoud Masouleh EP 2.2 Nell, Sebastian DD 12.8 Nemeth-Csoka, Marton 30.5 Netzwerk Teilchenwelt-Kollaboration T 24.3, T 24.1, T 52.1, T 52.4, T 24.2, T 52.3 Neuberger, Moritz 75.3 Neuberger, Moritz .
Nackenhorst, Olaf ST 3.4, T 34.5, T 70.7, T 107.8 Nadol, Lilli •EP 12.10 Nagar, Alessandro GR 9.4 Nagataki, Shigehiro EP 1.4 Näger, Paul M. •AGPhil 9.2 namasaka, wilson •T 47.6 Nandi, Gerrit DD 5.2 Naranjo, Pedro AGPhil 1.1 Nardi, Roberto GP 2.3 Nardo, Lorenzo ST 9.2 Narimani Charan, Abtin •T 79.4 Nast Esfahani, Alina •T 51.5 Nass, Christian •T 62.3 Nast, Natalie T 107.4 Nattland, Philipp T 33.2 Navarro, Jaume •GP 6.2 Nawrodt, Ronny DD 5.2, DD 12.7, D 31.1, DD 34.12, DD 36.5, DD 44.3 Nayaz, Abdullah •T 18.2 Nedora, Vsevolod •GR 11.3 Negar, Farjoud Masouleh EP 2.2 Netwerk Teilchenwelt-Kollaboration T 24.3, T 24.1, T 52.1, T 52.4, T 24.2, T 52.3 Neuberger, Moritz •T 75.3 Neuendorffer, Stephen T 69.7 Neuendorffer, Stephen T 69.7
Nackenhorst, Olaf ST 3.4, T 34.5, T 70.7, T 107.8 Nadol, Lilli EP 12.10 Nagar, Alessandro GR 9.4 Nagataki, Shigehiro EP 14.10 Nagar, Alessandro GR 9.4 Nagataki, Shigehiro EP 14.10 Nagar, Alessandro GR 9.4 Nagar, Alessandro AGPhil 9.2 namasaka, wilson T 47.6 Nardi, Gerrit DD 5.2 Naranjo, Pedro AGPhil 1.1 Nardi, Roberto GP 2.3 Nardo, Lorenzo ST 9.2 Narimani Charan, Abtin T 79.4 Nasr Esfahani, Alina T 75.15 Nass, Christian T 62.3 Nattland, Philipp T 33.2 Navaro, Jaume GP 6.2 Nawrodt, Ronny DD 5.2, DD 12.7, DD 31.1, DD 34.12, DD 36.5, DD 44.3 Nayaz, Abdullah T 16.4 Necker, Jannis T 16.4 Necker, Jannis T 18.2 Nedora, Vsevolod GR 11.3 Negar, Farjoud Masouleh EP 2.2 Nell, Sebastian DD 12.8 <t< td=""></t<>
Nackenhorst, Olaf ST 3.4, T 34.5, T 70.7, T 107.8 Nadol, Lilli EP 12.10 Nagar, Alessandro GR 9.4 Nagataki, Shigehiro EP 14.10 Näger, Paul M. AGPhil 9.2 namasaka, wilson T 47.6 Nandi, Gerrit D5.2 Naranjo, Pedro AGPhil 1.1 Nardi, Roberto GP 2.3 Nardo, Lorenzo ST 9.2 Narimani Charan, Abtin T 79.4 Nast, Schristian T 62.3 Nass, Christian T 62.3 Nast, Natalie T 107.4 Nattland, Philipp T 33.2 Navarro, Jaume GP 6.2 Nawrodt, Ronny D 5.2, DD 12.7, DD 31.1, DD 34.12, DD 36.5, DD 44.3 Nayaz, Abdullah Nedora, Vsevolod GR 11.3 Negar, Farjoud Masouleh EP 2.2 Nell, Sebastian D 12.8 Nemeth-Csoka, Marton T 30.5 Netzwerk Teilchenwelt-Kollaboration T 23.3 Neuendorffer, Stephen
Nackenhorst, Olaf ST 3.4, T 34.5, T 70.7, T 107.8 Nadol, Lilli •EP 12.10 Nagar, Alessandro GR 9.4 Nagataki, Shigehiro EP 1.4 Näger, Paul M. •AGPhil 9.2 namasaka, wilson •T 47.6 Nardi, Gerrit DD 5.2 Naranjo, Pedro AGPhil 1.1 Nardi, Roberto GP 2.3 Nardo, Lorenzo ST 9.2 Narimani Charan, Abtin •T 79.4 Nasr Esfahani, Alina •T 51.5 Nass, Christian •T 62.3 Natt, Natalie T 107.4 Nattland, Philipp T 33.2 Navarro, Jaume •GP 6.2 Nawrodt, Ronny DD 5.2, DD 12.7, DD 31.1, DD 34.12, DD 36.5, DD 44.3 Nayaz, Abdullah •T 16.4 Necker, Jannis •T 18.2 Nedora, Vsevolod •GR 11.3 Negar, Farjoud Masouleh EP 2.2 Nell, Sebastian DD 12.8 Nemeth-Csoka, Marton •T 30.2 Neuberger, Moritz •T 75.3 Neuendorffer, Stephen T 6
Nackenhorst, Olaf ST 3.4, T 34.5, T 70.7, T 107.8 Nadol, Lilli •EP 12.10 Nagar, Alessandro GR 9.4 Nagataki, Shigehiro .EP 1.4 Näger, Paul M. •AGPhil 9.2 namasaka, wilson •T 47.6 Nardi, Gerrit DD 5.2 Naranjo, Pedro AGPhil 1.1 Nardi, Roberto GP 2.3 Nardo, Lorenzo ST 9.2 Narimani Charan, Abtin •T 79.4 Nast Esfahani, Alina •T 51.5 Nass, Christian •T 62.3 Nattland, Philipp T 33.2 Navaro, Jaume •GP 6.2 Nawrot, Ronny DD 5.2, DD 12.7, DD 31.1, DD 34.12, DD 36.5, DD 44.3 Nayaz, Abdullah Netzwerk, Tonnis •T 16.4 Necker, Jannis •T 16.4 Negar, Farjoud Masouleh EP 2.2 Nell, Sebastian DD 12.8 Nemeth-Csoka, Marton •T 30.5 Netzwerk Teilchenwelt-Kollaboration T 24.3, T 24.1, T 52.1, T 52.4, T 24.2, T 53.3 Neuberger, Moritz •T 75.3 Neuendorffer, Stephen T 65.7, T 63.4 Neuhaus, Simon <td< td=""></td<>
Nackenhorst, Olaf ST 3.4, T 34.5, T 70.7, T 107.8 Nadol, Lilli EP 12.10 Nagar, Alessandro GR 9.4 Nagataki, Shigehiro EP 12.10 Nagar, Alessandro GR 9.4 Nagataki, Shigehiro EP 1.4 Näger, Paul M. AGPhil 9.2 namasaka, wilson T 47.6 Nandi, Gerrit D 5.2 Naranjo, Pedro AGPhil 1.1 Nardi, Roberto GP 2.3 Nardo, Lorenzo ST 9.2 Narimani Charan, Abtin T 79.4 Nasr Esfahani, Alina T 51.5 Nass, Christian T 62.3 Nast, Natalie T 107.4 Nattland, Philipp T 33.2 Navarro, Jaume GP 6.2 Navarro, Jaume GP 6.2 Navarot, Ronny D 5.2, DD 12.7, D 31.1, DD 34.12, DD 36.5, DD 44.3 Nayaz, Abdullah Negar, Farjoud Masouleh EP 2.2 Nedra, Vsevolod GR 11.3 Negar, Farjoud Masouleh EP 2.2 Nell, Sebastian D 12.8 Nemeth-Csoka, Marton .
Nackenhorst, Olaf ST 3.4, T 34.5, T 70.7, T 107.8 Nadol, Lilli •EP 12.10 Nagar, Alessandro GR 9.4 Nagataki, Shigehiro EP 1.4 Näger, Paul M. •AGPhil 9.2 namasaka, wilson •T 47.6 Naranjo, Pedro AGPhil 9.2 Naranjo, Pedro AGPhil 1.1 Nardi, Roberto GP 2.3 Nardo, Lorenzo ST 9.2 Narimani Charan, Abtin •T 79.4 Nasr Esfahani, Alina •T 51.5 Nass, Christian •T 62.3 Nattland, Philipp T 33.2 Navaro, Jaume •GP 6.2 Nawrodt, Ronny DD 5.2, DD 12.7, DD 31.1, DD 34.12, DD 36.5, DD 44.3 Nayaz, Abdullah •T 16.4 Necker, Jannis •T 16.4 Necker, Jannis •T 18.2 Nedora, Vsevolod •GR 11.3 Negar, Farjoud Masouleh EP 2.2 Nell, Sebastian DD 12.8 Netwerk Teilchenwelt-Kollaboration T 24.3, T 24.1, T 52.4, T 24.2, T 53.3 Neuberger, Moritz •T 75.3
Nackenhorst, Olaf ST 3.4, T 34.5, T 70.7, T 107.8 Nadol, Lilli •EP 12.10 Nagar, Alessandro GR 9.4 Nagataki, Shigehiro EP 1.4 Näger, Paul M. •AGPhil 9.2 namasaka, wilson •T 47.6 Nardi, Gerrit DD 5.2 Naranjo, Pedro AGPhil 1.1 Nardi, Roberto GP 2.3 Nardo, Lorenzo ST 9.2 Narimani Charan, Abtin •T 79.4 Nast Esfahani, Alina •T 51.5 Nass, Christian •T 62.3 Nast, Natalie T 107.4 Nattland, Philipp T 33.2 Navarro, Jaume •GP 6.2 Navarro, Jaume •GP 6.2 Navaro, Jaume •GP 6.2 Neusr, Karjoud Masouleh EP 2.2 Neldora, Vsevolod<
Nackenhorst, Olaf ST 3.4, T 34.5, T 70.7, T 107.8 Nadol, Lilli •EP 12.10 Nagar, Alessandro GR 9.4 Nagataki, Shigehiro EP 1.4 Näger, Paul M. •AGPhil 9.2 namasaka, wilson •T 47.6 Nardi, Gerrit DD 5.2 Naranjo, Pedro AGPhil 1.1 Nardi, Roberto GP 2.3 Nardo, Lorenzo ST 9.2 Narimani Charan, Abtin •T 79.4 Nasr Esfahani, Alina •T 51.5 Nass, Christian •T 62.3 Natt, Natalie T 107.4 Nattland, Philipp T 33.2 Navarro, Jaume •GP 6.2 Nawrodt, Ronny DD 5.2, DD 12.7, DD 31.1, DD 34.12, DD 36.5, DD 44.3 Nayaz, Abdullah •T 16.4 Necker, Jannis •T 18.2 Nedora, Vsevolod •GR 11.3 Negar, Farjoud Masouleh EP 2.2 Nell, Sebastian DD 12.8 Nemeth-Csoka, Marton •T 30.2 Neuberger, Moritz •T 75.3 Neuendorffer, Stephen T 6
Nackenhorst, Olaf ST 3.4, T 34.5, T 70.7, T 107.8 Nadol, Lilli
Nackenhorst, Olaf ST 3.4, T 34.5, T 70.7, T 107.8 Nadol, Lilli EP 12.10 Nagar, Alessandro GR 9.4 Nagataki, Shigehiro EP 1.4 Näger, Paul M. AGPhil 9.2 namasaka, wilson T 47.6 Nandi, Gerrit D 5.2 Naranjo, Pedro AGPhil 1.1 Nardi, Roberto GP 2.3 Nardo, Lorenzo ST 9.2 Narimani Charan, Abtin T 79.4 Nasr Esfahani, Alina T 51.5 Nass, Christian T 62.3 Nast, Natalie T 107.4 Nattland, Philipp T 33.2 Navarro, Jaume GP 6.2 Navarro, Jaume GP 6.2 Navarro, Jaume GP 6.2 Navarot, Ronny DD 5.2, DD 12.7, D 31.1, DD 34.12, DD 36.5, DD 44.3 Nayaz, Abdullah Nedra, Vsevolod GR 11.3 Negar, Farjoud Masouleh EP 2.2 Nell, Sebastian D 12.8 Nemeth-Csoka, Marton T 30.5 Netzwerk Teilchenwelt-Kollaboration T 24.3, T 24.1, T 52.1, T 52.4, T 24.2, T 52.3
Nackenhorst, Olaf ST 3.4, T 34.5, T 70.7, T 107.8 Nadol, Lilli -EP 12.10 Nagar, Alessandro GR 9.4 Nagataki, Shigehiro EP 1.4 Näger, Paul M. -AGPhil 9.2 namasaka, wilson -T 47.6 Nardi, Gerrit DD 5.2 Naranjo, Pedro AGPhil 1.1 Nardi, Roberto GP 2.3 Nardo, Lorenzo ST 9.2 Narimani Charan, Abtin -T 79.4 Nasr Esfahani, Alina -T 51.5 Nass, Christian -T 62.3 Natt, Natalie T 107.4 Nattland, Philipp T 33.2 Navarro, Jaume -GP 6.2 Nawrodt, Ronny DD 5.2, DD 12.7, DD 31.1, DD 34.12, DD 36.5, DD 44.3 Nayaz, Abdullah -T 16.4 Necker, Jannis -T 16.4 Necker, Jannis -T 18.2 Nedora, Vsevolod -GR 11.3 Negar, Farjoud Masouleh EP 2.2 Nell, Sebastian DD 12.8 Nemeth-Csoka, Marton T 33.6 Neuberger, Moritz -T 75.3
Nackenhorst, Olaf ST 3.4, T 34.5, T 70.7, T 107.8 Nadol, Lilli EP 12.10 Nagar, Alessandro GR 9.4 Nagataki, Shigehiro EP 1.4 Näger, Paul M. AGPhil 9.2 namasaka, wilson T 47.6 Nandi, Gerrit D 5.2 Naranjo, Pedro AGPhil 1.1 Nardi, Roberto GP 2.3 Nardo, Lorenzo ST 9.2 Narimani Charan, Abtin T 79.4 Nast Esfahani, Alina T 51.5 Nass, Christian T 62.3 Navarro, Jaume GP 6.2 Navarro, Jaume GP 6.2 Navarro, Jaume GP 6.2 Navaro, Jaume DD 5.2, DD 12.7 D 31.1, DD 34.12, DD 36.5, DD 44.
Nackenhorst, Olaf ST 3.4, T 34.5, T 70.7, T 107.8 Nadol, Lilli -EP 12.10 Nagar, Alessandro GR 9.4 Nagataki, Shigehiro EP 1.4 Näger, Paul M. -AGPhil 9.2 namasaka, wilson -T 47.6 Nardi, Gerrit DD 5.2 Naranjo, Pedro AGPhil 1.1 Nardi, Roberto GP 2.3 Nardo, Lorenzo ST 9.2 Narimani Charan, Abtin -T 79.4 Nasr Esfahani, Alina -T 51.5 Nass, Christian -T 62.3 Natt, Natalie T 107.4 Nattland, Philipp T 33.2 Navarro, Jaume -GP 6.2 Nawrodt, Ronny DD 5.2, DD 12.7, DD 31.1, DD 34.12, DD 36.5, DD 44.3 Nayaz, Abdullah -T 16.4 Necker, Jannis -T 18.2 Nedora, Vsevolod -GR 11.3 Negar, Farjoud Masouleh EP 2.2 Nell, Sebastian DD 12.8 Nemeth-Csoka, Marton -T 30.5 Neuberger, Moritz -T 75.3 Neuendorffer, Stephen T
Nackenhorst, OlafST 3.4, T 34.5, T 70.7, T 107.8Nadol, Lilli \cdot EP 12.10Nagar, AlessandroGR 9.4Nagataki, ShigehiroEP 1.4Näger, Paul M. \cdot AGPhil 9.2namasaka, wilson \cdot T 47.6Nandi, GerritD 5.2Naranjo, PedroAGPhil 1.1Nardi, RobertoGP 2.3Narinani Charan, Abtin \cdot T 79.4Nast Esfahani, Alina \cdot T 51.5Nass, Christian \cdot T 62.3Nast, NatalieT 107.4Nattland, PhilippT 33.2Navarro, Jaume \cdot GP 6.2Nawrott, RonnyD 5.2, DD 12.7,DD 31.1, DD 34.12, DD 36.5, DD 44.3Nayaz, Abdullah \cdot T 16.4Necker, Jannis \cdot T 18.2Nedora, Vsevolod \cdot GR 11.3Negar, Farjoud MasoulehEP 2.2Nell, SebastianDD 12.8Nemeth-Csoka, Marton \cdot T 30.5Netwerk Teilchenwelt-KollaborationT 24.3, T 24.1, T 52.1, T 52.4, T 24.2, T 52.3Neuberger, Moritz \cdot T 75.3Neuberger, Moritz \cdot T 75.3Neuberger, Moritz \cdot T 75.3Neuberger, Moritz \cdot T 75.3Neubaus, Simon \cdot T 84.6Neukirch, Thomas \cdot F 12.10Neukirch, Thomas \cdot T 87.3Nickel, Lukas \cdot T 17.4, T 77.5Nguyen, Le HoangT 77.4, T 77.5Nguyen, Le HoangT 77.4, T 77.5Nickeler, Dieter \cdot EP 1.2Nickler, Dieter \cdot EP 1.2Nickler, Dieter \cdot EP 1.2Nickler, Diete
Nackenhorst, OlafST 3.4, T 34.5, T 70.7, T 107.8Nadol, Lilli \cdot EP 12.10Nagar, AlessandroGR 9.4Nagataki, ShigehiroEP 1.4Näger, Paul M. \cdot AGPhil 9.2namasaka, wilson \cdot T 47.6Nandi, GerritD 5.2Naranjo, PedroAGPhil 1.1Nardi, RobertoGP 2.3Nardo, LorenzoST 9.2Narimani Charan, Abtin \cdot T 79.4Nast Esfahani, Alina \cdot T 51.5Nass, Christian \cdot T 62.3Nast, NatalieT 107.4Nattland, PhilippT 33.2Navarot, Jaume \cdot GP 6.2Nawrodt, RonnyDD 5.2, DD 12.7, D 31.1, DD 34.12, DD 36.5, DD 44.3Nayaz, Abdullah \cdot T 16.4Necker, Jannis \cdot T 18.2Nedora, Vsevolod \cdot GR 11.3Negar, Farjoud MasoulehEP 2.2Nell, SebastianDD 12.8Nemeth-Csoka, MartonT 30.5Netzwerk Teilchenwelt-KollaborationT 24.3, T 24.1, T 52.1, T 52.4, T 24.2, T 52.3Neundorffer, StephenT 65.5, T 65.6, T 65.7, T 65.4Neukirch, ThomasEP 12.10Neukurn, Max \cdot T 87.5Neundorff, Jonas \cdot T 87.5Nguyen, Le HoangT 77.4, T 77.5Nguyen, Le HoangT 77.4, T 77.5Nguyen, Le HoangT 77.4, T 77.5Nickeler, Dieter \cdot EP 12.10Nickeler, Dieter \cdot EP 12.10Nickeler, Dieter \cdot EP 12.10Nickeler, Dieter \cdot EP 12.10Nickeler, Dieter \cdot EP 12.10Nicke
Nackenhorst, OlafST 3.4, T 34.5, T 70.7, T 107.8Nadol, Lilli-EP 12.10Nagar, AlessandroGR 9.4Nagataki, ShigehiroEP 1.4Näger, Paul MAGPhil 9.2namasaka, wilson-T 47.6Nardi, GerritDD 5.2Naranjo, PedroAGPhil 1.1Nardi, RobertoGP 2.3Nardo, LorenzoST 9.2Narimani Charan, Abtin-T 79.4Nast Esfahani, Alina-T 51.5Nass, Christian-T 62.3Nator, Jaume-GP 6.2Navrodt, RonnyDD 5.2, DD 12.7, DD 31.1, DD 34.12, DD 36.5, DD 44.3Nayaz, Abdullah-T 16.4Necker, Jannis-T 18.2Nedora, Vsevolod-GR 11.3Negar, Farjoud MasoulehEP 2.2Nell, SebastianDD 12.8Neuberger, Moritz-T 75.3Neuendorffer, StephenT 69.7Neufeld, MariusT 65.5, T 65.6, T 65.7, T 94.4Neukorg, Jonas-T 63.9Nguyen, Le HoangT 77.4, T 77.5Nguyen, Minh-Duc-AKPIK 1.5Nickeler, Dieter-EP 1.2Nickeler, Dieter-EP 1.2Nickel, Lukas-T 17.3Nickeler, Dieter-EP 1.2Nickeler, Dieter
Nackenhorst, Olaf ST 3.4, T 34.5, T 70.7, T 107.8 Nadol, Lilli EP 12.10 Nagar, Alessandro GR 9.4 Nagataki, Shigehiro EP 12.10 Nagar, Alessandro GR 9.4 Nagataki, Shigehiro EP 1.4 Näger, Paul M. AGPhil 9.2 namasaka, wilson T 47.6 Nandi, Gerrit D 5.2 Naranjo, Pedro AGPhil 1.1 Nardi, Roberto GP 2.3 Naranjo, Lorenzo ST 9.2 Narimani Charan, Abtin T 79.4 Nast Esfahani, Alina T 51.5 Nass, Christian T 62.3 Navarro, Jaume GP 6.2 Navarro, Jaume GP 6.2 Navarot, Ronny DD 5.2, DD 12.7, D 31.1, DD 34.12, DD 36.5, DD 44.3 Nayaz, Abdullah Nedcra, Vsevolod GR 11.3 Negar, Farjoud Masouleh E 2.2 Nell, Sebastian D 12.8 Nemeth-Csoka, Marton T 30.5 Netzwerk Teilchenwelt-Kollaboration T 24.3, T 24.1, T 52.1, T 52.4, T 24.2, T 55.3 Neuendorffer, Stephen G6.7 </td
Nackenhorst, Olaf ST 3.4, T 34.5, T 70.7, T 107.8 Nadol, Lilli EP 12.10 Nagar, Alessandro GR 9.4 Nagataki, Shigehiro EP 12.10 Nagar, Alessandro GR 9.4 Nagataki, Shigehiro EP 1.4 Näger, Paul M. AGPhil 9.2 namasaka, wilson T 47.6 Nardi, Gerrit D 5.2 Naranjo, Pedro AGPhil 1.1 Nardi, Roberto GP 2.3 Nardo, Lorenzo ST 9.2 Narimani Charan, Abtin T 79.4 Nasr Esfahani, Alina T 51.5 Nass, Christian T 62.3 Nast, Natalie T 107.4 Nattland, Philipp T 33.2 Navarro, Jaume GP 6.2 Nawrodt, Ronny DD 5.2, DD 12.7, DD 31.1, DD 34.12, DD 36.5, DD 44.3 Nayaz, Abdullah Necker, Jannis T 18.2 Nedora, Vsevolod GR 11.3 Negar, Farjoud Masouleh EP 2.2 Nell, Sebastian D 12.8 Nemeth-Csoka, Marton T 30.5 Netzwerk Teilchenwelt-Kollaboration </td
Nackenhorst, OlafST 3.4, T 34.5, T 70.7, T 107.8Nadol, Lilli \cdot EP 12.10Nagar, AlessandroGR 9.4Nagataki, ShigehiroEP 1.4Näger, Paul M. \cdot AGPhil 9.2namasaka, wilson \cdot T 47.6Nandi, GerritDD 5.2Naranjo, PedroAGPhil 1.1Nardi, RobertoGP 2.3Nardo, LorenzoST 9.2Narimani Charan, AbtinT 79.4Nasr Esfahani, AlinaT 51.5Nass, ChristianT 62.3Nato, LorenzoST 9.2Narton, JaumeGP 6.2Navrodt, RonnyDD 5.2, DD 12.7, DD 31.1, DD 34.12, DD 36.5, DD 44.3Nayaz, Abdullah \cdot T 16.4Necker, Jannis \cdot T 18.2Nedora, VsevolodGR 11.3Negar, Farjoud MasoulehEP 2.2Nell, SebastianDD 12.8Neuberger, Moritz \cdot T 75.3Neuerger, Moritz \cdot T 75.3Neudorffer, StephenT 69.7Neufeld, MariusT 65.5, T 65.6, T 65.7, T 94.4Neubaus, Simon \cdot T 34.6Neukorf, Jonas \cdot T 63.9Nguyen, Le HoangT 77.4, T 77.5Nguyen, Le HoangT 77.4, T 77.5Nickeler, Dieter \cdot EP 12.0Nickeler, Dieter \cdot EP 12.2Nickeler, Dieter \cdot EP 12.0Nickeler, Dieter \cdot EP 12.0Nickeler, Dieter \cdot E

	•T 86.6
Nissan, Yuvai	•PV VII
Nitsche, Hannes	.•T 52.6
Nitschke, Anna-Katharina	•ST 2.4,
ST 3.2 Noethe Maximilian	T 101 2
Noethlichs, Michel . DD 30.1,	•DD 47.1
Nöll, Andreas EP 4.	3, EP 4.6
Noll, DennisT 61.5, •T 61	.6, T 61.7
Nolte, Niklas 1 26.	6, 1 43.6 T 79.6
Nordmeier, Volkhard DD 11.2	. DD 16.2
Noronha, Jorge	. MP 2.3
Northe, Christian	MP 1.2
Nothe for the CTA Consortium	, , , , , , , , , , , , , , , , , , ,
Nöthe for the CTA LST Project,	IXI IIX 2.0
Maximilian	T 17.3
Nöthe, Maximilian•A	KPIK 1.2,
Novak, Andrzei T 26.1, T 26.3	•T 89.1
T 89.1, T 89.2, T 90.6, T 105.	10
Novgorodova, Olga	. •ST 1.4
NOWOTNY, Fabian	•1 3/.2
Nürnberg, Andreas	T 94.4
O. Silva, Hector	.•GR 2.7
Obczovsky, Markus	•DD 11.5
Uberauer, Lothar I 20.	8, 1 48.7,
Oberlack. Uwe	DD 46.2
Odagiu, Patrick	T 10.9
Oder, Bernd	DD 34.4
Oebler Marie	I 94.3
Oertel, Jessica •DD 4.2.	DD 20.4
Oeser, Thomas	•T 4.7
Oez, Erdem	•T 105.2
Ogata, Yoshiko	MP 1.1
Ohme. Frank	.•GR 8.1
Ohtani, Yoshiki	T 17.8
Oikonomou, Foteini	T 71.9
Olbrich, Lena	. •ST 7.2
Oleksivuk Ivan	•T 93.4
Ulischlager, Hans	T 93.1
Olischlager, Hans	T 93.1 EP 1.6
Oliva, André	T 93.1 EP 1.6 •T 17.2
Olischiager, Hans Oliva, André Olivera-Nieto, Laura Olschewsky, Kevin Omeliukh, Anastasija	T 93.1 EP 1.6 •T 17.2 .•T 56.7 •EP 7.10
Oliva, André Oliva, André Olivera-Nieto, Laura Olschewsky, Kevin Omeliukh, Anastasiia Omidvari, Negar	T 93.1 EP 1.6 •T 17.2 T 56.7 •EP 7.10 ST 9.2
Oliva, André Oliva, André Olivera-Nieto, Laura Olschewsky, Kevin Omeliukh, Anastasiia Omidvari, Negar on Envision team, VenSpec	T 93.1 EP 1.6 •T 17.2 T 56.7 •EP 7.10 ST 9.2 EP 3.1
Olischlager, Hans Oliva, André Olivera-Nieto, Laura Olschewsky, Kevin Omeliukh, Anastasiia Omidvari, Negar on Envision team, VenSpec on VERTIAS team, VEM	T 93.1 EP 1.6 •T 17.2 •T 56.7 •EP 7.10 ST 9.2 EP 3.1 EP 3.1
Olischlager, Hans Oliva, André Olivera-Nieto, Laura Olschewsky, Kevin Omeliukh, Anastasiia Omidvari, Negar on Envision team, VenSpec on VERTIAS team, VEM Öncel, Ö. Oğul Onvisi, Peter T 60.	T 93.1 EP 1.6 •T 17.2 •T 56.7 •EP 7.10 ST 9.2 EP 3.1 EP 3.1 3.4, T 8.5 3. T 60.4
Olischager, Hans Oliva, André Olivera-Nieto, Laura Olschewsky, Kevin Omeliukh, Anastasiia Omidvari, Negar on Envision team, VenSpec on VERTIAS team, VEM Öncel, Ö. Oğul Onyisi, Peter Ordek, Serhat	T 93.1 EP 1.6 •T 17.2 •T 56.7 •EP 7.10 ST 9.2 EP 3.1 EP 3.1 3.4, T 8.5 3, T 60.4 T 89.4
Olischager, Hans Oliva, André Olivera-Nieto, Laura Olschewsky, Kevin Omeliukh, Anastasiia Omidvari, Negar on Envision team, VenSpec on VERTIAS team, VEM Öncel, Ö. Oğul Onyisi, Peter Ordek, Serhat Orekhov, Vsevolod	T 93.1 EP 1.6 •T 17.2 .•T 56.7 •EP 7.10 ST 9.2 EP 3.1 3.4, T 8.5 3, T 60.4 T 89.4 •T 102.9
Olischlager, Hans Oliva, André Olivera-Nieto, Laura Olschewsky, Kevin Omeliukh, Anastasiia Omidvari, Negar on Envision team, VenSpec on VERTIAS team, VEM Öncel, Ö. Oğul Onyisi, Peter Ordek, Serhat Orekhov, Vsevolod Ørsøe, Rasmus Ortiohan, Hans-Werner T 21	T 93.1 EP 1.6 •T 17.2 .•T 56.7 •EP 7.10 EP 3.1 EP 3.1 3.4, T 8.5 3, T 60.4 T 89.4 •T 102.9 KPIK 4.7 KPIK 4.7
Olischlager, Hans Oliva, André Olivera-Nieto, Laura Olschewsky, Kevin Omeliukh, Anastasiia Omidvari, Negar on Ervision team, VenSpec on VERTIAS team, VEM Öncel, Ö. Oğul Onyisi, Peter Ordek, Serhat Orekhov, Vsevolod Ørsøe, Rasmus Ortjohann, Hans-Werner T 21 Ortmann, Julia	T 93.1 EP 1.6 •T 17.2 T 56.7 •EP 7.10 ST 9.2 EP 3.1 EP 3.1 3.4, T 8.5 3, T 60.4 T 89.4 •T 102.9 KPIK 4.7 4, T 21.5 •DD 50.2
Olischlager, Hans Oliva, André Olivara-Nieto, Laura Olschewsky, Kevin Omeliukh, Anastasiia Omidvari, Negar on Envision team, VenSpec on VERTIAS team, VEM Öncel, Ö. Oğul Onyisi, Peter Ordek, Serhat Orekhov, Vsevolod Ørsøe, Rasmus Ortjohann, Hans-Werner T 21 Ortmann, Julia Ortmann, Julia	T 93.1 EP 1.6 •T 56.7 •EP 7.10 ST 9.2 EP 3.1 3.4, T 8.5 3, T 60.4 T 89.4 •T 102.9 KPIK 4.7 4, T 21.5 •DD 50.2 7, T 78.6,
Olischiager, Hans Oliva, André Olivara-Nieto, Laura Olschewsky, Kevin Omeliukh, Anastasiia Omidvari, Negar on Envision team, VenSpec on VERTIAS team, VEM Öncel, Ö. Oğul Onyisi, Peter Ordek, Serhat Orekhov, Vsevolod Ørsøe, Rasmus Ortghoann, Hans-Werner Ortmann, Tobias Ortmann, Tobias T74. T78.7, T 104.5, T 104.6, •T 10	T 93.1 EP 1.6 T 17.2 T 56.7 EP 7.10 EP 3.1 3.4, T 8.5 3, T 60.4 T 89.4 T 89.4 T 89.4 T 89.4 T 89.2 KPIK 4.7 .4, T 21.5 D 50.2 7, T 78.6, 4 7
Olischlager, Hans Oliva, André Olivara-Nieto, Laura Olschewsky, Kevin Omeliukh, Anastasiia Omidvari, Negar on Ervision team, VenSpec on VERTIAS team, VEM Öncel, Ö. Oğul Oryisi, Peter Ordek, Serhat Orekhov, Vsevolod Ørsøe, Rasmus Ortkono, Hans-Werner Ortmann, Tobias Ortmann, Tobias T74. T78.7, T 104.5, T 104.6, •T 10 Ossekine, Serguei	T 93.1 EP 1.6 T 17.2 T 56.7 EP 7.10 EP 3.1 3.4, T 8.5 3, T 60.4 T 89.4 T 89.4 VFIK 4.7 .4, T 21.5 DD 50.2 7, T 78.6, 0 4.7 GR 9.2 5 T 44.6
Olischlager, Hans Oliva, André Olivera-Nieto, Laura Olschewsky, Kevin Omeliukh, Anastasiia Omidvari, Negar on Ervision team, VenSpec on VERTIAS team, VEM Öncel, Ö. Oğul Oryisi, Peter Ordek, Serhat Orekhov, Vsevolod Ørsøe, Rasmus Ortkon, Hans-Werner Otrmann, Tobias Ortmann, Tobias Ortmann, Tobias T74. T78.7, T 104.5, T 104.6, •T 10 Ossokine, Serguei Osterhoff, Jens Cthman, Gulden	T 93.1 EP 1.6 T 17.2 T 56.7 EP 7.10 EP 3.1 3.4, T 8.5 3, T 60.4 T 89.4 T 89.4 VF1 K 4.7 .4, T 21.5 D 50.2 7, T 78.6, 0.4, 7 GR 9.2 5, T 44.6
Olischlager, Hans Oliva, André Olivara-Nieto, Laura Olschewsky, Kevin Omeliukh, Anastasiia Omidvari, Negar on Envision team, VenSpec on VERTIAS team, VEM Öncel, Ö. Oğul Onyisi, Peter Oreklov, Vsevolod Ørsøe, Rasmus Orekhov, Vsevolod Ørsøe, Rasmus Ortmann, Tobias Ortmann, Tobias Ortmann, Tobias Ortmann, Tobias Ortmann, Tobias Otrehoff, Jens Osterhoff, Jens Ott, Philipp	T 93.1 EP 1.6 • T 17.2. • E 97.10 ST 9.2 EP 3.1 3.4, T 8.5 3, T 60.4 T 89.4 • T 102.9 KPIK 4.7 4, T 21.5 • DD 50.2 7, T 78.6, 04.7 GR 9.2 5, T 44.6 T 5.8
Olischager, Hans Oliva, André Olivera-Nieto, Laura Olschewsky, Kevin Omeliukh, Anastasiia Omidvari, Negar on Ervision team, VenSpec on VERTIAS team, VEM Öncel, Ö. Oğul Onyisi, Peter Orekhov, Vsevolod Ørsøe, Rasmus Orekhov, Vsevolod Ørsøe, Rasmus Ortmann, Tobias Ortmann, Tobias Ortmann, Tobias Ortmann, Tobias Ortmann, Gulden Ott, Philipp Oughton, Sean	T 93.1 EP 1.6 • T 17.2.7 • EP 7.10 EP 3.1 3.4, T 8.5 3, T 60.4 T 89.4 • T 102.9 KPIK 4.7 4, T 21.5 • DD 50.2 7, T 78.6, 1.4.7 T 5.8 T 5.8 EP 12.3 T 5.8 EP 12.3 T 5.8 EP 12.3 T 5.8 EP 12.3 T 5.8 EP 12.3 T 5.8 EP 12.3 T 5.8 T 5.8
Olischlager, Hans Oliva, André Olivara-Nieto, Laura Olschewsky, Kevin Omeliukh, Anastasiia Omidvari, Negar on Ervision team, VenSpec on VERTIAS team, VEM Öncel, Ö. Oğul Onyisi, Peter Oreklov, Vsevolod Ørsøe, Rasmus Orekhov, Vsevolod Ørsøe, Rasmus Ortmann, Tobias Ortmann, Tobias Ortmann, Tobias Ortmann, Tobias Osterhoff, Jens Osterhoff, Jens Osterhoff, Jens Ott, Philipp Oughton, Sean Ovtscharenko, Nikolaus Ovtscharenko, Nikolaus	T 93.1 EP 1.6 • T 17.2.7 • EP 7.10 ST 9.2 EP 3.1 3.4, T 8.5 3, T 60.4 T 89.4 • T 102.9 KPIK 4.7 4, T 21.5 • DD 50.2 7, T 78.6, 0.4.7 T 5.8 EP 12.3 T 5.8 EP 12.3 T 68.42.6 T 68.7 T
Olischager, Hans Oliva, André Olivera-Nieto, Laura Olschewsky, Kevin Omeliukh, Anastasiia Omidvari, Negar on Envision team, VenSpec on VERTIAS team, VEM Öncel, Ö. Oğul Oryisi, Peter Orekhov, Vsevolod Ørsøe, Rasmus Orekhov, Vsevolod Ørsøe, Rasmus Ortmann, Julia Ortmann, Tobias Ortmann, Tobias Orthof, Jens Osterhoff, Jens Osterhoff, Jens Ott, Philipp Oughton, Sean Ovtscharenko, Nikolaus P2-Kollaboration Paasch, Alexander	T 93.1 EP 1.6 • T 17.2. • EP 7.10 ST 9.2 EP 3.1 3.4, T 8.5 3, T 60.4 T 89.4 • T 102.9 KPIK 4.7 4, T 21.5 • DD 50.2 7, T 78.6, J4.7 GR 9.2 5, T 44.6 T 5.8 EP 12.3 F 12.3 T 68.7 T 68.7 T 68.7 • T 7.2
Olischager, Hans Oliva, André Olivera-Nieto, Laura Olschewsky, Kevin Omeliukh, Anastasiia Omidvari, Negar on Envision team, VenSpec on VERTIAS team, VEM Onyisi, Peter Orekhov, Vsevolod Ørsøe, Rasmus Orekhov, Vsevolod Ørsøe, Rasmus Ortmann, Julia Ortmann, Julia Ortmann, Tobias Otran, Serguei Osterhoff, Jens Ott, Philipp Oughton, Sean Ovtscharenko, Nikolaus P2-Kollaboration Paasch, Alexander Paasch, Steven	T 93.1 EP 1.6 • T 17.2. • E 97.10 ST 9.2 EP 3.1 3.4, T 8.5 3, T 60.4 T 89.4 • T 102.9 KPIK 4.7 4, T 21.5 • DD 50.2 7, T 78.6, 14.7 GR 9.2 5, T 44.6 T 53.8 EP 12.3 F 23.1 • T 7.2 , • T 85.1,
Olischager, Hans Oliva, André Oliva, André Olivara-Nieto, Laura Olschewsky, Kevin Omeliukh, Anastasiia Omidvari, Negar on Envision team, VenSpec on VERTIAS team, VEM Onyisi, Peter Orekhov, Vsevolod Ørsøe, Rasmus Orekhov, Vsevolod Ørsøe, Rasmus Ortmann, Julia Ortmann, Julia Ortmann, Tobias Ottrann, Julia Ortmann, Gulden Ott, Philipp Oughton, Sean Ovtscharenko, Nikolaus P2-Kollaboration Paasch, Alexander Paasch, Steven Data Baiomin	T 93.1 EP 1.6 • T 17.2. • EP 7.10 ST 9.2 EP 3.1 3.4, T 8.5 3, T 60.4 T 89.4 • T 102.9 KPIK 4.7 4, T 21.5 • DD 50.2 7, T 78.6, 14.7 T 53.8 EP 12.3 • T 38.4 T 68.7 • T 78.5, 1, T 85.1, • T 85.1, • T 22.2
Olischager, Hans Oliva, André Oliva, André Olivara-Nieto, Laura Olschewsky, Kevin Omeliukh, Anastasiia Omidvari, Negar on Envision team, VenSpec on VERTIAS team, VEM Onyisi, Peter Orekiov, Vsevolod Ørsøe, Rasmus Orekhov, Vsevolod Ørsøe, Rasmus Ortmann, Julia Ortmann, Julia Ortmann, Tobias Ottrann, Julia Ortmann, Gulden Ott, Philipp Oughton, Sean Ovtscharenko, Nikolaus P2-Kollaboration Paasch, Alexander Paasch, Steven T 85.5 Page, Benjamin Paino. Jason	T 93.1 EP 1.6 T 17.2. T 56.7 EP 7.10 ST 9.2 EP 3.1 3.4, T 8.5 3, T 60.4 T 89.4 T 88.4 T 89.4 T 89.4 T 88.4 T 89.4 T 89.5 T 89.5 T 82.5 T 82.5 T 82.5 T 82.5 T 82.5 T 82.5 T 82.4 T 83.4 T
Olischager, Hans Oliva, André Olivara-Nieto, Laura Olschewsky, Kevin Omeliukh, Anastasiia Omidvari, Negar on Envision team, VenSpec on VERTIAS team, VEM Öncel, Ö. Oğul Tä Onyisi, Peter Orekhov, Vsevolod Ørsøe, Rasmus Orekhov, Vsevolod Ørsøe, Rasmus Ortmann, Julia Ortmann, Julia Ortmann, Tobias T74. T78.7, T 104.5, T 104.6, •T 10 Ossokine, Serguei Osterhoff, Jens Ott, Philipp Oughton, Sean Ovttscharenko, Nikolaus P2-Kollaboration Paasch, Alexander Paasch, Alexander Paasch, Steven T 85.5 Page, Benjamin Paino, Jason Pak, Sang-Il	T 93.1 EP 1.6 T 17.2.7 • EP 7.10 ST 9.2 EP 3.1 3.4, T 8.5 3, T 60.4 T 89.4 • T 102.9 KPIK 4.7 4, T 21.5 • DD 50.2 7, T 78.6, 14.7 GR 9.2 5, T 44.6 T 38.4 T 42.6 T 42.6 T 85.1, T 2.2 ST 3.4 T 9.3
Olischager, Hans Oliva, André Olivera-Nieto, Laura Olschewsky, Kevin Omeliukh, Anastasiia Omidvari, Negar on Envision team, VenSpec on VERTIAS team, VEM Öncel, Ö. Oğul Tä Onyisi, Peter Orekhov, Vsevolod Ørsøe, Rasmus Orekhov, Vsevolod Ørsøe, Rasmus Ortmann, Julia Ortmann, Julia Ortmann, Tobias T74. T78.7, T 104.5, T 104.6, •T 10 Ossokine, Serguei Osterhoff, Jens Ott, Philipp Oughton, Sean Ovttscharenko, Nikolaus P2-Kollaboration Paasch, Alexander Paasch, Alexander Paasch, Steven T 36.7 •T 85.5 Page, Benjamin Paino, Jason Pak, Sang-Il Pal, Divya Oslinka Sergian Ostanov Pase, Sang-Il Pal, Divya Ostanov Ostanov Ostanov Ostanov Categoria Cate	T 93.1 EP 1.6 T 17.2.7 • EP 7.10 ST 9.2 EP 3.1 3.4, T 8.5 3, T 60.4 T 89.4 • T 102.9 KPIK 4.7 4, T 21.5 • DD 50.2 7, T 78.6, 14.7 GR 9.2 5, T 44.6 T 38.4 T 42.6 T 42.6 T 85.1, T 2.2 , • T 85.1, T 2.2 ST 3.4 T 9.3 T 9.3
Olischager, Hans Oliva, André Olivera-Nieto, Laura Olschewsky, Kevin Omeliukh, Anastasiia Omidvari, Negar on Envision team, VenSpec on VERTIAS team, VEM Öncel, Ö. Oğul Tä Onyisi, Peter Orekhov, Vsevolod Ørsøe, Rasmus Orekhov, Vsevolod Ørsøe, Rasmus Ortmann, Julia Ortmann, Julia Ortmann, Tobias T74. T78.7, T 104.5, T 104.6, •T 10 Ossokine, Serguei Osterhoff, Jens Ott, Philipp Oughton, Sean Ovtscharenko, Nikolaus P2-Kollaboration Paasch, Alexander Paasch, Alexander Paasch, Steven Staf, Steven Paasch, Steven Paasc	T 93.1 EP 1.6 T 17.2.7 • EP 7.10 ST 9.2 EP 3.1 3.4, T 8.5 3, T 60.4 T 89.4 • T 102.9 KPIK 4.7 4, T 21.5 • DD 50.2 7, T 78.6, 14.7 GR 9.2 5, T 44.6 T 53.8 EP 12.3 T 42.6 T 85.1, T 2.2 ST 3.4 T 9.2 T 3.2 MP 7.5 MP
Olischlager, Hans Oliva, André Olivera-Nieto, Laura Olschewsky, Kevin Omeliukh, Anastasiia Omidvari, Negar on Envision team, VenSpec on VERTIAS team, VEM Öncel, Ö. Oğul Tä Onyisi, Peter Orekhov, Vsevolod Ørsøe, Rasmus Orekhov, Vsevolod Ørsøe, Rasmus Ortmann, Julia Ortmann, Julia Ortmann, Tobias T74. T78.7, T 104.5, T 104.6, •T 10 Ossokine, Serguei Osterhoff, Jens Ott, Philipp Oughton, Sean Ovtscharenko, Nikolaus P2-Kollaboration Paasch, Alexander Paasch, Alexander Paasch, Alexander Paasch, Steven Palaoino, Jason Paliwal, Ayush Palomo, F. Rogelio	T 93.1 EP 1.6 T 17.2 T 56.7 • EP 7.10 ST 9.2 EP 3.1 3.4, T 8.5 3, T 60.4 T 89.4 • T 102.9 KPIK 4.7 4, T 21.5 • DD 50.2 7, T 78.6, 14.7 GR 9.2 5, T 44.6 T 53.8 EP 12.3 T 42.6 T 85.1 T 2.2 ST 3.4 T 9.2 T 3.2 • MP 7.5 ST 9.4
Olischager, Hans Oliva, André Olivera-Nieto, Laura Olschewsky, Kevin Omeliukh, Anastasiia Omidvari, Negar on Envision team, VenSpec on VERTIAS team, VEM Öncel, Ö. Oğul Tä Onyisi, Peter Orekhov, Vsevolod Ørsøe, Rasmus Orekhov, Vsevolod Ørsøe, Rasmus Ortmann, Julia Ortmann, Julia Ortmann, Tobias T74. T78.7, T 104.5, T 104.6, •T 10 Ossokine, Serguei Osterhoff, Jens Ott, Philipp Oughton, Sean Ovtscharenko, Nikolaus P2-Kollaboration Paasch, Alexander Paasch, Alexander Paasch, Alexander Paasch, Steven Palaoin, Jason Pal, Divya Palomo, F. Rogelio Paneque, David Enterno Paneque, David Description Description Paneque, David Oterno Panegue, David Description Description Panegue, David Description Descri	T 93.1 EP 1.6 T 17.2 T 56.7 EP 3.1 EP 3.1 3.4, T 8.5 3, T 60.4 T 89.4 T 89.4 T 89.4 T 89.4 VT 102.9 KPIK 4.7 4, T 21.5 D 50.2 7, T 78.6, 14.7 GR 9.2 5, T 44.6 T 53.8 EP 12.3 T 42.6 T 85.1, T 2.2 ST 3.4 T 9.2 ST 3.4 T 9.2 ST 3.4 ST 9.2 ST 9.3 ST 9
Olischager, Hans Oliva, André Oliva, André Olivara-Nieto, Laura Olschewsky, Kevin Omeliukh, Anastasiia Omidvari, Negar on Envision team, VenSpec on VERTIAS team, VEM Öncel, Ö. Oğul Tä Onyisi, Peter Orekhov, Vsevolod Ørsøe, Rasmus Orekhov, Vsevolod Ørsøe, Rasmus Ortmann, Julia Ortmann, Julia Ortmann, Tobias T74. T78.7, T 104.5, T 104.6, •T 10 Ossokine, Serguei Osterhoff, Jens Ott, Philipp Oughton, Sean Ovttscharenko, Nikolaus P2-Kollaboration Paasch, Alexander Paasch, Alexander Paasch, Steven Stafe, Sang-II Pal, Divya Palomo, F. Rogelio Pane Queid Pane Devision Pane Devi	T 93.1 EP 1.6 T 17.2 F 56.7 • EP 7.10 ST 9.2 EP 3.1 3.4, T 8.5 3, T 60.4 T 89.4 • T 102.9 KPIK 4.7 4, T 21.5 • DD 50.2 7, T 78.6, 14.7 GR 9.2 5, T 44.6 T 53.8 EP 12.3 T 42.6 T 85.1, T 2.2 ST 3.4 T 9.3 T 9.3 T 2.2 ST 3.4 T 9.3 T 3.2 • MP 7.5 ST 9.3 2, T 83.2 GR 9.2 T 9.3 T 2.2 ST 3.4 T 9.3 T 3.2 • MP 7.5 ST 9.3 ST 9.4 ST 9.5 ST 9.4 ST 9.5 ST 9.4 ST 9.
Olischager, Hans Oliva, André Oliva, André Olivara-Nieto, Laura Olschewsky, Kevin Omeliukh, Anastasiia Omidvari, Negar on Envision team, VenSpec on VERTIAS team, VEM Öncel, Ö. Oğul Tä Onyisi, Peter Orekhov, Vsevolod Ørsøe, Rasmus Orekhov, Vsevolod Ørsøe, Rasmus Orekhov, Vsevolod Ørsøe, Rasmus Ortmann, Julia Ortmann, Julia Ortmann, Tobias T74. T78.7, T 104.5, T 104.6, •T 10 Ossokine, Serguei Osterhoff, Jens Ott, Philipp Oughton, Sean Ovtscharenko, Nikolaus P2-Kollaboration Paasch, Alexander Paasch, Alexander Paasch, Steven Palaions, Jason Pal, Divya Palomo, F. Rogelio Paneque, David Pano, Pano Paneque, David Pano, Pano Pano, Pano Pano Pano, Pano Pano Pano, Pano Pano Pano Pano Pano Pano Pano Pano	T 93.1 EP 1.6 F 16.7 F 56.7 EP 3.1 EP 3.1 3.4, T 8.5 3, T 60.4 T 89.4 T 9.2 T 7.2 S 7.3.4 T 9.2 S 7.3.4 S 7.3.2 G R 9.6 T 6.4 G R 9.4 G R 9.4 S 7.3 S 7
Olischager, Hans Oliva, André Olivara-Nieto, Laura Olschewsky, Kevin Omeliukh, Anastasiia Omidvari, Negar on Envision team, VenSpec on VERTIAS team, VEM Öncel, Ö. Oğul Tä Onyisi, Peter Orekhov, Vsevolod Ørsøe, Rasmus Orekhov, Vsevolod Ørsøe, Rasmus Orekhov, Vsevolod Ørsøe, Rasmus Ortmann, Julia Ortmann, Julia Ortmann, Tobias T74. T78.7, T 104.5, T 104.6, •T 10 Ossokine, Serguei Osterhoff, Jens Ott, Philipp Oughton, Sean Ovttscharenko, Nikolaus P2-Kollaboration Paasch, Alexander Paasch, Alexander Paasch, Alexander Paasch, Steven Palaions, San Paliwal, Ayush Palomo, F. Rogelio Paneque, David Pano, Papa, Maria Alessandra	T 93.1 EP 1.6 F 16.7 F 56.7 EP 3.1 EP 3.1 3.4, T 8.5 3, T 60.4 T 89.4 T 9.2 T 9.2 T 9.2 T 9.2 T 9.2 T 9.2 T 9.2 T 9.2 T 9.2 T 9.4 T 9.2 T 9.4 T 9.2 T 9.4 T 9.2 T 9.2 T 9.4 T 9.2 T 9.4 T 9.2 T 9.4 T 9.2 T 9.4 T 9.4 T 9.5 T
Olischager, Hans Oliva, André Olivara-Nieto, Laura Olschewsky, Kevin Omeliukh, Anastasiia Omidvari, Negar on Envision team, VenSpec on VERTIAS team, VEM Öncel, Ö. Oğul Tra Onyisi, Peter Ordek, Serhat Ortekhov, Vsevolod Ørsøe, Rasmus Ortakov, Vsevolod Ørsøe, Rasmus Ortmann, Hans-Werner Tra Ortmann, Julia Ortmann, Julia Ortmann, Tobias Tr4. Tr8.7, T 104.5, T 104.6, •T 10 Ossokine, Serguei Osterhoff, Jens Otthman, Gulden Ott, Philipp Oughton, Sean Owtscharenko, Nikolaus P2-Kollaboration Paasch, Alexander Paasch, Alexander Paasch, Steven T 39.8, T 68 Palacios Schweitzer, Sofia Paliwal, Ayush Palomo, F. Rogelio Paneque, David Panosso Macedo, Rodrigo Papa, Maria Alessandra Pape, Sebastian Sta 9.	T 93.1 EP 1.6 T 17.2 T 56.7 • EP 7.10 ST 9.2 EP 3.1 3.4, T 8.5 3, T 60.4 T 89.4 • T 102.9 KPIK 4.7 4, T 21.5 • DD 50.2 7, T 78.6, 14.7 GR 9.2 5, T 44.6 T 53.8 EP 12.3 T 42.6 T 85.1 T 2.2 ST 3.4 T 9.2 ST 3.4 T 9.2 ST 9.4 T 9.2 ST 9.4 T 6.4 T 9.2 ST 9.4 ST 9.4 T 6.4 ST 9.2 ST 9.4 ST 9.4 ST 9.4 ST 9.2 ST 9.4 ST 9.4
Olischager, Hans Oliva, André Olivara-Nieto, Laura Olschewsky, Kevin Omeliukh, Anastasiia on Envision team, VenSpec on VERTIAS team, VEM Öncel, Ö. Oğul Tra Ördek, Serhat Ortjohann, Hans-Werner Ortkov, Vsevolod Ørsøe, Rasmus Ortmann, Julia Ortmann, Julia Ortmann, Julia Ortmann, Tobias Tr4. Tr8.7, T104.5, T104.6, •T10 Ossokine, Serguei Osterhoff, Jens Osterhoff, Jens Osterhoff, Jens Osterhoff, Jens Otthman, Gulden Ott, Philipp Oughton, Sean Owtscharenko, Nikolaus P2-Kollaboration Paasch, Alexander Paasch, Alexander Paasch, Steven Pal, Divya Pal, Divya Palono, F. Rogelio Panel de Standar Panel Agelio Panel Scherent Panel Scherent P	T 93.1 EP 1.6 T 17.2 T 56.7 • EP 7.10 ST 9.2 EP 3.1 3.4, T 8.5 3, T 60.4 T 89.4 • T 102.9 KPIK 4.7 4, T 21.5 • DD 50.2 7, T 78.6, 14.7 GR 9.2 5, T 44.6 T 53.8 EP 12.3 T 42.6 T 68.7 T 7.2 ST 3.4 T 9.3 1, T 97.1 ST 3.4 T 9.3 1, T 97.1 ST 3.4 T 3.2 • MP 7.5 ST 3.4 T 68.7 ST 9.3 1, T 97.1 ST 9.2 ST 9.4 GR 9.2 ST 9.4 ST 9.3 T 6.8 T 6.8 T 6.8 T 7.2 ST 3.4 T 9.3 T 9.3
Oliscnager, Hans Oliva, André Olivara-Nieto, Laura Oschewsky, Kevin Omeliukh, Anastasiia on Envision team, VenSpec on VERTIAS team, VEM Öncel, Ö. Öğul Ördek, Serhat Ortjohann, Hans-Werner Ortkov, Vsevolod Ørsøe, Rasmus Ortmann, Julia Ortmann, Julia Ortmann, Julia Ortmann, Julia Ortmann, Tobias T74. T78.7, T104.5, T104.6, •T10 Ossokine, Serguei Osterhoff, Jens Osterhoff, Jens Osterhoff, Jens Otthman, Gulden Ott, Philipp Oughton, Sean Owtscharenko, Nikolaus P2-Kollaboration Paasch, Alexander Paasch, Alexander Paasch, Steven T 39.8, T68. Palacios Schweitzer, Sofia Paliwal, Ayush Palomo, F. Rogelio Panegue, David Panoso Macedo, Rodrigo Papa, Maria Alessandra Pape, Laszlo Darbo, T55. Pape, Laszlo Papa, Laszlo Papa, Laszlo Darbo, T75. Pardos, Carmen Diez	T 93.1 EP 1.6 T 17.2 T 56.7 • EP 7.10 ST 9.2 EP 3.1 3.4, T 8.5 3, T 60.4 T 89.4 • T 102.9 KPIK 4.7 4, T 21.5 • DD 50.2 7, T 78.6, 14.7 GR 9.2 5, T 44.6 T 53.8 EP 12.3 T 42.6 T 68.7 T 7.2 ST 3.4 T 9.2 ST 3.4 T 9.2 ST 3.4 T 9.2 ST 3.4 T 9.2 ST 3.4 T 3.2 GR 9.2 T 3.2 ST 3.4 T 9.2 ST 3.4 T 9.2 ST 9.2 ST 3.4 T 9.2 ST 9.2 ST 9.4 T 64.4 GR 9.2 T 64.4 GR 9.2 T 64.4 GR 9.2 T 64.4 GR 9.2 T 64.4 GR 9.2 T 64.4 GR 9.2 T 9.2 ST 9.4 T 9.3 T 9.2 T 9.2 ST 9.4 T 9.3 T 9.3
Olischarger, Hans Oliva, André Olivara-Nieto, Laura Oschewsky, Kevin Omeliukh, Anastasiia on Envision team, VenSpec on VERTIAS team, VEM Öncel, Ö. Oğul Oridek, Serhat Orekhov, Vsevolod Ørsøe, Rasmus Orthann, Hans-Werner Orthann, Hans-Werner Ortmann, Tobias Ortmann, Tobias Ortmann, Tobias Ortmann, Tobias Ortmann, Tobias Ortmann, Tobias Ortmann, Tobias Ortmann, Sear Osterhoff, Jens Osterhoff, Jens Osterhoff, Jens Oughton, Sean Owtscharenko, Nikolaus P2-Kollaboration Paasch, Alexander Paasch, Alexander Paasch, Steven Pal, Divya Pal, Divya Palono, F. Rogelio Paneque, David Panoso Macedo, Rodrigo Papa, Maria Alessandra Pape, Sebastian Pape, Laszlo Darto, Chiara Pank, George	T 93.1 EP 1.6 T 17.2 T 56.7 • EP 7.10 ST 9.2 EP 3.1 3.4, T 8.5 3, T 60.4 T 89.4 • T 102.9 KPIK 4.7 4, T 21.5 • DD 50.2 7, T 78.6, 14.7 GR 9.2 5, T 44.6 T 53.8 EP 12.3 T 5.8 T 42.6 T 68.7 T 7.2 ST 3.4 T 9.2 ST 3.4 T 9.2 ST 3.4 T 3.2 • MP 7.5 ST 3.4 T 3.2 • MP 7.5 ST 9.2 T 64.4 GR 9.2 T 3.2 • MP 7.5 ST 3.4 T 3.2 • MP 7.5 ST 9.4 • T 9.2 T 9.3 T 3.2 T 9.3 T 3.2 T 9.4 T 3.5 T 9.4 T 9.4
Oliscnager, Hans Oliva, André Olivara-Nieto, Laura Oschewsky, Kevin Omeliukh, Anastasiia on Envision team, VenSpec on VERTIAS team, VEM Öncel, Ö. Öğul Oridek, Serhat Orekhov, Vsevolod Ørsøe, Rasmus Orthann, Hans-Werner Ortmann, Hans-Werner Ortmann, Julia Ortmann, Julia Ortmann, Julia Ortmann, Tobias T74. T78.7, T104.5, T104.6, •T10 Ossokine, Serguei Osterhoff, Jens Otthoff, Jens Otthoff, Jens Otthoff, Jens Otthoff, Jens Oughton, Sean Owtscharenko, Nikolaus P2-Kollaboration Paasch, Alexander Paasch, Alexander Paasch, Steven T 39.8, T68 Palacios Schweitzer, Sofia Paliwal, Ayush Palomo, F. Rogelio Paneque, David Panoso Macedo, Rodrigo Papa, Maria Alessandra Pape, Sebastian Pape, Sebastian Pape, Laszlo Dartof, Katia Paoradi, Katia	T 93.1 EP 1.6 T 17.2 T 56.7 • EP 7.10 ST 9.2 EP 3.1 3.4, T 8.5 3, T 60.4 T 89.4 • T 102.9 KPIK 4.7 4, T 21.5 • DD 50.2 7, T 78.6, 14.7 GR 9.2 5, T 44.6 T 53.8 EP 12.3 T 42.6 T 38.4 T 7.2 ST 3.4 T 9.2 ST 3.4 T 9.2 ST 3.4 T 9.2 ST 3.4 T 3.2 • MP 7.5 ST 9.4 T 3.2 • GR 9.6 T 64.4 GR 9.2 ST 3.4 T 9.2 ST 3.4 T 9.2 ST 9.4 T 9.3 1, T 97.1 ST 9.2 GR 9.6 T 64.4 GR 5.2 3, ST 9.4 • T 93.10 4, T 78.5 T 3.2 T 9.4 T 3.2 T 9.4 T 3.2 T 9.4 T 3.2 T 9.4 T 3.2 T 9.4 T 9.4 T 3.2 T 9.4 T 9.4 T 3.2 T 9.4 T 3.2 T 9.4 T 3.2 T 9.4 T 3.2 T 9.4 T 3.2 T 9.4 T 3.2 T 9.4 T 9.4 T 2.2 T 9.4 T 3.2 T 9.4 T 3.2 T 9.4 T 3.2 T 9.4 T 2.1 T 3.2 T 9.4 T 3.2 T 9.4 T 2.1 T 3.2 T 9.4 T 2.1 T 9.4 T 2.1 T 9.4 T 9.5 T 9.5
Oliscnager, Hans Oliva, André Olivara-Nieto, Laura Oschewsky, Kevin Omeliukh, Anastasiia on Envision team, VenSpec on VERTIAS team, VEM Öncel, Ö. Oğul Ördek, Serhat Ortjohann, Hans-Werner Ortghann, Hans-Werner Ortmann, Julia Ortmann, Julia Ortmann, Tobias T74. T78.7, T104.5, T104.6, •T10 Ossokine, Serguei Osterhoff, Jens Osterhoff, Jens Otthman, Gulden Ott, Philipp Oughton, Sean Oughton, Sean Oughton, Sean Oughton, Sean Oughton, Steven Paasch, Alexander Paasch, Alexander Paasch, Steven Pal, Divya Palacios Schweitzer, Sofia Palacios Schweitzer, Sofia Paladival, Ayush Palomo, F. Rogelio Panoso Macedo, Rodrigo Papa, Maria Alessandra Papp, Laszlo Parodi, Katia Parzefall, Ulrich T13.6, T66. T94.1	T 93.1 EP 1.6 F 16.7 F 56.7 F 56.7 EP 3.1 EP 3.1 3.4, T 8.5 3, T 60.4 T 89.4 T 85.1 T 2.2 ST 3.4 T 9.2 ST 9.4 T 9.2 T 83.2 GR 9.6 T 64.4 ST 9.3 T 99.4 T 33.7 T 99.4 ST 2.1 6, T 66.7,

	Nippel, Kathrin	Paschen, Botho T 95.1, T 95.3, T 95.7
Muddu, Venkata Sai Saketh . •GR 5.4	Nissan, Yuval•T 9.1, T 9.2	Passero, Taimara•AGPhil 1.4
Mueller, AndreasDD 45.2	Nissanke, Samaya•PV VII	Passon, Oliver
Mueller, Stefan E	Nitsche, Hannes•T 52.6	Patil, Abhit•T 25.8
Mühlleitner, Margarete T 85.3, T 92.2,	Nitschke, Anna-Katharina •ST 2.4,	Pattavina, Luca T 74.7, T 78.6, T 78.7,
Т 92.3	ST 3.2	T 104.5, T 104.6, T 104.7
Mujagic, Jasmina	Noethe, Maximilian	Patton, Lydia•AGPhil 4.1
Mukherjee, Swagata • 1 1.3, 1 64.2,	Noethlichs, Michel DD 30.1, •DD 47.1	Patzwahl, Marcel
	Noll, Andreas EP 4.3, EP 4.6	Paul, Stephan 1 58./
Mukherjee, lista •EP II.I, •I 83.I	Noll, Dennis 1 61.5, •1 61.6, 1 61./	Pauls, Alexander 1 65.1, •1 65.2,
Mulder, Mick 14.4	Nonce, Nikids 1 20.0, 1 43.0	1 05.3, 1 05.4, 1 94.0
Müller Andreas DD 10.2 DD 36.1	Nordmeier Volkbard DD 11 2 DD 16 2	Pawiowsky, Jaillis
Müller Avel T 20 3 T 20 9 T 48 6	Noronha lorge MP 2.3	Paitzech Saecha T 2 6
T 48 8 T 102 3	Northe Christian MP12	Pelicci Luca T 20 2 T 20 5 T 43 1
Müller Fabienne DD 31 1	Nöthe for the CTA Consortium	T 48 4 T 103 7 T 103 9
Müller, Jakob	Maximilian	PEN-Kollaboration T 98.4
Müller, Jonas	Nöthe for the CTA LST Project,	Pena, Karla
Müller, Julia•T 50.2	Maximilian	Penek, Ömer
Müller, Martin•T 14.8	Nöthe, Maximilian •AKPIK 1.2,	Penger, SDD 34.3
Müller, Nils EP 3.1	AKPIK 2.1	Penger, SophieDD 44.2
Müller, Rainer DD 6.2, DD 13.2,	Novak, Andrzej T 26.1, T 26.3, •T 89.1,	Penski, Katrin •T 11.1, T 11.2, T 11.3,
DD 14.1, DD 33.2	T 89.1, T 89.2, T 90.6, T 105.10	T 11.4, T 11.5, T 11.6, •T 97.7
Müller, Sara•ST 1.5	Novgorodova, Olga•ST 1.4	Pereira de Lira, Jhonatan T 3.9
Müller, Saskia . ST 8.2, ST 8.3, ST 8.4	Nowotny, Fabian•T 37.2	Pereira Martins, Edyvania Emily
Müller, Thomas . T 7.5, T 65.5, T 65.6,	NUCLEUS-Kollaboration T 74.8	•1 73.4
1 65./, 1 8/.4, 1 8/.5, 1 94.4	Nurnberg, Andreas 1 94.4	Peric, IvanSI 6.4
Müller Gesewisch Jan Ols	U. SIIVA, HECTOF	Perinck, volker
T 65 7 T 01 1	Oberguer Lother T20 9 T 49 7	Pernice Wolfrom T 21 4 U.8
Multi-wavelength collaborators and the	T 48 9 T 48 10 T 00 5	DD 17 1
MAGIC and Fermi-I AT-Kollaboration		Perry Charlotte T 07 6
FP 11 2	Odagiu Patrick T 10 0	Pertl Tobias Andreas
Mund. Jens MP 7 2	Oder Bernd DD 34.4	Pertoldi. Luigi T 75.3 •T 75.6
Munz, Vera	Oechsle. Jan T 94.3	Peter. Stefanie
•DD 12.5. DD 12.6	Oehler, Marie	Petereit. Jelena
Muthukrishnan. Siddharth	Oertel, Jessica •DD 4.2, DD 20.4	Peters. Krisztian GR 14.1. T 10.3.
•AGPhil 6.4	Oeser, Thomas•T 4.7	T 38.1, T 38.7, T 92.6
Muthusamy, GowthamDD 6.2,	Oez, Erdem•T 105.2	Petricca, Federica
DD 38.1	Ogata, Yoshiko MP 1.1	Pfaff, Karl DD 18.1
NA61/SHINE-Kollaboration T 101.9	Ogrissek, Sara DD 41.3	Pfeffer, Emanuel T 88.6, T 88.7,
Naab, Richard	Ohme, Frank•GR 8.1	•T 88.8, T 88.9
Naaijkens, Pieter	Ohtani, YoshikiT 17.8	Pfeifer, BernhardT 39.1
Naber, Jonathan DD 27.2	Oikonomou, Foteini T 71.9	Pfeifer, Christian•GR 11.5
Nachman, Benjamin AKPIK 4.2	Olbrich, Lena•ST 7.2	Pfeiffer, Harald•GR 7.1, GR 13.3
Nachtmann, Otto	Oleksiuk, Ivan T 93.4	Pham, Binh
Nackenhorst, Olaf ST 3.4, T 34.5,	Oleksiyuk, Ivan•T 93.3	Phan, Vo Hong Minh . •EP 6.6, T 73.6
I /U./, I IU/.8	Olivo Andrá ED 16	Philipp, Dennis•GR 2.4, GR 2.6,
Nadol, Lilli•EP 12.10	Olivera Nieto Loura	GR 2.8 Dhilippon Sockia T 09.7
Nagar, Alessandro GR 9.4	Olivera-Inleto, Laura	Plilippen, Saskia
Nagataki Shigohiro ED14	Olechowsky Kovin T 56 7	
Nagataki, ShigehiroEP 1.4	Olschewsky, Kevin	
Nagataki, ShigehiroEP 1.4 Näger, Paul MAGPhil 9.2 namasaka wilson	Olschewsky, Kevin	Piergrossi, Joseph
Nagataki, Shigehiro	Olschewsky, Kevin	Piergrossi, Joseph
Nagataki, Shigehiro	Olschewsky, Kevin	Piergrossi, Joseph
Nagataki, Shigehiro	Olschewsky, Kevin	Piergrossi, Joseph
Nagataki, Shigehiro EP 1.4 Näger, Paul M. •AGPhil 9.2 namasaka, wilson •T 47.6 Nandi, Gerrit DD 5.2 Naranjo, Pedro AGPhil 1.1 Nardi, Roberto GP 2.3 Nardo, Lorenzo ST 9.2	Olschewsky, Kevin	Piergrossi, Joseph
Nagataki, Shigehiro EP 1.4 Näger, Paul M. •AGPhil 9.2 namasaka, wilson •T 47.6 Nandi, Gerrit DD 5.2 Naranjo, Pedro AGPhil 1.1 Nardi, Roberto GP 2.3 Nardo, Lorenzo ST 9.2 Narimani Charan, Abtin •T 79.4	Olschewsky, Kevin	Piergrossi, Joseph
Nagataki, Shigehiro EP 1.4 Näger, Paul M. •AGPhil 9.2 namasaka, wilson •T 47.6 Nandi, Gerrit DD 5.2 Naranjo, Pedro AGPhil 1.1 Nardi, Roberto GP 2.3 Nardo, Lorenzo ST 9.2 Narimani Charan, Abtin •T 79.4 Nasr Esfahani, Alina •T 51.5	Olschewsky, Kevin	 Piergrossi, Joseph
Nagataki, Shigehiro EP 1.4 Näger, Paul M. •AGPhil 9.2 namasaka, wilson •T 47.6 Nandi, Gerrit DD 5.2 Naranjo, Pedro AGPhil 1.1 Nardi, Roberto GP 2.3 Nardo, Lorenzo ST 9.2 Narimani Charan, Abtin •T 79.4 Nasr Esfahani, Alina •T 51.5 Nass, Christian •T 62.3	Olschewsky, Kevin •T 56.7 Omeliukh, Anastasiia •EP 7.10 Omidvari, Negar ST 9.2 on Envision team, VenSpec .EP 3.1 on VERTIAS team, VEM	Piergrossi, Joseph T 24.2 Piergrossi, Joseph T 25.5, T 72.6, T 101.7 Pierre Auger-Kollaboration T 19.1, T 19.2, T 19.4, T 19.7, AKPIK 4.6, AKPIK 4.9, T 71.4, T 19.8, T 19.5, T 98.5, T 71.1, T 71.2, T 71.3, T 51.2, T 19.3, T 51.3, T 47.1, T 47.4, T 47.3, T 51.1, T 47.2 Pietroni, Massimo GR 6.2
Nagataki, Shigehiro EP 1.4 Näger, Paul M. •AGPhil 9.2 namasaka, wilson •T 47.6 Nandi, Gerrit DD 5.2 Naranjo, Pedro AGPhil 1.1 Nardi, Roberto GP 2.3 Nardo, Lorenzo ST 9.2 Narimani Charan, Abtin •T 79.4 Nasr Esfahani, Alina •T 51.5 Nass, Christian •T 62.3 Nastian •T 79.4	Olschewsky, Kevin •T 56.7 Omeliukh, Anastasiia •EP 7.10 Omidvari, Negar ST 9.2 on Envision team, VenSpec .EP 3.1 on VERTIAS team, VEM	Piergrossi, Joseph T 24.2 Piergrossi, Joseph T 25.5, T 72.6, T 101.7 Pierre Auger-Kollaboration T 19.1, T 19.2, T 19.4, T 19.7, AKPIK 4.6, AKPIK 4.9, T 71.4, T 19.8, T 19.5, T 98.5, T 71.1, T 71.2, T 71.3, T 51.2, T 19.3, T 51.3, T 47.1, T 47.4, T 47.3, T 51.1, T 47.2 Pietroni, Massimo GR 6.2 Pilarczyk, Oliver T 48.5
Nagataki, Shigehiro EP 1.4 Näger, Paul M. •AGPhil 9.2 namasaka, wilson •T 47.6 Nandi, Gerrit DD 5.2 Naranjo, Pedro AGPhil 1.1 Nardi, Roberto GP 2.3 Nardo, Lorenzo ST 9.2 Narimani Charan, Abtin •T 79.4 Nass Esfahani, Alina •T 51.5 Nast, Natalie T 107.4 Nattland, Philipp T 33.2	Olschewsky, Kevin I 56.7 Omeliukh, Anastasiia I P7.10 Omidvari, Negar ST 9.2 on Envision team, VenSpec P 3.1 on VERTIAS team, VEM EP 3.1 Öncel, Ö. Oğul T 8.4, T 8.5 Onyisi, Peter T 60.3, T 60.4 Ördek, Serhat T 89.4 Orekhov, Vsevolod I T 102.9 Ørsøe, Rasmus AKPIK 4.7 Ortjohann, Hans-Werner T 21.4, T 21.5 Ortmann, Julia D 50.2	Piergrossi, Joseph
Nagataki, Shigehiro EP 1.4 Näger, Paul M. •AGPhil 9.2 namasaka, wilson •T 47.6 Nandi, Gerrit DD 5.2 Naranjo, Pedro AGPhil 1.1 Nardi, Roberto GP 2.3 Nardo, Lorenzo ST 9.2 Narimani Charan, Abtin •T 79.4 Nast, Schristian •T 62.3 Nast, Natalie T 107.4 Nattland, Philipp T 33.2 Navarro, Jaume •GP 6.2	Olschewsky, Kevin • T 56.7 Omeliukh, Anastasiia • EP 7.10 Omidvari, Negar ST 9.2 on Envision team, VenSpec EP 3.1 on VERTIAS team, VEM EP 3.1 Öncel, Ö. Oğul T 8.4, T 8.5 Onjsi, Peter T 60.3, T 60.4 Ördek, Serhat T 89.4 Orekhov, Vsevolod • T 102.9 Ørsøe, Rasmus • AKPIK 4.7 Ortjohann, Hans-Werner T 21.4, T 21.5 Ortmann, Julia • DD 50.2 Ortmann, Tobias T 74.7, T 78.6,	Piergrossi, Joseph
Nagataki, Shigehiro EP 1.4 Näger, Paul M. •AGPhil 9.2 namasaka, wilson •T 47.6 Nandi, Gerrit DD 5.2 Naranjo, Pedro AGPhil 1.1 Nardi, Roberto GP 2.3 Nardo, Lorenzo ST 9.2 Narimani Charan, Abtin •T 79.4 Nasr Esfahani, Alina •T 51.5 Nass, Christian •T 62.3 Nast, Natalie T 107.4 Nattland, Philipp T 33.2 Navarro, Jaume •GP 6.2 Navordt, Ronny DD 52.2 DD 12.7, DD 24.1 DD 24.4 12 DD 26.5 CD 44.2	Olschewsky, Kevin • T 56.7 Omeliukh, Anastasiia • EP 7.10 Omidvari, Negar ST 9.2 on Envision team, VenSpec EP 3.1 on VERTIAS team, VEM EP 3.1 Öncel, Ö. Oğul T 8.4, T 8.5 Onyisi, Peter T 60.3, T 60.4 Ördek, Serhat T 89.4 Orekhov, Vsevolod • T 102.9 Ørsøe, Rasmus • AKPIK 4.7 Ortjohann, Hans-Werner T 21.4, T 21.5 Ortmann, Julia • DD 50.2 Ortmann, Tobias T 74.7, T 78.6, T 78.7, T 104.5, T 104.6, • T 104.7 Ores 2	Piergrossi, Joseph
Nagataki, Shigehiro EP 1.4 Näger, Paul M. •AGPhil 9.2 namasaka, wilson •T 47.6 Nandi, Gerrit DD 5.2 Naranjo, Pedro AGPhil 1.1 Nardi, Roberto GP 2.3 Nardo, Lorenzo ST 9.2 Narimani Charan, Abtin •T 79.4 Nasr Esfahani, Alina •T 62.3 Nast, Natalie T 107.4 Nattland, Philipp T 33.2 Navaro, Jaume •GP 6.2 Nawrot, Ronny DD 5.2, DD 12.7, DD 31.1, DD 34.12, DD 36.5, DD 44.3 T 16.4	Olschewsky, Kevin	Piergrossi, Joseph
Nagataki, Shigehiro EP 1.4 Näger, Paul M. •AGPhil 9.2 namasaka, wilson •T 47.6 Nandi, Gerrit DD 5.2 Naranjo, Pedro AGPhil 1.1 Nardi, Roberto GP 2.3 Nardo, Lorenzo ST 9.2 Narimani Charan, Abtin •T 79.4 Nasr Esfahani, Alina •T 51.5 Nass, Christian •T 62.3 Nattland, Philipp T 33.2 Navarro, Jaume •GP 6.2 Nawordt, Ronny DD 5.2, DD 12.7, DD 31.1, DD 34.12, DD 36.5, DD 44.3 Nayaz, Abdullah Nayaz, Abdullah •T 16.4	Olschewsky, Kevin	Piergrossi, Joseph T 24.2 Piergog, Tanguy T 72.5, T 72.6, T 101.7 Pierre Auger-Kollaboration T 19.1, T 19.2, T 19.4, T 19.7, AKPIK 4.6, AKPIK 4.9, T 71.4, T 19.8, T 19.5, T 98.5, T 71.1, T 71.2, T 71.3, T 51.2, T 19.3, T 51.3, T 47.1, T 47.4, T 47.3, T 51.1, T 47.2 Pietroni, Massimo GR 6.2 Pilarczyk, Oliver T 48.5 Pimpalshende, Vivek GR 14.4, GR 14.5, •GR 14.6 Pinamonti, Nicola •GR 12.2, •MP 9.2 Piotter, Mona T 22.7 Pirnen, Pekka D 36.1
Nagataki, Shigehiro EP 1.4 Näger, Paul M. •AGPhil 9.2 namasaka, wilson •T 47.6 Nandi, Gerrit DD 5.2 Naranjo, Pedro AGPhil 1.1 Nardi, Roberto GP 2.3 Nardo, Lorenzo ST 9.2 Narimani Charan, Abtin •T 79.4 Nasr Esfahani, Alina •T 51.5 Nass, Christian •T 62.3 Nast, Natalie T 107.4 Natland, Philipp T 33.2 Navarro, Jaume •GP 6.2 Navarot, Ronny DD 5.2, DD 12.7, DD 31.1, DD 34.12, DD 36.5, DD 44.3 Nayaz, Abdullah Nayaz, Abdullah •T 18.2 Neekora, Vsevolord •GP 11.3	Olschewsky, Kevin •T 56.7 Omeliukh, Anastasiia •EP 7.10 Omidvari, Negar ST 9.2 on Envision team, VenSpec .EP 3.1 on VERTIAS team, VEM	Piergrossi, Joseph
Nagataki, Shigehiro EP 1.4 Näger, Paul M. •AGPhil 9.2 namasaka, wilson •T 47.6 Nandi, Gerrit DD 5.2 Naranjo, Pedro AGPhil 1.1 Nardi, Roberto GP 2.3 Narimani Charan, Abtin •T 79.4 Nasr Esfahani, Alina •T 51.5 Nass, Christian •T 62.3 Nast, Natalie T 107.4 Nattland, Philipp T 33.2 Navarro, Jaume •GP 6.2 Navarod, Lorenzo D 5.2, DD 12.7, DD 31.1, DD 34.12, DD 36.5, DD 44.3 Nayaz, Abdullah Necker, Jannis •T 16.4 Necker, Jannis •T 18.2 Nedora, Vsevolod •GR 11.3	Olschewsky, Kevin • T 56.7 Omeliukh, Anastasiia • EP 7.10 Omidvari, Negar ST 9.2 on Envision team, VenSpec EP 3.1 on VERTIAS team, VEM EP 3.1 Öncel, Ö. Oğul T 8.4, T 8.5 Onjsis, Peter T 60.3, T 60.4 Ördek, Serhat T 89.4 Orekhov, Vsevolod • T 102.9 Ørsøe, Rasmus • AKPIK 4.7 Ortjohann, Hans-Werner T 21.4, T 21.5 Ortmann, Julia • DD 50.2 Ortmann, Tobias T 74.7, T 78.6, T 78.7, T 104.5, T 104.6, • T 104.7 Ossokine, Serguei Ossokine, Serguei G 9.2 Osterhoff, Jens T 44.5, T 44.6 Othman, Gulden • T 38.4 Ott, Philipp T 5.8 Oughton, Sean FP 12.3	Piergrossi, Joseph
Nagataki, Shigehiro EP 1.4 Näger, Paul M. •AGPhil 9.2 namasaka, wilson •T 47.6 Nandi, Gerrit DD 5.2 Naranjo, Pedro AGPhil 1.1 Nardi, Roberto GP 2.3 Nardo, Lorenzo ST 9.2 Narimani Charan, Abtin •T 79.4 Nasr Esfahani, Alina •T 51.5 Nass, Christian •T 62.3 Nattland, Philipp T 33.2 Navarro, Jaume •GP 6.2 Nawrodt, Ronny DD 5.2, DD 12.7, DD 31.1, DD 34.12, DD 36.5, DD 44.3 Nayaz, Abdullah •T 18.2 Nedker, Jannis •T 18.2 Nedgar, Farjoud Masouleh EP 2.2 Nell Sebastian DD 12.8	Olschewsky, Kevin •• T 56.7 Omeliukh, Anastasiia •EP 7.10 Omidvari, Negar ST 9.2 on Envision team, VenSpec EP 3.1 on VERTIAS team, VEM EP 3.1 Öncel, Ö. Oğul T 8.4, T 8.5 Onyisi, Peter T 60.3, T 60.4 Ördek, Serhat T 89.4 Orekhov, Vsevolod •T 102.9 Ørsøe, Rasmus •AKPIK 4.7 Ortjohann, Hans-Werner T 21.4, T 21.5 Ortmann, Julia •D 50.2 Ortmann, Tobias T 74.7, T 78.6, T 78.7, T 104.5, T 104.6, •T 104.7 Ossokine, Serguei Ossokine, Serguei GR 9.2 Osterhoff, Jens T 44.5, T 44.6 Othman, Gulden •T 38.4 Ott, Philipp T 5.8 Oughton, Sean EP 12.3 Owtscharenko, Nikolaus •T 42 6	Piergrossi, Joseph
Nagataki, Shigehiro EP 1.4 Näger, Paul M. •AGPhil 9.2 namasaka, wilson •T 47.6 Nandi, Gerrit DD 5.2 Naranjo, Pedro AGPhil 1.1 Nardi, Roberto GP 2.3 Nardo, Lorenzo ST 9.2 Narimani Charan, Abtin •T 79.4 Nasr Esfahani, Alina •T 51.5 Nass, Christian •T 62.3 Nast, Natalie T 107.4 Nattland, Philipp T 33.2 Navarro, Jaume •GP 6.2 Nawrodt, Ronny DD 5.2, DD 12.7, DD 31.1, DD 34.12, DD 36.5, DD 44.3 Nayaz, Abdullah •T 16.4 Necker, Jannis •T 18.2 Nedora, Vsevolod •GR 11.3 Negar, Farjoud Masouleh EP 2.2 Nell, Sebastian DD 12.8 Nemeth-Csoka, Marton •T 30.5	Olschewsky, Kevin •• T 56.7 Omeliukh, Anastasiia •EP 7.10 Omidvari, Negar ST 9.2 on Envision team, VenSpec EP 3.1 on VERTIAS team, VEM EP 3.1 Öncel, Ö. Oğul T 8.4, T 8.5 Onyisi, Peter T 60.3, T 60.4 Ördek, Serhat T 89.4 Orekhov, Vsevolod •T 102.9 Ørsøe, Rasmus •AKPIK 4.7 Ortjohann, Hans-Werner T 21.4, T 21.5 Ortmann, Julia •D 50.2 Ortmann, Tobias T 74.7, T 78.6, T 78.7, T 104.5, T 104.6, •T 104.7 Ossokine, Serguei Osterhoff, Jens T 44.5, T 44.6 Othman, Gulden •T 38.4 Ott, Philipp T 5.8 Oughton, Sean EP 12.3 Owtscharenko, Nikolaus •T 42.6 P2-Kollaboration T 68 7	Piergrossi, Joseph
Nagataki, Shigehiro EP 1.4 Näger, Paul M. •AGPhil 9.2 namasaka, wilson •T 47.6 Nandi, Gerrit DD 5.2 Naranjo, Pedro AGPhil 1.1 Nardi, Roberto GP 2.3 Nardo, Lorenzo ST 9.2 Narimani Charan, Abtin •T 79.4 Nasr Esfahani, Alina •T 51.5 Nass, Christian •T 62.3 Nattland, Philipp T 33.2 Navaro, Jaume •GP 6.2 Nawrodt, Ronny DD 5.2, DD 12.7, DD 31.1, DD 34.12, DD 36.5, DD 44.3 Nayaz, Abdullah •T 16.4 Necker, Jannis •T 18.2 Nedora, Vsevolod •GR 11.3 Negar, Farjoud Masouleh EP 2.2 Nell, Sebastian DD 12.8 Nemeth-Csoka, Marton •T 30.5	Olschewsky, Kevin •• T 56.7 Omeliukh, Anastasiia •• EP 7.10 Omidvari, Negar ST 9.2 on Envision team, VenSpec .EP 3.1 on VERTIAS team, VEM	Piergrossi, Joseph T 24.2 Piergo, Tanguy T 72.5, T 72.6, T 101.7 Pierre Auger-Kollaboration T 19.1, T 19.2, T 19.4, T 19.7, AKPIK 4.6, AKPIK 4.9, T 71.4, T 19.8, T 19.5, T 98.5, T 71.1, T 71.2, T 71.3, T 51.2, T 19.3, T 51.3, T 47.1, T 47.4, T 47.3, T 51.1, T 47.2 Pietroni, Massimo GR 14.5, •GR 14.6 Pinamonti, Nicola •GR 12.2, •MP 9.2 Piotter, Mona T 22.7 Pirinen, Pekka D 36.1 Pirrer, Hans Jürgen •AGPhil 12.3 Pirsch, Johannes •MP 2.6 Piscopo, Maria Laura T 56.4 Pittermann. Martin •ST 6.4
Nagataki, Shigehiro EP 1.4 Näger, Paul M. •AGPhil 9.2 namasaka, wilson •T 47.6 Nandi, Gerrit DD 5.2 Naranjo, Pedro AGPhil 1.1 Nardi, Roberto GP 2.3 Nardo, Lorenzo ST 9.2 Narimani Charan, Abtin •T 79.4 Nasr Esfahani, Alina •T 51.5 Nass, Christian •T 62.3 Nast, Natalie T 107.4 Nattland, Philipp T 33.2 Navarro, Jaume •GP 6.2 Nawrodt, Ronny DD 5.2, DD 12.7, DD 31.1, DD 34.12, DD 36.5, DD 44.3 Nayaz, Abdullah •T 16.4 Nedker, Jannis •T 18.2 Nedar, Vsevolod •GR 11.3 Negar, Farjoud Masouleh EP 2.2 Nell, Sebastian DD 12.8 Netwerk Teilchenwelt-Kollaboration T 30.5 Netzwerk Teilchenwelt-Kollaboration T 24.3, T 24.1, T 52.1, T 52.4, T 24.2	Olschewsky, Kevin • T 56.7 Omeliukh, Anastasiia • EP 7.10 Omidvari, Negar ST 9.2 on Envision team, VenSpec EP 3.1 on VERTIAS team, VEM EP 3.1 Öncel, Ö. Oğul T 8.4, T 8.5 Onjsi, Peter T 60.3, T 60.4 Ördek, Serhat T 89.4 Orekhov, Vsevolod • T 102.9 Ørsøe, Rasmus • AKPIK 4.7 Ortjohann, Hans-Werner 7 21.4, T 21.5 Ortmann, Julia • DD 50.2 Ortmann, Tobias T 74.7, T 78.6, T 78.7, T 104.5, T 104.6, • T 104.7 Ossokine, Serguei Osterhoff, Jens T 44.5, T 44.6 Othman, Gulden • T 38.4 Ott, Philipp T 5.8 Oughton, Sean EP 12.3 Owtscharenko, Nikolaus • T 42.6 P2-Kollaboration T 68.7 Paasch, Alexander • T 7.2 Paasch, Steven T 36.7, • T 85.1	Piergrossi, Joseph T 24.2 Piergo, Tanguy T 72.5, T 72.6, T 101.7 Pierre Auger-Kollaboration T 19.1, T 19.2, T 19.4, T 19.7, AKPIK 4.6, AKPIK 4.9, T 71.4, T 19.8, T 19.5, T 98.5, T 71.1, T 71.2, T 71.3, T 51.2, T 19.3, T 51.3, T 47.1, T 47.4, T 47.3, T 51.1, T 47.2 Pietroni, Massimo GR 14.5, •GR 14.6 Pinamonti, Nicola •GR 12.2, •MP 9.2 Piotter, Mona •T 22.7 Pirinen, Pekka DD 36.1 Pirner, Hans Jürgen •AGPhil 12.3 Pirola, Giorgio •T 17.8 Pirscop, Maria Laura T 56.4 Pitt, Michael T 7.2, T 2.8, T 12.9, T 30.8
Nagataki, Shigehiro EP 1.4 Näger, Paul M. •AGPhil 9.2 namasaka, wilson •T 47.6 Nandi, Gerrit DD 5.2 Naranjo, Pedro AGPhil 1.1 Nardi, Roberto GP 2.3 Nardo, Lorenzo ST 9.2 Narimani Charan, Abtin •T 79.4 Nasr Esfahani, Alina •T 51.5 Nass, Christian •T 62.3 Nast, Natalie T 107.4 Nattland, Philipp T 33.2 Navarro, Jaume •GP 6.2 Nawordt, Ronny D 5.2, DD 12.7, D 31.1, DD 34.12, DD 36.5, DD 44.3 Nayaz, Abdullah •T 16.4 Necker, Jannis •T 18.2 Nedora, Vsevolod •GR 11.3 Negar, Farjoud Masouleh EP 2.2 Nemeth-Csoka, Marton •T 30.5 Netzwerk Teilchenwelt-Kollaboration T 24.3, T 24.1, T 52.1, T 52.4, T 24.2, T 52.3	Olschewsky, Kevin •• T 56.7 Omeliukh, Anastasiia •• EP 7.10 Omidvari, Negar ST 9.2 on Envision team, VenSpec EP 3.1 on VERTIAS team, VEM EP 3.1 Öncel, Ö. Oğul T 8.4, T 8.5 Onyisi, Peter T 60.3, T 60.4 Ördek, Serhat T 89.4 Orekhov, Vsevolod •T 102.9 Ørsøe, Rasmus •AKPIK 4.7 Ortjohann, Hans-Werner T 21.4, T 21.5 Ortmann, Julia •DD 50.2 Ortmann, Tobias T 74.7, T 78.6, T 78.7, T 104.5, T 104.6, •T 104.7 Ossokine, Serguei Ossokine, Serguei GR 9.2 Osterhoff, Jens T 44.5, T 44.6 Othman, Gulden •T 38.4 Ott, Philipp T 5.8 Oughton, Sean EP 12.3 Owtscharenko, Nikolaus T 42.6 P2-Kollaboration T 68.7 Paasch, Alexander T 7.2 Paasch, Alexander T 36.7, •T 85.1,	Piergrossi, Joseph
Nagataki, Shigehiro EP 1.4 Näger, Paul M. •AGPhil 9.2 namasaka, wilson •T 47.6 Nandi, Gerrit DD 5.2 Naranjo, Pedro AGPhil 1.1 Nardi, Roberto GP 2.3 Nardo, Lorenzo ST 9.2 Narimani Charan, Abtin •T 79.4 Nasr Esfahani, Alina •T 51.5 Nass, Christian •T 62.3 Nast, Natalie T 107.4 Nattland, Philipp T 33.2 Navarro, Jaume •GP 6.2 Nawrodt, Ronny DD 5.2, DD 12.7, DD 31.1, DD 34.12, DD 36.5, DD 44.3 Nayaz, Abdullah •T 18.2 Nedra, Vsevolod •GR 11.3 Negar, Farjoud Masouleh EP 2.2 Nell, Sebastian DD 12.8 Nemeth-Csoka, Marton •T 30.5 Netzwerk Teilchenwelt-Kollaboration T 24.3, T 24.1, T 52.1, T 52.4, T 24.2, T 52.3 Yet.3 Neuberger, Moritz •T 75.3	Olschewsky, Kevin •• T 56.7 Omeliukh, Anastasiia •EP 7.10 Omidvari, Negar ST 9.2 on Envision team, VenSpec EP 3.1 on VERTIAS team, VEM EP 3.1 Öncel, Ö. Oğul T 8.4, T 8.5 Onyisi, Peter T 60.3, T 60.4 Ördek, Serhat T 89.4 Orekhov, Vsevolod •T 102.9 Ørsøe, Rasmus •AKPIK 4.7 Ortjohann, Hans-Werner T 21.4, T 21.5 Ortmann, Julia •D 50.2 Ortmann, Tobias T 74.7, T 78.6, T 78.7, T 104.5, T 104.6, •T 104.7 Ossokine, Serguei Ossokine, Serguei GR 9.2 Osterhoff, Jens T 44.5, T 44.6 Othman, Gulden •T 38.4 Ott, Philipp T 5.8 Oughton, Sean EP 12.3 Owtscharenko, Nikolaus •T 42.6 P2-Kollaboration T 68.7 Paasch, Alexander •T 7.2 Paasch, Alexander •T 7.2 Paasch, Steven T 36.7, •T 85.1, •T 85.5 Page, Benjamin	Piergrossi, Joseph T 24.2 Piergo, Tanguy T 72.5, T 72.6, T 101.7 Pierre Auger-Kollaboration T 19.1, T 19.2, T 19.4, T 19.7, AKPIK 4.6, AKPIK 4.9, T 71.4, T 19.8, T 19.5, T 98.5, T 71.1, T 71.2, T 71.3, T 51.2, T 19.3, T 51.3, T 47.1, T 47.4, T 47.3, T 51.1, T 47.2 Pietroni, Massimo GR 14.5, •GR 14.6 Pinamonti, Nicola •GR 12.2, •MP 9.2 Piotter, Mona -T 22.7 Pirinen, Pekka DD 36.1 Pirrer, Hans Jürgen •AGPhil 12.3 Pirola, Giorgio -T 17.8 Pisch, Johannes •MP 2.6 Pistch, Johannes -MP 2.6 Pittermann, Martin •ST 6.4 Pitz, Joniel T 12.8, T 12.9, T 30.8 Pivovarov, Alexei A. T 56.5 Pizzella, Veronica T 78.3
Nagataki, Shigehiro EP 1.4 Näger, Paul M. •AGPhil 9.2 namasaka, wilson •T 47.6 Nandi, Gerrit DD 5.2 Naranjo, Pedro AGPhil 1.1 Nardi, Roberto GP 2.3 Nardo, Lorenzo ST 9.2 Narimani Charan, Abtin •T 79.4 Nasr Esfahani, Alina •T 51.5 Nass, Christian •T 62.3 Nattland, Philipp T 33.2 Navarro, Jaume •GP 6.2 Nawrodt, Ronny .DD 5.2, DD 12.7, DD 31.1, DD 34.12, DD 36.5, DD 44.3 Nayaz, Abdullah •T 16.4 Necker, Jannis •T 18.2 Nedora, Vsevolod •GR 11.3 Negar, Farjoud Masouleh EP 2.2 Nell, Sebastian DD 12.8 Nemeth-Csoka, Marton •T 30.5 Netzwerk Teilchenwelt-Kollaboration T 24.3, T 24.1, T 52.1, T 52.4, T 24.2, T 52.3 Neuberger, Moritz •T 75.3 Neuberger, Moritz •T 75.3	Olschewsky, Kevin •• T 56.7 Omeliukh, Anastasiia •EP 7.10 Omidvari, Negar ST 9.2 on Envision team, VenSpec EP 3.1 on VERTIAS team, VEM EP 3.1 Öncel, Ö. Oğul T 8.4, T 8.5 Onyisi, Peter T 60.3, T 60.4 Ördek, Serhat T 89.4 Orekhov, Vsevolod •T 102.9 Ørsøe, Rasmus •AKPIK 4.7 Ortjohann, Hans-Werner T 21.4, T 21.5 Ortmann, Julia •D 50.2 Ortmann, Jolia •D 50.2 Ortmann, Julia •D 50.2 Ortmann, Julia •D 50.2 Osterhoff, Jens T 44.5, T 44.6 Othman, Gulden •T 38.4 Ott, Philipp T 5.8 Oughton, Sean EP 12.3 Owtscharenko, Nikolaus •T 42.6 P2-Kollaboration T 68.7 Paasch, Alexander •T 7.2 Paasch, Steven T 36.7, •T 85.1, •T 85.5 Page, Benjamin T 2.2 Paino, Jason ST 3.4	Piergrossi, Joseph T 24.2 Piergossi, Joseph T 24.2 Piergossi, Joseph T 25, T 72.6, T 101.7 Pierre Auger-Kollaboration T 19.1, T 19.2, T 19.4, T 19.7, AKPIK 4.6, AKPIK 4.9, T 71.4, T 19.8, T 19.5, T 98.5, T 71.1, T 71.2, T 71.3, T 51.2, T 19.3, T 51.3, T 47.1, T 47.4, T 47.3, T 51.1, T 47.2 Pietroni, Massimo GR 6.2 Pilarczyk, Oliver T 48.5 Pimpalshende, Vivek GR 14.4, GR 14.5, •GR 14.6 Pinamonti, Nicola •GR 12.2, •MP 9.2 Piotter, Mona T 22.7 Pirnen, Pekka DD 36.1 Pirner, Hans Jürgen •AGPhil 12.3 Pirola, Giorgio •T 17.8 Pirsch, Johannes •MP 2.6 Pistch Michael T 7.5 Pittermann, Martin •ST 6.4 Pitz, Daniel T 12.8, T 12.9, T 30.8 Pivovarov, Alexei A. T 56.5 Pizzella, Veronica T 78.3 Planinic, Maja DD 18.2
Nagataki, Shigehiro EP 1.4 Näger, Paul M. •AGPhil 9.2 namasaka, wilson •T 47.6 Nandi, Gerrit DD 5.2 Naranjo, Pedro AGPhil 1.1 Nardi, Roberto GP 2.3 Nardo, Lorenzo ST 9.2 Narimani Charan, Abtin •T 79.4 Nasr Esfahani, Alina •T 51.5 Nass, Christian •T 62.3 Nattland, Philipp T 33.2 Navaro, Jaume •GP 6.2 Nawrodt, Ronny DD 5.2, DD 12.7, DD 31.1, DD 34.12, DD 36.5, DD 44.3 Nayaz, Abdullah •T 16.4 Necker, Jannis •T 18.2 Nedora, Vsevolod •GR 11.3 Negar, Farjoud Masouleh EP 2.2 Nell, Sebastian DD 12.8 Nemeth-Csoka, Marton •T 30.5 Netzwerk Teilchenwelt-Kollaboration T 24.3, T 24.1, T 52.1, T 52.4, T 24.2, T 53.3 Neuberger, Moritz •T 75.3 Neuberger, Moritz •T 75.3 Neuberger, Moritz •T 65.5, T 65.6,	Olschewsky, Kevin •• T 56.7 Omeliukh, Anastasiia •EP 7.10 Omidvari, Negar ST 9.2 on Envision team, VenSpec .EP 3.1 on VERTIAS team, VEM .EP 3.1 Öncel, Ö. Öğul T 8.4, T 8.5 Onyisi, Peter	Piergrossi, Joseph T 24.2 Piergossi, Joseph T 24.2 Piergo, Tanguy T 72.5, T 72.6, T 101.7 Pierre Auger-Kollaboration T 19.1, T 19.2, T 19.4, T 19.7, AKPIK 4.6, AKPIK 4.9, T 71.4, T 19.8, T 19.5, T 98.5, T 71.1, T 71.2, T 71.3, T 51.2, T 19.3, T 51.3, T 47.1, T 47.4, T 47.3, T 51.1, T 47.2 Pietroni, Massimo GR 6.2 Pilarczyk, Oliver •T 48.5 Pimpalshende, Vivek GR 14.4, GR 14.5, •GR 14.6 Pinamonti, Nicola •GR 12.2, •MP 9.2 Piotter, Mona •T 22.7 Pirinen, Pekka D 36.1 Pirner, Hans Jürgen •AGPhil 12.3 Pirsch, Johannes •MP 2.6 Piscopo, Maria Laura T 56.4 Pitt, Michael T 7.5 Pitzella, Veronica •T 24.2 Pivovarov, Alexei A. T 56.4 Pitz, Daniel T 12.8, T 12.9, T 30.8 Pivovarov, Alexei A. T 56.5 Pizzella, Veronica •T 78.3 Planinic, Maja DD 18.2 Plaschke, Ferdinand EP 3.4, EP 3.5
Nagataki, Shigehiro EP 1.4 Näger, Paul M. •AGPhil 9.2 namasaka, wilson •T 47.6 Nandi, Gerrit DD 5.2 Naranjo, Pedro AGPhil 9.2 Naranjo, Pedro AGPhil 1.1 Nardi, Roberto GP 2.3 Nardo, Lorenzo ST 9.2 Narimani Charan, Abtin •T 79.4 Nasr Esfahani, Alina •T 51.5 Nass, Christian •T 62.3 Nattland, Philipp T 33.2 Navaro, Jaume •GP 6.2 Nawrodt, Ronny DD 5.2, DD 12.7, DD 31.1, DD 34.12, DD 36.5, DD 44.3 Nayaz, Abdullah •T 16.4 Necker, Jannis •T 18.2 Nedora, Vsevolod •GR 11.3 Negar, Farjoud Masouleh EP 2.2 Nell, Sebastian DD 12.8 Nemeth-Csoka, Marton •T 30.5 Netzwerk Teilchenwelt-Kollaboration T 24.3, T 24.1, T 52.1, T 52.4, T 24.2, T 53.3 Neuberger, Moritz •T 75.3 Neuberger, Moritz •T 65.5, T 65.6, T 65.7, T 94.4	Olschewsky, Kevin •• T 56.7 Omeliukh, Anastasiia •• EP 7.10 Omidvari, Negar ST 9.2 on Envision team, VenSpec .EP 3.1 on VERTIAS team, VEM EP 3.1 Öncel, Ö. Oğul T 8.4, T 8.5 Onsij, Peter T 60.3, T 60.4 Ördek, Serhat T 8.4, T 8.5 Orekhov, Vsevolod •T 102.9 Ørsøe, Rasmus •AKPIK 4.7 Ortjohann, Hans-Werner T 21.4, T 21.5 Ortmann, Julia DD 50.2 Ortmann, Tobias T 74.7, T 78.6, T 78.7, T 104.5, T 104.6, •T 104.7 Ossokine, Serguei G 9.2 Osterhoff, Jens T 44.5, T 44.6 Otthman, Gulden •T 38.4 Ott, Philipp T 5.8 Oughton, Sean EP 12.3 Owtscharenko, Nikolaus T 42.6 P2-Kollaboration T 68.7 Paasch, Alexander T 7.2 Paasch, Alexander T 7.2 Paasch, Steven T 36.7, •T 85.1, •T 85.5 Page, Benjamin T 2.2 Paino, Jason ST 3.4 Pak, Sang-II	Piergrossi, Joseph
Nagataki, Shigehiro EP 1.4 Näger, Paul M. •AGPhil 9.2 namasaka, wilson •T 47.6 Nandi, Gerrit DD 5.2 Naranjo, Pedro AGPhil 1.1 Nardi, Roberto GP 2.3 Nardo, Lorenzo ST 9.2 Narimani Charan, Abtin •T 79.4 Nasr Esfahani, Alina •T 51.5 Nass, Christian •T 62.3 Navarro, Jaume •GP 6.2 Nawrodt, Ronny DD 5.2, DD 12.7, D 31.1, DD 34.12, DD 36.5, DD 44.3 Nayaz, Abdullah •T 16.4 Necker, Jannis •T 18.2 Nedora, Vsevolod •GR 11.3 Negar, Farjoud Masouleh EP 2.2 Nemeth-Csoka, Marton •T 30.5 Netzwerk Teilchenwelt-Kollaboration T 24.3, T 24.1, T 52.1, T 52.4, T 24.2, T 52.3 Neuberger, Moritz •T 75.3 Neuendorffer, Stephen T 69.7 Neufled, Marius T 65.5, T 65.6, T 65.7, T 94.4	Olschewsky, Kevin •• T 56.7 Omeliukh, Anastasiia •• EP 7.10 Omidvari, Negar ST 9.2 on Envision team, VenSpec EP 3.1 on VERTIAS team, VEM EP 3.1 Öncel, Ö. Oğul T 8.4, T 8.5 Ongisi, Peter T 60.3, T 60.4 Ördek, Serhat T 89.4 Orekhov, Vsevolod •T 102.9 Ørsøe, Rasmus •AKPIK 4.7 Ortjohann, Hans-Werner T 21.4, T 21.5 Ortmann, Julia •DD 50.2 Ortmann, Tobias T 74.7, T 78.6, T 78.7, T 104.5, T 104.6, •T 104.7 Ossokine, Serguei Ossokine, Serguei GR 9.2 Osterhoff, Jens T 44.5, T 44.6 Othman, Gulden •T 38.4 Ott, Philipp T 5.8 Oughton, Sean EP 12.3 Owtscharenko, Nikolaus •T 42.6 P2-Kollaboration T 68.7 Paasch, Alexander •T 7.2 Paasch, Steven T 36.7, •T 85.1, Page, Benjamin T 2.2 Paino, Jason S 13.4 Pak, Sang-II T 9.3 Pal, Divya •T	Piergrossi, Joseph T 24.2 Piergo, Tanguy T 72.5, T 72.6, T 101.7 Pierre Auger-Kollaboration T 19.1, T 19.2, T 19.4, T 19.7, AKPIK 4.6, AKPIK 4.9, T 71.4, T 19.8, T 19.5, T 98.5, T 71.1, T 71.2, T 71.3, T 51.2, T 19.3, T 51.3, T 47.1, T 47.4, T 47.3, T 51.1, T 47.2 Pietroni, Massimo GR 14.5, •GR 14.6 Pinamonti, Nicola •GR 12.2, •MP 9.2 Pioter, Mona •T 22.7 Pirinen, Pekka DD 36.1 Pirner, Hans Jürgen •AGPhil 12.3 Pirola, Giorgio •T 17.8 Piscopo, Maria Laura T 56.4 Pitt, Michael T 7.5 Pitzermann, Martin ST 6.4 Pitz, Daniel T 12.8, T 12.9, T 30.8 Pivovarov, Alexei A. T 56.5 Pizzella, Veronica •T 78.3 Planinic, Maja D D 18.2 Plaschke, Ferdinand EP 3.4, EP 3.5 Plehn, Tilman T 2.6, T 2.7, T 2.8,
Nagataki, Shigehiro EP 1.4 Näger, Paul M. •AGPhil 9.2 namasaka, wilson •T 47.6 Nandi, Gerrit DD 5.2 Naranjo, Pedro AGPhil 1.1 Nardi, Roberto GP 2.3 Nardo, Lorenzo ST 9.2 Narimani Charan, Abtin •T 79.4 Nasr Esfahani, Alina •T 51.5 Nass, Christian •T 62.3 Nast, Natalie T 107.4 Nattland, Philipp T 33.2 Navarro, Jaume •GP 6.2 Navarro, Jaume •GP 6.2 Navarro, Jaume •GP 6.2 Navarod, Ronny DD 5.2, DD 12.7, DD 31.1, DD 34.12, DD 36.5, DD 44.3 Nayaz, Abdullah •T 16.4 Necker, Jannis •T 18.2 Nedora, Vsevolod •GR 11.3 Negar, Farjoud Masouleh EP 2.2 Nell, Sebastian DD 12.8 Nemeth-Csoka, Marton •T 30.5 Netwerk Teilchenwelt-Kollaboration T 24.3, T 24.1, T 52.1, T 52.4, T 24.2, T 52.3 Neuendorffer, Stephen T 69.7 Neuendorffer, Stephen T 65.6, T 65.6, T 65.6, T 65.6, T 65.6, T 65.7, T 94.4	Olschewsky, Kevin •• T 56.7 Omeliukh, Anastasiia •EP 7.10 Omidvari, Negar ST 9.2 on Envision team, VenSpec EP 3.1 on VERTIAS team, VEM EP 3.1 Öncel, Ö. Oğul T 8.4, T 8.5 Ongisi, Peter T 60.3, T 60.4 Ördek, Serhat T 89.4 Orekhov, Vsevolod •T 102.9 Ørsøe, Rasmus •AKPIK 4.7 Ortjohann, Hans-Werner T 21.4, T 21.5 Ortmann, Julia •DD 50.2 Ortmann, Tobias T 74.7, T 78.6, T 78.7, T 104.5, T 104.6, •T 104.7 Ossokine, Serguei Ossokine, Serguei GR 9.2 Osterhoff, Jens T 44.5, T 44.6 Othman, Gulden •T 38.4 Ott, Philipp T 5.8 Oughton, Sean EP 12.3 Owtscharenko, Nikolaus •T 42.6 P2-Kollaboration T 68.7 Paasch, Alexander •T 7.2 Paasch, Steven T 36.7, •T 85.1, •T 85.5 Page, Benjamin T 2.2 Paino, Jason ST 3.4 Pak, Sang-II T 9.3 Pal, Di	Piergrossi, Joseph
Nagataki, Shigehiro EP 1.4 Näger, Paul M. •AGPhil 9.2 namasaka, wilson •T 47.6 Nandi, Gerrit DD 5.2 Naranjo, Pedro AGPhil 1.1 Nardi, Roberto GP 2.3 Nardo, Lorenzo ST 9.2 Narimani Charan, Abtin •T 79.4 Nasr Esfahani, Alina •T 51.5 Nass, Christian •T 62.3 Nast, Natalie T 107.4 Nattland, Philipp T 33.2 Navarro, Jaume •GP 6.2 Nawrodt, Ronny DD 5.2, DD 12.7, DD 31.1, DD 34.12, DD 36.5, DD 44.3 Nayaz, Abdullah •T 18.2 Nedora, Vsevolod •GR 11.3 Negar, Farjoud Masouleh EP 2.2 Nell, Sebastian DD 12.8 Nemeth-Csoka, Marton •T 30.5 Netzwerk Teilchenwelt-Kollaboration T 24.3, T 24.1, T 52.1, T 52.4, T 24.2, T 52.3 Neuberger, Moritz •T 75.3 Neuendorffer, Stephen T 69.7 Neurdorffer, Stephen T 65.5, T 65.6, T 65.7, T 94.4 Neuhaus, Simon •T 34.6 Neukirch, Thomas EP 12.10 Neukum, Max <t< td=""><td>Olschewsky, Kevin •• T 56.7 Omeliukh, Anastasiia •EP 7.10 Omidvari, Negar ST 9.2 on Envision team, VenSpec EP 3.1 on VERTIAS team, VEM EP 3.1 Öncel, Ö. Oğul T 8.4, T 8.5 Onyisi, Peter T 60.3, T 60.4 Ördek, Serhat T 89.4 Orekhov, Vsevolod •T 102.9 Ørsøe, Rasmus •AKPIK 4.7 Ortjohann, Hans-Werner T 21.4, T 21.5 Ortmann, Julia •D 50.2 Ortmann, Julia •D 50.2 Ortmann, Julia •D 50.2 Ortmann, Tobias T 74.7, T 78.6, T 78.7, T 104.5, T 104.6, •T 104.7 Ossokine, Serguei Ossokine, Serguei GR 9.2 Osterhoff, Jens T 44.5, T 44.6 Othman, Gulden •T 38.4 Ott, Philipp T 5.8 Oughton, Sean EP 12.3 Owtscharenko, Nikolaus •T 42.6 P2-Kollaboration T 68.7 Paasch, Alexander •T 7.2 Paasch, Alexander •T 7.2 Paasch, Alexander •T 7.2 Paasch, Jason <t< td=""><td>Piergrossi, Joseph T 24.2 Piergo, Tanguy T 72.5, T 72.6, T 101.7 Pierre Auger-Kollaboration T 19.1, T 19.2, T 19.4, T 19.7, AKPIK 4.6, AKPIK 4.9, T 71.4, T 19.8, T 19.5, T 98.5, T 71.1, T 71.2, T 71.3, T 51.2, T 19.3, T 51.3, T 47.1, T 47.4, T 47.3, T 51.1, T 47.2 Pietroni, Massimo GR 14.5, •GR 14.6 Pinamonti, Nicola •GR 12.2, •MP 9.2 Piotter, Mona Pirola, Giorgio Pirner, Pekka D 36.1 Pirner, Hans Jürgen •AGPhil 12.3 Pirola, Giorgio Pitt, Michael T 7.5 Pittermann, Martin •ST 6.4 Pitz, Janiel Pizella, Veronica •T 78.3 Planinic, Maja DD 18.2 Plaschke, Ferdinand EP 3.4, EP 3.5 Piefka, Jan Piscopo, 2, T 86.7, T 90.3, T 93.1, T 93.4, T 93.5, AKPIK 4.2</td></t<></td></t<>	Olschewsky, Kevin •• T 56.7 Omeliukh, Anastasiia •EP 7.10 Omidvari, Negar ST 9.2 on Envision team, VenSpec EP 3.1 on VERTIAS team, VEM EP 3.1 Öncel, Ö. Oğul T 8.4, T 8.5 Onyisi, Peter T 60.3, T 60.4 Ördek, Serhat T 89.4 Orekhov, Vsevolod •T 102.9 Ørsøe, Rasmus •AKPIK 4.7 Ortjohann, Hans-Werner T 21.4, T 21.5 Ortmann, Julia •D 50.2 Ortmann, Julia •D 50.2 Ortmann, Julia •D 50.2 Ortmann, Tobias T 74.7, T 78.6, T 78.7, T 104.5, T 104.6, •T 104.7 Ossokine, Serguei Ossokine, Serguei GR 9.2 Osterhoff, Jens T 44.5, T 44.6 Othman, Gulden •T 38.4 Ott, Philipp T 5.8 Oughton, Sean EP 12.3 Owtscharenko, Nikolaus •T 42.6 P2-Kollaboration T 68.7 Paasch, Alexander •T 7.2 Paasch, Alexander •T 7.2 Paasch, Alexander •T 7.2 Paasch, Jason <t< td=""><td>Piergrossi, Joseph T 24.2 Piergo, Tanguy T 72.5, T 72.6, T 101.7 Pierre Auger-Kollaboration T 19.1, T 19.2, T 19.4, T 19.7, AKPIK 4.6, AKPIK 4.9, T 71.4, T 19.8, T 19.5, T 98.5, T 71.1, T 71.2, T 71.3, T 51.2, T 19.3, T 51.3, T 47.1, T 47.4, T 47.3, T 51.1, T 47.2 Pietroni, Massimo GR 14.5, •GR 14.6 Pinamonti, Nicola •GR 12.2, •MP 9.2 Piotter, Mona Pirola, Giorgio Pirner, Pekka D 36.1 Pirner, Hans Jürgen •AGPhil 12.3 Pirola, Giorgio Pitt, Michael T 7.5 Pittermann, Martin •ST 6.4 Pitz, Janiel Pizella, Veronica •T 78.3 Planinic, Maja DD 18.2 Plaschke, Ferdinand EP 3.4, EP 3.5 Piefka, Jan Piscopo, 2, T 86.7, T 90.3, T 93.1, T 93.4, T 93.5, AKPIK 4.2</td></t<>	Piergrossi, Joseph T 24.2 Piergo, Tanguy T 72.5, T 72.6, T 101.7 Pierre Auger-Kollaboration T 19.1, T 19.2, T 19.4, T 19.7, AKPIK 4.6, AKPIK 4.9, T 71.4, T 19.8, T 19.5, T 98.5, T 71.1, T 71.2, T 71.3, T 51.2, T 19.3, T 51.3, T 47.1, T 47.4, T 47.3, T 51.1, T 47.2 Pietroni, Massimo GR 14.5, •GR 14.6 Pinamonti, Nicola •GR 12.2, •MP 9.2 Piotter, Mona Pirola, Giorgio Pirner, Pekka D 36.1 Pirner, Hans Jürgen •AGPhil 12.3 Pirola, Giorgio Pitt, Michael T 7.5 Pittermann, Martin •ST 6.4 Pitz, Janiel Pizella, Veronica •T 78.3 Planinic, Maja DD 18.2 Plaschke, Ferdinand EP 3.4, EP 3.5 Piefka, Jan Piscopo, 2, T 86.7, T 90.3, T 93.1, T 93.4, T 93.5, AKPIK 4.2
Nagataki, Shigehiro EP 1.4 Näger, Paul M. •AGPhil 9.2 namasaka, wilson •T 47.6 Nandi, Gerrit DD 5.2 Naranjo, Pedro AGPhil 9.2 Naranjo, Pedro AGPhil 1.1 Nardi, Roberto GP 2.3 Nardo, Lorenzo ST 9.2 Narimani Charan, Abtin •T 79.4 Nasr Esfahani, Alina •T 51.5 Nass, Christian •T 62.3 Nast, Natalie T 107.4 Nattland, Philipp T 33.2 Navarro, Jaume •GP 6.2 Nawrodt, Ronny DD 5.2, DD 12.7, DD 31.1, DD 34.12, DD 36.5, DD 44.3 Nayaz, Abdullah •T 16.4 Necker, Jannis •T 18.2 Nedora, Vsevolod •GR 11.3 Negar, Farjoud Masouleh EP 2.2 Nell, Sebastian DD 12.8 Neuberger, Moritz •T 75.3 Neuberger, Moritz •T 75.3 Neuberger, Moritz •T 75.3 Neuberger, Moritz •T 65.5, T 65.6, T 65.7, T 94.4 Neubaus, Simon •T 34.6 Neukirch, Thomas EP 12.10 Neukum, Max	Olschewsky, Kevin •• T 56.7 Omeliukh, Anastasiia •EP 7.10 Omidvari, Negar ST 9.2 on Envision team, VenSpec EP 3.1 on VERTIAS team, VEM EP 3.1 Onel, Ö. Oğul T 8.4, T 8.5 Onyisi, Peter T 60.3, T 60.4 Ördek, Serhat T 89.4 Orekhov, Vsevolod •T 102.9 Ørsøe, Rasmus •AKPIK 4.7 Ortjohann, Hans-Werner T 21.4, T 21.5 Ortmann, Julia •D 50.2 Ortmann, Gulden T 38.4 Ott, Philipp T 5.8 Oughton, Sean EP 12.3 Owtscharenko, Nikolaus •T 42.6 P2-Kollaboration T 68.7 Paasch, Alexander •T 7.2 Paasch, Steven T 36.7, •T 85.1, •T 85.5 Page, Benjamin T 2.2 Paino, Jason ST 3.4 Pak, Sang-II T 9.3 Pal, Divya •T 39.8, T 68.1, T 97.1	Piergrossi, Joseph
Nagataki, Shigehiro EP 1.4 Näger, Paul M. •AGPhil 9.2 namasaka, wilson •T 47.6 Nandi, Gerrit DD 5.2 Naranjo, Pedro AGPhil 9.2 Naranjo, Pedro AGPhil 1.1 Nardi, Roberto GP 2.3 Nardo, Lorenzo ST 9.2 Narimani Charan, Abtin •T 79.4 Nasr Esfahani, Alina •T 51.5 Nass, Christian •T 62.3 Nattland, Philipp T 33.2 Navaro, Jaume •GP 6.2 Nawrodt, Ronny DD 5.2, DD 12.7, DD 31.1, DD 34.12, DD 36.5, DD 44.3 Nayaz, Abdullah •T 16.4 Necker, Jannis •T 18.2 Nedora, Vsevolod •GR 11.3 Negar, Farjoud Masouleh EP 2.2 Nell, Sebastian DD 12.8 Nemeth-Csoka, Marton •T 30.5 Netzwerk Teilchenwelt-Kollaboration T 24.3, T 24.1, T 52.1, T 52.4, T 24.2, T 53.3 Neuberger, Moritz •T 75.3 Neuberger, Moritz •T 65.5, T 65.6, T 65.7, T 94.4 Neuhaus, Simon •T 34.6 Neukirch, Thomas EP 12.10	Olschewsky, Kevin •• T 56.7 Omeliukh, Anastasiia •EP 7.10 Omidvari, Negar ST 9.2 on Envision team, VenSpec EP 3.1 on VERTIAS team, VEM EP 3.1 Oncel, Ö. Ögul T 8.4, T 8.5 Ongids, Serhat T 89.4 Orekhov, Vsevolod •T 102.9 Ørsøe, Rasmus •AKPIK 4.7 Ortghonn, Hans-Werner T 21.4, T 21.5 Ortmann, Julia •D 50.2 Ortmann, Tobias T 74.7, T 78.6, T 78.7, T 104.5, T 104.6, •T 104.7 Ossokine, Serguei G 9.2 Osterhoff, Jens T 44.5, T 44.6 Otthon, Sean EP 12.3 Owtscharenko, Nikolaus •T 42.6 Oughton, Sean EP 12.3 Owtscharenko, Nikolaus •T 42.6 P2-Kollaboration T 68.7 Paasch, Steven T 36.7, •T 85.1, •T 85.5 Page, Benjamin Paasch, Steven T 36.7, •T 85.1, •T 85.5 Page, Benjamin Palacios Schweitzer, Sofia •T 3.2 Paino, Jason ST 3.4 Pak, Sang-II T 9.3 Pal	Piergrossi, Joseph T 24.2 Piergossi, Joseph T 24.2 Piergossi, Joseph T 24.2 Piergossi, Joseph T 25.5, T 72.6, T 101.7 Pierre Auger-Kollaboration T 19.1, T 19.2, T 19.4, T 19.7, AKPIK 4.6, AKPIK 4.9, T 71.4, T 19.8, T 19.5, T 98.5, T 71.1, T 71.2, T 71.3, T 51.2, T 19.3, T 51.3, T 47.1, T 47.4, T 47.3, T 51.1, T 47.2 Pietroni, Massimo GR 6.2 Pilarczyk, Oliver •T 48.5 Pimpalshende, Vivek GR 14.4, GR 14.5, •GR 14.6 F 14.6 Pinamonti, Nicola •GR 12.2, •MP 9.2 Piotter, Mona •T 22.7 Pirinen, Pekka D 0 36.1 Pinner, Hans Jürgen •AGPhil 12.3 Pirsch, Johannes •MP 2.6 Piscopo, Maria Laura T 56.4 Pitt, Michael T 7.5 Pitzermann, Martin •ST 6.4 Pitz, Daniel T 12.8, T 12.9, T 30.8 Pivovarov, Alexei A. T 56.5 Pizzella, Veronica •T 78.3 Plashke, Ferdinand EP 3.4, EP 3.5 Plefka, Jan •MP 2.2 Plehn, Tilman T 2.6, T 2.7
Nagataki, Shigehiro EP 1.4 Näger, Paul M. •AGPhil 9.2 namasaka, wilson •T 47.6 Nandi, Gerrit DD 5.2 Naranjo, Pedro AGPhil 9.2 Narda, Kolernit DD 5.2 Naranjo, Pedro AGPhil 1.1 Nardi, Roberto GP 2.3 Nardo, Lorenzo ST 9.2 Narimani Charan, Abtin •T 79.4 Nasr Esfahani, Alina •T 51.5 Nass, Christian •T 62.3 Nast, Natalie T 107.4 Nattland, Philipp T 33.2 Navarro, Jaume •GP 6.2 Nawrodt, Ronny D 5.2, DD 12.7, D 31.1, DD 34.12, DD 36.5, DD 44.3 Nayaz, Abdullah •T 16.4 Necker, Jannis •T 18.2 Nedora, Vsevolod •GR 11.3 Negar, Farjoud Masouleh EP 2.2 Nell, Sebastian DD 12.8 Nemeth-Csoka, Marton •T 30.5 Netzwerk Teilchenwelt-Kollaboration T 24.3, T 24.1, T 52.1, T 52.4, T 24.2, T 52.3 Neuberger, Moritz •T 75.3 Neuendorffer, Stephen T 69.7 Neukirch, Thomas	Olschewsky, Kevin •• 156.7 Omeliukh, Anastasiia •• EP 7.10 Omidvari, Negar ST 9.2 on Envision team, VenSpec EP 3.1 on VERTIAS team, VEM EP 3.1 Öncel, Ö. Oğul T 8.4, T 8.5 Ongisi, Peter T 60.3, T 60.4 Ördek, Serhat T 89.4 Orekhov, Vsevolod •T 102.9 Ørsøe, Rasmus •AKPIK 4.7 Ortjohann, Hans-Werner T 21.4, T 21.5 Ortmann, Julia •DD 50.2 Ortmann, Tobias T 74.7, T 78.6, T 78.7, T 104.5, T 104.6, •T 104.7 Ossokine, Serguei Ossokine, Serguei GR 9.2 Osterhoff, Jens T 44.5, T 44.6 Othman, Gulden •T 38.4 Ott, Philipp T 5.8 Oughton, Sean EP 12.3 Owtscharenko, Nikolaus •T 42.6 P2-Kollaboration T 68.7 Paasch, Alexander •T 7.2 Paasch, Steven T 36.7, •T 85.1, •T 85.5 Page, Benjamin T 2.2 Paino, Jason ST 3.4 Pak, Sang-II T 9.3 Pal, D	Piergrossi, Joseph T 24.2 Piergossi, Joseph T 24.2 Piergossi, Joseph T 2.5, T 72.6, T 101.7 Pierre Auger-Kollaboration T 19.1, T 19.2, T 19.4, T 19.7, AKPIK 4.6, AKPIK 4.9, T 71.4, T 19.8, T 19.5, T 98.5, T 71.1, T 71.2, T 71.3, T 51.2, T 19.3, T 51.3, T 47.1, T 47.4, T 47.3, T 51.1, T 47.2 Pietroni, Massimo GR 6.2 Pilarczyk, Oliver T 48.5 Pimpalshende, Vivek GR 14.4, GR 14.5, •GR 14.6 Pinamonti, Nicola •GR 12.2, •MP 9.2 Piotter, Mona T 22.7 Pirnen, Pekka DD 36.1 Pirner, Hans Jürgen •AGPhil 12.3 Pirola, Giorgio •T 17.8 Pirsch, Johannes •MP 2.6 Piscopo, Maria Laura T 56.4 Pitt, Michael T 7.5 Pittermann, Martin •ST 6.4 Pitz, Daniel T 12.8, T 12.9, T 30.8 Pivovarov, Alexei A. T 56.5 Pizzella, Veronica •T 78.3 Planinic, Maja D D 18.2 Plehn, Tilman T 2.6, T 2.7, T 2.8, T 31.2, T 86.7, T 90.3, T 93.1, T 93.4, T 93.5, AKPIK 4.2 Plesanovs, Vladislavs T 39.1, T 39.2, •T 39.3 Plicht, Katja •DD 34.11
Nagataki, Shigehiro EP 1.4 Näger, Paul M. •AGPhil 9.2 namasaka, wilson •T 47.6 Nandi, Gerrit DD 5.2 Naranjo, Pedro AGPhil 1.1 Nardi, Roberto GP 2.3 Nardo, Lorenzo ST 9.2 Narimani Charan, Abtin •T 79.4 Nasr Esfahani, Alina •T 51.5 Nass, Christian •T 62.3 Nast, Natalie T 107.4 Nattland, Philipp T 33.2 Navarro, Jaume •GP 6.2 Navarro, Jaume •GP 6.2 Navarro, Jaume •GP 6.2 Navarod, Ronny DD 5.2, DD 12.7, DD 31.1, DD 34.12, DD 36.5, DD 44.3 Nayaz, Abdullah •T 16.4 Necker, Jannis •T 18.2 Nedora, Vsevolod •GR 11.3 Negar, Farjoud Masouleh EP 2.2 Nell, Sebastian DD 12.8 Nemeth-Csoka, Marton •T 30.5 Netzwerk Teilchenwelt-Kollaboration T 24.3, T 24.1, T 52.1, T 52.4, T 24.2, T 52.3 Neubarger, Moritz •T 75.3 Neuendorffer, Stephen T 69.7 Neuhaus, Simon	Olschewsky, Kevin •• T 56.7 Omeliukh, Anastasiia •EP 7.10 Omidvari, Negar ST 9.2 on Envision team, VenSpec EP 3.1 on VERTIAS team, VEM EP 3.1 Öncel, Ö. Oğul T 8.4, T 8.5 Ongis, Peter T 60.3, T 60.4 Ördek, Serhat T 8.4, T 8.5 Orekhov, Vsevolod •T 102.9 Ørsøe, Rasmus •AKPIK 4.7 Ortjohann, Hans-Werner T 2.1, T 21.5 Ortmann, Julia •DD 50.2 Ortmann, Julia •DD 50.2 Ortmann, Julia •D 50.2 Ortmann, Tobias T 74.7, T 78.6, T 78.7, T 104.5, T 104.6, •T 104.7 Ossokine, Serguei Ossokine, Serguei GR 9.2 Osterhoff, Jens T 44.5, T 44.6 Othman, Gulden •T 38.4 Ott, Philipp T 5.8 Oughton, Sean EP 12.3 Owtscharenko, Nikolaus •T 42.6 P2-Kollaboration T 68.7 Paasch, Alexander •T 7.2 Paasch, Steven T 30.7, •T 85.1, •T 85.5 Page, Benjamin T 2.2	Piergrossi, Joseph T 24.2 Piergossi, Joseph T 24.2 Piergo, Tanguy T 72.5, T 72.6, T 101.7 Pierre Auger-Kollaboration T 19.1, T 19.2, T 19.4, T 19.7, AKPIK 4.6, AKPIK 4.9, T 71.4, T 19.8, T 19.5, T 98.5, T 71.1, T 71.2, T 71.3, T 51.2, T 19.3, T 51.3, T 47.1, T 47.4, T 47.3, T 51.1, T 47.2 Pietroni, Massimo GR 6.2 Pilarczyk, Oliver T 48.5 Pimpalshende, Vivek GR 14.4, GR 14.5, •GR 14.6 Pinamonti, Nicola •GR 12.2, •MP 9.2 Piotter, Mona T 22.7 Pirinen, Pekka DD 36.1 Pirner, Hans Jürgen •AGPhil 12.3 Pirola, Giorgio T 17.8 Piscopo, Maria Laura T 56.4 Pitt, Michael T 7.5 Pittermann, Martin ST 6.4 Pitz, Daniel T 12.8, T 12.9, T 30.8 Pivovarov, Alexei A. T 56.5 Pizzella, Veronica T 78.3 Planinic, Maja D 18.2 Pleschke, Ferdinand EP 3.4, EP 3.5 Plefka, Jan •MP 2.2 Plehn, Tilman T 2.6, T 2.7, T 2.8, T 31.2, T
Nagataki, Shigehiro $EP 1.4$ Näger, Paul M. \bullet AGPhil 9.2namasaka, wilson \bullet T 47.6Nandi, GerritDD 5.2Naranjo, PedroAGPhil 1.1Nardi, RobertoGP 2.3Nardo, LorenzoST 9.2Narimani Charan, Abtin \bullet T 79.4Nasr Esfahani, Alina \bullet T 51.5Nass, Christian \bullet T 62.3Natdau, PhilippT 33.2Navarro, Jaume \bullet GP 6.2Nawrodt, RonnyDD 5.2, DD 12.7,DD 31.1, DD 34.12, DD 36.5, DD 44.3Nayaz, Abdullah \bullet T 16.4Necker, Jannis \bullet T 18.2Nedra, Vsevolod \bullet GR 11.3Negar, Farjoud MasoulehE P 2.2Nett-Csoka, Marton \bullet T 30.5Netzwerk Teilchenwelt-KollaborationT 24.3, T 24.1, T 52.1, T 52.4, T 24.2, T 52.3Neuhorffer, StephenT 69.7Neurderld, MariusT 65.5, T 65.6, T 65.7, T 94.4Neuhaus, Simon \bullet T 34.6Neukirch, Thomas \bullet T 12.10Neukum, Max \bullet T 63.9Nguyen, Le HoangT 77.4, T 77.5Nguyen, Le HoangT 77.4, T 77.5Nickeler, Dieter \bullet EP 1.2Nickleir, Dieter \bullet EP 1.2Nickleir, Dieter \bullet EP 1.2	Olschewsky, Kevin •• T 56.7 Omeliukh, Anastasiia •EP 7.10 Omidvari, Negar ST 9.2 on Envision team, VenSpec EP 3.1 on VERTIAS team, VEM EP 3.1 Öncel, Ö. Oğul T 8.4, T 8.5 Onyisi, Peter T 60.3, T 60.4 Ördek, Serhat T 89.4 Orekhov, Vsevolod •T 102.9 Ørsøe, Rasmus •AKPIK 4.7 Ortjohann, Hans-Werner T 21.4, T 21.5 Ortmann, Julia •D 50.2 Ortmann, Julia •D 50.2 Ortmann, Julia ·D 50.2 Ortmann, Tobias T 74.7, T 78.6, T 78.7, T 104.5, T 104.6, •T 104.7 Ossokine, Serguei Ossokine, Serguei GR 9.2 Osterhoff, Jens T 44.5, T 44.6 Othman, Gulden •T 38.4 Ott, Philipp T 5.8 Oughton, Sean EP 12.3 Owtscharenko, Nikolaus •T 42.6 P2-Kollaboration T 68.7 Paasch, Alexander •T 7.2 Paasch, Alexander •T 7.2 Paasch, Alexander •T 3.2 Paino, Jason <tr< td=""><td>Piergrossi, Joseph T 24.2 Piergossi, Joseph T 24.2 Piergossi, Joseph T 25, T 72.6, T 101.7 Pierre Auger-Kollaboration T 19.1, T 19.2, T 19.4, T 19.7, AKPIK 4.6, AKPIK 4.9, T 71.4, T 19.8, T 19.5, T 98.5, T 71.1, T 71.2, T 71.3, T 51.2, T 19.3, T 51.3, T 47.1, T 47.4, T 47.3, T 51.1, T 47.2 Pietroni, Massimo GR 6.2 Pilarczyk, Oliver T 48.5 Pimpalshende, Vivek GR 14.4, GR 14.5, •GR 14.6 Pinamonti, Nicola •GR 12.2, •MP 9.2 Piotter, Mona T 22.7 Pirnen, Pekka DD 36.1 Pirner, Hans Jürgen •AGPhil 12.3 Pirola, Giorgio T 17.8 Piscopo, Maria Laura T 56.4 Pitt, Michael T 7.5 Pittermann, Martin •ST 6.4 Pitz, Daniel T 12.8, T 12.9, T 30.8 Piovarov, Alexei A. T 56.5 Pizzella, Veronica T 78.3 Planinic, Maja DD 18.2 Plaschke, Ferdinand EP 3.4, EP 3.5, Plefka, Jan MP 2.2 Plehn, Tilman T 2.6, T 2.7, T 2.8, T 31.2, T</td></tr<>	Piergrossi, Joseph T 24.2 Piergossi, Joseph T 24.2 Piergossi, Joseph T 25, T 72.6, T 101.7 Pierre Auger-Kollaboration T 19.1, T 19.2, T 19.4, T 19.7, AKPIK 4.6, AKPIK 4.9, T 71.4, T 19.8, T 19.5, T 98.5, T 71.1, T 71.2, T 71.3, T 51.2, T 19.3, T 51.3, T 47.1, T 47.4, T 47.3, T 51.1, T 47.2 Pietroni, Massimo GR 6.2 Pilarczyk, Oliver T 48.5 Pimpalshende, Vivek GR 14.4, GR 14.5, •GR 14.6 Pinamonti, Nicola •GR 12.2, •MP 9.2 Piotter, Mona T 22.7 Pirnen, Pekka DD 36.1 Pirner, Hans Jürgen •AGPhil 12.3 Pirola, Giorgio T 17.8 Piscopo, Maria Laura T 56.4 Pitt, Michael T 7.5 Pittermann, Martin •ST 6.4 Pitz, Daniel T 12.8, T 12.9, T 30.8 Piovarov, Alexei A. T 56.5 Pizzella, Veronica T 78.3 Planinic, Maja DD 18.2 Plaschke, Ferdinand EP 3.4, EP 3.5, Plefka, Jan MP 2.2 Plehn, Tilman T 2.6, T 2.7, T 2.8, T 31.2, T
Nagataki, Shigehiro $EP 1.4$ Näger, Paul M. \bullet AGPhil 9.2namasaka, wilson \bullet T 47.6Nandi, GerritDD 5.2Naranjo, PedroAGPhil 1.1Nardi, RobertoGP 2.3Nardo, LorenzoST 9.2Narimani Charan, Abtin \bullet T 79.4Nasr Esfahani, Alina \bullet T 51.5Nass, Christian \bullet T 62.3Nattland, PhilippT 33.2Navaro, Jaume \bullet GP 6.2Nawrodt, RonnyDD 5.2, DD 12.7,DD 31.1, DD 34.12, DD 36.5, DD 44.3Nayaz, Abdullah \bullet T 16.4Necker, Jannis \bullet T 18.2Nedora, Vsevolod \bullet GR 11.3Negar, Farjoud Masouleh $EP 2.2$ Nell, SebastianDD 12.8Neuberger, Moritz \bullet T 75.3Neuendorffer, StephenT 69.7Neufeld, MariusT 65.5, T 65.6,T 65.7, T 94.4Neuhaus, SimonNeukorf, Homas \bullet T 12.10Neukum, Max \bullet T 87.5Neundorf, Jonas \bullet T 30.5Nickeler, Dieter \bullet EP 12.10Neuker, Dieter \bullet EP 12.10Nickeler, Dieter \bullet EP 12.10<	Olschewsky, Kevin •• T 56.7 Omeliukh, Anastasiia •EP 7.10 Omidvari, Negar ST 9.2 on Envision team, VenSpec EP 3.1 on VERTIAS team, VEM EP 3.1 Öncel, Ö. Oğul T 8.4, T 8.5 Onyisi, Peter T 60.3, T 60.4 Ördek, Serhat T 89.4 Orekhov, Vsevolod •T 102.9 Ørsøe, Rasmus •AKPIK 4.7 Ortjohann, Hans-Werner T 21.4, T 21.5 Ortmann, Julia 0D 50.2 Ortmann, Julia 0D 50.2 Ortmann, Tobias T 74.7, T 78.6, T 78.7, T 104.5, T 104.6, •T 104.7 Ossokine, Serguei GR 9.2 Osterhoff, Jens T 44.5, T 44.6 Othman, Gulden •T 38.4 Ott, Philipp T 5.8 Oughton, Sean EP 12.3 Owtscharenko, Nikolaus •T 42.6 P2-Kollaboration T 68.7 Paasch, Alexander •T 7.2 Paasch, Alexander •T 7.2 Paasch, Steven T 36.7, •T 85.1, •T 85.5 Page, Benjamin T 2.2 Paino, Jason ST 3.4	Piergrossi, Joseph T 24.2 Piergossi, Joseph T 24.2 Piergossi, Joseph T 25, T 72.6, T 101.7 Pierre Auger-Kollaboration T 19.1, T 19.2, T 19.4, T 19.7, AKPIK 4.6, AKPIK 4.9, T 71.4, T 19.8, T 19.5, T 98.5, T 71.1, T 71.2, T 71.3, T 51.2, T 19.3, T 51.3, T 47.1, T 47.4, T 47.3, T 51.1, T 47.2 Pietroni, Massimo GR 6.2 Pilarczyk, Oliver •T 48.5 Pimpalshende, Vivek GR 14.4, GR 14.5, •GR 14.6 Pinamonti, Nicola •GR 12.2, •MP 9.2 Piotter, Mona •T 22.7 Pirinen, Pekka D 36.1 Pirrer, Hans Jürgen •AGPhil 12.3 Pirsch, Johannes •MP 2.6 Piscopo, Maria Laura T 56.4 Pittermann, Martin •ST 6.4 Pitzl, Daniel T 12.8, T 12.9, T 30.8 Pivovarov, Alexei A. T 56.5 Pizzella, Veronica •T 78.3 Planinic, Maja DD 18.2 Plesanovs, Vladislavs T 39.1, T 39.4, T 31.2, T 86.7, T 90.3, T 93.1, T 93.4, T 93.5, AKPIK 4.2 Plesanovs, Vladislavs T 39.1, T 39.2, •T 39.3 Plic
Nagataki, Shigehiro $EP 1.4$ Näger, Paul M. \bullet AGPhil 9.2namasaka, wilson \bullet T 47.6Nandi, GerritDD 5.2Naranjo, PedroAGPhil 1.1Nardi, Roberto $GP 2.3$ Nardo, LorenzoST 9.2Narimani Charan, Abtin \bullet T 79.4Nasr Esfahani, Alina \bullet T 62.3Nast, NatalieT 107.4Nast, NatalieT 107.4Nattland, PhilippT 33.2Navaro, Jaume \bullet GP 6.2Nawrodt, RonnyDD 5.2, DD 12.7,DD 31.1, DD 34.12, DD 36.5, DD 44.3Nayaz, Abdullah \bullet T 16.4Necker, Jannis \bullet T 18.2Nedora, Vsevolod \bullet GR 11.3Negar, Farjoud Masouleh $EP 2.2$ Nell, SebastianDD 12.8Neuberger, Moritz \bullet T 53.3Neuberger, Moritz \bullet T 53.3Neuberger, Moritz \bullet T 65.5, T 65.6,T 65.7, T 94.4Neuhaus, SimonNeudorf, Jonas \bullet 787.5Neundorf, Jonas \bullet 734.6Neukum, Max \bullet T 87.5Neundorf, Jonas \bullet 73.5Nickel, Lukas \bullet T 17.3Nickeler, Dieter \bullet EP 1.2Nickler, Dieter \bullet EP 1.2<	Olschewsky, Kevin •• T 56.7 Omeliukh, Anastasiia •• EP 7.10 Omidvari, Negar ST 9.2 on Envision team, VenSpec EP 3.1 on VERTIAS team, VEM EP 3.1 Öncel, Ö. Oğul T 8.4, T 8.5 Ongisi, Peter T 60.3, T 60.4 Ördek, Serhat T 89.4 Orekhov, Vsevolod •T 102.9 Ørsøe, Rasmus •AKPIK 4.7 Ortjohann, Hans-Werner 72.1, T 78.6, T 78.7, T 104.5, T 104.6, •T 104.7 Osokine, Serguei GR 9.2 Osterhoff, Jens T 44.5, T 44.6 Othman, Gulden •T 38.4 Ott, Philipp T 5.8 Oughton, Sean EP 12.3 Owtscharenko, Nikolaus •T 42.6 P2-Kollaboration T 68.7 Paasch, Alexander •T 7.2 Paasch, Alexander •T 7.2 Paasch, Steven T 36.7, •T 85.1, •T 85.5 Page, Benjamin T 2.2 Paino, Jason ST 3.4 Pak, Sang-II T 9.3 Pal, Divya •T 39.8, T 68.1, T 97.1 Palacios Schweitzer, Sofia	Piergrossi, Joseph T 24.2 Piergossi, Joseph T 24.2 Piergossi, Joseph T 25.5, T 72.6, T 101.7 Pierre Auger-Kollaboration T 19.1, T 19.2, T 19.4, T 19.7, AKPIK 4.6, AKPIK 4.9, T 71.4, T 19.8, T 19.5, T 98.5, T 71.1, T 71.2, T 71.3, T 51.2, T 19.3, T 51.3, T 47.1, T 47.4, T 47.3, T 51.1, T 47.2 Pietroni, Massimo GR 6.2 Pilarczyk, Oliver •T 48.5 Pimpalshende, Vivek GR 14.4, GR 14.5, •GR 14.6 Finamonti, Nicola Pinter, Mona •T 22.7 Pirinen, Pekka D 0 36.1 Pirner, Hans Jürgen •AGPhil 12.3 Pirsch, Johannes •MP 2.6 Piscopo, Maria Laura T 56.4 Pitt, Michael T 7.5 Pitzel, Daniel T 12.8, T 12.9, T 30.8 Pivovarov, Alexei A. T 56.5 Pizzella, Veronica •T 78.3 Plaschke, Ferdinand EP 3.4, EP 3.5 Plefka, Jan •MP 2.2 Plehn, Tilman T 2.6, T 2.7, T 2.8, T 31.2, T 86.7, T 90.3, T 93.1, T 93.4, T 93.5, AKPIK 4.2 Plesanovs, Vladislavs T 39.1, T 39.2,
Nagataki, Shigehiro $EP 1.4$ Näger, Paul M. $\cdot AGPhil 9.2$ namasaka, wilson $\cdot T 47.6$ Nandi, GerritDD 5.2Naranjo, Pedro $AGPhil 1.1$ Nardi, Roberto $GP 2.3$ Narido, LorenzoST 9.2Narimani Charan, Abtin $\cdot T 79.4$ Nasr Esfahani, Alina $\cdot T 51.5$ Nass, Christian $\cdot T 62.3$ Nast, Natalie $T 107.4$ Nattland, Philipp $T 33.2$ Navarro, Jaume $\cdot GP 6.2$ Nawrodt, RonnyDD 5.2, DD 12.7,DD 31.1, DD 34.12, DD 36.5, DD 44.3Nayaz, Abdullah $\cdot T 16.4$ Necker, Jannis $\cdot T 18.2$ Nedora, Vsevolod $\cdot GR 11.3$ Negar, Farjoud Masouleh $EP 2.2$ Nemeth-Csoka, Marton $\cdot T 30.5$ Netzwerk Teilchenwelt-Kollaboration $T 24.3, T 24.1, T 52.1, T 52.4, T 24.2, T 52.3$ Neuberger, Moritz $\cdot T 75.3$ Neuendorffer, StephenT 69.7Neufled, Marius $\cdot T 65.5, T 65.6, T 65.7, T 94.4$ Neukinch, Thomas $EP 12.10$ Neukum, Max $\cdot T 33.5$ Nickeler, Dieter $\cdot EP 1.2$ Nickel, Lukas $\cdot T 17.3$ Nickeler, Dieter $\cdot EP 1.2$ Nickeler, Dieter $\cdot ET 3.6$	Olschewsky, Kevin •• 156.7 Omeliukh, Anastasiia •EP 7.10 Omidvari, Negar ST 9.2 on Envision team, VenSpec EP 3.1 on VERTIAS team, VEM EP 3.1 Öncel, Ö. Oğul T 8.4, T 8.5 Ongisi, Peter T 60.3, T 60.4 Ördek, Serhat T 89.4 Orekhov, Vsevolod •T 102.9 Ørsøe, Rasmus •AKPIK 4.7 Ortighann, Hans-Werner T 21.4, T 21.5 Ortmann, Julia •DD 50.2 Ortmann, Tobias T 74.7, T 78.6, T 78.7, T 104.5, T 104.6, •T 104.7 Ossokine, Serguei Ossokine, Serguei GR 9.2 Osterhoff, Jens T 44.5, T 44.6 Othman, Gulden •T 38.4 Ott, Philipp T 5.8 Oughton, Sean EP 12.3 Owtscharenko, Nikolaus •T 42.6 P2-Kollaboration T 68.7 Paasch, Alexander •T 7.2 Paasch, Steven T 36.7, •T 85.1, •T 85.5 Page, Benjamin T 2.2 Paino, Jason ST 3.4 Pak, Sang-II T 9.3 Pal, Div	Piergrossi, Joseph T 24.2 Piergo, Tanguy T 72.5, T 72.6, T 101.7 Pierre Auger-Kollaboration T 19.1, T 19.2, T 19.4, T 19.7, AKPIK 4.6, AKPIK 4.9, T 71.4, T 19.8, T 19.5, T 98.5, T 71.1, T 71.2, T 71.3, T 51.2, T 19.3, T 51.3, T 47.1, T 47.4, T 47.3, T 51.1, T 47.2 Pietroni, Massimo GR 6.2 Pilarczyk, Oliver •T 48.5 Pimpalshende, Vivek GR 14.4, GR 14.5, •GR 14.6 •T 22.7 Piotter, Mona •T 22.7 Pirinen, Pekka DD 36.1 Pirner, Hans Jürgen •AGPhil 12.3 Pirola, Giorgio •T 17.8 Pirsch, Johannes •MP 2.6 Piscopo, Maria Laura T 56.4 Pitt, Michael T 7.5 Pittermann, Martin •ST 6.4 Pitz, Daniel T 12.8, T 12.9, T 30.8 Pivovarov, Alexei A. T 56.5 Pizzella, Veronica •T 78.3 Plaschke, Ferdinand EP 3.4, EP 3.5 Plefka, Jan •MP 2.2 Plehn, Tilman T 2.6, T 2.7, T 2.8, T 31.2, T 86.7, T 90.3, T 39.1, T 39.4, T 39.5, AKPIK 4.2 Plesanovs, Vladislavs T 39.1, T 39.2, •T 39.3<
Nagataki, Shigehiro EP 1.4 Näger, Paul M. •AGPhil 9.2 namasaka, wilson •T 47.6 Nandi, Gerrit DD 5.2 Naranjo, Pedro AGPhil 1.1 Nardi, Roberto GP 2.3 Nardo, Lorenzo ST 9.2 Narimani Charan, Abtin •T 79.4 Nasr Esfahani, Alina •T 51.5 Nass, Christian •T 62.3 Nast, Natalie T 107.4 Nattland, Philipp T 33.2 Navarro, Jaume •GP 6.2 Navarro, Jaume •GP 6.2 Navarro, Jaume •GP 6.2 Navarodt, Ronny DD 5.2, DD 12.7, DD 31.1, DD 34.12, DD 36.5, DD 44.3 Nayaz, Abdullah •T 16.4 Necker, Jannis •T 18.2 Nedora, Vsevolod •GR 11.3 Negar, Farjoud Masouleh EP 2.2 Nell, Sebastian DD 12.8 Nemeth-Csoka, Marton •T 30.5 Netzwerk Teilchenwelt-Kollaboration T 24.3, T 24.1, T 52.1, T 52.4, T 24.2, T 52.3 Neubarger, Moritz •T 75.3 Neuendorffer, Stephen T 65.7, T 65.6, T 65.6, T 65.7, T 65.4, T 65.7, T 65.4, T 24.3,	Olschewsky, Kevin •• T 56.7 Omeliukh, Anastasiia •EP 7.10 Omidvari, Negar ST 9.2 on Envision team, VenSpec EP 3.1 on VERTIAS team, VEM EP 3.1 Öncel, Ö. Oğul T 8.4, T 8.5 Ongisi, Peter T 60.3, T 60.4 Ördek, Serhat T 8.4, T 8.5 Orekhov, Vsevolod •T 102.9 Ørsøe, Rasmus •AKPIK 4.7 Ortjohann, Hans-Werner T 2.1, T 21.5 Ortmann, Julia •DD 50.2 Ortmann, Julia •DD 50.2 Ortmann, Tobias T 74.7, T 78.6, T 78.7, T 104.5, T 104.6, •T 104.7 Ossokine, Serguei Ossokine, Serguei GR 9.2 Osterhoff, Jens T 44.5, T 44.6 Othman, Gulden •T 38.4 Ott, Philipp T 5.8 Oughton, Sean EP 12.3 Owtscharenko, Nikolaus •T 42.6 P2-Kollaboration T 68.7 Paasch, Alexander •T 7.2 Paasch, Steven T 30.7, •T 39.8, T 68.1, T 97.1 Palacios Schweitzer, Sofia •T 3.2 Paliwal, Ayush •MP 7.5	Piergrossi, Joseph T 24.2 Piergossi, Joseph T 24.2 Piergossi, Joseph T 25, T 72.6, T 101.7 Pierre Auger-Kollaboration T 19.1, T 19.2, T 19.4, T 19.7, AKPIK 4.6, AKPIK 4.9, T 71.4, T 19.8, T 19.5, T 98.5, T 71.1, T 71.2, T 71.3, T 51.2, T 19.3, T 51.3, T 47.1, T 47.4, T 47.3, T 51.1, T 47.2 Pietroni, Massimo GR 6.2 Pilarczyk, Oliver T 48.5 Pimpalshende, Vivek GR 14.4, GR 14.5, •GR 14.6 Pinamonti, Nicola •GR 12.2, •MP 9.2 Piotter, Mona T 22.7 Pirinen, Pekka DD 36.1 Pirner, Hans Jürgen •AGPhil 12.3 Pirola, Giorgio T 17.8 Piscopo, Maria Laura T 56.4 Pitt, Michael T 7.5 Pittermann, Martin ST 6.4 Pitz, Daniel T 12.8, T 12.9, T 30.8 Pivovarov, Alexei A. T 56.5 Pizzella, Veronica T 78.3 Planinic, Maja D 10 18.2 Pleschke, Ferdinand EP 3.5 Piefka, Jan •MP 2.2 Plehn, Tilman T 2.6, T 2.7,T 2.8, T 31.2, T 86.7,
Nagataki, Shigehiro EP 1.4 Näger, Paul M. •AGPhil 9.2 namasaka, wilson •T 47.6 Nandi, Gerrit DD 5.2 Naranjo, Pedro AGPhil 1.1 Nardi, Roberto GP 2.3 Nardo, Lorenzo ST 9.2 Narimani Charan, Abtin •T 79.4 Nasr Esfahani, Alina •T 51.5 Nass, Christian •T 62.3 Nast, Natalie T 107.4 Nasts, Natalie T 107.4 Nattland, Philipp T 33.2 Navarro, Jaume •GP 6.2 Nawrodt, Ronny DD 5.2, DD 12.7, DD 31.1, DD 34.12, DD 36.5, DD 44.3 Nayaz, Abdullah •T 16.4 Necker, Jannis •T 18.2 Nedora, Vsevolod •GR 11.3 Negar, Farjoud Masouleh EP 2.2 Nell, Sebastian DD 12.8 Nemeth-Csoka, Marton •T 30.5 Netzwerk Teilchenwelt-Kollaboration T 24.3, T 24.1, T 52.1, T 52.4, T 24.2, T 52.3 Neuberger, Moritz •T 75.3 Neuendorffer, Stephen T 65.5, T 65.6, T 65.7, T 94.4 Neuhaus, Simon <td>Olschewsky, Kevin •• T 56.7 Omeliukh, Anastasiia •EP 7.10 Omidvari, Negar ST 9.2 on Envision team, VenSpec EP 3.1 on VERTIAS team, VEM EP 3.1 Öncel, Ö. Oğul T 8.4, T 8.5 Ongis, Peter T 60.3, T 60.4 Ördek, Serhat T 8.4, T 8.5 Ördek, Serhat T 8.9, 4 Orekhov, Vsevolod •T 102.9 Ørsøe, Rasmus •AKPIK 4.7 Ortjohann, Hans-Werner T 21.4, T 21.5 Ortmann, Julia •D 50.2 Ortmann, Tobias T 74.7, T 78.6, T 78.7, T 104.5, T 104.6, •T 104.7 Ossokine, Serguei Ossokine, Serguei GR 9.2 Osterhoff, Jens T 44.5, T 44.6 Othman, Gulden •T 38.4 Ott, Philipp T 5.8 Oughton, Sean EP 12.3 Owtscharenko, Nikolaus •T 42.6 P2-Kollaboration T 68.7 Paasch, Alexander •T 7.2 Paasch, Alexander •T 7.2 Paasch, Steven T 36.7, •T 85.1, •T 85.5 Page, Benjamin T 2.2</td> <td>Piergrossi, Joseph T 24.2 Piergo, Tanguy T 72.5, T 72.6, T 101.7 Pierre Auger-Kollaboration T 19.1, T 19.2, T 19.4, T 19.7, AKPIK 4.6, AKPIK 4.9, T 71.4, T 19.8, T 19.5, T 98.5, T 71.1, T 71.2, T 71.3, T 51.2, T 19.3, T 51.3, T 47.1, T 47.4, T 47.3, T 51.1, T 47.2 Pietroni, Massimo GR 6.2 Pilarczyk, Oliver T 48.5 Pimpalshende, Vivek GR 14.4, GR 14.5, •GR 14.6 Pinamonti, Nicola •GR 12.2, •MP 9.2 Piotter, Mona T 22.7 Pirinen, Pekka DD 36.1 Pirner, Hans Jürgen •AGPhil 12.3 Pirola, Giorgio T 17.8 Piscopo, Maria Laura T 56.4 Pitt, Michael T 7.5 Pittermann, Martin •ST 6.4 Pitz, Daniel T 12.8, T 12.9, T 30.8 Pivovarov, Alexei A T 56.5 Pizzella, Veronica •T 78.3 Planinic, Maja DD 18.2 Plaschke, Ferdinand EP 3.4, EP 3.5, Pleha, Jan •M 2.2, T 19.3, T 39.1, T 39.2, •T 31.2, T 86.7, T 90.3, T 93.1, T 93.4, T 93.5, AKPIK 4.2 Plesanovs, Vladislavs <</td>	Olschewsky, Kevin •• T 56.7 Omeliukh, Anastasiia •EP 7.10 Omidvari, Negar ST 9.2 on Envision team, VenSpec EP 3.1 on VERTIAS team, VEM EP 3.1 Öncel, Ö. Oğul T 8.4, T 8.5 Ongis, Peter T 60.3, T 60.4 Ördek, Serhat T 8.4, T 8.5 Ördek, Serhat T 8.9, 4 Orekhov, Vsevolod •T 102.9 Ørsøe, Rasmus •AKPIK 4.7 Ortjohann, Hans-Werner T 21.4, T 21.5 Ortmann, Julia •D 50.2 Ortmann, Tobias T 74.7, T 78.6, T 78.7, T 104.5, T 104.6, •T 104.7 Ossokine, Serguei Ossokine, Serguei GR 9.2 Osterhoff, Jens T 44.5, T 44.6 Othman, Gulden •T 38.4 Ott, Philipp T 5.8 Oughton, Sean EP 12.3 Owtscharenko, Nikolaus •T 42.6 P2-Kollaboration T 68.7 Paasch, Alexander •T 7.2 Paasch, Alexander •T 7.2 Paasch, Steven T 36.7, •T 85.1, •T 85.5 Page, Benjamin T 2.2	Piergrossi, Joseph T 24.2 Piergo, Tanguy T 72.5, T 72.6, T 101.7 Pierre Auger-Kollaboration T 19.1, T 19.2, T 19.4, T 19.7, AKPIK 4.6, AKPIK 4.9, T 71.4, T 19.8, T 19.5, T 98.5, T 71.1, T 71.2, T 71.3, T 51.2, T 19.3, T 51.3, T 47.1, T 47.4, T 47.3, T 51.1, T 47.2 Pietroni, Massimo GR 6.2 Pilarczyk, Oliver T 48.5 Pimpalshende, Vivek GR 14.4, GR 14.5, •GR 14.6 Pinamonti, Nicola •GR 12.2, •MP 9.2 Piotter, Mona T 22.7 Pirinen, Pekka DD 36.1 Pirner, Hans Jürgen •AGPhil 12.3 Pirola, Giorgio T 17.8 Piscopo, Maria Laura T 56.4 Pitt, Michael T 7.5 Pittermann, Martin •ST 6.4 Pitz, Daniel T 12.8, T 12.9, T 30.8 Pivovarov, Alexei A T 56.5 Pizzella, Veronica •T 78.3 Planinic, Maja DD 18.2 Plaschke, Ferdinand EP 3.4, EP 3.5, Pleha, Jan •M 2.2, T 19.3, T 39.1, T 39.2, •T 31.2, T 86.7, T 90.3, T 93.1, T 93.4, T 93.5, AKPIK 4.2 Plesanovs, Vladislavs <
Nagataki, Shigehiro $EP 1.4$ Näger, Paul M. \bullet AGPhil 9.2namasaka, wilson \bullet T 47.6Nandi, GerritDD 5.2Naranjo, PedroAGPhil 1.1Nardi, Roberto $GP 2.3$ Nardo, LorenzoST 9.2Narimani Charan, Abtin \bullet T 79.4Nasr Esfahani, Alina \bullet T 51.5Nass, Christian \bullet T 62.3Nast, NatalieT 107.4Nattland, PhilippT 33.2Navaro, Jaume \bullet GP 6.2Nawrodt, RonnyDD 5.2, DD 12.7,DD 31.1, DD 34.12, DD 36.5, DD 44.3Nayaz, Abdullah \bullet T 16.4Necker, Jannis \bullet T 18.2Nedora, Vsevolod \bullet GR 11.3Negar, Farjoud Masouleh $EP 2.2$ Nell, SebastianDD 12.8Neuberger, Moritz \bullet T 75.3Neuendorffer, StephenT 69.7Neufeld, MariusT 65.5, T 65.6,T 65.7, T 94.4 \bullet T 87.5Neundorf, Jonas \bullet T 87.5Neundorf, Jonas \bullet T 63.9Nguyen, Le HoangT 77.4, T 77.5Nguyen, Le HoangT 77.4, T 77.5Nguyen, Le HoangT 77.4, T 77.5Nguyen, Le HoangT 10.6Nickeler, Dieter \bullet EP 1.2Nicchciol, MarcusT 101.6Niechciol, MarcusT 101.6Niechciol, MarcusT 101.6Niechciol, MarcusT 103.5Neukire, Dieter \bullet EP 1.2Nicolini, PieroGR 3.4Nicchciol, MarcusT 101.6Niechciol, MarcusT 101.6N	Olschewsky, Kevin •• T 56.7 Omeliukh, Anastasiia •EP 7.10 Omidvari, Negar ST 9.2 on Envision team, VenSpec EP 3.1 on VERTIAS team, VEM EP 3.1 Öncel, Ö. Oğul T 8.4, T 8.5 Onyisi, Peter T 60.3, T 60.4 Ördek, Serhat T 89.4 Orekhov, Vsevolod •T 102.9 Ørsøe, Rasmus •AKPIK 4.7 Ortjohann, Hans-Werner T 21.4, T 21.5 Ortmann, Julia 0D 50.2 Ortmann, Julia 0D 50.2 Ortmann, Tobias T 74.7, T 78.6, T 78.7, T 104.5, T 104.6, •T 104.7 Ossokine, Serguei Ogg 9.2 Osterhoff, Jens T 44.5, T 44.6 Othman, Gulden •T 38.4 Ott, Philipp T 5.8 Oughton, Sean EP 12.3 Owtscharenko, Nikolaus •T 42.6 P2-Kollaboration T 68.7 Paasch, Alexander •T 7.2 Paasch, Alexander •T 7.2 Paasch, Steven T 36.7, •T 85.1, •T 85.5 Page, Benjamin T 2.2 Paino, Jason ST 3.4 <	Piergrossi, Joseph T 24.2 Piergo, Tanguy T 72.5, T 72.6, T 101.7 Pierre Auger-Kollaboration T 19.1, T 19.2, T 19.4, T 19.7, AKPIK 4.6, AKPIK 4.9, T 71.4, T 19.8, T 19.5, T 98.5, T 71.1, T 71.2, T 71.3, T 51.2, T 19.3, T 51.3, T 47.1, T 47.4, T 47.3, T 51.1, T 47.2 Pietroni, Massimo GR 6.2 Pilarczyk, Oliver •T 48.5 Pimpalshende, Vivek GR 14.4, GR 14.5, •GR 14.6 Finamonti, Nicola Piotter, Mona •T 22.7 Pirinen, Pekka D 0.61 Pirrola, Giorgio •T 17.8 Pirsch, Johannes •MP 2.6 Piscopo, Maria Laura T 56.4 Pittermann, Martin •ST 6.4 Pitzl, Daniel T 12.8, T 12.9, T 30.8 Pivovarov, Alexei A. T 56.5 Pizzella, Veronica •T 78.3 Plaschke, Ferdinand EP 3.4, EP 3.5 Plefka, Jan •MP 2.2 Plesanovs, Vladislavs T 39.1, T 39.2, •T 39.3 Plicht, Katja •DD 34.11 Plotko, Pavlo •EP 7.2 Pöder, Kristjan T 44.5, T 44.6 Podszus,
Nagataki, Shigehiro EP 1.4 Näger, Paul M. •AGPhil 9.2 namasaka, wilson •T 47.6 Nandi, Gerrit DD 5.2 Naranjo, Pedro AGPhil 9.2 Naranjo, Pedro AGPhil 1.1 Nardi, Roberto GP 2.3 Nardo, Lorenzo ST 9.2 Narimani Charan, Abtin •T 79.4 Nasr Esfahani, Alina •T 51.5 Nass, Christian •T 62.3 Nast, Natalie T 107.4 Nattland, Philipp T 33.2 Navarro, Jaume •GP 6.2 Nawrodt, Ronny DD 5.2, DD 12.7, DD 31.1, DD 34.12, DD 36.5, DD 44.3 Nayaz, Abdullah •T 16.4 Necker, Jannis •T 18.2 Nedora, Vsevolod •GR 11.3 Negar, Farjoud Masouleh EP 2.2 Nell, Sebastian DD 12.8 Neuberger, Moritz •T 75.3 Neuberger, Moritz •T 75.3 Neuberger, Moritz •T 65.5, T 65.6, T 65.7, T 94.4 Neubaus, Simon •T 34.6 Neukirch, Thomas EP 12.0 Neukum, Max •T 87.5 Neudorff, Jonas	Olschewsky, Kevin •• T 56.7 Omeliukh, Anastasiia •EP 7.10 Omidvari, Negar ST 9.2 on Envision team, VenSpec EP 3.1 on VERTIAS team, VEM EP 3.1 Öncel, Ö. Oğul T 8.4, T 8.5 Onyisi, Peter T 60.3, T 60.4 Ördek, Serhat T 89.4 Orekhov, Vsevolod •T 102.9 Ørsøe, Rasmus •AKPIK 4.7 Ortjohann, Hans-Werner T 21.4, T 21.5 Ortmann, Julia >DD 50.2 Ortmann, Tobias T 74.7, T 78.6, T 78.7, T 104.5, T 104.6, •T 104.7 Ossokine, Serguei Ossokine, Serguei GR 9.2 Osterhoff, Jens T 44.5, T 44.6 Otthman, Gulden •T 38.4 Ott, Philipp T 5.8 Oughton, Sean EP 12.3 Owtscharenko, Nikolaus •T 42.6 P2-Kollaboration T 68.7 Paasch, Steven T 36.7, •T 85.1, •T 85.5 Page, Benjamin T 2.2 Paino, Jason ST 3.4 Pak, Sang-II T 3.2 Palivya •T 39.8, T 68.1, T 97.1	Piergrossi, Joseph T 24.2 Piergossi, Joseph T 24.2 Piergossi, Joseph T 24.2 Piergossi, Joseph T 25, T 72.6, T 101.7 Pierre Auger-Kollaboration T 19.1, T 19.2, T 19.4, T 19.7, AKPIK 4.6, AKPIK 4.9, T 71.4, T 19.8, T 19.5, T 98.5, T 71.1, T 71.2, T 71.3, T 51.2, T 19.3, T 51.3, T 47.1, T 47.4, T 47.3, T 51.1, T 47.2 Pietroni, Massimo GR 6.2 Pilarczyk, Oliver •T 48.5 Pimpalshende, Vivek GR 14.4, GR 14.5, •GR 14.6 Pinamonti, Nicola •GR 12.2, •MP 9.2 Piotter, Mona •T 22.7 Pirinen, Pekka DD 36.1 Pirner, Hans Jürgen •AGPhil 12.3 Piroch, Giorgio •T 17.8 Pirsch, Johannes •MP 2.6 Piscopo, Maria Laura T 56.4 Pitt, Michael T 7.5 Pitzella, Veronica •T 78.3 Plaschke, Ferdinand EP 3.4, EP 3.5 Plefka, Jan •MP 2.2 Plehn, Tilman T 2.6, T 2.7, T 2.8, T 31.2, T 86.7, T 90.3, T 93.1, T 93.4, T 39.3 Plicht, Katja •DD 34.11 Plottk, Pavlo •EP 7.2 Pöder, Kristjan T 44.5, T 44.6 Podsus, T 92.3

GR 14.5, GR 14.6, T 65.1, T 65.2,
T 65.3, T 65.4, T 68.8, T 94.6
Popov, Ivan•T 68.6
Popp, Jennifer •T 44.5, T 44.6
Porod, Werner
Potzel, Walter T 104.5, T 104.6,
Pozdnyakov, Andrey T 26.3, T 89.1,
•T 89.2, T 90.6
Prechelt. Remv
Preiß, Stephan
Preitnacher, Jonathan •ST 9.5, T 66.4
•T 79.1
Priemer, Burkhard DD 4.1
Prim, Markus . 1 16.6, 1 30.6, •1 64.7,
Pröbst, Franz
Probst, Uwe
Protect 8-Kollaboration T 49 1 T 49 2
T 49.3, T 49.4
Prokop, Axel-Thilo
Prottung. Janik
Prouvé, Claire T 26.6
Prouvost, Nathan
Pueschel, M.J
Pump, Kristin
PUNCH4NFDI-Kollaboration
Pürrer, Michael GR 9.2
Pusch, Alexander DD 13.3, DD 32.1
Pyras, Lilly
Qin, Liqing
Quadfasel, Tobias
T 12.3. T 12.5. T 12.6. T 14.6. T 33.4.
Т 33.6, Т 33.8, Т 35.5, Т 70.4, Т 88.5,
•DD 3.1 Ouartin Miguel •GR 6.5
Quast Guenter T 251
quuer, euclider frifficher i zeit
Quast, Günter T 15.3, T 15.4, T 15.6,
Quast, Günter T 15.3, T 15.4, T 15.6, T 15.7, T 25.2, T 32.3, T 32.5, T 36.4, T 36.5, DD 35.6, DD 36.3
Quast, Günter T 15.3, T 15.4, T 15.6, T 15.7, T 25.2, T 32.3, T 32.5, T 36.4, T 36.5, DD 35.6, DD 36.3 Quick, Thomas
Quast, Günter T 15.3, T 15.4, T 15.6, T 15.7, T 25.2, T 32.3, T 32.5, T 36.4, T 36.5, DD 35.6, DD 36.3 Quick, ThomasDD 19.4 Quintana, ChristianST 9.3 Pabbetz, Klaus T 32 3, T 32 5
Quast, Günter T 15.3, T 15.4, T 15.6, T 15.7, T 25.2, T 32.3, T 32.5, T 36.4, T 36.5, DD 35.6, DD 36.3 Quick, ThomasDD 19.4 Quintana, ChristianST 9.3 Rabbertz, KlausT 32.3, T 32.5 Rabe, ChristianDD 34.9
Quast, Günter T 15.3, T 15.4, T 15.6, T 15.7, T 25.2, T 32.3, T 32.5, T 36.4, T 36.5, DD 35.6, DD 36.3 Quick, ThomasDD 19.4 Quintana, ChristianST 9.3 Rabbertz, KlausT 32.3, T 32.5 Rabe, ChristianDD 34.9 Rabe, ThoridDD 35.8
Quast, Günter T 15.3, T 15.4, T 15.6, T 15.7, T 25.2, T 32.3, T 32.5, T 36.4, T 36.5, DD 35.6, DD 36.3 Quick, Thomas
Quast, Günter T 15.3, T 15.4, T 15.6, T 15.7, T 25.2, T 32.3, T 32.5, T 36.4, T 36.5, DD 35.6, DD 36.3 Quick, Thomas DD 19.4 Quintana, Christian ST 9.3, T 32.5, T 36.4, T 32.5, T 36.4, T 32.3, T 32.5 Rabbertz, Klaus DD 34.9 Rabe, Christian DD 34.9 Rabe, Thorid DD 35.8 Rabusov, Andrei T 58.7 Rachamin, Reuven T 96.4 Rack-Helleis, John T 23.5, •T 23.6,
Quast, Günter T 15.3, T 15.4, T 15.6, T 15.7, T 25.2, T 32.3, T 32.5, T 36.4, T 36.5, DD 35.6, DD 36.3 Quick, Thomas DD 19.4 Quintana, Christian ST 9.3 Rabbertz, Klaus T 32.3, T 32.5, T 36.4, T 32.3, T 32.5 Rabe, Christian DD 34.9 Rabe, Christian DD 35.8 Rabusov, Andrei T 58.7 Rachamin, Reuven T 96.4 T 93.7 Radermacher, Thomas T 39.5, T 39.5, T 39.6, T 23.5, •T 23.6,
Quast, Günter T 15.3, T 15.4, T 15.6, T 15.7, T 25.2, T 32.3, T 32.5, T 36.4, T 36.5, DD 35.6, DD 36.3 Quick, Thomas DD 19.4 Quintana, Christian DD 19.4 Quintana, Christian DD 35.8, DD 35.8, DD 36.9, T 32.3, T 32.5, T 36.4, T 32.3, T 32.5, T 32.3, T 32.5, T 32.3, T 32.5, T 32.3, T 32.5, Rabusov, Andrei DD 34.9 Rabe, Christian DD 34.9 Rabe, Thorid DD 35.8 Rabusov, Andrei T 58.7 Rachamin, Reuven T 96.4 Rack-Helleis, John T 23.5, •T 23.6, T 93.7 Radermacher, Thomas T 39.5, T 39.6, T 39.7, *T 102.8 Padlkheureri Vacarri T 44.2, T 00.9 T 44.2, T 00.9
Quast, Günter T 15.3, T 15.4, T 15.6, T 15.7, T 25.2, T 32.3, T 32.5, T 36.4, T 36.5, DD 35.6, DD 36.3 Quick, Thomas DD 19.4 Quintana, Christian ST 9.3 Rabbertz, Klaus T 32.3, T 32.5, T 36.4, T 32.3, T 32.5 Raber, Christian ST 9.3 Rabe, Christian DD 34.9 Rabe, Thorid DD 35.8 Rabusov, Andrei T 58.7 Rack-Helleis, John T 23.5, •T 23.6, T 93.7 Radermacher, Thomas T 39.5, T 39.6, T 39.7, •T 102.8 Radkhorrami, Yasser T 44.3, T 90.8 Rafanoharana, Dimbiniaina •T 73.8
Quast, Günter T 15.3, T 15.4, T 15.6, T 15.7, T 25.2, T 32.3, T 32.5, T 36.4, T 36.5, DD 35.6, DD 36.3 Quick, Thomas DD 19.4 Quintana, Christian ST 9.3 Rabbertz, Klaus D3.7, 32.5, T 36.4, T 32.3, T 32.5, T 36.4, T 32.3, T 32.5 Rabertz, Klaus D1 34.9 Rabe, Christian DD 34.9 Rabe, Thorid DD 35.8 Rabusov, Andrei T 58.7 Rack-Helleis, John T 23.5, •T 23.6, T 93.7 Radermacher, Thomas T 39.5, T 39.6, T 39.7, T 102.8 Radahorrami, Yasser T 44.3, T 90.8 Rafanoharana, Dimbiniaina T 7.3 Rafeek, Rufa T 88.7, T 88.8, T 88.9
Quast, Günter T 15.3, T 15.4, T 15.6, T 15.7, T 25.2, T 32.3, T 32.5, T 36.4, T 36.5, DD 35.6, DD 36.3 Quick, Thomas
Quast, Günter T 15.3, T 15.4, T 15.6, T 15.7, T 25.2, T 32.3, T 32.5, T 36.4, T 36.5, DD 35.6, DD 36.3 Quick, Thomas •DD 19.4 Quintana, Christian ST 9.3, T 32.5, T 36.4, T 32.3, T 32.5, T 36.4, T 3
Quast, Günter T 15.3, T 15.4, T 15.6, T 15.7, T 25.2, T 32.3, T 32.5, T 36.4, T 36.5, DD 35.6, DD 36.3 Quick, Thomas DD 19.4 Quintana, Christian ST 9.3, Rabbertz, Klaus T 32.3, T 32.5 Rabe, Christian DD 34.9 Rabe, Thorid DD 34.9 Rabe, Thorid DD 34.9 Rabe, Thorid DD 34.9 Rabe, Thorid
Quast, Günter T 15.3, T 15.4, T 15.6, T 15.7, T 25.2, T 32.3, T 32.5, T 36.4, T 36.5, DD 35.6, DD 36.3 Quick, Thomas DD 19.4 Quintana, Christian ST 9.3, Rabbertz, Klaus T 32.3, T 32.5 Rabe, Christian DD 34.9 Rabe, Thorid DD 34.9 Rabe, Thorid DD 34.9 Rabe, Thorid DD 35.8 Rabusov, Andrei T 58.7 Rachamin, Reuven T 96.4 Rack-Helleis, John T 23.5, •T 23.6, T 93.7 Radermacher, Thomas T 39.5, T 39.6, T 39.7, •T 102.8 Radkhorrami, Yasser T 44.3, T 90.8 Rafanoharana, Dimbiniaina•T 7.3 Rafeek, Rufa•T 88.7, T 88.8, T 88.9 Rahman, Ninoy F 56.8 Rahman, Ninoy
Quast, Günter T 15.3, T 15.4, T 15.6, T 15.7, T 25.2, T 32.3, T 32.5, T 36.4, T 36.5, DD 35.6, DD 36.3 Quick, Thomas
Quast, Günter T 15.3, T 15.4, T 15.6, T 15.7, T 25.2, T 32.3, T 32.5, T 36.4, T 36.5, DD 35.6, DD 36.3 Quick, ThomasDD 19.4 Quintana, ChristianST 9.3, Rabbertz, KlausT 32.3, T 32.5 Rabe, ChristianDD 34.9 Rabe, ThoridDD 34.9 Rabe, ThoridDD 35.8 Rabusov, AndreiT 32.3, T 32.5 Rachamin, ReuvenT 96.4 Rack-Helleis, JohnT 23.5, •T 23.6, T 93.7 Radermacher, Thomas T 39.5, T 39.6, T 39.7, •T 102.8 Radkhorrami, YasserT 44.3, T 90.8 Rafanoharana, DimbiniainaT 7.3 Rafeek, Rufa•T 88.7, T 88.8, T 88.9 Rahimi, MuslemT 56.8 Rahman, Ninoy
Quast, Günter T 15.3, T 15.4, T 15.6, T 15.7, T 25.2, T 32.3, T 32.5, T 36.4, T 36.5, DD 35.6, DD 36.3 Quick, Thomas DD 19.4 Quintana, Christian ST 9.3, ST 32.5, T 36.4, T 32.3, T 32.5, T 36.4,
Quast, Günter T 15.3, T 15.4, T 15.6, T 15.7, T 25.2, T 32.3, T 32.5, T 36.4, T 36.5, DD 35.6, DD 36.3 Quick, Thomas DD 19.4 Quintana, Christian ST 9.3 Rabbertz, Klaus T 32.3, T 32.5 Rabe, Christian DD 34.9 Rabe, Christian DD 34.9 Rabe, Christian DD 35.8 Rabe, Christian DD 35.8 Rabusov, Andrei T 58.7, T 23.6, T 23.6, T 93.7 Rack-Helleis, John T 23.5, •T 23.6, T 93.7 Radermacher, Thomas T 39.5, T 39.6, T 39.7, •T 102.8 Radkhorrami, Yasser T 44.3, T 90.8 Rafanoharana, Dimbiniaina T 7.3 Rafeek, Rufa •T 88.7, T 88.8, T 88.9 Rahimi, Muslem •T 56.6 Rahman, Ninoy •E 15.5 Raine, Johnny •T 33.3 Ramachandran, Shivani •T 86.7, T 86.7 Ramme, Fabian •D 7.2 Ramos-buades, antoni •G R 13.3 Rani, Erika • T 86.2, •T 86.3 Ramy, John •G Phil 9.4 Ratit, Alireza •G R 13.2 Ratki, Alireza •G R
Quast, Günter T 15.3, T 15.4, T 15.6, T 15.7, T 25.2, T 32.3, T 32.5, T 36.4, T 36.5, DD 35.6, DD 36.3 Quick, ThomasDD 19.4 Quintana, ChristianST 9.3 Rabbertz, Klaus T 32.3, T 32.5 Rabe, ChristianDD 34.9 Rabe, ThoridDD 34.9 Rabe, ThoridDD 35.8 Rabusov, AndreiT 32.3, T 25.7 Rachamin, ReuvenT 96.4 Rack-Helleis, JohnT 23.5, •T 23.6, T 93.7 Radermacher, Thomas T 39.5, T 39.6, T 39.7, •T 102.8 Radkhorrami, YasserT 44.3, T 90.8 Rafanoharana, DimbiniainaT 7.3 Rafeek, Rufa•T 88.7, T 88.8, T 88.9 Rahimi, MuslemT 58.7, T 88.8, T 88.9 Rahimi, MuslemT 58.7, T 88.8, T 88.9 Rahima, Ninoy
Quast, Günter T 15.3, T 15.4, T 15.6, T 15.7, T 25.2, T 32.3, T 32.5, T 36.4, T 36.5, DD 35.6, DD 36.3 Quick, Thomas DD 19.4 Quintana, Christian ST 9.3 Rabbertz, Klaus T 32.3, T 32.5 Rabe, Christian Rabe, Christian DD 34.9 Rabe, Christian T 32.3, T 32.5 Rabe, Christian DD 34.9 Rabusov, Andrei T 58.7 Rachamin, Reuven T 58.7 Rachamin, Reuven T 96.4 Radermacher, Thomas T 39.5, T 39.6, T 39.5, T 39.6, T 39.7, •T 102.8 Radathorrami, Yasser T 44.3, T 90.8 Rafanoharana, Dimbiniaina T 7.3 Rafeek, Rufa T 88.7, T 88.8, T 88.9 Rahimi, Muslem T 56.8 Rahman, Ninoy T 86.2, •T 86.3 Ramosi, Burim T 25.4 ra
Quast, Günter T 15.3, T 15.4, T 15.6, T 15.7, T 25.2, T 32.3, T 32.5, T 36.4, T 36.5, DD 35.6, DD 36.3 Quick, Thomas DD 19.4 Quintana, Christian DD 34.9 Rabbertz, Klaus T 32.3, T 32.5 T 36.5, DD 35.8 Rabe, Christian DD 34.9 Rabusov, Andrei T 58.7 Rachamin, Reuven Rachamin, Reuven T 96.4 Rack-Helleis, John T 39.7, achamin, Reuven T 23.5, •T 23.6, T 39.5, T 39.6, T 39.7, •T 102.8 Radermacher, Thomas T 39.5, T 39.6, T 39.7, •T 102.8 Radarnoharana, Dimbiniaina Rafanoharana, Dimbiniaina T 68.8, T 88.9 Rahimi, Muslem T 56.8 Rahman, Ninoy T 98.7 R 33.3 Ramosaj, Burim T 86.2, *T 86.3 R 33.3 Rame, Fabian G R 13.2 Raine, Johnny
Quast, Günter T 15.3, T 15.4, T 15.6, T 15.7, T 25.2, T 32.3, T 32.5, T 36.4, T 36.5, DD 35.6, DD 36.3 Quick, Thomas DD 19.4 Quintana, Christian DD 34.9 Rabbertz, Klaus T 32.3, T 32.5 Rabe, Christian DD 34.9 Rabe, Christian DD 34.9 Rabe, Christian DD 34.9 Rabusov, Andrei T 58.7, T 23.6, T 23.6, T 39.6, T 39.7, Rachamin, Reuven T 96.4 Rack-Helleis, John T 23.5, •T 23.6, T 39.7, T 102.8 Radermacher, Thomas Radermacher, Thomas T 39.7, •T 102.8 Radakhorrami, Yasser T 44.3, T 90.8 Rafanoharana, Dimbiniaina •T 7.3 Rafeek, Rufa •T 88.7, T 88.8, T 88.9 Rahimi, Muslem •T 56.8 Rahman, Ninoy EP 1.5 Raine, Johnny •T 98.7, T 88.8, T 88.9 Rahman, Ninoy EP 1.5 Rame, Fabian •D 0.7.2 Ramosaj, Burim •T 25.4 Ramosaj, Burim •T 28.4 •G R 13.3 Rarity, John •G R 13.2 Rahit, Alireza GR 13.2 Ratke, Alexander •ST 2.3, ST 3.1 Rätz, Johanna •D 12.2, DD 12.4, DD 12.4, DD 12.5, DD 12.6 </td
Quast, Günter T 15.3, T 15.4, T 15.6, T 15.7, T 25.2, T 32.3, T 32.5, T 36.4, T 36.5, DD 35.6, DD 36.3 Quick, Thomas Quintana, Christian Rabbertz, Klaus Rabe, Christian Rabe, Christian Rabe, Christian Rabe, Christian Rabusov, Andrei Rack-Helleis, John T 23.5, •T 23.6, T 39.6, T 39.7, •T 102.8 Radermacher, Thomas T 39.5, T 39.6, T 39.7, •T 102.8 Rafanoharana, Dimbiniaina Rafanoharana, Dimbiniaina Rafanoharana, Ninoy Rahman, Ninoy Ramosaj, Burim Ramosaj, Burim Ramosaj, Burim
Quast, Günter T 15.3, T 15.4, T 15.6, T 15.7, T 25.2, T 32.3, T 32.5, T 36.4, T 36.5, DD 35.6, DD 36.3 Quick, Thomas DD 19.4 Quintana, Christian DD 19.4 Quintana, Christian DD 34.9 Rabe, Christian DD 34.9 Rabe, Christian DD 35.8 Rabe, Christian DD 35.8 Rabusov, Andrei T 58.7 Rachamin, Reuven T 96.4 Rack-Helleis, John T 23.5, •T 23.6, T 39.6, T 39.7, •T 102.8 Radermacher, Thomas T 39.5, T 39.6, T 39.7, •T 102.8 Rafanoharana, Dimbiniaina Rafanoharana, Dimbiniaina T 7.3 Rafanoharana, Dimbiniaina T 7.3 Rafanoharana, Ninoy F 15.5 Raine, Johnny T 33.3 Ramachandran, Shivani T 98.6 Ramme, Fabian <
Quast, Günter T 15.3, T 15.4, T 15.6, T 15.7, T 25.2, T 32.3, T 32.5, T 36.4, T 36.5, DD 35.6, DD 36.3 Quick, Thomas DD 19.4 Quintana, Christian ST 9.3 Rabbertz, Klaus ST 9.3 Rabe, Christian DD 34.9 Rabe, Christian DD 35.8 Rabe, Christian DD 35.8 Rabusov, Andrei Rack-Helleis, John T 23.5, •T 23.6, T 39.6, T 39.7, •T 102.8 Radermacher, Thomas T 39.5, T 39.6, T 39.7, •T 102.8 Radarmachar, T 38.8, T 88.9 Rafanoharana, Dimbiniaina Rafanoharana, Dimbiniaina Rafanoharana, Ninoy Rahman, Ninoy Ramosaj, Burim Ramosaj, Burim Ramosaj, Burim
Quast, Günter T 15.3, T 15.4, T 15.6, T 15.7, T 25.2, T 32.3, T 32.5, T 36.4, T 36.5, DD 35.6, DD 36.3 Quick, Thomas DD 19.4 Quintana, Christian ST 9.3 Rabbertz, Klaus ST 9.3 Rabe, Christian DD 19.4 Rabe, Christian DD 35.8 Rabe, Christian DD 35.8 Rabusov, Andrei T 58.7 Rachamin, Reuven T 96.4 Rack-Helleis, John T 23.5, •T 23.6, T 39.7 Radermacher, Thomas T 39.5, T 39.6, T 39.7, •T 102.8 Radkhorrami, Yasser T 44.3, T 90.8 Rafanoharana, Dimbiniaina T 7.3 Rafanoharana, Dimbiniaina T 7.3 Rafanoharana, Shivani T 88.8, T 88.9 Rahmi, Muslem T 86.8, T 83.3 Rather, Johnny T 33.3 Ramachandran, Shivani
Quast, Günter T 15.3, T 15.4, T 15.6, T 15.7, T 25.2, T 32.3, T 32.5, T 36.4, T 36.5, DD 35.6, DD 36.3 Quick, Thomas DD 19.4 Quintana, Christian DD 19.4 Quintana, Christian DD 34.9 Rabbertz, Klaus T 32.3, T 32.5 Rabe, Christian DD 35.8 Rabusov, Andrei T 58.7 Rachamin, Reuven T 96.4 Rack-Helleis, John T 23.5, •T 23.6, T 93.7 Radermacher, Thomas T 39.5, T 39.6, T 39.7, •T 102.8 Radkhorrami, Yasser T 44.3, T 90.8 Rafanoharana, Dimbiniaina Rafanoharana, Dimbiniaina T 7.3 Rafeek, Rufa T 88.8, T 88.9 Rahman, Ninoy EP 1.5 Raine, Johnny T 33.3 Ramchandran, Shivani T 98.2, •T 88.3 Ra13.3 Rating, Johnny T 86.2, •T 86.3 Ra11.2 Ratek, Alexander S 37.3 Ratek, Alex
Quast, Günter T 15.3, T 15.4, T 15.6, T 15.7, T 25.2, T 32.3, T 32.5, T 36.4, T 36.5, DD 35.6, DD 36.3 Quick, Thomas DD 19.4 Quintana, Christian ST 9.3, Rabbertz, Klaus T 32.3, T 32.5 Rabe, Christian DD 34.9 Rabe, Thorid DD 34.9 Rabe, Thorid DD 34.9 Rabe, Thorid DD 34.9 Rabe, Christian DD 34.9 Rabe, Thorid DD 35.8 Rabe, Christian DD 34.9 Rabe, Thorid DD 35.8 Rack-Helleis, John T 23.5, •T 23.6, T 93.7 Radermacher, Thomas T 39.5, T 39.6, T 39.7, •T 102.8 Radkhorrami, Yasser T 44.3, T 90.8 Rafanoharana, Dimbiniaina •T 7.3 Rafeek, Rufa •T 88.7, T 88.8, T 88.9 Rahimi, Muslem •T 56.8 Rahman, Ninoy EP 1.5 Raine, Johnny T 25.4 ramos-buades, antoni •GR 13.3 Rani, Erika T 86.2, •T 86.3 Rarity, John AGPhil 9.4 Rashti, Alireza GR 13.2 Ratke, Alexander •ST 2.3, ST 3.1 Rätzel, Dennis •GR 4.2, GR 6.4 Rauber, Saskia •DD 12.2, DD 12.4, DD 12.5, DD 12.6 Rätzel, Dennis •GR 4.2, GR 6.4 Rauber, Saskia •DD 35.3 Raudzis, Carsten •DD 5.2 Rauls, Noreen
Quast, Günter T 15.3, T 15.4, T 15.6, T 15.7, T 25.2, T 32.3, T 32.5, T 36.4, T 36.5, DD 35.6, DD 36.3 Quick, Thomas DD 19.4 Quintana, Christian ST 9.3 Rabbertz, Klaus T 32.3, T 32.5 Rabe, Christian DD 34.9 Rabe, Thorid DD 34.9 Rabe, Thorid DD 35.8 Rabusov, Andrei T 58.7, T 36.4, T 93.7 Rachamin, Reuven T 96.4 Rack-Helleis, John T 23.5, •T 23.6, T 93.7 Radermacher, Thomas T 39.5, T 39.6, T 39.7, •T 102.8 Radkhorrami, Yasser T 44.3, T 90.8 Rafanoharana, Dimbiniaina •T 7.3 Rafeek, Rufa •T 88.7, T 88.8, T 88.9 Rahimi, Muslem •T 58.7, T 88.8, T 88.9 Rahimi, Muslem •T 58.7, T 88.8, T 88.9 Rahimi, Muslem •T 96.6 Ramme, Fabian •DD 7.2, Ramos-buades, antoni •GR 13.3 Rani, Erika •T 86.2, •T 86.3 Rarity, John •GR 13.2 Ratke, Alexander •ST 2.3, ST 3.1 Rätze, Johanna •DD 12.2, DD 12.4, DD 12.5, DD 12.6 Rätzel, Dennis •GR 4.2, GR 6.4 Rauber, Saskia DD 5.2 Rauls, Noreen *T 84.9 Rautenberg, Julian *T 47.4 Recchia, Sarah
Quast, Günter T 15.3, T 15.4, T 15.6, T 15.7, T 25.2, T 32.3, T 32.5, T 36.4, T 36.5, DD 35.6, DD 36.3 Quick, Thomas DD 19.4 Quintana, Christian DJ 34.9 Rabe, Christian DD 34.9 Rabe, Christian DD 34.9 Rabe, Christian DD 34.9 Rabe, Christian DD 35.8 Rabe, Christian DD 35.8 Rabusov, Andrei T 58.7, T 23.6, T 23.6, T 93.7 Rack-Helleis, John T 23.5, •T 23.6, T 93.7, •T 102.8 Radkmorani, Yasser T 44.3, T 90.8 Rafanoharana, Dimbiniaina T 7.3 Rafeek, Rufa •T 88.7, T 88.8, T 88.9 Rahimi, Muslem •T 56.6 Rahman, Ninoy •EP 1.5 Raine, Johnny 33.3 Ramachandran, Shivani •T 86.3 Rame, Fabian •DD 7.2 Ramos-buades, antoni @R 13.3 Raity, John AGPhil 9.4 Ratit, Alireza @R 13.2 Raity, John AGPhil 9.4 Ratit, Alireza @R 13.2 Raity, John PD 12.2, DD 12.4, DD 12.2, DD 12.4, DD 12.2, DD 12.4

Reiber, Thomas•GR 1.3
Reichel Niklas DD 14 1
Deichenhach Leenhard T 442
Reichenbach, Leonnaid•1 44.5
Reichert, Paula
Reichherzer. Patrick •T 73.3. T 99.6.
T 101 1 T 101 2
Reid, Jack
Reidelstürz, Joshua•T 6.2
Reif Markus T 30.3
Deifenberger Andress T1026
Relienberger, Andreas 1 102.0
Reimann, René
Reimer Anita FP 7.5
Deimero Arno Christenh T 20 2
Reimers, Ame Chinstophi•1 20.2
Rein, Constantin GR 3.2
Reindl, Florian
Reinecke Mathias T 41 3
Reinnardt, Christoph•GR 14.1
Reinhold, Peter DD 34.6, DD 34.7
Reininghaus Maximilian T 72.6
Reinsch, Tobias•DD 34.12
Reith, Aaron
Reithler Hans T 67 1
Rej, Amartya 1 33.9
Relancio, Javier GR 11.5
Rendel, Marian • T 9.4. T 67.3. T 67.4.
•T 67 5 T 67 6
Relikitoff, Saran
Renner, Frederic•T 70.5
Renz Roland •MP 8.1
Penik Matei T142 T 45 2
Repik, Malej 1 14.3, 1 05.2
Repke, Julian DD 20.3
Resconi Elisa T 46 8 T 71 9 •T 109 3
Deusch Simeon
Revathi, Purnalingam
Reville, Brian MP 11.1, T 45.1, T 73.7,
T 73 8 T 101 5
Densei Estehanen Mahaan TCO
Rezael Estabragh, Monsen 1 6.2
Riahinia, Ali T 105.10
Richter Christiane •DD 201
RIEDISCH, Matthias AKPIK 2.9
Rieck, Ina DD 5.2
Rieck Patrick T 67 3 T 67 4 T 67 5
т 67 6
1 0/.0
Rieger, Marcel
Rieger Peter DD 35 7 DD 35 10
Diegoh Johann T1047
Riescii, Jonann
Riese, JosefDD 49.1, DD 49.2
Riesen, Timm
Rievers Benny GR 2.8
Rievers, Benny GR 2.8
Rievers, Benny GR 2.8 Rifai, Mariam . T 20.2, •T 20.5, T 43.1,
Rievers, Benny GR 2.8 Rifai, Mariam . T 20.2, •T 20.5, T 43.1, T 48.4
Rievers, Benny GR 2.8 Rifai, Mariam . T 20.2, •T 20.5, T 43.1, T 48.4 Rimoldi, Marco •T 76.4
Rievers, Benny
Rievers, Benny
Rievers, Benny GR 2.8 Rifai, Mariam . T 20.2, •T 20.5, T 43.1, T 48.4 Rimoldi, Marco •T 76.4 Rincke, Philipp •T 69.9 Ringwald, Andreas GR 14.1
Rievers, Benny GR 2.8 Rifai, Mariam . T 20.2, •T 20.5, T 43.1, T 48.4 Rimoldi, Marco Rincke, Philipp •T 76.4 Ringwald, Andreas GR 14.1 Rink, Thomas •T 74.2
Rievers, Benny GR 2.8 Rifai, Mariam .T 20.2, •T 20.5, T 43.1, T 48.4 Rimoldi, Marco •T 76.4 Rincke, Philipp •T 69.9 Ringwald, Andreas GR 14.1 Rink, Thomas •T 74.2 Rinnagel, Maximilian T 11.1, T 11.2
Rievers, Benny GR 2.8 Rifai, Mariam . T 20.2, •T 20.5, T 43.1, T 48.4 Rimoldi, Marco •T 76.4 Rincke, Philipp •T 69.9 Ringwald, Andreas GR 14.1 Rink, Thomas •T 74.2 Rinnagel, Maximilian T 114 T 114
Rievers, Benny GR 2.8 Rifai, Mariam . T 20.2, •T 20.5, T 43.1, T 48.4 Rimoldi, Marco •T 76.4 Rincke, Philipp •T 69.9 Ringwald, Andreas GR 14.1 Rink, Thomas •T 74.2 Rinnagel, Maximilian T 11.1, T 11.2, T 11.3, T 11.4, •T 11.5, T 11.6, T 97.7
Rievers, Benny GR 2.8 Rifai, Mariam . T 20.2, •T 20.5, T 43.1, T 48.4 Rimoldi, Marco •T 76.4 Rincke, Philipp •T 69.9 Ringwald, Andreas GR 14.1 Rink, Thomas •T 74.2 Rinnagel, Maximilian T 11.1, T 11.2, T 11.3, T 11.4, •T 11.5, T 11.6, T 97.7 Risse, Markus
Rievers, Benny GR 2.8 Rifai, Mariam . T 20.2, •T 20.5, T 43.1, T 48.4 Rimoldi, Marco •T 76.4 Rincke, Philipp •T 69.9 Ringwald, Andreas GR 14.1 Rink, Thomas •T 74.2 Rinnagel, Maximilian T 11.1, T 11.2, T 11.3, T 11.4, •T 11.5, T 11.6, T 97.7 Risse, Markus
Rievers, Benny GR 2.8 Rifai, Mariam . T 20.2, •T 20.5, T 43.1, T 48.4 Rimoldi, Marco •T 76.4 Rincke, Philipp •T 69.9 Ringwald, Andreas GR 14.1 Rink, Thomas •T 74.2 Rinnagel, Maximilian T 11.1, T 11.2, T 11.3, T 11.4, •T 11.5, T 11.6, T 97.7 Risse, Markus
Rievers, Benny GR 2.8 Rifai, Mariam . T 20.2, •T 20.5, T 43.1, T 48.4 Rimoldi, Marco •T 76.4 Rincke, Philipp •T 69.9 Ringwald, Andreas GR 14.1 Rink, Thomas •T 74.2 Rinagel, Maximilian T 11.3, T 11.4, •T 11.5, T 11.6, T 97.7 Risse, Markus T 93.10, T 101.6 Ritter, Sebastian •T 59.2
Rievers, Benny GR 2.8 Rifai, Mariam . T 20.2, •T 20.5, T 43.1, T 48.4 Rimoldi, Marco •T 76.4 Rincke, Philipp •T 69.9 Ringwald, Andreas GR 14.1 Rink, Thomas •T 74.2 Rinnagel, Maximilian T 11.1, T 11.2, T 11.3, T 11.4, •T 11.5, T 11.6, T 97.7 Risse, Markus
Rievers, Benny GR 2.8 Rifai, Mariam . T 20.2, •T 20.5, T 43.1, T 48.4 Rimoldi, Marco . T 76.4 Rincke, Philipp . T 69.9 Ringwald, Andreas . GR 14.1 Rink, Thomas
Rievers, Benny GR 2.8 Rifai, Mariam . T 20.2, •T 20.5, T 43.1, T 48.4 Rimoldi, Marco •T 76.4 Rincke, Philipp •T 69.9 Ringwald, Andreas GR 14.1 Rink, Thomas •T 74.2 Rinagel, Maximilian T 11.3, T 11.4, •T 11.5, T 11.6, T 97.7 Risse, Markus Ritter, Martin AKPIK 1.9 Ritter, Sebastian •T 59.2 Rittsteiger, Rieka T 24.1, •T 18.4 RNO-G Collaboration-Kollaboration T 71.5, T 12, T 71.5
Rievers, Benny GR 2.8 Rifai, Mariam . T 20.2, •T 20.5, T 43.1, T 48.4 Rimoldi, Marco •T 76.4 Rincke, Philipp •T 69.9 Ringwald, Andreas GR 14.1 Rink, Thomas •T 74.2 Rinnagel, Maximilian T 11.1, T 11.2, T 11.3, T 11.4, •T 11.5, T 11.6, T 97.7 Risse, Markus
Rievers, Benny GR 2.8 Rifai, Mariam . T 20.2, •T 20.5, T 43.1, T 48.4 Rimoldi, Marco •T 76.4 Rincke, Philipp •T 69.9 Ringwald, Andreas GR 14.1 Rink, Thomas •T 74.2 Rinagel, Maximilian T 11.1, T 11.2, T 11.3, T 11.4, •T 11.5, T 11.6, T 97.7 Risse, Markus T 93.10, T 101.6 Ritter, Martin AKPIK 1.9 Ritter, Sebastian •T 59.2 Rittseiger, Rieka T 84.1, •T 84.4 RNO-G Collaboration-Kollaboration T 17.7, T 71.5, T 71.6 Robles Manzano, Maria de la Soledad
Rievers, Benny GR 2.8 Rifai, Mariam . T 20.2, •T 20.5, T 43.1, T 48.4 Rimoldi, Marco . T 76.4 Rincke, Philipp . T 69.9 Ringwald, Andreas . GR 14.1 Rink, Thomas . • T 74.2 Rinnagel, Maximilian . 11.1, T 11.2, T 11.3, T 11.4, •T 11.5, T 11.6, T 97.7 Risse, Markus
Rievers, Benny GR 2.8 Rifai, Mariam . T 20.2, •T 20.5, T 43.1, T 48.4 Rimoldi, Marco . T 76.4 Rincke, Philipp •T 69.9 Ringwald, Andreas . GR 14.1 Rink, Thomas . •T 74.2 Rinnagel, Maximilian . T 11.1, T 11.2, T 11.3, T 11.4, •T 11.5, T 11.6, T 97.7 Risse, Markus . T 93.10, T 101.6 Ritter, Sebastian . •T 59.2 Rittsteiger, Rieka . T 84.1, •T 84.4 RNO-G Collaboration-Kollaboration T 71.7, T 71.5, T 71.6 Robles Manzano, Maria de la Soledad •T 97.5 Robles Manzano, Marisol T 41.1,
Rievers, Benny GR 2.8 Rifai, Mariam . T 20.2, •T 20.5, T 43.1, T 48.4 Rimoldi, Marco •T 76.4 Rincke, Philipp •T 69.9 Ringwald, Andreas GR 14.1 Rink, Thomas •T 74.2 Rinagel, Maximilian T 11.1, T 11.2, T 11.3, T 11.4, •T 11.5, T 11.6, T 97.7 Risse, Markus T 93.10, T 101.6 Ritter, Martin KAPIK 1.9 Ritter, Sebastian •T 59.2 Rittseiger, Rieka T 84.1, •T 84.4 RNO-G Collaboration-Kollaboration T 71.7, T 71.5, T 71.6 Robles Manzano, Maria de la Soledad •T 97.5 Robles Manzano, Marisol T 41.1, 1
Rievers, Benny GR 2.8 Rifai, Mariam . T 20.2, •T 20.5, T 43.1, T 48.4 Rimoldi, Marco •T 76.4 Rincke, Philipp •T 69.9 Ringwald, Andreas GR 14.1 Rink, Thomas •T 74.2 Rinnagel, Maximilian T 11.1, T 11.2, T 11.3, T 11.4, •T 11.5, T 11.6, T 97.7 Risse, Markus Ritter, Martin AKPIK 1.9 Ritter, Sebastian •T 59.2 Rittsteiger, Rieka T 84.1, •T 84.4 RNO-G Collaboration-Kollaboration T 71.7, T 71.5, T 71.6 Robles Manzano, Maria de la Soledad •T 97.5 Robles Manzano, Marisol T 41.1, T 59.2 Robles Manzano, Marisol Packetuble Carcter DD 21.1 DD 21.2
Rievers, Benny GR 2.8 Rifai, Mariam . T 20.2, •T 20.5, T 43.1, T 48.4 Rimoldi, Marco •T 76.4 Rincke, Philipp •T 69.9 Ringwald, Andreas GR 14.1 Rink, Thomas •T 74.2 Rinagel, Maximilian T 11.1, T11.2, T 11.3, T 11.4, •T 11.5, T 11.6, T 97.7 Risse, Markus T 93.10, T 101.6 Ritter, Martin
Rievers, Benny GR 2.8 Rifai, Mariam . T 20.2, •T 20.5, T 43.1, T 48.4 Rimoldi, Marco . T 76.4 Rincke, Philipp . T 69.9 Ringwald, Andreas . GR 14.1 Rink, Thomas . • T 74.2 Rinnagel, Maximilian . 11.1, T 11.2, T 11.3, T 11.4, •T 11.5, T 11.6, T 97.7 Risse, Markus Ritter, Martin
Rievers, Benny GR 2.8 Rifai, Mariam . T 20.2, •T 20.5, T 43.1, T 48.4 Rimoldi, Marco . T 76.4 Rincke, Philipp •T 69.9 Ringwald, Andreas . GR 14.1 Rink, Thomas . •T 74.2 Rinnagel, Maximilian . •T 11.1, T 11.2, T 11.3, T 11.4, •T 11.5, T 11.6, T 97.7 Risse, Markus . T 93.10, T 101.6 Ritter, Martin
Rievers, Benny GR 2.8 Rifai, Mariam . T 20.2, •T 20.5, T 43.1, T 48.4 Rimoldi, Marco . T 76.4 Rincke, Philipp . T 69.9 Ringwald, Andreas . GR 14.1 Rink, Thomas . •T 74.2 Rinnagel, Maximilian . T 11.1, T 11.2, T 11.3, T 11.4, •T 11.5, T 11.6, T 97.7 Risse, Markus . T 93.10, T 101.6 Ritter, Martin
Rievers, Benny GR 2.8 Rifai, Mariam . T 20.2, •T 20.5, T 43.1, T 48.4 Rimoldi, Marco •T 76.4 Rincke, Philipp •T 69.9 Ringwald, Andreas GR 14.1 Rink, Thomas •T 74.2 Rinnagel, Maximilian T 11.1, T 11.2, T 11.3, T 11.4, •T 11.5, T 11.6, T 97.7 Risse, Markus
Rievers, Benny GR 2.8 Rifai, Mariam . T 20.2, •T 20.5, T 43.1, T 48.4 Rimoldi, Marco • T 76.4 Rincke, Philipp • T 69.9 Ringwald, Andreas GR 14.1 Rink, Thomas • T 74.2 Rinnagel, Maximilian T 11.1, T 11.2, T 11.3, T 11.4, • T 11.5, T 11.6, T 97.7 Risse, Markus T 93.10, T 101.6 Ritter, Martin AKPIK 1.9 Ritter, Sebastian • T 59.2 Robles Manzano, Marisol T 41.1, T 97.5 Robles Manzano, Marisol T 41.1, T 59.2 Rockstuhl, Carsten DD 31.1, DD 31.2, DD 36.2 Rode, Stefan F 87.7 Rodigues, Xavier EP 7.2, •EP 11.3, • T 83.3 • T 83.3
Rievers, Benny GR 2.8 Rifai, Mariam . T 20.2, •T 20.5, T 43.1, T 48.4 Rimoldi, Marco •T 76.4 Rincke, Philipp •T 69.9 Ringwald, Andreas GR 14.1 Rink, Thomas •T 74.2 Rinnagel, Maximilian .T 11.1, T 11.2, T 11.3, T 11.4, •T 11.5, T 11.6, T 97.7 Risse, Markus T 93.10, T 101.6 Ritter, Martin AKPIK 1.9 Ritter, Sebastian •T 59.2 Rittsteiger, Rieka T 84.1, •T 84.4 RNO-G Collaboration-Kollaboration T 71.7, T 71.5, T 71.6 Robles Manzano, Maria de la Soledad •T 97.5 Robles Manzano, Mariaol T 41.1, T 59.2 Rockstuhl, Carsten .DD 31.1, DD 31.2, DD 36.2 Rode, Stefan EP 7.2, •EP 11.3, Rodsiguez, Xavier EP 7.2, •EP 11.3, •T 83.3 Rodriguez, Aturo T 13 6 T 94 1
Rievers, Benny GR 2.8 Rifai, Mariam . T 20.2, •T 20.5, T 43.1, T 48.4 Rimoldi, Marco . T 76.4 Rincke, Philipp . T 69.9 Ringwald, Andreas . GR 14.1 Rink, Thomas . T 74.2 Rinnagel, Maximilian . T 11.1, T 11.2, T 11.3, T 11.4, •T 11.5, T 11.6, T 97.7 Risse, Markus . T 93.10, T 101.6 Ritter, Martin . AKPIK 1.9 Ritter, Sebastian . T 59.2 Robles Manzano, Marisol . T 41.1, T 59.2 Rockstuhl, Carsten . DD 31.1, DD 31.2, D 36.2 Rode, Stefan . T 87.7 Rodrigues, Xavier . EP 7.2, •EP 11.3, •T 83.3 Rodrigues, Xavier . EP 7.2, •EP 11.3, •T 20.4
Rievers, Benny GR 2.8 Rifai, Mariam . T 20.2, •T 20.5, T 43.1, T 48.4 Rimoldi, Marco . T 76.4 Rincke, Philipp . T 69.9 Ringwald, Andreas . GR 14.1 Rink, Thomas . • T 74.2 Rinnagel, Maximilian . T 11.1, T 11.2, T 11.3, T 11.4, •T 11.5, T 11.6, T 97.7 Risse, Markus
Rievers, Benny GR 2.8 Rifai, Mariam . T 20.2, •T 20.5, T 43.1, T 48.4 Rimoldi, Marco . T 76.4 Rincke, Philipp . T 69.9 Ringwald, Andreas . GR 14.1 Rink, Thomas . T 74.2 Rinnagel, Maximilian . T 11.1, T 11.2, T 11.3, T 11.4, •T 11.5, T 11.6, T 97.7 Risse, Markus . T 93.10, T 101.6 Ritter, Martin . AKPIK 1.9 Ritter, Sebastian . T 59.2 Robles Manzano, Maria de la Soledad . T 97.5 Robles Manzano, Marisol T 41.1, T 59.2 Rockstuhl, Carsten . DD 31.1, DD 31.2, DD 36.2 Rode, Stefan . T 87.7 Rodriguez, Aturo T 13.6, T 94.1 Rodriguez Gómez, Laura T 39.8, T 68.1, •T 97.1
Rievers, Benny GR 2.8 Rifai, Mariam . T 20.2, •T 20.5, T 43.1, T 48.4 Rimoldi, Marco . T 76.4 Rimcke, Philipp . T 69.9 Ringwald, Andreas . GR 14.1 Rink, Thomas . • T 74.2 Rinagel, Maximilian . T 11.1, T 11.2, T 11.3, T 11.4, •T 11.5, T 11.6, T 97.7 Risse, Markus . T 93.10, T 101.6 Ritter, Martin
Rievers, Benny GR 2.8 Rifai, Mariam . T 20.2, •T 20.5, T 43.1, T 48.4 Rimoldi, Marco •T 76.4 Rincke, Philipp •T 69.9 Ringwald, Andreas GR 14.1 Rink, Thomas •T 74.2 Rinnagel, Maximilian .T 11.1, T 11.2, T 11.3, T 11.4, •T 11.5, T 11.6, T 97.7 Risse, Markus
Rievers, Benny GR 2.8 Rifai, Mariam . T 20.2, •T 20.5, T 43.1, T 48.4 Rimoldi, Marco . T 76.4 Rincke, Philipp . T 69.9 Ringwald, Andreas . GR 14.1 Rink, Thomas . T 74.2 Rinnagel, Maximilian . T 11.1, T 11.2, T 11.3, T 11.4, •T 11.5, T 11.6, T 97.7 Risse, Markus . T 93.10, T 101.6 Ritter, Martin . AKPIK 1.9 Ritter, Sebastian . T 59.2 Robles Manzano, Maria de la Soledad . T 97.5 Robles Manzano, Marisol . T 41.1, T 59.2 Rockstuhl, Carsten . D 31.1, DD 31.2, D 36.2 Rode, Stefan . T 87.7 Rodriguez, Xavier . EP 7.2, •EP 11.3, • T 83.3 Rodriguez, Arturo T 13.6, T 94.1 Rodriguez, Gómez, Laura . T 39.8, T 68.1, •T 97.1 Roemer, Jonas . T 39.4 Roggel, Jenss . T 10.2, •T 37.7, T 88.1
Rievers, Benny GR 2.8 Rifai, Mariam . T 20.2, •T 20.5, T 43.1, T 48.4 Rimoldi, Marco . T 76.4 Rincke, Philipp . T 69.9 Ringwald, Andreas . GR 14.1 Rink, Thomas . T 74.2 Rinnagel, Maximilian . T 11.1, T 11.2, T 11.3, T 11.4, •T 11.5, T 11.6, T 97.7 Risse, Markus . T 93.10, T 101.6 Ritter, Martin . AKPIK 1.9 Rittsteiger, Rieka . T 84.1, •T 84.4 RNO-G Collaboration-Kollaboration T 71.7, T 71.5, T 71.6 Robles Manzano, Maria de la Soledad •T 97.5 Robles Manzano, Maria de la Soledad •T 97.5 Rockstuhl, Carsten . DD 31.1, DD 31.2, . DD 36.2 Rode, Stefan . T 87.7 Rodriguez, Xavier . EP 7.2, •EP 11.3, •T 83.3 Rodriguez, Arturo T 13.6, T 94.1 Rodriguez Gómez, Laura T 39.8, T 68.1, •T 97.1 Roemer, Jonas . T 10.2, •T 37.7, T 88.1 Rogel, Jens Rodriguez, Jens . T 10.2, •T 37.7, T 88.1
Rievers, Benny GR 2.8 Rifai, Mariam . T 20.2, •T 20.5, T 43.1, T 48.4 Rimoldi, Marco . T 76.4 Rincke, Philipp . T 69.9 Ringwald, Andreas . GR 14.1 Rink, Thomas . T 74.2 Rinnagel, Maximilian . T 11.1, T 11.2, T 11.3, T 11.4, •T 11.5, T 11.6, T 97.7 Risse, Markus . T 93.10, T 101.6 Ritter, Martin . AKPIK 1.9 Ritter, Sebastian . T 59.2 Robles Manzano, Maria de la Soledad . T 97.5 Robles Manzano, Marisol . T 41.1, T 59.2 Rockstuhl, Carsten . DD 31.1, DD 31.2, DD 36.2 . T 83.3 Rodriguez, Arturo . T 13.6, T 94.1 Rodriguez Gómez, Laura . T 39.4 R 39.4 Rodgel, Jens . T 10.2, •T 37.7, T 88.1 Roguer, Victor
Rievers, Benny GR 2.8 Rifai, Mariam . T 20.2, •T 20.5, T 43.1, T 48.4 Rimoldi, Marco . T 76.4 Rimcke, Philipp . T 69.9 Ringwald, Andreas . GR 14.1 Rink, Thomas
Rievers, Benny GR 2.8 Rifai, Mariam . T 20.2, •T 20.5, T 43.1, T 48.4 Rimoldi, Marco •T 76.4 Rincke, Philipp •T 69.9 Ringwald, Andreas GR 14.1 Rink, Thomas •T 74.2 Rinnagel, Maximilian T 11.1, T 11.2, T 11.3, T 11.4, •T 11.5, T 11.6, T 97.7 Risse, Markus
Rievers, Benny GR 2.8 Rifai, Mariam . T 20.2, •T 20.5, T 43.1, T 48.4 Rimoldi, Marco . T 76.4 Rimcke, Philipp . T 69.9 Ringwald, Andreas . GR 14.1 Rink, Thomas
Rievers, Benny GR 2.8 Rifai, Mariam . T 20.2, •T 20.5, T 43.1, T 48.4 Rimoldi, Marco . T 76.4 Rincke, Philipp . T 69.9 Ringwald, Andreas . GR 14.1 Rink, Thomas . T 74.2 Rinnagel, Maximilian . T 11.1, T 11.2, T 11.3, T 11.4, •T 11.5, T 11.6, T 97.7 Risse, Markus . T 93.10, T 101.6 Ritter, Martin . AKPIK 1.9 Ritter, Sebastian . •T 59.2 Rittsteiger, Rieka . T 84.1, •T 84.4 RNO-G Collaboration-Kollaboration T 71.7, T 71.5, T 71.6 Robles Manzano, Maria de la Soledad •T 97.5 Robles Manzano, Mariasol T 41.1, T 59.2 Rockstuhl, Carsten . DD 31.1, DD 31.2, D 36.2 Rode, Stefan . T 87.7 Rodriguez, Arturo . T 13.6, T 94.1 Rodriguez, Acturo . T 13.6, T 94.1 Rodriguez, Gómez, Laura . T 39.8, T 68.1, •T 97.1 Roemer, Jonas . T 39.4 Roggel, Jens . T 10.2, •T 37.7, T 88.1 Rodguez, Gómez, Laura . T 39.4 Roggel, Jens . T 10.2, •T 37.7, T 88.1
Rievers, Benny GR 2.8 Rifai, Mariam . T 20.2, •T 20.5, T 43.1, T 48.4 Rimoldi, Marco . T 76.4 Rincke, Philipp . T 69.9 Ringwald, Andreas . GR 14.1 Rink, Thomas
Rievers, Benny GR 2.8 Rifai, Mariam . T 20.2, •T 20.5, T 43.1, T 48.4 Rimoldi, Marco . T 76.4 Rincke, Philipp . T 69.9 Ringwald, Andreas . GR 14.1 Rink, Thomas . T 74.2 Rinnagel, Maximilian . T 11.1, T 11.2, T 11.3, T 11.4, •T 11.5, T 11.6, T 97.7 Risse, Markus . T 93.10, T 101.6 Ritter, Martin . AKPIK 1.9 Ritter, Sebastian . T 59.2 Rittsteiger, Rieka . T 84.1, •T 84.4 RNO-G Collaboration-Kollaboration T 71.7, T 71.5, T 71.6 Robles Manzano, Maria de la Soledad •T 97.5 Robles Manzano, Maria de la Soledad •T 97.5 Rockstuhl, Carsten . DD 31.1, DD 31.2, DD 36.2 Rode, Stefan . T 87.7 Rodriguez, Xavier . E P 7.2, •E P 11.3, •T 83.3 Rodriguez, Arturo . T 13.6, T 94.1 Roderiguez, Anturo . T 39.4 Roggel, Jens . T 10.2, •T 3.7, T 88.1 Roggle, Jens . T 10.2, •T 3.7, T 88.1 Rodriguez, Génez, Laura . T 39.4 Roggle, Jens . T 10.2, •T 3.7, T 88.1 Rogley, Ru
Rievers, Benny GR 2.8 Rifai, Mariam . T 20.2, •T 20.5, T 43.1, T 48.4 Rimoldi, Marco . T 76.4 Rincke, Philipp . T 69.9 Ringwald, Andreas . GR 14.1 Rink, Thomas . T 74.2 Rinnagel, Maximilian . T 11.1, T 11.2, T 11.3, T 11.4, •T 11.5, T 11.6, T 97.7 Risse, Markus . T 93.10, T 101.6 Ritter, Martin . AKPIK 1.9 Ritter, Sebastian . T 59.2 Robles Manzano, Maria de la Soledad . T 97.5 Robles Manzano, Marisol . T 41.1, T 59.2 Rockstuhl, Carsten . DD 31.1, DD 31.2, DD 36.2 Rode, Stefan . T 87.7 Rodriguez, Arturo
Rievers, Benny GR 2.8 Rifai, Mariam . T 20.2, •T 20.5, T 43.1, T 48.4 Rimoldi, Marco . T 76.4 Rincke, Philipp . T 69.9 Ringwald, Andreas . GR 14.1 Rink, Thomas
Rievers, Benny GR 2.8 Rifai, Mariam . T 20.2, •T 20.5, T 43.1, T 48.4 Rimoldi, Marco . T 76.4 Rincke, Philipp . T 69.9 Ringwald, Andreas . GR 14.1 Rink, Thomas . T 74.2 Rinnagel, Maximilian . T 11.1, T 11.2, T 11.3, T 11.4, •T 11.5, T 11.6, T 97.7 Risse, Markus . T 93.10, T 101.6 Ritter, Martin . AKPIK 1.9 Ritter, Sebastian . •T 59.2 Robles Manzano, Maria de la Soledad . T 97.5 Robles Manzano, Maria de la Soledad . T 97.5 Rockstuhl, Carsten . DD 31.1, DD 31.2, . DD 36.2 Rode, Stefan . T 87.7 Rodriguez, Aturo . T 13.6, T 94.1 Rodriguez, Acturo . T 13.6, T 94.1 Rodriguez, Acturo . T 39.4, Roggel, Jens . T 10.2, • T 37.7, T 88.1 Rodguez Gómez, Laura . T 39.4 Roggel, Jens . T 10.2, • T 37.7, T 88.1 Roggel, Jens . T 10.2, • T 37.7, T 88.1 Roggel, Jens . T 10.2, • T 37.7, T 88.3 Rodriguez Gómez, Laura T 39.4 Roggel, Jens <
Rievers, Benny GR 2.8 Rifai, Mariam . T 20.2, •T 20.5, T 43.1, T 48.4 Rimoldi, Marco . T 76.4 Rincke, Philipp . T 69.9 Ringwald, Andreas . GR 14.1 Rink, Thomas
Rievers, Benny GR 2.8 Rifai, Mariam . T 20.2, •T 20.5, T 43.1, T 48.4 Rimoldi, Marco . T 76.4 Rincke, Philipp . T 69.9 Ringwald, Andreas . GR 14.1 Rink, Thomas . T 74.2 Rinnagel, Maximilian . T 11.1, T 11.2, T 11.3, T 11.4, •T 11.5, T 11.6, T 97.7 Risse, Markus . T 93.10, T 101.6 Ritter, Martin . AKPIK 1.9 Ritter, Sebastian . T 59.2 Rittsteiger, Rieka . T 84.1, •T 84.4 RNO-G Collaboration-Kollaboration T 71.7, T 71.5, T 71.6 Robles Manzano, Maria de la Soledad •T 97.5 Robles Manzano, Maria de la Soledad •T 97.5 Rockstuhl, Carsten . DD 31.1, DD 31.2, DD 36.2 Rodriguez, Xavier . EP 7.2, •EP 11.3, •T 83.3 Rodriguez, Arturo . T 13.6, T 94.1 Rodriguez, Gómez, Laura T 39.4 Roggel, Jens . T 10.2, •T 3.7, T 88.1 Roggle, Jens . T 10.2, •T 3.7, T 88.1 Roggle, Jens . T 10.2, •T 3.7, T 88.1 Rodriguez, Arturo . MP 12.1, MP 12.2 Röhrig, Lars . T 34.7 Rodri
Rievers, Benny GR 2.8 Rifai, Mariam . T 20.2, •T 20.5, T 43.1, T 48.4 Rimoldi, Marco . T 76.4 Rincke, Philipp . T 69.9 Ringwald, Andreas . GR 14.1 Rink, Thomas . T 74.2 Rinnagel, Maximilian . T 11.1, T 11.2, T 11.3, T 11.4, •T 11.5, T 11.6, T 97.7 Risse, Markus . T 93.10, T 101.6 Ritter, Martin . AKPIK 1.9 Ritter, Sebastian . T 59.2 Robles Manzano, Maria de la Soledad . T 97.5 Robles Manzano, Marisol . T 41.1, T 59.2 Rockstuhl, Carsten . DD 31.1, DD 31.2, DD 36.2 Rode, Stefan . T 87.7 Rodriguez, Arturo
Rievers, Benny GR 2.8 Rifai, Mariam . T 20.2, •T 20.5, T 43.1, T 48.4 Rimoldi, Marco . T 76.4 Rincke, Philipp . T 69.9 Ringwald, Andreas . GR 14.1 Rink, Thomas
Rievers, Benny GR 2.8 Rifai, Mariam . T 20.2, •T 20.5, T 43.1, T 48.4 Rimoldi, Marco . T 76.4 Rincke, Philipp . T 69.9 Ringwald, Andreas . GR 14.1 Rink, Thomas . T 74.2 Rinnagel, Maximilian . T 11.1, T 11.2, T 11.3, T 11.4, •T 11.5, T 11.6, T 97.7 Risse, Markus . T 93.10, T 101.6 Ritter, Martin . AKPIK 1.9 Ritter, Sebastian . •T 59.2 Robles Manzano, Maria de la Soledad . T 97.5 Robles Manzano, Marisol T 41.1, T 59.2 Rockstuhl, Carsten . DD 31.1, DD 31.2, DD 36.2 Rode, Stefan . T 87.7 Rodriguez, Aturo T 13.6, T 94.1 Rodriguez, Acturo T 13.6, T 94.1 Rodriguez, Gómez, Laura T 39.4, Roggel, Jens T 10.2, •T 37.7, T 88.1 Rodguez, Gómez, Laura T 39.4 Roggel, Jens T 10.2, •T 37.7, T 88.1 Rodyal, Lars T 34.7 Rolad, Benoit T 89.6 Rodriguez, Arturo T 13.6, T 94.1 Rodgel, Jack .
Rievers, Benny GR 2.8 Rifai, Mariam T 20.2, •T 20.5, T 43.1, T 48.4 Rimoldi, Marco T 76.4 Rimcke, Philipp T 69.9 Ringwald, Andreas GR 14.1 Rink, Thomas •T 74.2 Rinnagel, Maximilian T 11.1, T 11.2, T 11.3, T 11.4, •T 11.5, T 11.6, T 97.7 Risse, Markus T 93.10, T 101.6 Ritter, Martin AKPIK 1.9 Ritter, Sebastian •T 59.2 Rittes, Banzano, Maria de la Soledad •T 97.5 Robles Manzano, Maria de la Soledad •T 97.5 Rockstuhl, Carsten DD 31.1, DD 31.2, DD 36.2 Rodriguez, Axivier .EP 7.2, •EP 11.3, •T 83.3 Rodriguez, Axivier .EP 7.2, •EP 11.3, •T 83.3 Rodriguez, Axivier .EP 7.2, •EP 11.3, •T 86.1, •T 97.1 Roemer, Jonas .T 39.4 Rogle, Jens .T 10.2, •T 37.7, T 88.1 Rogly, Rudolph .T 78.6 Rogner, Victor .MP 12.1, MP 12.2 Röhrig,
Rievers, Benny GR 2.8 Rifai, Mariam . T 20.2, •T 20.5, T 43.1, T 48.4 Rimoldi, Marco . T 76.4 Rincke, Philipp . T 69.9 Ringwald, Andreas . GR 14.1 Rink, Thomas . T 74.2 Rinnagel, Maximilian . T 11.1, T 11.2, T 11.3, T 11.4, •T 11.5, T 11.6, T 97.7 Risse, Markus . T 93.10, T 101.6 Ritter, Martin . AKPIK 1.9 Ritter, Sebastian . T 59.2 Robles Manzano, Maria de la Soledad . T 97.5 Robles Manzano, Maria de la Soledad . T 97.5 Robles Manzano, Maria de la Soledad . T 97.5 Rockstuhl, Carsten . DD 31.1, DD 31.2, . D 36.2 Rode, Stefan . T 87.7 Rodriguez, Xavier . EP 7.2, •EP 11.3, • T 83.3 Rodriguez, Arturo . T 13.6, T 94.1 Rodgel, Jens . T 10.2, •T 37.7, T 88.1 Rogle, Jens . T 10.2, •T 37.7, T 88.1 Rogle, Jens . T 10.2, •T 37.7, T 88.1 Rodriguez, Arturo . T 13.6, T 94.1 Rodriguez, Arturo . T 36.4 Roggel, Jens . T 10.2, •T 37.7, T 88.1 <
Rievers, Benny GR 2.8 Rifai, Mariam . T 20.2, •T 20.5, T 43.1, T 48.4 Rimoldi, Marco . T 76.4 Rincke, Philipp . T 69.9 Ringwald, Andreas . GR 14.1 Rink, Thomas . T 74.2 Rinnagel, Maximilian . T 11.1, T 11.2, T 11.3, T 11.4, •T 11.5, T 11.6, T 97.7 Risse, Markus . T 93.10, T 101.6 Ritter, Martin . AKPIK 1.9 Ritter, Sebastian . T 59.2 Robles Manzano, Maria de la Soledad . T 97.5 Robles Manzano, Marisol . T 41.1, T 59.2 Rockstuhl, Carsten . DD 31.1, DD 31.2, DD 36.2 Rode, Stefan . T 87.7 Rodriguez, Arturo
Rievers, Benny GR 2.8 Rifai, Mariam . T 20.2, •T 20.5, T 43.1, T 48.4 Rimoldi, Marco . T 76.4 Rincke, Philipp . T 69.9 Ringwald, Andreas . GR 14.1 Rink, Thomas
Rievers, Benny GR 2.8 Rifai, Mariam . T 20.2, •T 20.5, T 43.1, T 48.4 Rimoldi, Marco . T 76.4 Rincke, Philipp . T 69.9 Ringwald, Andreas . GR 14.1 Rink, Thomas . T 74.2 Rinnagel, Maximilian . T 11.1, T 11.2, T 11.3, T 11.4, •T 11.5, T 11.6, T 97.7 Risse, Markus . T 93.10, T 101.6 Ritter, Martin . AKPIK 1.9 Ritter, Sebastian . T 59.2 Robles Manzano, Maria de la Soledad . T 97.5 Robles Manzano, Marisol T 41.1, T 13.6, T 97.5 Robles Manzano, Marisol T 41.1, Rockstuhl, Carsten . DD 31.1, DD 31.2, . DD 36.2 Rode, Stefan T 13.6, T 94.1 Rodriguez, Arturo T 13.6, T 94.1 Rodriguez, Acturo T 13.6, T 94.1 Rodriguez, Gómez, Laura T 39.4 Roggel, Jens T 10.2, •T 37.7, T 88.1 Rodriguez, Gómez, Laura T 39.4 Roggel, Jens T 10.2, •T 37.7, T 88.1 Rodriguez, MP 12.2, MP 12.2 Röhrig, Lars T 34.7 Roland, Benoit T 89.6 Rogner, Victor MP 12.1, MP
Rievers, Benny GR 2.8 Rifai, Mariam T 20.2, •T 20.5, T 43.1, T 48.4 Rimoldi, Marco T 76.4 Rimcke, Philipp T 69.9 Ringwald, Andreas GR 14.1 Rink, Thomas •T 74.2 Rinnagel, Maximilian T 11.1, T 11.2, T 11.3, T 11.4, •T 11.5, T 11.6, T 97.7 Risse, Markus T 93.10, T 101.6 Ritter, Martin AKPIK 1.9 Ritter, Sebastian •T 59.2 Rittes, Banzano, Maria de la Soledad •T 97.5 Robles Manzano, Maria de la Soledad •T 97.5 Robles Manzano, Maria de la Soledad •T 97.5 Rockstuhl, Carsten DD 31.1, DD 31.2, DD 36.2 Rode, Stefan •T 87.7 Rodriguez, Xavier EP 7.2, •EP 11.3, •T 83.3 Rodriguez, Axier T 39.4 Rodgel, Jens T 10.2, •T 37.7, T 88.1 Rogly, Rudolph T 78.6 Rogner, Victor •M 12.1, MP 12.2 Röhrig, Lars •T 34.7 Rolad, Benoit T 89.6 Rodriguez, Anturo •T 18.3 Rodriguez, Gómez, Laura T 39.4 <td< td=""></td<>

	Roth, Markus EP 7.3, T 19.3, T 47.1,
	T 47.3, T 73.4, T 101.7, AKPIK 4.8,
	Roth. Stefan T 39.5. T 39.6. T 39.7.
	T 102.8
	Rothe, Johannes T 74.7, •T 74.8,
	T 104 7
	Rottler, Benjamin•T 89.6
	Rottoli, LucaT 2.4
	Rousselot, Armand
	Rövert Nicolas •T 65 1 T 65 2
	T 65.3, T 65.4, T 94.6
	Rübenach, Jonas •T 36.1, T 76.7
	Rubiera Gimeno, Jose Alejandro
	Rücker. Franziska
	Ruderer, Rainer MP 4.6
	Ruehl, Philip•T 19.1, T 93.10
	Ruhe. Tim EP 11.8. T 46.5. T 83.8
	Ruiz Daza, Sara•T 40.6
	Ruohan, Li
	Rupp, Natascha I /8.2 Puprecht Daniel T 53.8
	Rurikova, Zuzana
	Rusiecka, Katarzyna ST 1.5
	Rusov, Aleksey •T 56.2, T 56.4, T 85.2
	Rütter, Hannes GR 13.4
	Rymaszewski, PiotrT 40.9
	Rzhevskij, Vladimir
	Salari, MonammadSI 2.1 Saffer Julian
	Sahakyan, Narek EP 11.2, T 83.2
	Sahoo, Niladri T 4.4
	Saimpert, Matthias
	•DD 46.3
	Sakhare, Shreyans EP 4.3, EP 4.6
	Salewski, Janine
	Salmhofer, Manfred
	Salomon, Richard T 21.4, •T 21.5
	Salzburger, Andreas T 16.7, T 16.8
	Sam, Zeyd•GR 1.4 Sammel Dirk •T 15 2
	Sampaio-Kronister, Iva N DD 10.7
	Sampathkumar, Pranav T 47.8,
	Sánchez. Federico
	Sanchez Garcia, Edgar•T 74.1
	Sanchioni, Marco•AGPhil 3.1 Sandrock Alexander T 51.6 •T 101.8
	Sanduk, Mohammed •AGPhil 12.4
	Santiago, Jessica GR 2.3
	Santimaria, Marco 1 4.4 Santos Rui T 85.3 T 92.2
	Sara, El-Beit Shawish
	Saraswati, Ayu ST 3.4
	Sarbicki, Gniewomir MP 5.4 Sarbis Gediminas T 57.8
	Sauer, Christof
	Sauerburger, Frank
	Sauerland, Dennis
	Saur, Miroslav
	Savastano, Stefano
	Saveiberg, Joelle 1 100.1, •1 100.3 Savoiu, Daniel•T 32.3. T 32.7
	Savu, Vladimir
	SBT-Kollaboration T 42.1, T 42.2,
	Schaa, Elena
	Schaaf, Magnus C
	Schaefer, Ulrich
	Schael, StefanT 4.1.T 57 4 T 57 8
	Schäfer, Björn Malte
	Schäfer, Marlin Benedikt •GR 9.1
	Schafmeister Johanna T 104.3
	Schalk, Lennart
	Schall, Lars
	scham, moritzAKPIK 2.4 Scharf Christian T 68.2 T 04.2
	Schatz, Carina DD 46.1
	Schaufel, Merlin
1	Scheer Stina DD 13 2 DD 26 2
Ţ	
	Scheiger, Philipp DD 6.3, •DD 44.3
	Scheiger, Philipp DD 6.3, •DD 44.3 Scheilhaas, Timo . T 70.2, AKPIK 2.3,
	Scheiger, Philipp DD 6.3, •DD 24.3 Scheilhaas, Timo T 70.2, AKPIK 2.3, •AKPIK 2.7 Schenk Lena

Scherer, Klaus	
Schermer, Nicole•1 /4./, 1 /8.6, T 78.7, T 104.5, T 104.7	
Scheulen, Chris•T 14.6, •T 35.5	
Scheumann, Jan	
Schiffer, Tobias •T 11.8, T 14.2, T 38.5	
Schifferer, Florian	
Schilling, Isabelle ST 4.3, •ST 6.2	
Schimmasek Rudolf ST 6 4	
Schindler. Günther AKPIK 1.8	
Schindler, Jörg T 37.9	
Schindler, Sebastian	
Schinnerl, Luca•T 53.2	
Schirrmacher, Arne	
Schlag, Bastian	
T 101.2	
Schleicher, Katharina •T 8.8, T 8.9	
Schleichert, RalfGR 14.3	
T 91.6	
Schlichting, Sören	
Schlickeiser, Reinhard EP 12.2	
Schlingmann, Jacqueline •ST 1.2,	
ST 4.3	
Schlummer, Paul •DD 3.3, DD 17.1.	
DD 37.2	
Schlüß, Jennifer•ST 7.1	
Schlüter, Berit	
Schmallzl, Wolfgang . ST 9.5, •1 66.4,	
Schmeing, Jonas	
Schmeißer, Nick	
Schmeißer, Nick Jannis•T 93.7	
Schmeling, Sascha . DD 10.2, DD 11.7,	
DD 33.1	
Schmelling, Michael I 3.7, I 58.4 Schmidt Alexander T 26.1 T 26.3	
T 89 2 T 90 6 T 105 10	
Schmidt, David EP 7.3, T 19.3, T 47.3,	
T 51.4, AKPIK 4.8, AKPIK 4.9	
O I . II EI: T 40.0	
Schmidt, Ella 1 43.8	
Schmidt, Ella	
Schmidt, Ella	
Schmidt, Jakob • T 42.2, T 64.9, T 68.2 Schmidt, Kevin AKPIK 1.6, AKPIK 4.3, AKPIK 4.5	
Schmidt, Ella 1 43.8 Schmidt, Jakob 1 43.8 Schmidt, Jakob 1 43.8 T 68.2 Schmidt, Kevin AKPIK 4.5 Schmidt, Kustafa	
Schmidt, Ella 1 43.8 Schmidt, Jakob • T 42.2, T 64.9, T 68.2 Schmidt, Kevin Schmidt, Kevin AKPIK 1.6, AKPIK 4.3, AKPIK 4.5 Schmidt, Mustafa Schmidt, Sebastian • T 3.9	
Schmidt, Ella 1 43.8 Schmidt, Jakob • T 42.2, T 64.9, T 68.2 Schmidt, Kevin AKPIK 1.6, AKPIK 4.3, AKPIK 4.5 Schmidt, Mustafa • T 3.9 Schmidt, Sebastian • T 11.8, T 38.5 Schmidt, Ristof • T 10.7, • T 29.2 Schmidt ar Delf • T 30.4 T 26.5	
Schmidt, Ella 143.8 Schmidt, Jakob • T 42.2, T 64.9, T 68.2 Schmidt, Kevin Schmidt, Kevin AKPIK 1.6, AKPIK 4.3, AKPIK 4.5 Schmidt, Sebastian Schmidt, Sebastian • T 11.8, T 38.5 Schmidt, Kristof T 10.7, • T 29.2 Schmieder, Ralf • T 36.4, T 36.5	
Schmidt, Ella 143.8 Schmidt, Jakob • T 42.2, T 64.9, T 68.2 Schmidt, Kevin Schmidt, Kevin AKPIK 1.6, AKPIK 4.3, AKPIK 4.5 Schmidt, Sebastian Schmidt, Sebastian • T 13.9 Schmidt, Sebastian • T 10.7, • T 29.2 Schmieder, Kristof • T 36.4, T 36.5 Schmienann, Philipp DD 5.3 Schmitt, Caspar • T 30.1	
Schmidt, Ella 143.8 Schmidt, Jakob • T 42.2, T 64.9, T 68.2 Schmidt, Kevin Schmidt, Kevin AKPIK 1.6, AKPIK 4.3, AKPIK 4.5 Schmidt, Mustafa Schmidt, Sebastian • T 11.8, T 38.5 Schmidt, Sebastian • T 10.7, • T 29.2 Schmieden, Kristof • T 36.4, T 36.5 Schmieder, Ralf • T 36.4, T 36.5 Schmitt, Caspar • T 30.1 Schmitt, Christian • T 16.8, T 41.1	
Schmidt, Ella 143.8 Schmidt, Jakob • T 42.2, T 64.9, T 68.2 Schmidt, Kevin AKPIK 1.6, AKPIK 4.3, AKPIK 4.5 Schmidt, Mustafa • T 3.9 Schmidt, Sebastian • T 11.8, T 38.5 Schmidt, Sebastian • T 10.7, • T 29.2 Schmieden, Kristof • T 10.7, • T 29.2 Schmieder, Ralf • T 36.4, T 36.5 Schmitt, Caspar • T 30.1 Schmitt, Christian • T 16.8, T 41.1, T 59.2, T 89.5, T 90.7	
Schmidt, Ella 143.8 Schmidt, Jakob • T 42.2, T 64.9, T 68.2 Schmidt, Kevin AKPIK 1.6, AKPIK 4.3, AKPIK 4.5 Schmidt, Mustafa • T 3.9 Schmidt, Sebastian • T 11.8, T 38.5 Schmidt, Sebastian • T 10.7, • T 29.2 Schmieden, Kristof • T 10.7, • T 29.2 Schmieder, Ralf • T 36.4, T 36.5 Schmiet, Caspar • T 30.1 Schmitt, Caspar • T 30.1 Schmitt, Kevin • D 16.8, T 41.1, T 59.2, T 89.5, T 90.7 Schmitt, Kevin • D 5.1	
Schmidt, Ella 143.8 Schmidt, Jakob • T 42.2, T 64.9, T 68.2 Schmidt, Kevin AKPIK 1.6, AKPIK 4.3, AKPIK 4.5 Schmidt, Mustafa • T 3.9 Schmidt, Sebastian • T 11.8, T 38.5 Schmidt, Sebastian • T 10.7, • T 29.2 Schmieden, Kristof • T 10.7, • T 29.2 Schmieder, Ralf • T 36.4, T 36.5 Schmieder, Ralf • T 30.1 Schmitt, Caspar • T 30.1 Schmitt, Christian • T 16.8, T 41.1, T 59.2, T 89.5, T 90.7 Schmitt, Kevin • D 5.1 Schmitt, Niklas • T 89.5	
Schmidt, Ella 143.8 Schmidt, Jakob • T 42.2, T 64.9, T 68.2 Schmidt, Kevin AKPIK 1.6, AKPIK 4.3, AKPIK 4.5 Schmidt, Mustafa • T 3.9 Schmidt, Sebastian • T 11.8, T 38.5 Schmidt, Sebastian • T 10.7, • T 29.2 Schmieden, Kristof • T 10.7, • T 29.2 Schmieden, Kristof • T 36.4, T 36.5 Schmitt, Caspar • T 30.1 Schmitt, Christian • T 16.8, T 41.1, T 59.2, T 89.5, T 90.7 Schmitt, Kevin • D 5.1 Schmitt, Sebastian • T 49.5 Schmitt, Stafa • T 49.4	
Schmidt, Ella 143.8 Schmidt, Jakob • T 42.2, T 64.9, T 68.2 Schmidt, Kevin AKPIK 1.6, AKPIK 4.3, AKPIK 4.5 Schmidt, Mustafa • T 3.9 Schmidt, Sebastian • T 11.8, T 38.5 Schmidt, Sebastian • T 10.7, • T 29.2 Schmieden, Kristof • T 10.7, • T 29.2 Schmieden, Kristof • T 36.4, T 36.5 Schmitt, Caspar • T 30.1 Schmitt, Christian • T 16.8, T 41.1, T 59.2, T 89.5, T 90.7 Schmitt, Kevin • DD 5.1 Schmitt, Kevin • DD 5.1 Schmitt, Sebastian • T 4.1 Schmitt, Stefan • T 4.1 Schmitt, Stefan • T 94.3	
Schmidt, Ella 143.8 Schmidt, Jakob • T 42.2, T 64.9, T 68.2 Schmidt, Kevin Schmidt, Kevin AKPIK 1.6, AKPIK 4.3, AKPIK 4.5 Schmidt, Sebastian Schmidt, Sebastian • T 39.5 Schmidt, Sebastian • T 10.7, • T 29.2 Schmieden, Kristof • T 10.7, • T 29.2 Schmieder, Ralf • T 36.4, T 36.5 Schmitt, Caspar • T 30.1 Schmitt, Caspar • T 30.1 Schmitt, Caspar • T 16.8, T 41.1, T 59.2, T 89.5, T 90.7 Schmitt, Kevin Schmitt, Sebastian • T 49.5 Schmitt, Niklas • T 89.5 Schmitt, Stefan • T 4.1 Schmitz, Jannes • T 95.3, T 95.7 Schmitz, Sermaire, Felix • T 17.7	
Schmidt, Ella 143.8 Schmidt, Jakob • T 42.2, T 64.9, T 68.2 Schmidt, Kevin AKPIK 1.6, AKPIK 4.3, AKPIK 4.5 Schmidt, Sebastian • T 39.5 Schmidt, Sebastian • T 11.8, T 38.5 Schmidt, Sebastian • T 10.7, • T 29.2 Schmieden, Kristof • T 10.7, • T 29.2 Schmieder, Ralf • T 36.4, T 36.5 Schmiet, Caspar • T 30.1 Schmitt, Christian • T 16.8, T 41.1, T 59.2, T 89.5, T 90.7 Schmitt, Kevin • D 5.1 Schmitt, Niklas • T 89.5 Schmitt, Sebastian • T 4.1 Schmitt, Stefan • T 94.3 Schmitz, Jannes • T 95.3, T 95.7 Schmitz, Jannes • T 17.7 Schnabel, Jutta • T 25.6	
Schmidt, Ella 143.8 Schmidt, Jakob • T 42.2, T 64.9, T 68.2 Schmidt, Kevin Schmidt, Kevin AKPIK 1.6, AKPIK 4.3, AKPIK 4.5 Schmidt, Sebastian Schmidt, Sebastian • T 39.9 Schmidt, Sebastian • T 10.7, • T 29.2 Schmieden, Kristof • T 10.7, • T 29.2 Schmieder, Ralf • T 36.4, T 36.5 Schmitt, Caspar • T 30.1 Schmitt, Caspar • T 30.1 Schmitt, Christian • T 16.8, T 41.1, T 59.2, T 89.5, T 90.7 Schmitt, Sebastian Schmitt, Stefan • T 4.1 Schmitt, Stefan • T 4.1 Schmitt, Stefan • T 95.3, T 95.7 Schmitt, Stefan • T 17.7 Schmitt, Stefan • T 17.7 Schmitz, Jannes • T 25.6 Schmitz, Jannes • T 25.3 Schmitz, Stefan • T 17.7	
Schmidt, Ella 143.8 Schmidt, Jakob • T 42.2, T 64.9, T 68.2 Schmidt, Kevin AKPIK 1.6, AKPIK 4.3, AKPIK 4.5 Schmidt, Kevin AKPIK 1.6, AKPIK 4.3, AKPIK 4.5 Schmidt, Sebastian • T 39.5 Schmidt, Sebastian • T 10.7, • T 29.2 Schmieden, Kristof • T 10.7, • T 29.2 Schmieder, Ralf • T 36.4, T 36.5 Schmitt, Caspar • T 30.1 Schmitt, Christian • T 16.8, T 41.1, T 59.2, T 89.5, T 90.7 Schmitt, Kevin • DD 5.1 Schmitt, Kevin • DD 5.1 Schmitt, Sebastian • T 4.1 Schmitt, Stefan • T 41.3 Schmitt, Stefan • T 45.3, T 95.7 Schmitt, Stefan • T 47.3 Schmitt, Stefan • T 47.7 Schmitt, Stefan • T 47.7 Schmitt, Stefan • T 47.7 Schmitt, Stefan • T 45.3, T 95.7 Schmake, Simon • T 53.4 Schneeweiß, Robert D D 41.3 Schapeter Elevice • T 25.6	
Schmidt, Ella 143.8 Schmidt, Jakob • T 42.2, T 64.9, T 68.2 Schmidt, Kevin Schmidt, Kevin AKPIK 1.6, AKPIK 4.3, AKPIK 4.5 Schmidt, Sebastian Schmidt, Sebastian • T 11.8, T 38.5 Schmidt, Sebastian • T 10.7, • T 29.2 Schmieden, Kristof • T 10.7, • T 29.2 Schmidt, Sebastian • T 36.4, T 36.5 Schmitt, Raspar • T 30.1 Schmitt, Christian • T 16.8, T 41.1, T 59.2, T 89.5, T 90.7 Schmitt, Kevin Schmitt, Sebastian • T 4.1 Schmitt, Sebastian • T 41.3 Schmitt, Stefan • T 95.3, T 95.7 Schmitt, Stefan • T 47.3 Schmitt, Stefan • T 17.7 Schnabel, Jutta • T 25.6 Schnabel, Jutta • T 25.4 Schneeweiß, Robert DD 41.3 Schneider, Florian • D 24.8	
Schmidt, Ella 143.8 Schmidt, Jakob • T 42.2, T 64.9, T 68.2 Schmidt, Kevin Schmidt, Kevin AKPIK 1.6, AKPIK 4.3, AKPIK 4.5 Schmidt, Sebastian Schmidt, Sebastian • T 13.9 Schmidt, Sebastian • T 11.8, T 38.5 Schmidt, Sebastian • T 10.7, • T 29.2 Schmieden, Kristof • T 10.7, • T 29.2 Schmidt, Caspar • T 36.4, T 36.5 Schmitt, Caspar • T 30.1 Schmitt, Christian • T 16.8, T 41.1, T 59.2, T 89.5, T 90.7 Schmitt, Kevin Schmitt, Sebastian • T 4.1 Schmitt, Sebastian • T 4.1 Schmitt, Stefan • T 95.3, T 95.7 Schmitt, Stefan • T 45.4 Schmitz, Jannes • T 95.3, T 95.7 Schmitz, Jannes • T 95.3, T 95.7 Schmitz, Jannes • T 53.4 Schneeweiß, Robert DD 41.3 Schneider, Irina • DD 34.8 Schneider, Jörg • D 41.1	
Schmidt, Ella 143.8 Schmidt, Jakob • T 42.2, T 64.9, T 68.2 Schmidt, Kevin Schmidt, Kevin AKPIK 1.6, AKPIK 4.3, AKPIK 4.5 Schmidt, Sebastian Schmidt, Sebastian • T 18, T 38.5 Schmidt, Sebastian • T 10.7, • T 29.2 Schmieden, Kristof • T 10.7, • T 29.2 Schmieden, Kristof • T 10.7, • T 29.2 Schmieden, Kristof • T 10.7, • T 29.2 Schmidt, Gaspar • T 36.4, T 36.5 Schmitt, Caspar • T 30.1 Schmitt, Christian • T 16.8, T 41.1, T 59.2, T 89.5, T 90.7 Schmitt, Kevin Schmitt, Sebastian • T 4.1 Schmitt, Stefan • T 4.1 Schmitt, Stefan • T 95.3, T 95.7 Schmuckermaier, Felix • T 17.7 Schnabel, Jutta • T 25.6 Schneider, Florian ST 2.1 Schneider, Florian ST 2.1 Schneider, Florian • DD 41.1 Schneider, Jörg • DD 41.1	
Schmidt, Ella 143.8 Schmidt, Jakob • T 42.2, T 64.9, T 68.2 Schmidt, Kevin Schmidt, Kevin AKPIK 1.6, AKPIK 4.3, AKPIK 4.5 Schmidt, Mustafa Schmidt, Sebastian • T 18, T 38.5 Schmidt, Sebastian 11.8, T 38.5 Schmidt, Sebastian • T 10.7, • T 29.2 Schmieden, Kristof • T 10.7, • T 29.2 Schmieder, Ralf • T 36.4, T 36.5 Schmit, Caspar • T 30.1 Schmitt, Christian • T 16.8, T 41.1, T 59.2, T 89.5, T 90.7 Schmitt, Kevin Schmitt, Sebastian • T 4.1 Schmitt, Stefan • T 94.3 Schmitt, Stefan • T 95.3, T 95.7 Schmuckermaier, Felix • T 17.7 Schnabel, Jutta • T 25.4 Schneider, Florian • D 41.3 Schneider, Florian • D 34.8 Schneider, Jörg • D 41.1 Schneider, Jörg • D 41.1 Schneider, Martin • T 99.2 Schneider, Mike D. • AGPhil 5.3	
Schmidt, Ella 143.8 Schmidt, Jakob • T 42.2, T 64.9, T 68.2 Schmidt, Kevin Schmidt, Kevin AKPIK 1.6, AKPIK 4.3, AKPIK 4.5 Schmidt, Sebastian Schmidt, Sebastian • T 18, T 38.5 Schmidt, Sebastian • T 10.7, • T 29.2 Schmieden, Kristof • T 10.7, • T 29.2 Schmitt, Sebastian • T 30.1 Schmitt, Caspar • T 30.1 Schmitt, Christian • T 16.8, T 41.1, T 59.2, T 89.5, T 90.7 Schmitt, Kevin Schmitt, Kevin • DD 5.1 Schmitt, Sebastian • T 4.1 Schmitt, Stefan • T 4.1 Schmitt, Jannes • T 95.3, T 95.7 Schmuckermaier, Felix • T 17.7 Schnabel, Jutta • T 25.6 Schneider, Florian ST 2.1 Schneider, Florian • D 34.8 Schneider, Jörg • D 41.1 Schneider, Marti	
Schmidt, Ella 143.8 Schmidt, Jakob • T 42.2, T 64.9, T 68.2 Schmidt, Kevin Schmidt, Kevin AKPIK 1.6, AKPIK 4.3, AKPIK 4.5 Schmidt, Mustafa Schmidt, Sebastian . T 13.9 Schmidt, Sebastian . T 11.8, T 38.5 Schmidt, Sebastian . T 10.7, •T 29.2 Schmieden, Kristof . T 10.7, •T 29.2 Schmieder, Ralf . •T 30.1 Schmitt, Caspar . •T 30.1 Schmitt, Caspar . •T 30.1 Schmitt, Kevin . •D 5.1 Schmitt, Kevin . •D 5.1 Schmitt, Kevin . •D 5.1 Schmitt, Stefan . T 4.1 Schmitt, Stefan . T 94.3 Schmitz, Jannes . •T 95.3, T 95.7 Schmitz, Jannes . •T 95.3, 4 Schneeker, Simon . T 53.4 Schneeweiß, Robert . DD 41.3 Schneider, Florian . S 2.1 Schneider, Jörg . DD 41.1 Schneider, Martin . •T 99.2 Schneider, Martin . •T 99.2 Schneider, Susanne AGPhil 5.3 Sc	
Schmidt, Ella 143.8 Schmidt, Jakob • T 42.2, T 64.9, T 68.2 Schmidt, Kevin Schmidt, Kevin AKPIK 1.6, AKPIK 4.3, AKPIK 4.5 Schmidt, Sebastian Schmidt, Sebastian • T 39.9 Schmidt, Sebastian • T 30.1 Schmitt, Caspar • T 30.1 Schmitt, Kevin • DD 5.1 Schmitt, Kevin • DD 5.1 Schmitt, Sebastian • T 4.1 Schmitt, Stefan • T 94.3 Schmitt, Stefan • T 95.3, T 95.7 Schmitz, Jannes • T 95.3, T 95.7 Schmitz, Sobert DD 41.3 Schneider, Florian S 2.1 Schneider, Florian S 2.1 Schneider, Florian S 2.1 Schneider, Martin • D 9.4.8 Schneider, Martin • D 9.2.2 DD 32.2 Schneider, Susanne DD 10.4, DD 25.2,	
Schmidt, Ella 143.8 Schmidt, Jakob • T 42.2, T 64.9, T 68.2 Schmidt, Kevin AKPIK 1.6, AKPIK 4.3, AKPIK 4.5 Schmidt, Kevin AKPIK 1.6, AKPIK 4.3, AKPIK 4.5 Schmidt, Sebastian • T 39.5 Schmidt, Sebastian • T 10.7, • T 29.2 Schmieden, Kristof • T 10.7, • T 29.2 Schmieder, Ralf • T 36.4, T 36.5 Schmieder, Kristof • T 10.7, • T 29.2 Schmitt, Sebastian • T 30.1 Schmitt, Caspar • T 30.1 Schmitt, Caspar • T 30.1 Schmitt, Kevin • DD 5.1 Schmitt, Kevin • DD 5.1 Schmitt, Sebastian • T 4.1 Schmitt, Stefan • T 94.3 Schmitt, Stefan • T 47.7 Schnake, Simon • T 53.4 Schneeweiß, Robert DD 41.3 Schneider, Florian • T 21.4 Schneider, Martin • T 99.2 Schneider, Martin • D 34.8 Schneider, Susanne DD 10.4, DD 25.2, DD 32.2 DD 32.2 Schneider, Susanne D 10.4, DD 25.2, DD 32.2	
Schmidt, Ella 143.8 Schmidt, Jakob • T 42.2, T 64.9, T 68.2 Schmidt, Kevin Schmidt, Kevin AKPIK 1.6, AKPIK 4.3, AKPIK 4.5 Schmidt, Mustafa Schmidt, Sebastian T 13.9 Schmidt, Sebastian T 10.7, •T 29.2 Schmieden, Kristof T 10.7, •T 29.2 Schmieder, Ralf •T 36.4, T 36.5 Schmiemann, Philipp DD 5.3 Schmitt, Caspar •T 30.1 Schmitt, Caspar •T 30.1 Schmitt, Kevin DD 5.1 Schmitt, Kevin DD 5.1 Schmitt, Sebastian T 4.1 Schmitt, Stefan T 94.3 Schmitt, Stefan T 17.7 Schnabel, Jutta T 25.6 Schneeweiß, Robert DD 41.3 Schneider, Florian T 21.4 Schneider, Jörg D0 41.1 Schneider, Jörg DD 41.1 Schneider, Jörg DD 41.1 Schneider, Susanne DD 10.4, DD 25.2, DD 32.2 Schneider, Susanne DD 10.4, DD 25.2, DD 32.2	
Schmidt, Ella 143.8 Schmidt, Jakob • T 42.2, T 64.9, T 68.2 Schmidt, Kevin Schmidt, Kevin AKPIK 1.6, AKPIK 4.3, AKPIK 4.5 Schmidt, Mustafa Schmidt, Sebastian . T 13.9 Schmidt, Sebastian . T 10.7, •T 29.2 Schmieden, Kristof . T 10.7, •T 29.2 Schmieder, Ralf •T 36.4, T 36.5 Schmieder, Ralf . •T 30.1 Schmitt, Caspar •T 30.1 Schmitt, Caspar . •T 30.1 Schmitt, Kevin . •D 5.1 Schmitt, Kevin . •D 5.1 Schmitt, Sebastian . T 4.1 Schmitt, Sebastian . T 4.1 Schmitt, Stefan . T 94.3 Schmitt, Stefan . T 17.7 Schnake, Simon . T 53.4 Schneweiß, Robert . D 41.3 Schneider, Florian	
Schmidt, Ella 143.8 Schmidt, Jakob • T 42.2, T 64.9, T 68.2 Schmidt, Kevin Schmidt, Kevin AKPIK 1.6, AKPIK 4.3, AKPIK 4.5 Schmidt, Mustafa Schmidt, Sebastian . T 11.8, T 38.5 Schmidt, Sebastian . T 10.7, •T 29.2 Schmieden, Kristof . T 10.7, •T 29.2 Schmieder, Ralf •T 36.4, T 36.5 Schmiett, Caspar •T 30.1 Schmitt, Caspar •T 30.1 Schmitt, Caspar •T 30.1 Schmitt, Caspar •T 16.8, T 41.1, T 59.2, T 89.5, T 90.7 Schmitt, Kevin Schmitt, Sebastian . T 14.1 Schmitt, Sebastian . T 4.1 Schmitt, Stefan . T 94.3 Schmitt, Stefan . T 17.7 Schnake, Simon . T 53.4 Schneweiß, Robert . D 41.3 Schneweiß, Robert . D 41.3 Schneider, Irina . D 24.8 Schneider, Jörg . D 41.1 Schneider, Mathin . T 99.2 Schneider, Susanne DD 10.4, DD 25.2, DD 32.2 Schneider, Susanne DD 10.4, D	
Schmidt, Ella 143.8 Schmidt, Jakob • T 42.2, T 64.9, T 68.2 Schmidt, Kevin AKPIK 1.6, AKPIK 4.3, AKPIK 4.5 Schmidt, Mustafa • T 3.9 Schmidt, Sebastian . T 11.8, T 38.5 Schmidt, Sebastian . T 10.7, • T 29.2 Schmidt, Sebastian . T 10.7, • T 29.2 Schmidt, Caspar • T 36.4, T 36.5 Schmitt, Caspar • T 30.1 Schmitt, Caspar • T 30.1 Schmitt, Caspar • D 5.1 Schmitt, Kevin • DD 5.1 Schmitt, Kevin • DD 5.1 Schmitt, Sebastian • T 4.1 Schmitt, Sebastian • T 4.1 Schmitt, Stefan • T 95.3, T 95.7 Schmitt, Stefan • T 4.3 Schmitt, Stefan • T 25.6 Schnake, Simon • T 53.4 Schneeweiß, Robert DD 41.3 Schneider, Florian • DD 34.8 Schneider, Martin • T 99.2 Schneider, Martin • T 99.2 Schneider, Susanne DD 10.4, DD 25.2, DD 32.2 <td></td>	
Schmidt, Ella 143.8 Schmidt, Jakob • T 42.2, T 64.9, T 68.2 Schmidt, Kevin Schmidt, Kevin AKPIK 1.6, AKPIK 4.3, AKPIK 4.5 Schmidt, Mustafa Schmidt, Sebastian • T 13.9 Schmidt, Sebastian • T 10.7, • T 29.2 Schmieden, Kristof • T 10.7, • T 29.2 Schmidt, Sebastian • T 36.4, T 36.5 Schmieden, Kristof • T 10.7, • T 29.2 Schmidt, Caspar • T 30.1 Schmitt, Caspar • T 30.1 Schmitt, Christian • T 16.8, T 41.1, T 59.2, T 89.5, T 90.7 Schmitt, Niklas Schmitt, Kevin • DD 5.1 Schmitt, Stefan • T 4.1 Schmitt, Stefan • T 4.1 Schmitt, Stefan • T 4.3 Schmitt, Stefan • T 45.3 Schmitt, Stefan • T 45.3 Schmitt, Stefan • T 25.6 Schnake, Simon • T 53.4 Schneider, Irina • DD 41.3 Schneider, Irina • DD 34.8 Schneider, Martin • T 99.2	
Schmidt, Ella 143.8 Schmidt, Jakob • T 42.2, T 64.9, T 68.2 Schmidt, Kevin Schmidt, Kevin AKPIK 1.6, AKPIK 4.3, AKPIK 4.5 Schmidt, Mustafa Schmidt, Sebastian • T 13.9 Schmidt, Sebastian • T 10.7, • T 29.2 Schmidt, Sebastian • T 36.4, T 36.5 Schmieden, Kristof • T 10.7, • T 29.2 Schmidt, Caspar • T 30.1 Schmitt, Caspar • T 30.1 Schmitt, Christian • T 16.8, T 41.1, T 59.2, T 89.5, T 90.7 Schmitt, Kevin Schmitt, Kevin • DD 5.1 Schmitt, Sebastian • T 4.1 Schmitt, Stefan • T 4.1 Schmitt, Stefan • T 45.3 Schmitt, Stefan • T 95.3, T 95.7 Schmitt, Stefan • T 45.3 Schmitt, Stefan • T 4.1 Schmitt, Stefan • T 4.1 Schmitt, Stefan • T 45.3 Schmitt, Stefan • T 25.6 Schmitt, Stefan • T 25.6 Schneider, Jürg • DD 41.3 Schneider, Kastin • T 29.2 Sch	
Schmidt, Ella 143.8 Schmidt, Jakob • T 42.2, T 64.9, T 68.2 Schmidt, Kevin Schmidt, Kevin AKPIK 1.6, AKPIK 4.3, AKPIK 4.5 Schmidt, Mustafa Schmidt, Sebastian . T 13.9 Schmidt, Sebastian . T 10.7, •T 29.2 Schmidt, Sebastian . T 30.1 Schmitt, Caspar •T 30.1 Schmitt, Caspar •T 30.1 Schmitt, Caspar •T 30.1 Schmitt, Kevin •D 5.1 Schmitt, Kevin •D 5.1 Schmitt, Kevin •D 5.1 Schmitt, Sebastian •T 4.1 Schmitt, Stefan •T 94.3 Schmitt, Stefan •T 17.7 Schnake, Simon •T 53.4 Schneider, Florian ST 2.1 Schneider, Florian ST 2.1 Schneider, Florian ST 2.1 Schneider, Martin •T 99.2 Schneider, Martin •T 99.2 Schneider, Susanne DD 10.4, DD 25.2, DD 32.2 Schneider, Susanne DD 10.4, DD 25.2, DD 32.2 Schneider, Steffen T 97.5 <t< td=""><td></td></t<>	
Schmidt, Ella 143.8 Schmidt, Jakob • T 42.2, T 64.9, T 68.2 Schmidt, Kevin AKPIK 1.6, AKPIK 4.3, AKPIK 4.5 Schmidt, Kevin AKPIK 1.6, AKPIK 4.3, AKPIK 4.5 Schmidt, Sebastian • T 39.5 Schmidt, Sebastian • T 10.7, • T 29.2 Schmidt, Sebastian • T 30.1 Schmitt, Caspar • T 30.1 Schmitt, Caspar • T 30.1 Schmitt, Caspar • T 30.1 Schmitt, Kevin • DD 5.1 Schmitt, Kevin • DD 5.1 Schmitt, Kevin • DD 5.1 Schmitt, Sebastian • T 41.1 Schmitt, Stefan • T 94.3 Schmitt, Jita • T 25.6 Schmitz, Jannes • T 95.3, T 95.7 Schmitz, Jannes • T 95.3, 45.4 Schneider, Florian • T 17.7 Schnake, Simon • T 53.4 Schneider, Florian • T 2.1 Schneider, Jörg • DD 41.3 Schneider, Martin • T 99.2 Schneider, Susanne DD 10.4, DD 25.2, DD 32.2 Schneider, Susanne DD 10.4, DD 25.2, DD 32.2 Schneider, Steffen <td< td=""><td></td></td<>	
Schmidt, Ella 143.8 Schmidt, Jakob • T 42.2, T 64.9, T 68.2 Schmidt, Kevin AKPIK 1.6, AKPIK 4.3, AKPIK 4.5 Schmidt, Kevin AKPIK 1.6, AKPIK 4.3, AKPIK 4.5 Schmidt, Sebastian T 13.9 Schmidt, Sebastian T 10.7, •T 29.2 Schmieden, Kristof T 10.7, •T 29.2 Schmieder, Ralf •T 36.4, T 36.5 Schmietz, Sebastian T 16.8, T 41.1, T 59.2, T 89.5, T 90.7 Schmitt, Caspar •T 30.1 Schmitt, Kevin •D 5.1 Schmitt, Kevin •D 5.1 Schmitt, Sebastian T 4.1 Schmitt, Stefan T 94.3 Schmitt, Juanes T 95.3, T 95.7 Schmitz, Jannes T 25.6 Schnake, Simon T 53.4 Schneider, Florian T 21.7 Schneider, Florian T 21.1 Schneider, Martin •D 94.1.1 Schneider, Martin •D 41.3 Schneider, Susanne D 10.4, DD 25.2, DD 32.2 DD 32.2 Schneider, Steffen 197.5 Scholer, Patrick T 39.1, T 39.2, T 39.3 Scholer, Patric	
Schmidt, Ella 143.8 Schmidt, Jakob • T 42.2, T 64.9, T 68.2 Schmidt, Kevin Schmidt, Kevin AKPIK 1.6, AKPIK 4.3, AKPIK 4.5 Schmidt, Mustafa Schmidt, Sebastian . T 13.9 Schmidt, Sebastian . T 10.7, •T 29.2 Schmieden, Kristof . T 10.7, •T 29.2 Schmieder, Ralf . •T 36.4, T 36.5 Schmitt, Caspar . •T 30.1 Schmitt, Caspar . •T 30.1 Schmitt, Caspar . •T 30.1 Schmitt, Kevin . •D 5.1 Schmitt, Kevin . •D 5.1 Schmitt, Sebastian . •T 4.1 Schmitt, Stefan . T 47.3 Schmitt, Stefan . •T 47.7 Schnecker, Simon . •T 53.4 Schneeweiß, Robert . D 41.3 Schneider, Florian	
Schmidt, Ella 143.8 Schmidt, Jakob • T 42.2, T 64.9, T 68.2 Schmidt, Kevin AKPIK 1.6, AKPIK 4.3, AKPIK 4.5 Schmidt, Kevin AKPIK 1.6, AKPIK 4.3, AKPIK 4.5 Schmidt, Sebastian • T 13.9 Schmidt, Sebastian • T 10.7, • T 29.2 Schmieden, Kristof • T 10.7, • T 29.2 Schmieder, Ralf • T 36.4, T 36.5 Schmitt, Caspar • T 30.1 Schmitt, Caspar • T 30.1 Schmitt, Caspar • T 10.7, • T 29.2 Schmitt, Caspar • T 30.1 Schmitt, Caspar • T 10.7, • T 29.2 Schmitt, Caspar • T 30.1 Schmitt, Caspar • T 10.7, • T 29.2 Schmitt, Kevin • D D 5.1 Schmitt, Kevin • D D 5.1 Schmitt, Kevin • D 5.3 Schmitt, Sebastian • T 4.1 Schmitt, Stefan • T 4.1 Schmitt, Stefan • T 4.1 Schmitt, Sebastian • T 17.7 Schneider, Jörg • D 41.3 Schneider, Irina • D 25.3 Schneider, Jörg • D 41.1 Schneider, Mathin • T 99.2<	
Schmidt, Ella 143.8 Schmidt, Jakob • T 42.2, T 64.9, T 68.2 Schmidt, Kevin AKPIK 1.6, AKPIK 4.3, AKPIK 4.5 Schmidt, Kevin AKPIK 1.6, AKPIK 4.3, AKPIK 4.5 Schmidt, Sebastian • T 13.9 Schmidt, Sebastian • T 10.7, • T 29.2 Schmieden, Kristof • T 10.7, • T 29.2 Schmieden, Kristof • T 10.7, • T 29.2 Schmiett, Caspar • T 30.1 Schmitt, Caspar • T 30.1 Schmitt, Caspar • T 10.7, • T 29.2 Schmitt, Caspar • T 30.1 Schmitt, Caspar • T 10.7, • T 29.2 Schmitt, Caspar • T 30.1 Schmitt, Caspar • T 10.7 Schmitt, Kevin • D D 5.1 Schmitt, Kevin • D D 5.1 Schmitt, Stefan • T 4.1 Schmitt, Stefan • T 4.1 Schmitt, Stefan • T 95.3, T 95.7 Schmitt, Stefan • T 17.7 Schneake, Simon • T 53.4 Schneeweiß, Robert D 41.3 Schneider, Horian • D 94.13 Schneider, Mike D • AGPhil 5.3 Schneider, Susanne <t< td=""><td></td></t<>	
Schmidt, Ella 143.8 Schmidt, Jakob • T 42.2, T 64.9, T 68.2 Schmidt, Kevin AKPIK 1.6, AKPIK 4.3, AKPIK 4.5 Schmidt, Kevin AKPIK 1.6, AKPIK 4.3, AKPIK 4.5 Schmidt, Sebastian • T 13.9 Schmidt, Sebastian • T 10.7, • T 29.2 Schmieden, Kristof • T 10.7, • T 29.2 Schmidt, Caspar • T 36.4, T 36.5 Schmitt, Caspar • T 30.1 Schmitt, Caspar • T 30.1 Schmitt, Caspar • T 0.7, • T 29.2 Schmitt, Caspar • T 30.1 Schmitt, Caspar • D 5.1 Schmitt, Kevin • DD 5.1 Schmitt, Kevin • DD 5.1 Schmitt, Stefan • T 41.1 Schmitt, Stefan • T 41.3 Schmitt, Stefan • T 41.3 Schmitt, Stefan • T 47.7 Schmeckermaier, Felix • T 17.7 Schmeckermaier, Felix • T 17.7 Schneke, Simon • T 53.4 Schneeweiß, Robert DD 41.3 Schneider, Irina • DD 34.8 Schneider, Martin • T 99.2 Schneider, Susanne DD 10.4, DD 25.2,	
Schmidt, Ella 143.8 Schmidt, Jakob • T 42.2, T 64.9, T 68.2 Schmidt, Kevin AKPIK 1.6, AKPIK 4.3, AKPIK 4.5 Schmidt, Kevin AKPIK 1.6, AKPIK 4.3, AKPIK 4.5 Schmidt, Sebastian • T 13.9 Schmidt, Sebastian • T 10.7, • T 29.2 Schmieden, Kristof • T 10.7, • T 29.2 Schmieden, Kristof • T 10.7, • T 29.2 Schmidt, Caspar • T 30.1 Schmitt, Caspar • T 30.1 Schmitt, Caspar • T 30.1 Schmitt, Caspar • D 5.1 Schmitt, Kevin • DD 5.1 Schmitt, Kevin • DD 5.1 Schmitt, Kevin • DD 5.1 Schmitt, Sebastian • T 4.1 Schmitt, Stefan • T 4.3 Schneider, Joing • D 41.3 Schneider, Kobert D 41.3 Schneider, Kistof • D 92.2 Schneider, Martin • T 99.2 <t< td=""><td></td></t<>	
Schmidt, Ella 143.8 Schmidt, Jakob • T 42.2, T 64.9, T 68.2 Schmidt, Kevin AKPIK 1.6, AKPIK 4.3, AKPIK 4.5 Schmidt, Kustafa • T 39.9 Schmidt, Sebastian . T 11.8, T 38.5 Schmidt, Sebastian . T 10.7, • T 29.2 Schmieden, Kristof . T 10.7, • T 29.2 Schmidt, Sebastian . T 16.4, T 36.5 Schmidt, Caspar • T 30.1 Schmitt, Caspar • T 30.1 Schmitt, Caspar • D 5.3 Schmitt, Kevin • DD 5.1 Schmitt, Stefan • T 4.1 Schmitt, Sebastian • T 4.1 Schmitt, Stefan • T 4.3 Schmitt, Sebastian • T 4.1 Schmitt, Stefan • T 25.3 Schmitt, Stefan • T 25.3 Schneider, Joörg • DD 41.3 Schneider, Irin	

T 10 5 T 10 8 T 61 8 T 62 6	5 T 62 7
T 69 5	, i 02.7,
Schröder. René	DD 2.2
Schröder, Sarah	SYMD 1.4
Schroeder, Frank	T 101.7
Schroeder, Sascha	DD 10.5
Schroller, Marcel •T 9	9.6, T 99.7
Schrön, Martin	EP 5.3
Schubatzky, Thomas . DD 4	.3, DD 8.1,
DD 11.4, DD 11.5, DD 19.2, D	D 35.3,
•DD 43.3, DD 49.1, •DD 49.2	2
Schubert, Jan Lukas	•T 17.6
Schuck, Carsten DD 3.	3, DD 17.1,
DD 37.2	
Schug, David	ST 1.5
Schuler, Christoph	•1 64.3
Schuller, Frederic EP 11.7	', •EP 12.4,
I 83./	
	UU 40.1
Schulte, Denny	•I ZZ.0
Schulte Nicolo	T 12 6
Schulte Philipp	T 22 6
Schulte-Herbrüggen Thomas	
•MP 5.2 MP 5.3	
Schulz Andreas	•DD 2 2
Schulz, Oliver	T 25.5
Schulz, Tobias	•T 19.3
Schulz, Volkmar	ST 1.5
Schulze Eißing, Henning	•T 22.3,
Т 22.6	
Schulze, Florian	EP 6.6
Schulz-Schaeffer, Reinhard	DD 3.3,
DD 17.1, DD 37.2	
Schumacher, Lisa•T 23	s.4, •T 71.9
Scnumacher, Markus	5./, T 8.3,
	5.2,
I 15.5, I 89.6	T 100 0
Schumann, Jonannes	•I 103.3
Schufflählt, Martin	1 04.8
Schuster Sebestian	• GP 2 3
Schütter-Kerndl Britta	DD 34 4
Schütze Paul T 4	07 T 65 8
Schwäbig Patrick	•T 69 7
Schwandt, Joern	5 2 T 66 8
Schwandt, Jörn – T 12.8, T 12	2.9, T 13.3,
Schwandt, Jörn T 12.8, T 12 T 66.1, T 66.5	2.9, T 13.3,
Schwandt, Jörn T 12.8, T 12 T 66.1, T 66.5 Schwanenberger, Christian	2.9, T 13.3, T 33.1,
Schwandt, Jörn T 12.8, T 12 T 66.1, T 66.5 Schwanenberger, Christian T 36.1, T 76.7	2.9, T 13.3, T 33.1,
Schwandt, Jorn T 12.8, T 12 T 66.1, T 66.5 Schwanenberger, Christian T 36.1, T 76.7 Schwanke, Hagen	2.9, T 13.3, T 33.1, . •DD 17.2
Schwandt, Jorn T 12.8, T 12 T 66.1, T 66.5 Schwanenberger, Christian T 36.1, T 76.7 Schwanke, Hagen Schwartländer, Daniel	T 33.1, T 33.1, DD 17.2 T 31.3
Schwandt, Jorn T 12.8, T 12 T 66.1, T 66.5 Schwanenberger, Christian T 36.1, T 76.7 Schwanke, Hagen Schwartländer, Daniel Schwarz, DennisT	T 33.1, T 33.1, T 31.3 T 31.3 7.2, T 32.6
Schwandt, Jorn T 12.8, T 12 T 66.1, T 66.5 Schwanenberger, Christian T 36.1, T 76.7 Schwanke, Hagen Schwartländer, Daniel Schwarz, Dennis Schwarz, Mario	T 33.1, T 33.1, T 31.3, T 31.3 7.2, T 32.6 3.7, T 78.4,
Schwandt, Jorn 12.8, F12 T 66.1, T 66.5 Schwanenberger, Christian T 36.1, T 76.7 Schwanke, Hagen Schwartländer, Daniel Schwarz, Dennis Schwarz, Mario T 102.7	T 33.1, T 33.1, T 31.3 T 31.3 7.2, T 32.6 3.7, T 78.4,
Schwandt, Jorn T 12.8, T 12 T 66.1, T 66.5 Schwanenberger, Christian T 36.1, T 76.7 Schwanke, Hagen Schwartländer, Daniel Schwarz, Dennis ochwarz, Mario T 102.7 Schwarzwälder, Stefan	2.9, T 13.3, T 33.1, T 33.1, T 31.3 7.2, T 32.6 3.7, T 78.4, DD 5.2
Schwandt, Jorn 12.8, F12 T 66.1, T 66.5 Schwanenberger, Christian T 36.1, T 76.7 Schwanke, Hagen Schwartländer, Daniel Schwarz, Dennis Schwarz, Mario •T 102.7 Schwarzwälder, Stefan Schwarzwälder, Stefan	T 33.1, T 33.1, DD 17.2 T 31.3 7.2, T 32.6 3.7, T 78.4, DD 5.2 MP 11.1
Schwandt, Jorn 12.8, T12 T 66.1, T 66.5 Schwanenberger, Christian T 36.1, T 76.7 Schwanke, Hagen Schwartländer, Daniel Schwarz, Dennis Chwarz, Mario T 102.7 Schwarzwälder, Stefan Schween, Nils Walter Schwefer, Georg T 46.7 + T 72.4 T 102.8	2.9, T 13.3, T 33.1, T 33.1, T 31.3 7.2, T 32.6 3.7, T 78.4, DD 5.2 MP 11.1 3.8, T 46.1,
Schwandt, Jorn T 12.8, T 12 T 66.1, T 66.5 Schwanenberger, Christian T 36.1, T 76.7 Schwanke, Hagen Schwartländer, Daniel Schwarz, Dennis T Schwarz, Mario T 48 •T 102.7 Schwarzwälder, Stefan Schween, Nils Walter Schwefer, Georg T 18 T 46.7, •T 72.4, T 102.8 Schweizer Konstantin T 20	2.9, T 13.3, T 33.1, T 33.1, T 31.3 7.2, T 32.6 8.7, T 78.4, DD 5.2 •MP 11.1 3.8, T 46.1,
Schwandt, Jorn 112.8, F12 T 66.1, T 66.5 Schwanenberger, Christian T 36.1, T 76.7 Schwarke, Hagen Schwartländer, Daniel Schwarz, Dennis Schwarz, Mario T 102.7 Schwarz, Mario T 102.7 Schwarzwälder, Stefan Schwefer, Georg Schwefer, Georg T 46.7, •T 72.4, T 102.8 Schweizer, Konstantin T 20 T 99.5	2.9, T 13.3, T 33.1, DD 17.2 T 31.3 7.2, T 32.6 3.7, T 78.4, DD 5.2 MP 11.1 3.8, T 46.1, 0.8, T 48.7,
Schwandt, Jorn 112.8, F12 T 66.1, T 66.5 Schwanenberger, Christian T 36.1, T 76.7 Schwarke, Hagen Schwartländer, Daniel Schwarz, Mario +T 102.7 Schwarz, Mario +T 102.7 Schwarzwälder, Stefan Schween, Nils Walter T 46.7, +T 72.4, T 102.8 Schweizer, Konstantin T 20 T 99.5 Schweizer, Luca Schweizer, Schweizer, Luca Schweizer, Schweizer, Schweize	2.9, T 13.3, T 33.1, DD 17.2 T 31.3 7.2, T 32.6 3.7, T 78.4, DD 5.2 MP 11.1 3.8, T 46.1, 0.8, T 48.7, 0.8, T 48.9,
Schwandt, Jorn T 12.8, T 12 T 66.1, T 66.5 Schwanenberger, Christian T 36.1, T 76.7 Schwanke, Hagen Schwartländer, Daniel Schwarz, Dennis Schwarz, Mario Schwarz, Mario Schwarz, Mario Schwarz, Mario Schwerz, Mario Schwerz, Mario Schwarzwälder, Stefan Schweren, Nils Walter Schweizer, Konstantin 7 20 T 46.7, •T 72.4, T 102.8 Schweizer, Konstantin Schweizer, Konstantin 7 20 T 99.5 Schweizer, Luca	, T 33.3, T 33.1, T 33.1, T 31.3, 7.2, T 31.3, 7.2, T 32.6 3.7, T 78.4, DD 5.2 MP 11.1, 3.8, T 46.1, 0.8, T 48.9,
Schwandt, Jorn 12.8, F12 T 66.1, T 66.5 Schwanenberger, Christian T 36.1, T 76.7 Schwarke, Hagen Schwart, Dennis Schwarz, Mario •T 102.7 Schwarz, Mario •T 102.7 Schwarzwälder, Stefan Schween, Nils Walter T 46.7, •T 72.4, T 102.8 Schweizer, Konstantin T 20 T 99.5 Schweizer, Luca C 1 99.5 Schweizer, Luca Schweizer, Christina	2.9, T 13.3, T 33.1, .•DD 17.2, T 31.3, 7.2, T 32.6 3.7, T 78.4, DD 5.2 .•MP 11.1, 3.8, T 46.1, 0.8, T 48.7, 0.8, T 48.9, T 66.6,
Schwandt, Jorn 12.8, F12 T 66.1, T 66.5 Schwanenberger, Christian T 36.1, T 76.7 Schwarke, Hagen Schwart, Dennis Schwarz, Mario •T 102.7 Schwarz, Mario •T 102.7 Schwarz, Mario Schwarz, Mario •T 102.7 Schwarz, Mario Schwefer, Georg T 46.7, •T 72.4, T 102.8 Schweizer, Konstantin T 20 T 99.5 Schweizer, Luca C T 99.5 Schweizer, Luca T 48.10, T 99.5 Schwembauer, Christina T 66.7	2.9, T 13.3, T 33.1, .•DD 17.2, T 31.3, 7.2, T 32.6 3.7, T 78.4, DD 5.2 .•MP 11.1, 3.8, T 46.1, 0.8, T 48.7, 0.8, T 48.9, T 66.6,
Schwandt, Jorn 112.8, F12 T 66.1, T 66.5 Schwanenberger, Christian T 36.1, T 76.7 Schwarke, Hagen Schwartländer, Daniel Schwarz, Dennis Schwarz, Mario T 102.7 Schwarz, Mario Schwarzwälder, Stefan Schweren, Nils Walter Schwefer, Georg Schweizer, Konstantin T 20 T 99.5 Schweizer, Luca T 48.10, T 99.5 Schweizer, Luca T 48.10, T 99.5 Schweizer, Christina T 66.7 Schwemmbauer, Christina T 66.7	, T 13.3, T 33.1, T 33.1, T 31.3, 7.2, T 32.6 3.7, T 78.4, DD 5.2 MP 11.1 3.8, T 46.1, 0.8, T 48.7, 1.8, T 48.9, T 66.6, T 21.7
Schwandt, Jorn 112.8, F12 T 66.1, T 66.5 Schwanenberger, Christian T 36.1, T 76.7 Schwanke, Hagen Schwartländer, Daniel Schwarz, Dennis Schwarz, Mario T 102.7 Schwarz, Mario T 102.7 Schweren, Nils Walter Schwefer, Georg Schwefer, Georg Schweizer, Konstantin T 20 T 48.10, T 99.5 Schwembauer, Christina T 66.7 Schwenker, Benjamin T 84 T 40.7 Schwenker, Benjamin T 84 T 46.7 Schwenker, Benjamin T 84 T 46 T 46.7 Schwenker, Benjamin T 84 T 46 T 46 T 48 T 46 T 48 T	2.9, T 13.3, T 33.1, T 33.1, T 31.3, 7.2, T 32.6 3.7, T 78.4, DD 5.2 MP 11.1 3.8, T 46.1, 9.8, T 48.9, 9.8, T 48.9, T 66.6, T 21.7 4.9, T 95.1,
Schwandt, Jorn 112.8, F12 T 66.1, T 66.5 Schwanenberger, Christian T 36.1, T 76.7 Schwarke, Hagen Schwarz, Dennis Schwarz, Mario •T 102.7 Schwarz, Mario •T 102.7 Schwarz, Mario •T 102.7 Schwerer, Georg Schweizer, Konstantin T 46.7, •T 72.4, T 102.8 Schweizer, Konstantin T 46.7, •T 72.4, T 102.8 Schweizer, Konstantin T 48.10, T 99.5 Schwemmbauer, Christina T 66.7 Schwenmer, Alessandro Schwenker, Benjamin T 84 T 95.4, T 98.3 Schwenker, Sensanta Schwenker, Sensanta	2.9, T 13.3, T 33.1, T 33.1, T 31.3, 7.2, T 31.3, 7.2, T 32.6 3.7, T 78.4, DD 5.2 MP 11.1 3.8, T 46.1, 0.8, T 48.9, 0.8, T 48.9, T 66.6, T 21.7 1.9, T 95.1,
Schwandt, Jorn 112.8, F12 T 66.1, T 66.5 Schwanenberger, Christian T 36.1, T 76.7 Schwanke, Hagen Schwartländer, Daniel Schwarz, Mario Schwarz, Mario T 102.7 Schwarz, Mario Schwarz, Mario Schwerz, Mario Schweizer, Kostantin T 20 T 46.7, •T 72.4, T 102.8 Schweizer, Konstantin T 20 Schweizer, Luca T 48.10, T 99.5 Schwenmbauer, Christina T 66.7 Schwenmer, Alessandro Schwenker, Benjamin T 95.4, T 98.3 Schwerdt, C. Schwerdt, Caralia	2.9, T 13.3, T 33.1, T 33.1, T 31.3, 7.2, T 32.6 3.7, T 78.4, DD 5.2 MP 11.1, 3.8, T 46.1, 0.8, T 48.9, 0.8, T 48.9, T 66.6, T 21.7 4.9, T 95.1, EP 5.2
Schwandt, Jorn 112.8, T12 T 66.1, T 66.5 Schwanenberger, Christian T 36.1, T 76.7 Schwanke, Hagen Schwark, Hagen Schwartländer, Daniel Schwarz, Mario Schwerz, Mario Schwere, Nils Walter	2.9, T 13.3, T 33.1, T 33.1, T 31.3, 7.2, T 31.3, 7.2, T 32.6, 3.7, T 78.4, DD 5.2 MP 11.1, 3.8, T 46.1, 0.8, T 48.7, 0.8, T 48.7, 0.8, T 48.7, 0.8, T 48.9, T 66.6, T 21.7, 4.9, T 95.1, EP 5.2 2.3, T 52.4
Schwandt, Jorn 112.8, I12 T 66.1, T 66.5 Schwanenberger, Christian T 36.1, T 76.7 Schwanke, Hagen Schwartländer, Daniel Schwarz, Dennis Schwarz, Mario • T 102.7 Schwarzwälder, Stefan Schwerz, Georg Schweizer, Georg Schweizer, Konstantin T 46.7, • T 72.4, T 102.8 Schweizer, Konstantin Schweizer, Luca T 99.5 Schwenker, Benjamin T 66.7 Schwenker, Benjamin T 95.4, T 98.3 Schwerdt, Carolin Schwerdt, Carolin Schwerdt, Carolin	2.9, T 13.3, T 33.1, T 33.1, T 31.3, 7.2, T 31.3, 7.2, T 32.6 3.7, T 78.4, DD 5.2 MP 11.1 3.8, T 48.7, 0.8, T 48.7, 0.8, T 48.7, 0.8, T 48.9, T 66.6, T 21.7 1.9, T 95.1, EP 5.2 2.3, T 52.4 GR 6.2 1.7 52.4
Schwandt, Jorn 112.8, F12 T 66.1, T 66.5 Schwanenberger, Christian T 36.1, T 76.7 Schwanke, Hagen Schwanke, Hagen Schwartländer, Daniel Schwarz, Dennis T 12.7 Schwarz, Mario T 44 •T 102.7 Schwarzwälder, Stefan Schwerz, Mario T 44 •T 102.7 Schwerzwälder, Stefan Schween, Nils Walter Schwefer, Georg Schweizer, Konstantin T 20 Y 99.5 Schweizer, Konstantin T 20 Schweizer, Luca T 20 T 48.10, T 99.5 Schwemmbauer, Christina T 66.7 Schwenker, Benjamin T 84 Schwerdt, C arolin T 83 Schwerdt, C arolin T 55.4, T 98.3 Schwerdt, Carolin T 55.5 Schwersenz, Nico T 55.5 Schwersenz, Nico T 55.5 Schwersenz, Mico T 55.5 Schwerdt, Carolin T 55.5 Schwerdt, Carolin T 55.5 Schwerdt, Carolin T 55.5 Schwerdt, Kristref <td< td=""><td>, T 13.3, , T 13.3, , T 13.3, , T 13.3, , T 31.3, 7.2, T 32.6 3.7, T 78.4, , DD 5.2 , MP 11.1 3.8, T 48.7, 0.8, T 48.7, 0.8, T 48.7, 0.8, T 48.9, , T 66.6, , T 21.7 1.9, T 95.1, , EP 5.2 2.3, T 52.4 , GR 6.2 1.1, T 95.4 T 10.5</td></td<>	, T 13.3, , T 13.3, , T 13.3, , T 13.3, , T 31.3, 7.2, T 32.6 3.7, T 78.4, , DD 5.2 , MP 11.1 3.8, T 48.7, 0.8, T 48.7, 0.8, T 48.7, 0.8, T 48.9, , T 66.6, , T 21.7 1.9, T 95.1, , EP 5.2 2.3, T 52.4 , GR 6.2 1.1, T 95.4 T 10.5
Schwandt, Jorn 112.8, F12 T 66.1, T 66.5 Schwanenberger, Christian T 36.1, T 76.7 Schwanke, Hagen Schwarz, Dennis Schwarz, Mario •T 102.7 Schwarz, Mario Schwarz, Mario •T 102.7 Schwarz, Mario Schwerz, Mario Schwer, Stefan Schwer, Feorg Schweizer, Konstantin 7 99.5 Schweizer, Luca 7 99.5 Schwenmer, Alessandro Schwenmer, Alessandro Schwendt, Carolin Schwerdt, Carolin<	2.9, T 13.3, T 33.1, T 33.1, T 31.3, 7.2, T 32.6 3.7, T 78.4, DD 5.2 MP 11.1 3.8, T 48.7, 1.8, T 48.7, 1.8, T 48.9, T 66.6, T 21.7 1.9, T 95.1, EP 5.2 2.3, T 52.4 • GR 6.2 • T 95.4 T 10.6 • T 471
Schwandt, Jorn 112.8, F12 T 66.1, T 66.5 Schwanenberger, Christian T 36.1, T 76.7 Schwanke, Hagen Schwartländer, Daniel Schwarz, Mario Schwarz, Mario •T 102.7 Schwarz, Mario Schwerz, Mario •T 102.7 Schwerz, Mario Schweizer, Korstantin T 20, T 99.5 Schweizer, Konstantin Schweizer, Luca T 48.10, T 99.5 Schwemmbauer, Christina T 66.7 Schwemmer, Alessandro Schwerdt, C Schwerdt, Carolin Schwerdt, Carolin Schwerdt, Carolin Schwerker, Penjamin Schwerdt, Carolin Schwerdt, Carolin Schwerdt, Carolin Schwerdt, Kistof Schweiker, Marina Scoriavacche, Marina	2.9, T 13.3, T 33.1, T 33.1, T 31.3, 7.2, T 31.3, 7.2, T 32.6 3.7, T 78.4, DD 5.2 MP 11.1, 3.8, T 48.7, 1.8, T 48.7, 1.8, T 48.7, 1.8, T 48.7, 1.9, T 95.1, F 5.2 2.3, T 52.4 F 5.2 2.3, T 52.4 F 95.4 T 10.6 T 10.6 T 10.7
Schwandt, Jorn 112.8, I12 T 66.1, T 66.5 Schwanenberger, Christian T 36.1, T 76.7 Schwanke, Hagen Schwarke, Hagen Schwartländer, Daniel Schwarz, Mario Schwerz, Mario	2.9, T 13.3, T 33.1, T 33.1, T 31.3, 7.2, T 31.3, 7.2, T 32.6, 3.7, T 78.4, DD 5.2 MP 11.1, 3.8, T 46.1, 0.8, T 48.7, 0.8, T 48.7, 0.8, T 48.7, 0.8, T 48.9, T 66.6, • T 21.7, 4.9, T 52.4 • GR 6.2, • GR 6.2, • GR 6.2, • T 10.6 • T 10.6 • T 10.6 • T 10.2 • T
Schwandt, Jorn 112.8, I12 T 66.1, T 66.5 Schwanenberger, Christian T 36.1, T 76.7 Schwanke, Hagen Schwartländer, Daniel Schwarz, Dennis Schwarz, Mario •T 102.7 Schwarzwälder, Stefan Schwarzwälder, Stefan Schwerz, Georg Schweizer, Konstantin T 46.7, •T 72.4, T 102.8 Schweizer, Konstantin Schweizer, Konstantin T 95.5 Schweizer, Konstantin T 95.5 Schwenker, Benjamin T 66.7 Schwenker, Benjamin T 67.7 Schwenker, Benjamin Schwerdt, C. Schwerdt, Carolin Schwerdt, Carolin Schwerdt, Carolin Schwerdt, Kristof Schwerdt, Marike Schwersenz, Nico Schwersenz, Nico Schwersenz, Nico Schwersenz, Nico Schwersenz, Nico Schwersenz, Marike Schwersenz, Marike Schwersenz	2.9, T 13.3, T 33.1, T 33.1, T 31.3, 7.2, T 32.6 3.7, T 78.4, DD 5.2 MP 11.1, 3.8, T 46.1, 0.8, T 48.7, 0.8, T 48.7, 0.8, T 48.9, T 66.6, •T 21.7, 1.9, T 95.1, • GR 6.2 • GR 6.2 • T 95.4 T 10.6 • T 47.1 T 11.2 AGPhil 4.3 • T 31.2
Schwandt, Jorn 112.8, I12 T 66.1, T 66.5 Schwanenberger, Christian T 36.1, T 76.7 Schwanke, Hagen Schwanke, Hagen Schwartländer, Daniel Schwarz, Dennis T 12.7 Schwarz, Mario T 44 •T 102.7 Schwarzwälder, Stefan Schwarzwälder, Stefan Schwerz, Georg Schweizer, Georg T 18 T 46.7, •T 72.4, T 102.8 Schweizer, Konstantin Schweizer, Konstantin T 20 T 99.5 Schweizer, Kuca Schweizer, Kuca T 20 T 48.10, T 99.5 Schwemmbauer, Christina T 66.7 Schwenker, Benjamin T 84 T 95.4, T 98.3 Schwerdt, C Schwerdt, Schwert, Schwerdt, Schwerdt, Schwert, Schwerdt, Schwert, Schwerdt, Schwert, Schwerdt, Schwert, Schwerdt, Schwer	, T 13.3, , T 13.3, , T 13.3, , T 13.3, , T 13.3, 7.2, T 32.6 3.7, T 78.4, , DD 5.2 , MP 11.1 3.8, T 48.7, 0.8, T 48.7, 0.8, T 48.7, 0.8, T 48.7, 0.8, T 48.9, , T 66.6, , T 21.7 1.9, T 95.1, , EP 5.2 2.3, T 52.4 , GR 6.2 1.1, •T 95.4 , T 91.2 AGPhil 4.3 , T 91.2 AGPhil 4.3 , T 91.2
Schwandt, Jorn 112.8, I12 T 66.1, T 66.5 Schwanenberger, Christian T 36.1, T 76.7 Schwanke, Hagen Schwartländer, Daniel Schwarz, Dennis Schwarz, Mario •T 102.7 Schwarz, Mario Schwarz, Mario •T 102.7 Schwarz, Mario Schwerz, Mario Schwer, Nils Walter Schwer, Georg T 46.7, •T 72.4, T 102.8 Schweizer, Konstantin T 20 T 99.5 Schweizer, Konstantin T 20 T 48.10, T 99.5 Schwemmbauer, Christina T 66.7 Schwenker, Benjamin T 84 T 95.4, T 98.3 Schwerdt, Carolin T 55 Schwerdt, Carolin T 55 Schwerdt, Kristof Scornavacche, Marina Scyboz, Ludovic \$ Seebar, Charles • Seebarder, Norbert \$ Sefword, Felix \$	2.9, T 13.3, T 33.1, T 33.1, T 33.3, 7.2, T 32.6 3.7, T 78.4, DD 5.2 MP 11.1 3.8, T 48.7, 0.8, T 48.7, 0.8, T 48.7, 0.8, T 48.9, T 66.6, T 21.7 1.9, T 95.1, EP 5.2 2.3, T 52.4 GR 6.2 2.3, T 52.4 T 95.4 T 10.6 T 95.4 T 95.4
Schwandt, Jorn 112.8, F12 T 66.1, T 66.5 Schwanenberger, Christian T 36.1, T 76.7 Schwanke, Hagen Schwartländer, Daniel Schwarz, Mario Schwarz, Mario •T 102.7 Schwarz, Mario Schwarz, Mario •T 102.7 Schwerz, Mario Schweizer, Korstantin 7 46.7 Schweizer, Luca T 48.10, T 99.5 Schwemmer, Alessandro Schwerdt, C. Schwerdt, C. Schwerdt, Carolin Schwerdt, Carolin Schwerdt, Marike Schwerdt, Marina Scyboz, Ludovic Sethen, Charles seybaz, Ludovic Sefkow, Felix Sefkow, Felix	2.9, T 13.3, T 33.1, T 33.1, T 31.3, 7.2, T 31.3, 7.2, T 32.6 3.7, T 78.4, DD 5.2 MP 11.1, 3.8, T 48.7, 1.8, T 48.7, 1.8, T 48.7, 1.8, T 48.7, 1.8, T 48.7, 1.9, T 95.1, F 5.2 2.3, T 52.4 F 10.6 F 12.1 F 12.1 AGPhil 4.3 F 13.4, T 34.1 F 12.1 F 12.1 F 13.4 F 12.1 F 12.1 F 12.1 F 13.1 F 12.1 F 12.1 F 13.1 F 12.1
Schwandt, Jorn 112.8, I12 T 66.1, T 66.5 Schwanenberger, Christian T 36.1, T 76.7 Schwanke, Hagen Schwartländer, Daniel Schwarz, Mario Schwarz, Mario T 102.7 Schwarz, Mario Schwarz, Mario Schwerz, Mario Schweizer, Korstantin T 20.7 Schweizer, Konstantin Schweizer, Luca T 48.0, T 99.5 Schwemmer, Alessandro Schwendt, Carolin Schwerdt, Carolin Schwerdt, Carolin Schwerdt, Carolin Schwerdt, Marike Schwerker, Benjamin Schwerker, Benjamin Schwerdt, Carolin Schwerdt, Carolin Schwerdt, Carolin Schwerdt, Carolin Schwerdt, Marike Schwerdt, Marina Scyboz, Ludovic Seehanfer,	2.9, T 13.3, T 33.1, T 33.1, T 31.3, 7.2, T 31.3, 7.2, T 32.6 8.7, T 78.4, DD 5.2 MP 11.1, 3.8, T 48.7, 0.8, T 47.1, 0.9, T 47.1
Schwandt, Jorn 112.8, I12 T 66.1, T 66.5 Schwanenberger, Christian T 36.1, T 76.7 Schwanke, Hagen Schwartländer, Daniel Schwarz, Dennis Schwarz, Mario •T 102.7 Schwarz, Mario Schwarz, Mario •T 102.7 Schwarz, Mario Schwerz, Monstantin T 46.7, •T 72.4, T 102.8 Schweizer, Konstantin Schweizer, Konstantin T 95.5 Schweizer, Konstantin T 66.7 Schwenter, Benjamin T 66.7 Schwenker, Benjamin Schwersenz, Nico Schwersenz, Nico Schwersenz, Nico Schwerdt, C Schwersenz, Nico Schwerdt, Marike Schwersenz, Nico Schwersenz, Nico Schwersenz, Nico <	2.9, T 13.3, T 33.1, T 33.1, T 31.3, 7.2, T 32.6 3.7, T 78.4, DD 5.2 MP 11.1 3.8, T 46.1, 0.8, T 48.7, 0.8, T 48.7, 0.8, T 48.9, T 66.6, T 66.6, T 66.6, T 66.6, T 66.6, T 66.6, T 66.6, T 10.6 T 10.6 T 10.7 T 91.2 AGPhil 4.3 T 11.2 T 14.3 MP 4.3 5.2, T 76.2
Schwandt, Jorn 112.8, I12 T 66.1, T 66.5 Schwanenberger, Christian T 36.1, T 76.7 Schwanke, Hagen Schwartländer, Daniel Schwarz, Dennis Schwarz, Mario •T 102.7 Schwarz, Mario Schwarz, Mario •T 102.7 Schwarz, Mario Schwerz, Mario Schwert, Stefan Schwerzer, Konstantin T 20.7 Schwerzer, Konstantin T 48.10, T 99.5 Schwermbauer, Christina T 66.7 Schwenker, Benjamin Schwerdt, C Schwerdt, C Schwerdt, Carolin	2.9, T 13.3, T 33.1, T 33.1, T 31.3, 7.2, T 32.6 3.7, T 78.4, DD 5.2 MP 11.1 3.8, T 46.1, 0.8, T 48.7, 0.8, T 48.7, 0.8, T 48.7, 0.8, T 48.7, 0.8, T 48.7, 0.8, T 48.7, 0.8, T 48.7, 1.9, T 95.1, F 5.2 2.3, T 52.4 GR 6.2 (1, •T 95.4, T 10.6 (1, •T 95.4, T 95.4, T 95.4 T 91.2 AGPhil 4.3 T 71.4 T 91.2 AGPhil 4.3 T 71.4 T 91.2 AGPhil 4.3 T 71.4 T 91.5 T 95.4 T 95.4 T 10.5 T 11.5 T 11.5
Schwandt, Jorn 112.8, I12 T 66.1, T 66.5 Schwanenberger, Christian T 36.1, T 76.7 Schwanke, Hagen Schwartländer, Daniel Schwarz, Dennis Schwarz, Mario •T 102.7 Schwarz, Mario Schwarz, Mario •T 102.7 Schwarz, Mario Schwer, Nils Walter Schween, Nils Walter Schweizer, Konstantin T 46.7, •T 72.4, T 102.8 Schweizer, Konstantin T 99.5 Schweizer, Konstantin T 48.10, T 99.5 Schwemmbauer, Christina T 66.7 Schwenker, Benjamin Schwerdt, C Schwerdt, Carolin Schwerdt, Carolin Schwerdt, Carolin Schwerdt, Carolin Schwerdt, Kristof Scornavacche, Marina Scyboz, Ludovic Seehafer, Norbert Sefkow, Felix Sehgal, Srijan Seitz, Claudia Seitz, Claudia Seitz, Nicolas	2.9, T 13.3, T 33.1, T 33.1, DD 17.2 T 31.3 7.2, T 32.6 3.7, T 78.4, DD 5.2 MP 11.1 3.8, T 46.1, 0.8, T 48.9, T 66.6, T 21.7 1.9, T 95.1, GR 6.2 2.3, T 52.4 2.3, T 52.4 GR 6.2 2.3, T 52.4 GR 6.2 2.3, T 52.4 F 95.4 T 10.6 T 10.6 T 11.7 AGPhil 4.3 T 13.4 T 1.4 T 1.4
Schwandt, Jorn 112.8, I12 T 66.1, T 66.5 Schwanenberger, Christian T 36.1, T 76.7 Schwanke, Hagen Schwartländer, Daniel Schwarz, Mario Schwarz, Mario •T 102.7 Schwarz, Mario Schwarz, Mario •T 102.7 Schwerz, Mario Schwerz, Mario Schwer, Nils Walter Schwer, Stefan Schwer, Feorg T 46.7, •T 72.4, T 102.8 Schweizer, Konstantin T 20 T 99.5 Schweizer, Luca T 99.5 Schwemmbauer, Christina T 66.7 Schwenker, Benjamin Schwerdt, C Schwerdt, Carolin Schwerdt, Carolin Schwerdt, Kristof Scornavacche, Marina Scyboz, Ludovic Sefkow, Felix Sefkow, Felix Sehgal, Srijan Seitz, Nicolas Seitz, Nicolas Sekmen, Sezen Serverent Seftow, Feliz	2.9, T 13.3, T 33.1, T 33.1, DD 17.2 T 31.3 7.2, T 32.6 3.7, T 78.4, DD 5.2 MP 11.1 3.8, T 46.1, 0.8, T 48.9, T 66.6, T 21.7 1.9, T 95.1, GR 6.2 2.3, T 52.4 GR 6.2 2.3, T 52.4 T 91.2 4GPhil 4.3 T 91.2 AGPhil 4.3 T 71.4 MP 4.3 5.2, T 76.2 T 9.3 T 51.3 T 51.3 T 5.1
Schwandt, Jorn 112.8, I12 T 66.1, T 66.5 Schwanenberger, Christian T 36.1, T 76.7 Schwanke, Hagen Schwartländer, Daniel Schwarz, Mario Schwarz, Mario •T 102.7 Schwarz, Mario Schwarz, Mario •T 102.7 Schwerz, Mario Schwerz, Mario Schwerz, Mario Schwerz, Mario Schwerz, Mario Schwerz, Mario Schweizer, Konstantin T 20 T 46.7, •T 72.4, T 102.8 Schweizer, Konstantin Schweizer, Konstantin T 20 T 48.10, T 99.5 Schwemmbauer, Christina T 66.7 Schwemmer, Alessandro Schwerdt, Carolin Schwerdt, Carolin Schwerdt, Carolin Schwerdt, Marike Schwerdt, Marina Scyboz, Ludovic Seehanfer, Norbert Seehafer, Norbert Sefkow, Felix Sehens, Charles Seitz, Nicolas Seitz, Nicolas	2.9, T 13.3, T 33.1, T 33.1, T 31.3, T 31.3, 7.2, T 31.3, 7.2, T 32.6 3.7, T 78.4, DD 5.2 MP 11.1, 3.8, T 48.7, 1.8, T 48.7, 1.8, T 48.7, 1.8, T 48.7, 1.8, T 48.7, 1.8, T 48.7, 1.8, T 48.7, 1.9, T 95.1, F 5.2, F 5.2, F 5.2, T 91.2, AGPhil 4.3, T 91.2, AGPhil 4.3, T 91.2, AGPhil 4.3, T 71.4, T 71.4, T 71.4, T 71.4, T 71.4, T 71.4, T 71.2, T 71.2, T 91.2, T 91.2
Schwandt, Jorn 112.8, I12 T 66.1, T 66.5 Schwanenberger, Christian T 36.1, T 76.7 Schwanke, Hagen Schwartländer, Daniel Schwarz, Dennis Schwarz, Mario •T 102.7 Schwarz, Mario Schwarz, Mario •T 102.7 Schwarz, Mario Schwerz, Morstantin Schweizer, Konstantin T 46.7, •T 72.4, T 102.8 Schweizer, Konstantin Schweizer, Konstantin T 95.5 Schweizer, Konstantin T 66.7 Schwenter, Benjamin T 66.7 Schwenter, Christina T 66.7 Schwenter, Carolin Schwerker, Benjamin T 82 T 95.4, T 98.3 Schwerdt, Carolin Schwerdt, Carolin Schwerdt, Carolin Schwerdt, Marike	2.9, T 13.3, T 33.1, T 33.1, T 33.1, T 31.3, 7.2, T 32.6 3.7, T 78.4, DD 5.2 MP 11.1 3.8, T 46.1, 0.8, T 48.7, 0.8, T 48.7, 0.8, T 48.9, T 66.6, T 21.7 4.9, T 95.1, EP 5.2 2.3, T 52.4 GR 6.2 2.3, T 52.4 1, -T 95.4 4.1, -T 95.4 4.1, -T 95.4 4.3, -T 10.6 T 10.6 T 13.1 FP 12.11 T 13.1 T 13.1 T 13.1 T 13.3 T 15.1 T 13.3 T 15.1 T 13.3 T 13.1 T 13.3 T 13.1 T 13.3 T 13.1 T 13.3 T 13.1 T 13.1
Schwandt, Jorn 112.8, I12 T 66.1, T 66.5 Schwanenberger, Christian T 36.1, T 76.7 Schwanke, Hagen Schwartländer, Daniel Schwarz, Dennis Schwarz, Mario •T 102.7 Schwarz, Mario Schwarz, Mario •T 102.7 Schwarz, Mario Schwerz, Mario Schwert, Georg T 46.7, •T 72.4, T 102.8 Schweizer, Konstantin Schweizer, Konstantin T 20.5 Schwert, Stefan Schwert, Carolin Schwert, C Schwerdt, C Schwerdt, C Schwerdt, Carolin Schwerdt, Carolin <tr< td=""><td>2.9, T 13.3, T 33.1, T 33.1, T 33.1, T 31.3, 7.2, T 32.6 3.7, T 78.4, DD 5.2 MP 11.1 3.8, T 46.1, 0.8, T 48.7, 0.8, T 48.7, 0.8, T 48.7, 0.8, T 48.9, T 66.6, T 21.7, 1.9, T 95.1, EP 5.2 2.3, T 52.4 GR 6.2 (.1, •T 95.4, T 10.6 (•T 47.1 T 91.2 AGPhil 4.3 5.2, T 76.2 T 15.1 MP 4.3 5.2, T 76.2 T 15.3 AGPhil 7.4 .8, T 33.4, A T 90 E</td></tr<>	2.9, T 13.3, T 33.1, T 33.1, T 33.1, T 31.3, 7.2, T 32.6 3.7, T 78.4, DD 5.2 MP 11.1 3.8, T 46.1, 0.8, T 48.7, 0.8, T 48.7, 0.8, T 48.7, 0.8, T 48.9, T 66.6, T 21.7, 1.9, T 95.1, EP 5.2 2.3, T 52.4 GR 6.2 (.1, •T 95.4, T 10.6 (•T 47.1 T 91.2 AGPhil 4.3 5.2, T 76.2 T 15.1 MP 4.3 5.2, T 76.2 T 15.3 AGPhil 7.4 .8, T 33.4, A T 90 E
Schwandt, Jorn 112.8, I12 T 66.1, T 66.5 Schwanenberger, Christian T 36.1, T 76.7 Schwanke, Hagen Schwartländer, Daniel Schwarz, Dennis Schwarz, Mario •T 102.7 Schwarz, Mario Schwarz, Mario •T 102.7 Schwarz, Mario Schwerz, Mario Schwerzer, Koorstantin 7 99.5 Schwerzer, Luca 7 99.5 Schwermbauer, Christina 7 66.7 Schwendt, Carolin Schwerdt, Carolin <	2.9, T 13.3, T 33.1, T 33.1, T 33.1, T 31.3, 7.2, T 32.6 3.7, T 78.4, DD 5.2 MP 11.1 3.8, T 46.1, 0.8, T 48.9, T 66.6, T 21.7 1.9, T 95.1, GR 62, • GR 62, • GR 62, • GR 62, T 91.2, AGPhil 4.3 T 13.4, T 13.4, T 15.1, T 9.3, T 15.1, T 9.3, T 0.1 3, GP HI 7.4, T 9.3, T 15.1, T 9.3, T 0.1 3, GP HI 7.4, T 9.3, T 0.1 3, GP HI 7.4, T 9.3, T 15.1, T 0.1 3, GP HI 7.4, T 0.1 4, GP HI 7.4, T 0.1, T 0.1, T 0.1, T 0.1, T 0.1, T 0.1, T 0.1,
Schwandt, Jorn 112.8, I12 T 66.1, T 66.5 Schwanenberger, Christian T 36.1, T 76.7 Schwanke, Hagen Schwartländer, Daniel Schwarz, Mario Schwerz, Mario Schwerz, Mario Schwerz, Mario Schwer, Nils Walter Schwer, Stefan Schweizer, Konstantin T 46.7, •T 72.4, T 102.8 Schweizer, Luca T 20 T 48.10, T 99.5 Schweizer, Luca T 20 T 48.10, T 99.5 Schwenmer, Alessandro Schwendt, C. Schwendt, Carolin T 52 Schwerdt, Carolin T 52 Schwerelt, Kristof Scornavacche, Marina	2.9, T 13.3, T 33.1, T 33.1, T 33.1, T 31.3, 7.2, T 32.6 3.7, T 78.4, DD 5.2 MP 11.1 3.8, T 48.7, 1.8, T 48.7, 1.8, T 48.9, T 66.6, •T 21.7, 1.9, T 95.1, • GR 6.2 2.3, T 52.4 • GR 6.2 2.3, T 52.4 • GR 6.2 2.3, T 52.4 • GR 6.2 2.3, T 52.4 • T 95.4 • T 10.6 • T 10.6 • T 34.1 • T 10.4 • T 14.1 • T 14.1 • T 14.1 • T 15.1 • T 9.3 • T 15.1 • T 9.3 • T 15.1 • T 9.3 • T 15.1 • T 9.3 • T 26.7, 1, • T 26.7, • T 26.7,
Schwandt, Jorn 112.8, I12 T 66.1, T 66.5 Schwanenberger, Christian T 36.1, T 76.7 Schwanke, Hagen Schwartländer, Daniel Schwarz, Mario Schwarz, Mario •T 102.7 Schwarz, Mario Schwerz, Mario •T 102.7 Schwerz, Mario Schweizer, Konstantin 7 46.7 T 46.7, •T 72.4, T 102.8 Schweizer, Luca T 48.10, T 99.5 Schwemmer, Alessandro Schwerdt, C. Schwerdt, Carolin Schwerdt, Carolin Schwerdt, Carolin Schwerdt, Marike Schwerdt, Marina Scyboz, Ludovic Seebens, Charles Setiz, Nicolas Sethwer, Reix Sethwerdt, Strian Seitz, Nicolas Setrer, Robert <tr< td=""><td>2.9, T 13.3, T 33.1, T 33.1, T 31.3, 7.2, T 31.3, 7.2, T 32.6 3.7, T 78.4, DD 5.2 MP 11.1, 3.8, T 48.7, 1.8, T 48.7, 1.8, T 48.7, 1.8, T 48.7, 1.8, T 48.7, 1.8, T 48.7, 1.9, T 95.1, F 5.2, 2.3, T 52.4, F 0.1, 4.9, T 95.1, 2.3, T 52.4, F 0.1, 4.9, T 95.1, 2.3, T 52.4, F 0.1, 3 T 10.6, T 10.6, T 10.6, T 10.7, 3 T 11.7, AGPhil 4.3, 3 T 12.1, T 11.4, T 10.1, T 11.4, T 10.1, T 11.4, T 11.5, T 12.7, T 13.7, T 13.7, </td></tr<>	2.9, T 13.3, T 33.1, T 33.1, T 31.3, 7.2, T 31.3, 7.2, T 32.6 3.7, T 78.4, DD 5.2 MP 11.1, 3.8, T 48.7, 1.8, T 48.7, 1.8, T 48.7, 1.8, T 48.7, 1.8, T 48.7, 1.8, T 48.7, 1.9, T 95.1, F 5.2, 2.3, T 52.4, F 0.1, 4.9, T 95.1, 2.3, T 52.4, F 0.1, 4.9, T 95.1, 2.3, T 52.4, F 0.1, 3 T 10.6, T 10.6, T 10.6, T 10.7, 3 T 11.7, AGPhil 4.3, 3 T 12.1, T 11.4, T 10.1, T 11.4, T 10.1, T 11.4, T 11.5, T 12.7, T 13.7, T 13.7,
Schwandt, Jorn 112.8, I12 T 66.1, T 66.5 Schwanenberger, Christian T 36.1, T 76.7 Schwanke, Hagen Schwartländer, Daniel Schwarz, Dennis Schwarz, Mario •T 102.7 Schwarz, Mario Schwarz, Mario •T 102.7 Schwarz, Mario Schwarz, Mario •T 102.7 Schwarz, Mario Schwerz, Mario Schwerz, Mario Schwerz, Mario Schwerz, Morstantin Schweizer, Konstantin T 46.7, •T 72.4, T 102.8 Schweizer, Konstantin Schweizer, Konstantin T 95.5 Schweizer, Konstantin Schweizer, Konstantin T 95.5 Schweizer, Konstantin Schweizer, Konstantin T 95.5 Schweizer, Konstantin Schweizer, Konstantin Schweizer, Koristina T 66.7 Schwenth, Carolin Schwersenz, Nico Schwersenz, Nico Schwersenz, Nico Schwersenz, Nic	2.9, T 13.3, T 33.1, T 33.1, T 33.1, T 31.3, 7.2, T 32.6 3.7, T 78.4, DD 5.2 MP 11.1, 3.8, T 46.1, 0.8, T 48.9, 0.8, T 48.7, 0.8, T 48.7, 1.9, T 95.1, F 95.4, T 66.6, • T 21.7 4.9, T 95.1, • GR 6.2, 1.4, • T 95.4, • GR 6.2, 1.4, • T 95.4, • T 11.4, • T 91.2, AGPhil 4.3, • T 11.4, • T 91.2, AGPhil 7.4, • T 13.3, • T 14.1, • T 14.1, • T 14.4, • T 14.4, • T 14.3, • T 14.4, • T 26.7, • T 1.3, • ST 1.3, 3, T 6.0 4
Schwandt, Jorn 112.8, I12 T 66.1, T 66.5 Schwanenberger, Christian T 36.1, T 76.7 Schwanke, Hagen Schwanke, Hagen Schwartländer, Daniel Schwarz, Dennis Schwarz, Mario Schwarz, Mario Schwarz, Mario Schwarz, Mario Schwarz, Mario	2.9, T 13.3, T 33.1, T 33.1, T 33.1, T 31.3, 7.2, T 32.6 3.7, T 78.4, DD 5.2 MP 11.1 3.8, T 46.1, 0.8, T 48.7, 0.8, T 48.7, 0.8, T 48.9, T 66.6, T 166.6, T 166.6, T 66.6, T 66.6, T 66.6, T 66.6, T 10.6 T 9.12, AGPhil 4.3 T 10.1, MP 4.3, 56.2, T 76.2 T 15.1, MP 4.3, 36.2, T 76.2 T 15.1, DD 18.3 AGPhil 7.4, 4, T 88.5 1, • T 26.7, ST 1.3, 0.3, T 60.4

Sharshunova, Maria	T 23.4
Shaw, Vasundhara	•T 73.1
Shchedrolosiev, Mykyta	••••••••••••••••••••••••••••••••••••••
Shetali, Shetali	•T 100.8
CP 1 5	ivienai
Shekhzadeh Imahn	
Shigarov Alexev	AKPIK 1.5
Shprits. Yuri	EP 5.7
Shtembari, Lolian •T 10	5.5, •T 107.1
Sich, A	DD 34.3
Sich, Annemarie	DD 44.2
Sichau, Christian •DD 1	5.3, •GP 1.3
Siegert, Frank	T 88.3
Siegmann, Daniel	•T 95.8
Silva, Hector U	GR 5.5
Silva, Pedro	I /.5
Simon Hardy T 22.5 T	10.1, •1 40.3
Simon Frank T /1.2 T	22.7, 1 70.3 50 1 T 68 6
Simon Jan	•DD 23 2
Simon. Stevens	EP 2.2
Simonis, Hans Jürgen ST	6.4, T 65.5,
T 65.6, T 65.7, T 94.4	
Sindhu, Sreelakshmi •T	7.8, •T 70.4
Singh, Parampreet GR 1	I2.1, MP 9.1,
MP 10.2	
Singhal, Apeksha T2	20.2, T 20.5,
1 43.1, 1 103.7, •1 103.9 Sinha Supriva	
Sinnhuber Miriam	•1/./ 28 • EP 2 2
FP 5.5 FP 5.6	2.0, •LF 3.2,
Sirletti, Salvatore Samuele	•GR 3 4
Sirrenberg. Filip	DD 46 2
Sitnikova, Elizaveta	•T 10.3
Skambraks, Sebastian	T 43.8
Skidmore, Nicola	T 56.2
Skorenok, Oleksandr	T 79.1
Skorupa, Justin	•T 30.3
Skulberg, Emilie	•GP 9.4
Smilga, Walter	•GR 10.3
Smith, Eluned	
Sillyczek, David 1 3 Specke Welter	9.5, •1 39.0 T 40 9
Snoeys, Walter	I 40.8
T 72 3	31.0, 1 72.2,
Solanki Jav	GR 1 1
Solanki. Sami	EP 12.7
Soldin Philipp •T	
	48.1, T 48.2
Sole, Silvia	48.1, T 48.2
Sole, Silvia	48.1, T 48.2 T 4.4 T 76.2
Sole, Silvia	48.1, T 48.2 T 4.4 T 76.2 T 39.4
Sole, Silvia Solis, Alvaro Lopez Solis, Michelle Solomon, Shalu	48.1, T 48.2 T 4.4 T 76.2 T 39.4 •T 87.2
Sole, Silvia Solis, Alvaro Lopez Solis, Michelle Solomon, Shalu Solovieva, Ksenia	48.1, T 48.2 T 4.4 T 76.2 T 39.4 •T 87.2 39.1, •T 39.2
Sole, Silvia Solis, Alvaro Lopez Solis, Michelle Solomon, Shalu Solovieva, Ksenia Sommer, Philip	48.1, T 48.2 T 4.4 T 76.2 T 39.4 T 87.2 39.1, •T 39.2 T 29.1 T 29.1
Sole, Silvia Solis, Alvaro Lopez Solis, Michelle Solomon, Shalu Solovieva, Ksenia Sommer, Philip Sommerhalder, Manuel	48.1, T 48.2 T 4.4 T 76.2 T 39.4 T 87.2 39.1, •T 39.2 T 29.1 T 93.2 T 93.2 T 93.2
Sole, Silvia Sole, Silvia Solis, Alvaro Lopez Solis, Michelle Solovieva, Ksenia Somer, Philip Sommerhalder, Manuel Sonnabend, Robin	48.1, T 48.2 T 4.4 T 76.2 T 39.4 T 87.2 39.1, •T 39.2 •T 29.1 •T 93.2 •T 101.4
Sole, Silvia Sole, Silvia Solis, Alvaro Lopez Solis, Michelle Solovieva, Ksenia Solovieva, Ksenia Sommer, Philip Sommer, Abdin Sonnabend, Robin Sonntag, Dörte	48.1, T 48.2 T 4.4 T 76.2 T 39.4 T 87.2 T 87.2 T 29.1 T 29.1 T 93.2 T 101.4 DD 3.2 24.4 T 35.1
Sole, Silvia Sole, Silvia Solis, Alvaro Lopez Solis, Michelle Solovieva, Ksenia Sonore, Philip Sommer, Philip Sonnabend, Robin Sonntag, Dörte Sopczak, André T 10.9, T T 35.6, T 35.7, T 52.2	48.1, T 48.2 T 4.4 T 76.2 T 39.4 T 87.2 T 39.2 T 93.2 T 93.2 T 101.4 DD 3.2 24.4, T 35.1,
Sole, Silvia Sole, Silvia Solis, Alvaro Lopez Solis, Michelle Solovieva, Ksenia Solovieva, Ksenia Sommer, Philip Sommerhalder, Manuel Sonnabend, Robin Sonntag, Dörte Sopczak, André T 10.9, T T 35.6, T 35.7, T 52.2 Sorgenfrei, Niels	48.1, T 48.2
Sole, Silvia Sole, Silvia Solis, Alvaro Lopez Solis, Michelle Solovieva, Ksenia Sonorer, Philip Sommerhalder, Manuel Sonnabend, Robin Sonntag, Dörte Sopczak, André T 10.9, T T 35.6, T 35.7, T 52.2 Sorgenfrei, Niels Sorrenson, Peter T 10	48.1, T 48.2
Sole, Silvia Sole, Silvia Solis, Alvaro Lopez Solis, Michelle Solovieva, Ksenia Sonoreva, Ksenia Sommer, Philip Sommerhalder, Manuel Sonnabend, Robin Sonntag, Dörte Sopczak, André T 10.9, T T 35.6, T 35.7, T 52.2 Sorgenfrei, Niels Sorrenson, Peter Sothilingam, Varsiha	48.1, T 48.2 T 4.4 T 76.2
Sole, Silvia Sole, Silvia Solis, Alvaro Lopez Solis, Michelle Solovieva, Ksenia Sonorer, Philip Sommer, Philip Sommerhalder, Manuel Sonnabend, Robin Sonntag, Dörte Sopczak, André T 10.9, T T 35.6, T 35.7, T 52.2 Sorgenfrei, Niels Sorrenson, Peter Sothilingam, Varsiha Sotnikov, Vasily	48.1, T 48.2
Sole, Silvia Sole, Silvia Solis, Alvaro Lopez Solis, Michelle Solomon, Shalu Solovieva, Ksenia Sommer, Philip Sommerhalder, Manuel Sonnabend, Robin Sonntag, Dörte Sopczak, André T 10.9, T T 35.6, T 35.7, T 52.2 Sorgenfrei, Niels Sorrenson, Peter T 9 Sothilingam, Varsiha Sotnikov, Vasily Sowa, Janis	48.1, T 48.2
Sole, Silvia Sole, Silvia Solis, Alvaro Lopez Solis, Michelle Solovieva, Ksenia Solovieva, Ksenia Sommer, Philip Sommer, Philip Sonnabend, Robin Sonnabend, Robin Sonnabend, Robin Sontag, Dörte Sopczak, André T 10.9, T T 35.6, T 35.7, T 52.2 Sorgenfrei, Niels Sorrenson, Peter Sothilingam, Varsiha Sotnikov, Vasily Sowa, Janis	48.1, T 48.2
Sole, Silvia Sole, Silvia Solis, Alvaro Lopez Solis, Michelle Solorion, Shalu Solovieva, Ksenia Sommer, Philip Sommer, Philip Sommer, Robin Sonnabend, Robin Sonnabend, Robin Sonnabend, Robin Sontag, Dörte Sopczak, André T 10.9, T T 35.6, T 35.7, T 52.2 Sorgenfrei, Niels Sorrenson, Peter Sothilingam, Varsiha Sothikov, Vasily Sowa, Janis Sowa, Lars Soyke Daniel	48.1, T 48.2
Sole, Silvia Sole, Silvia Solis, Alvaro Lopez Solis, Michelle Solovieva, Ksenia Sonore, Philip Sommerhalder, Manuel Sonnabend, Robin Sonntag, Dörte Sopczak, André T 10.9, T T 35.6, T 35.7, T 52.2 Sorgenfrei, Niels Sorrenson, Peter Sothilingam, Varsiha Sothikov, Vasily Sowa, Janis Sowa, Lars Soybelman, Nathalie Sonan Bernhard	48.1, T 48.2
Sole, Silvia Sole, Silvia Solis, Alvaro Lopez Solis, Michelle Solovieva, Ksenia Solovieva, Ksenia Sommer, Philip Sommerhalder, Manuel Sonnabend, Robin Sonnabend, Robin Sontag, Dörte Sopczak, André T 10.9, T T 35.6, T 35.7, T 52.2 Sorgenfrei, Niels Sorrenson, Peter Sothilingam, Varsiha Sothilingam, Varsiha Sothilingam, Varsiha Sothilingam, Varsiha Sowa, Janis Soybelman, Nathalie Soyk, Daniel Span, Bernhard ST 8.2, ST 8.3, ST 8.4	48.1, T 48.2
Sole, Silvia Sole, Silvia Solis, Alvaro Lopez Solis, Michelle Solorieva, Ksenia Sonoreva, Ksenia Sommer, Philip Sommerhalder, Manuel Sonnabend, Robin Sonntag, Dörte Sopczak, André T 10.9, T T 35.6, T 35.7, T 52.2 Sorgenfrei, Niels Sorrenson, Peter Sothilingam, Varsiha Sotnikov, Vasily Sowa, Janis Sowa, Lars Soybelman, Nathalie Soyk, Daniel Spaan, Bernhard ST 8.2, ST 8.3, ST 8.4 Späh, Jan Lukas	48.1, T 48.2
Sole, Silvia Sole, Silvia Solis, Alvaro Lopez Solis, Michelle Solorieva, Ksenia Solovieva, Ksenia Sommer, Philip Sommerhalder, Manuel Sonnabend, Robin Sonntag, Dörte Sopczak, André T 10.9, T T 35.6, T 35.7, T 52.2 Sorgenfrei, Niels Sorrenson, Peter Sotnilingam, Varsiha Sotnikov, Vasily Sowa, Janis Sowa, Lars Soybelman, Nathalie Soyk, Daniel Spaan, Bernhard ST 8.2, ST 8.3, ST 8.4 Späh, Jan Lukas Spane, Felix EP 7.4, EP 7.4, EP 7.4, EP 7.4, EP	48.1, T 48.2
Sole, Silvia Sole, Silvia Solis, Alvaro Lopez Solis, Michelle Solowev, Ksenia Solovieva, Ksenia Sommer, Philip Sommer, Palip Sommer, André Sonnabend, Robin Sonnabend, Robin Sonnabend, Robin Sontag, Dörte Sopczak, André T 10.9, T T 35.6, T 35.7, T 52.2 Sorgenfrei, Niels Sorrenson, Peter Sothilingam, Varsiha Sotnikov, Vasily Sowa, Janis Sowa, Lars Soybelman, Nathalie Soyk, Daniel Spaan, Bernhard ST 8.2, ST 8.3, ST 8.4 Späh, Jan Lukas Sor EP 7.4, EP 7 Spannagel, Simon	48.1, T 48.2
Sole, Silvia Sole, Silvia Solis, Alvaro Lopez Solis, Michelle Solovieva, Ksenia Sonore, Philip Sommerhalder, Manuel Sonnabend, Robin Sonntag, Dörte Sopczak, André T 10.9, T T 35.6, T 35.7, T 52.2 Sorgenfrei, Niels Sorrenson, Peter Sothilingam, Varsiha Sotnikov, Vasily Sowa, Janis Sowa, Lars Soybelman, Nathalie Soyk, Daniel Spaan, Bernhard ST 8.2, ST 8.3, ST 8.4 Späh, Jan Lukas Spannagel, Simon Spannfellner, Christian	48.1, T 48.2
Sole, Silvia Sole, Silvia Solis, Alvaro Lopez Solis, Michelle Solorieva, Ksenia Solovieva, Ksenia Sommer, Philip Sommerhalder, Manuel Sonnabend, Robin Sonnabend, Robin Sontag, Dörte Sopczak, André T 10.9, T T 35.6, T 35.7, T 52.2 Sorgenfrei, Niels Sorrenson, Peter Sothilingam, Varsiha Sothilingam, Varsiha Sothikov, Vasily Sowa, Janis Sowa, Lars Soybelman, Nathalie Soybelman, Nathalie Soybelman, Nathalie Spaan, Bernhard ST 8.2, ST 8.3, ST 8.4 Späh, Jan Lukas Spannagel, Simon Spannfellner, Christian Spataro, David	48.1, T 48.2
Sole, Silvia Sole, Silvia Solis, Alvaro Lopez Solis, Michelle Solorieva, Ksenia Solovieva, Ksenia Sommer, Philip Sommerhalder, Manuel Sonnabend, Robin Sonnabend, Robin Sonnabend, Robin Sontag, Dörte Sopczak, André T 10.9, T T 35.6, T 35.7, T 52.2 Sorgenfrei, Niels Sorrenson, Peter Sothilingam, Varsiha Sotnikov, Vasily Sowa, Janis Sova, Janis Soya, Lars Soybelman, Nathalie Soyk, Daniel Spaan, Bernhard ST 8.2, ST 8.3, ST 8.4 Späh, Jan Lukas Spannagel, Simon Spannfellner, Christian Spatz, Verena DD 5.1, DD	48.1, T 48.2
Sole, Silvia Sole, Silvia Solis, Alvaro Lopez Solis, Michelle Solorieva, Ksenia Sonomer, Philip Sommer, Philip Sommer, Robin Sonnabend, Robin Sonnabend, Robin Sontag, Dörte Sopczak, André T 10.9, T T 35.6, T 35.7, T 52.2 Sorgenfrei, Niels Sorrenson, Peter Sothilingam, Varsiha Sotnikov, Vasily Sowa, Janis Sowa, Lars Soybelman, Nathalie Soyk, Daniel Spaan, Bernhard ST 8.2, ST 8.3, ST 8.4 Späh, Jan Lukas Spannagel, Simon Spannfellner, Christian Spataro, David Spat., Verena DD 19.2, DD 22.3, DD 25.1	48.1, T 48.2
Sole, Silvia Sole, Silvia Solis, Alvaro Lopez Solis, Michelle Solovieva, Ksenia Sonomer, Philip Sommer, Philip Sommer, Robin Sonnabend, Robin Sonnabend, Robin Sonzak, André T 10.9, T T 35.6, T 35.7, T 52.2 Sorgenfrei, Niels Sorrenson, Peter Sotrilingam, Varsiha Sotnikov, Vasily Sowa, Janis Sowa, Lars Soybelman, Nathalie Soyk, Daniel Soyk, Daniel St 8.2, ST 8.3, ST 8.4 Spän, Jan Lukas T 82, ST 8.3, ST 8.4 Spännfellner, Christian Spatz, Verena D 51, DD DD 19.2, DD 22.3, DD 25.1 Spector, Aaron	48.1, T 48.2
Sole, Silvia Sole, Silvia Solis, Alvaro Lopez Solis, Michelle Soloron, Shalu Solovieva, Ksenia Sommer, Philip Sommer, Philip Sommer, Robin Sonnabend, Robin Sonnabend, Robin Sonrag, Dörte Sopczak, André T 10.9, T T 35.6, T 35.7, T 52.2 Sorgenfrei, Niels Sorrenson, Peter Sotrilingam, Varsiha Sotnikov, Vasily Sowa, Janis Sowa, Lars Soybelman, Nathalie Soyk, Daniel Spaan, Bernhard ST 8.2, ST 8.3, ST 8.4 Späh, Jan Lukas Spanifellner, Christian Spanter, Felix Spantz, David Spanter, Christian Spatz, Verena D 19.2, DD 22.3, DD 25.1 Spector, Aaron Spenier, Hendrik	48.1, T 48.2
Sole, Silvia Sole, Silvia Solis, Alvaro Lopez Solis, Michelle Soloron, Shalu Solovieva, Ksenia Sommer, Philip Sommer, Philip Sommer, Robin Sonnabend, Robin Sonnabend, Robin Sontag, Dörte Sopczak, André T 10.9, T T 35.6, T 35.7, T 52.2 Sorgenfrei, Niels Sorrenson, Peter Sorrenson, Peter Sortalingam, Varsiha Sotnikov, Vasily Sowa, Janis Sowa, Janis Sowa, Lars Soybelman, Nathalie Soyk, Daniel Spaan, Bernhard ST 8.2, ST 8.3, ST 8.4 Späh, Jan Lukas Stater, Selix Spannfellner, Christian Spantel, Simon Spanfellner, Christian Spataro, David DD 19.2, DD 22.3, DD 25.1 Spector, Aaron Speer, Jannis Spencer, Benjamin A	48.1, T 48.2
Sole, Silvia Sole, Silvia Solis, Alvaro Lopez Solis, Michelle Solovieva, Ksenia Solovieva, Ksenia Sommer, Philip Sommerhalder, Manuel Sonnabend, Robin Sonnabend, Robin Sonnabend, Robin Sopczak, André T 10.9, T: T 35.6, T 35.7, T 52.2 Sorgenfrei, Niels Sorrenson, Peter Sothilingam, Varsiha Sotnikov, Vasily Sowa, Janis Sowa, Lars Soybelman, Nathalie Soyk, Daniel Spaan, Bernhard ST 8.2, ST 8.3, ST 8.4 Späh, Jan Lukas Spanier, Felix Spataro, David Spanfellner, Christian Spataro, David D 19.2, DD 22.3, DD 25.1 Spector, Aaron Speer, Jannis Spengler, Felix	48.1, T 48.2
Sole, Silvia Sole, Silvia Solis, Alvaro Lopez Solis, Michelle Solovieva, Ksenia Solovieva, Ksenia Sommer, Philip Sommerhalder, Manuel Sonnabend, Robin Sonnabend, Robin Sonnabend, Robin Sonzes, André T 10.9, T T 35.6, T 35.7, T 52.2 Sorgenfrei, Niels Sorrenson, Peter Sothilingam, Varsiha Sothilingam, Varsiha Sothilingam, Varsiha Sothilingam, Varsiha Sothilingam, Varsiha Sowa, Janis Sowa, Janis Sowa, Lars Soybelman, Nathalie Soyk, Daniel Spaan, Bernhard ST 8.2, ST 8.3, ST 8.4 Späh, Jan Lukas Spataro, David Spataro, David Spataro, David Spet, Jannis Speter, Hendrik Spencer, Benjamin A. Spenger, Felix	48.1, T 48.2
Sole, Silvia Sole, Silvia Solis, Alvaro Lopez Solis, Michelle Solorieva, Ksenia Solovieva, Ksenia Sommer, Philip Sommerhalder, Manuel Sonnabend, Robin Sonnabend, Robin Sonnabend, Robin Sontag, Dörte Sopczak, André T 10.9, T T 35.6, T 35.7, T 52.2 Sorgenfrei, Niels Sorrenson, Peter Sothilingam, Varsiha Sotnikov, Vasily Sowa, Janis Sowa, Lars Soybelman, Nathalie Soyk, Daniel Spaan, Bernhard ST 8.2, ST 8.3, ST 8.4 Späh, Jan Lukas Spannagel, Simon Spannfellner, Christian Spataro, David Spat, Verena DD 5.1, DD DD 19.2, DD 22.3, DD 25.1 Spector, Aaron Speer, Jannis Speiser, Hendrik Spenza, Enrico Speri, Lorenzo	48.1, T 48.2 T 44. T 4.4 T 76.2 .99.1, •T 39.2 T 29.1
Sole, Silvia Sole, Silvia Solis, Alvaro Lopez Solis, Michelle Solovieva, Ksenia Sonmer, Philip Sommer, Philip Sommer, Robin Sonnabend, Robin Sonnabend, Robin Sontag, Dörte Sopczak, André T 10.9, T T 35.6, T 35.7, T 52.2 Sorgenfrei, Niels Sorrenson, Peter Sotrilingam, Varsiha Sotnikov, Vasily Sowa, Janis Sowa, Lars Soybelman, Nathalie Soyk, Daniel Spaan, Bernhard ST 8.2, ST 8.3, ST 8.4 Späh, Jan Lukas T 6 Spanier, Felix Spannagel, Simon Spannfellner, Christian Spataro, David Spato, Javid Speetor, Aaron Speer, Jannis Speiser, Hendrik Spenier, Felix Spenza, Enrico Sperilch, Dennis T	48.1, T 48.2
Sole, Silvia Sole, Silvia Solis, Alvaro Lopez Solis, Michelle Solovieva, Ksenia Sonmer, Philip Sommer, Philip Sommer, Robin Sonnabend, Robin Sonnabend, Robin Sonrabend, Robin Sopczak, André T 10.9, T T 35.6, T 35.7, T 52.2 Sorgenfrei, Niels Sorrenson, Peter Sotrilingam, Varsiha Sotnikov, Vasily Sowa, Janis Sowa, Lars Soybelman, Nathalie Soyk, Daniel Soya, Lars Soybelman, Nathalie Soyk, Daniel Soya, Lars Soybelman, Nathalie Soyk, Daniel Spaan, Bernhard ST 8.2, ST 8.3, ST 8.4 Späh, Jan Lukas T 6 Spanier, Felix EP 7.4, EP 7 Spannagel, Simon Spantellner, Christian Spatz, Verena DD 5.1, DD DD 19.2, DD 22.3, DD 25.1 Spector, Aaron Speri, Jannis Speiser, Hendrik Spenzer, Benjamin A. Spengler, Felix Speranza, Enrico Speri, Lorenzo Speri, Lorenzo Speri, Lorenzo Speri, P4.1	48.1, T 48.2
Sole, Silvia Sole, Silvia Solis, Alvaro Lopez Solis, Michelle Solovieva, Ksenia Solovieva, Ksenia Sommer, Philip Sommerhalder, Manuel Sonnabend, Robin Sonnabend, Robin Sonntag, Dörte Sopczak, André T 10.9, T T 35.6, T 35.7, T 52.2 Sorgenfrei, Niels Sorrenson, Peter Sothilingam, Varsiha Sotnikov, Vasily Sowa, Janis Sowa, Lars Soybelman, Nathalie Soyk, Daniel Spaan, Bernhard ST 8.2, ST 8.3, ST 8.4 Späh, Jan Lukas Sotar, Ser 7.4, EP 7.4, EP 7 Spannagel, Simon Spannfellner, Christian Spatz, Verena DD 5.1, DD DD 19.2, DD 22.3, DD 25.1 Spector, Aaron Speer, Jannis Speiser, Hendrik Speranza, Enrico Speri, Lorenzo Sperlich, Dennis T 66.7, T 94.1 Sphicas, Paris	48.1, T 48.2
Sole, Silvia Sole, Silvia Solis, Alvaro Lopez Solis, Michelle Solovieva, Ksenia Solovieva, Ksenia Sommer, Philip Sommerhalder, Manuel Sonnabend, Robin Sonnabend, Robin Sonnabend, Robin Sopczak, André T 10.9, T: T 35.6, T 35.7, T 52.2 Sorgenfrei, Niels Sorrenson, Peter Sothilingam, Varsiha Sotnikov, Vasily Sowa, Janis Sowa, Lars Soybelman, Nathalie Soyk, Daniel Spaan, Bernhard ST 8.2, ST 8.3, ST 8.4 Späh, Jan Lukas Spainer, Felix Spanngel, Simon Spannfellner, Christian Spataro, David D 19.2, DD 22.3, DD 25.1 Sper, Jannis Speiser, Hendrik Sper, Jannis Speiser, Hendrik Sper, Jannis Speiser, Hendrik Sper, Lorenzo Sperlich, Dennis T 66.7, T 94.1 Sphicas, Paris	48.1, T 48.2
Sole, Silvia Sole, Silvia Solis, Alvaro Lopez Solis, Michelle Soloron, Shalu Solovieva, Ksenia Sommer, Philip Sommerhalder, Manuel Sonnabend, Robin Sonnabend, Robin Sonnabend, Robin Sontag, Dörte Sopczak, André T 10.9, T: T 35.6, T 35.7, T 52.2 Sorgenfrei, Niels Sorrenson, Peter Sothilingam, Varsiha Sothikov, Vasily Sowa, Janis Sowa, Lars Soybelman, Nathalie Soyk, Daniel Soya, Daniel Spaan, Bernhard ST 8.2, ST 8.3, ST 8.4 Späh, Jan Lukas Spataro, David Spanter, Felix Spanzel, Simon Spanter, Felix Spataro, David Sper, Jannis Speiser, Hendrik Sper, Jannis Speiser, Hendrik Sper, Lorenzo Speri, Lorenzo Speri, Lorenzo Speri, Lorenzo Sperich, Dennis T 7 Spiecker, Fleur Spiezer, Alexandra	48.1, T 48.2
Sole, Silvia Sole, Silvia Solis, Alvaro Lopez Solis, Michelle Soloron, Shalu Solovieva, Ksenia Sommer, Philip Sommerhalder, Manuel Sonnabend, Robin Sonnabend, Robin Sonnabend, Robin Sonranson, Peter Sopczak, André T 10.9, T T 35.6, T 35.7, T 52.2 Sorgenfrei, Niels Sorrenson, Peter Sortenison, Peter Sothilingam, Varsiha Sotnikov, Vasily Sowa, Janis Sowa, Lars Soybelman, Nathalie Soya, Daniel Spaan, Bernhard ST 8.2, ST 8.3, ST 8.4 Späh, Jan Lukas Spanagel, Simon Spanfellner, Christian Spatz, Verena DD 5.1, DD DD 19.2, DD 22.3, DD 25.1 Spector, Aaron Speer, Jannis Speiser, Hendrik Speranza, Enrico Speri, Lorenzo Speri, Lorenzo Sperich, Dennis T T 66.7, T 94.1 Spitzer, Alexandra Spitzer, Alexandra Spitzer, Philipp Snercklespe Elorian	48.1, T 48.2
Sole, Silvia Sole, Silvia Solis, Alvaro Lopez Solis, Michelle Soloron, Shalu Solovieva, Ksenia Sommer, Philip Sommerhalder, Manuel Sonnabend, Robin Sonnabend, Robin Sontag, Dörte Sopczak, André T 10.9, T T 35.6, T 35.7, T 52.2 Sorgenfrei, Niels Sorrenson, Peter Sothilingam, Varsiha Sotnikov, Vasily Sowa, Janis Sowa, Lars Soybelman, Nathalie Soyk, Daniel Spaan, Bernhard ST 8.2, ST 8.3, ST 8.4 Späh, Jan Lukas Stat, Verena Spannfellner, Christian Spataro, David Spaetor, Aaron Speetor, Aaron Speet, Jannis Speiser, Hendrik Sperik, Dennis Speri, Chernzo Speri, Lorenzo Sperik, Dennis Sperik, Paris Speizer, Paris Spitzer, Alexandra Spitzer, Alexandra Spitzer, Philipp Spreckelsen, Florian AKPIK 14	48.1, T 48.2
Sole, Silvia Sole, Silvia Solis, Alvaro Lopez Solis, Michelle Soloron, Shalu Solovieva, Ksenia Sommer, Philip Sommerhalder, Manuel Sonnabend, Robin Sonnabend, Robin Sonrag, Dörte Sopczak, André T 10.9, T T 35.6, T 35.7, T 52.2 Sorgenfrei, Niels Sorrenson, Peter Sotrilingam, Varsiha Sotnikov, Vasily Sowa, Janis Sowa, Lars Soybelman, Nathalie Soyk, Daniel Soya, Daniel Soya, Lars Soybelman, Nathalie Soyk, Daniel Soya, Lars Soybelman, Nathalie Soyk, Daniel Spaan, Bernhard ST 8.2, ST 8.3, ST 8.4 Späh, Jan Lukas T 6 Spanier, Felix EP 7.4, EP 7 Spannagel, Simon Spanfellner, Christian Spatz, Verena DD 5.1, DD DD 19.2, DD 22.3, DD 25.1 Spector, Aaron Speer, Jannis Speiser, Hendrik Speranza, Enrico Sperilch, Dennis Speris, Speris Speicer, Fleur Spitzer, Alexandra Spitzer, Alexandra Spitzer, Philipp Spreckelsen, Florian AKPIK 1.4	48.1, T 48.2
Sole, Silvia Sole, Silvia Solis, Alvaro Lopez Solis, Michelle Solovieva, Ksenia Sonmer, Philip Sommerhalder, Manuel Sonnabend, Robin Sonnabend, Robin Sonrag, Dörte Sopczak, André T 10.9, T T 35.6, T 35.7, T 52.2 Sorgenfrei, Niels Sorrenson, Peter Sotrilingam, Varsiha Sotnikov, Vasily Sowa, Janis Sowa, Lars Soybelman, Nathalie Soyk, Daniel Soya, Lars Soybelman, Nathalie Soyk, Daniel Soya, Lars Soybelman, Nathalie Soyk, Daniel Soya, Lars Soybelman, Nathalie Soyk, Daniel Span, Bernhard ST 8.2, ST 8.3, ST 8.4 Späh, Jan Lukas T 6 Spanier, Felix EP 7.4, EP Spannagel, Simon Spantfellner, Christian Spatz, Verena D D 19.2, DD 22.3, DD 25.1, DD DD 19.2, DD 22.3, DD 25.1, DD DD 19.2, DD 22.3, DD 25.1, Spector, Aaron Speer, Jannis Speiser, Hendrik Spencar, Benjamin A. Spengler, Felix Speranza, Enrico Speri, Lorenzo Speri, Lorenzo Sperich, Dennis Spitzer, Pleur Spitzer, Alexandra Spitzer, Alexandra Spitzer, Philipp Spreckelsen, Florian AKPIK 1.4 Staacks, Sebastian DD 35	48.1, T 48.2

Stadler, Robert T 104.3 Stafford, Dominic T 36.1, •T 76.7 Stahl, Achim GR 14.3, GR 14.4, GR 14.5, GR 14.6, ST 1.3, ST 1.5, GR 14.6, ST 1.3, ST 1.5, GR 14.6, ST 1.3, ST 1.5, GR 14.6, ST 1.5, ST 1.5
ST 3.5, 1 8.6, 1 48.1, 1 48.2, AKPIK 2.2, AKPIK 4.4 Stahlberg, Martin •T 104.1, T 104.3, T 104.6
Stallmach, Frank DD 44.1 Stamou, Emmanuel T 60.2 Stampekis, Alexios T 60.5 Stampfer, Christoph DD 35.13, DD 41.2 DD 35.13,
Standke, Mark
Stanitzki, Marcel T 94.3 Stapf, Birgit •T 62.1 Stappers, Maik T 21.4 Starauschek, Erich .DD 39.1, DD 47.2 Staszewski, Rafal T 10.9 Stauber, Josef •T 74.3 Stecker, Polina •ST 7.3 Steck, Pia ST 6.4, T 65.5, T 65.6, T 65.7, T 94.4
Stefan, Rapp EP 2.2 Stefanov, Nicole T 76.7 Stehr, Felix T 44.5, •T 44.6 Steiger, Hans T 20.8, T 48.5, T 48.7, T 102.2
Steiger, Hans Th. J. . T 48.9, •T 48.10 Steiger, Hans Theodor Josef . T 20.4 Steiger, Christian •EP 5.8 Stein, Annika . T 26.1, T 26.3 Steinberg, Dominik . T 107.5 Steinbrück, Georg . T 12.9, •T 109.2 Steindl, Stephanie . DD 10.12 Steinhoff, Jan . GR 2.7, •GR 8.2 Steinmassl, Simon . T 45.5 Steinmetz, Tilmann . DD 47.2 Sterr, Tobias . T 20.3, T 20.9, •T 48.6,
T 48.8, T 102.3 Stevens, Holger T 13.8, T 14.1, T 68.4 Stiller, Gabrielle
T 41.9, T 62.9, T 69.1 Strahl, Alexander Strahl, Alexander
Streit, Achim
stroth, Michelle ST 8.2, ST 8.3, •ST 8.4 Strzys, Marcel Sturys, Marcel T 17.8 Stuhrmann, Jochen DD 37.2 Sturdy, Jared T 39.4 Stürwald, Timo •T 93.6 Stütz, Katharina •DD 36.5 Styles, Nicholas T 16.3 Suárez Fernández, Isabel GR 13.4 Suda, Yusuke T 17.8 Sührig, Laura •DD 24.2 Sulivan, Sean T 87.8 Summ, Benjamin T 31.2 Supanitsky, Daniel T 51.7 SuperCDMS-Kollaboration T 77.1,
L //.2, L //.3 Suri, Vinit •DD 35.5 Susac, AnaDD 18.2, DD 36.1 Sutcliffe, William T 3.4, T 79.6, T 84.8 Suveyzdis, RomanT 18.8, T 46.1, •T 46.3, T 46.7

Suvorov, Arthur	. GR 11.2
Svidras, Henrikas	. •T 84.5
Swallow, Paul	T 4.4
SWGO-Kollaboration	T 45.6
Swirski, Thorben T 11.7	7, •T 67.2
Szabo-Roberts, Matyas	EP 5.7
SZCZUREK, ANTONI	1 85.8
Tada Aika	1 89.5
Takatech Chrie	.•192.5 T 3 0
Tanaka lunichi	T 10 9
Tangerine-Kollaboration	T 40 1
T 40.2. T 40.3. T 40.4. T 40.6	
Tapprogge. Stefan	6. T 69.3
Taronna, Massimo	.MP 2.4
Tasdan, Ufuk A	GPhil 6.1
Tasevsky, Marek	T 10.9
Tastekin, Mesut Ibrahim	•DD 12.8
Tateno, Gen	T 10.9
Tauqeer, Komal	.•T 87.4
Tayade, Akhilesh	•1 16.3
Taylor, Andrew	1 /3.1
Taylor, Stuart F	. •EP Z.4 T 10 6
te Vrugt Michael	2011 10.0
Teichrew Albert DD 16.3	DD 17 4
DD 24.2	55
Teissier, Jean-Mathieu EP 6.1	, •EP 6.2
Televantos Ubeda, Lucas Carlo	S
•DD 11.6	
Tenbruck, Anna-Sophia	•T 106.8
Tepe, Enes	. •GP 8.2
Terashi, Koji	1 10.9
Ternerde Marius	4, 1 50.5
Tercerue, Marius	i 33.2 / 22 חח
Tessalitz, Amelle	T Q 2
Tews Ingo	GR 9.6
Thamm Nick •T 39 5 T 39	6 T 39 7
Thampi. Ashish	T 63.2
Thara Hariharan, Vidhya	. •T 20.7
Thébault, Karim•A	GPhil 6.1
Thees, Michael	. DD 12.3
Theisen, Eric	T 48.5
Therolf, Sascha	. DD 46.1
Theyßen, Heike DD 5.3,	DD 24.1,
DD 33.4, DD 39.1	
Theyssen, Helke	DD 50.3
Thill Datrick	0, •1 00.4
Thirolf Peter G	ST 2 1
Thomas, Florian	T 49.4
Thorne. Larisa	•T 49.1
Threlfall, James	. EP 12.5
Thürigen, Johannes	•MP 7.4
Tian, Yusong	•T 12.1
Tichy, Wolfgang	. GR 13.2
Tietzsch, Alexander T 20.3	3, T 20.9,
T 48.6, T 48.8, T 102.3	
Tim, Kunibusch	•GR 14.5
Timmormonn Booto	51 Z.3 6 9T 6 2
Tischer Jonas	0,310.2
Tkachenko Olena	•T 19 4
Tluczykont, Martin	7. T 77.4
Tobias, Jürgen	
Talaansa Viataala A	
Tokareva, victoria•A	KPIK 1.5
Tokareva, victoria•A Tolba, Tamer	KPIK 1.5 . •T 59.6
Tolkareva, victoriaA Tolba, Tamer Tomrlin, Bruno	KPIK 1.5 . •T 59.6 . DD 36.1
Tokareva, victoriaA Tolba, Tamer Tomrlin, Bruno Tönnies, Florian	KPIK 1.5 . •T 59.6 . DD 36.1 . •T 50.6
Tokareva, victoria •A Tolba, Tamer Tomrlin, Bruno Tönnies, Florian Torndal, Julie Török Tibar	KPIK 1.5 • T 59.6 • DD 36.1 • T 50.6 • T 90.8 EP 12.11
Tokareva, Victoria A Tolba, Tamer Tomrlin, Bruno Tönnies, Florian Torndal, Julie Török, Tibor Torroe Hoberth	KPIK 1.5 •T 59.6 DD 36.1 •T 50.6 •T 90.8 EP 12.11 •T 28 2
Tokareva, victoria	KPIK 1.5 . •T 59.6 . DD 36.1 . •T 50.6 . •T 90.8 EP 12.11 . •T 28.3 3 T 60 4
Tokareva, Victoria A Tolba, Tamer Tomrlin, Bruno Tönnies, Florian Torndal, Julie Török, Tibor Torres, Heberth Tost, Marc T 60. Trabs Mathias A	KPIK 1.5 . •T 59.6 . DD 36.1 . •T 50.6 . •T 90.8 EP 12.11 . •T 28.3 3, T 60.4 KPIK 4 2
Tokareva, victoria •A Tolba, Tamer •A Tomrlin, Bruno • Tönnies, Florian • Torndal, Julie • Török, Tibor • Torres, Heberth • Tost, Marc • Trabs, Mathias • Trefzger, Thomas DD 8.2	KPIK 1.5 . •T 59.6 . DD 36.1 . •T 50.6 . •T 90.8 EP 12.11 . •T 28.3 3, T 60.4 KPIK 4.2 , DD 11.3.
Tokareva, Victoria A Tolba, Tamer Tomrlin, Bruno Törnies, Florian Torrdal, Julie Torres, Heberth Torst, Marc T 60. Trabs, Mathias A Trefzger, Thomas DD 8.2, DD 17.2, DD 17.3, DD 18.3, DD	KPIK 1.5 . •T 59.6 . DD 36.1 . •T 50.6 . •T 90.8 EP 12.11 . •T 28.3 3, T 60.4 KPIK 4.2 , DD 11.3, 9 35.11,
Tokareva, Victoria	KPIK 1.5 . •T 59.6 . DD 36.1 . •T 50.6 . •T 90.8 EP 12.11 . •T 28.3 3, T 60.4 KPIK 4.2 ,DD 11.3, 9 35.11,
Tokareva, Victoria	KPIK 1.5 • T 59.6 DD 36.1 • T 50.6 • T 90.8 EP 12.11 • T 28.3 3, T 60.4 KPIK 4.2 ,DD 11.3, 9 35.11, . DD 17.4
Tokareva, Victoria	KPIK 1.5 • T 59.6 DD 36.1 • T 50.6 • T 90.8 EP 12.11 • T 28.3 3, T 60.4 KPIK 4.2 ,DD 11.3, 9 35.11, . DD 17.4
Tokareva, Victoria	KPIK 1.5 • T 59.6 DD 36.1 • T 50.6 • T 90.8 EP 12.11 • T 28.3 3, T 60.4 KPIK 4.2 DD 11.3, 9 35.11, . DD 17.4
Tokareva, Victoria	KPIK 1.5 • T 59.6 DD 36.1 • T 50.6 • T 90.8 EP 12.11 • T 28.3 3, T 60.4 KPIK 4.2 DD 11.3, 35.11, . DD 17.4 • T 45.1
Tokareva, Victoria A Tolba, Tamer Tomrlin, Bruno Törnies, Florian Török, Tibor Torres, Heberth Tost, Marc T 60. Trabs, Mathias A Trefzger, Thomas DD 8.2, DD 17.2, DD 17.3, DD 18.3, DC DD 48.1, DD 48.2 Tremmel, Michael TRIPLE-nanoAUV-Kollaboratior EP 4.3 Tsaklidis, Ilias T 3. Tsirou, Michelle	KPIK 1.5 • T 59.6 DD 36.1 • T 50.6 • T 90.8 EP 12.11 • T 28.3 3, T 60.4 KPIK 4.2 DD 11.3, 9 35.11, . DD 17.4 4, T 79.6 • • T 45.1 ST 2 4
Tokareva, Victoria	KPIK 1.5 • T 59.6 D 36.1 • T 50.6 • T 90.8 EP 12.11 • T 28.3 3, T 60.4 KPIK 4.2 , DD 11.3, 0 35.11, . DD 17.4 4, T 79.6 • T 45.1 ST 3.4 T 4.4
Tokareva, Victoria	KPIK 1.5 . •T 59.6 . DD 36.1 . •T 50.6 . •T 90.8 EP 12.11 . •T 28.3 3, T 60.4 KPIK 4.2 . DD 11.3, 0 35.11, . DD 17.4 4, T 79.6 . •T 45.1 ST 3.4 T 4.4 •T 2.8
Tokareva, Victoria	KPIK 1.5 • T 59.6 DD 36.1 • T 50.6 • T 90.8 EP 12.11 • T 28.3 3, T 60.4 KPIK 4.2 DD 11.3, 9 35.11, . DD 17.4 4, T 79.6 •T 45.1 ST 3.4 T 4.4 • T 72.8 • T 72.8 • T 72.8 • T 77.8
Tokareva, Victoria •A Tolba, Tamer Tomrlin, Bruno Tönnies, Florian Torrdal, Julie Török, Tibor Torres, Heberth Tost, Marc T 60. Trabs, Mathias A Trefzger, Thomas DD 8.2, DD 17.2, DD 17.3, DD 18.3, DD DD 48.1, DD 48.2 Tremmel, Michael TRIPLE-nanoAUV-Kollaboratior Tsirou, Michelle Staklidis, Ilias Tsirou, Michelle Tosi, Ah Chung Tuning, Niels Turkovic, Timur Turkovic, Cimur Tivsuy, Cenk	KPIK 1.5 • T 59.6 DD 36.1 • T 50.6 • T 90.8 EP 12.11 • T 28.3 3, T 60.4 KPIK 4.2 DD 11.3, 0 35.11, . DD 17.4 4, T 79.6 . • T 45.1 ST 3.4 T 4.4 T 72.8 T 4.5 T 4.5 T 9.5 T 9.
Tokareva, Victoria	KPIK 1.5 • T 59.6 DD 36.1 • T 50.6 • T 90.8 EP 12.11 • T 28.3 3, T 60.4 KPIK 4.2 DD 11.3, 0 35.11, DD 17.4 4, T 79.6 • T 72.8 • T 4.4 • T 72.8 • T 4.4 • T 72.8 • T 67.7 8, T 79.9 2, T 34.3,
Tokareva, Victoria	KPIK 1.5 • T 59.6 · DD 36.1 • T 50.6 • T 90.8 EP 12.11 • T 28.3 3, T 60.4 KPIK 4.2 ,DD 11.3, 35.11, . DD 17.4 · T 45.1 T 45.1 T 4.4 • T 72.8 • T 67.7 8, T 79.9 2, T 34.3,
Torrareva, Victoria	KPIK 1.5 . •T 59.6 DD 36.1 . •T 90.8 EP 12.11 . •T 28.3 3, T 60.4 KPIK 4.2 ,DD 11.3, 0 35.11, . DD 17.4 4, T 79.6 . •T 45.1 T 4.4 T 4.4 T 4.4 T 4.4 T 79.8 T 4.4 T 4.4 T 79.9 2, T 34.3, DD 32.1,
Tokareva, Victoria •A Toka, Tamer •Toka, Tamer Tomrlin, Bruno •Tonnies, Florian Torndal, Julie •Torradal, Julie Torros, Heberth •Torros, Marc Tost, Marc •Tof0, Trabs, Mathias Tost, Marc •D Do 17.2, DD 17.3, DD DD 48.1, DD 48.2 Tremmel, Michael •T TRIPLE-nanoAUV-Kollaboratior EP EP 4.3 Tsaklidis, Ilias •T Tsirou, Michelle •T Tourotte, Roxanne •T Turkovic, Timur T Tuscute, Roxanne •T Turkovic, Timur T Tuscute, Assance •T Tuscute, Rozana, Eftychia T T 3.5.2, •T Ubben, Malte	KPIK 1.5 . •T 59.6 DD 36.1 . •T 50.6 . •T 90.8 EP 12.11 . •T 28.3 3, T 60.4 KPIK 4.2 DD 11.3, 35.11, . DD 17.4 . •T 45.1 ST 3.4 •T 45.1 ST 3.4 •T 45.1 ST 3.4 •T 45.1 ST 3.4 •T 45.1 T 79.9 2, T 34.3, DD 32.1, DD 32.1,
Tokareva, Victoria •A Toka, Tamer •Toka, Tamer Tomrlin, Bruno •Tornal, Julie Torndal, Julie •Torras, Florian Torrad, Julie •Torras, Heberth Tost, Marc •Tofo, Trabs, Mathias A Trefzger, Thomas DD 8.2, DD 17.2, DD 17.3, DD 18.3, DD DD 48.1, DD 48.2 Tremmel, Michael	KPIK 1.5 • T 59.6 DD 36.1 • T 50.6 • T 90.8 EP 12.11 • T 28.3 3, T 60.4 KPIK 4.2 DD 11.3, 9 35.11, . DD 17.4 4, T 79.6 • T 45.1 ST 3.4 T 4.4 • T 72.8 • T 4.5 • T 72.8 • T 4.3, DD 32.1, DD 11.9 • T 4.5 • T 9.6 • • • • • • • • • • • • • • • • • •
Tokareva, Victoria	KPIK 1.5 • T 59.6 DD 36.1 • T 50.6 • T 90.8 EP 12.11 • T 28.3 3, T 60.4 KPIK 4.2 DD 11.3, 0 35.11, . DD 17.4 4, T 79.6 T 45.1 ST 3.4 T 4.4 • T 72.8 0, T 72.8 T 4.5 T 4.5

Ullrich, Mark DD 17.4, DD 24.2
Ulrich, Ralf T 47.8, T 72.5, T 72.6
Ulrichs, Johannes•T 77.5
ULTRASAT-Kollaboration T 13.7
Unger, Daniel
Unger, Kai Lukas I 43.8
Unger, Michael 19.6
Unverzant Daniel
Unzicker Alexander •MP 11 4
•GP 10.1
Urban, Korbinian•T 21.8
Urban, Petr
Usoskin, Ilya EP 5.5, EP 5.6
Utehs, Julian•T 41.6
Uwer, Ulrich
Uzeiroska, Rukije•T 51.2
Vagts, Alexander
Vanidinia, Monammad Hasan GR I.5
Valderania Chryspotemas T 11 1
T 11 2 T 11 2 T 11 4 T 11 5 T 11 6
T 14 5 T 97 7
Valeriani-Kaminski Barbara T 52 1
Vallecorsa. Sofia
van de meent, maarten GR 13.3
van de Ven, Antoine A.J • MP 11.5
van de Ven, Christiaan •MP 7.6
van der Graaf, Aaron•T 107.8
van der Linden, Jan •T 88.6, T 88.7,
Т 88.8, Т 88.9
van der Spek, Trienke •DD 15.1,
•GP 1.1
Van Dyk, Danny I 56.8
Van Vliet, Arjen EP /.2, 1 /3.1
Vara Carbonell Francisco Javiar
vardanyan tatevik •GR 3.6
Vassallo. Antonio
Vauterin, Antoine T 24.4, •T 52.2
Vauth, Annika
Veatch, Jason
Veberic, Darko . EP 7.3, T 19.3, T 47.3,
T 51.4, T 73.4, AKPIK 4.8, AKPIK 4.9
Veghel, Maarten
Veghel, Maarten
Veghel, Maarten
Veghel, Maarten T 4.4 Vellasco, Murillo • T 63.7 Vellidis, Costas T 9.5 Vent, Sophia T 2.8 Ventura Barroso Ana • T 65.7
Veghel, Maarten T 4.4 Vellasco, Murillo • T 63.7 Vellidis, Costas T 9.5 Vent, Sophia T 2.8 Ventura Barroso, Ana • T 65.8 Verheek Nico ST 16
Veghel, Maarten T 4.4 Vellasco, Murillo • T 63.7 Vellidis, Costas T 9.5 Vent, Sophia T 2.8 Ventura Barroso, Ana • T 65.8 Verbeek, Nico ST 1.6 Verbytski, Andrii T 58.3
Veghel, Maarten
Veghel, Maarten
Veghel, Maarten
Veghel, Maarten T 4.4 Vellasco, Murillo •T 63.7 Vellidis, Costas T 9.5 Vent, Sophia T 2.8 Verbuek, Nico ST 1.6 Verbeek, Nico ST 1.6 Verbytski, Andrii T 58.3 Verch, Rainer GR 3.2 Vernizzi, Filippo GR 5.3 Verstege, Cedric •T 32.5 Verstraete, Frank •MP 3.1
Veghel, Maarten T 4.4 Vellasco, Murillo •T 63.7 Vellidis, Costas T 9.5 Vent, Sophia T 2.8 Verbuek, Nico ST 1.6 Verbytski, Andrii T 58.3 Verbytski, Andrii T 58.3 Verch, Rainer GR 3.2 Verstrate, Filippo GR 5.3 Verstrate, Frank •MP 3.1 Vetter, Sebastian •T 50.1
Veghel, Maarten T 4.4 Vellasco, Murillo • T 63.7 Vellidis, Costas T 9.5 Vent, Sophia T 2.8 Vertura Barroso, Ana • T 65.8 Verbeek, Nico ST 1.6 Verbytski, Andrii T 58.3 Verch, Rainer GR 3.2 Vernizzi, Filippo GR 5.3 Verstege, Cedric • T 32.5 Verstraete, Frank • MP 3.1 Vetter, Sebastian • T 50.1 Viau, João T 92.2
Veghel, Maarten
Veghel, Maarten T 4.4 Vellasco, Murillo •T 63.7 Vellidis, Costas T 9.5 Vent, Sophia T 2.8 Ventura Barroso, Ana •T 65.8 Verbeek, Nico ST 1.6 Verbytski, Andrii T 58.3 Verch, Rainer GR 3.3 Verstege, Cedric •T 32.5 Verstraete, Frank •MP 3.1 Vetter, Sebastian •T 50.1 Viana, João T 92.2 Vieu, Thibault •T 73.8 Vigranda Gianpiaro •T 64.1
Veghel, Maarten T 4.4 Vellasco, Murillo •T 63.7 Vellidis, Costas T 9.5 Vent, Sophia T 2.8 Ventura Barroso, Ana •T 65.8 Verbeek, Nico ST 1.6 Verbytski, Andrii T 58.3 Verch, Rainer GR 3.2 Vernizzi, Filippo GR 5.3 Verstege, Cedric •T 32.5 Verstraete, Frank •MP 3.1 Vetter, Sebastian •T 50.1 Viana, João T 92.2 Vieu, Thibault •T 73.8 Vigilante, Lorenzo •T 64.1 Vignola, Gianpiero •T 40.2 Viana ST 9.3
Veghel, Maarten T 4.4 Vellasco, Murillo •T 63.7 Vellidis, Costas T 9.5 Vent, Sophia T 2.8 Ventura Barroso, Ana •T 65.8 Verbeek, Nico ST 1.6 Verbytski, Andrii T 58.3 Verchek, Nico ST 1.6 Verbytski, Andrii T 58.3 Verch, Rainer GR 5.3 Verstege, Cedric •T 32.5 Verstraete, Frank •MP 3.1 Vetter, Sebastian •T 50.1 Viana, João T 92.2 Vieu, Thibault •T 73.8 Vigilante, Lorenzo •T 64.1 Vignola, Gianpiero •T 40.2 Vila, Ivan ST 9.3 Villa, Emilio T 2.7
Veghel, Maarten T 4.4 Vellasco, Murillo •T 63.7 Vellidis, Costas T 9.5 Vent, Sophia T 2.8 Ventura Barroso, Ana •T 65.8 Verbeek, Nico ST 1.6 Verbytski, Andrii T 58.3 Verch, Rainer GR 3.2 Vernizzi, Filippo GR 5.3 Verstege, Cedric •T 32.5 Verstraete, Frank •MP 3.1 Vieter, Sebastian •T 50.1 Viana, João T 92.2 Vieu, Thibault •T 73.8 Vigilante, Lorenzo •T 64.1 Vianol, Gianpiero •T 40.2 Vial, Ivan ST 9.3 Vila, Ivan T 2.7 Villaba Petro, Carmen Victoria *T 2.7
Veghel, Maarten T 4.4 Vellasco, Murillo •T 63.7 Vellidis, Costas T 9.5 Vent, Sophia T 2.8 Verbeek, Nico ST 1.6 Verbeek, Nico ST 1.6 Verbytski, Andrii T 58.3 Verch, Rainer GR 3.2 Vernizzi, Filippo GR 5.3 Verstege, Cedric •T 32.5 Verstraete, Frank •MP 3.1 Vetter, Sebastian •T 50.1 Viana, João T 92.2 Vieu, Thibault •T 73.8 Vigilante, Lorenzo •T 64.1 Vignola, Gianpiero •T 40.2 Vila, Ivan ST 9.3 Villa, Emilio T 2.7 Villalba Petro, Carmen Victoria •T 66.1
Veghel, Maarten
Veghel, Maarten T 4.4 Vellasco, Murillo T 63.7 Vellidis, Costas T 9.5 Vent, Sophia T 2.8 Ventura Barroso, Ana T 65.8 Verbeek, Nico ST 1.6 Verbytski, Andrii T 58.3 Verch, Rainer GR 3.3 Verstege, Cedric T 32.5 Verstraete, Frank MP 3.1 Vetter, Sebastian •T 50.1 Viana, João T 92.2 Vieu, Thibault •T 73.8 Vignola, Gianpiero •T 64.1 Vignola, Gianpiero •T 40.2 Villa, Emilio T 2.7 Villaba Petro, Carmen Victoria •T 66.1 •T 66.1 Vines, Justin GR 5.4 Virto, Javier T 56.1
Veghel, Maarten T 4.4 Vellasco, Murillo •T 63.7 Vellidis, Costas T 9.5 Vent, Sophia T 2.8 Ventura Barroso, Ana •T 65.8 Verbeek, Nico ST 1.6 Verbytski, Andrii T 58.3 Verch, Rainer GR 3.2 Vernizzi, Filippo GR 5.3 Verstege, Cedric •T 32.5 Vestraete, Frank •MP 3.1 Vetter, Sebastian •T 50.1 Viana, João T 92.2 Vieu, Thibault •T 73.8 Vignola, Gianpiero •T 64.1 Vignola, Gianpiero •T 40.2 Vila, Ivan ST 9.3 Villa, Emilio T 2.7 Villaba Petro, Carmen Victoria •T 66.1 Vines, Justin GR 5.4 Virto, Javier T 56.4 Visser, Matt GR 2.3
Veghel, Maarten T 4.4 Vellasco, Murillo T 63.7 Vellidis, Costas T 9.5 Vent, Sophia T 2.8 Ventura Barroso, Ana T 65.8 Verbeek, Nico ST 1.6 Verbytski, Andrii T 58.3 Verch, Rainer GR 5.3 Verstege, Cedric T 32.5 Verstege, Cedric T 50.1 Viana, João T 92.2 Vieu, Thibault T 73.8 Vigilante, Lorenzo T 64.1 Vingola, Gianpiero T 40.2 Villa, Emilio T 2.7 Villaba Petro, Carmen Victoria T 56.1 Vireser, Matti GR 2.3 Virey, Matthieu T 78.6
Veghel, Maarten T 4.4 Vellasco, Murillo •T 63.7 Vellidis, Costas T 9.5 Vent, Sophia T 2.8 Ventura Barroso, Ana •T 65.8 Verbeek, Nico ST 1.6 Verbytski, Andrii T 58.3 Verch, Rainer GR 3.2 Vernizzi, Filippo GR 5.3 Verstege, Cedric •T 32.5 Verstraete, Frank •MP 3.1 Vetter, Sebastian •T 50.1 Viana, João T 92.2 Vieu, Thibault •T 73.8 Vigilante, Lorenzo •T 64.1 Vignola, Gianpiero •T 40.2 Vila, Ivan ST 9.3 Villa, Emilio T 2.7 Villaba Petro, Carmen Victoria •T 66.1 •T 66.1 GR 5.4 Virto, Javier T 56.1 Viseer, Matt GR 2.3 Vivier, Matthieu T 78.6 Voevodina, Elena T 14.5, •T 67.3,
Veghel, Maarten T 4.4 Vellasco, Murillo • T 63.7 Vellidis, Costas T 9.5 Vent, Sophia T 2.8 Ventura Barroso, Ana • T 65.8 Verbeek, Nico ST 1.6 Verbytski, Andrii T 58.3 Verch, Rainer GR 3.2 Vernizzi, Filippo GR 5.3 Verstege, Cedric • T 32.5 Verstraete, Frank • MP 3.1 Vetter, Sebastian • T 50.1 Viana, João T 92.2 Vieu, Thibault • T 73.8 Vigilante, Lorenzo • T 64.1 Vignola, Gianpiero • T 40.2 Vila, Ivan ST 9.3 Villa, Emilio T 2.7 Villaba Petro, Carmen Victoria • T 66.1 • T 66.1 GR 5.4 Virto, Javier T 56.1 Vires, Justin GR 5.4 Virto, Javier T 56.1 Viser, Matthieu T 78.6 Voevodina, Elena T 14.5, • T 67.3, T 67.5, T 67.6 Voreal Fabian
Veghel, Maarten
Veghel, Maarten T 4.4 Vellasco, Murillo T 63.7 Vellidis, Costas T 9.5 Vent, Sophia T 2.8 Ventura Barroso, Ana T 65.8 Verbeek, Nico ST 1.6 Verbytski, Andrii T 58.3 Verch, Rainer GR 5.3 Verstege, Cedric T 32.5 Verstraete, Frank MP 3.1 Vetter, Sebastian •T 50.1 Viana, João T 92.2 Vieu, Thibault •T 73.8 Vignola, Gianpiero •T 64.1 Vignola, Gianpiero •T 40.2 Villa, Van ST 9.3 Villa, Emilio T 2.7 Villaba Petro, Carmen Victoria •T 56.1 Viros, Justin GR 5.4 Virto, Javier T 56.1 Viser, Matthieu T 78.6 Voevodina, Elena .T 14.5, •T 67.3, T 67.5, T 67.6 Vogel, Fabian .T 11.1, T 11.2, T 11.3, T 11.4, T 11.5, •T 11.6, T 97.7 Vogel, Lorenz •T 93.1
Veghel, Maarten T 4.4 Vellasco, Murillo T 63.7 Vellidis, Costas T 9.5 Vent, Sophia T 2.8 Ventura Barroso, Ana T 65.8 Verbeek, Nico ST 1.6 Verbytski, Andrii T 58.3 Verbeek, Nico ST 1.6 Verbytski, Andrii T 58.3 Verch, Rainer GR 5.3 Verstege, Cedric T 32.5 Verstraete, Frank MP 3.1 Vetter, Sebastian •T 50.1 Viana, João T 92.2 Vieu, Thibault •T 73.8 Vignola, Gianpiero •T 44.2 Vila, Ivan ST 9.3 Villa, Emilio T 2.7 Villa, Emilio T
Veghel, Maarten T 4.4 Vellasco, Murillo T 63.7 Vellidis, Costas T 9.5 Vent, Sophia T 2.8 Ventura Barroso, Ana T 65.8 Verbeek, Nico ST 1.6 Verbytski, Andrii T 58.3 Verchek, Nico ST 1.6 Verbytski, Andrii T 58.3 Verch, Rainer GR 5.3 Verstege, Cedric T 32.5 Verstraete, Frank •MP 3.1 Vetter, Sebastian •T 50.1 Viana, João T 92.2 Vieu, Thibault •T 73.8 Vignola, Gianpiero •T 64.1 Vignola, Gianpiero •T 40.2 Vila, Ivan ST 9.3 Villa, Emilio T 2.7 Villaba Petro, Carmen Victoria •T 66.1 Vines, Justin GR 5.4 Virto, Javier T 56.1 Visser, Matti GR 2.3 Vivier, Matthieu T 78.6 Voevodina, Elena T 14.5, •T 67.3, T 67.5, T 67.6 Vogel, Lorenz •T 93.1 Vogel, Lorenz •T 93.1 Vogel, Lorenz
Veghel, Maarten T 4.4 Vellasco, Murillo T 63.7 Vellidis, Costas T 9.5 Vent, Sophia T 2.8 Ventura Barroso, Ana T 65.8 Verbeek, Nico ST 1.6 Verbytski, Andrii T 58.3 Verchek, Nico ST 1.6 Verbytski, Andrii T 58.3 Verchzi, Filippo GR 5.3 Verstege, Cedric T 32.5 Verstraete, Frank •MP 3.1 Vetter, Sebastian •T 50.1 Viana, João T 92.2 Vieu, Thibault •T 73.8 Vignola, Gianpiero •T 40.2 Vila, Ivan ST 9.3 Villa, Emilio T 2.7 Villalba Petro, Carmen Victoria •T 66.1 Viros, Javier T 56.1 Visser, Matt GR 2.3 Vivire, Matthieu T 78.6 Voevodina, Elena T 14.5, •T 67.3, Y 10.4, T 11.5, •T 11.6, T 97.7 Vogel, Lorenz Vogel, Naomi •T 51.8 Vogl, Christoph •T 78.4 Vogl, Stefan •T 54.4
Veghel, Maarten T 4.4 Vellasco, Murillo T 63.7 Vellidis, Costas T 9.5 Vent, Sophia T 2.8 Ventura Barroso, Ana T 65.8 Verbeek, Nico ST 1.6 Verbytski, Andrii T 58.3 Verch, Rainer GR 3.2 Vernizzi, Filippo GR 5.3 Verstege, Cedric T 32.5 Verstege, Cedric T 32.5 Verstege, Cedric T 32.5 Verstege, Cedric T 73.8 Vietter, Sebastian T 50.1 Viana, João T 92.2 Vieu, Thibault T 73.8 Vigilante, Lorenzo T 64.1 Vignola, Gianpiero T 40.2 Vila, Ivan ST 9.3 Villa, Emilio T 2.7 Villalba Petro, Carmen Victoria •T 66.1 Visser, Matt GR 2.3 Vivier, Matthieu T 78.6 Voecodina, Elena T 14.5, •T 67.3, T 67.5 Vogel, Fabian T 11.1, T 11.2, T 11.3, T 11.4, T 11.5, •T 11.6, T 97.7 Vogel, Lorenz •T 93.1 Vogel, Naomi •T 51.8
Veghel, Maarten T 4.4 Vellasco, Murillo • T 63.7 Vellidis, Costas T 9.5 Vent, Sophia T 2.8 Ventura Barroso, Ana • T 65.8 Verbeek, Nico ST 1.6 Verbytski, Andrii T 58.3 Verch, Rainer GR 3.2 Vernizzi, Filippo GR 5.3 Verstege, Cedric • T 32.5 Verstraete, Frank • MP 3.1 Vetter, Sebastian • T 50.1 Viana, João T 92.2 Vieu, Thibault • T 73.8 Vigilante, Lorenzo • T 64.1 Vignola, Gianpiero • T 40.2 Vila, Ivan ST 9.3 Villa, Emilio T 2.7 Viser, Matthi
Veghel, Maarten
Veghel, Maarten T 4.4 Vellasco, Murillo T 63.7 Vellidis, Costas T 9.5 Vent, Sophia T 2.8 Ventura Barroso, Ana T 65.8 Verbeek, Nico ST 1.6 Verbytski, Andrii T 58.3 Verbeek, Nico ST 1.6 Verbytski, Andrii T 58.3 Verch, Rainer GR 5.3 Verstege, Cedric T 32.5 Verstraete, Frank MP 3.1 Vetter, Sebastian •T 50.1 Viana, João T 92.2 Vieu, Thibault •T 73.8 Vigialante, Lorenzo •T 64.1 Vignola, Gianpiero •T 40.2 Vila, Ivan ST 9.3 Villa, Emilio T 2.7 Villalba Petro, Carmen Victoria •T 66.1 Vines, Justin GR 5.4 Virto, Javier T 56.1 Viser, Matthieu T 78.6 Voevodina, Elena T 14.5, •T 67.3, T 67.5, T 67.6 Vogel, Fabian T 51.8 Vogl, Christoph •T 78.4 Vogl, Stefan •T 54.4 Vogt, Marco
Veghel, Maarten T 4.4 Vellasco, Murillo T 63.7 Vellidis, Costas T 9.5 Vent, Sophia T 2.8 Ventura Barroso, Ana T 65.8 Verbeek, Nico ST 1.6 Verbytski, Andrii T 58.3 Verbeek, Nico ST 1.6 Verbytski, Andrii T 58.3 Verch, Rainer GR 5.3 Verstege, Cedric T 32.5 Verstraete, Frank MP 3.1 Vetter, Sebastian •T 50.1 Viana, João T 92.2 Vieu, Thibault •T 73.8 Viginola, Gianpiero •T 40.2 Vila, Ivan ST 9.3 Villa, Emilio T 2.7 Villa, Emilio T
Veghel, Maarten T 4.4 Vellasco, Murillo T 63.7 Vellidis, Costas T 9.5 Vent, Sophia T 2.8 Ventura Barroso, Ana T 65.8 Verbeek, Nico ST 1.6 Verbytski, Andrii T 58.3 Verch, Rainer GR 3.2 Vernizzi, Filippo GR 5.3 Verstege, Cedric T 32.5 Verstraete, Frank •MP 3.1 Vetter, Sebastian •T 50.1 Viana, João T 92.2 Vieu, Thibault •T 73.8 Vigialante, Lorenzo •T 64.1 Vignola, Gianpiero •T 40.2 Vila, Ivan ST 9.3 Villa, Emilio T 2.7 Villaba Petro, Carmen Victoria •T 66.1 Viros, Javier T 56.4 Virito, Javier T 56.4 Viros, Javier T 56.5 Vogel, Fabian T 11.5, •T 11.6, T 37.4 Vogel, Lorenz •T 93.1 Vogel, Lorenz •T 91.3 Vogel, Johann Christoph •T 78.4 Vogl, Marco T 15.6, T 36.4 Voigtidinder, Tim </td
Veghel, Maarten T 4.4 Vellasco, Murillo • T 63.7 Vellidis, Costas T 9.5 Vent, Sophia T 2.8 Ventura Barroso, Ana • T 65.8 Verbeek, Nico ST 1.6 Verbytski, Andrii T 58.3 Verch, Rainer GR 3.2 Vernizzi, Filippo GR 5.3 Verstege, Cedric • T 32.5 Verstraete, Frank • MP 3.1 Vetter, Sebastian • T 50.1 Viana, João T 92.2 Vieu, Thibault • T 73.8 Vigilante, Lorenzo • T 64.1 Vigola, Gianpiero • T 40.2 Vila, Ivan ST 9.3 Villa, Emilio T 2.7 Villaba Petro, Carmen Victoria • T 66.1 Vines, Justin GR 5.4 Virto, Javier T 56.1 Viser, Matthieu T 78.6 Voevodina, Elena T 14.5, • T 67.3, T 67.5, T 67.6 Vogel, Lorenz • T 93.1 Vogel, Naomi • T 51.8 Vogl, Christoph • T 78.4 Vogl, Christoph • T 78.4 <t< td=""></t<>
Veghel, Maarten
Veghel, Maarten
Veghel, Maarten T 4.4 Vellasco, Murillo T 63.7 Vellidis, Costas T 9.5 Vent, Sophia T 2.8 Ventura Barroso, Ana T 65.8 Verbeek, Nico ST 1.6 Verbytski, Andrii T 58.3 Verch, Rainer GR 5.3 Verstege, Cedric T 32.5 Verstraete, Frank MP 3.1 Vetter, Sebastian •T 50.1 Viana, João T 92.2 Vieu, Thibault •T 73.8 Vigola, Gianpiero •T 64.1 Vignola, Gianpiero •T 40.2 Villa, Van ST 9.3 Villa, Emilio T 2.7 Villaba Petro, Carmen Victoria •T 66.1 Viros, Justin GR 5.4 Virto, Javier T 56.1 Viseser, Matt GR 2.3 Vivier, Matthieu T 78.6 Voevodina, Elena .T 14.5, •T 67.3, T 67.5, T 67.6 Vogel, Lorenz •T 93.1 Vogel, Naomi .T 51.8 Vogl, Christoph .T 78.4 Vogl, Stefan .T 54.4 Vogt, Marco
Veghel, Maarten T 4.4 Vellasco, Murillo T 63.7 Vellidis, Costas T 9.5 Vent, Sophia T 2.8 Ventura Barroso, Ana T 65.8 Verbeek, Nico ST 1.6 Verbytski, Andrii T 58.3 Verbeek, Nico ST 1.6 Verbytski, Andrii T 58.3 Verch, Rainer GR 5.3 Verstege, Cedric T 32.5 Verstraete, Frank MP 3.1 Vetter, Sebastian •T 50.1 Viana, João T 92.2 Vieu, Thibault •T 73.8 Viginola, Gianpiero •T 40.2 Vila, Ivan ST 9.3 Vila, Emilio T 2.7 Villa, Emilio T
Veghel, Maarten T 4.4 Vellasco, Murillo T 63.7 Vellidis, Costas T 9.5 Vent, Sophia T 2.8 Ventura Barroso, Ana T 65.8 Verbeek, Nico ST 1.6 Verbytski, Andrii T 58.3 Verchek, Nico ST 1.6 Verbytski, Andrii T 58.3 Verch, Rainer GR 5.3 Verstege, Cedric T 32.5 Verstraete, Frank MP 3.1 Vetter, Sebastian •T 50.1 Viana, João T 92.2 Vieu, Thibault •T 73.8 Viginola, Gianpiero •T 64.1 Vignola, Gianpiero •T 40.2 Vila, Ivan ST 9.3 Villa, Emilio T 2.7 Villalba Petro, Carmen Victoria •T 66.1 Viros, Javier T 56.1 Visser, Matthieu T 78.6 Voevodina, Elena T 14.5, •T 67.3 T 67.5, T 67.6 Vogel, Lorenz •T 93.1 Vogel, Naomi •T 51.8 Vogl, Christoph •T 78.4 Vogl, Stefan •T 15.6, T 36.4 Voitz-Wi
Veghel, Maarten T 4.4 Vellasco, Murillo • T 63.7 Vellidis, Costas T 9.5 Vent, Sophia T 2.8 Ventura Barroso, Ana • T 65.8 Verbeek, Nico ST 1.6 Verbytski, Andrii T 58.3 Verch, Rainer GR 3.2 Vernizzi, Filippo GR 5.3 Verstege, Cedric • T 32.5 Verstraete, Frank • MP 3.1 Vetter, Sebastian • T 50.1 Viana, João T 92.2 Vieu, Thibault • T 73.8 Vigilante, Lorenzo • T 64.1 Vignola, Gianpiero • T 40.2 Vila, Ivan ST 9.3 Villa, Bernilio T 2.7 Villalba Petro, Carmen Victoria • T 66.1 Vines, Justin GR 5.4 Virto, Javier T 56.1 Viser, Matthieu T 78.6 Voevodina, Elena T 14.5, • T 67.3, Vila, Varoo T 78.4 Vogel, Lorenz • T 93.1 Vogel, Naomi • T 51.8 Vogl, Christoph • T 78.4 Vogl, Christoph <td< td=""></td<>

von Cube, R. Florian
von Cube, Ralf Florian•T 15.4
von den Driesch, Jost . T 26.7, •T 76.8
von Jasienicki, Larissa•T 95.7
von Krosigk, Belina
von Oy, Johanna•T 38.5
Vorholzer, Andreas
Vormwald Benedikt T 12 2
Voß Katharina •T 26.5
Voss Tobias DD 6.2
Vovk levgen T 17.8
Wash Repedikt
Wackermann Deiner DD 91 DD 42.2
Wackernialli, Railer DD 6.1, DD 45.5
wagner, victoria •1 /4.6, 1 /4./,
1 /8.6, 1 /8./, 1 104.5, 1 104./
Wagner, Wolfgang 1 6.2, •1 6.9,
Т 14.7, Т 34.4, Т 60.3, Т 60.4, Т 88.1
Wahdan, Shayma•T 34.4
Waitzmann, Moritz
Walbersloh, Jörg ST 7.1, ST 7.2,
ST 7.3, ST 7.4
Wald, Marcel T 2.5, •T 56.9
Walker, Rodney T 15.1
Walkowiak, Wolfgang T 26.5
Wall, Daniel•T 25.4
Wallis, Eckhard
Wallmann. C EP 5.2
Walpert, Daniel
Walter David •T 6.4
Walter Dominik •MP 4.4
Walter M EP 5 2
Walter Olga DD 50 1
Walther Cyrue T 35 1 • T 35 7
Walther, Cyrus
Wang Chichuci
Wang Dedeng
Wang Oian
wang, Hanyang 113.5, 140.8, 140.9
Wanke, Rainer 1 41./
Warmuth, Alexander
EP 12.4, •T 83.7
Warttmann, Erik•T 41.5
Warzel, Simone MP 4.6
Wäscher, Thomas•GR 10.4
Waßmer, Michael T 26.7, T 36.3,
T 76.8, T 88.7, T 88.8
Watson, Miriam T 60.5
Watson, Nigel
Weatherby, Thomas Sean •DD 16.1
Weber, Alena ST 6.4
Weber, Alfons
Weber Jannis •DD 35 4
Weber, Tom •AKPIK 2.9
Weidemann Lucas •T 63.4
Weidemüller Matthias ST 2 4 ST 2 2
Weigel Anna-Lena MD / 1
Weigel Hagen T 102 g
Weigl Stefan Andreas
Weiglein Georg T 21 / T 21 5 T 05 5
T 02 6
Woller David DD 40.1 DD 40.2
WEIIEL DAVID

Weingarten, Jens ST 1.1, ST 1.2, ST 1.6, ST 3.3, ST 3.4, ST 4.3, ST 4.4, ST 4.5, ST 6.2, ST 7.1, ST 7.2, ST 7.3, ST 7.4, ST 7.5, ST 9.4, T 13.6, T 53.3 Weinheimer, Christian T 21.4, T 21.5,
Veinstock, Lars Steffen•T 97.3
Weisspach, Anna
Weißler Biern ST 15
Weise Emanuel
Weitzel Ouirin T 41 1 T 59 2 T 97 5
Welberg Julia •DD 291
Welde Leonard T 3 9
Welsch Maximilian •T 84.6
Welzel-Breuer Manuela DD 21
DD 13.1. •DD 24.3
Wemmer, Florian
Wendland, Biörn •T 6.7, T 6.8, T 34.5.
T 70.8
Wendrich. Enrico•T 10.1
Wendt, Klaus T 102.6, DD 11.6,
DD 23.1, DD 23.2, DD 23.3, DD 30.2,
DD 34.10, DD 46.2, DD 49.3
Weniger, Christoph
Wennlöf, Håkan T 40.1, •T 40.7
Wenz, Daniel•T 22.4
Wermes, Norbert ST 6.1, T 13.5,
T 40.8, T 40.9, T 95.2
Werthebach, Johannes EP 11.8,
T 72.2, T 72.3, T 83.8, T 106.1,
•T 106.2, T 106.3
Werthmann, Clemens•T 2.9
Weßel, Antonia •T 103.7, T 103.9
Wesely, Bastian
Weßnigk, SusanneDD 6.1
Wessel, Christian
Westhoff, Susanne 1 34.2, 1 34.3
Westrich, Lukas
wex, Alexander 1 /4./, •1 /8./,
I 104.3, I 104.7 Wight Johannes ED 2.4
Wickert long ED 5.1
Wickert, Jells EP 3.1 Wickurch Christophor ED 4.2 ED 4.4
FD / 5 FD / 6 T 18 8 T / 6 1 T / 6 2
T 46 3 T 46 7 T 48 1 T 48 2 T 72 4
T 72 9 T 93 9 T 98 7 T 103 4
T 106 1 T 106 2 T 106 3 T 106 9
Wiedenbeck Sebastian •T 64.2
Wiedner Dirk T 14 1
Wiegelmann Thomas FP 12.8
•EP 12.9
Wiehe. Moritz ST 9.3
Wieland, Sebastian
Wiener, Jeff
Wiens, Lucas
Wiesemann, Marius T 2.4, T 90.2
Wiesinger, Christoph . T 21.7, •T 23.3,
•T 29.3, T 75.5, T 102.7
Wiesinger, Christph
Wilhelm, Sara•DD 7.1
Wilhelm Thomas DD 7 1 DD 10 1

DD 35.4. DD 35.5. DD 39.1	l w
Will, Martin	: W
Willers, Michael	2 X
Wilson, Matthew	
Wimberger, Sandro GR 6.2	2 X
Wimmer-Schweingruber, Robert	X
EP 10.1, T 81.1	Y
Windau, Michael•T 70.7	' Y
Windel, Hendrik T 68.6	6 Y
Wing, Matthew	- Y
Winkelmann, Jan DD 7.2, DD 17.4,	, Y
DD 24.2	
Winkens, Tobias DD 10.6	6 Y
Winter, Benedict T 61.2, T 61.3	; Y
Winter, Laura	- Y
Winter, Walter EP 7.2	2 Y
Winterhalder, Ramon•T 2.7	' Y
Wintersteller, Markus DD 10.10,	, Z
DD 10.11, DD 18.4	<u>Z</u>
Wintz, Johannes ST 2.3	8 <u>Z</u>
Wirth, Rosmarie•T 20.6	2 <u>Z</u>
Wissing, Jan-Maik EP 5.5	
Witola, Lukas•T 96.6	2 <u>Z</u>
Wittbrodt, Jonas	
Witte, Wanda DD 44.2	2 Z
Witthaus, Lucas •1 /2.2, 1 /2.3	· _
Wittig, Florian . 1 65.5, 1 65.6, 1 65.7,	, 4
•1 94.4	
Wittweg, Christian	
Witzel, Uliver	
Wiochal, Michael 1 65.2, 1 65.3	
Wochele, Doris	
Wodzinaki Dita	
WODZINSKI, RITA DD 11.8, DD 35.9,	, 4
Weeste Inge DD 12.2 DD 12.4	
NOESIE, IIIgaDD 12.2, •DD 12.4,	, 2
Wöhlko Corino DD 91 DD 42.2	
12 2 42.3	, 2
Wohllohon Fradarik T 51 9	
Woithe Julia	
Wolf Moritz	
Wolf Pascal	
Wolf Poger T 15 6 T 25 1 T 25 2	' -
T 36 3 T 36 4 T 36 5	' 7
Wolf Tim Michael Heinz •T 75 9	, ~
Wondrak Michael Florian •GR 16	7
Wong Ming Liang ST 1.5	
Wongel Alicia T 10.3	
Wonsak Biörn T 102 1	
Wozniewski. Sebastian	Ż
•T 82.4	' z
Wronska, Aleksandra ST 1.5. ST 3.5	; z
Wuchterl. Sebastian	l z
Wulff, Johan•T 46.4	Ż
Wulff, Jörg ST 1.6	5 z
Wulff, Peter	; z
Wunderlich, Fabian EP 2.8. EP 3.2	:
Wundheiler, Brian AKPIK 4.8	: z
Wunsch, Stefan	2 -
Wurm, Michael T 48.5, T 59.5, T 99.4	, z
T 99.9, T 102.2, T 102.9	Z
Wutzig, Michael EP 5.7	' Z
Wyngaard, Jane	2 Z

wynine, ounies
Wysotzki, Christian•T 48.3
XENON-Kollaboration . T 22.2, T 75.9
T 78.1, T 22.3, T 75.8
Xie, Zhaoheng ST 9.2, •ST 10.1
Xu, Zigong EP 10.1, T 81.1
Yang, Qian
Yang, Yewon•T 16.1, T 58.1
Yap, Yee Chinn
Ye, Hua T 12.1, T 12.2
Ye, Yuting EP 4.3, EP 4.4, EP 4.5,
EP 4.6
Yeo, Beomki T 16.7
Yoon, Peter
Young, Christopher T 8.1, •T 28.1
Yu, Boyang•T 53.1, T 53.2
Yu, Jia EP 10.1, T 81.1
Zacharias, Michael
Zacik, Peter•T 24.4, T 52.2
Zahid, Awais Bin•T 76.9
Zaleski, Shawn T 67.8, •T 67.9
Zambanini, André
Zanderighi, Giulia
Zanoli, Silvia•T 90.2
Zantis, Franz-Peter . GR 14.4, GR 14.5
GR 14.6
Zanzi, Daniele•T 28.4
Zappa, Francesco GR 9.4, •GR 9.7
Zaunick, Hans-Georg ST 9.1, ST 9.6
Zaytsev, Alexander
Zech. Andreas EP 7.5
Zeier, Robert MP 5.2
Zema, Vanessa T 104.3, T 104.6
Zhang, Chao•T 101.7
Zhang, Shenyi EP 10.1, T 81.1
Zhang, Sinuo•T 40.5
Zhao, Suting MP 1.2
Zhong, Tiancheng•T 96.3
Zhou, Jennifer ST 2.1
Zhukov, Valery T 57.8
Ziegler, Felix
Ziegler, Julia•T 36.8
Ziemons, Tim T 65.1, T 65.2, T 65.3
Ziemons, Tim T 65.1, T 65.2, T 65.3, T 65.4, •T 94.6
Ziemons, Tim T 65.1, T 65.2, T 65.3, T 65.4, •T 94.6 Zierke, Simon EP 4.3, EP 4.4, •EP 4.5
Ziemons, Tim T 65.1, T 65.2, T 65.3, T 65.4, •T 94.6 Zierke, Simon EP 4.3, EP 4.4, •EP 4.5 EP 4.6, T 106.9
Ziemons, Tim T 65.1, T 65.2, T 65.3, T 65.4, •T 94.6 Zierke, Simon EP 4.3, EP 4.4, •EP 4.5 EP 4.6, T 106.9 Zimermmane-Santos, Andre •T 33.1
Ziemons, Tim T 65.1, T 65.2, T 65.3, T 65.4, •T 94.6 Zierke, Simon EP 4.3, EP 4.4, •EP 4.5 EP 4.6, T 106.9 Zimermmane-Santos, Andre T 33.1 Zimmermann, Tim
Ziemons, Tim T 65.1, T 65.2, T 65.3, T 65.4, •T 94.6 Zierke, Simon EP 4.3, EP 4.4, •EP 4.5, EP 4.6, T 106.9 Zimermmane-Santos, Andre •T 33.1 Zimmermann, Tim
Ziemons, Tim T 65.1, T 65.2, T 65.3, T 65.4, •T 94.6 Zierke, Simon EP 4.3, EP 4.4, •EP 4.5, EP 4.6, T 106.9 Zimermane-Santos, Andre •T 33.1 Zimmermann, Tim GR 6.2 Zink, Adrian
Ziemons, Tim T 65.1, T 65.2, T 65.3, T 65.4, •T 94.6 Zierke, Simon EP 4.3, EP 4.4, •EP 4.5 EP 4.6, T 106.9 Zimermane-Santos, Andre •T 33.1 Zimmermann, Tim GR 6.2 Zink, Adrian •T 51.8 Zinßer, Joachim •T 69.8 Zirnbauer, Martin •MP 3.3.
Ziemons, Tim T 65.1, T 65.2, T 65.3, T 65.4, •T 94.6 Zierke, Simon EP 4.3, EP 4.4, •EP 4.5 EP 4.6, T 106.9 Zimermmane-Santos, Andre •T 33.1 Zimmermann, Tim GR 6.2 Zink, Adrian •T 69.8 Zirnbauer, Martin •MP 3.3 Zmija, Andreas T 51.8
Ziemons, Tim T 65.1, T 65.2, T 65.3, T 65.4, •T 94.6 Zierke, Simon EP 4.3, EP 4.4, •EP 4.5, EP 4.6, T 106.9 Zimmermane-Santos, Andre •T 33.1 Zimmermann, Tim GR 6.2 Zink, Adrian T 51.8 Zirnbauer, Martin •MP 3.3 Zmia, Andreas T 51.8 Zoch, Knut T 33.2
Ziemons, Tim T 65.1, T 65.2, T 65.3, T 65.4, •T 94.6 Zierke, Simon EP 4.3, EP 4.4, •EP 4.5, EP 4.6, T 106.9 Zimermmane-Santos, Andre T 33.1 Zimk, Adrian SR 6.2 Zink, Adrian T 51.8 Zinßer, Joachim F 69.8 Zimiga, Andreas T 51.8 Zoch, Knut T 33.3 Zöggeler, Marion DD 42.2
Ziemons, Tim T 65.1, T 65.2, T 65.3, T 65.4, •T 94.6 Zierke, Simon EP 4.3, EP 4.4, •EP 4.5, EP 4.6, T 106.9 Zimermmane-Santos, Andre T 33.1 Zimmermann, Tim GR 6.2 Zink, Adrian T 51.8 Zinker, Joachim F 69.8 Zinbauer, Martin MP 3.3 Zmija, Andreas T 51.8 Zodgeler, Martin DD 42.2 Zol, Irene T 10.1, T 12.9
Ziemons, Tim T 65.1, T 65.2, T 65.3, T 65.4, •T 94.6 Zierke, Simon EP 4.3, EP 4.4, •EP 4.5, EP 4.6, T 106.9 Zimermmane-Santos, Andre T 33.1 Zimmermann, Tim GR 6.2 Zinker, Joachim T 69.8 Zinker, Joachim T 69.8 Zinkauer, Martin MP 33.3 Zmija, Andreas T 51.8 Zoch, Knut T 33.3 Zöggeler, Martion DD 42.2 Zoi, Irene T 10.1, T 12.9 Zoska, M. EP 5.2
Ziemons, Tim T 65.1, T 65.2, T 65.3, T 65.4, •T 94.6 Zierke, Simon EP 4.3, EP 4.4, •EP 4.5, EP 4.6, T 106.9 Zimermmane-Santos, Andre T 33.1 Zimmermann, Tim GR 6.2 Zink, Adrian T 51.8 Zink, Adrian T 51.8 Zink, Adrian T 51.8 Zinka, Andreas T 51.8 Zoch, Knut T 33.3 Zöggeler, Marion DD 42.2 Zoi, Irene T 10.1, T 12.9 Zoska, M. EP 5.2 Zotz, Alexander T 86.9
Ziemons, Tim T 65.1, T 65.2, T 65.3, T 65.4, •T 94.6 Zierke, Simon EP 4.3, EP 4.4, •EP 4.5, EP 4.6, T 106.9 Zimermmane-Santos, Andre •T 33.1 Zimmermann, Tim GR 6.2 Zink, Adrian T 51.8 Zirnbauer, Martin •MP 3.3 Zmia, Andreas T 51.8 Zoch, Knut T 33.3 Zöggeler, Marion •D 42.2 Zoi, Irene T 10.1, T 12.9 Zosta, Alexander •T 84.5 Zuber, Kai T 49.5
Ziemons, Tim T 65.1, T 65.2, T 65.3, T 65.4, •T 94.6 Zierke, Simon EP 4.3, EP 4.4, •EP 4.5, EP 4.6, T 106.9 Zimermmane-Santos, Andre T 33.1 Zimmermann, Tim GR 6.2 Zink, Adrian T 51.8 Zinker, Joachim F 69.8 Zirnbauer, Martin MP 3.3 Zöggeler, Martin DD 42.2 Zoi, Knut T 10.1, T 12.9 Zoska, M. EP 5.2 Zotz, Alexander T 8.6 Zuber, Kai T 49.5 Züfle, Jan
Ziemons, Tim T 65.1, T 65.2, T 65.3, T 65.4, •T 94.6 Zierke, Simon EP 4.3, EP 4.4, •EP 4.5, EP 4.6, T 106.9 Zimermmane-Santos, Andre T 33.1 Zimk, Adrian T 51.8 Zinker, Joachim F 09.8 Zinkgr, Joachim MP 33.3 Zimbauer, Martin MP 33.3 Zöggeler, Marton DD 42.2 Zoi, Irene T 101, T 12.9 Zoska, M. EP 5.2 Zutz, Alexander T 8.6 Zuifle, Jan T 20.3, T 20.9, T 48.6, T 102.3
Ziemons, Tim T 65.1, T 65.2, T 65.3, T 65.4, •T 94.6 Zierke, Simon EP 4.3, EP 4.4, •EP 4.5, EP 4.6, T 106.9 Zimermmane-Santos, Andre T 33.1 Zimmermann, Tim GR 6.2 Zink, Adrian T 51.8 Zinker, Joachim T 69.8 Zinker, Joachim T 69.8 Zinkauer, Martin MP 33.3 Zöggeler, Martin D42.2 Zoi, Irene T 10.1, T 12.9 Zoska, M. EP 5.2 Zotz, Alexander T 8.6 Zuife, Jan T 20.3, T 20.9, T 48.6 T 48.8, T 102.3 Zumalacarregui, Miguel Zumalacarregui, Miguel GR 5.3
Ziemons, Tim T 65.1, T 65.2, T 65.3, T 65.4, •T 94.6 Zierke, Simon EP 4.3, EP 4.4, •EP 4.5, EP 4.6, T 106.9 Zimermmane-Santos, Andre T 33.1 Zimmermann, Tim GR 6.2 Zink, Adrian T 51.8 Zinker, Joachim T 69.8 Zirnbauer, Martin MP 3.3 Zöggeler, Marton D 42.2 Zoi, Irene T 10.1, T 12.9 Zoska, M. EP 5.2 Zotz, Alexander T 84.9, T 48.6 Zühle, Jan T 20.3, T 20.9, T 48.6 T 48.8, T 102.3 Zumalacarregui, Miguel Zundel, Dorina T 48.9, T 48.0
Ziemons, Tim T 65.1, T 65.2, T 65.3, T 65.4, •T 94.6 Zierke, Simon EP 4.3, EP 4.4, •EP 4.5, EP 4.6, T 106.9 Zimermmane-Santos, Andre T 33.1 Zimmermann, Tim GR 6.2 Zink, Adrian T 51.8 Zinkr, Joachim G 9.8 Zirnbauer, Martin MP 3.3 Zöggeler, Marion DD 42.2 Zoi, Irene T 10.1, T 12.9 Zoska, M. EP 5.2 Zotz, Alexander T 8.6 Zuifle, Jan T 20.3, T 20.9, T 48.6 T 48.8, T 102.3 Zundel, Dorina T 48.9, T 48